

# Mapping the landscape of gravity theories

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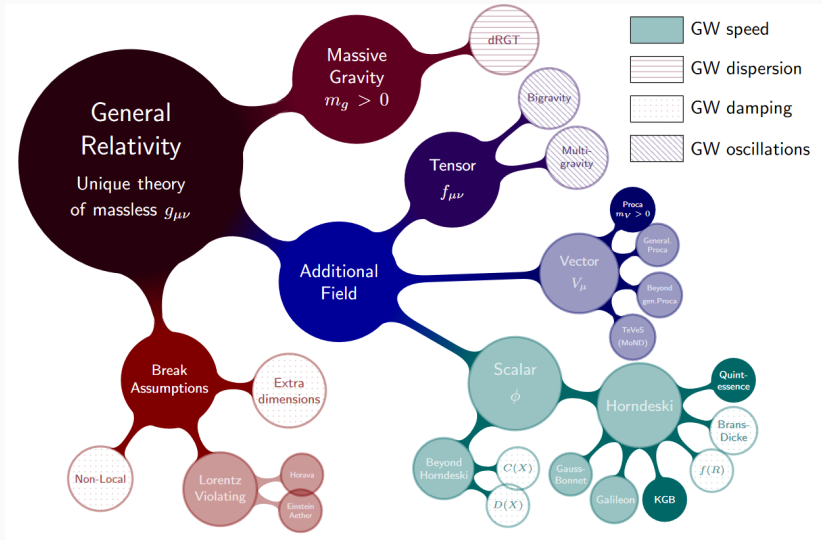
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Dark Matter & Stars, Lisbon

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# (A portion of) the landscape



Source: <sup>1</sup>

<sup>1</sup>J. M. Ezquiaga, M. Zumalacarregui, Front. Astron. Space Sci. 5:44 (2018) - arXiv:1807.09241

Explore generalisations of GR



Understand GR as a special case in a broader framework



Attempt to overcome limitations of GR (and learn about universe in the process)

# Thermodynamics and gravity

- Thermodynamics  $\Leftrightarrow$  gravity?
- GR  $\Leftrightarrow$  thermodynamical equilibrium <sup>2</sup>  
Modified gravity  $\Leftrightarrow$  non-equilibrium <sup>3</sup>
- "Thermodynamics of gravitational theories"
- **Question:** what exactly is the dissipative process leading to equilibrium from non-equilibrium?
- **Question:** what is the order parameter measuring closeness to equilibrium?

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<sup>2</sup>T. Jacobson, PRL 75 (1995) - arXiv:gr-qc/9504004

<sup>3</sup>C. Eling, R. Guedens, T. Jacobson, PRL 96 (2006) - arXiv:gr-qc/0602001

# The plan

- Take **scalar-tensor** gravity, containing  $f(\mathcal{R})$  as subclass <sup>4</sup>
- Model scalar contribution as effective **dissipative** fluid <sup>5</sup>
- Apply **irreversible non-equilibrium** thermodynamical description, such as Eckart's <sup>6</sup>
- Extract thermodynamical quantities such as temperature and viscosity
- Understand dissipative process

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<sup>4</sup>T. Sotiriou & V. Faraoni, Rev. Mod. Phys 82 (2010) - arXiv:0805.1726

<sup>5</sup>V. Faraoni & J. Coté, PRD 98 (2018) - arXiv:1808.02427

<sup>6</sup>C. Eckart, Phys. Rev. 58 (1940)

# Imperfect fluid description of modified gravity

- Brans-Dicke scalar-tensor action (Jordan frame) ( $G_{\text{eff}} = 1/\phi$ )

$$S_{\text{ST}} = \frac{1}{16\pi} \int d^4x \sqrt{-g} \left[ \phi \mathcal{R} - \frac{\omega(\phi)}{\phi} \nabla^c \phi \nabla_c \phi - V(\phi) \right] + S^{(m)}$$

- Effective Einstein equations

$$G_{ab} = 8\pi G_{\text{eff}} T_{ab}^{(m)} + 8\pi T_{ab}^{(\phi)}$$

- Study thermodynamics of imperfect fluid

$$T_{ab}^{(\phi)} = \rho^{(\phi)} u_a u_b + q_a^{(\phi)} u_b + q_b^{(\phi)} u_a + \Pi_{ab}^{(\phi)}$$

- $\nabla_a \phi$  timelike + future-oriented  $\Rightarrow$  natural fluid interpretation

- 4-velocity  $u^a = \frac{\nabla^a \phi}{\sqrt{-\nabla^e \phi \nabla_e \phi}}$  with  $u^a u_a = -1$

# Eckart's thermodynamics

$$P_{\text{vis}}^{(\phi)} = -\zeta \theta$$

$$q_a^{(\phi)} = -\mathcal{K} (h_{ab} \nabla^b \mathcal{T} + \mathcal{T} \dot{u}_a)$$

$$\pi_{ab}^{(\phi)} = -2\eta \sigma_{ab}$$

- **Non-equilibrium** (irreversible) thermodynamics
- "1st order": simplest (linear) assumptions to satisfy  $S_{;\alpha}^{\alpha} \geq 0$
- **Relativistic** generalisations of Stokes' law, Fourier's law and Newton's law of viscosity

# Temperature and viscosity of $\phi$ -fluid

- Comparing  $\dot{u}_a$  and  $q_a^{(\phi)}$

$$q_a^{(\phi)} = -\frac{\sqrt{-\nabla^c \phi \nabla_c \phi}}{8\pi\phi} \dot{u}_a \quad \Rightarrow \quad \mathcal{KT} = \frac{\sqrt{-\nabla^c \phi \nabla_c \phi}}{8\pi\phi} \geq 0$$

- Temperature "of scalar-tensor gravity", relative to GR <sup>7</sup>
- GR limit:  $\mathcal{KT} \rightarrow 0$  when  $\phi = \text{const.} \Rightarrow G_{\text{eff}} = G_{\text{N}}$ , no  $\phi$ -fluid

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<sup>7</sup>V. Faraoni, A. Giusti, PRD 103 (2021) - arXiv:2103.05389



# Fixed points of dissipation equation

- Approach to equilibrium governed by **effective heat equation** for  $\phi$ -fluid

$$\frac{d(\mathcal{KT})}{d\tau} = 8\pi(\mathcal{KT})^2 - \theta\mathcal{KT} + \frac{\square\phi}{8\pi\phi}$$



$$\mathcal{KT} = 0$$

GR + theories with  
**non-dynamical  $\phi$** <sup>8</sup>



$$\mathcal{KT} = \text{const.}$$

Scalar-tensor  
**stealth solutions**<sup>9</sup>

- GR seems to be the only stable equilibrium!

<sup>8</sup>V. Faraoni, A. Giusti, S. Jose, S. Giardino, PRD 106 (2022) - arXiv:2206.02046

<sup>9</sup>S. Giardino, A. Giusti, V. Faraoni - arXiv:2302.08550

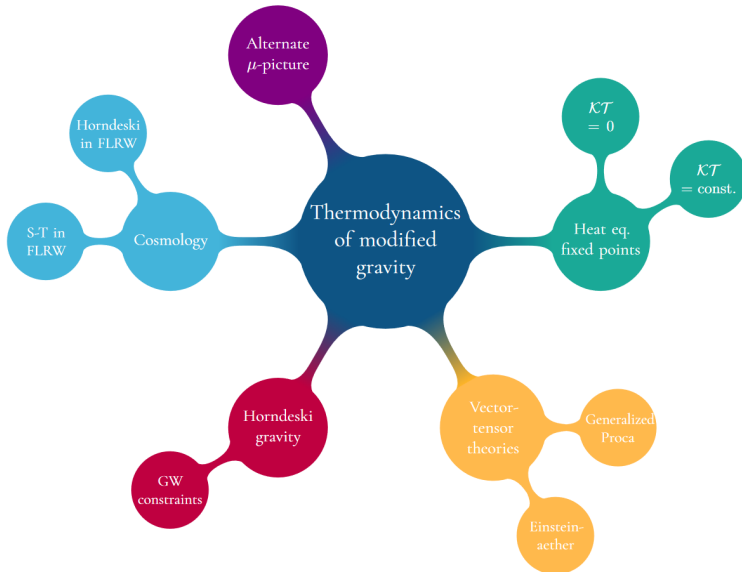
# Mapping the landscape

- Thermodynamical analogy  $\Rightarrow$  map of gravity theories landscape
- GR: special role as the **only stable equilibrium** at  $\mathcal{KT} = 0$
- Modified theories with additional dynamical degrees of freedom always have  $\mathcal{KT} > 0$
- Nordström gravity with less degrees of freedom than GR has  $\mathcal{KT} < 0$  (pathological)
- Dissipative relaxation process to GR in most cases, but not guaranteed

# Intuitive picture



# Further developments



# Summary

- Connection between **thermodynamics** and **gravity**
- Jacobson's idea: GR equilibrium state, modified gravity **non-equilibrium**  $\Rightarrow$  open questions
- **New perspective**: effective fluid approach to S-T theories + Eckart's non-equilibrium thermodynamics  $\rightarrow \mathcal{T}$  of modified gravity, relative to GR
- Mapping the landscape of gravity theories with goal of understanding GR in more general framework
- $\mathcal{KT} > 0$  whenever additional dynamical dof
- GR special as the **only stable equilibrium** state!

Thank you for your attention!