

The XVIII International Conference on Topics in Astroparticle
and Underground Physics (TAUP2023)



The status and prospects for China Jinping Underground Lab(CJPL)

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on behalf of CJPL team group
Tsinghua University
Aug. 28, 2023 , Vienna



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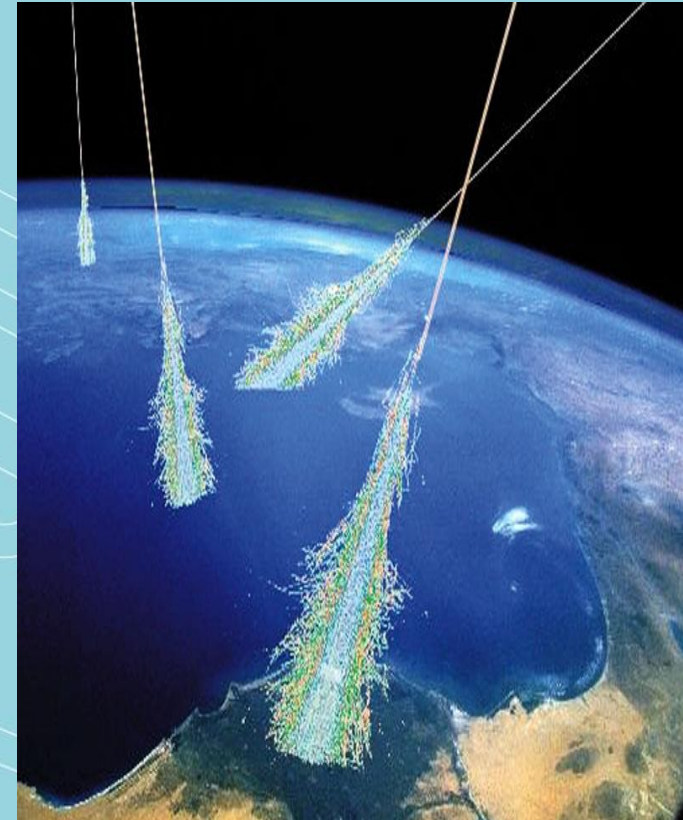
- I. Overview**
- II. The status of CJPL-I**
- III. Current status of CJPL-II**
- IV. Experiments in CJPL-II**
- V. Summary**

An isometric line drawing of a cable management system. It shows a network of white pipes and rectangular cable trays. Some pipes have small rectangular components attached to them. The entire system is set against a light blue background that features a diagonal line, creating a sense of depth and perspective.

I. Overview

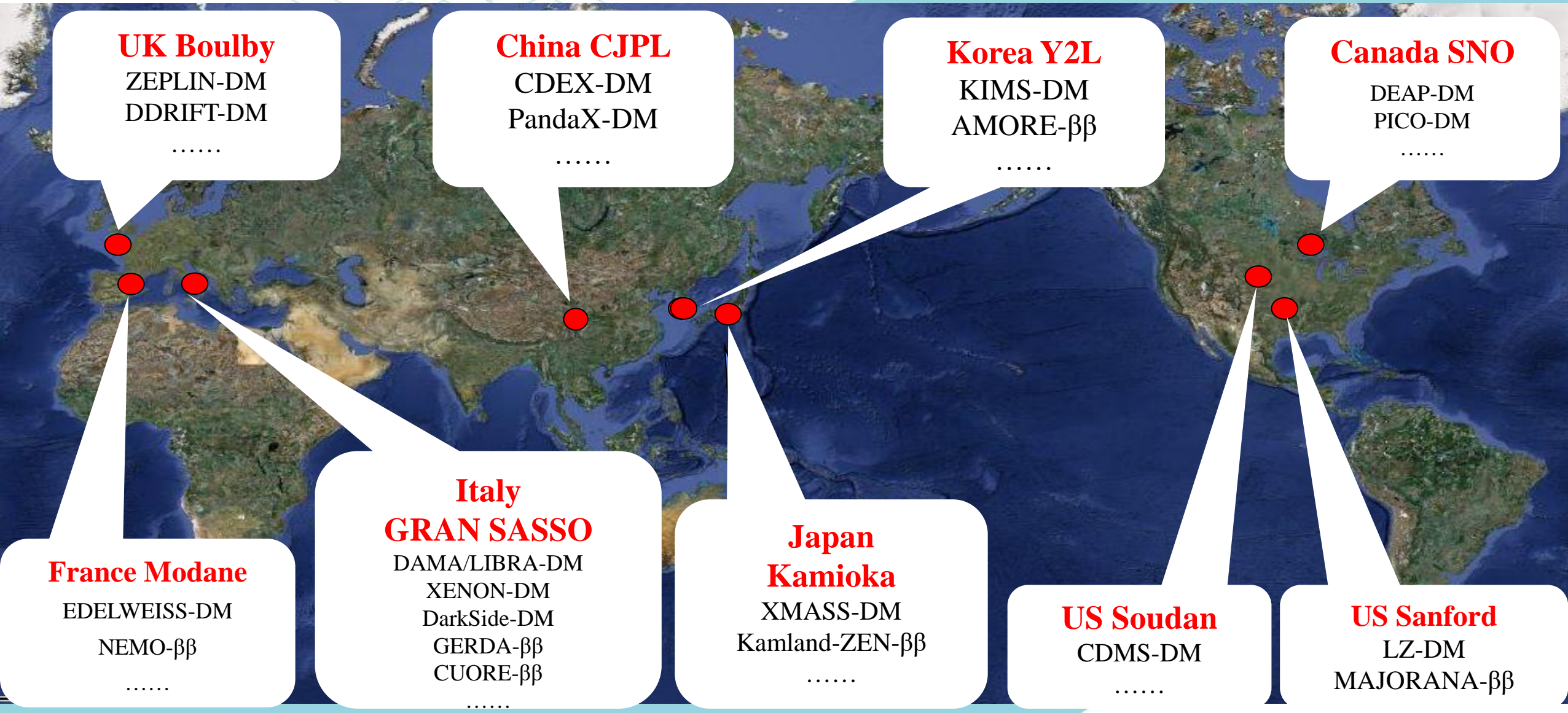
Go underground for frontier sciences

- Dark matter
- Neutrino
- Nuclear astrophysics
- Proton decay
- Biology
- Seismology
- Rock mechanics
-

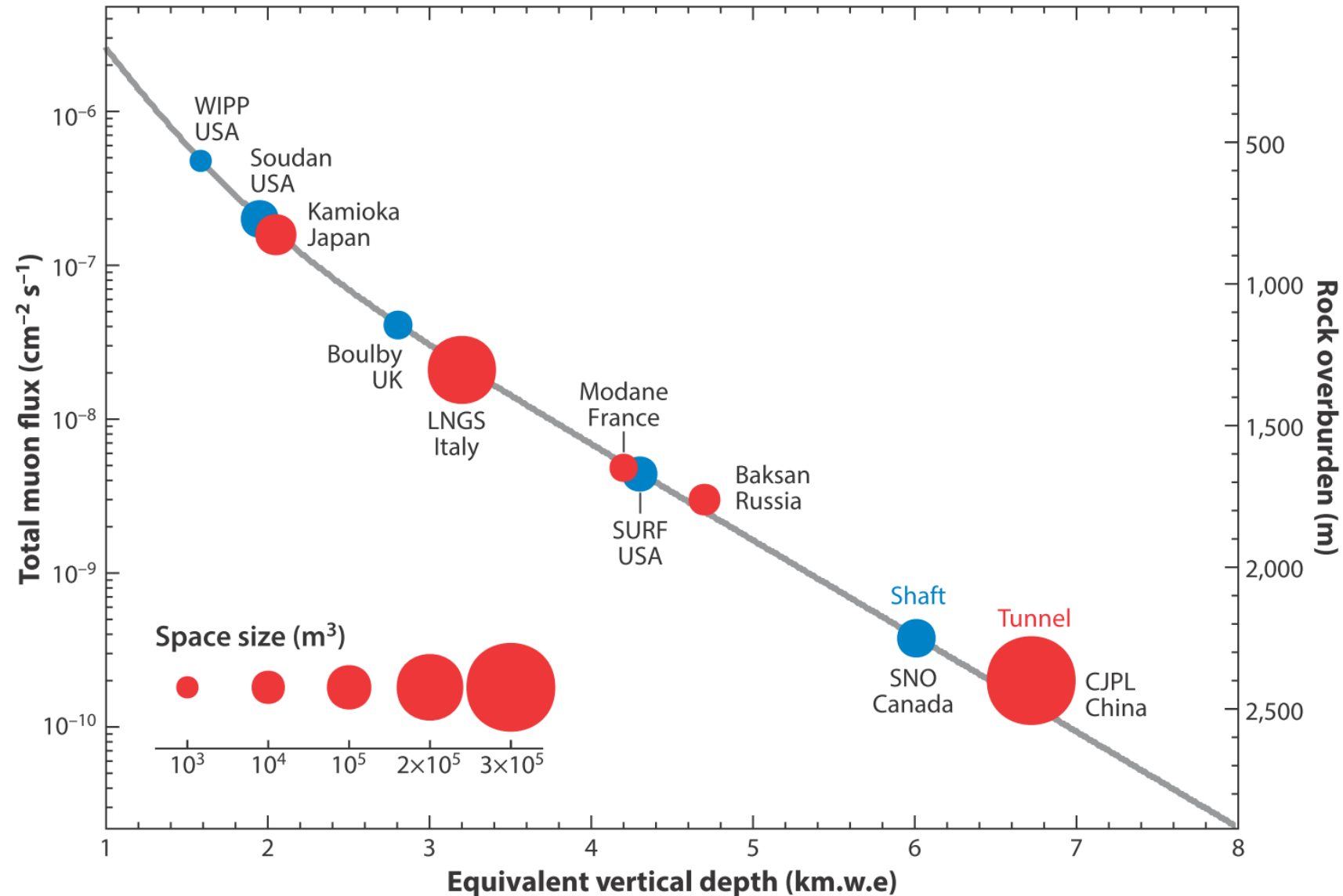


Shielding Cosmic rays

Main underground laboratories in the world



Main underground laboratories in the world



Jinping Hydroelectric Power Plants Project



Yalong river meets Jinping Mountain:

- The largest tributary of Yangzi River
- Jinping river bend: 150km long



Jinping traffic tunnel

- 17.5km long x 2
- **Overburden:** max. 2400m, 73% of length >1500m.
- Finished on Aug. 8, 2008

Tunnel Layout inside Jinping Mountain

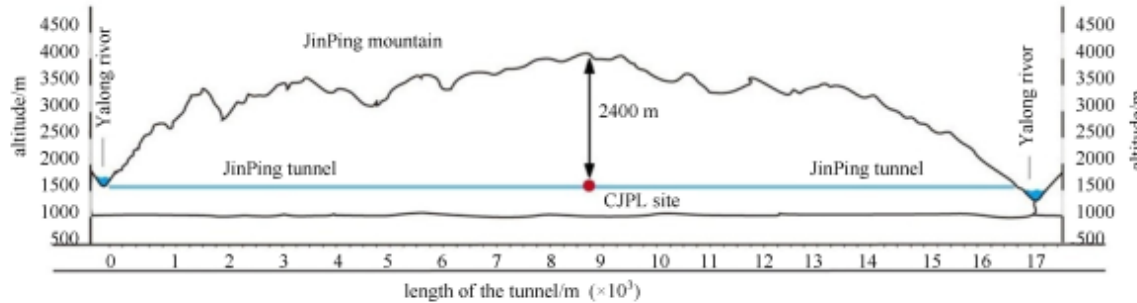
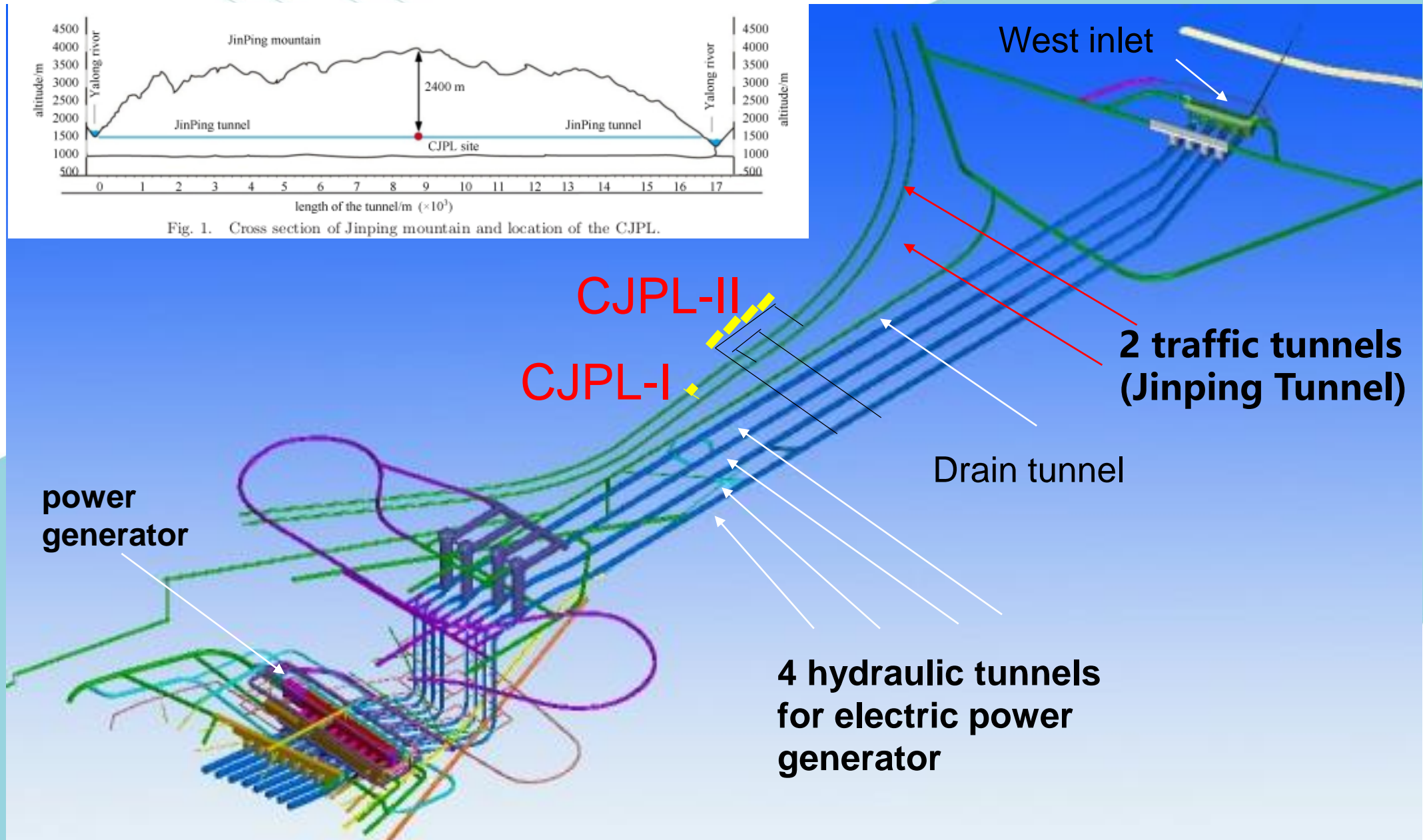
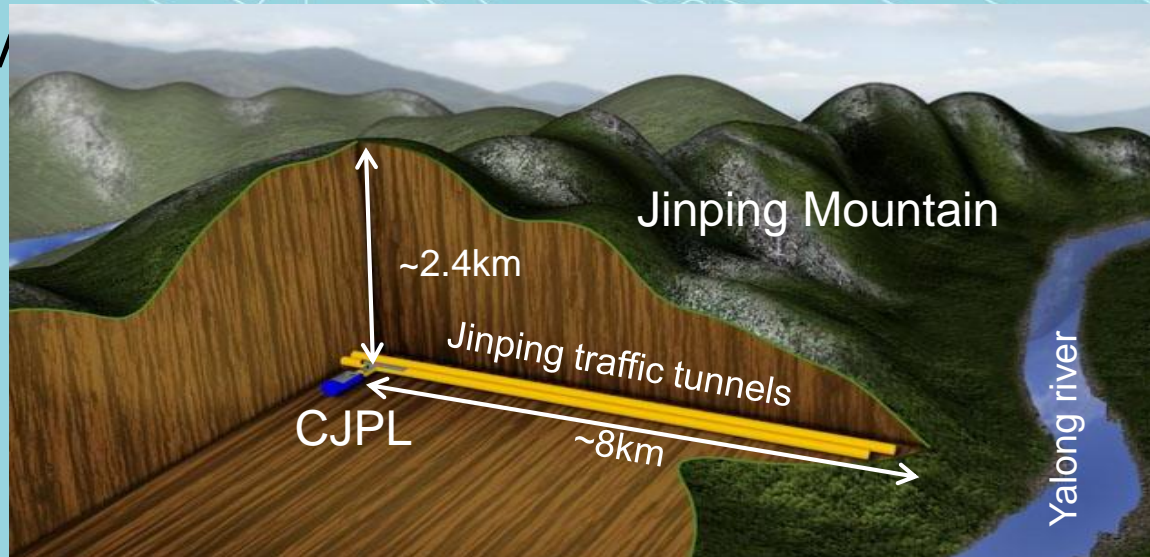
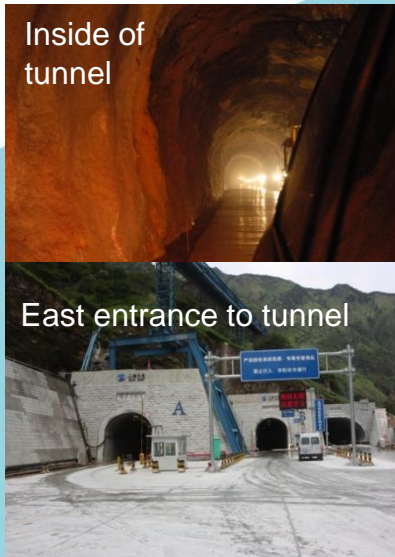


Fig. 1. Cross section of Jinping mountain and location of the CJPL.



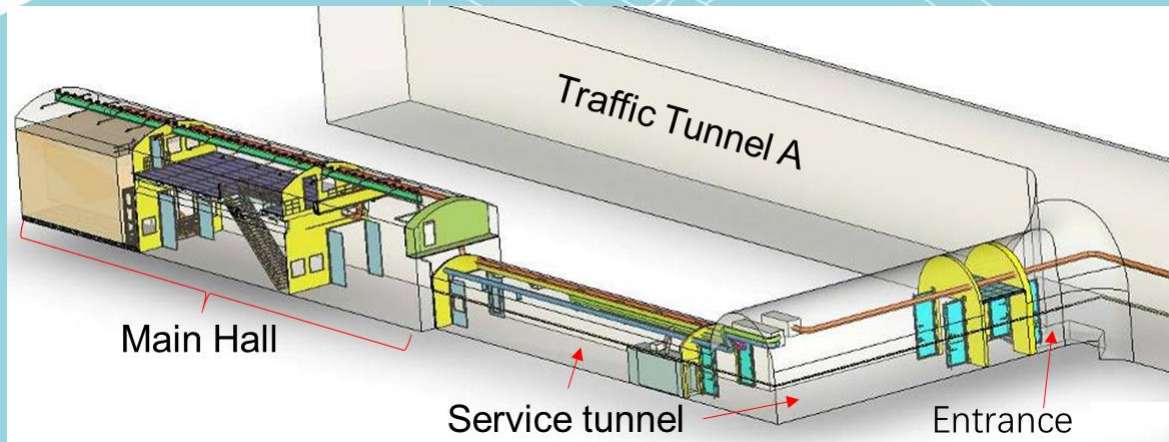
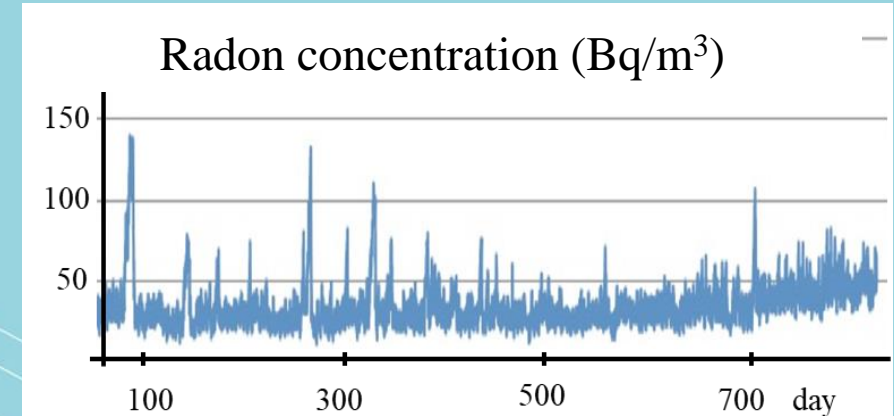
China JinPing Underground Laboratory(CJPL)

- An ideal site for an underground laboratory!
- THU-YHPC cooperation on a new underground lab started in May 2009.
- CJPL-I site selected in Aug. 2009.
- Rock sampling and in-situ measurement to study environmental



CJPL Features

- Open on Dec. 12, 2010.
- Total space: $\sim 4000 \text{ m}^3$.
- Main Hall: $6.5\text{m(W)} \times 6.5\text{m(H)} \times 42\text{m(L)}$.
- Deepest underground laboratory with a rock overburden of 2400m.
- Low Muon flux and low environmental background.



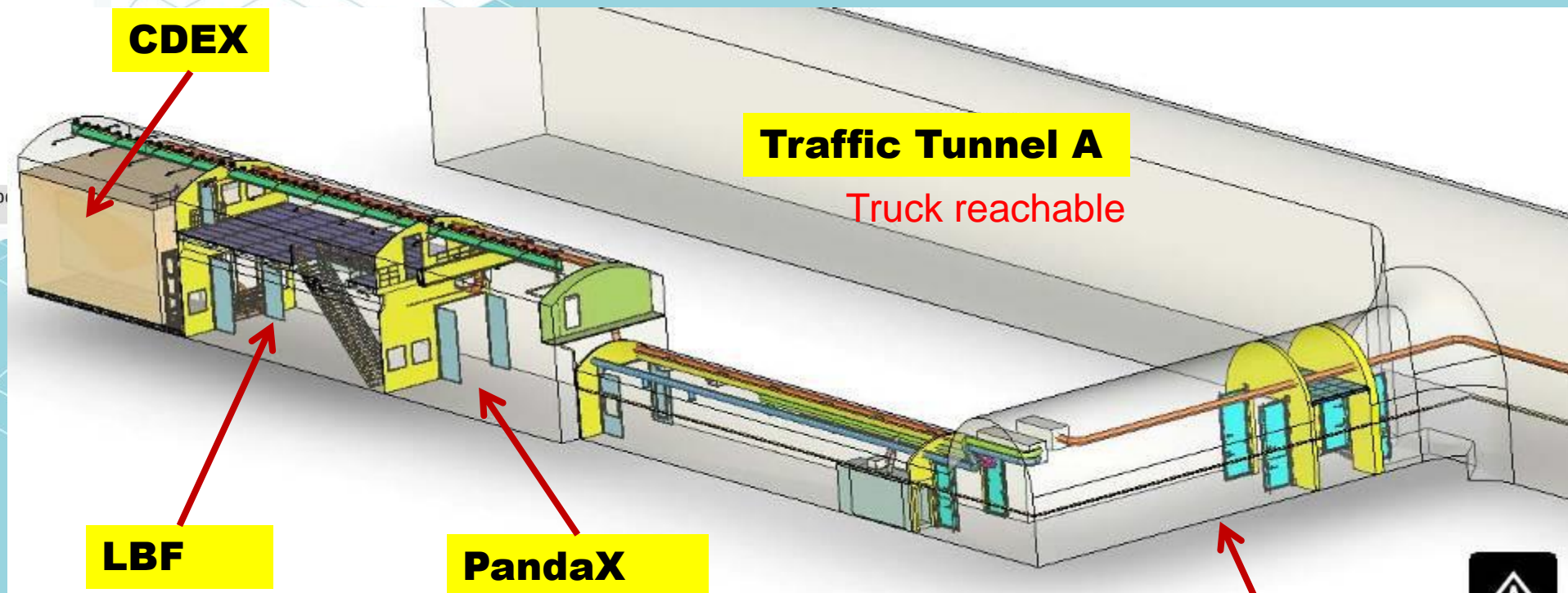
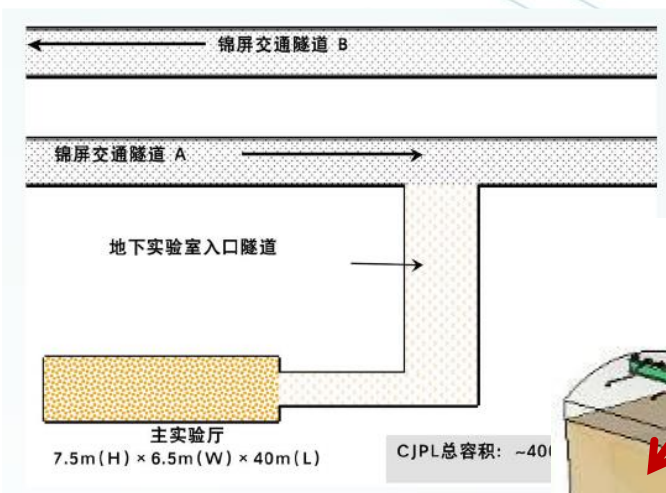
Rock sample(Marble) measurement by Germanium

(Unit : Bq/kg)	K-40	Ra-226 (609keV)	Th-232 (911keV)
Rock Sample	< 1.1	1.8 ± 0.2	< 0.27
Ground Level(Beijing)	~ 600	~ 25	~ 50



II. The status of CJPL-I

Layout of CJPL-I



- Total space: 4000 m³
- Main Lab Space: 6.5(W) x 6.5(H) x 42(L)

Experiments in CJPL-I



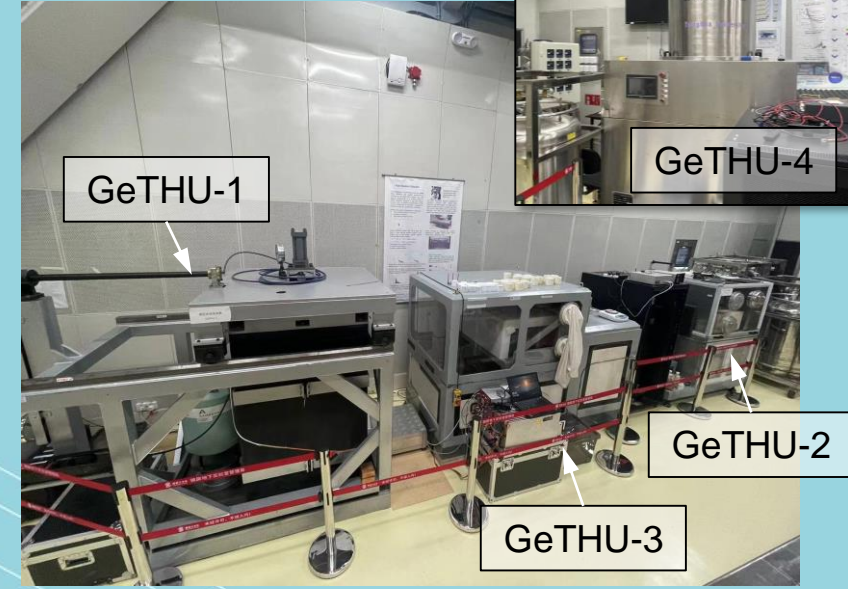
CDEX experiment



PandaX(removed)



Jinping neutrino



Low-background γ spectrometers

❑ Physics experiments:

- 2 dark matter experiments: CDEX, PandaX (now to CJPL-II)
- 1 neutrino experiment: Jinping Neutrino experiment

❑ Low background counting facilities:

- 4 low-background gamma spectrometers: GeTHU1, 2, 3, 4



III. Current status of CJPL-II

for frontier physics

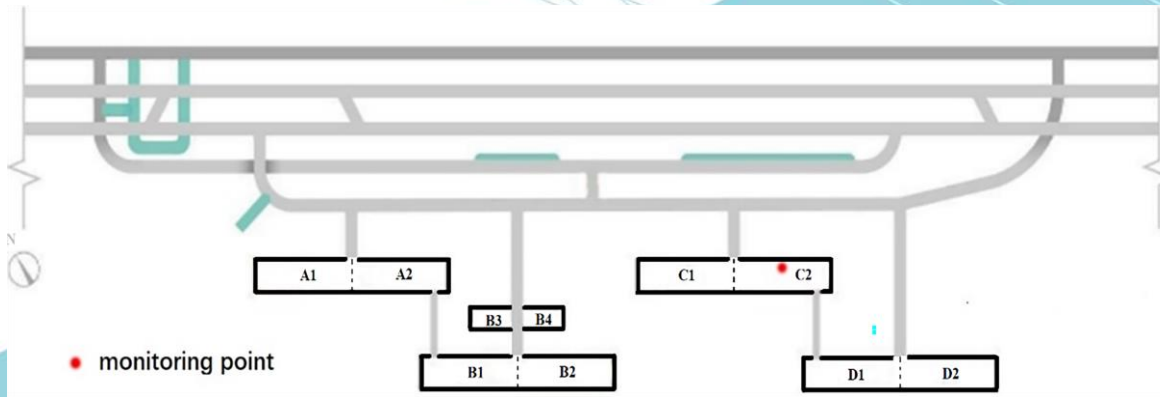


- **Internal Construction started at 2020/12**
- **Civil engineering will be finished in Oct. 2023**
- **All constructions will be completed at 2024**
- **4 experiment halls (A-D), total space of $>300,000 \text{ m}^3$**
- **Will be the deepest and largest underground lab worldwide**

Key Project: water-resistant and radon suppression

□ Monitoring Rn-222 in Hall-C2 in bare cavern state

- Rn-222 concentration in 14~776 Bq/m³ (average 201 Bq/m³)
- Data fitting demonstrates a Rn-222 change cycle of 12.7 month

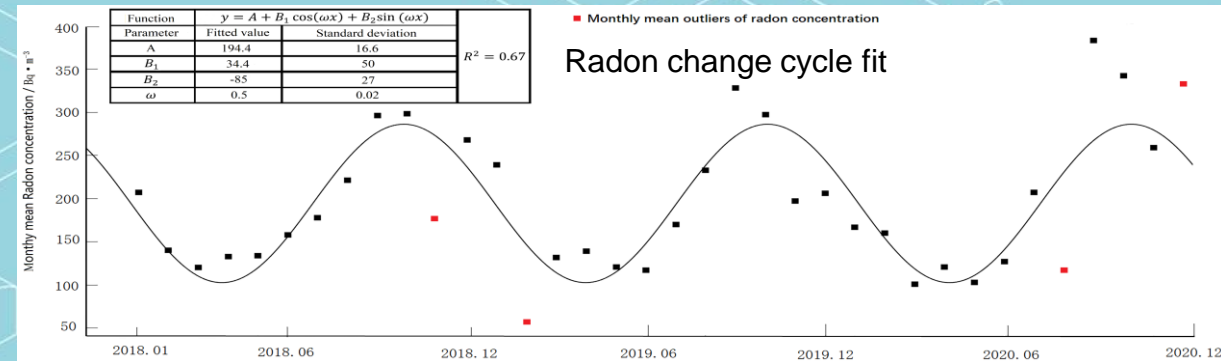
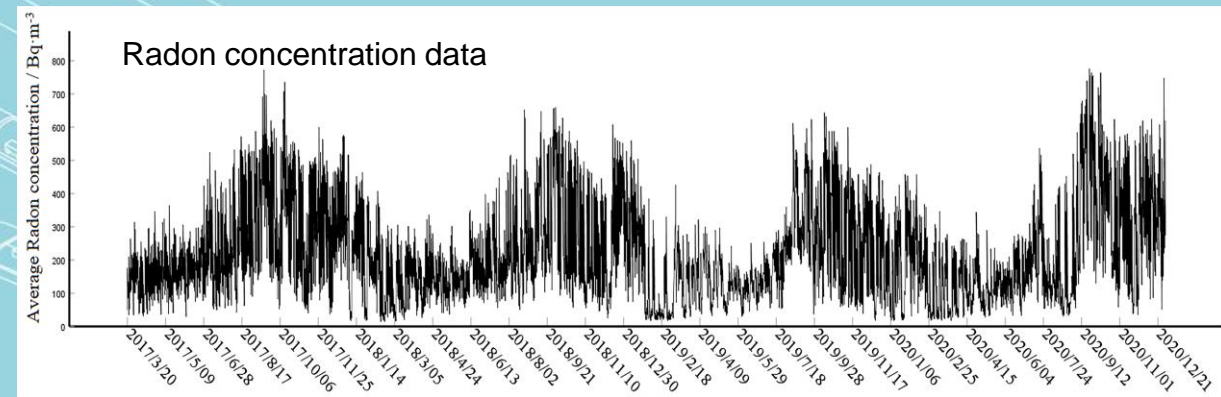


Det: Alpha-GUARD PQ2000

Loc: CJPL-II hall-C2

Date: 2017.3-2020.12

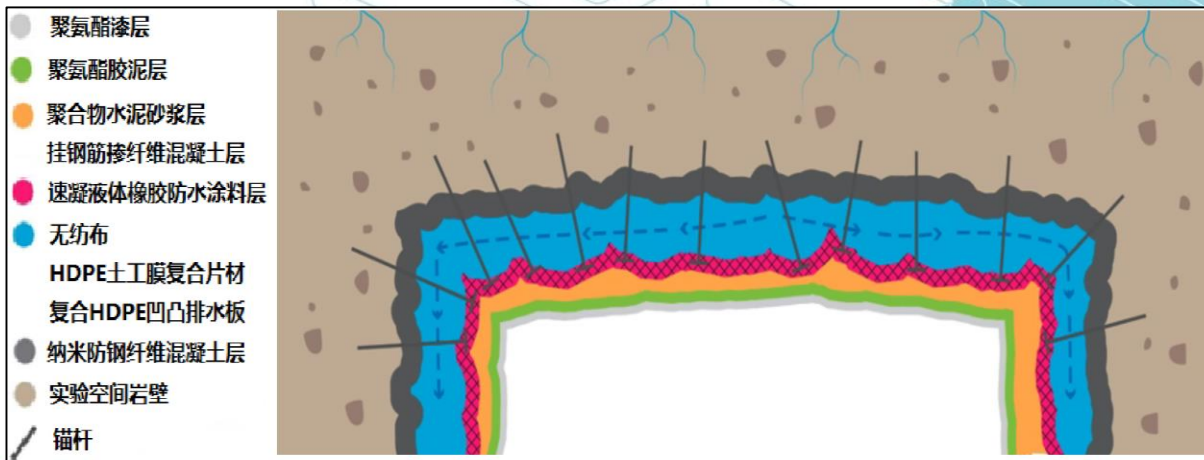
Mea-time: 60 min/point



Key Project: water-resistant and radon suppression

□ Water-Resistant and Radon Suppression (WRRS) layer

- A dedicated structure using multi-layer protection against water and radon



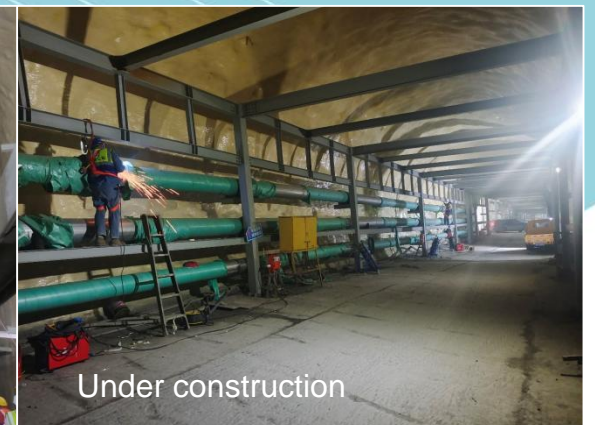
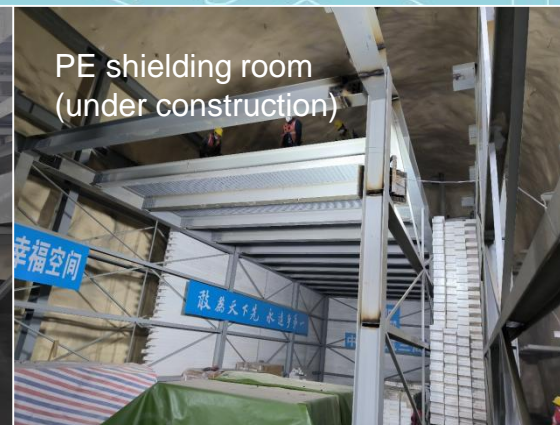
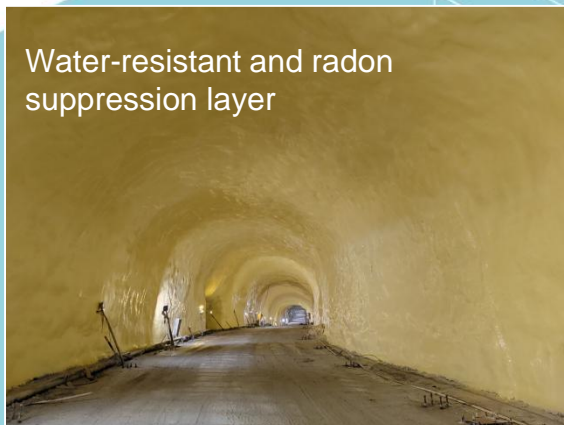
Features of WRRS layer:

- Preventing water and radon permeating from the rock into lab space simultaneously
- Covering the full-space of each experiment hall (wall and floor)
- Using low-background materials in WRRS
- 7 layers, 9 processing, dozens of materials



Civil engineering of CJPL-II

- ❑ **Water-resistant and radon suppression:** construction on wall is completed except for B2 hall, and processing on floor is ongoing
- ❑ **Steel structure:** Completed for A/B halls, and 70% completion for the rest
- ❑ **PE shielding room:** 1427 m³ PE blocks entered, and 40% completion for construction
- ❑ **Mechanical and electrical installation:** 60% completion for hall-A and 15% for 1/2 auxiliary tunnels



Key Project: Large Nitrogen vessel shielding

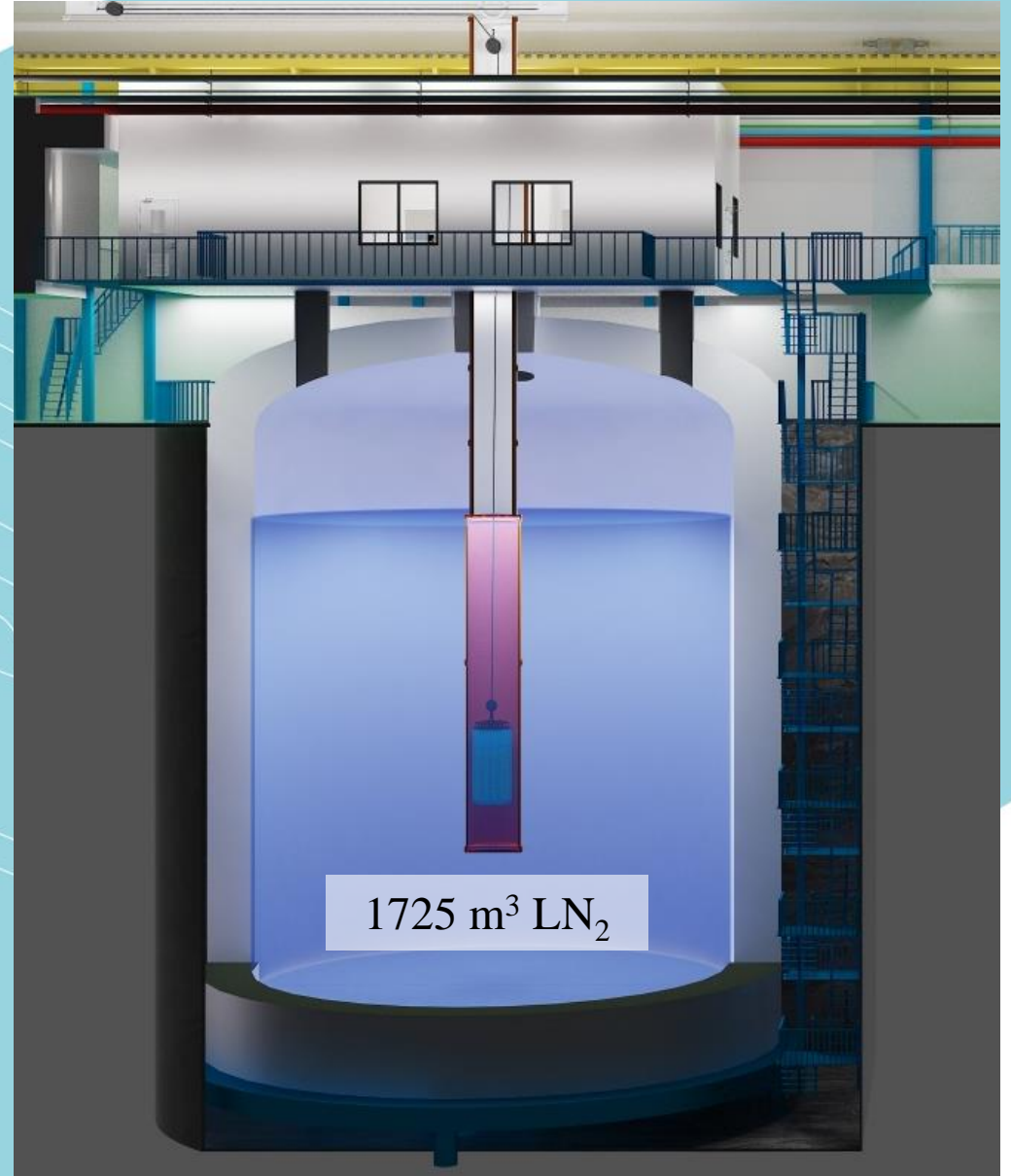
- Large LN tank completed
- Clean room under construction
- LN volume: $\Phi 13\text{m} \times \text{H}13\text{m}$, $\sim 1725 \text{ m}^3$
- LN filling planed



Pit in Hall C1



1725 m³ LN₂ tank in Hall C1

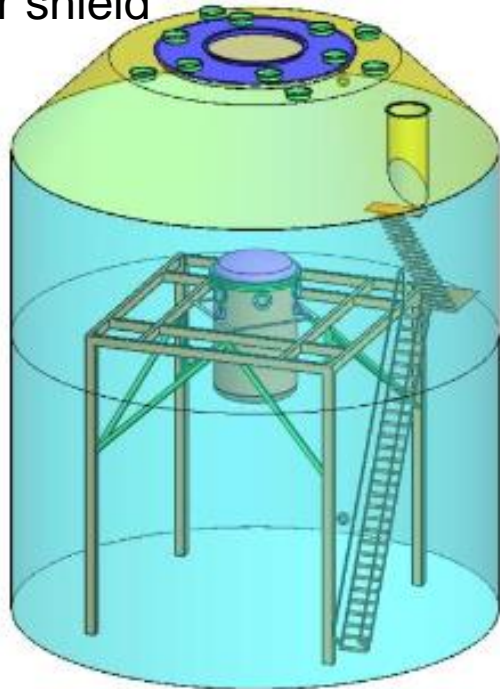


Key Project: Large pure water tank shielding

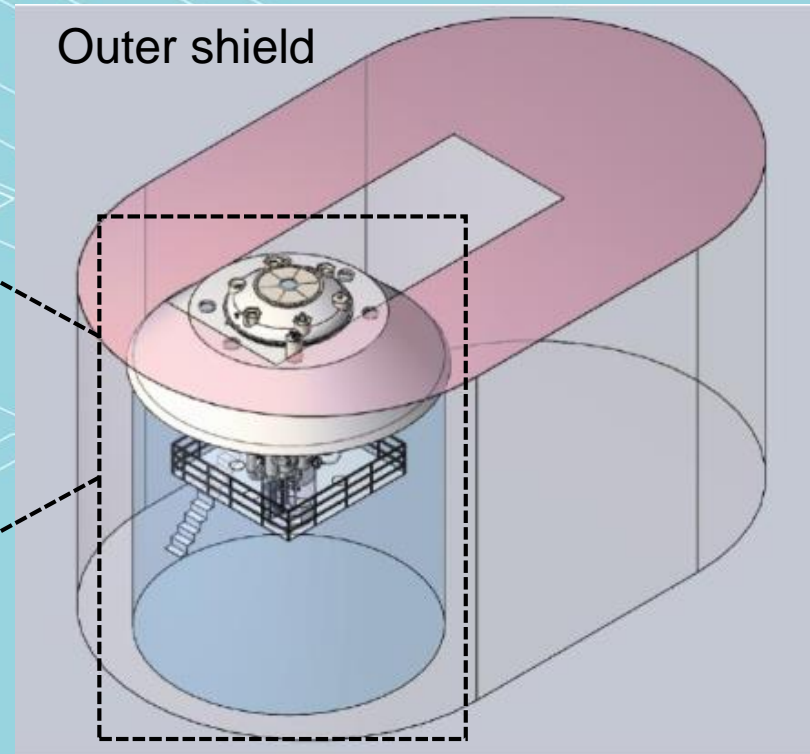
❑ Located in Hall-B2:

- Outer shield (4500 m³ water) and inner shield (1000 m³ water)
- 12 L low-level radon measure device completed
- Water-radon system pressure test completed

Inner shield



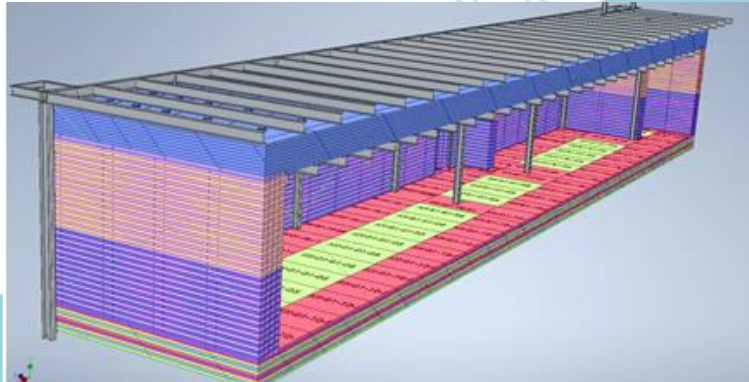
Outer shield



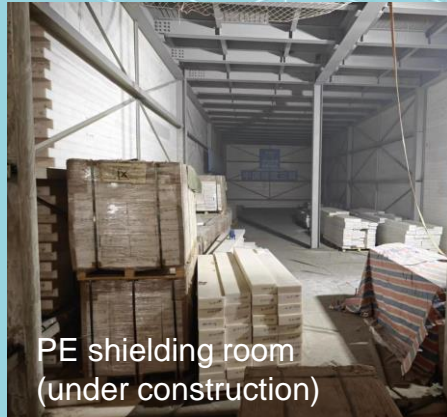
Key Project: Assemble combine shielding

❑ Multi-purpose PE, Copper, and Lead shield

- PE room under construction, 40% completion
- Inner copper and lead combined shielding module under preparation



PE shielding room
(under construction)



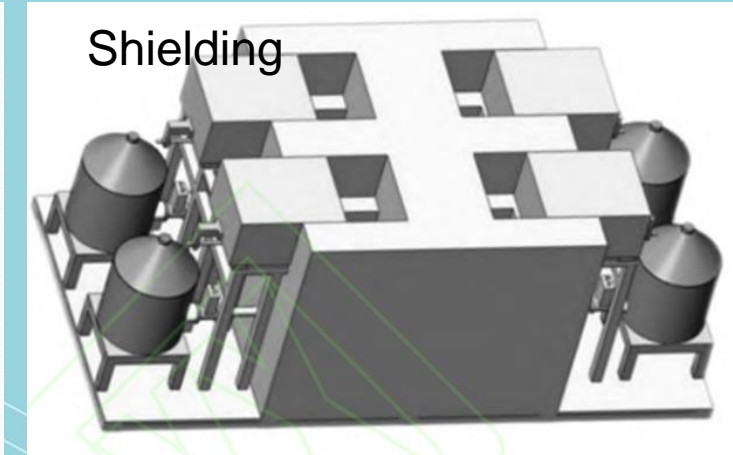
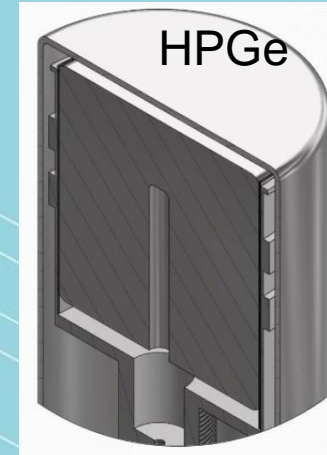
PE shielding room
(under construction)



Key Project: Ultra-low-background γ spectrometers

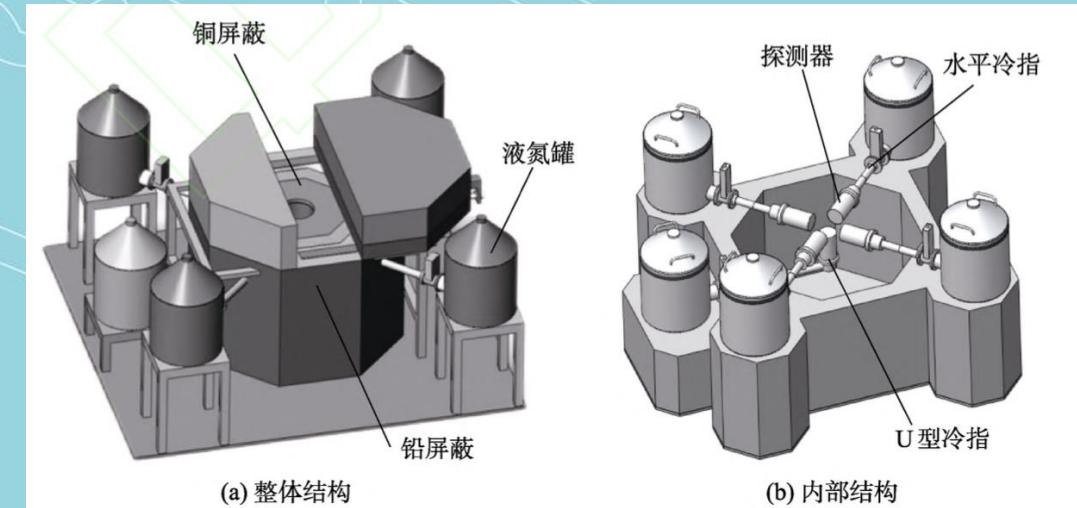
□ mBq/kg spectrometers (GeTHU)

- Extension of current GeTHU-1/2/3/4
- Detection limit: mBq/kg scale
- Total 15 HPGe γ spectrometers
- Commercial HPGe detectors



□ μ Bq/kg spectrometers (ARGUS)

- Detection limit: μ Bq/kg scale
- 5 commercial HPGe detectors
- Low-background shielding



Key Project: Cavern safety monitor system

□ Detectors:

- Multi-detectors to monitor the rock vibration and displacement.



microseismic monitoring system



gravity meter



seismograph



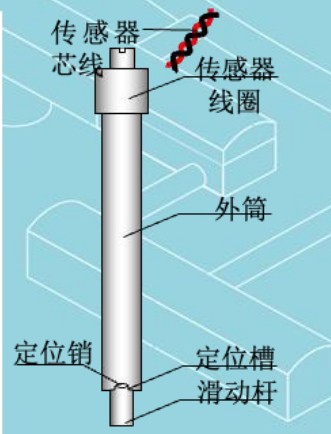
Bolt stress meter



osmometer



Disturbed Stress monitor



displacement sensor



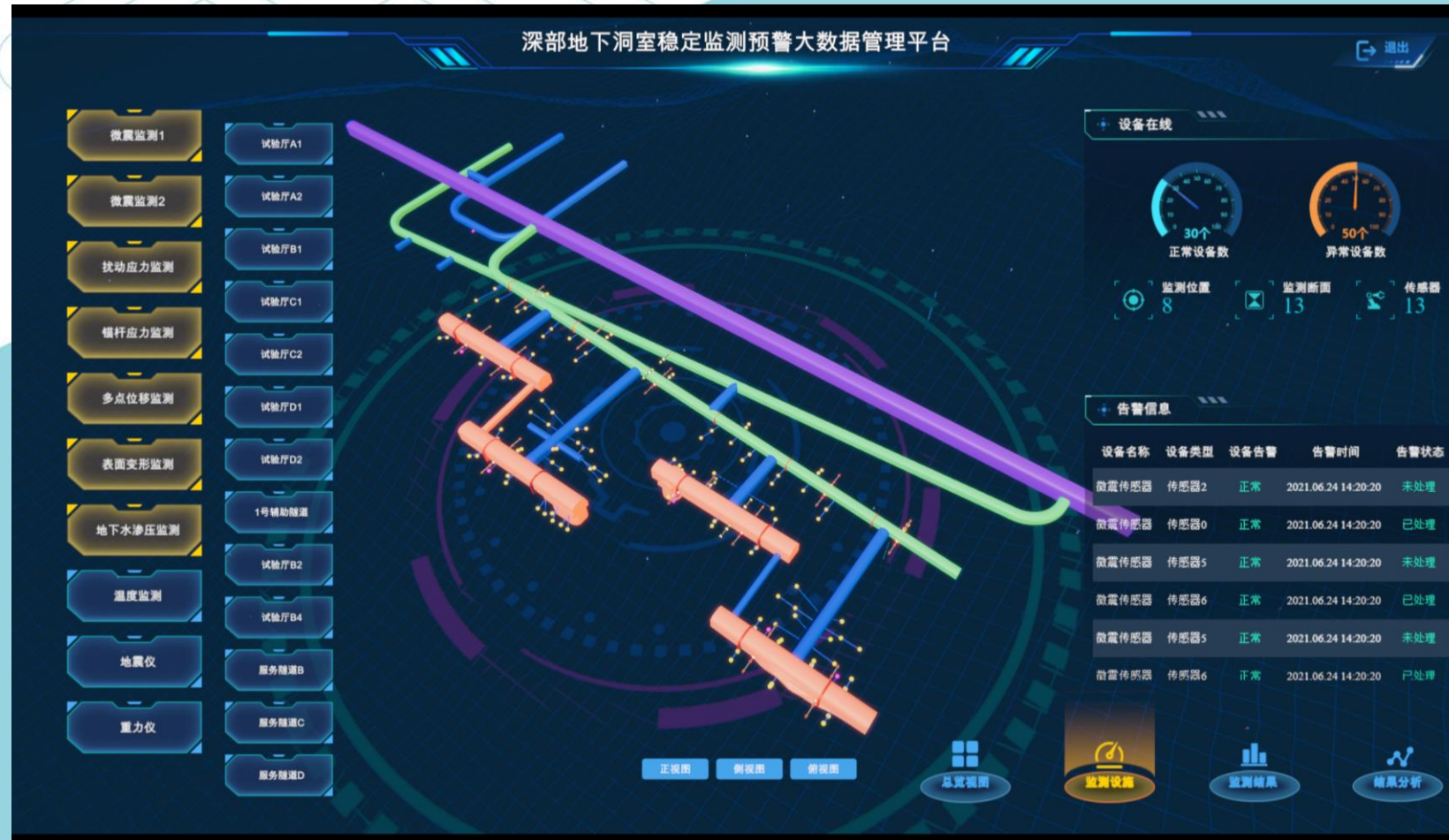
Surface deformation monitor



rock cracking monitor

Key Project: Cavern safety monitor system

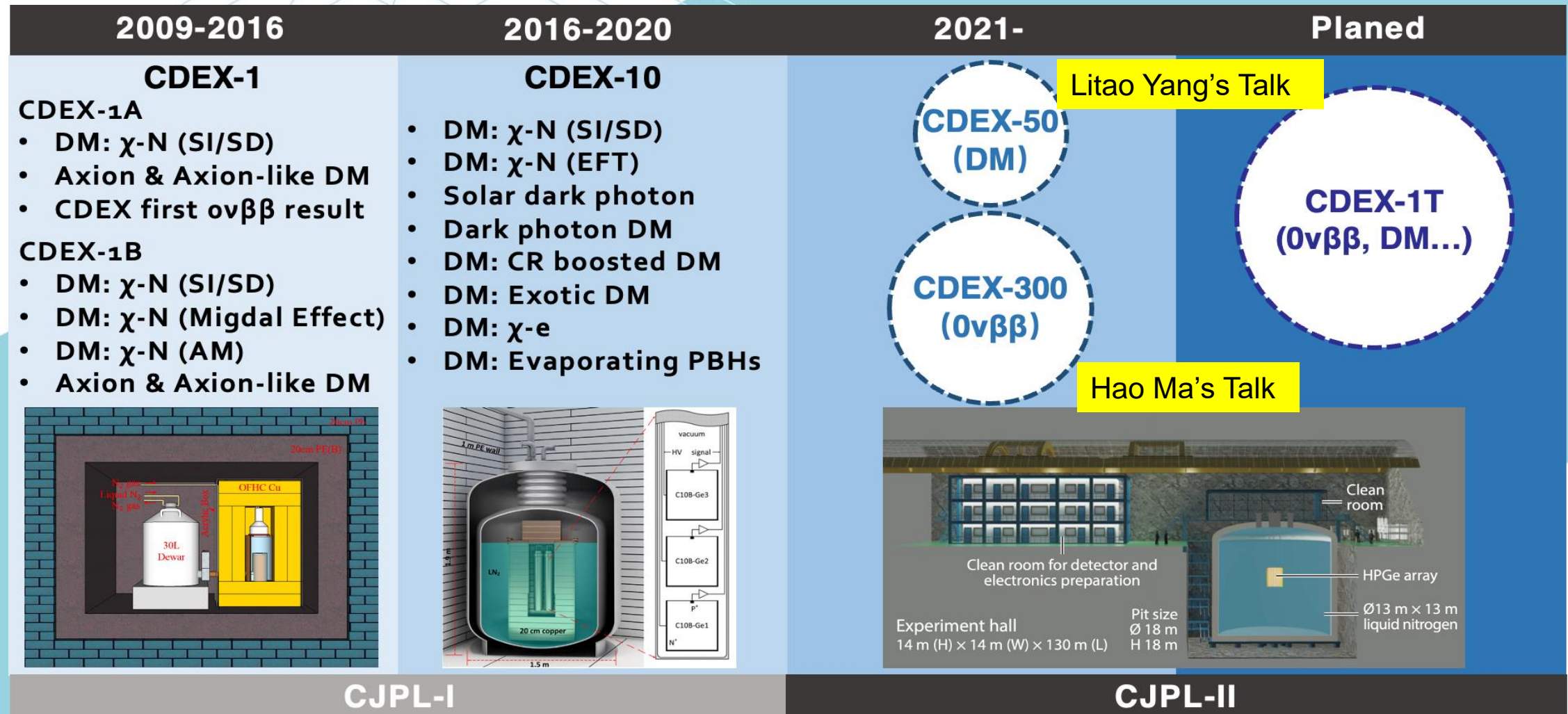
□ User Interface:



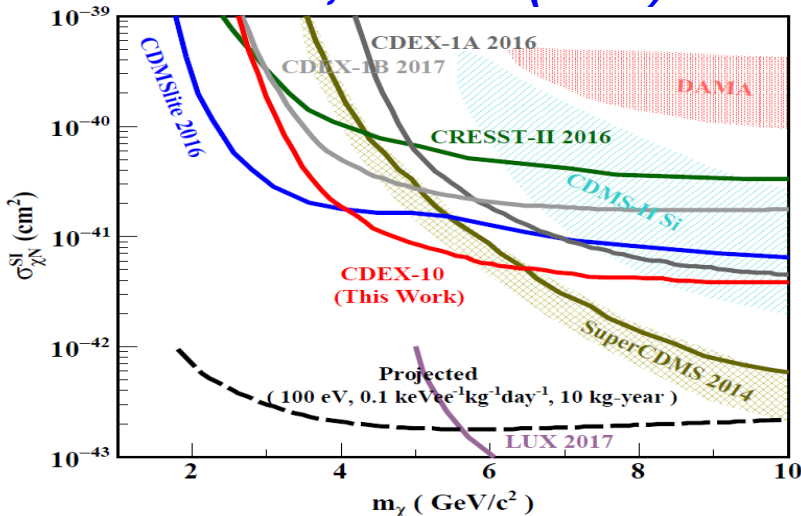
IV. Experiments in CJPL-II

- CDEX ($DM+0\nu\beta\beta$)
- PandaX ($DM+0\nu\beta\beta$)
- JUNA (Nuclear Astroparticle)
- Jinping Neutrino Program (Solar and Geoneutrino)
- SER (integrated circuit Soft ErroR Research)
- GeoDEX (Deep underground geologic experiment)
- CUPID-CJPL ($0\nu\beta\beta$)
- NvDEx ($0\nu\beta\beta$)
- More coming.....

- Founded in 2009, 11 institutions, more than 100 people now
- Focused on Dark Matter detection and Ge-76 $0\nu\beta\beta$ search using HPGe technology

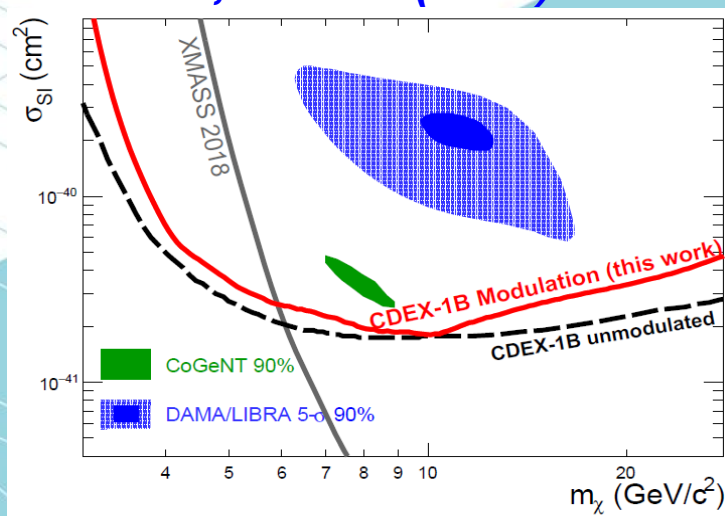


PRL 120, 241301 (2018)



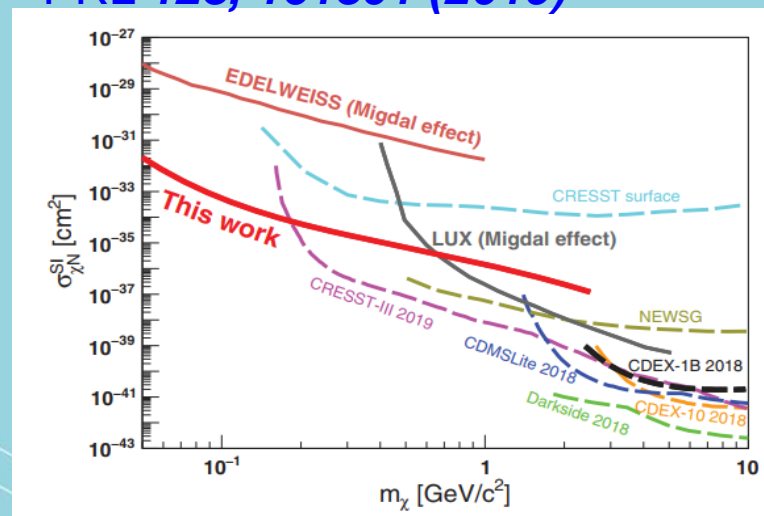
The best SI sensitivity at 4-5GeV energy region from CDEX in 2018

PRL 123, 221301 (2019)

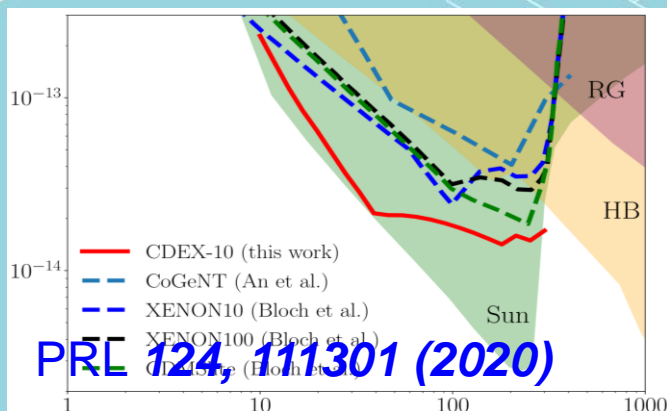


Exclude the possible region favored by CoGeNT and DAMA based on 4yr Data

PRL 123, 161301 (2019)

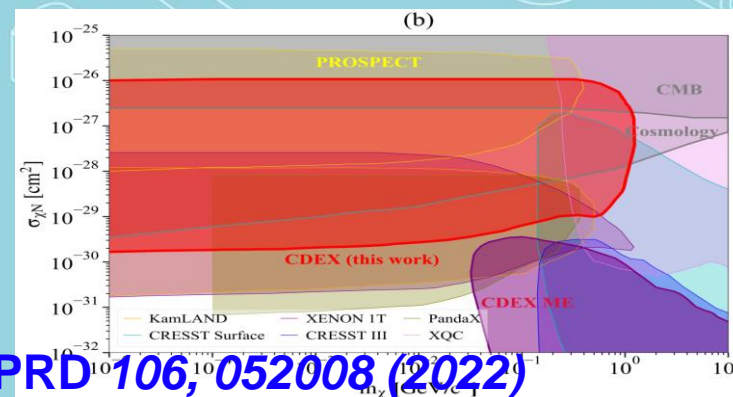


the best results at 50-180 MeV region With Migdal Effect



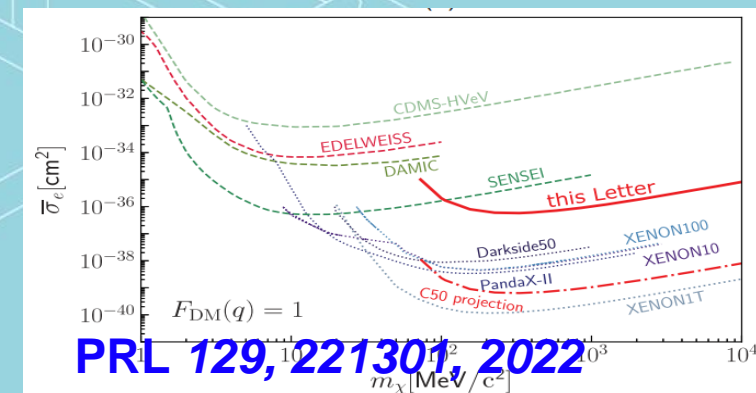
PRL 124, 111301 (2020)

Leading sensitivity in $m_V \sim 10-300$ eV for solar dark photon



PRD 106, 052008 (2022)

Detected CR- accelerated DM and scanned the new light-mass parameter region



PRL 129, 221301, 2022

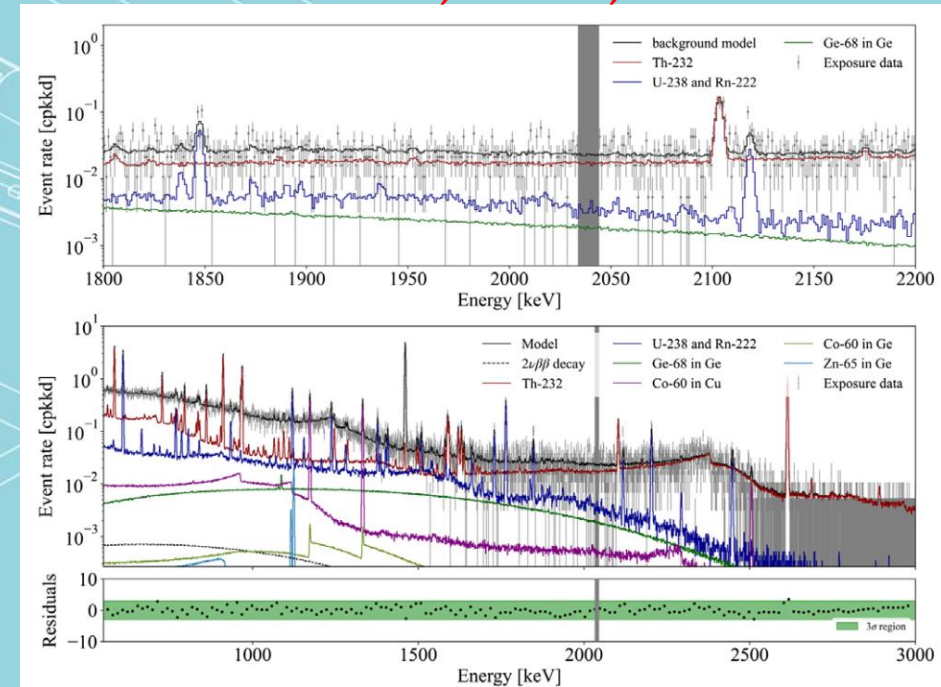
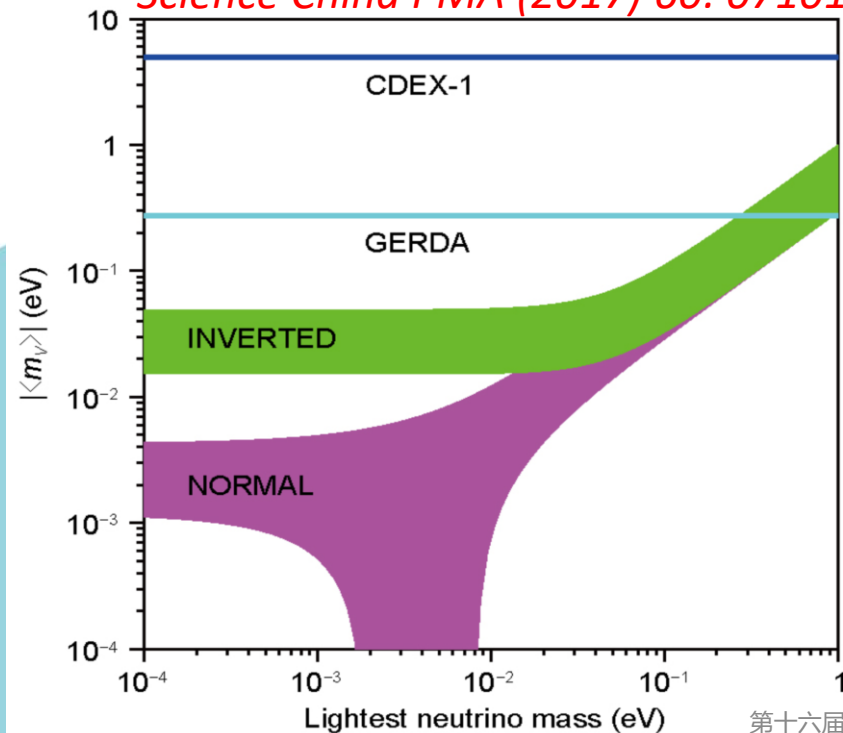
DM-Electron recoil with HPGe

CDEX ^{76}Ge $0\nu\beta\beta$ results

- Based on the CDEX-1 natural PCGe detector, first ^{76}Ge $0\nu\beta\beta$ result in 2017.
- $\langle m_{\beta\beta} \rangle$ sensitivity: one order less than GERDA-I at that time.
- Develop new BEGe detector for CDEX- $0\nu\beta\beta$ experiment.

Science China PMA (2017) 60: 071011

PRD 106, 032012, 2022



PandaX

- Starting in 2009, consists of dozens of universities and research Institutions
- Goals: Increase LXe detector mass for DM and neutrino studies

Jianglai Liu's Talk

PandaX



2009

Phase-I
120 kg



2010-2014

Phase-II
580 kg



2015-2019

PandaX-4T
3.7 ton



2020-

PandaX

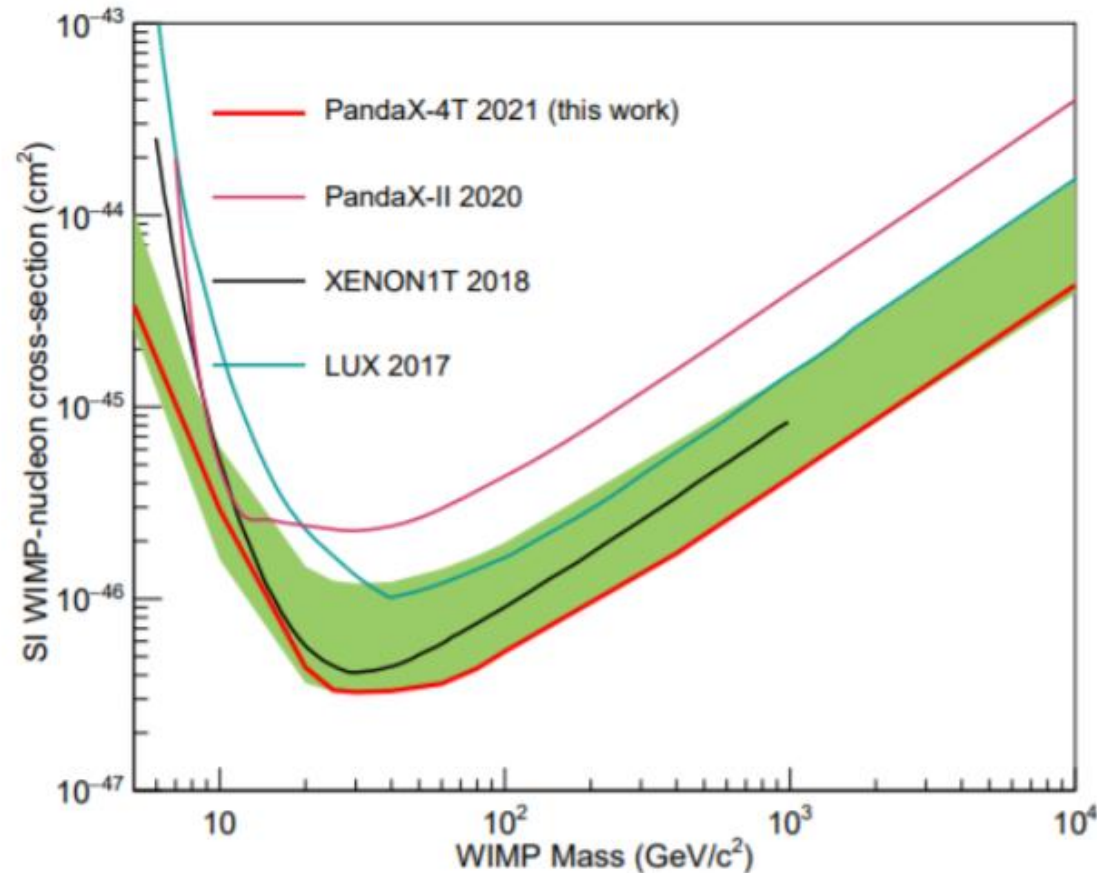
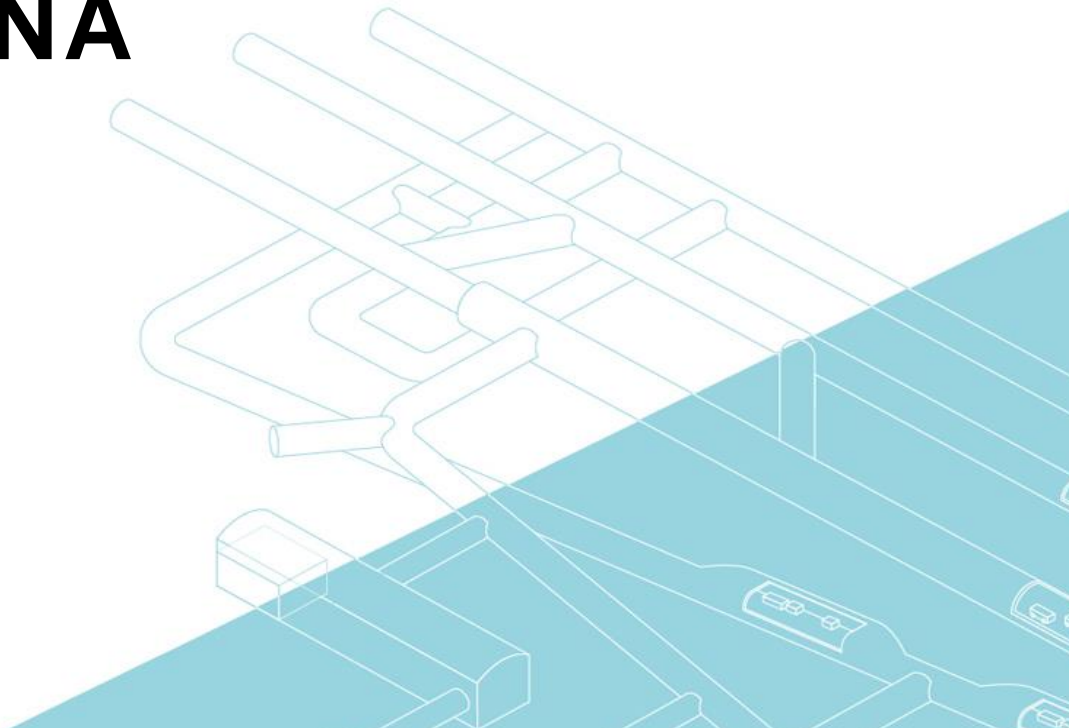


Fig. 9. The 90% C.L. upper limit vs. m_χ for the SI WIMP-nucleon elastic cross section from PandaX-4T commissioning data (red), overlaid with that from LUX 2017,⁴ XENON1T 2018³ and PandaX-II 2020.⁵ The green band represents the $\pm 1\sigma$ sensitivity band.

JUNA

- Starting in 2013, 7 universities and research Institutions
- Goals: Nuclear astrophysics study using underground accelerator





PHYSICAL REVIEW LETTERS 127, 152702 (2021)

Editors' Suggestion

Featured in Physics

Direct Measurement of the Astrophysical $^{19}\text{F}(p, \alpha\gamma)^{16}\text{O}$ Reaction in the Deepest Operational Underground Laboratory

L. Y. Zhang,¹ J. Su,¹ J. J. He^{1,*}, M. Wiescher,^{2,†} R. J. deBoer,² D. Kahl,³ Y. J. Chen,¹ X. Y. Li,¹ J. G. Wang,⁴ L. Zhang,⁵ F. Q. Cao,⁵ H. Zhang,⁵ Z. C. Zhang,⁶ T. Y. Jiao,⁴ Y. D. Sheng,¹ L. H. Wang,¹ L. Y. Song,¹ X. Z. Jiang,¹ Z. M. Li,¹ E. T. Li,⁶ S. Wang,⁷ G. Lian,⁵ Z. H. Li,⁵ X. D. Tang,⁴ H. W. Zhao,⁴ L. T. Sun,⁴ Q. Wu,⁴ J. Q. Li,⁴ B. Q. Cui,⁵ L. H. Chen,⁵ R. G. Ma,⁵ B. Guo,⁵ S. W. Xu,⁴ J. Y. Li,⁴ N. C. Qi,⁸ W. L. Sun,⁸ X. Y. Guo,⁸ P. Zhang,⁸ Y. H. Chen,⁸ Y. Zhou,⁸ J. F. Zhou,⁸ J. R. He,⁸ C. S. Shang,⁸ M. C. Li,⁸ X. H. Zhou,⁴ Y. H. Zhang,⁴ F. S. Zhang,¹ Z. G. Hu,⁴ H. S. Xu,⁴ J. P. Chen,^{1,g} and W. P. Liu^{5,‡}

Article

Measurement of $^{19}\text{F}(p, \gamma)^{20}\text{Ne}$ reaction suggests CNO breakout in first stars

<https://doi.org/10.1038/s41586-022-05230-x>

Received: 28 February 2022

Accepted: 11 August 2022

Published online: 26 October 2022

Check for updates

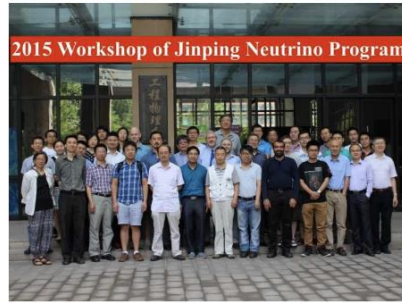
Liyong Zhang¹, Jianjun He^{1,‡}, Richard J. deBoer², Michael Wiescher^{2,‡}, Alexander Heger³, Daid Kahl⁴, Jun Su¹, Daniel Odell⁵, Yinji Chen¹, Xinyue Li¹, Jianguo Wang⁶, Long Zhang⁷, Fuqiang Cao⁷, Hao Zhang¹, Zhicheng Zhang⁸, Xinzhi Jiang¹, Luohuan Wang¹, Ziming Li¹, Luyang Song¹, Hongwei Zhao⁶, Liangting Sun⁶, Qi Wu⁶, Jiaqing Li⁶, Baoqun Cui⁷, Lihua Chen⁷, Ruigang Ma⁷, Ertao Li⁸, Gang Lian⁷, Yaode Sheng¹, Zhihong Li⁷, Bing Guo⁷, Xiaohong Zhou⁶, Yuhu Zhang⁶, Hushan Xu⁶, Jianping Cheng¹ & Weiping Liu^{7,‡}

Proposed mechanisms for the production of calcium in the first stars (population III stars)—primordial stars that formed out of the matter of the Big Bang—are at odds with observations¹. Advanced nuclear burning and supernovae were thought to be the dominant source of the calcium production seen in all stars². Here we suggest a qualitatively different path to calcium production through breakout from the ‘warm’ carbon–nitrogen–oxygen (CNO) cycle through a direct experimental measurement of the $^{19}\text{F}(p, \gamma)^{20}\text{Ne}$ breakout reaction down to a very low energy point of 186 kiloelectronvolts, reporting a key resonance at 225 kiloelectronvolts. In the domain of astrophysical interest³, at around 0.1 gigaelectronvolts, this thermonuclear $^{19}\text{F}(p, \gamma)^{20}\text{Ne}$ rate is up to a factor of 7.4 larger than the previous recommended rate³. Our stellar models show a stronger breakout during stellar hydrogen burning than previously thought^{4,5}, and may reveal the nature of calcium production in population III stars imprinted on the oldest known ultra-iron-poor star, SMSS0313-6708⁶. Our experimental result was obtained in the China Jinping Underground Laboratory⁷, which offers an environment with an extremely low cosmic-ray-induced background⁸. Our rate showcases the effect that faint population III star supernovae can have on the nucleosynthesis observed in the oldest known stars and first galaxies, which are key mission targets of the James Webb Space Telescope⁹.

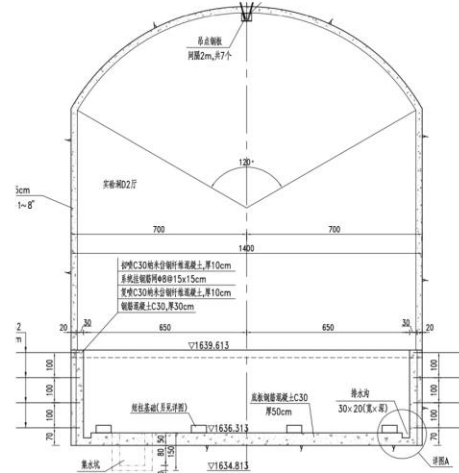
Jinping Neutrino Program

Goal:

- Solar neutrino detection
- Geoneutrino detection
- Supernova neutrino search



Pit for 500 ton Neutrino Observatory

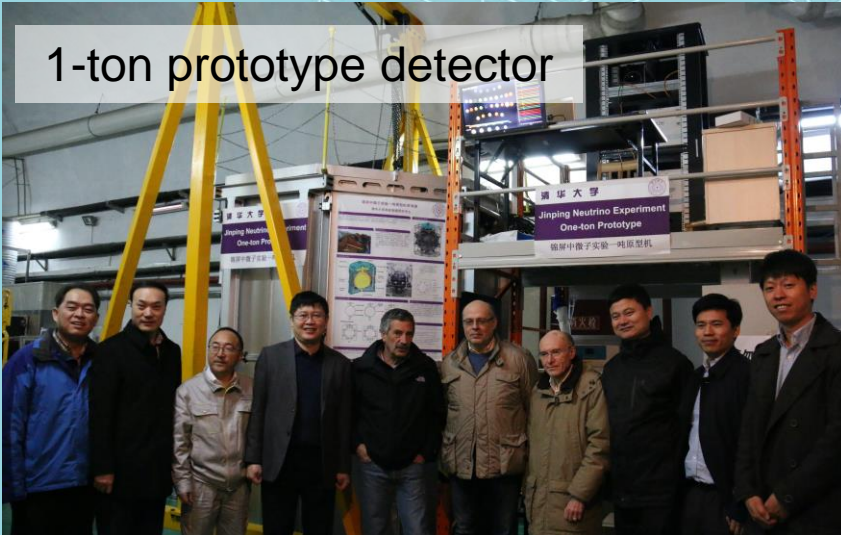


实验设备基坑设计图（华东设计院）

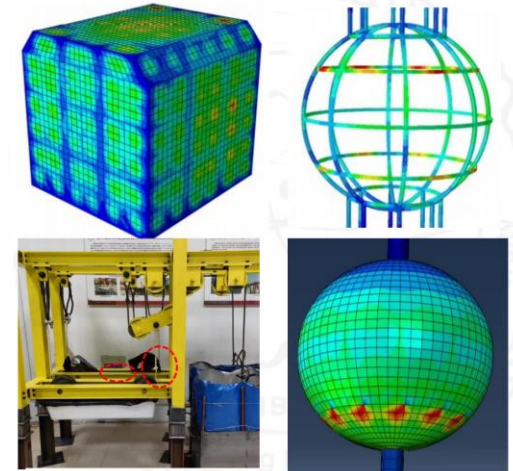
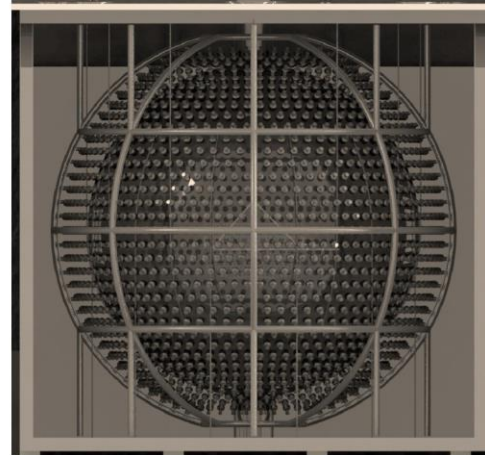


实验设备基坑，长16米，宽14米，深3.8米，开挖工作完成（水电5局）

1-ton prototype detector

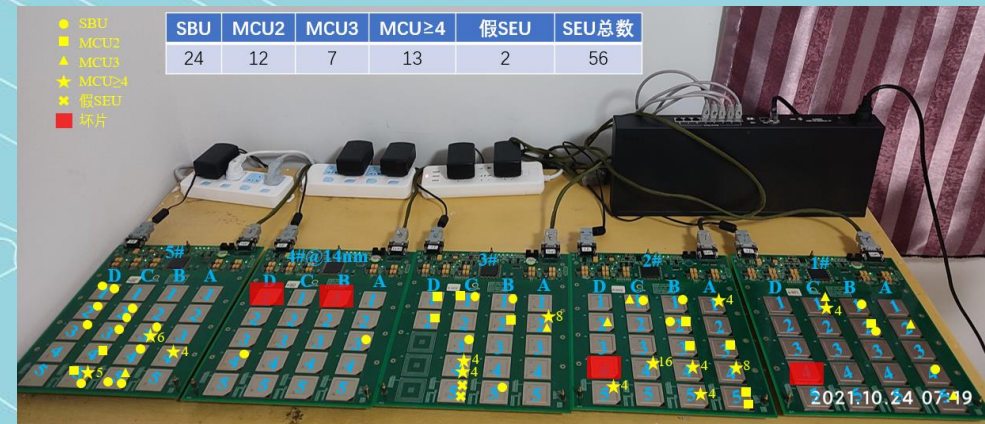
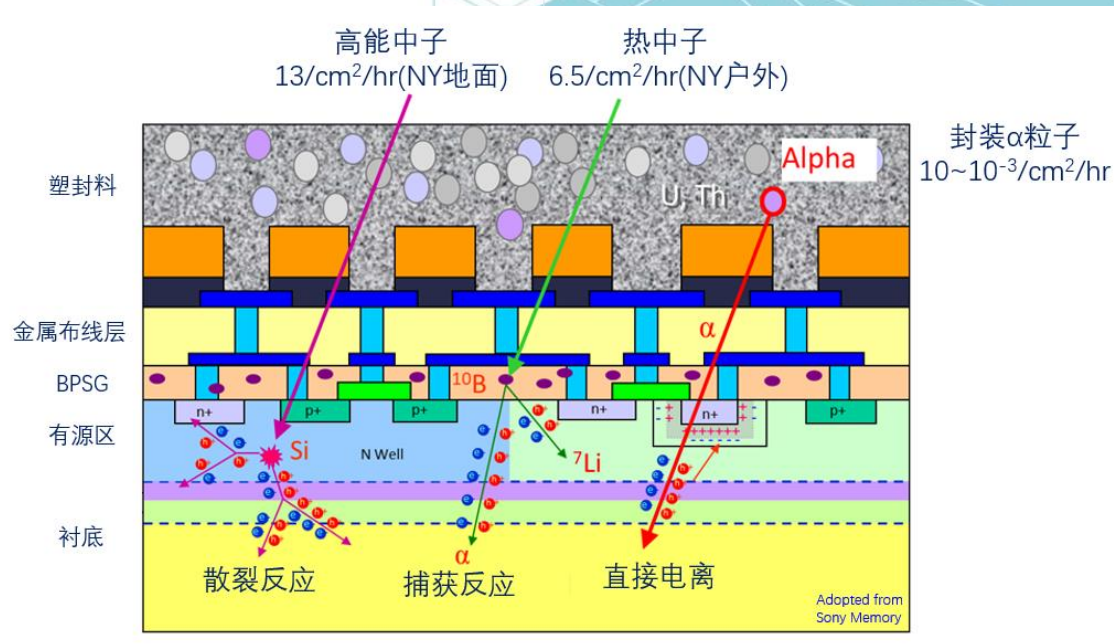


Neutrino detector design



SER: deep underground integrated circuit Soft Error Research

- Radioisotopes in IC could cause soft error (SE) by emitting alpha particles
- Study the SE rate in CJPL to prevent interference from atmosphere neutrons
- Compare test data in CJPL and Lasa to acquire “Golden data” for IC SE rate



Test at Lasa

- 6651 hr

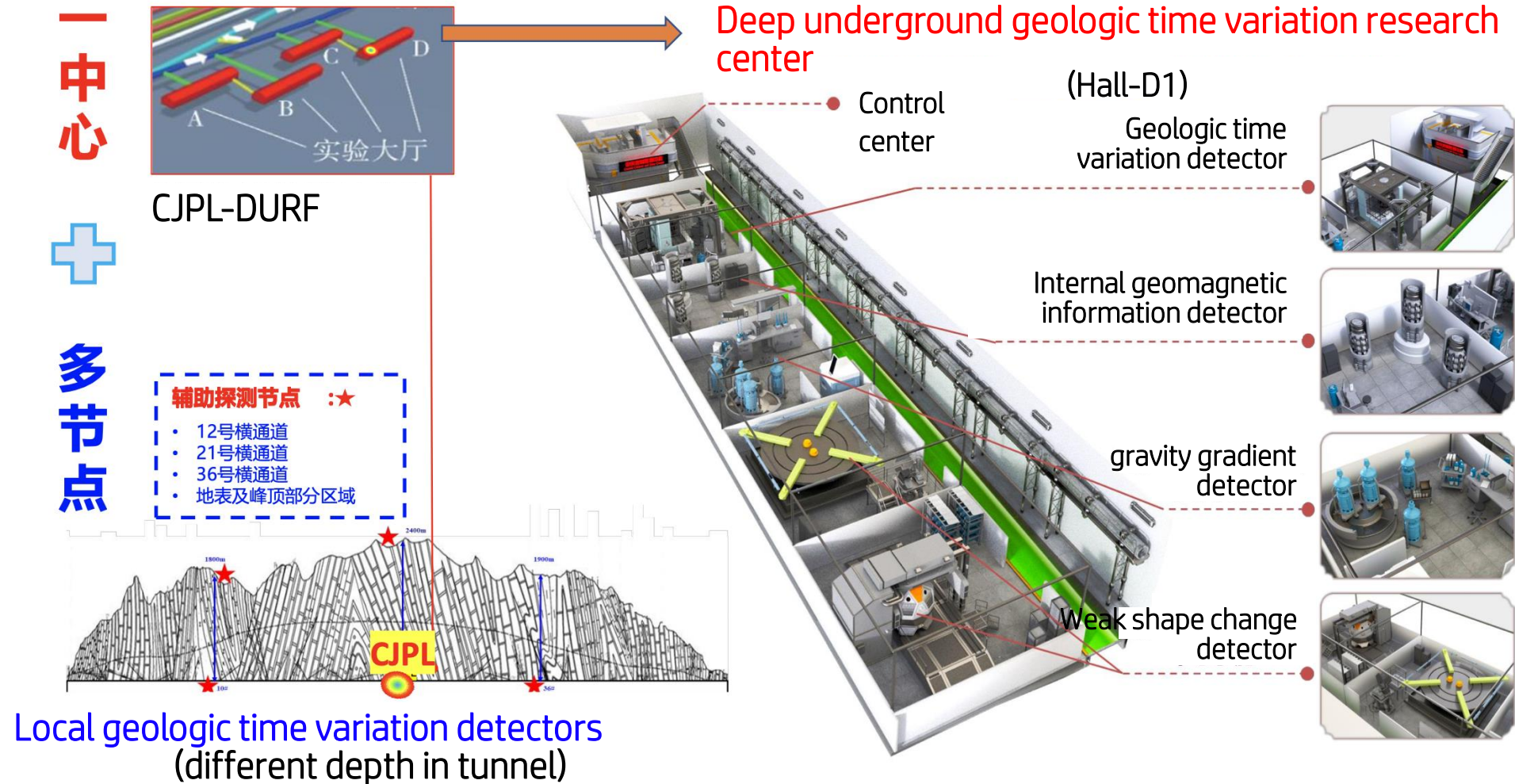


Test at CJPL

- 6000 hr

GeoDEX

- Deep underground geologic time variation in-situ detector experiment



CUPID-CJPL

- $0\nu\beta\beta$ experiment based on Mo-100 isotope and LMO crystal calorimeter array.

Participate in CUPID project:
crystal validation and electronics

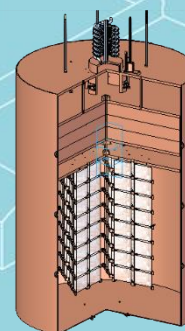
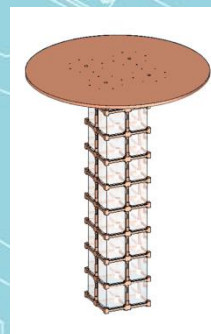
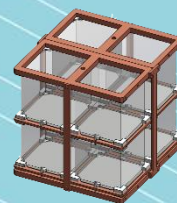
Crystal Testing
(2023-2028)

Develop CUPID-reach/1T Technology

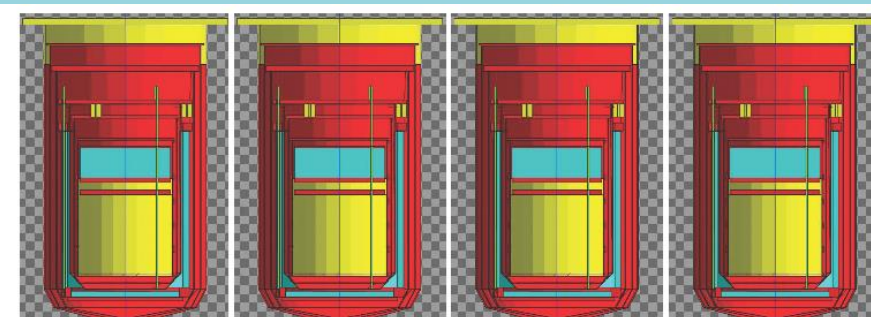
CUPID-CJPL Demo
(2026-2028)
10 kg enriched crystals

Build CUPID-reach or CUPID-1T at CJPL

CUPID-CJPL-200/1T
(2028+)
> 250 kg enriched crystals



Possible CUPID-1T configuration



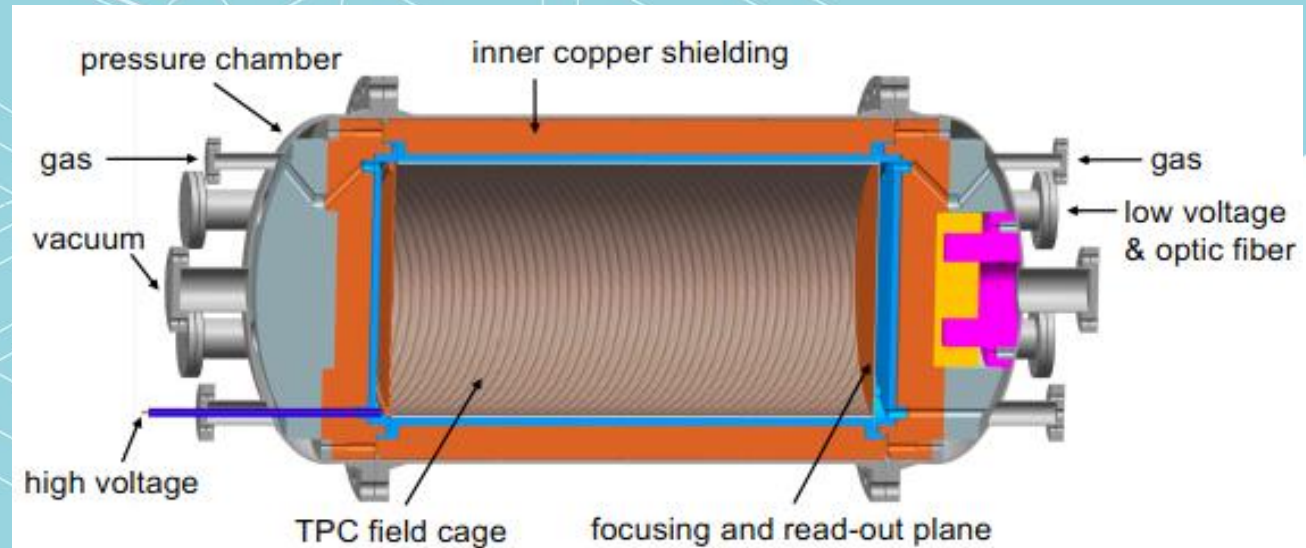
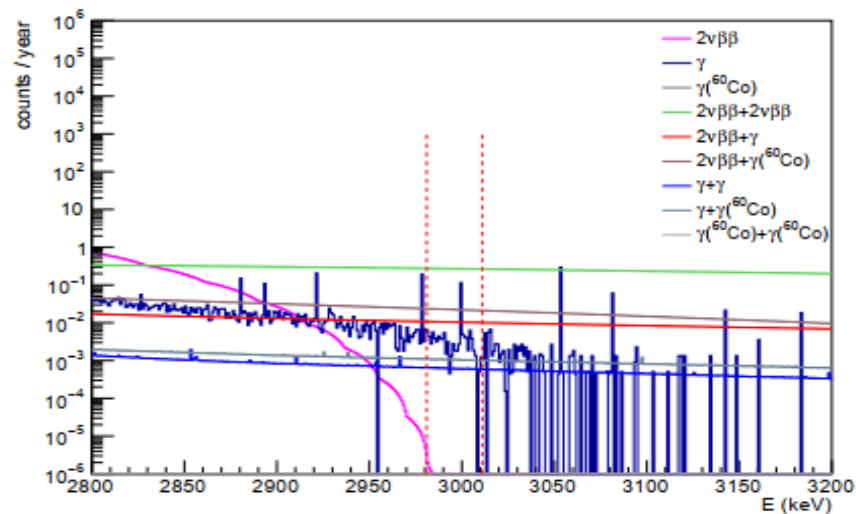
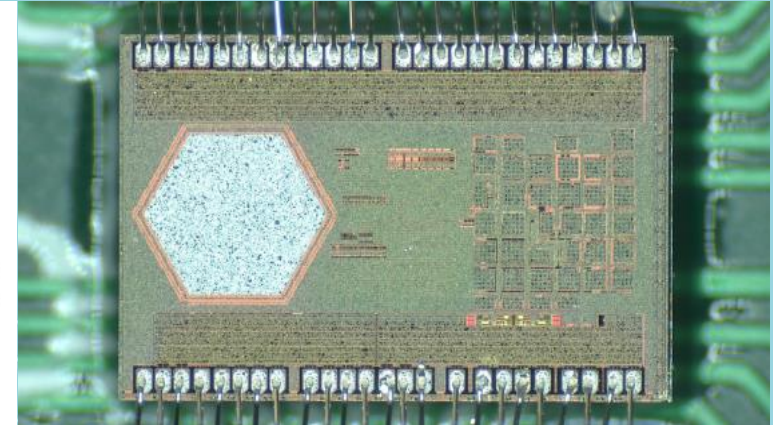
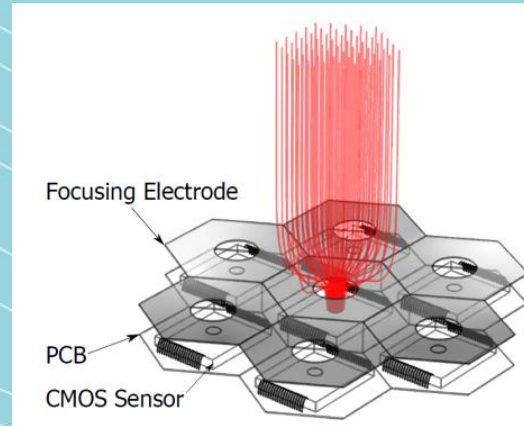
4x250 kg of ^{100}Mo

A “flock” of CUPIDs around the world

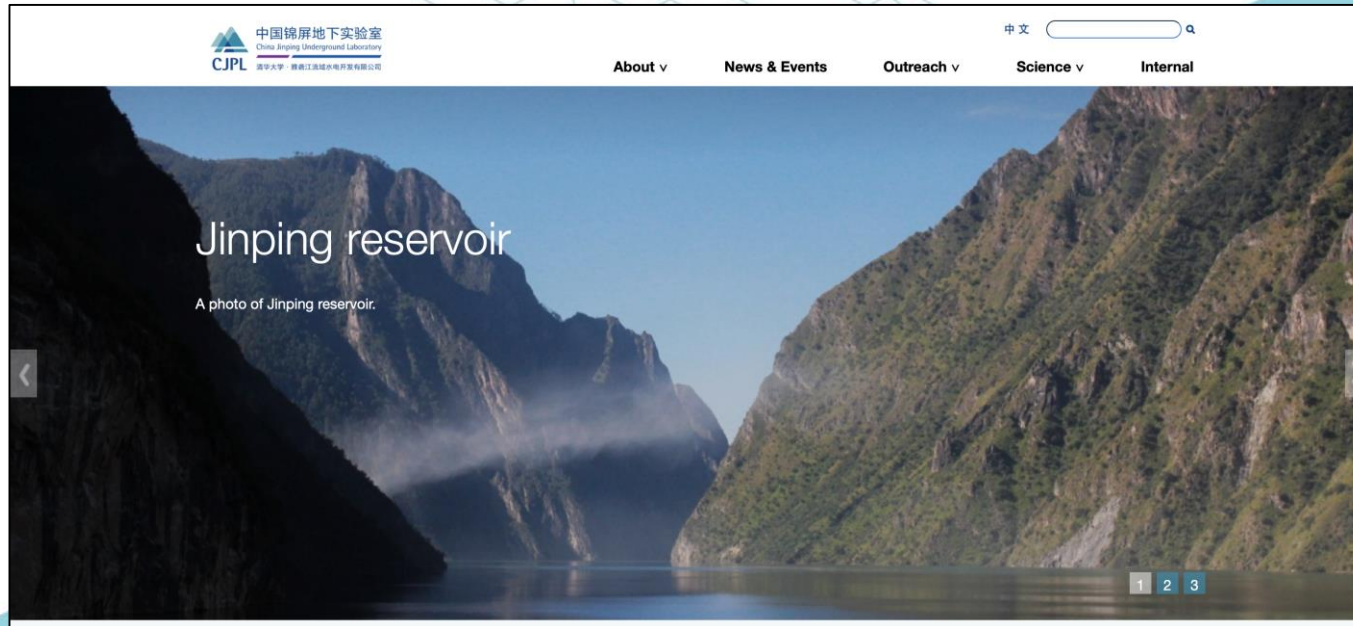
- 10 kg experiment
 - low enrichment
 - demonstrate key technologies
- 250 kg+ experiment
 - high enrichment
 - probe $m_{\beta\beta}$ below 10meV

NvDEX

- $0\nu\beta\beta$ experiment based on Se-82 isotope and High pressure TPC.
- Topmetal technique to readout the ion signal.
- High pressure $^{82}\text{SeF}_6$ gas as the source and target.
- 200kg per module * 5



Welcome New Experiments and Collaborations

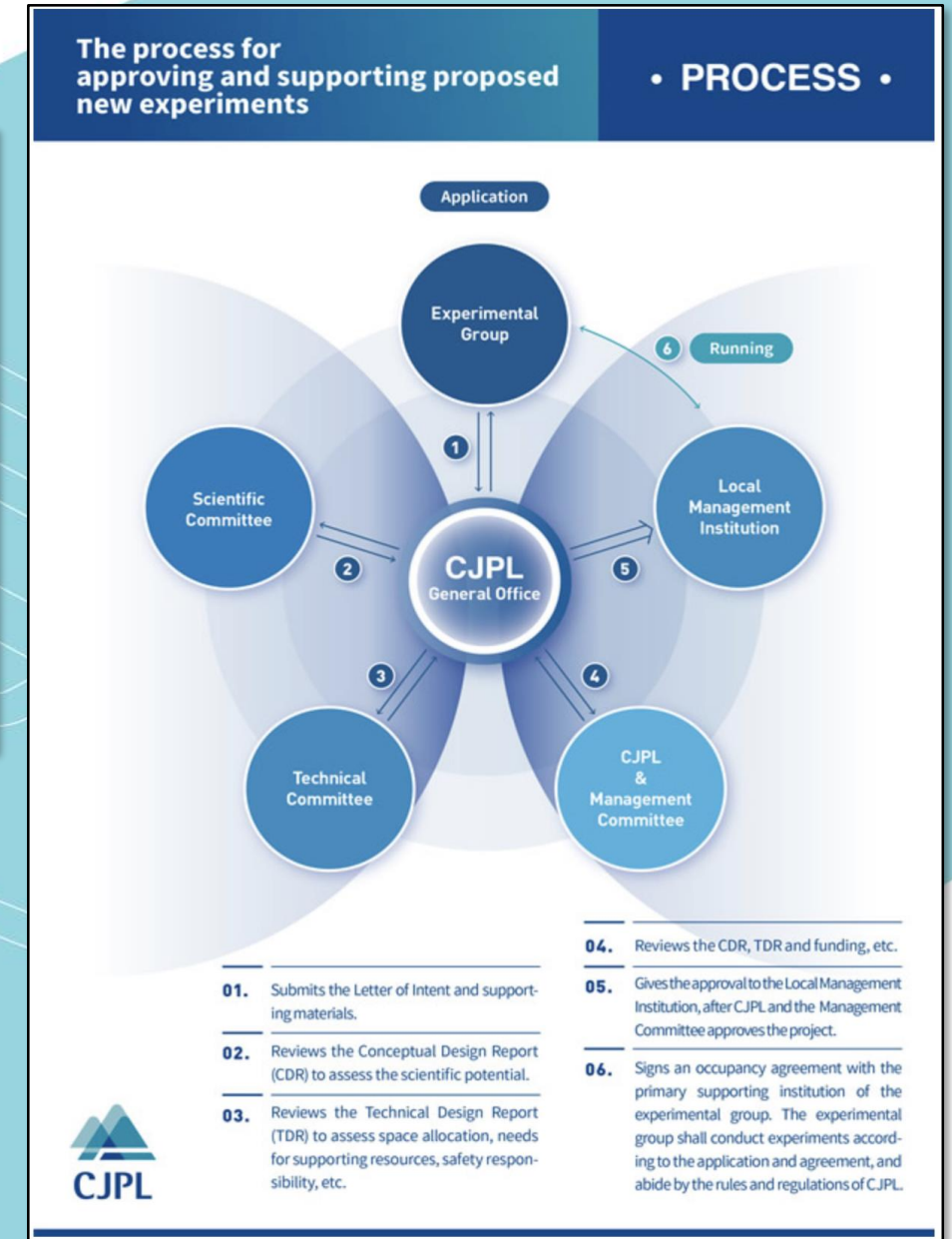


Visit our web for more info:

<https://cjpl.tsinghua.edu.cn/>

Contact:

cjpl@tsinghua.edu.cn



Summary:

- Series of important physical results on DM, nuclear astroparticle physics has been achieved based on CJPL deep underground platform.
- CJPL-II will be the deepest (2400m rock) and largest ($>300,000 \text{ m}^3$ space) underground Lab worldwide;
- CJPL-II plan to start operation in 2024.
- More collaborations have planned to start their experiments in CJPL-II.

CJPL welcomes more projects worldwide to apply for using UL space !



CJPL
中国锦屏地下实验室
China Jinping Underground Laboratory

Welcome!