

XII International Conference on New Frontiers in Physics 10-23 July 2023



Latest results and outlook of the KM3NeT neutrino telescope

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Neutrino astronomy

Unique neutrino properties:

- Electrically neutral, stable, and weakly interacting particles.
- **No deflection** in magnetic field (unlike cosmic rays).
- No absorption in background light/CMB (unlike gamma-rays).



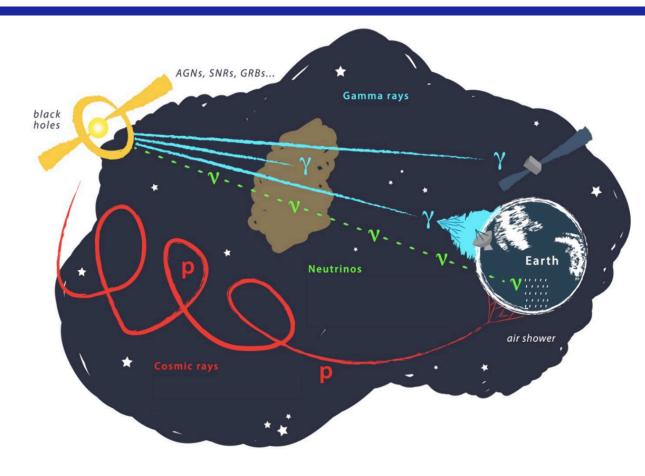
- Access to dense environments
- Access to cosmological distances
- Astronomy over full energy range

"Smoking gun" signature for hadronic processes:

$$p + p/\gamma \to X + \pi^0 \to \gamma \gamma$$

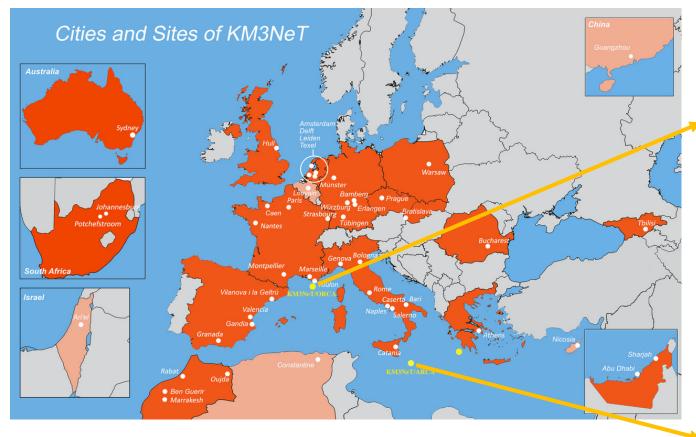
$$\to X + \pi^{\pm} \to \mu^{\pm} + \nu_{\mu}^{(-)}$$

$$\mu^{\pm} \to e^{\pm} + \nu_{e}^{(-)} + \nu_{\mu}^{(-)}$$



Correlation with electromagnetic and gravitational waves = Multimessenger approach

The KM3NeT Collaboration



~40 km

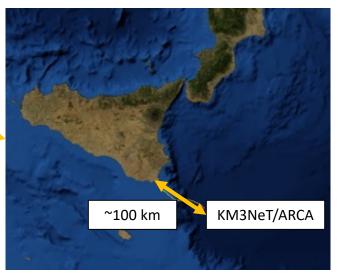
KM3NeT/ORCA

Oscillation
Research with
Cosmics in the
Abyss

GeV-TeV neutrinos

- ~250 members from more than 50 institutes
- Two deep-sea neutrino telescopes
- Single Collaboration and single technology

doi:10.1088/0954-3899/43/8/084001



Astroparticle
Research with
Cosmics in the
Abyss

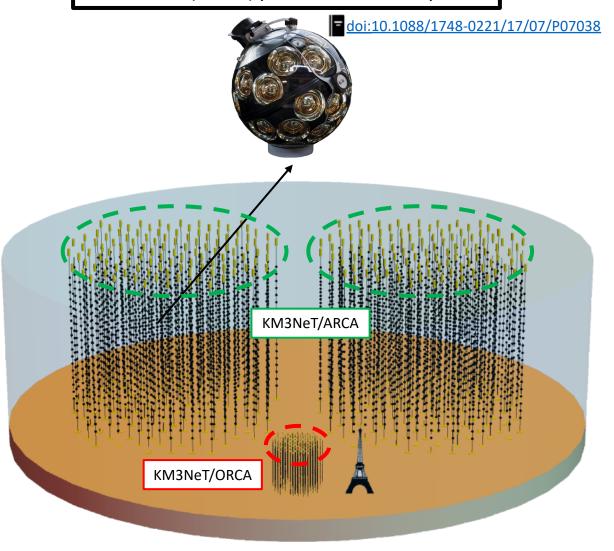
> TeV neutrinos



	KM3NeT/ARCA	KM3NeT/ORCA
#PMTs/DOM	31	31
#DOMs/DU	18	18
DOM vertical spacing	36 m	9 m
DU horizontal spacing	80 m	20 m
#DUs	115x2	115
Depth	3450 m	2450 m
Instrumented volume	~1 Gton	~7 Mton
Deployed DUs	21	18

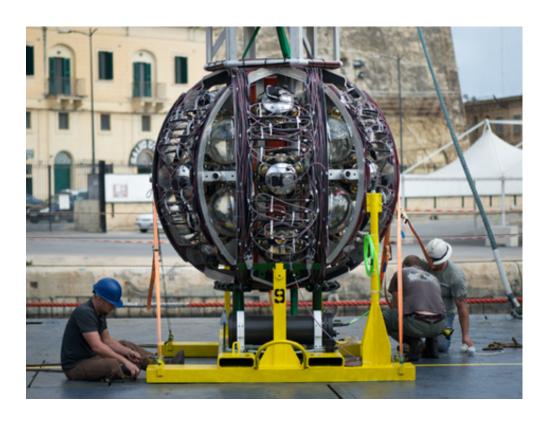
The data taking is already ongoing!

Digital Optical Module (DOM): multi-PMT (31x3") pressure-resistant sphere

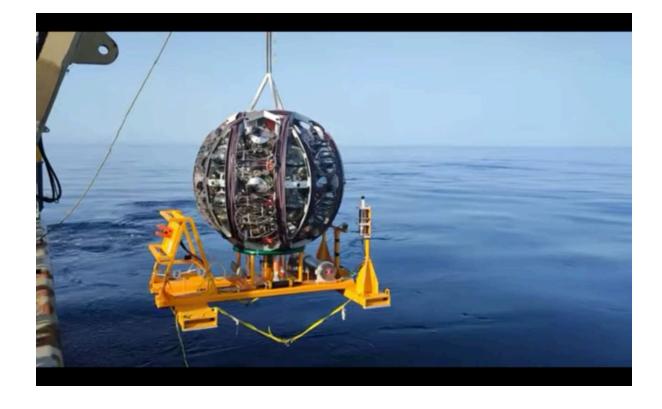




Reusable launcher modules **Multiple** DUs per sea campaign

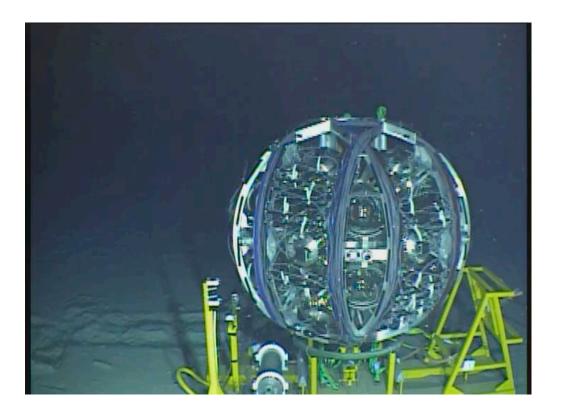


Deployment via ship crane
Sink to the sea floor using an anchor





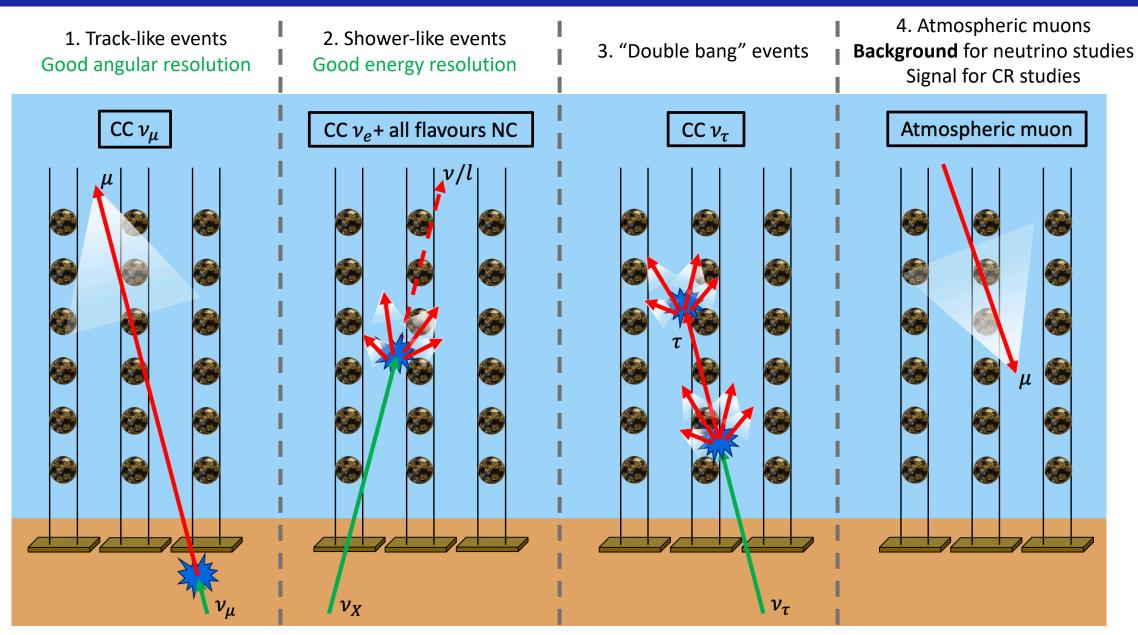
Autonomous/ROV unfurling



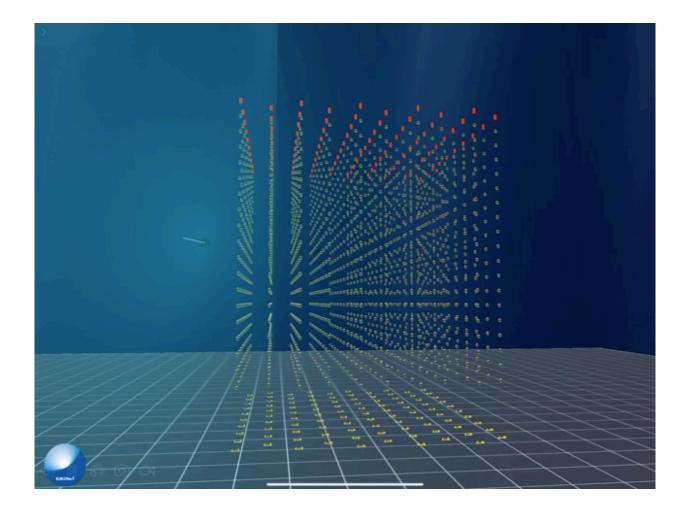
Visual inspection of DOMs with ROV





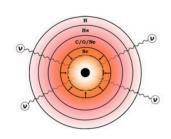


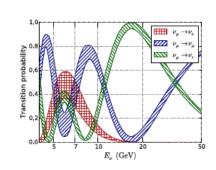


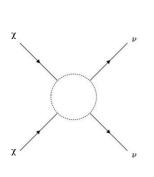


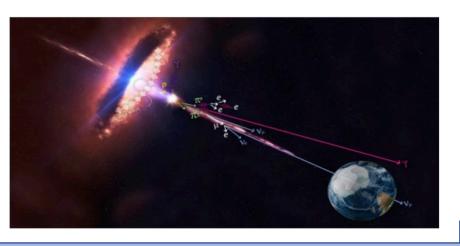
Supernovae Solar flares Atmospheric neutrinos Neutrino oscillations Mass hierarchy

Dark matter Monopoles, Nuclearites Cosmic neutrinos
Multi-messenger observations









MeV GeV — TeV — TeV — PeV

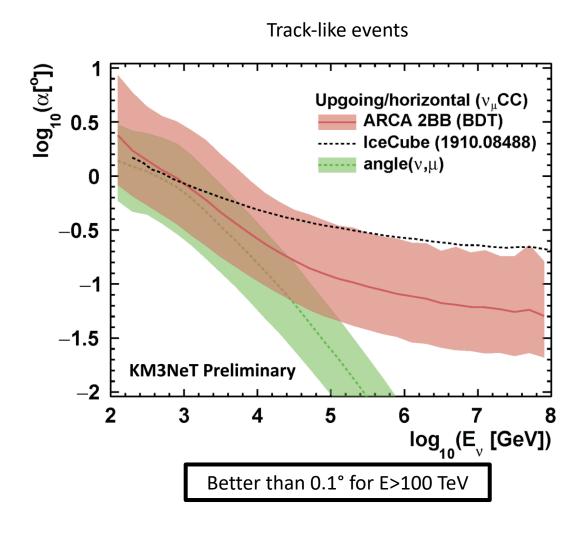


KM3NeT/ORCA



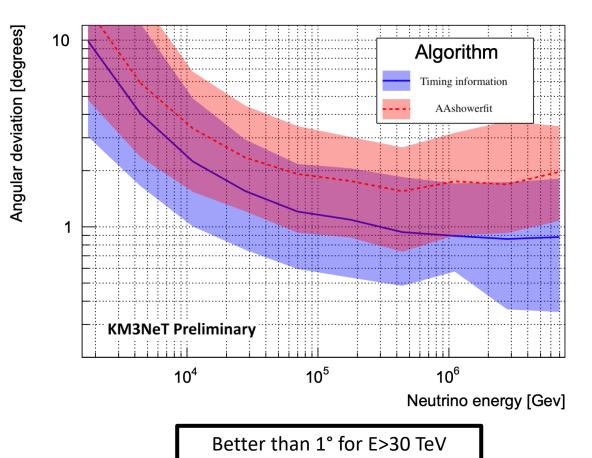
KM3NeT/ARCA





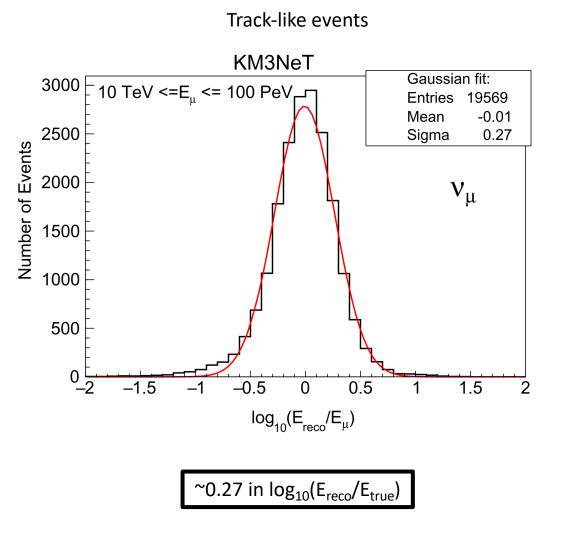
Timing information – includes additional info on hit times
Aashowerfit – standard shower reconstruction in KM3NeT/ARCA

Shower-like events

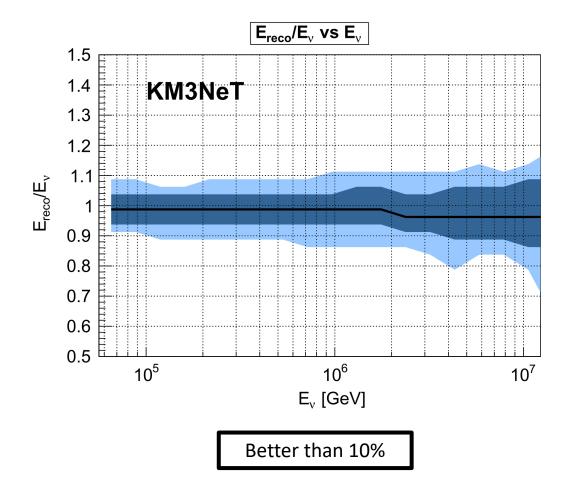


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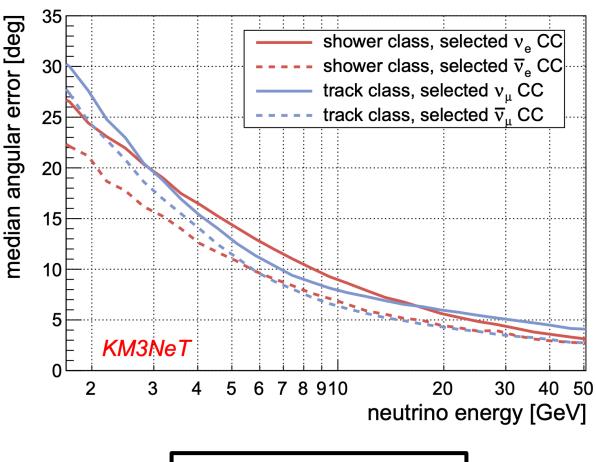






doi:10.1088/0954-3899/43/8/084001



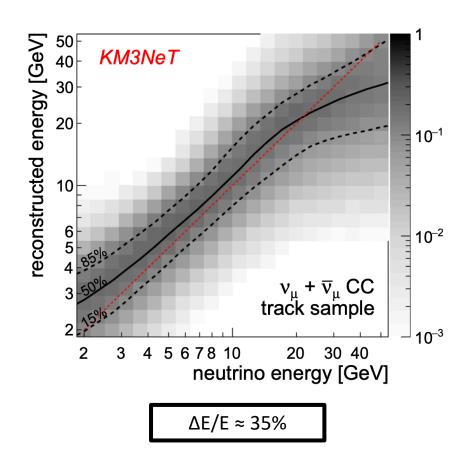


Better than 15° for E>5 GeV

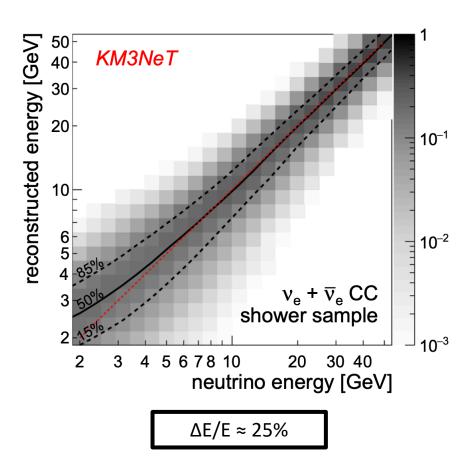
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Track-like events



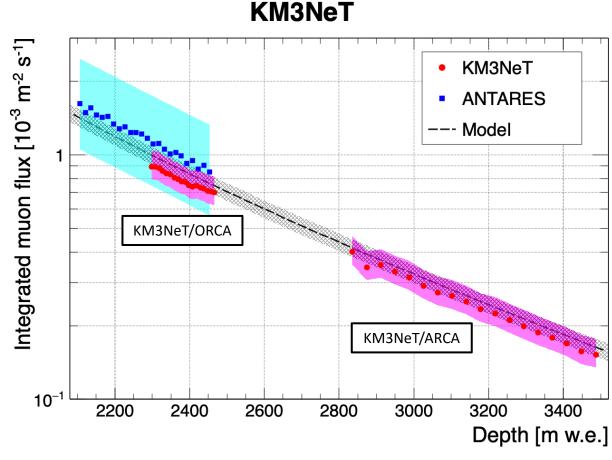
Shower-like events



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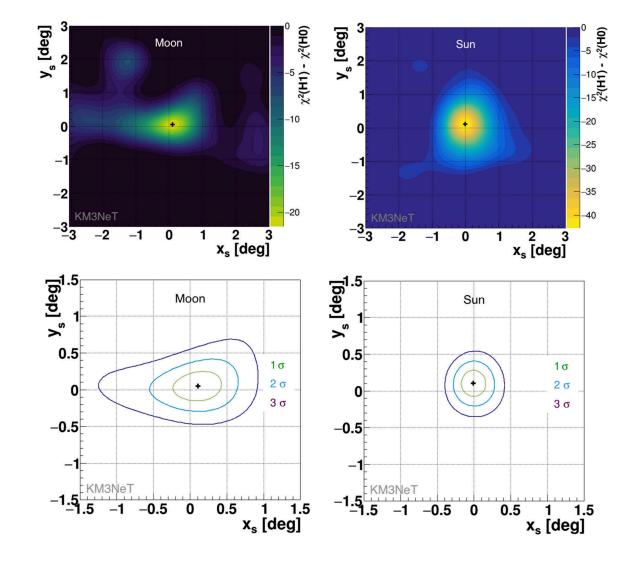
- Underwater muon flux as a function of depth.
- The KM3NeT/ARCA detector with 2 working DUs (23/12/2016-2/3/2017) and KM3NeT/ORCA with 1 DU (9/11/2017-13/12/2017)
- The flux is compared to the ANTARES result and to the Bugaev model



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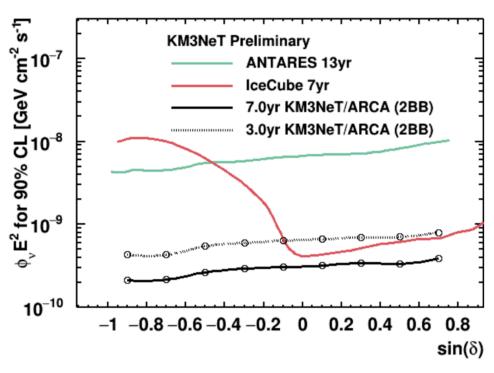
- The shadow is the deficit in atmospheric muon flux induced by the absorption of cosmic rays by the Moon/Sun.
- The KM3NeT/ORCA data taken between February 2020 and November 2021 with 6 DUs
- The shadow detection significance is 4.2σ and 6.2σ for the Moon and Sun, correspondingly.
- The angular resolution values are $\sigma_{res}=0.49^\circ$ for the Moon and $\sigma_{res}=0.66^\circ$ for the Sun.
- The pointing accuracy is within the 68% contour.



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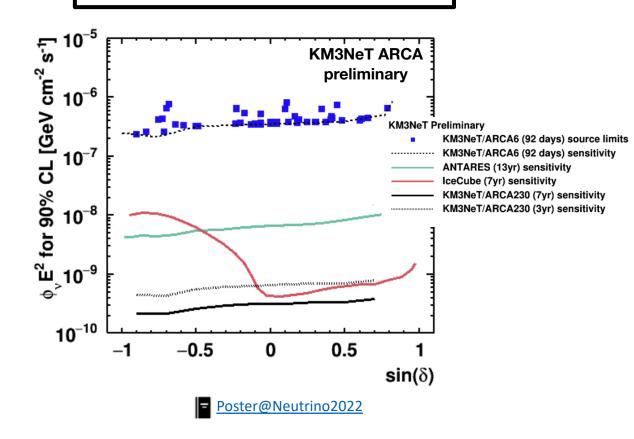


90% CL sensitivity to E⁻² neutrino flux Upgoing/horizontal $v_{\mu} + \bar{v}_{\mu}$ tracks



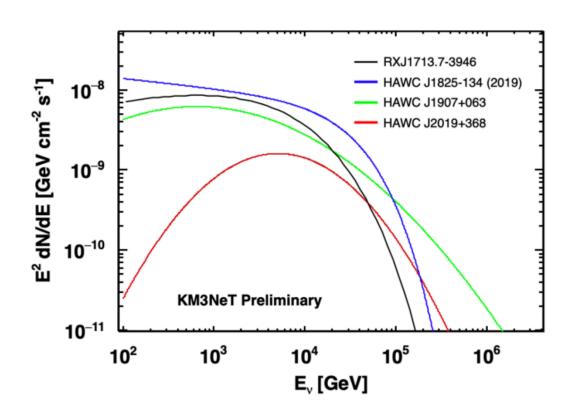
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Results obtained with the KM3NeT/ARCA detector with 6 working DUs

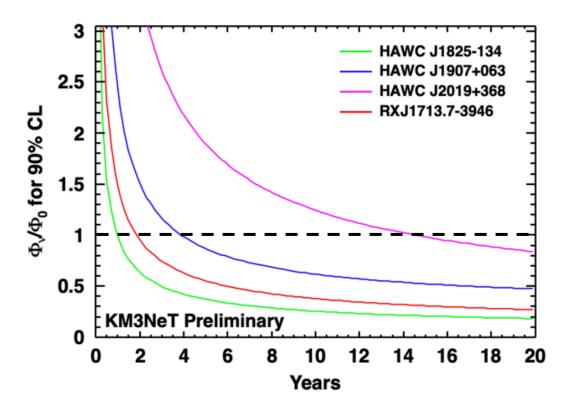




Expected ν fluxes assuming 100% hadronic emission mechanism



Sensitivity at 90% confidence level (Φ_0 is the expected flux) Several extended Galactic sources will be observable in a few years of operation of KM3NeT/ARCA with 2 BB.



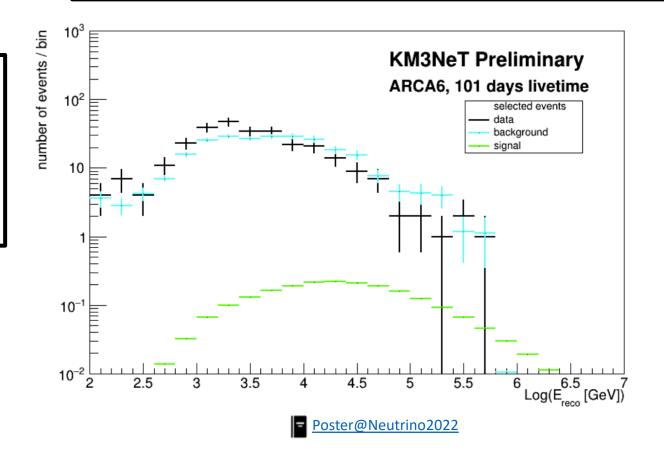


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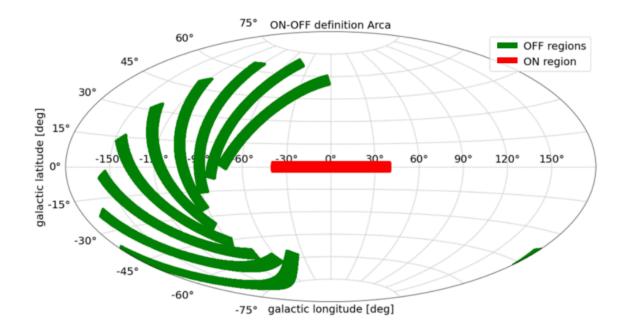
- KM3NeT/ARCA with 6 DUs, lifetime is 101 days
- Simulated signal 1.44 x 10⁻¹⁸ (E/100 TeV)^{-2.28} (IceCube diffuse flux)
- Obtained sensitivity for the corresponding flux: $17.3 \times 10^{-18} [\text{GeV}^{-1} \text{ cm}^{-2} \text{ s}^{-1} \text{ sr}^{-1}]$

Final energy distribution obtained with the reduced atmospheric muon contamination, keeping a high signal efficiency



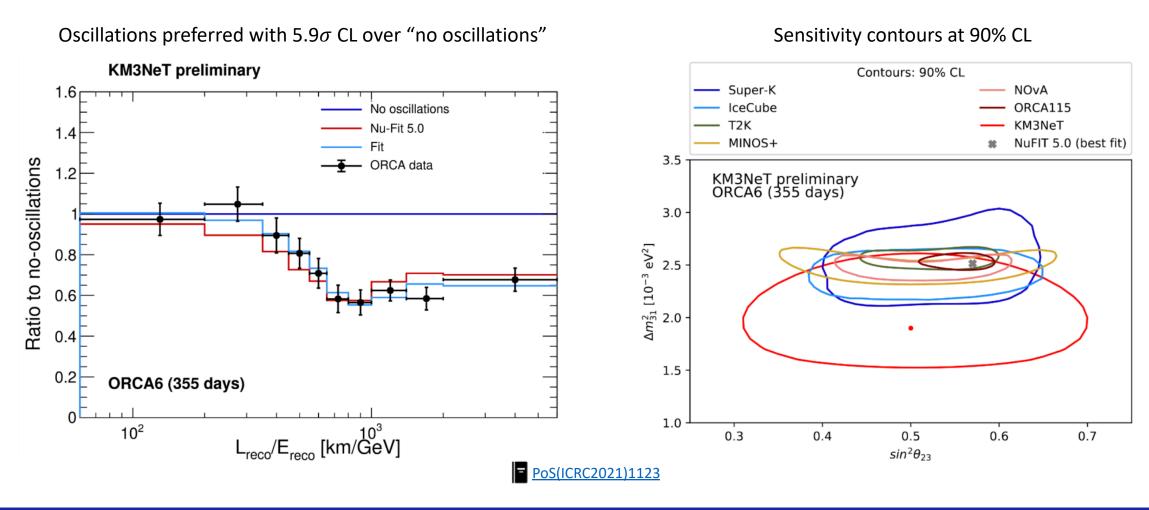


- KM3NeT/ARCA with 6 DUs, lifetime is 101 days
- ON region is the Galactic ridge: $|L_{gal}| < 40^{\circ}$, $|B_{gal}| < 3^{\circ}$
- OFF regions are shifted in time avoiding the Fermi bubbles
- The simulated signal within the ON region follows a power-law spectra: $1.2 \times 10^{-8} (E^{-2.4}/1 \text{ GeV}) [\text{GeV}^{-1} \text{ cm}^{-2} \text{ s}^{-1} \text{ sr}^{-1}]$
- No statistically significant excess found.
 Upper limit: 6.2 x 10⁻⁴ [GeV⁻¹ cm⁻² s⁻¹ sr⁻¹]



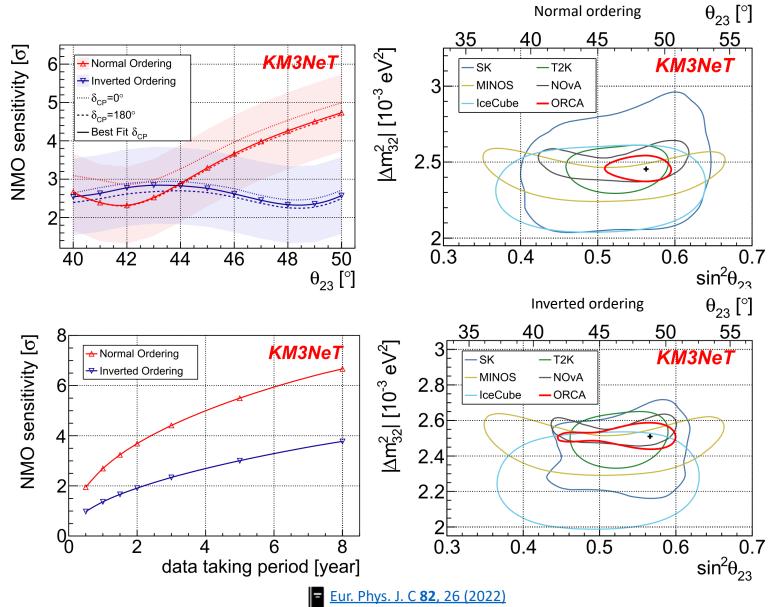
Poster@Neutrino2022

- The first KM3NeT/ORCA oscillation results with only 6 DUs!
- Clear effect of oscillations observed





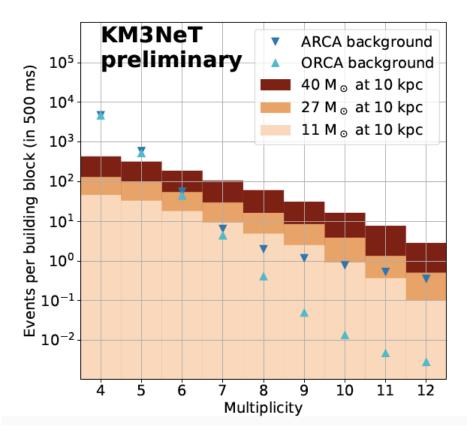
- KM3NeT/ORCA in its full configuration. Three years of data taking
- Sensitivity for neutrino mass ordering at 4.4σ (normal) and 2.3σ (inverted)
- Competitive sensitivity to Δm_{32}^2 and θ_{23}

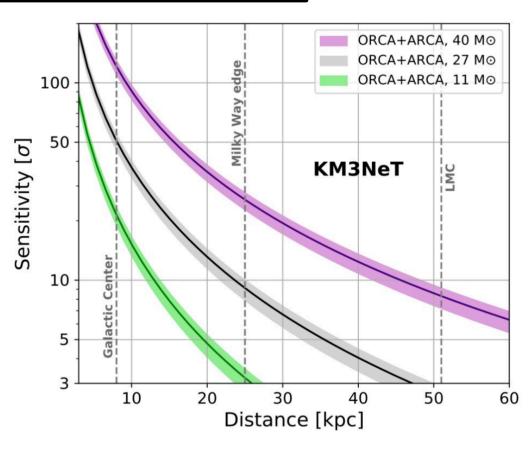




Supernova monitoring

- Neutrinos with E<100 MeV expected at several stages of core collapse
- Low-energy neutrino detection is based on the coincidences in single DOMs thanks to the multi-PMT DOM configuration





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- The KM3NeT research infrastructure is being built in the Mediterranean Sea and includes two neutrino telescope: KM3NeT/ORCA (Toulon, France) and KM3NeT/ARCA (Capo Passero, Italy)
- Very broad energy range (MeV PeV) to explore core-collapse supernovae, neutrino oscillations and astronomy
- To date, 21/230 DUs were deployed in KM3NeT/ARCA and 18/115 DUs in KM3NeT/ORCA
- Several physics results were already achieved and confirmed the expectations
- New exciting results will be presented at the summer conferences, ICRC2023

Stay tuned!