

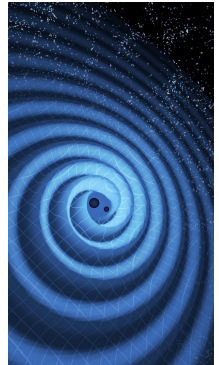
Searching higher-derivative extensions to GR in black holes' ringdown

Simon Maenaut

Pablo Cano, Thomas Hertog,
Kwinten Fransen, Tjonnie Li
[based on arXiv:2110.11378]

BE-NL GW-Meeting 2022

KU LEUVEN



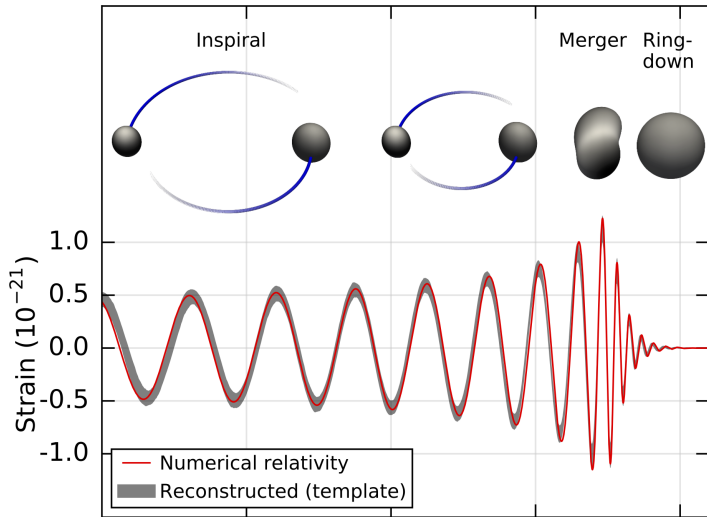
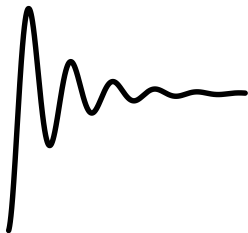


Figure: Theoretical signal and phases of a binary black hole merger [LIGO Collaboration]

Ringdown frequencies

- Perturbations of black hole decay over time
- GWs come from space around the black hole
- Fluctuations of a damped harmonic oscillator
- Boundary conditions set a dissipative system
- Resonance modes have complex frequencies



$$\sum A_k \exp(i\omega_k t)$$

Higher derivative Gravity

- Extend GR as an effective field theory of higher derivatives
- An approximation to study phenomena at one energy scale
- The complete theory at a higher energy scale is unknown

$$S = \int d^4x \sqrt{g} \left[R + \sum_{n=2}^{\infty} \ell^{2n-2} \mathcal{L}_{(n)} \right]$$

$$\mathcal{L}_{(3)} \ell^4 = \lambda_{\text{ev}} R_{\mu\nu}{}^{\rho\sigma} R_{\rho\sigma}{}^{\delta\gamma} R_{\delta\gamma}{}^{\mu\nu}$$

Corrections from Cubic Curvature Term

- Corrections to GR are expected to be suppressed ($\lambda \ll 1$)
- GWs are perturbations of the background spacetime ($\varepsilon \ll 1$)
- Metric perturbations with odd and even parity and a range of l
- Compute the fundamental resonance frequencies for small λ
- Represent leading-order corrections of cubic curvature term

$$g_{\mu\nu}^{(\lambda)} = \bar{g}_{\mu\nu}^{(\lambda)} + \varepsilon h_{\mu\nu}^{(\lambda)} + \mathcal{O}(\varepsilon^2, \lambda^2)$$

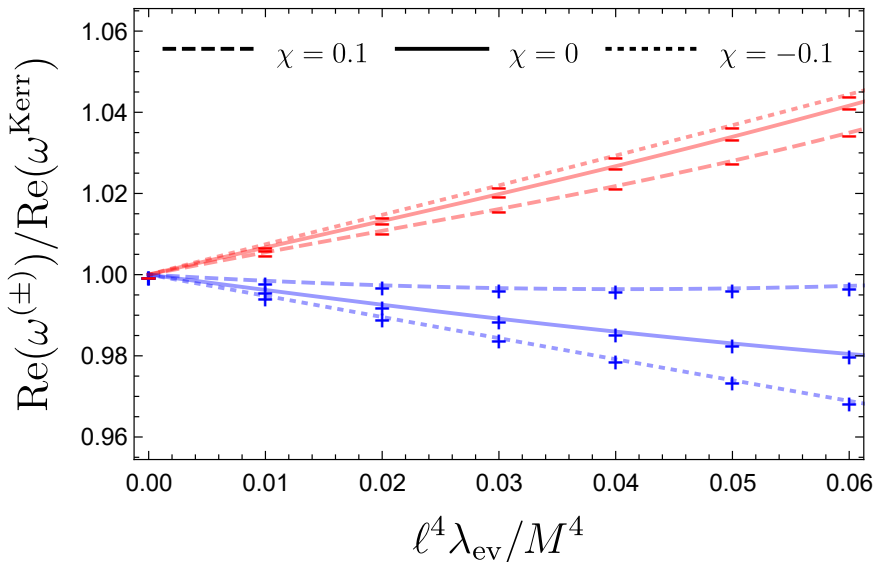


Figure: Fundamental QNM frequencies $l = m = 2$ relative to Kerr [arXiv:2110.11378]

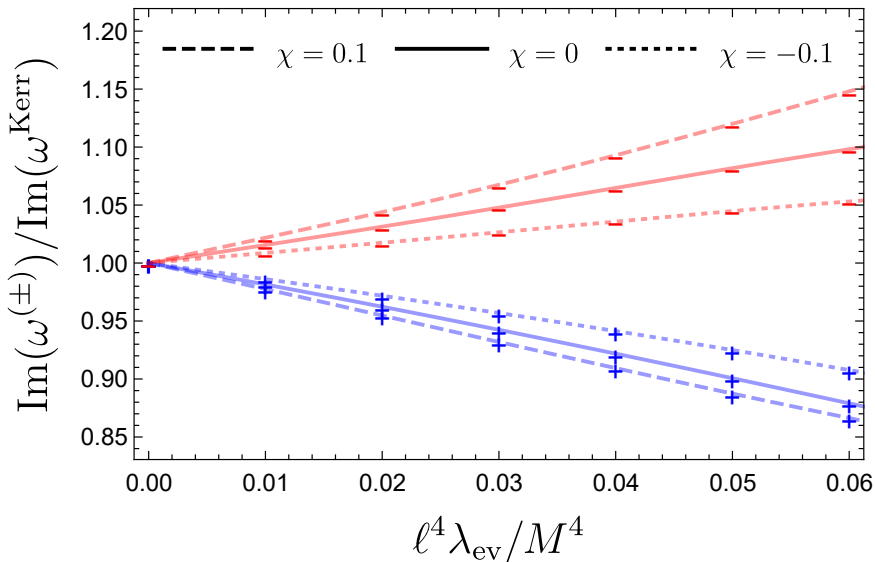


Figure: Fundamental QNM frequencies $l = m = 2$ relative to Kerr [arXiv:2110.11378]

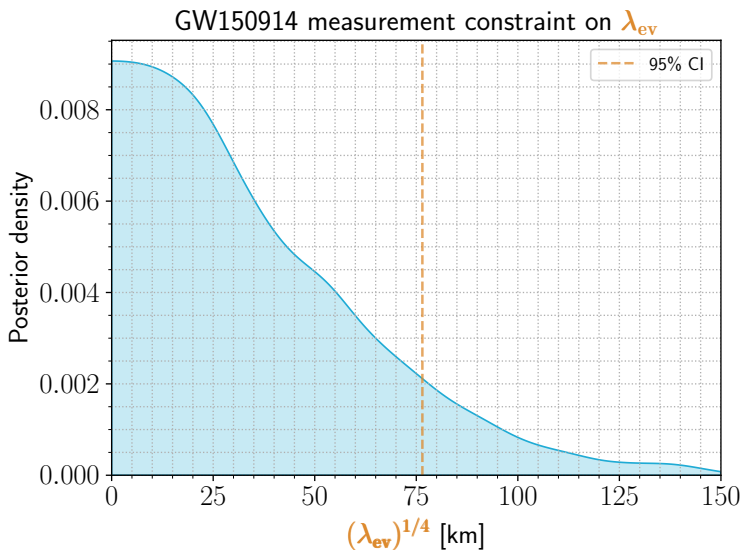


Figure: Calculated constraints on λ_{ev} from the ringdown analysis of GW150914

Future considerations

- Add parity-breaking higher-derivative terms
- Consider scalar field coupling (EdGB, dCS)
- Include possible inspiral and merger effects
- Frequency splitting as beyond GR signature
- Simulate sensitivity for future GW detectors

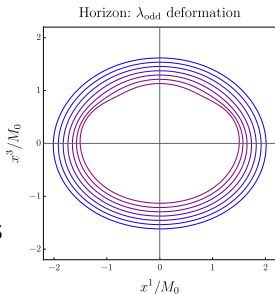
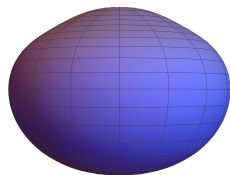


Figure: Example of the black hole horizon surface in a parity-breaking theory [Cano & Ruipérez 2019]

Future considerations

- Add parity-breaking higher-derivative terms
- Consider scalar field coupling (EdGB, dCS)
- Include possible inspiral and merger effects
- Frequency splitting as beyond GR signature
- Simulate sensitivity for future GW detectors

Thank you for your attention!

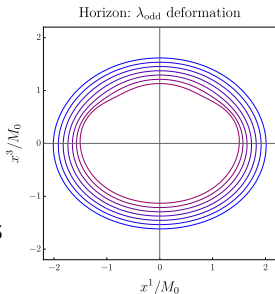
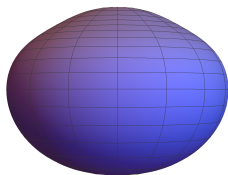


Figure: Example of the black hole horizon surface in a parity-breaking theory [Cano & Ruipérez 2019]