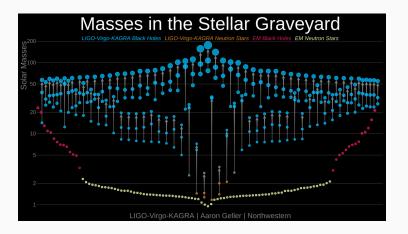


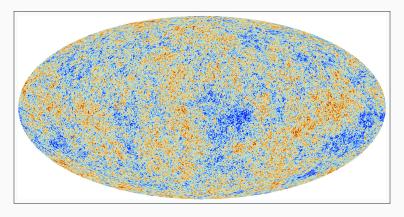
1) Next milestone in GW astronomy: detection of a stochastic GWB



- Status: LVK now routinely observe transient GW signals of astrophysical origin
- Next milestone: Detection of a stochastic GW background (GWB)
- Big news on June 29: Compelling evidence for a GWB reported by several teams!

20th century

[PLANCK Collaboration]

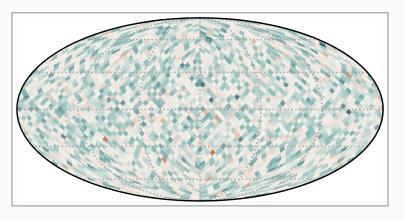


CMB: Cosmic microwave background Relic photons from the early Universe

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21th century

[Sato-Polito, Kamionkowski: 2305.05690]



GWB: Gravitational-wave background Relic gravitational waves from the early Universe $\sim or \sim$ astrophysical signal

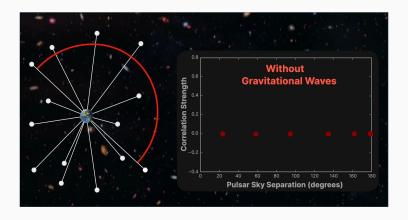
3) PTAs are galaxy-sized GW detectors that allow us to search for nHz GWs



Array of pulsars across the Milky Way \rightarrow GW detector of galactic dimensions!

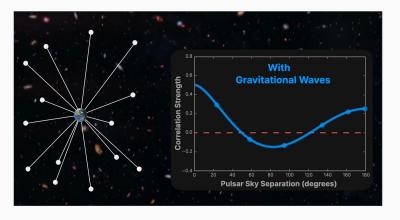
- Look for tiny distortions in pulse travel times caused by nanohertz GWs.
- Signal builds up over time; monitor PTA over years and decades.

4) Hallmark signature in cross-correlation of timing residuals of pulsar pairs



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4) Hallmark signature in cross-correlation of timing residuals of pulsar pairs

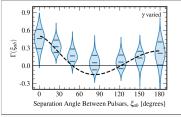


Quadrupolar correlations described by Hellings–Downs (HD) curve [Hellings, Downs: Astrophys. J. 265 (1983) L39]

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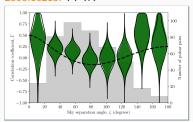
5) Major announcment on June 29: compelling evidence for HD correlations

2306.16213: NANOGrav



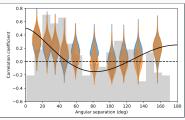
68 pulsars, 16 yr of data, HD at \sim 3 \cdots 4 σ

2306 16215: PPTA



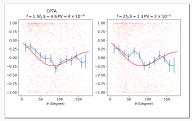
32 pulsars, 18 yr of data, HD at $\sim 2 \sigma$

2306.16214: EPTA+InPTA



25 pulsars, 25 yr of data, HD at \sim 3 σ

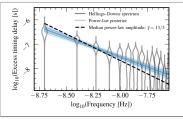
2306.16216: CPTA



57 pulsars, 3.5 yr of data, HD at \sim 4.6 σ

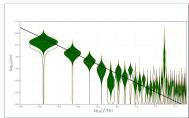
6) Spectral characterization: at present, in terms of a simple power-law fit

2306.16213: NANOGrav



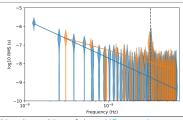
68 pulsars, 16 yr of data, HD at \sim 3 \cdots 4 σ

2306.16215: PPTA



32 pulsars, 18 yr of data, HD at \sim 2 σ

2306.16214: EPTA+InPTA



25 pulsars, 25 yr of data, HD at \sim 3 σ

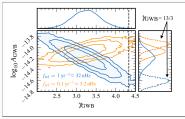
2306.16216: CPTA



57 pulsars, 3.5 yr of data, HD at \sim 4.6 σ

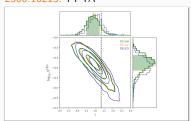
7) Results are consistent among the most recent PTA data sets

2306.16213: NANOGrav



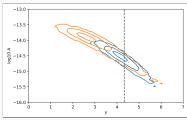
68 pulsars, 16 yr of data, HD at \sim 3 \cdots 4 σ

2306.16215: PPTA



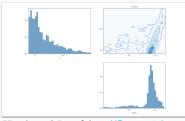
32 pulsars, 18 yr of data, HD at \sim 2 σ

2306.16214: EPTA+InPTA



25 pulsars, 25 yr of data, HD at \sim 3 σ

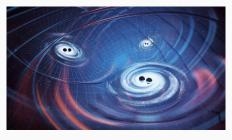
2306.16216: CPTA



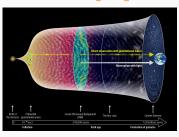
57 pulsars, 3.5 yr of data, HD at \sim 4.6 σ

8) Interpretation: SMBHBs (realistic) or new physics (speculative)

Supermassive black-hole binaries



9 GWs from the Big Bang



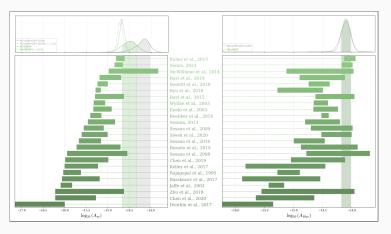
- SMBHBs: No SMBHB mergers observed → data-driven field thanks to PTAs
- New physics: Probe cosmology at early times, particle physics at high energies

BSM scenarios: Inflationary gravitational waves, scalar-induced gravitational waves, cosmological phase transition, cosmic strings, domain walls, axions, and many more

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9) SMBHBs: viable explanation, but unexpected corners of parameter space

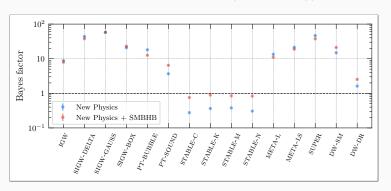
[NANOGrav 2306.16220]



- Parameter shifts towards larger GWB amplitudes than previously expected
- Self-consistent phenomenological description of environmental interactions
- Generally higher binary masses or densities, or highly efficient binary mergers

10) New physics: many models can fit the data, but situation inconclusive

[NANOGrav 2306.16219] [See also: EPTA 2306.16227]



Bayesian model comparison

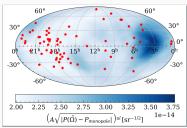
Reference model: $\mathcal{H}_0 = \{SMBHBs \text{ only}\}\$

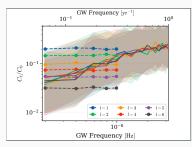
- Many BSM models reach Bayes of order 10 · · · 100.
- Interesting but not conclusive. Lots of uncertainties in SMBHB and BSM models.
- Bayes factors are sensitive to prior choices. No unique null distribution for \mathcal{H}_0 .

11) Complementary observables: anisotropies, continuous-wave signals



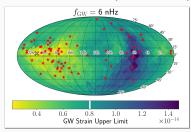


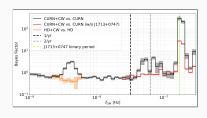




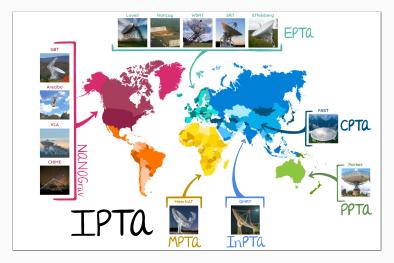
Continuous wave

[NANOGrav 2306.16222]





12) Promise: deep insights into galaxy and BH evolution, or BSM physics



- Status: Common-spectrum process; $2\cdots 4.6\,\sigma$ evidence for HD correlations
- Next: HD correlations at 5 σ , spectral shape, anisotropies across the sky, ...

Summary: Twelve key messages in twelve minutes

- 1. Next milestone in GW astronomy: detection of a stochastic GWB
- 2. GWB: 21st-century equivalent of the 20th-century discovery of the CMB
- 3. PTAs are galaxy-sized GW detectors that allow us to search for nHz GWs
- 4. Hallmark signature in cross-correlation of timing residuals of pulsar pairs
- 5. Major announcment on June 29: compelling evidence for HD correlations
- 6. Spectral characterization: at present, in terms of a simple power-law fit
- 7. Results are consistent among the most recent PTA data sets
- 8. Interpretation: SMBHBs (realistic) or new physics (speculative)
- 9. SMBHBs: viable explanation, but unexpected corners of parameter space
- 10. New physics: many models can fit the data, but situation inconclusive
- 11. Complementary observables: anisotropies, continuous-wave signals
- 12. Bright future: deep insights into galaxy and BH evolution, or BSM physics

Stay tuned!

And thanks a lot for your attention