Development and Testing of Composite Vacuum Tubes for Einstein Telescope

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Vacuum Tubes for ET

- * Total vacuum pipe requirement :
 - 3 detectors
 - 2 interferometers per detector
 - 2 arms per interferometer
 - Each arm is 10 km

=> 3 x 2 x 2 x 10 km = 120 km

- Ultra High Vacuum (UHV) is required,
 - Reduce the noise due to excess gas density fluctuations along the beam path
 - Reduction of test mass motion and thermal isolation
 - Preserving cleanliness of optical elements









Vacuum Tubes of Current Gravitational Wave Detectors





- Stainless steel is the standard due to,
 - UHV compatibility and mechanical integrity
- With diameter up to 1 m and 120 km of vacuum pipes, Einstein Telescope requires,
 - Total volume ~ 100,000 m^3 (~ 10 times greater than LIGO)
 - Total surface area ~ 800,000 m²
 - 4 mm thick stainless steel required ~ 100 kg/m (12000 tons in total)
- Decreasing of stainless steel would help in,
 - Reduction of vacuum firing cost (reduce H₂ outgassing)
 - Help in logistics of moving and assembly underground





Composite Vacuum Tube - First Prototype

- Composite Vacuum Tubes : metal liner + fiber composite have the potential of
 - Stainless steel reduction by a factor 10 (e.g., 4 to 0.4 mm)
 - Lighter tubes which help in logistics of moving and assembly underground
 - On-site production







- Metal liner thickness 0.8 mm
- Composite 6 mm



Manufacturing of the First Prototype

- Stainless Steel Liner (304L, t = 0.8 mm)
 - Laser structuring of the liner : 20 µm deep groves
 - To improve the bonding between the liner and the composite material







Manufacturing of the First Prototype

- Composite material Glass Fiber Reinforced Plastic (GFRP)
 - Glass fiber with epoxy matrix (circumferential winding)
 - Curing Temperature : 120° C (limits the bake-out temp)





Vacuum System and Testing



UHV

gauge

Pressure



The vacuum test :

- Helium leak test
- With turbo molecular pump after baking the system for 7 days at 100° C

 $P = 7 \times 10^{-10} \text{ mbar}$

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Residual Gas Spectrum





Delamination of the First Prototype

- Operation for more than 6 months under vacuum and survived multiple bake-outs (100 °C)
- * The liner got delaminated from fiber composite tube
- Partial destruction of laser structures while rolling the stainless-steel sheet into a tube



Initial position of the GFRP tube

Position after delamination



Possible Reasons for the Delamination

 Mismatch of coefficient of thermal expansion (CTE) of SS304L and fiber composite

Material	Typical CTE (μm / m °C)	
SS304L	17.3	
Fiber Composite	7.1 (along fibers) 31.3 (transverse to fibers)	



Initial position of the GFRP tube

Position after delamination



Possible Reasons for the Delamination

 Differential thermal expansion due to temperature gradient along and across the composite vacuum tube



- Temperature of 120 °C at the stainless steel ends
- Insulation using mineral wool of 15 cm
- Convection boundary condition with coefficient of 5 Wm⁻² °C⁻¹





Conclusions

- * The first prototype was under vacuum for more than 6 months
- * The fiber composite got delaminated from the stainless-steel liner
- Choice of composite :
 - Epoxy with randomly oriented fibers to make it isotropic and matching CTE with metal
- Simulation :
 - Minimize the material budget
 - Scaling of composite vacuum tube for ET dimensions
- Tests :
 - Shearing tests of chosen composites along with heat cycling test
 - UHV compatibility of composite tubes
 - Overpressure tests (safety margin)
- Design a heating strategy



Buckled Tube





Thank You!

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Appendix





Ultra High Vacuum(UHV) Requirement

- Reduce the noise due to excess gas density fluctuations along the laser beam path
- Reduction of test mass motion and thermal isolation
- Preserving cleanliness of optical elements
- Hydrogen (1 x 10⁻¹⁰ mbar) :
 - From the bulk during stainless steel production
- Water (< 5 x 10⁻¹¹ mbar) :
 - Moisture from air
 - Removed using heating/bake-out
- Hydrocarbons (< 10⁻¹⁴ mbar) :
 - Deposited on the surface e.g., oil during machining
 - Careful cleaning using detergents and assembly



1 order sensitivity increase requires 1 order improvement in vacuum



Vacuum Tubes of Current Gravitational Wave Detectors

LIGO, USA



Credits: Caltech/MIT/LIGO Laborator

- Material : Stainless Steel 304L
- Diameter : 1.2 m
- Arm length : 4 km each
- Achieved pressure : 2 x 10⁻⁹ mbar

Virgo, Italy



Credits: The Virgo Collaboration

- Material : Stainless Steel 304L
- Diameter : 1.2 m
- Arm length : 3 km each
- Achieved pressure : 1 x 10⁻⁹ mbar

KAGRA, Japan



Credits: KAGRA Observatory, ICRR, The University of Tokyo

- Material : Stainless Steel 304L
- Diameter : 0.8 m
- Arm length : 3 km each
- Achieved pressure : 1 x 10⁻⁷ mbar



History of Composite Vacuum Tubes at CERN





Critical Buckling Pressure (16.0 ± 0.5 bar)







Original Tube

Buckled Tube

 Only the SS304 liner bucked and the GFRP tube stayed intact





Hoop/Circumferential Stress in a Cylinder





- Hoop/Tangential stress : $\sigma_{\theta}(r) = \frac{p_0 r_0^2}{r_0^2 r_1^2} \left(\frac{r_i^2}{r^2} + 1\right)$ as d/t <20 (thick tube)
- Yield Stress = 205 MPa

Design	Thickness (mm)	Hoop Stress max at atmospheric pressure (MPa)
Prototype 1.0	0.8 + 6.08	- 0.88



- As the hoop stresses are compressive, epoxy matrix should support the load and prevent the E-glass fibers from buckling
- * Longitudinal Compressive strength of E-glass Epoxy ($V_f = 0.6$) = 620 MPa
- Typical Compressive strength of an unmodified epoxy = 70 MPa



Overpressure Test

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- Overpressure test was performed in order to
 - Verify/set a baseline for the simulation results
 - Effect of the delamination





