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ECFA HTE meeting on Z pole physics

September 23, 2022

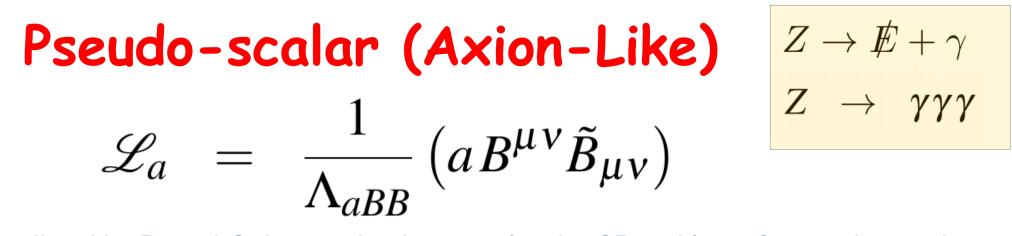
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Exotic Z decays

Sensitivity on BSM physics at the Z-pole: Giga-Z and Tera-Z factories

- correspond to several <u>BSM physics scenarios</u> : different final states (2,3,4 particles)
- benchmark models required (most inspired by dark-matter and dark-sector scenarios)
- axion-like particle (ALP)
- light-vector X coupled to anomalous currents
- magnetic inelastic DM, anomalous magnetic moment of neutrinos
- dark-photon (massive, massless scenarios)
- (light) massive spin-2 particles effectively coupled to SM sector
- Rare decays in SM particles (**ggg**, **gg+γ**, **γγγ**) sensitive to NP.

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QCD axions predicted by Peccei-Quinn mechanism to solve the CP problem of strong interactions

Axion-like particles (ALP) as portal connecting DM with SM sector

ultralight ALP as DM candidate

ullet couplings also with fermions $\partial_\mu a ar{\psi} \gamma^\mu \psi$ but chirally suppressed

ALP predicted by many UV theories (string-theory, SUSY, etc)

Z–γ–a mixed vertex induced from SM hypercharge B-field after SM SSB

Two main characteristic signatures at Z-pole : 2-body and 3-body final states

$$\Gamma(a \to \gamma \gamma) = \frac{1}{64\pi} \frac{1}{\Lambda_{aBB}^2} \cos \theta_w^4 m_a^3$$

$$\Gamma(Z \to \gamma a) = \frac{1}{96\pi} \frac{1}{\Lambda_{aBB}^2} \cos \theta_w^2 \sin \theta_w^2 m_Z^3 \left(1 - \frac{m_a^2}{m_Z^2} \right)^3$$

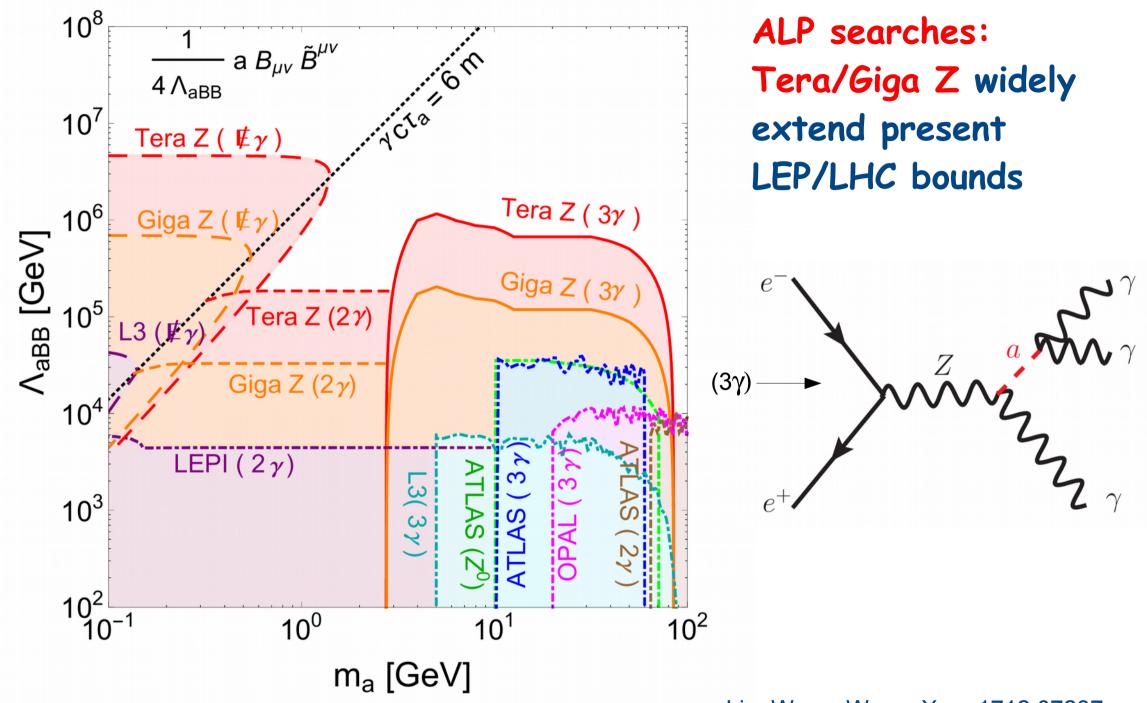
2-bodies

 $Z \rightarrow \gamma + \not{\!\!\!E}$ ALP as LLP detected as missing energy $- \text{light ALP decaying in } \gamma \gamma$ outside detector $- \text{ or } \mathbf{a} \rightarrow \chi \chi (\chi \text{ invisible DM})$ **3-bodies**

 $Z \rightarrow \gamma \gamma \gamma$

ALP decaying inside detector

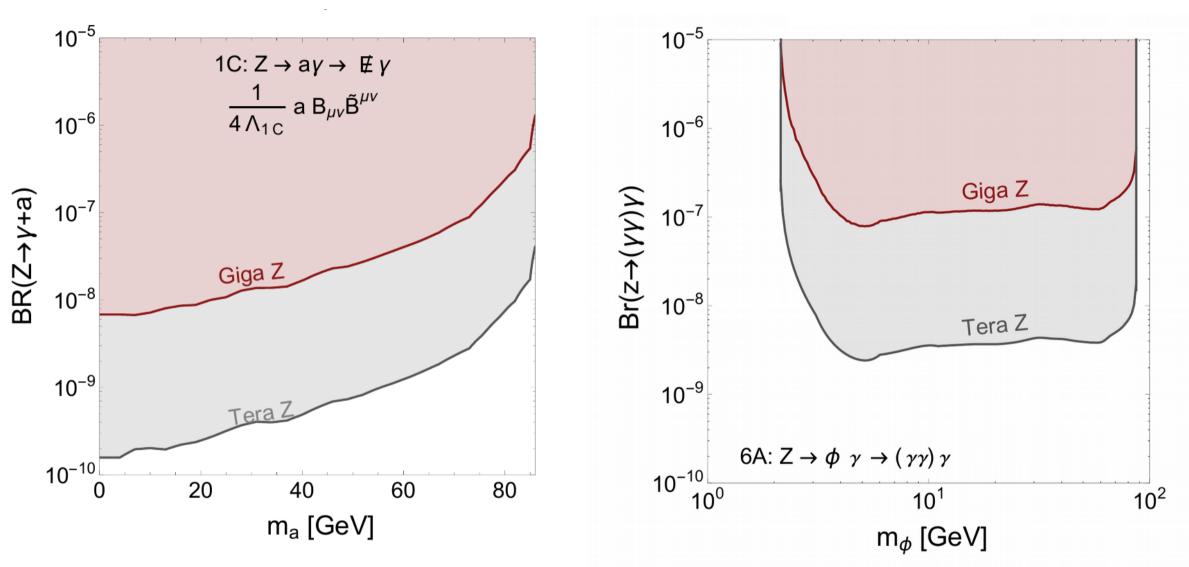
<u>3 \rightarrow 2-bodies</u> $Z \rightarrow \gamma \gamma$ if a-mass too small to resolve 2 photons ECFA HTE meeting on Z pole physics, Sep 23, 2022 4



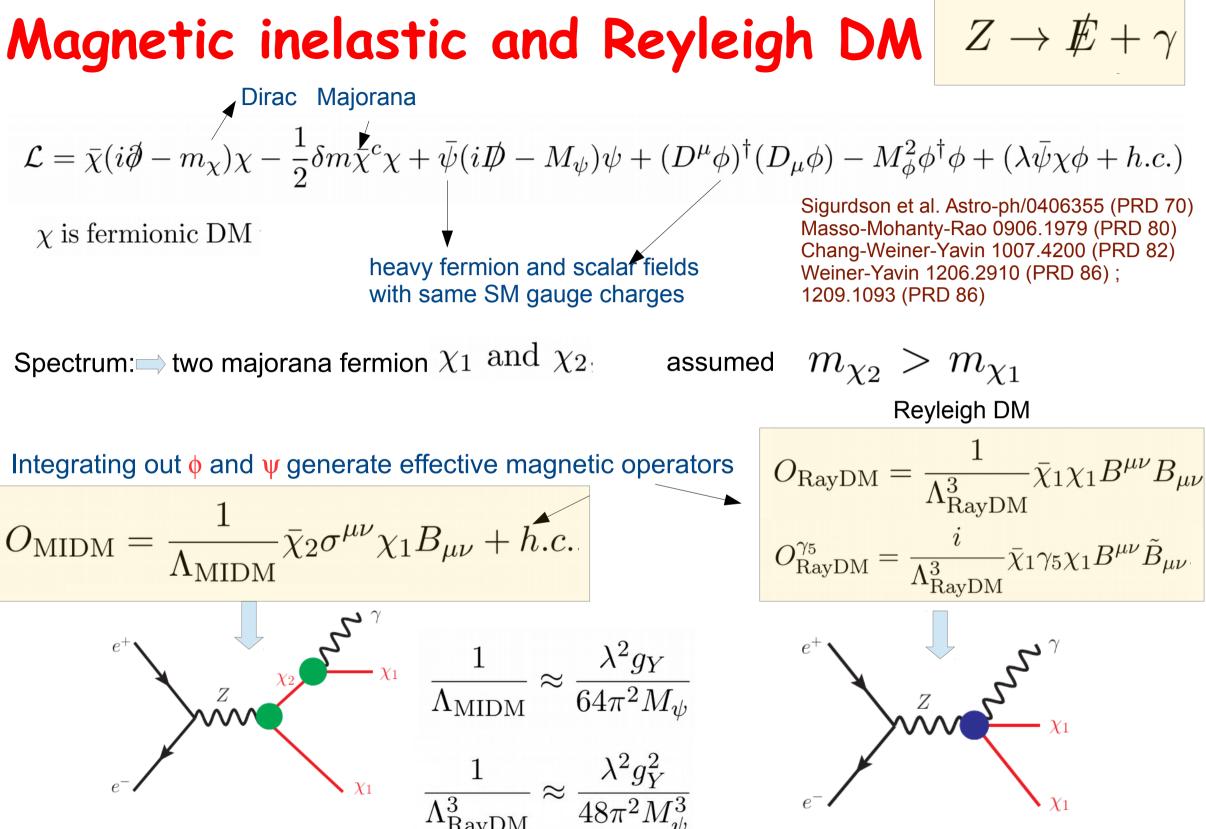
Liu, Wang, Wang, Xue, 1712.07237

 $(2\gamma) \rightarrow$ refers to the case in which the mass of ALP is too small to resolve 2-photons seen as \rightarrow 1 photon E. Gabrielli ECFA HTE meeting on Z pole physics, Sep 23, 2022 5

 $Z \to \gamma \gamma \gamma$



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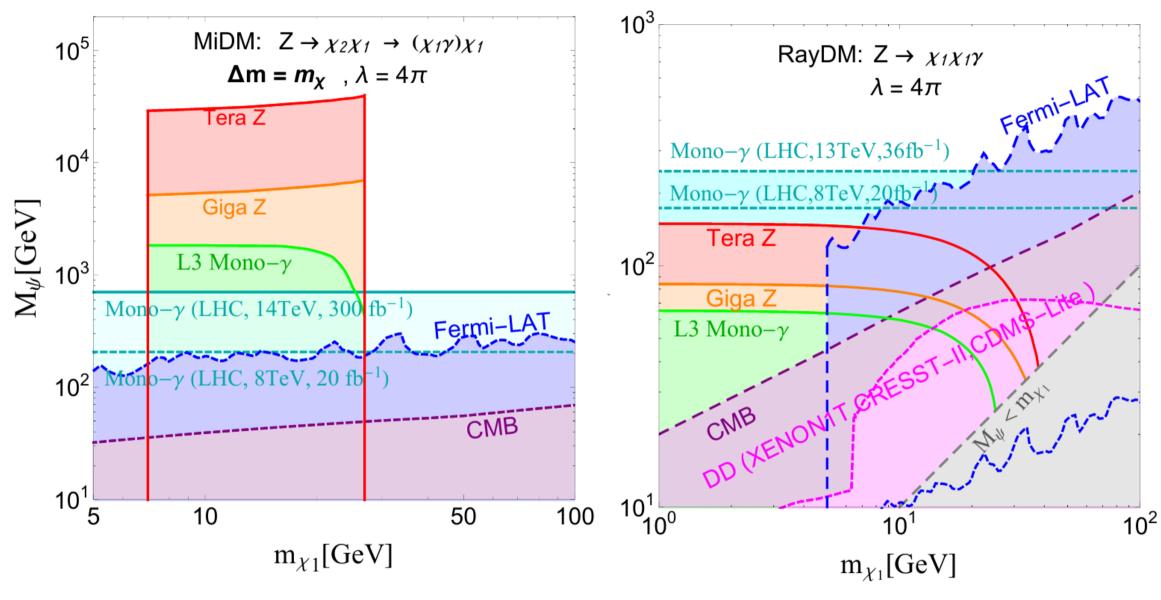
– focus only on region where splitting is large for Z decay.

– in this case the relevant annihilation initial state is only χ_1

– relevant annhilation cross section for $\ m_{\chi_1} < m_Z$

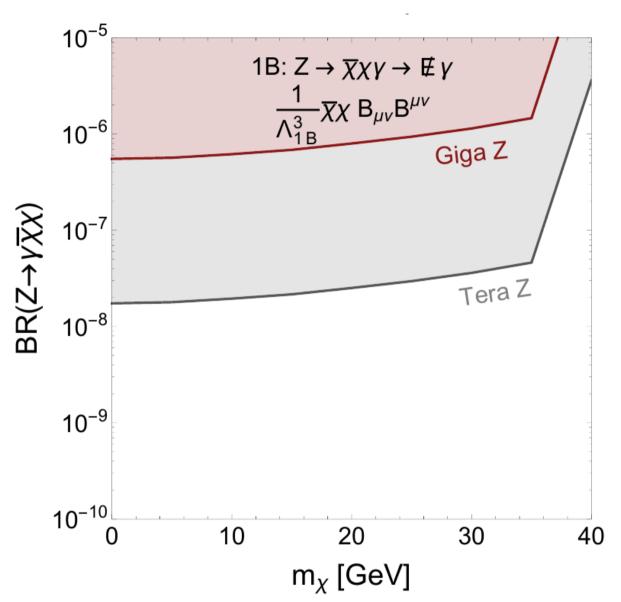
$$\begin{split} \sigma v(\chi_1 \chi_1 \to \gamma \gamma)_{\text{MIDM}} &= \frac{\cos^2 \theta_w m_{\chi_1}^2}{\pi \Lambda_{\text{MIDM}}^4} \frac{16y^6 - 9y^4 - 2y^2 - 2}{y^4 (y^2 + 2)^2} ,\\ \sigma v(\chi_1 \chi_1 \to \gamma \gamma)_{\text{RayDM}} &= \frac{\cos^2 \theta_w}{\pi} \frac{m_{\chi_1}^4}{\Lambda_{\text{RayDM}}^6} v_{\text{rel}}^2,\\ \sigma v(\chi_1 \chi_1 \to \gamma \gamma)_{\text{RayDM}}^{\gamma_5} &= \frac{16\cos^2 \theta_w}{\pi} \frac{m_{\chi_1}^4}{\Lambda_{\text{RayDM}}^6},\\ \frac{1}{\Lambda_{\text{MIDM}}} &\approx \frac{\lambda^2 g_Y}{64\pi^2 M_\psi} \qquad \frac{1}{\Lambda_{\text{RayDM}}^3} \approx \frac{\lambda^2 g_Y^2}{48\pi^2 M_\psi^3} \qquad y \equiv m_{\chi_2}/m_{\chi_1} \end{split}$$

 $Z \to \not\!\!\!E + \gamma$



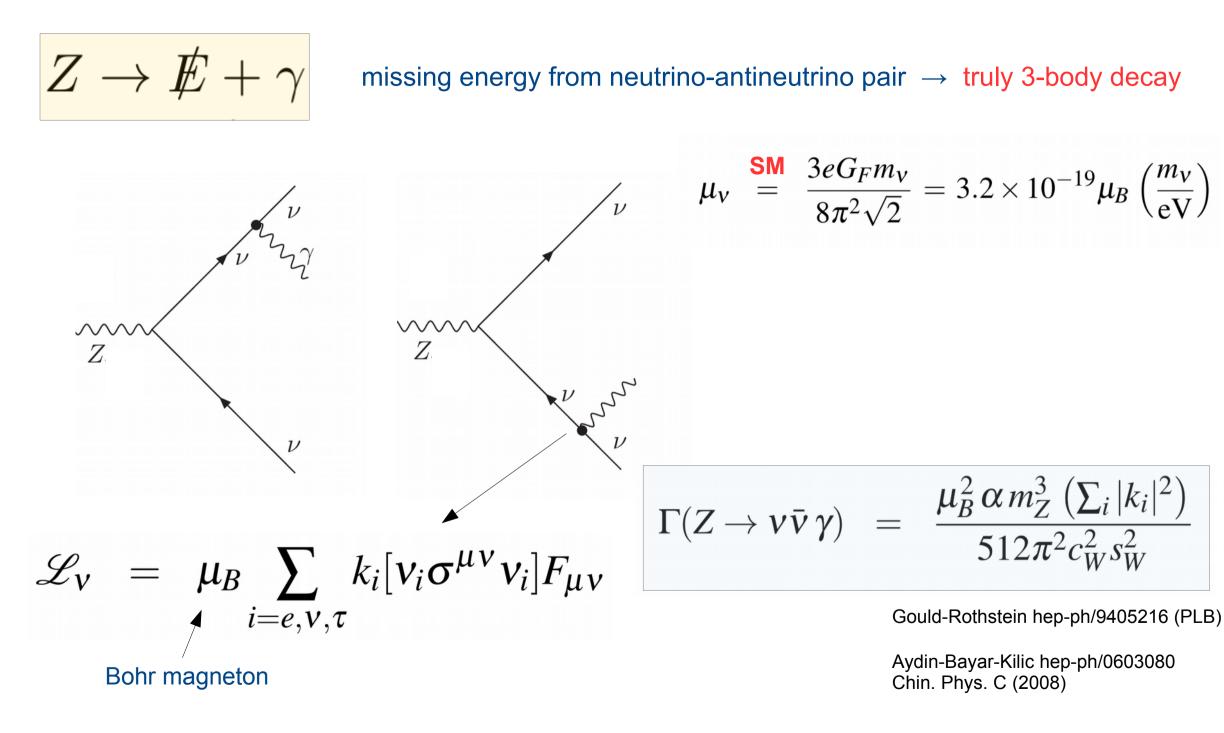
Liu, Wang, Wang, Xue, 1712.07237

This signature refers to a 2-body decay with "on-shell" decay of $\chi_2 \rightarrow \chi_1 + \gamma$

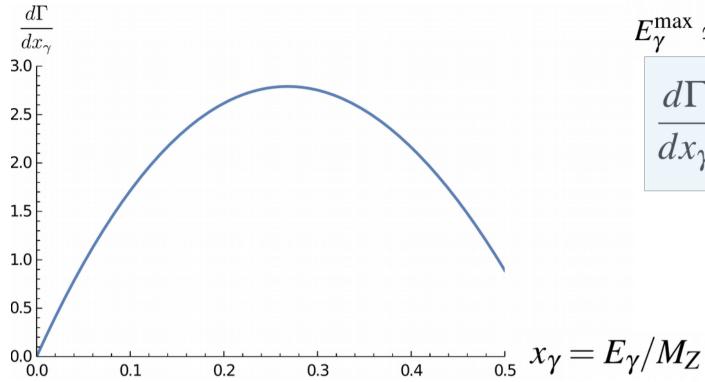


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Anomalous Magnetic moment of neutrinos



photon energy spectrum at the Z-pole



$$E_{\gamma}^{\max} \simeq 24 \,\text{GeV}$$
$$\frac{d\Gamma}{dx_{\gamma}} = \frac{\mu_B^2 \alpha m_Z^3 \left(\sum_i |k_i|^2\right)}{24\pi^2 c_W^2 s_W^2} F(x_{\gamma})$$

$$F(x) = x(1 - 2x + \frac{x^2}{3})$$

$$k^2 \equiv \sum_i |k_i|^2$$

From PDG

$$k_e < 2.9 \times 10^{-11}$$

 $k_\mu < 6.8 \times 10^{-10}$
 $k_\tau < 3.9 \times 10^{-7}$

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$$BR(Z \to v \bar{v} \gamma) = 2.3 \times 10^5 k^2$$

Improving bounds on MDM of neutrino-tau

naive estimations Giga Z \rightarrow $k < 6.6 \times 10^{-8}$ Tera Z \rightarrow $k < 2.1 \times 10^{-9}$ elevant SM bokg from tale of pon-resonant etc. $\rightarrow x x x$

relevant SM bckg from tale of non-resonant e+e- $\rightarrow \nu \; \nu \; \gamma$ E. Gabrielli

Light vectors coupled to anomalous currents

- Consider vector field X coupled to anomalous currents (like B-L)
- X has vectorial couplings to SM fermions
- UV completion require anomaly cancellation at high energy
- Wess-Zumino term can arise at low energy

$$\mathcal{L} \supset C_B g_X g'^2 \epsilon^{\mu\nu\rho\sigma} X_\mu B_\nu \partial_\rho B_\sigma + C_W g_X g^2 \epsilon^{\mu\nu\rho\sigma} X_\mu (W^a_\nu \partial_\rho W^a_\sigma + \frac{1}{3} g \epsilon^{abc} W^a_\nu W^b_\rho W^c_\sigma)$$

 $Z \to \not\!\!\!\! E + \gamma$

Dror-Lasenby-Pospelov 1705.06726 (PRL)

- No choice for C_{WZ} that preserves both U(1)_X and SM gauge group \rightarrow assumed to break U(1)_X • Massive X in the spectrum
- The longitudinal component (ϕ) of X acts as X=ALP coupled to Z and photon .

$$Z \to \gamma X$$

$$\Gamma_{Z \to \gamma X} \simeq \frac{\mathcal{A}^2}{1536\pi^5} g_X^2 g^2 g'^2 \frac{m_Z^3}{m_X^2}$$

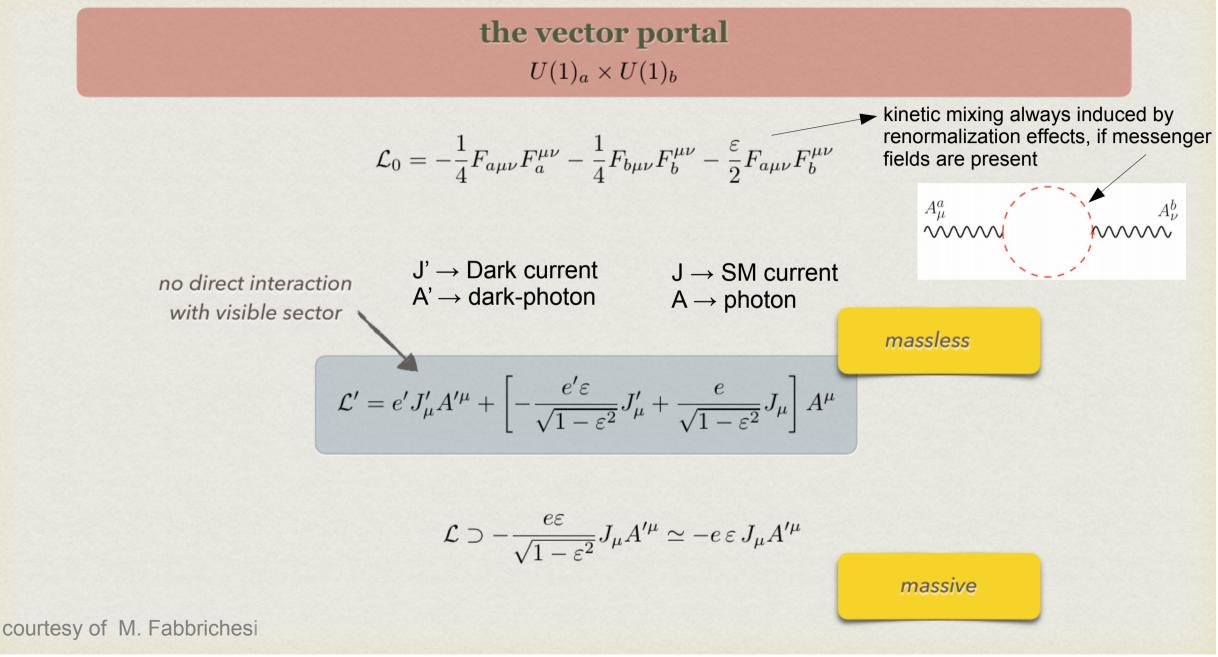
$$\frac{\mathcal{A}}{16\pi^2} \frac{g_X \varphi}{m_X} (g^2 W^a \tilde{W}^a - g'^2 B\tilde{B}) = Z, \gamma, \text{ field}$$

$$\frac{\mathcal{A}}{16\pi^2} \frac{g_X \varphi}{m_X} \left(g^2 (W^+ \tilde{W}^- + W^- \tilde{W}^+) \right) \quad \text{Strengths}$$

$$+ gg' (\cot \theta_w - \tan \theta_w) Z\tilde{Z} + 2gg' Z\tilde{F})$$

$$- ieg^2 \tilde{F}^{\mu\nu} (W^+_\mu W^-_\nu - W^+_\nu W^-_\mu) + \dots)$$

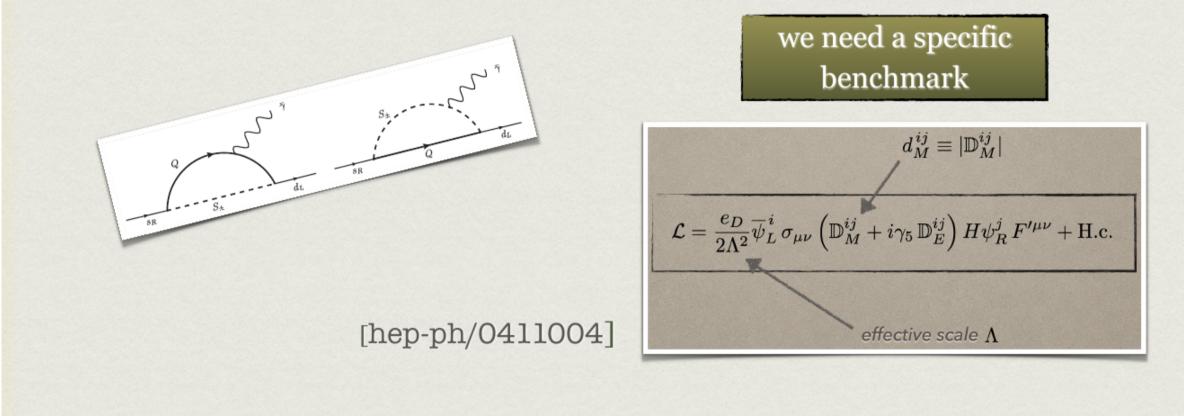
Dark Photon scenario



- B.Holdom, PLB 166B, 196 (1986)
 - B.A. Dobreascu, PRL 94 151802 (2005); [hep-ph/0411004]
 - EG, M. Fabbrichesi, G. Lanfranchi, "The dark photon" SpringerBriefs in Physics 2020; arXiv:2005.01515]

massless dark photon

the **massless** dark photon is not the massless limit of the **massive** dark photon



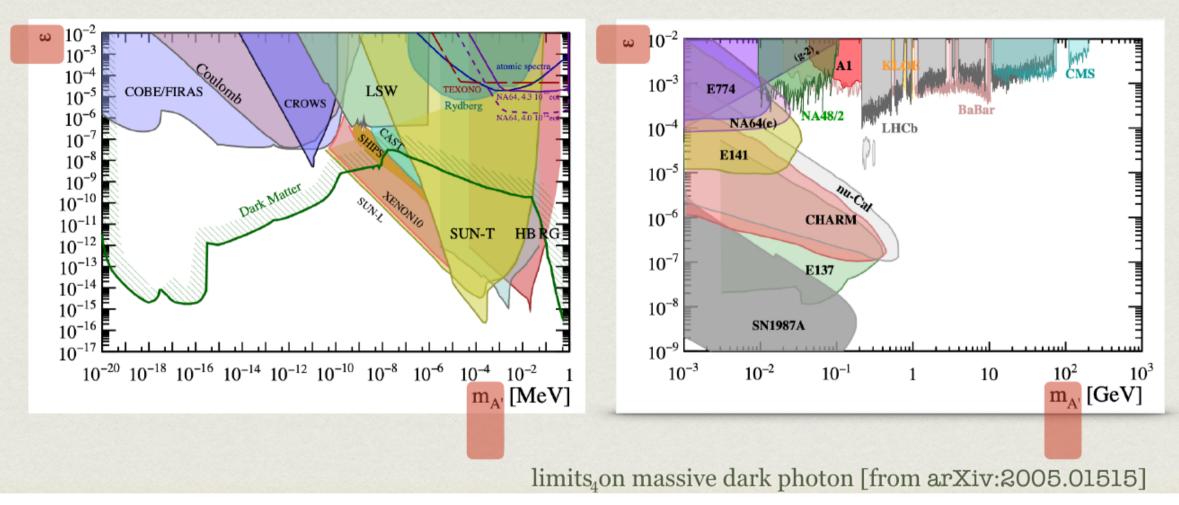
Massless dark photon interacts with SM particles via high dimensional operators → dim-5 (magnetic-dipole moments) leading contributions

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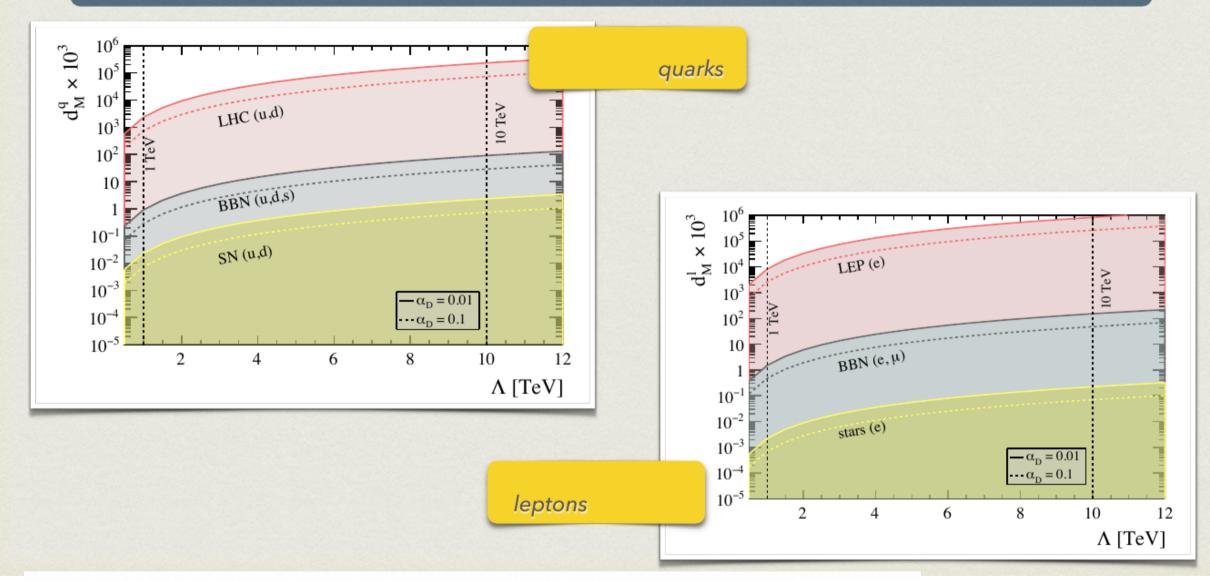
massive dark photon

invisible

visible



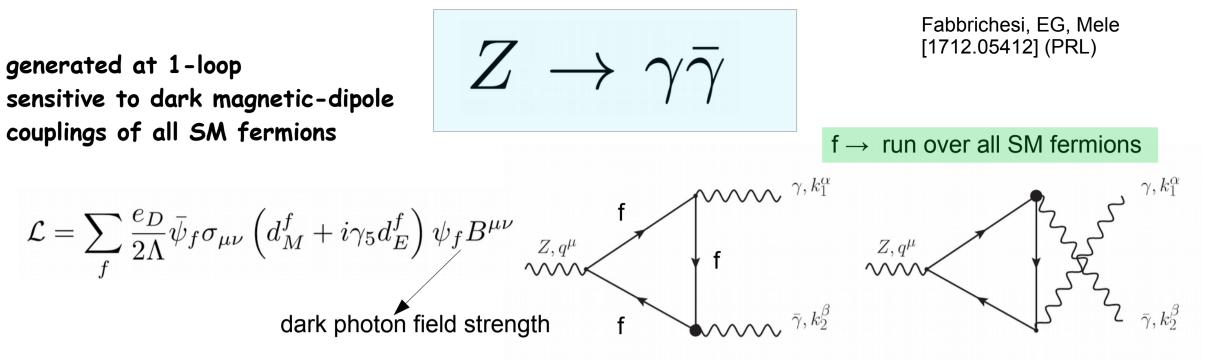
massless dark photon



BBN • Big bang nucleosynthesis. A cosmological bound for the dark photon operator comes from the determination of the effective number of relativistic species in addition to those of the SM partaking in the thermal bath—the same way the number of neutrinos is constrained.

SN • Supernovae. An additional limit is found from the neutrino signal of supernova 1987A, for which the length of the burst constrains anomalous energy losses in the explosion.

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- Landau-Yang theorem forbids Z /> 2 photons -> amplitude vanishes
 avoided due to distinguishability of photon and dark-photon interaction (blob)
- massive dark-photon couples also via magnetic dipole interaction, its tree-level coupling via mixing with photon is vanishing due to LY theorem

$$\mathcal{L}_{eff} = \frac{e}{\Lambda M_Z} \sum_{i=1}^{3} C_i \mathcal{O}_i(x)$$

C_i finite due
 gauge invariance

dimension-six operators \mathcal{O}_i are

 $\mathcal{O}_1(x) = Z_{\mu\nu}\tilde{B}^{\mu\alpha}A^{\nu}_{\ \alpha},$ $\mathcal{O}_2(x) = Z_{\mu\nu} B^{\mu\alpha} \tilde{A}^{\nu}{}_{\alpha} \,,$ $\mathcal{O}_3(x) = \tilde{Z}_{\mu\nu} B^{\mu\alpha} A^{\nu}_{\ \alpha} \,.$

$$\mathcal{L} = \sum_{f} \frac{e_D}{2\Lambda} \bar{\psi}_f \sigma_{\mu\nu} \left(d_M^f + i\gamma_5 d_E^f \right) \psi_f B^{\mu\nu}$$

$$BR(Z \to \gamma \bar{\gamma}) \simeq \frac{2.52 \ \alpha_D}{(\Lambda/\text{TeV})^2} \ (|d_M|^2 + |d_E|^2) \times 10^{-8}$$

$$\frac{10^{-9}}{\Lambda} \xrightarrow{\alpha_D \to 0.1} \Lambda \xrightarrow{\alpha_D \to 0.1} \Lambda \xrightarrow{\alpha_D \to 1} \text{TeV}$$

$$\frac{10^{-9}}{\Lambda} \xrightarrow{\alpha_D \to 0.1} \Lambda \xrightarrow{\alpha_D \to 0.1$$

10¹³ of Z boson events at the FCC-ee expected 10^2 -10⁴ of $Z \rightarrow \gamma \bar{\gamma}$ events g

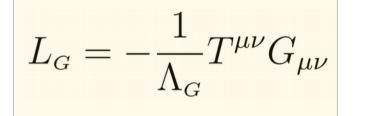
Spin-2 scenario

$$Z \to \not\!\!\!\! E + \gamma$$

 $G_{\mu\nu}$ massive spin 2 field \rightarrow Fierz-Pauli Lagrangian with mass m_G

Assume an effective coupling

 $T_{\mu\nu} \rightarrow \text{SM energy-momentum tensor}$ (gravity $m_G = 0$ $\Lambda_C^{-1} = \sqrt{8\pi G_N}$)



$$\Lambda_G$$
 \rightarrow universal coupling

Lower masses below eV severely constrained by test on deviation from gravity law

we restrict to the scenario $ightarrow \, {\rm eV} \lesssim m_G \lesssim 1 \, {
m GeV}$

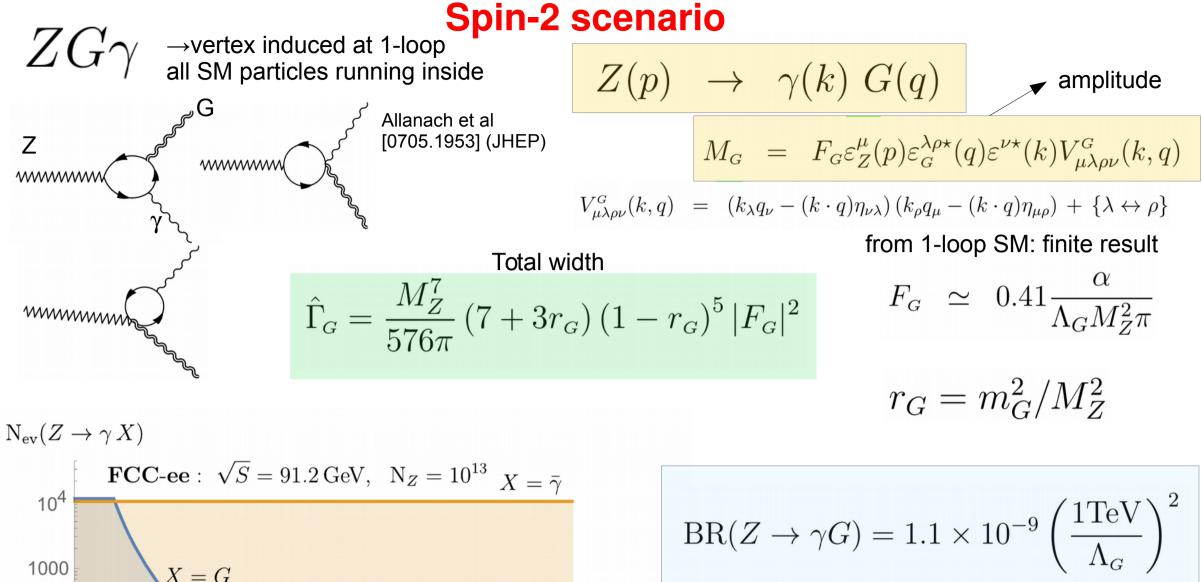
Requiring spin-2 not to decay inside detector (L=10m)

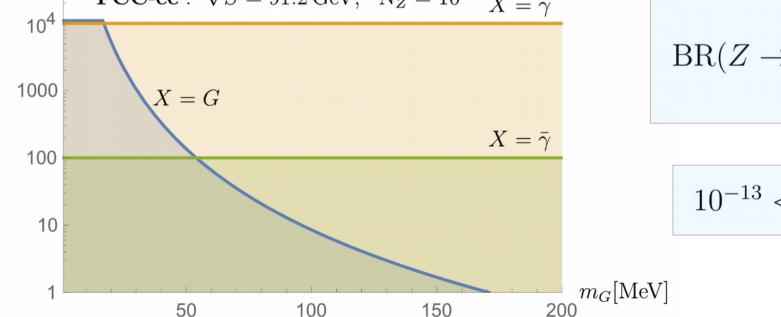
Comelato-EG [2006.00973] (PRD)

$$\Lambda_G \gtrsim 36 \left(\frac{m_G}{100 \text{MeV}}\right)^2 \text{TeV} , \qquad 1 \text{eV} \lesssim m_G \lesssim 2m_\mu$$

$$\Lambda_G \gtrsim 113 \left(\frac{m_G}{100 \text{MeV}}\right)^2 \text{TeV} , \qquad 2m_\mu \lesssim m_G \lesssim 1 \text{GeV}$$

 $\Lambda_G > \mathcal{O}(1\text{TeV})$ for all masses below 10MeV





$$BR(Z \to \gamma G) = 1.1 \times 10^{-9} \left(\frac{1 \text{TeV}}{\Lambda_G}\right)^2$$

$$10^{-13} < {\rm BR}(Z \to \gamma G) < 10^{-9}$$

[2006.00973]

 $Z \to \not\!\!\!E + \gamma$

signature with monochromatic photon (2 body decay)

Would it be possible to disentangle the spin S(X)=0,1,2 ?

[2006.00973]

 $\cos\theta$

<u>Polarized decay</u> $Z \rightarrow \gamma + X$

Polarizations: $R \rightarrow transverse$, $L \rightarrow longitudinal \beta_z \rightarrow 0$ Spin-1 (massless)

$$\frac{1}{\hat{\Gamma}} \frac{d\Gamma^{(T)}}{dz} = \frac{3}{4} \left(1 - z^2\right)$$
$$\frac{1}{\hat{\Gamma}} \frac{d\Gamma^{(L)}}{dz} = \frac{3}{2} z^2$$

Spin-O (massive)

Z =

$$\frac{1}{\hat{\Gamma}_{I}} \frac{d\Gamma_{I}^{(T)}}{dz} = \frac{3}{8} \left(1 + z^{2}\right)$$
$$\frac{1}{\hat{\Gamma}_{I}} \frac{d\Gamma_{I}^{(L)}}{dz} = \frac{3}{4} \left(1 - z^{2}\right)$$

(mass dependence cancels out)

$$\frac{1}{\hat{\Gamma}_{G}} \frac{d\Gamma_{G}^{(T)}}{dz} = \frac{3}{8} \frac{(1+z^{2}(1-2\delta_{G})+2\delta_{G})}{1+\delta_{G}}$$
$$\frac{1}{\hat{\Gamma}_{G}} \frac{d\Gamma_{G}^{(L)}}{dz} = \frac{3}{4} \frac{(1-z^{2}(1-2\delta_{G}))}{1+\delta_{G}},$$

Spin-2 (massive)

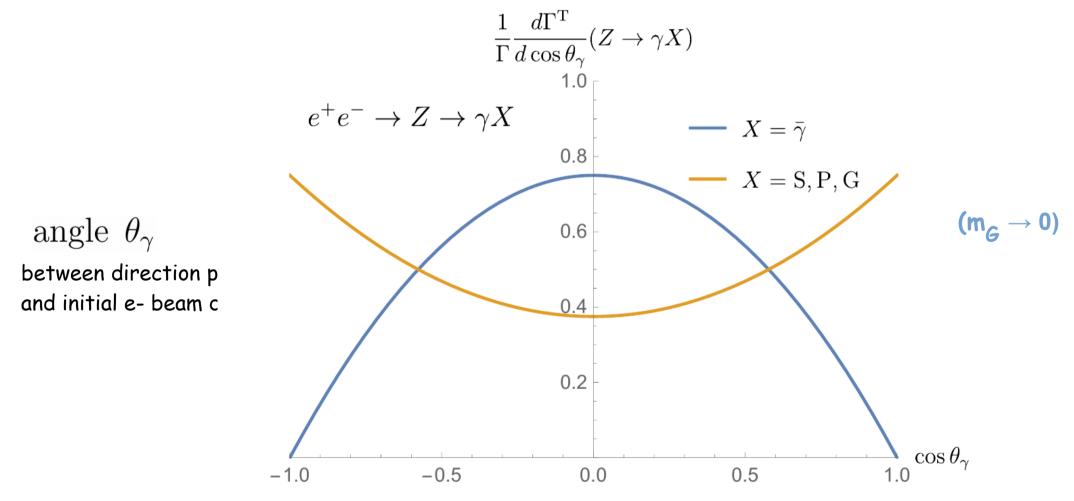
 $\delta_G = \frac{3}{7}r_G$

 $r_G = m_G^2 / M_Z^2$

Z decays at e^+e^- colliders

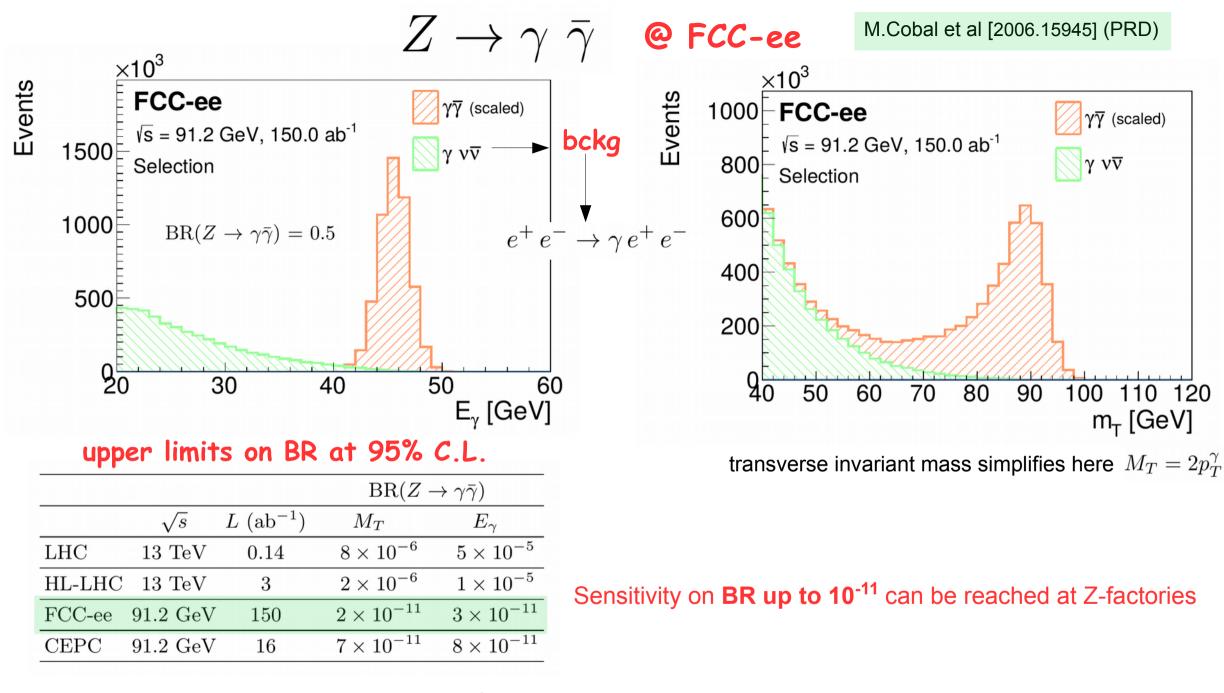
Z boson comes out mainly transverse (T) polarized \rightarrow due to angular momentum conservation

Possibility to disentangle spin-nature of X boson



X spin can be disentangled — spin-1: photon mainly produced <u>central and at large angles</u>

→ **spin-0/2:** photon mainly produced along FB directions



Spin analysis using test statistics

N=6 (N=17) \rightarrow lower bound for expected (observed) N. of signal events needed to exclude the

hypothesis under the $p_0(J^P = 1^-)$ assumption at 95% C.L.

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other interesting signatures

 $Z \to E + \gamma \gamma$

 $Z \to E + \ell^+ \ell^-$

N. resonances

		V	
	$Z \to \phi_d A', \phi_d \to (\gamma \gamma), A' \to (\bar{\chi} \chi)$	2	Vector portal
$Z \to \not\!\!\!\!\!\!\!\!\!\!\!\!\!\!\!\!\!\!\!\!\!\!\!\!\!\!\!\!\!\!\!\!\!\!\!$	$\begin{bmatrix} Z \to \phi_H \phi_A, \ \phi_H \to (\gamma \gamma), \ \phi_A \to (\bar{\chi} \chi) \end{bmatrix}$	2	2HDM extension
	$Z \to \chi_2 \chi_1, \ \chi_2 \to \chi_1 \phi, \ \phi \to (\gamma \gamma)$	1	Inelastic DM
	$Z \to \chi_2 \chi_2, \chi_2 \to \gamma \chi_1$	0	MIDM

Liu, Wang, Wang, Xue, 1712.07237

 χ is assumed to be Majorana (DM like) \rightarrow stable or decaying outside detector Vector-portal A' (DP like)

$$\mathcal{L}_F = \bar{\chi} i \partial \!\!\!/ \chi + g_D \bar{\chi} A'_\mu \gamma^\mu \chi - m_D \bar{\chi} \chi + (\Phi^* (y_L \bar{\chi}^c P_L \chi + y_R \bar{\chi}^c P_R \chi) + h.c.)$$

once ϕ gets a vev χ gets a Majorana mass term , + Dirac mass \rightarrow 2 Majorana $\chi_1 \chi_2$

n=1,2 resonances, depending on the masses of the res. BR sensitivity of Giga Z is around $\begin{bmatrix} 10^{-8.4}, 10^{-6.7} \end{bmatrix}$ for Tera Z $\begin{bmatrix} 10^{-11}, 10^{-9.7} \end{bmatrix}$ contact term \rightarrow zero resonance

 $[10^{-10.3}, 10^{-9.2}]$ (Tera-Z)

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 $[10^{-8.4}, 10^{-7.4}]$ (Giga-Z)

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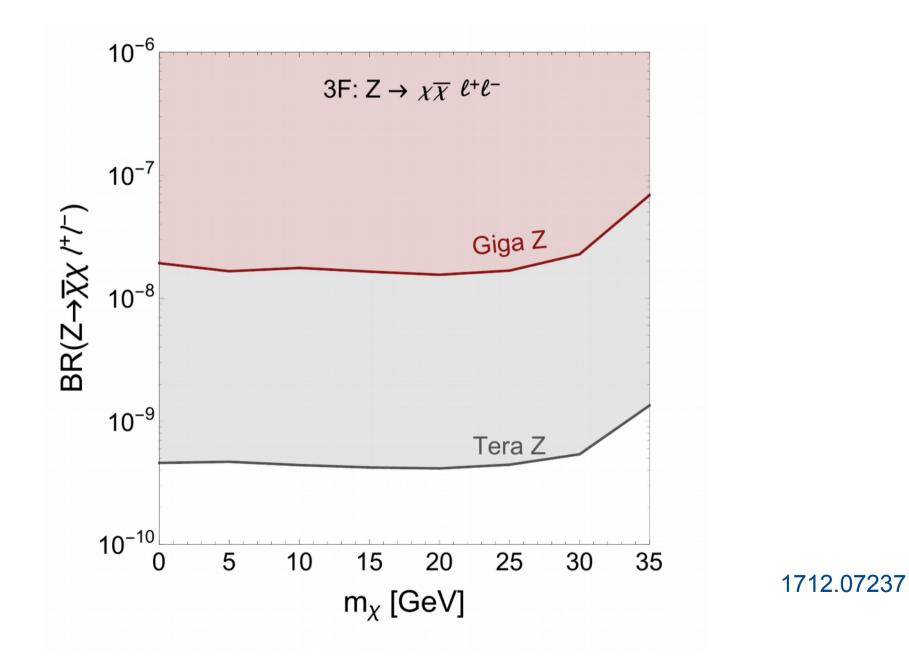
		N. reso	nances		
$Z \to \not\!\!\!\!\!\!\!\!\!\!\!\!\!\!\!\!\!\!\!\!\!\!\!\!\!\!\!\!\!\!\!\!\!\!\!$	$ \begin{array}{cccc} Z \to \phi_d A', \ A' \to (\ell^+ \ell^-), \ \phi_d \to \\ (\bar{\chi}\chi) \end{array} $	2	Vector portal		
	$Z \to A'SS \to (\ell\ell)SS$	1	Vector portal		
	$Z \to \phi(Z^*/\gamma^*) \to \phi \ell^+ \ell^-$	1	Long-lived ALP, Higgs portal		
	$Z \to \chi_2 \chi_1 \to \chi_1 A' \chi_1 \to (\ell^+ \ell^-) \not \!\!\! E$	1	Vector portal and Inelastic DM		
	$Z \to \chi_2 \chi_1, \chi_2 \to \chi_1 \ell^+ \ell^-$	0	MIDM, SUSY		
	$Z \to \bar{\chi} \chi \ell^+ \ell^-$	0	RayDM, slepton, heavy lepton mixing		

for Giga-Z $BR(Z \rightarrow \not E \ell^+ \ell^-)$ probed down to $\sim 10^{-8.5}$ for n. res > 0 $\sim 10^{-8}$ for n. res = 0 for Tera-Z \rightarrow sensitivity better by a factor $10^{1.5}$

1712.07237

 $Z \to \not\!\!\!E + \ell^+ \ell^-$

Contact term \rightarrow no resonances

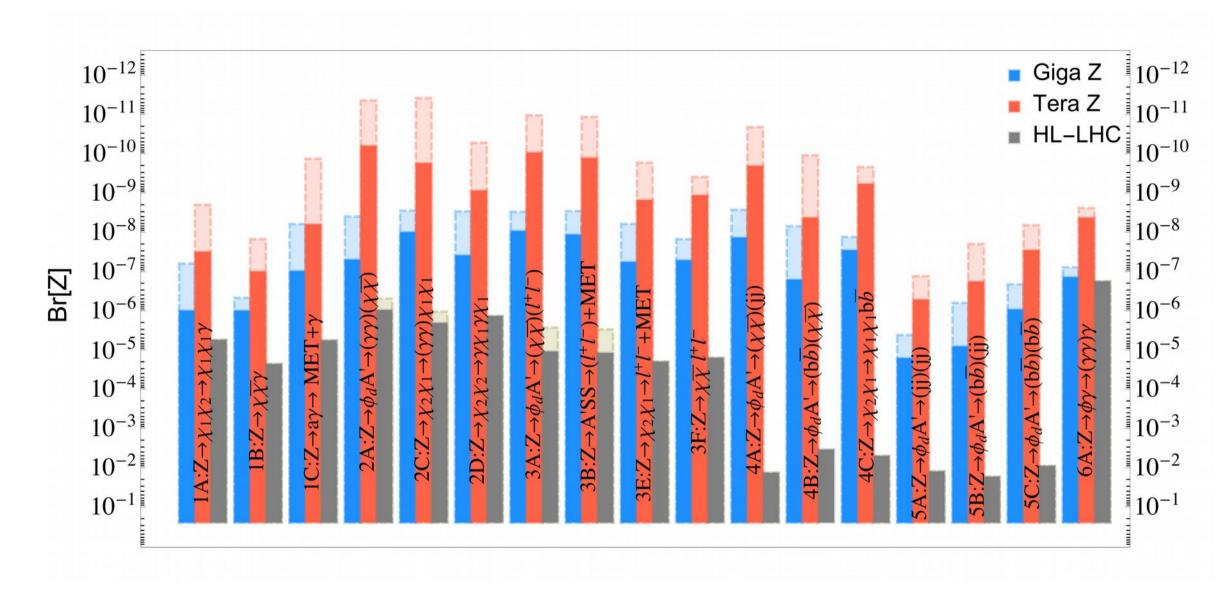


		N. res ↓	sonances
	$Z \to \phi_d A', \phi_d \to jj, A' \to jj$	2	Vector portal + Higgs portal
$Z \to (JJ)(JJ)$	$Z \to \phi_d A', \phi_d \to b\bar{b}, A' \to jj$	2	vector portal + Higgs portal
	$Z \to \phi_d A', \phi_d \to b\bar{b}, A' \to b\bar{b}$	2	vector portal + Higgs portal

for Giga-Z BR(Z
$$\rightarrow$$
 (JJ)(JJ))probed down to
 $\sim 10^{-5}$ for $(jj)(jj)$
 $\sim 10^{-6}$ for $(jj)(bb)$
 $\sim 10^{-6.5}$ for $(bb)(bb)$

for Tera-Z $\rightarrow\,$ sensitivity better by a factor $10^{1.5}$

1712.07237



Exotic scenarios: Z → ff + G SM EM tensor Massive spin-2 (Light) massive spin-2 G resonance, effectively (universally) coupled to SM \Longrightarrow L_G = f=SM fermions NNN G_{uv} EG-Mele, hep-ph/0205099, NPB 647 (2002)

Signature expected

done for QG scenarios (ADD) \rightarrow continuous spectrum

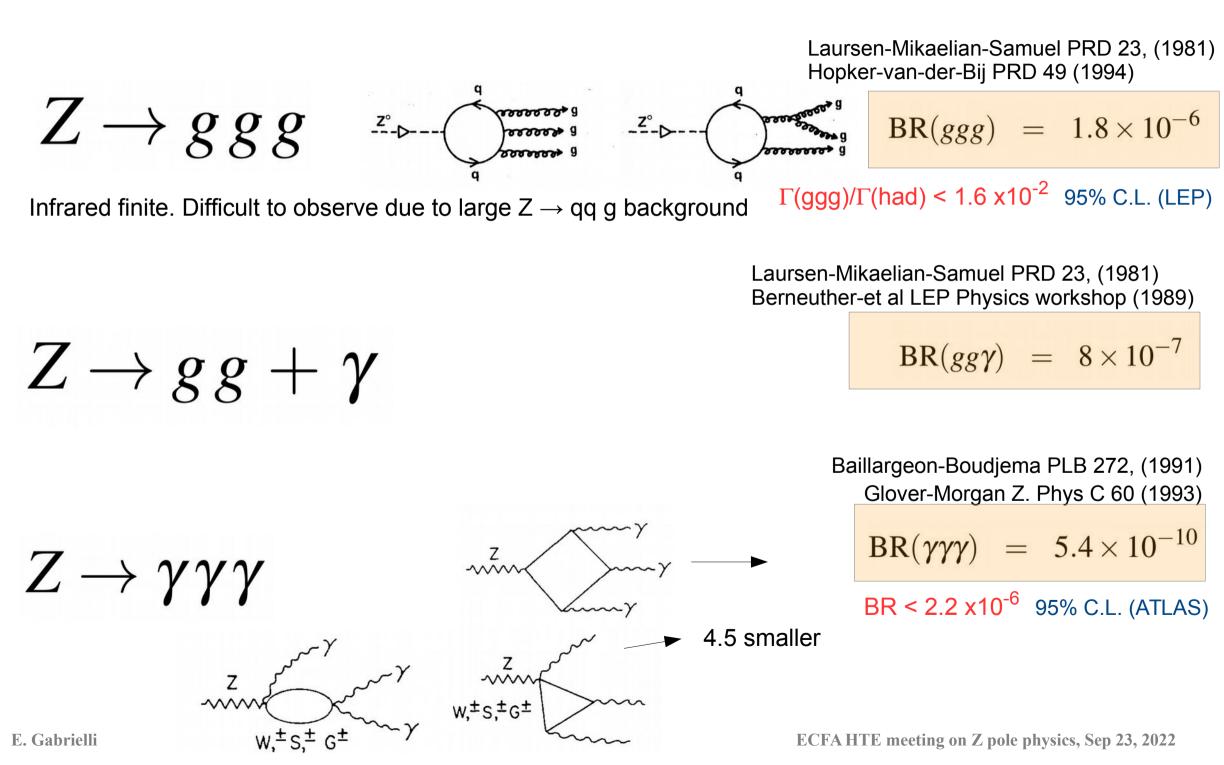
 $Z \to \not\!\!\!E + \ell^+ \ell^-$

missing analysis for a light (invisible) spin-2 scenario, with low-energy couplings

New lines of investigation

- study of sensitivity to the Λ scale including SM background
- analysis to disentangle other spin-X = ALP, massless-DP scenarios (behaving as missing energy)

Loop induced decays



Loop induced decays: sensitivity to NP

Few studies in the literature, mainly in contexts of specific models.

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Expected effective Lagrangians, dim. 8 operators with field-strengths (F,G,Z) couplings

$$\mathscr{L}_{Z}^{eff} = \frac{1}{\Lambda_{1}^{2}} (FFFZ) + \frac{1}{\Lambda_{2}^{2}} (G^{a}G^{a}FZ) + \frac{1}{\Lambda_{3}^{2}} (G^{a}G^{b}G^{c}Z) f^{abc}$$

$$\mathbf{F} \rightarrow \mathbf{photon} \qquad \mathbf{Z} \rightarrow \mathbf{Z}^{0} \qquad \mathbf{G} \rightarrow \mathbf{gluons} \qquad \text{(contractions of Lorentz indices)}$$

studied in the context of 331 minimal model $SU(3)_c \times SU(3)_L \times U(1)_Y$

in all possible ways)

Backup slides

$$\mathcal{O}_{1}(x) = Z_{\mu\nu}B^{\mu\alpha}A^{\nu}{}_{\alpha}$$
$$\mathcal{O}_{2}(x) = Z_{\mu\nu}B^{\mu\alpha}\tilde{A}^{\nu}{}_{\alpha}$$
$$\mathcal{O}_{3}(x) = \tilde{Z}_{\mu\nu}B^{\mu\alpha}A^{\nu}{}_{\alpha}$$

	b	t	s	c	τ	μ	
X_f	4.80	0.82	0.014	4.78	1.30	0.017	$\times 10^{-9}$

$$X_f \equiv \frac{m_f}{M_Z} N_c^f g_A^f Q_f e_D$$
$$N_c = 1(3) \text{ for leptons (quarks)}$$

 $C_2 = -3\sum_f \frac{d_M^f X_f}{4\pi^2} \left(2 + B_f\right),$

 $C_3 = 2\sum_{f} \frac{d_M^f X_f}{4\pi^2} \left(4 + 2B_f + C_f M_Z^2\right).$

 $C_1 = -\sum_{f} \frac{d_M^f X_f}{4\pi^2} \left(5 + 2B_f + 2C_f \left(m_f^2 + M_Z^2 \right) \right)$

 $B_f \equiv \text{Disc}[B_0(M_Z^2, m_f, m_f)],$ $C_f \equiv C_0(0, 0, M_Z^2, m_f, m_f, m_f)$

 B_0 and C_0 are the scalar two- and three-point Passarino-Veltman functions