

Search for the critical point of strongly interacting matter by NA61/SHINE at the $\check{CERN}SPS$

Haradhan Adhikary¹, Maja Pawlowska² for the NA61/SHINE Collaboration ¹Jan Kochanowski University (PL), ²Warsaw University of Technology (PL)

(1)

(2)

(3)

Motivation

The goal of this work is to search the **critical point** (a hypothetical endpoint of first-order phase transition line (QGP-HM) that has properties of secondorder phase transition) of the strongly interacting matter by measuring second scaled factorial moments of proton multiplicity distribution from a selection of Pb+Pb collisions at beam momentum of 13A GeV/c ($\sqrt{s_{NN}} \approx 5.1$ GeV) and 30A GeV/c ($\sqrt{s_{NN}} \approx 7.5$ GeV), and Ar+Sc collisions at beam momenta of $13A-150A \text{ GeV}/c \ (\sqrt{s_{NN}} \approx 5.1-17 \text{ GeV}) \text{ using cumulative variables and statistically independent data points.}$

Second scaled factorial moments

In NA61/SHINE [1], intermittency analysis [2] is performed at mid-rapidity, and particle fluctuations are studied in the transverse-momentum plane to search the QCD critical point [3, 4] by measuring scaled factorial moments of multiplicity distribution:

New approach to intermittency analysis

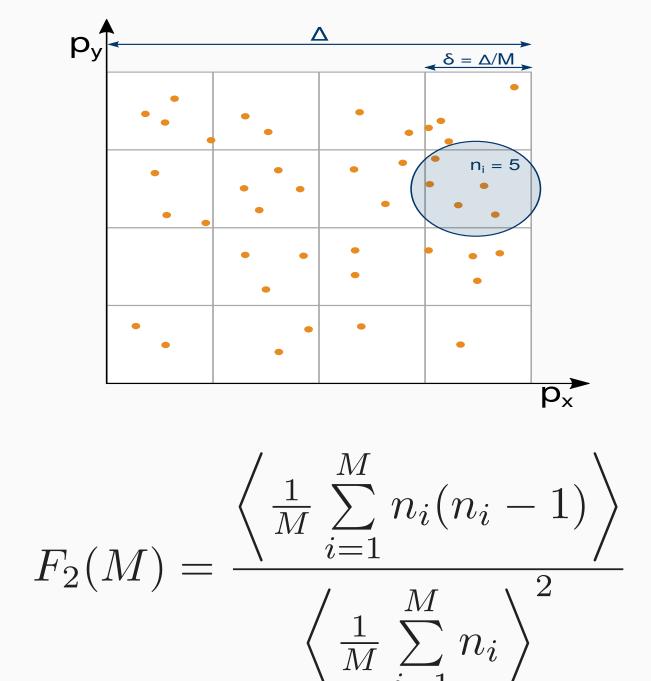
Cumulative variables: The SFMs are sensitive to the shape of the single-particle momentum distribution, which biases the signal of critical fluctuations. In our approach, to eliminate the bias, the cumulative transformation technique [9] is used, which has the following properties:

=|=|=

۲₁

XQCD

2023

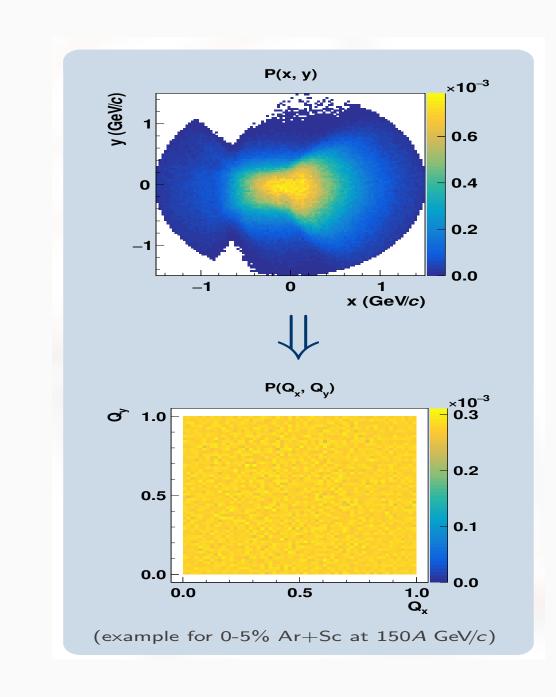


 n_i : numbers of particles in ith bin $< \dots >:$ averaging over events

M: number of subdivision intervals of the selected range Δ When the system is a simple fractal and $F_2(M)$ follows a power law dependence [5, 6, 7, 8]:

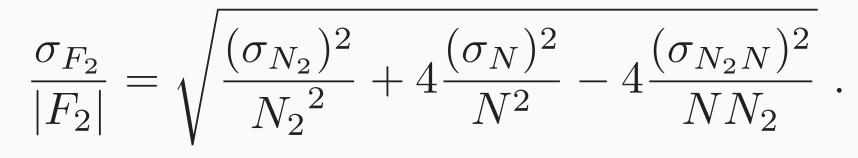
 $F_2(M) = F_2(\Delta) \cdot (M^2)^{\varphi_2}$

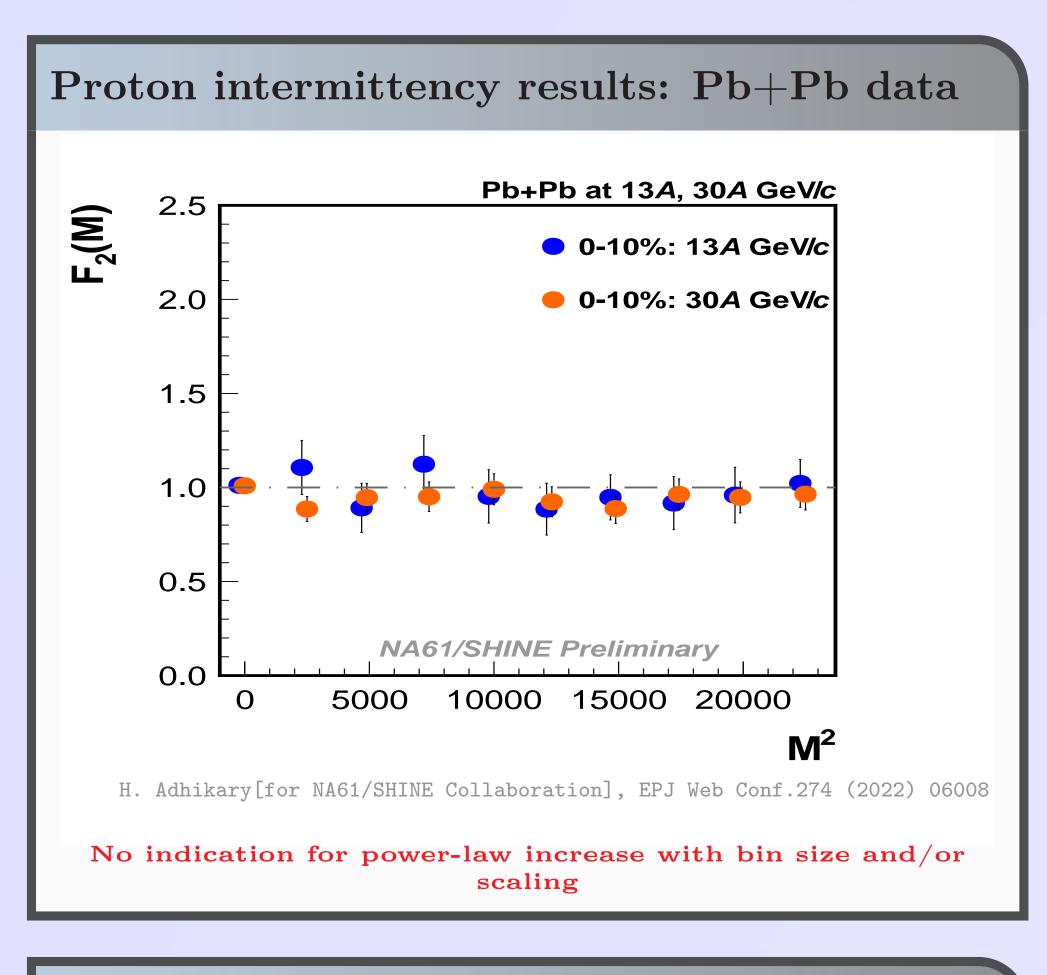
where critical exponent or intermittency index for proton intermittency, $\varphi_2 = \frac{5}{6} \approx 0.83$ [6] The statistical uncertainties can be calculated using the statistical uncertainty propagation:

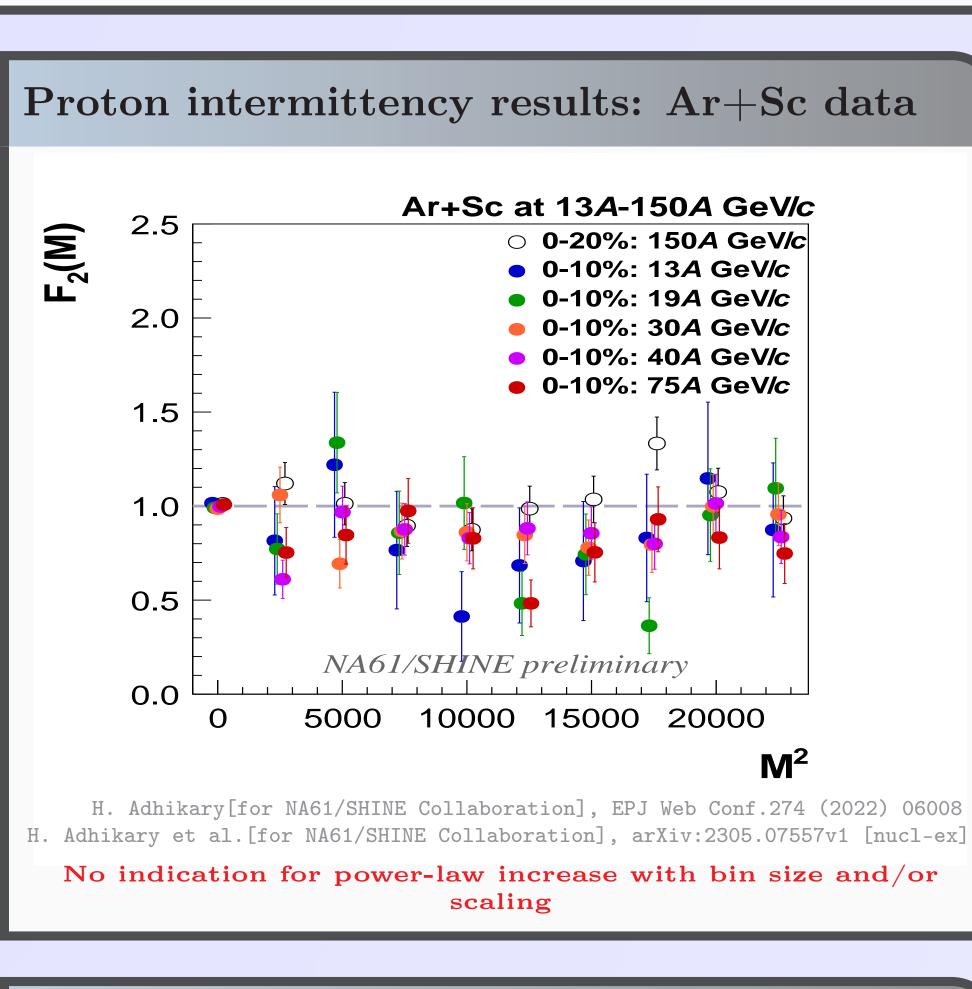


- transforms single-particle distribution into a uniform one ranging from 0 to 1
- remove the dependence of $F_2(M)$ on the shape of the single-particle distribution
- intermittency index of an ideal power-law correlation function was proven to remain invariant after the transformation

(4)

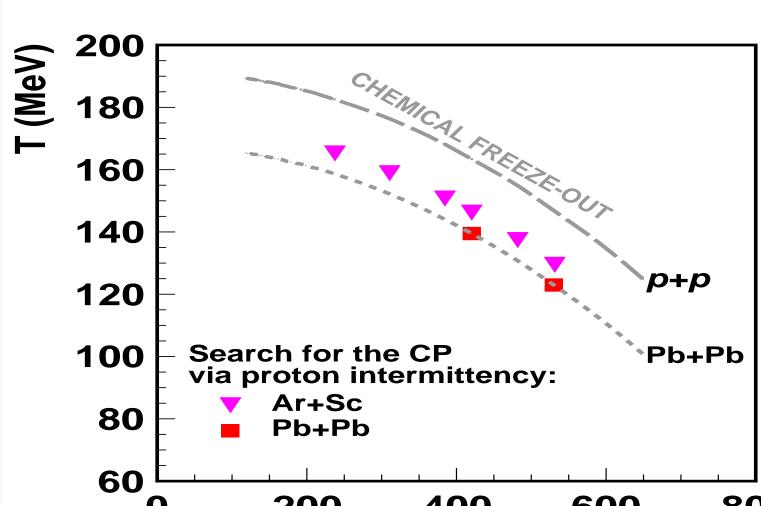






Statistically independent data points: Statistically-independent data subsets are used to obtain results for each subdivision number. As a result,

- for different subdivision numbers, results are statistically independent
- only diagonal elements of the covariance matrix are non-zero, and the complete relevant information needed to interpret the results is easy to present graphically



Proton intermittency summary

Power-Law model

The Power-Law model generates events that reproduce the experimental multiplicity and transverse momentum distributions of particles. Correlatedparticle pairs' transverse momentum difference follows a power-law distribution:

 $\rho(|\Delta \vec{p}_T|) = (|\Delta \vec{p}_T|)^{-\varphi_2}$

It has two main parameters:

- ratio of correlated to uncorrelated particles
- power-law exponent (φ_2)

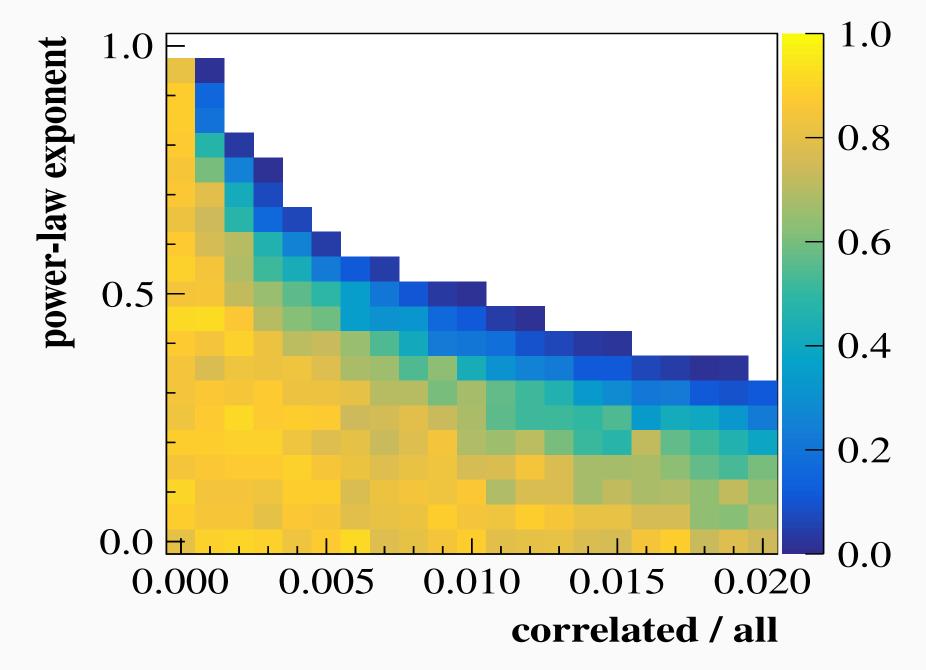
Lots of model data sets are generated:

- correlated-to-all ratio: vary from 0.0 to 2.0%,
- power-law-exponent: vary from 0.0 to 1.0,

and compared with the experimental data.

H. Adhikary et al. [for NA61/SHINE Collaboration], arXiv:2305.07557v1

Exclusion plot (for Ar+Sc at 150A GeV)



white area: p-value < 0.01exclusion plots for parameters of the Power-Law model

200 400 **600** 800 0 μ_{R} (MeV)

Becattini, Manninen and Gazdzicki, Phys.Rev.C73(2006) 044905 Summarize the ongoing NA61/SHINE critical point search program via proton intermittency of Pb+Pb and Ar+Sc data sets on the diagram of chemical freeze-out temperature and chemical potential

- - no indication of a power-law increase - -

References

1. NA61/SHINE web page:https://shine.web.cern.ch

- 2. Bialas, Peschanski, NPB 273(1986) 703
- 3. Stephanov, Rajagopal and Shuryak, PRL.81, 4816 (1998)
- M. A. Stephanov, Phys. Rev. Lett. 102, 032301 (2009)
- Barducci, Casalbuoni, Curtis, Gatto, Pettini, PLB 231 (1989)
- Wosiek, APPB 19 (1988) 863
- Antoniou, Diakonos, Kapoyannis, Kousouris, PRL97(2006)
- Bialas, Hwa, PLB 253 (1991) 436
- 9. Bialas, Gazdzicki, PLB 252 (1990) 483

Acknowledgements

This work is supported by the National Science Centre, Poland, under grant no. 2018/30/A/ST2/0026 & the NA61/SHINE SPS CERN Collaboration