

TAUP 2023  
Vienna, 29 August 2023

# Neutrinoless double-beta decay search with SNO+

*SNO+ Background Performance  
& Preparations for  $0\nu\beta\beta$*

Valentina Lozza for the SNO+ Collaboration

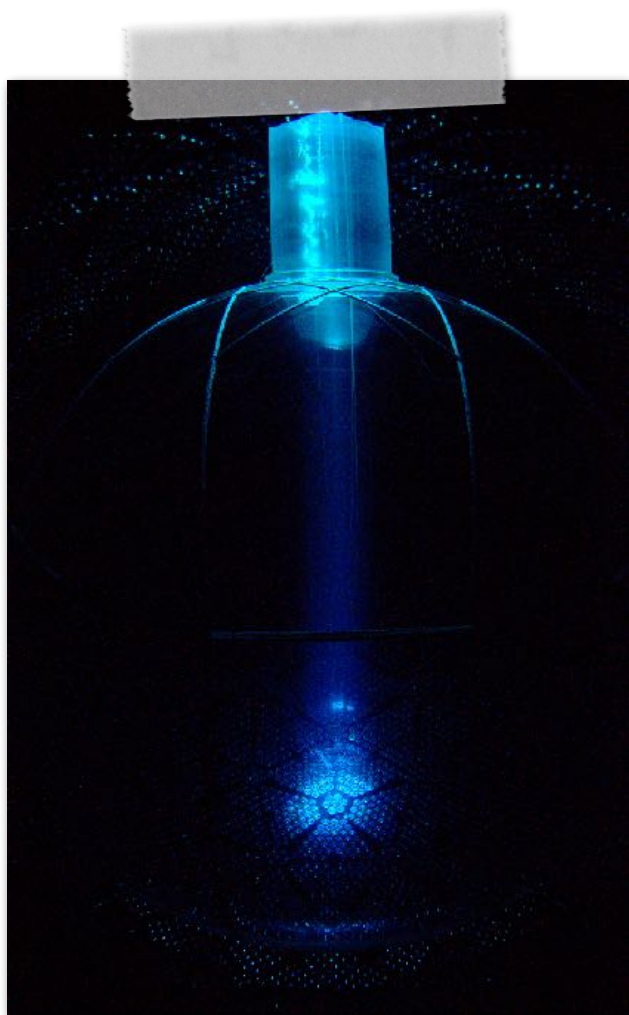
**FCT** Fundação  
para a Ciência  
e a Tecnologia



**LIP** Lisbon



- Multi-purpose neutrino experiment with the primary goal to search for the neutrinoless double-beta decay of  $^{130}\text{Te}$ .



*JINST 16 (2021) 08, P08059*



Acrylic Vessel (AV)  
12 m diam., 5 cm thick

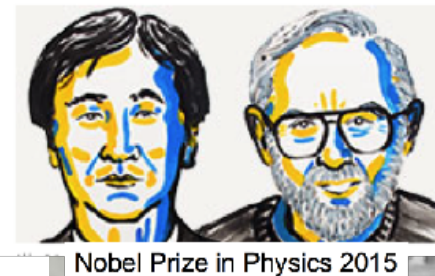
Support Structure  
with ~9400 PMTs

Light water (H<sub>2</sub>O) shielding

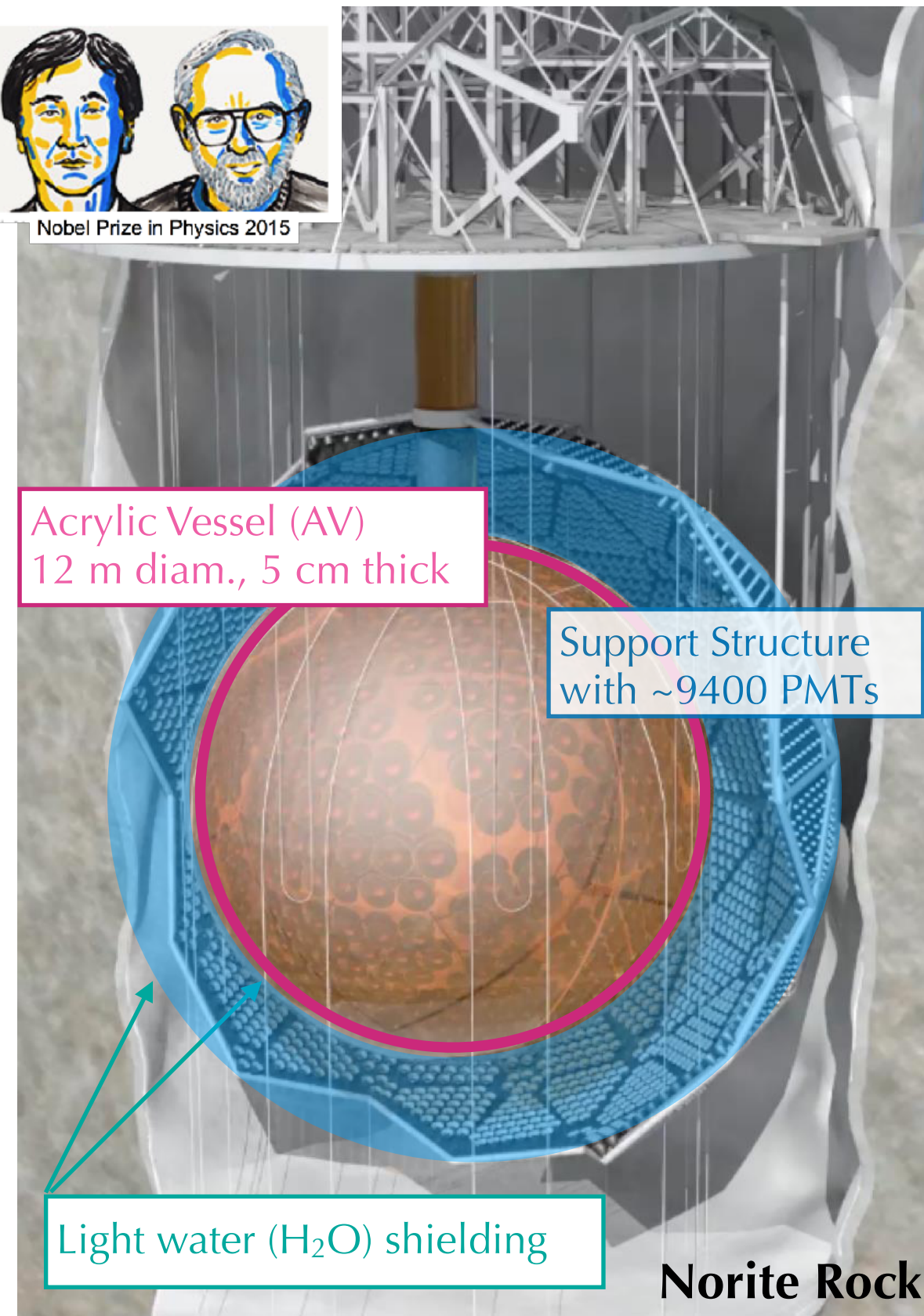
**Norite Rock**



- ◎ Multi-purpose neutrino experiment with the primary goal to search for the neutrinoless double-beta decay of  $^{130}\text{Te}$ .
- ◎ Three target's materials:
  - ▶ 905 t of *ultra-pure water* (2017-2019);
  - ▶ ~780 t of *high purity liquid scintillator* + 2.2 g/L PPO +bisMSB (2022-2024);
  - ▶ 3.9 t  $^{nat}\text{Te}$ -loaded scintillator (2024 - ).

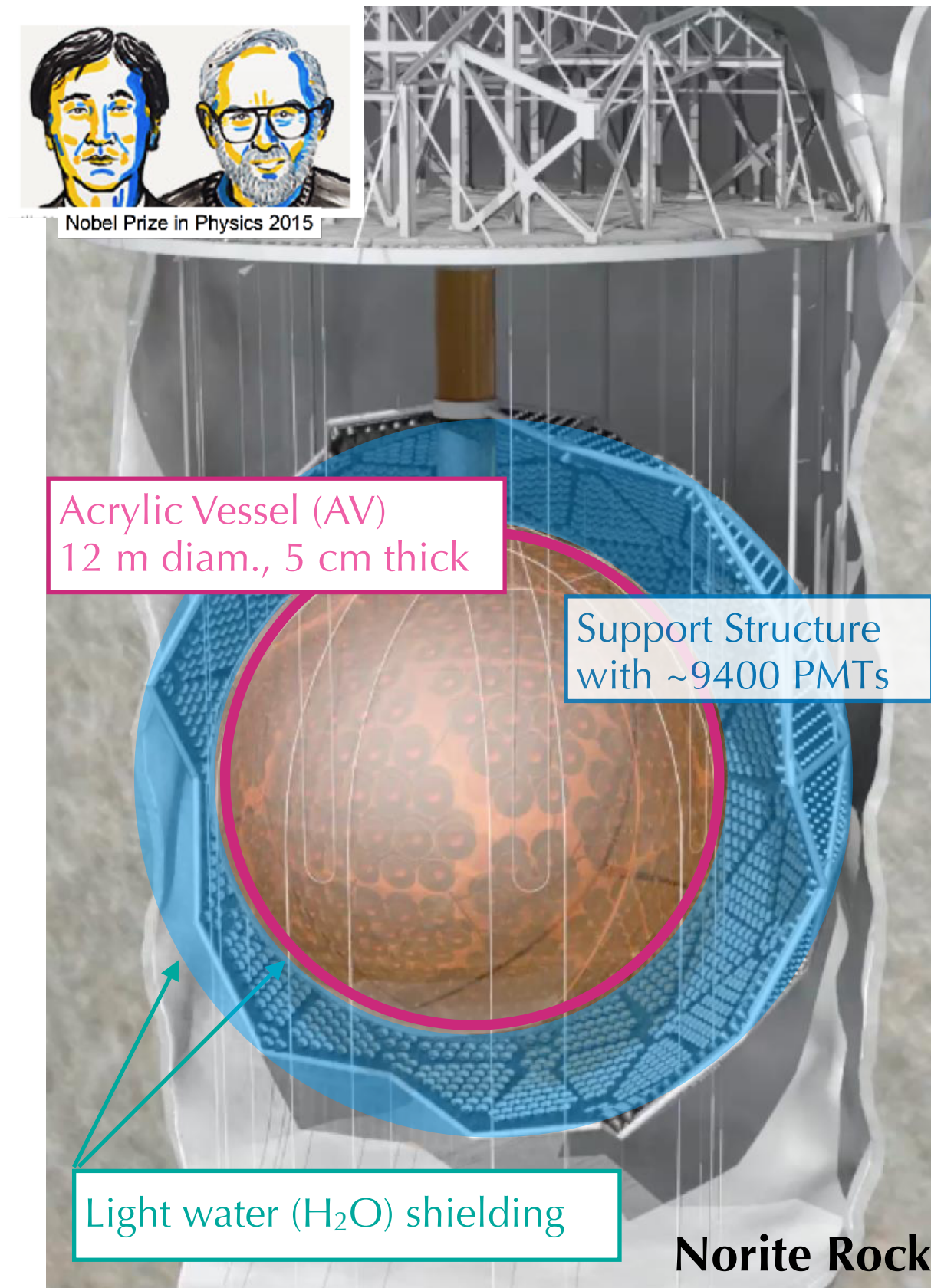


Nobel Prize in Physics 2015





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  - ▶ 3.9 t  *$^{nat}\text{Te}$ -loaded scintillator* (2024 - ).
- ◎ Several background reduction layers:
  - ▶ 7 kt of high-purity water shield;
  - ▶  $\text{N}_2$  Cover Gas blanket across the entire detector;
  - ▶ Radon-impermeable plastic covering the cavity walls.

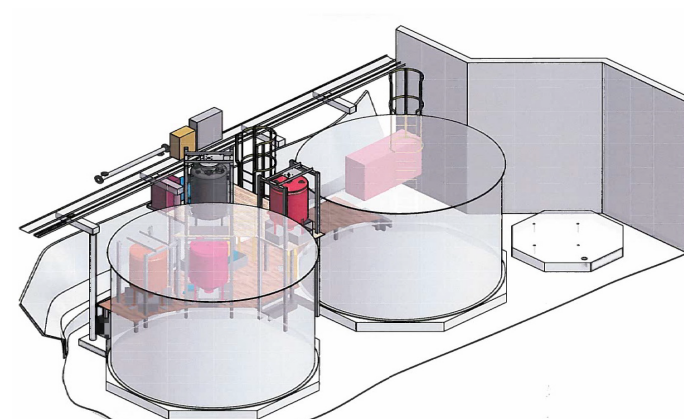




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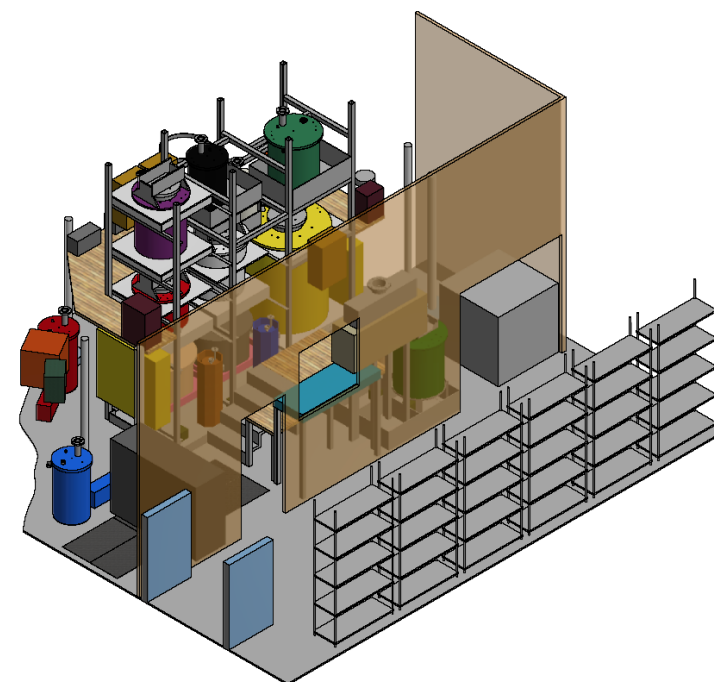
**TeDiol purification plant**



**LS purification plant**

- Material purification:

- ▶ 4 purification plants UG for water, scintillator and Te;
- ▶ Possibility to recirculate and repurify water and scintillator;
- ▶ Extensive QA campaigns before, during and after filling/loading.



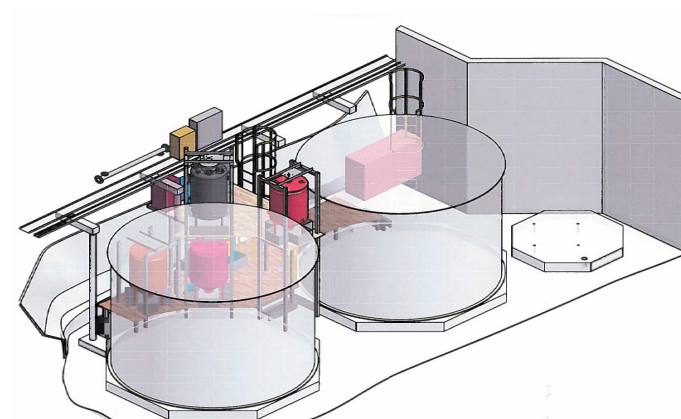
**TeA purification plant**



- Multi-purpose neutrino experiment with the primary goal to search for the neutrinoless double-beta decay of  $^{130}\text{Te}$ .

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TeDiol purification plant



LS purification plant

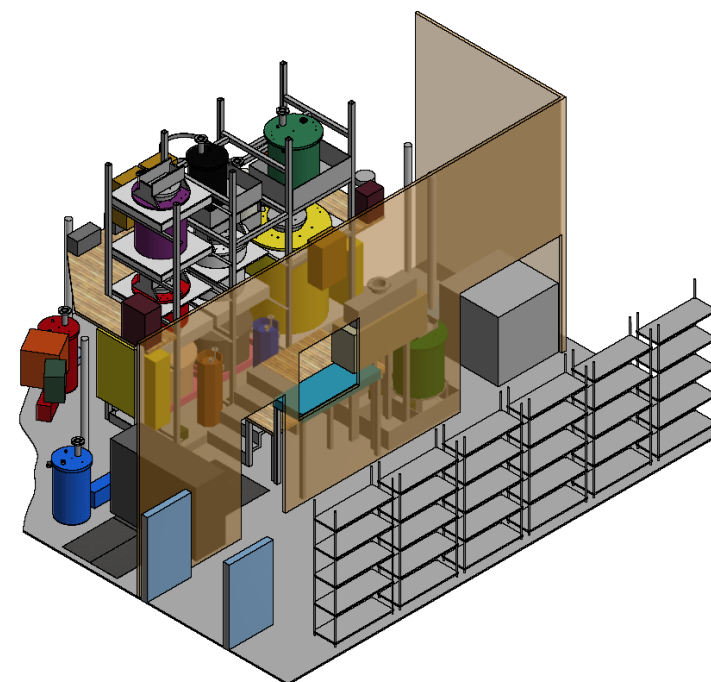
- Material

See poster

S. Manecki, S. Biller

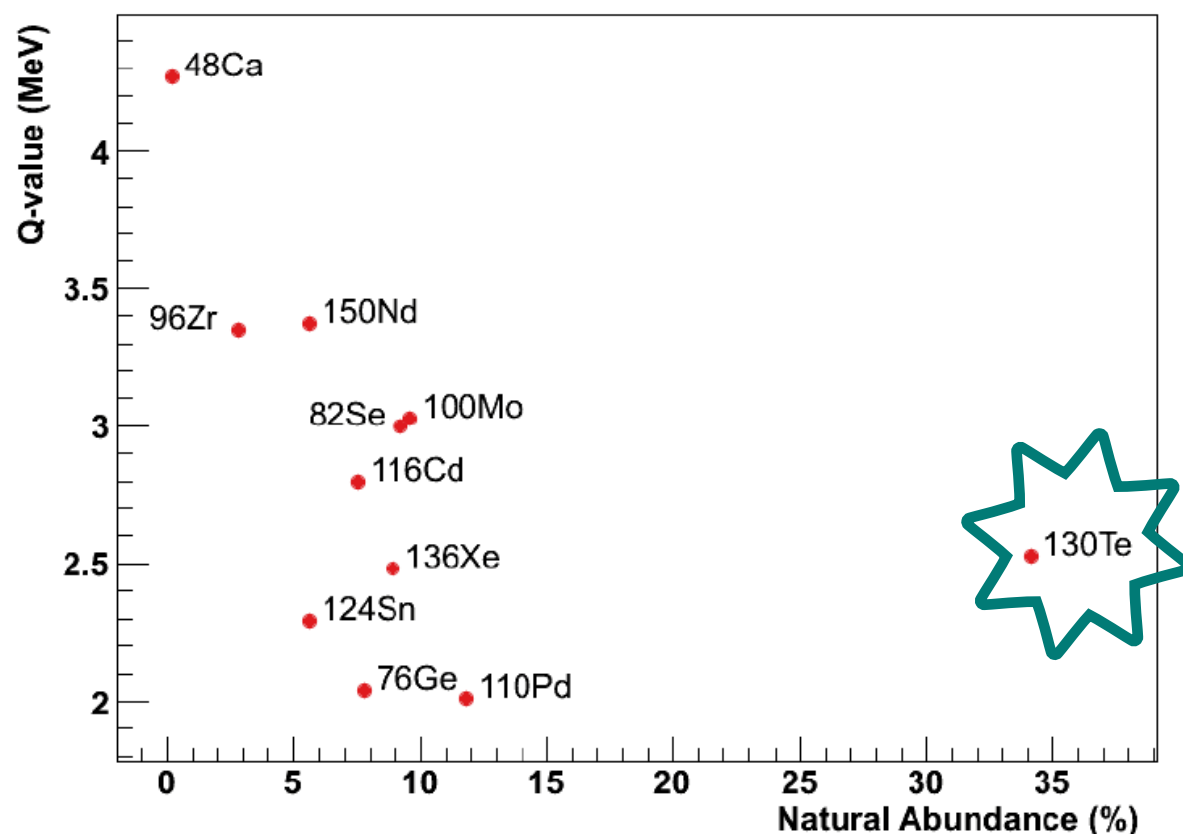
***SNO+ Tellurium Purification and Loading for Neutrinoless Double Beta Decay Search***

- Possibility to recirculate and repurify water and scintillator;
- Extensive QA campaigns before, during and after filling/loading.



TeA purification plant





## Major Advantages of Te

- ❖ No need for enrichment
- ❖ Long  $2\nu\beta\beta$  half-life ( $7.9 \times 10^{20}$  yrs)
- ❖ High Q-value of 2.5 MeV

## Major Advantages of SNO+

- ❖ Large size allows rejection of external backgrounds
- ❖ Fast timing allows rejection of U and Th chain background (+ alpha,n)
- ❖ High light yield for good resolution = targeting 460 PMT hits/MeV
- ❖ *Target-out measurements before and while adding Te*





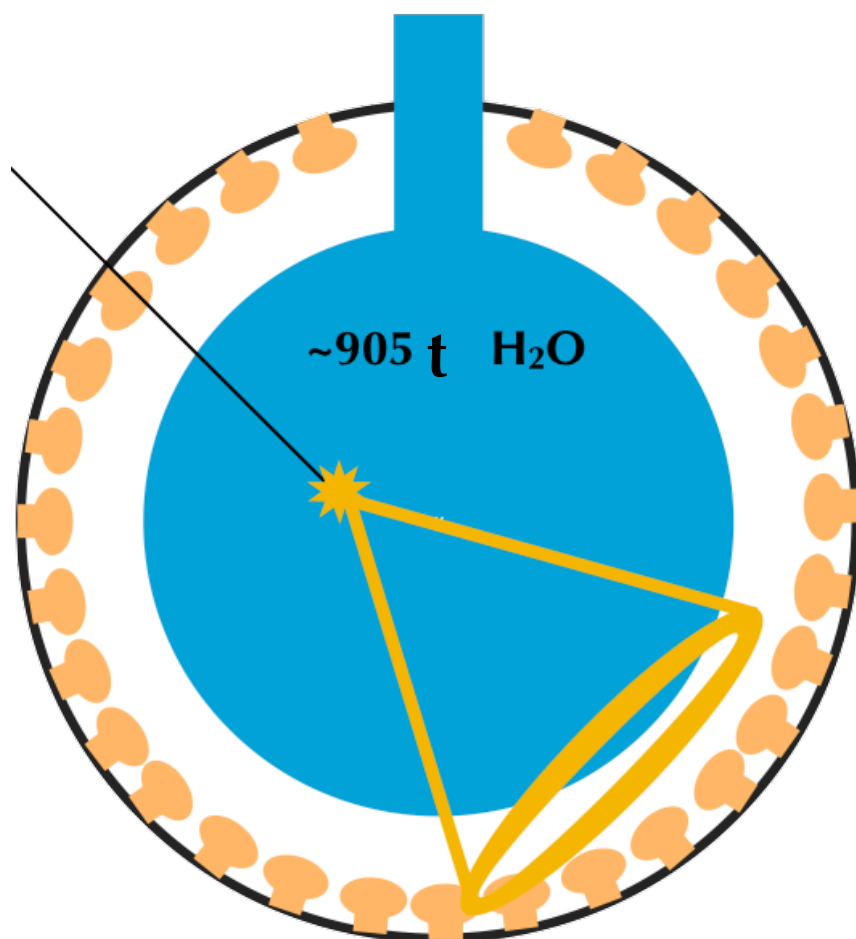
May - December 2017  
(~115 gold physics days)  
First SNO+ water phase  
*Phys.Rev.D 99, 032008 (ND)*  
*Phys.Rev.D 99, 012012 (Solar)*  
*Phys.Rev.C 102, 014002 (Antinu)*

October 2018 - June 2019  
(~185 gold physics days)  
Second SNO+ water phase  
*Phys.Rev.D 105, 112012 (ND)*

Additional Cover gas  
shielding to reduce Rn  
ingresses in water

## Major Outcomes

- ❖ Nucleon decay modes into invisible channels
- ❖ Solar neutrinos
- ❖ Reactor antineutrinos
  - ❖ **First measurement** of reactor antineutrinos using pure water *Phys.Rev.Lett 130, 091801*



See talk  
T. Kaptanoglu  
Aug 31, 2023, 3:15PM  
Hörsaal 21 lecture hall



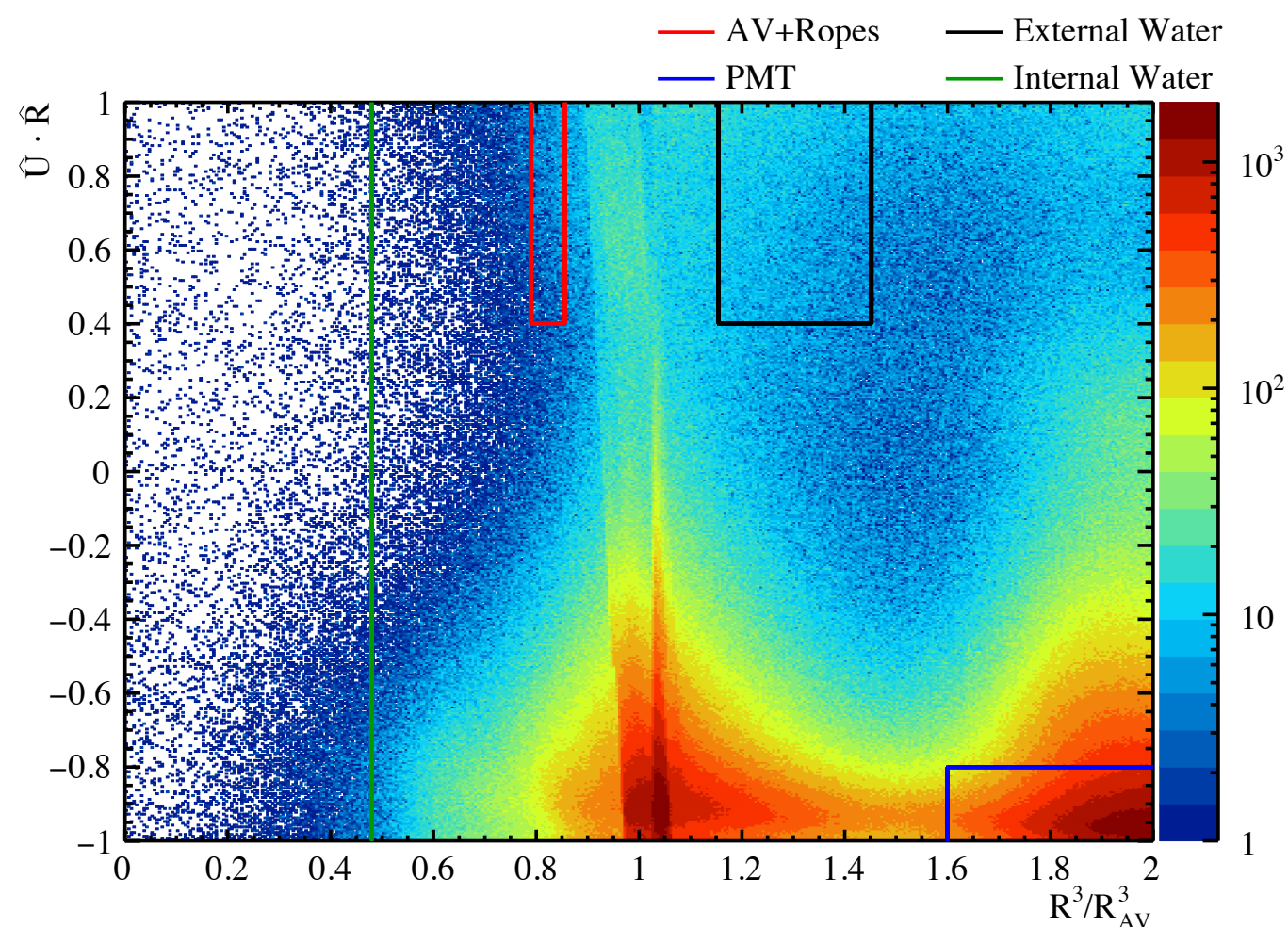
# SNO+ EXTERNAL BACKGROUNDS

❖ External background measurement during the water phase allows to use a directional cut

- $3.0 < E < 5.0$
- $-5.0 \text{ m} < Z < 5.0 \text{ m}$

Box	Cuts
AV	$5.55 \text{ m} < R_{AV} < 5.7 \text{ m}$ $U \cdot R_{AV} > 0.4$
External Water	$6.3 \text{ m} < R < 6.8 \text{ m}$ $U \cdot R > 0.4$
PMT	$1.6 < R^3 < 2.0$ $U \cdot R < -0.8$
Internal Water	$R_{AV} < 4.7 \text{ m}$

Background	Rate (Fraction of Nominal)
AV+Ropes	$0.21 \pm 0.009^{+0.64}_{-0.21}$
External Water	$0.44 \pm 0.003^{+0.32}_{-0.27}$
PMT	$1.48 \pm 0.002^{+1.65}_{-0.60}$



Contribution to the  $0\nu\beta\beta$  ROI is 50% smaller than with nominal values



# SNO+ SCINTILLATOR



2017

2018

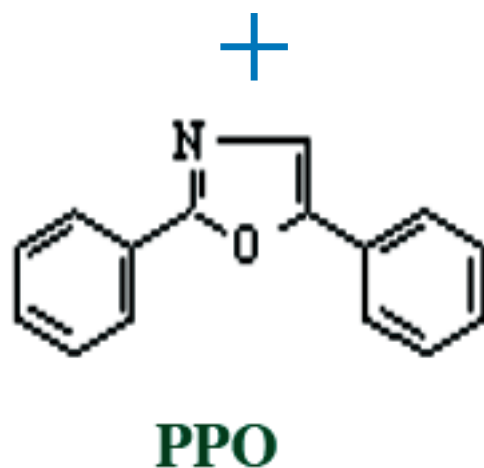
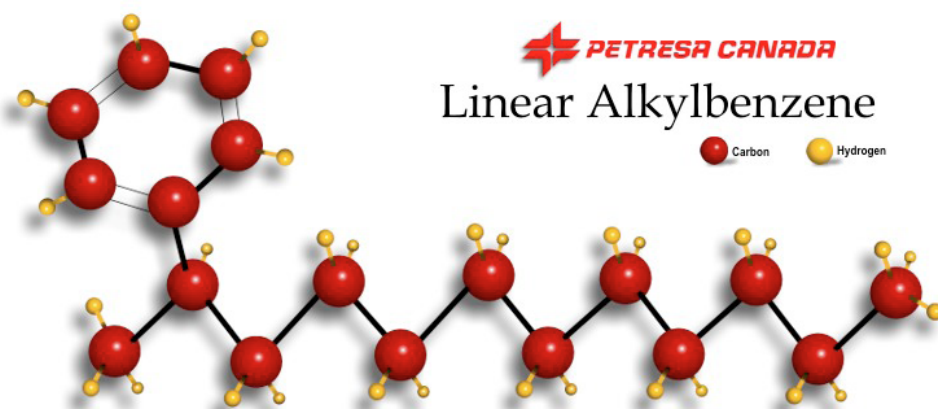
2019

2020

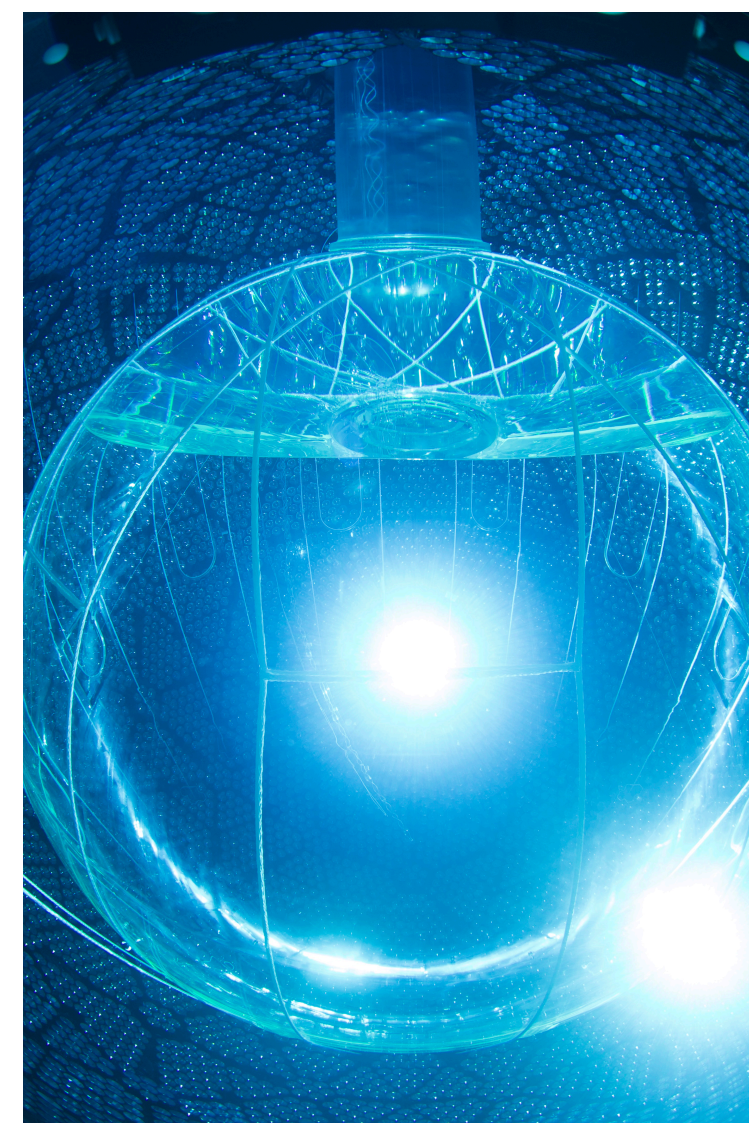
2021

↑  
Filling the detector  
with liquid scintillator  
*JINST 16 P05009*

↑  
April - October 2020  
(92 gold physics days)  
Bonus phase: half-filled  
detector with scintillator  
0.6 g/L PPO



- ❖ Developed by SNO+
- ❖ From water Cherenkov detector to scintillator detector!
- ❖ Allows isotopic loading





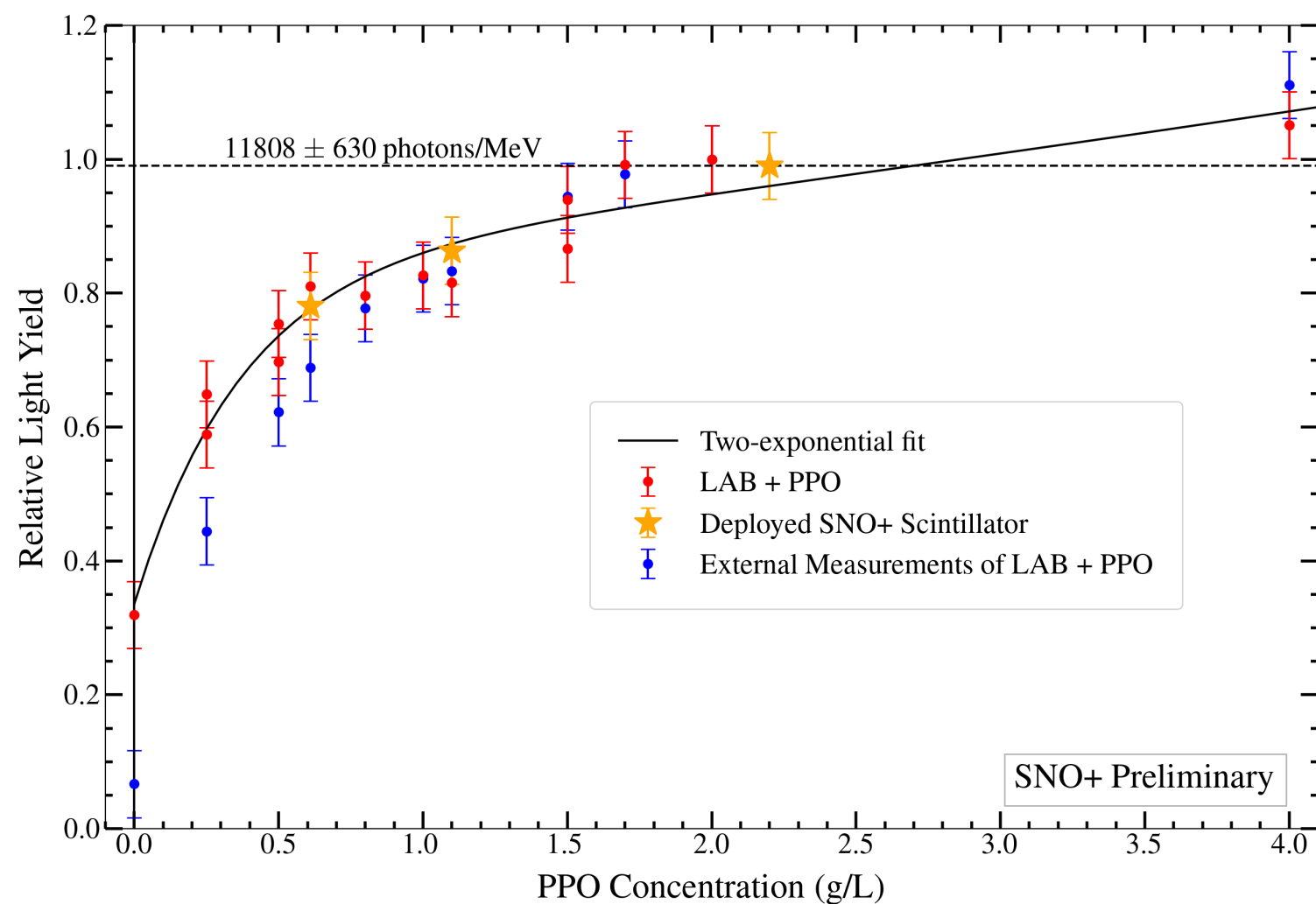
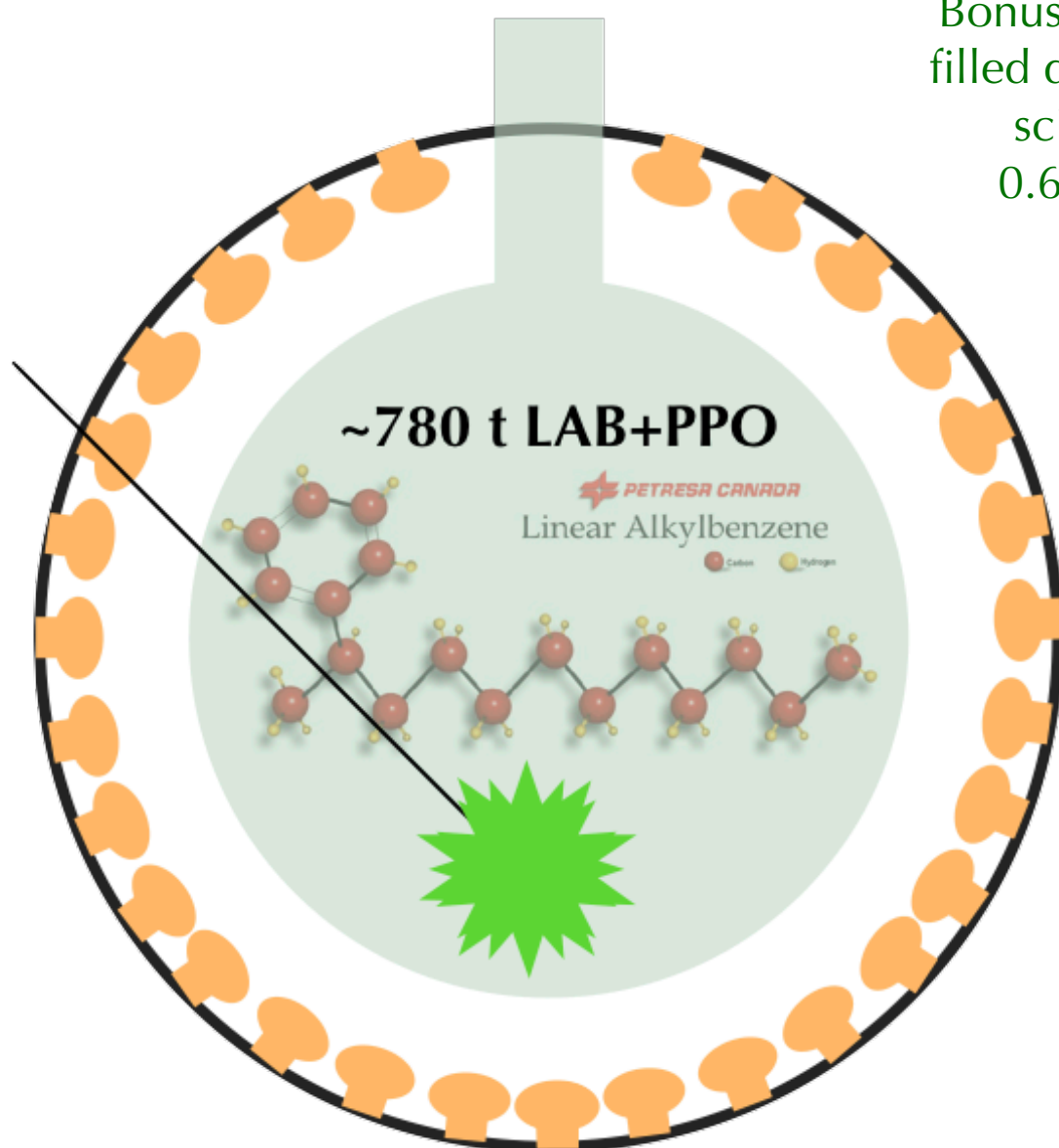
# SNO+ SCINTILLATOR



2017 2018 2019 2020 2021 2022 2023

April - June 2021  
Bonus phase: full-  
filled detector with  
scintillator  
0.6 g/L PPO

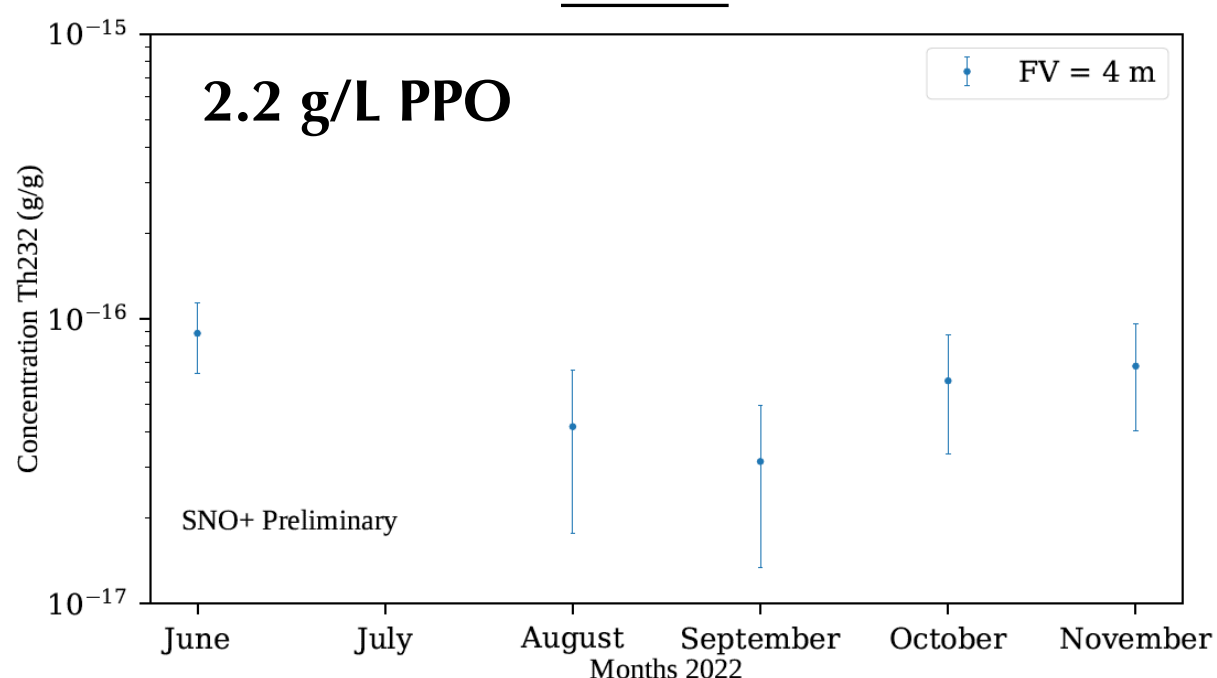
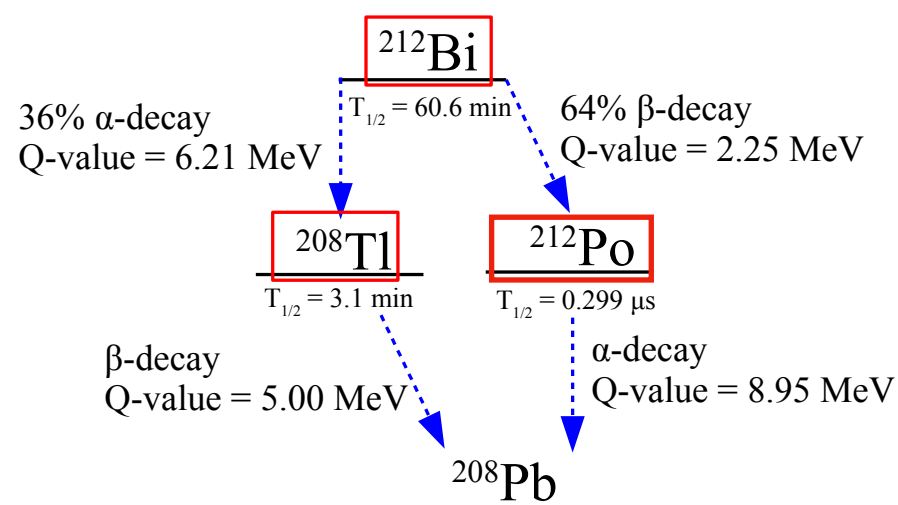
April 2022 - March 2023  
Full fill detector with  
scintillator  
2.2 g/L PPO



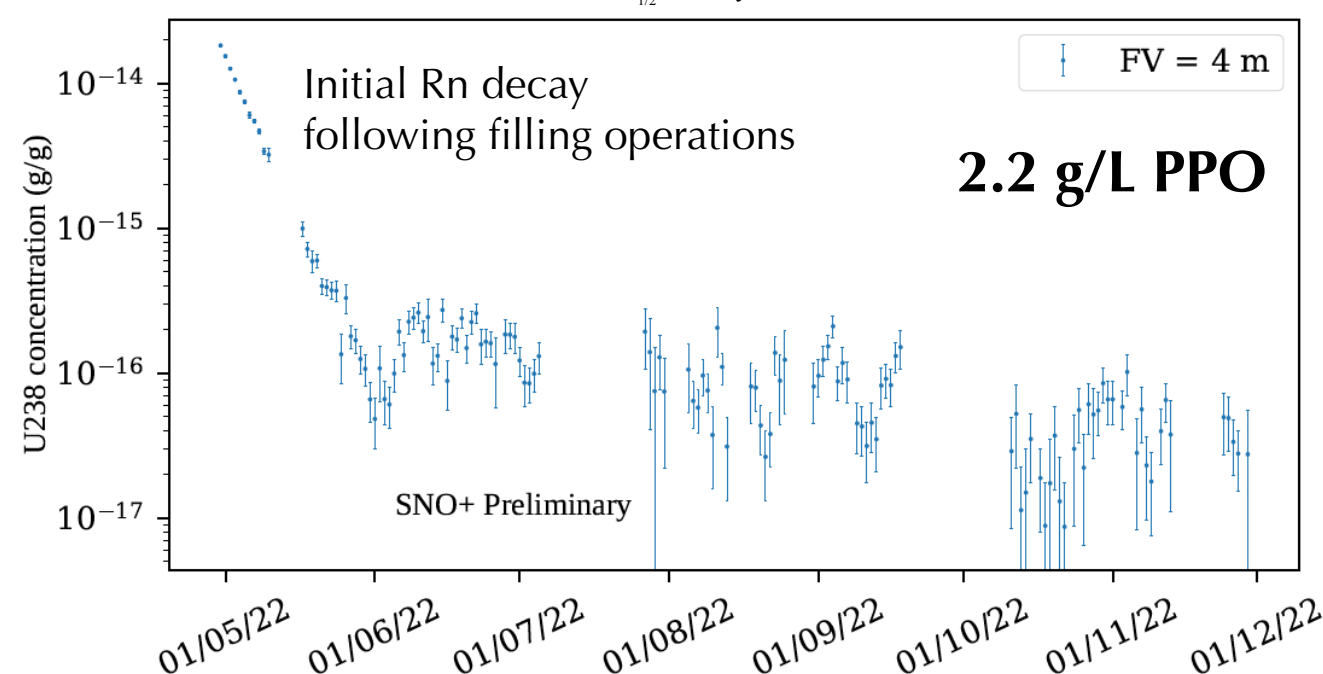
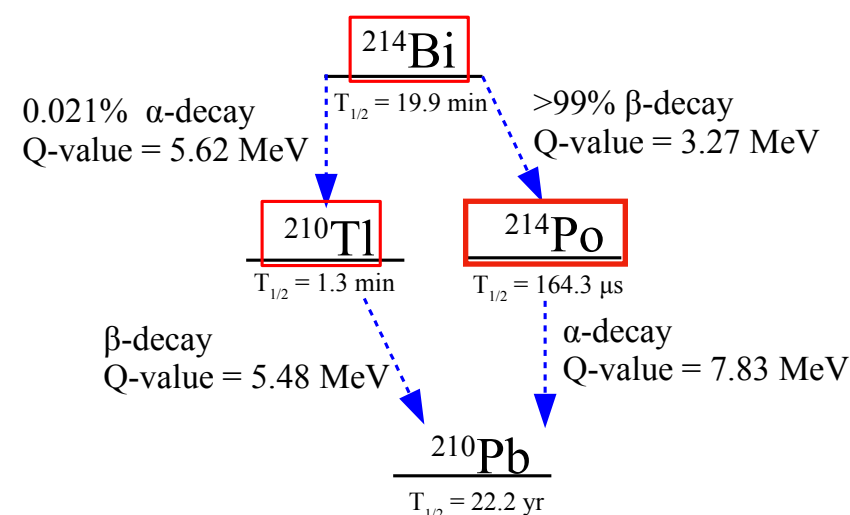


- ❖ Preparation for the double-beta decay phase: background and target-out measurement

## $^{232}\text{Th}$ via $^{212}\text{BiPo}$



## $^{238}\text{U}$ via $^{214}\text{BiPo}$



$$^{232}\text{Th} = (5.7 \pm 0.3) 10^{-17} \text{ g/g}$$

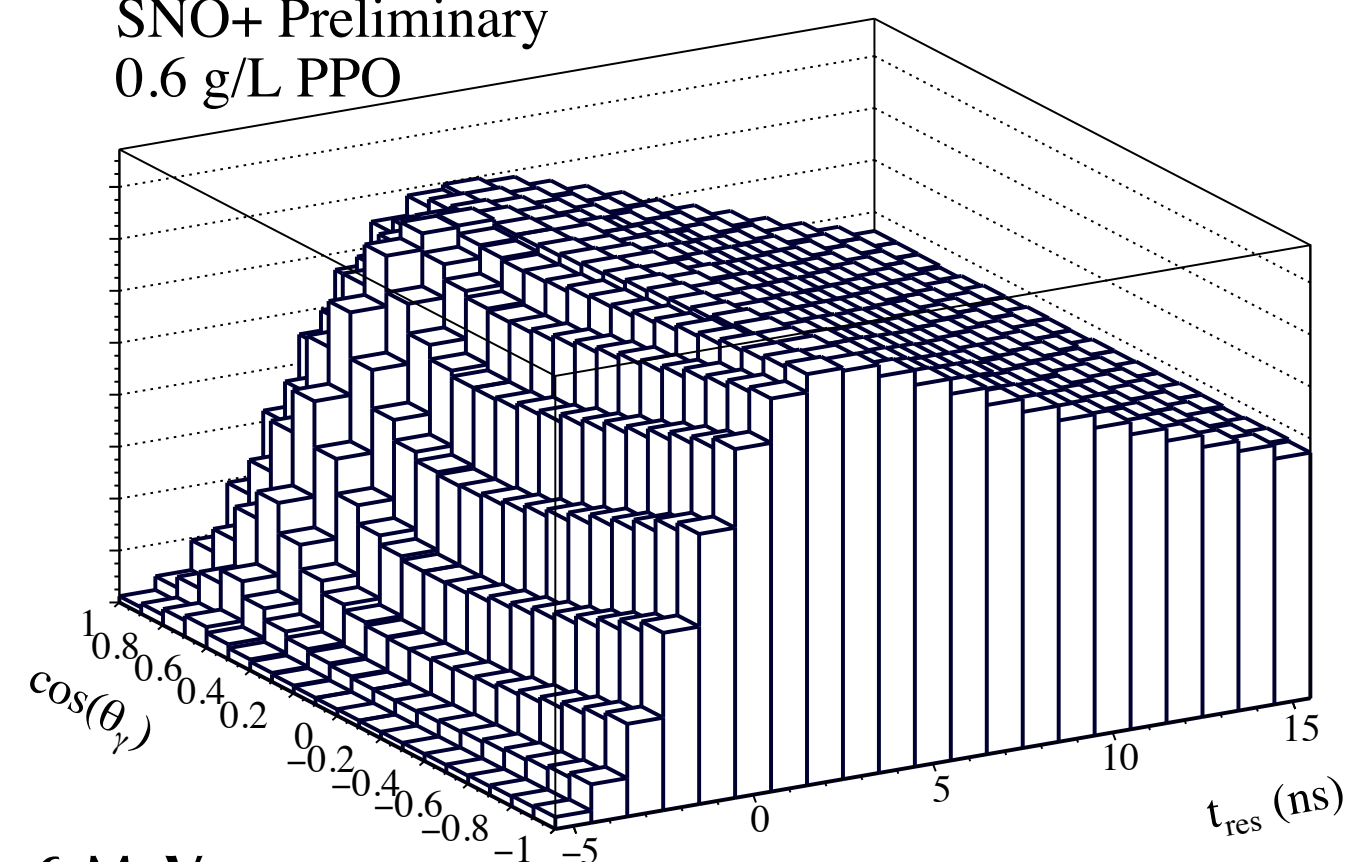
**Below DBD-phase  
requirements!**

$$^{238}\text{U} = (5.3 \pm 0.1) 10^{-17} \text{ g/g}$$



- ❖ Investigation of mitigation strategies for solar neutrinos: Directionality of  $^8\text{B}$
- ❖ Very promising results for the 0.6 g/L PPO scintillator:
  - Determined by fitting prompt timing profiles to combined Cherenkov-scintillation 2D PDFs

SNO+ Preliminary  
0.6 g/L PPO

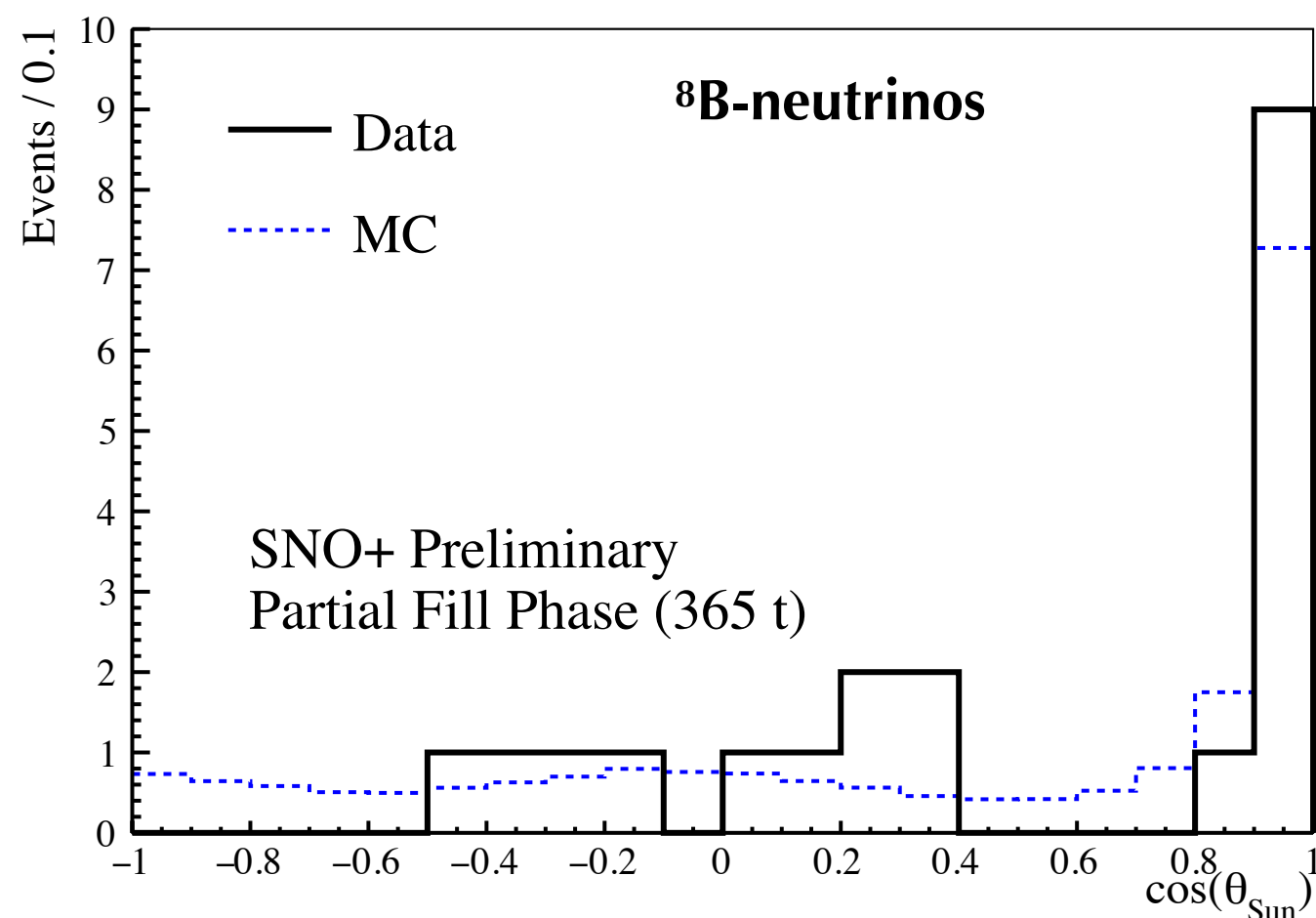
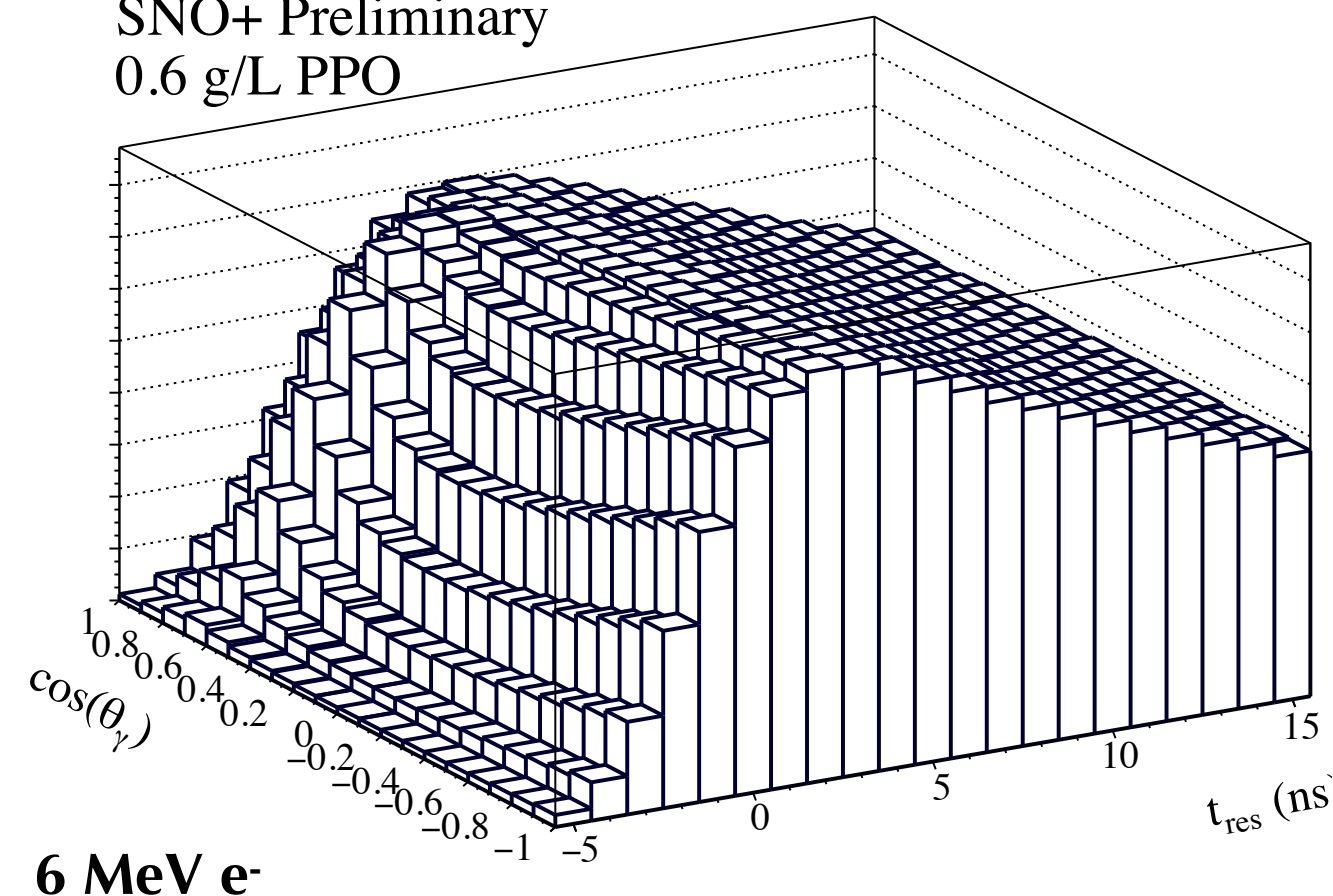


6 MeV  $e^-$



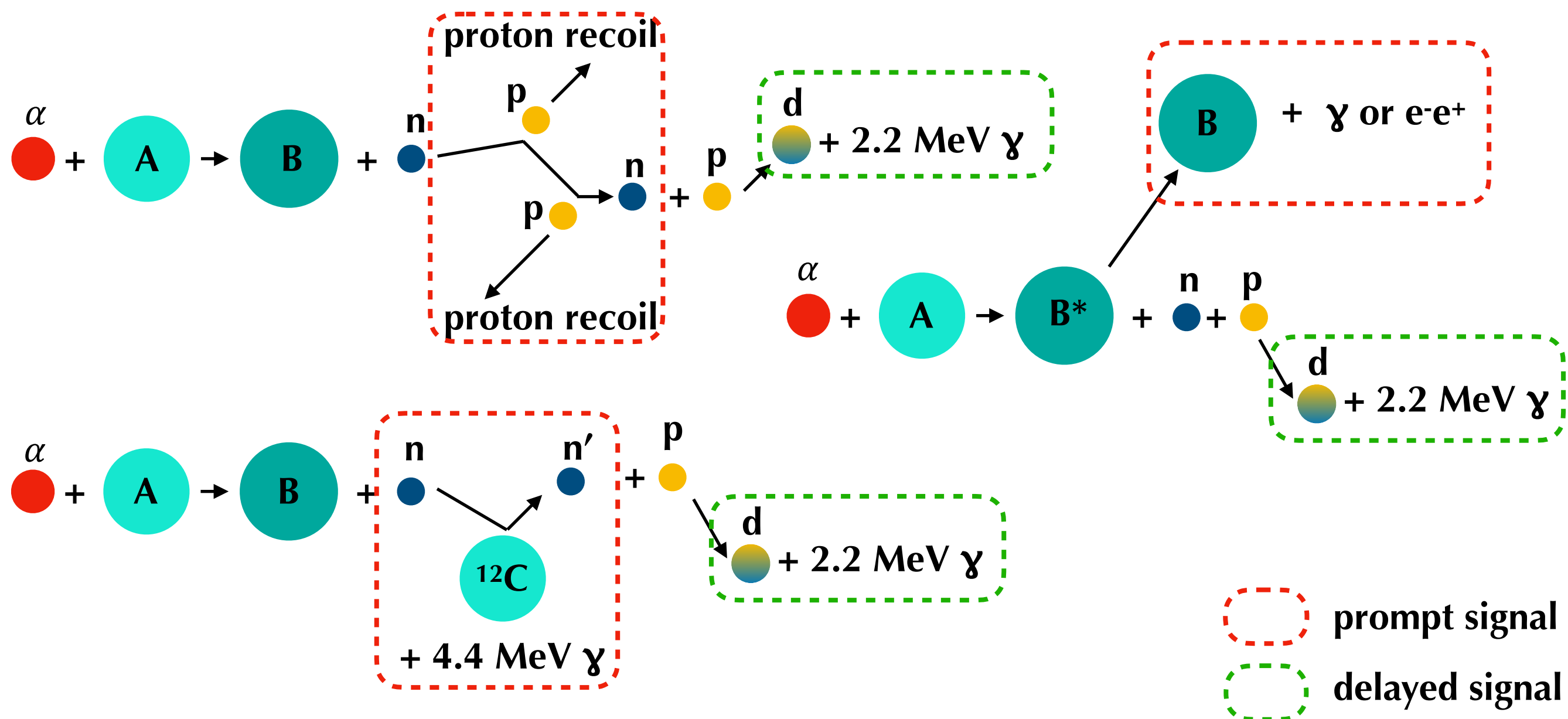
- ❖ Investigation of mitigation strategies for solar neutrinos: Directionality of  $^8\text{B}$
- ❖ Very promising results for the 0.6 g/L PPO scintillator:
  - Determined by fitting prompt timing profiles to combined Cherenkov-scintillation 2D PDFs
  - Event-by-event direction reconstruction compared to Borexino Correlated and Integrated Directionality
    - **first time in liquid scintillation experiment —> Paper in preparation!**

SNO+ Preliminary  
0.6 g/L PPO



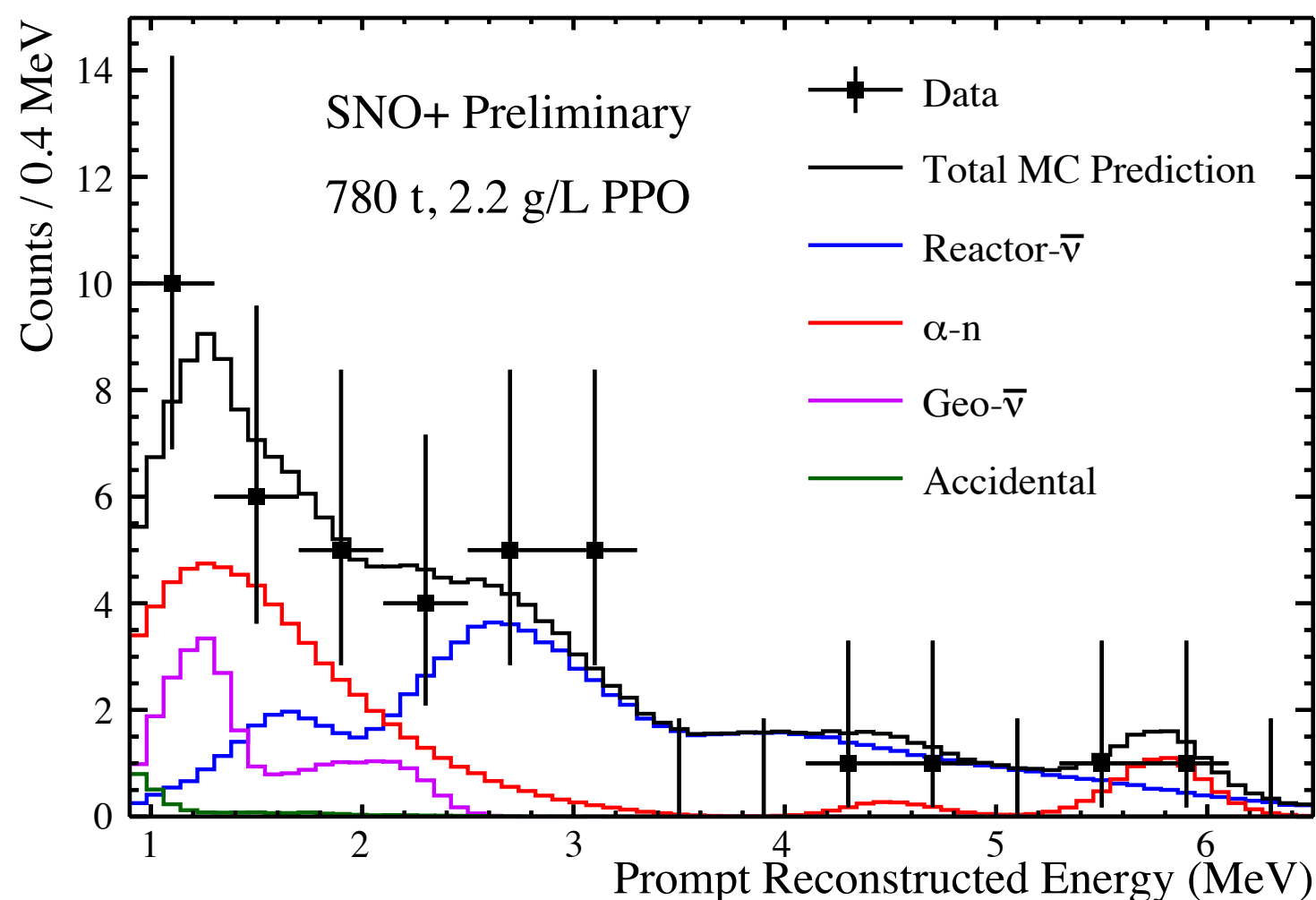
# SNO+ ALPHA, N REACTIONS

- ❖ Major background for the antineutrino analysis
- ❖ Prompt (scattered off protons) and delay (2.2 MeV gamma from neutron capture) can fall in the DBD ROI





- ❖ Major source of alphas is  $^{210}\text{Po}$ 
  - ❖ Highly reduced (70%)  $^{210}\text{Po}$  background from partial fill to the 2.2 g/L full fill phase
- ❖ Developed a classifier to separate  $^{13}\text{C}(\alpha,n)$  reactions from anti-neutrinos



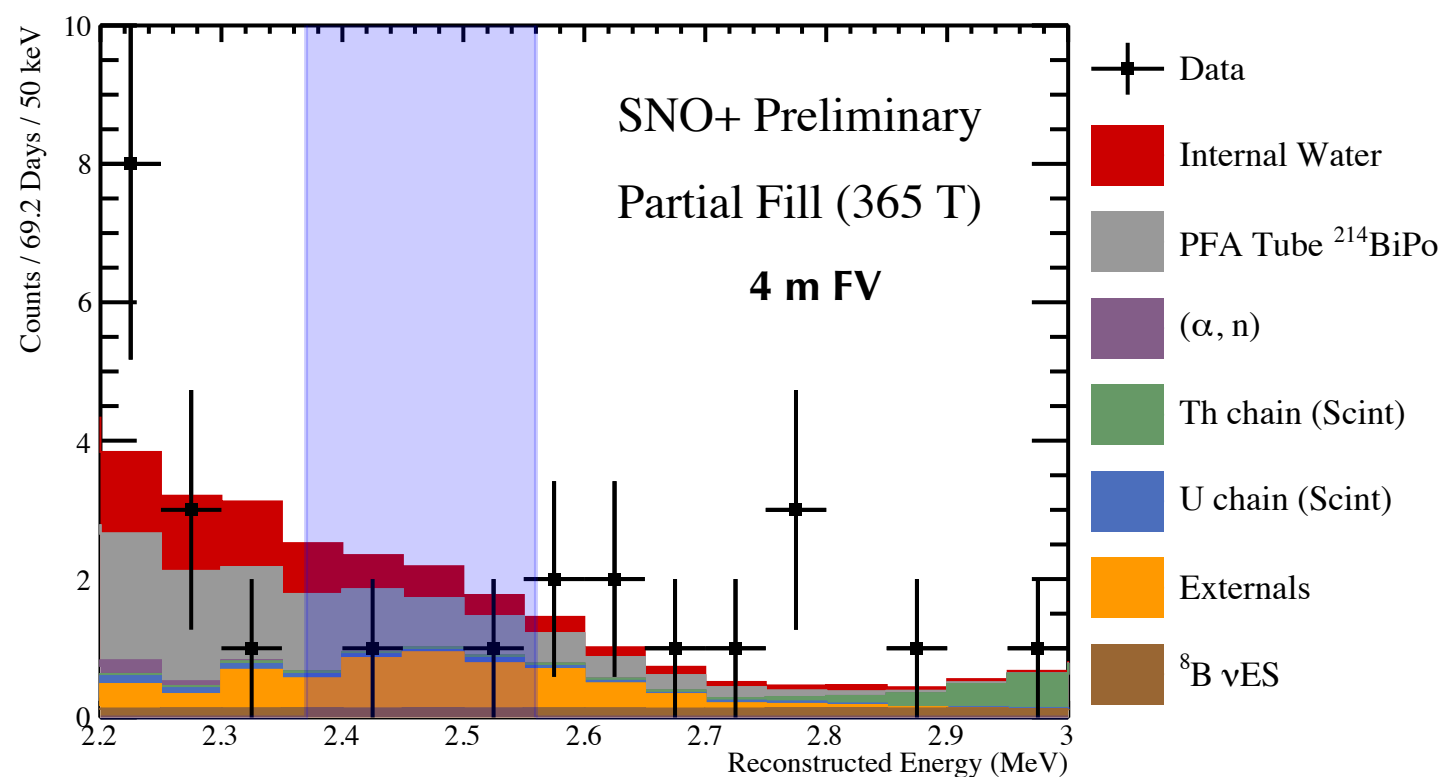
See talk  
T. Kaptanoglu  
Aug 31, 2023, 3:15PM,  
Hörsaal 21 lecture hall

Reactor IBD oscillated using:

- ❖  $\Delta m^2_{21} = 7.53 \times 10^{-5} \text{ eV}^2$
- ❖  $\sin^2(\theta_{12}) = 0.307$



- ❖ Preparation for the double-beta decay phase: background and target-out measurement



**Partial fill:**

- ❖ Expected 8 events, seen 2

**Full fill + 2.2 g/L PPO:**

- ❖ Analysis in progress

June - October 2020  
(~69 gold physics days)

Bonus phase:  
partial fill  
0.6 g/L PPO

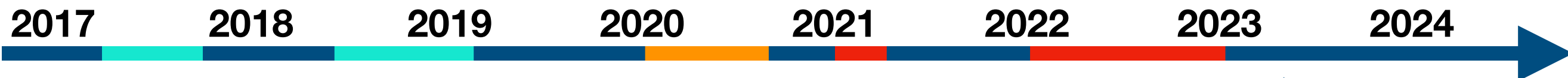
See poster  
B. Tam

*$0\nu\beta\beta$  Target Out Analysis for the SNO+ Experiment*



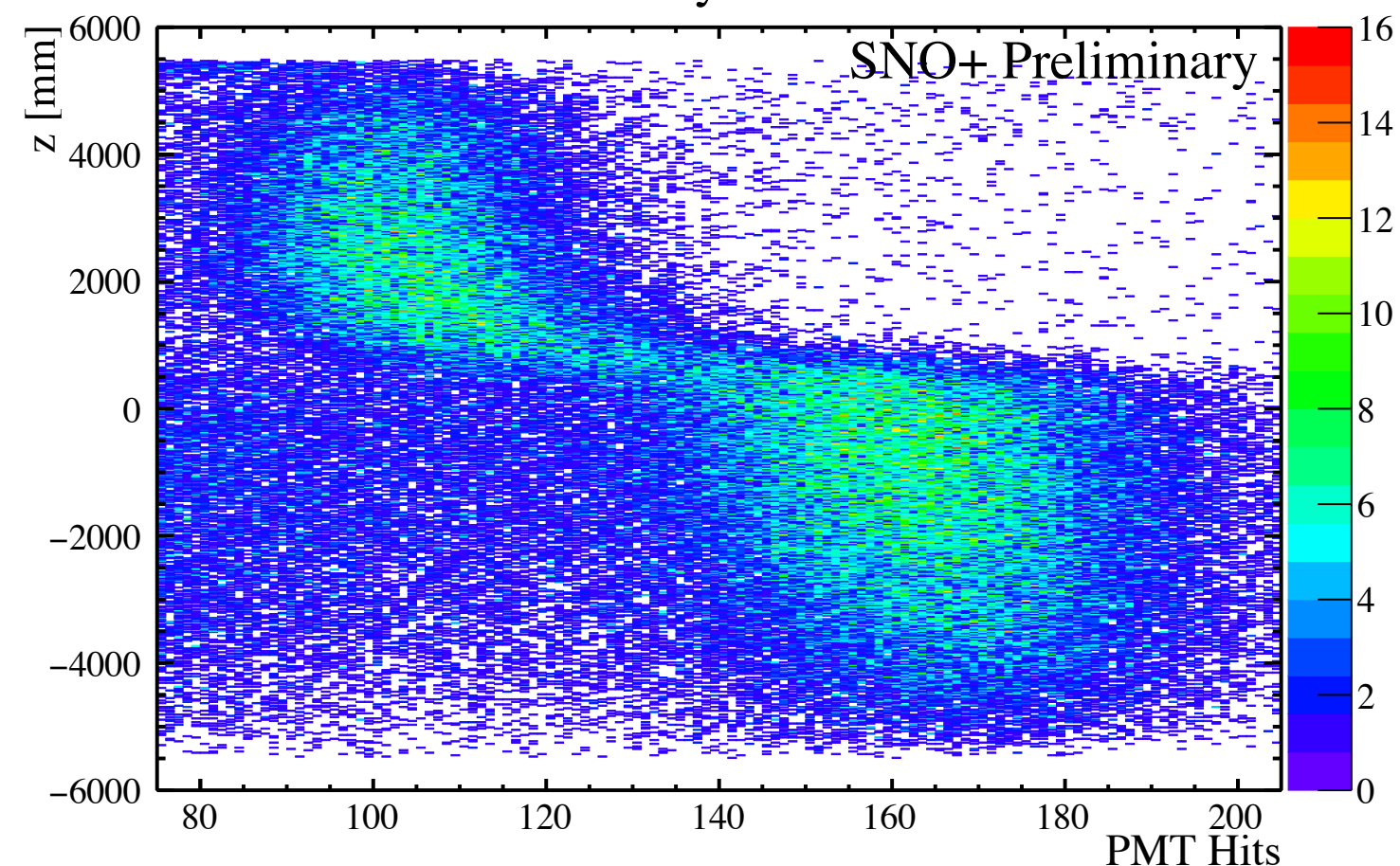


# SNO+ BisMSB

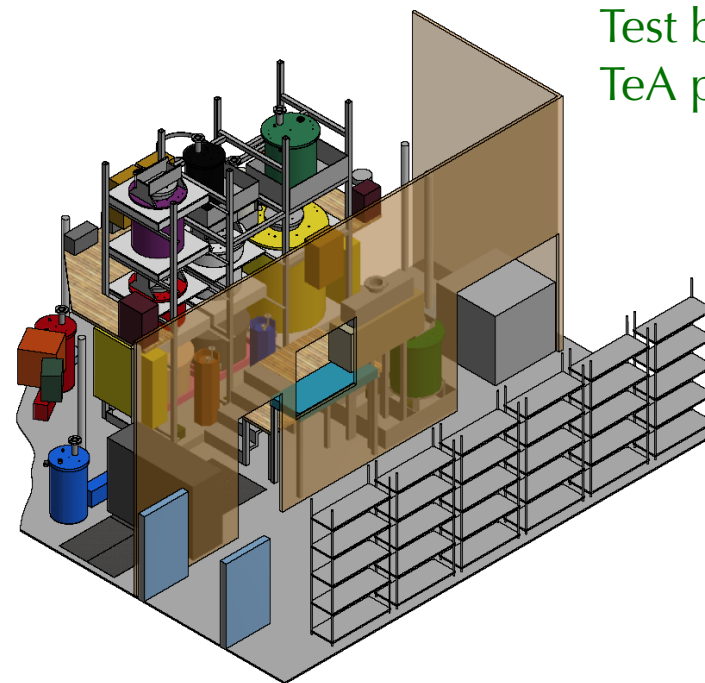
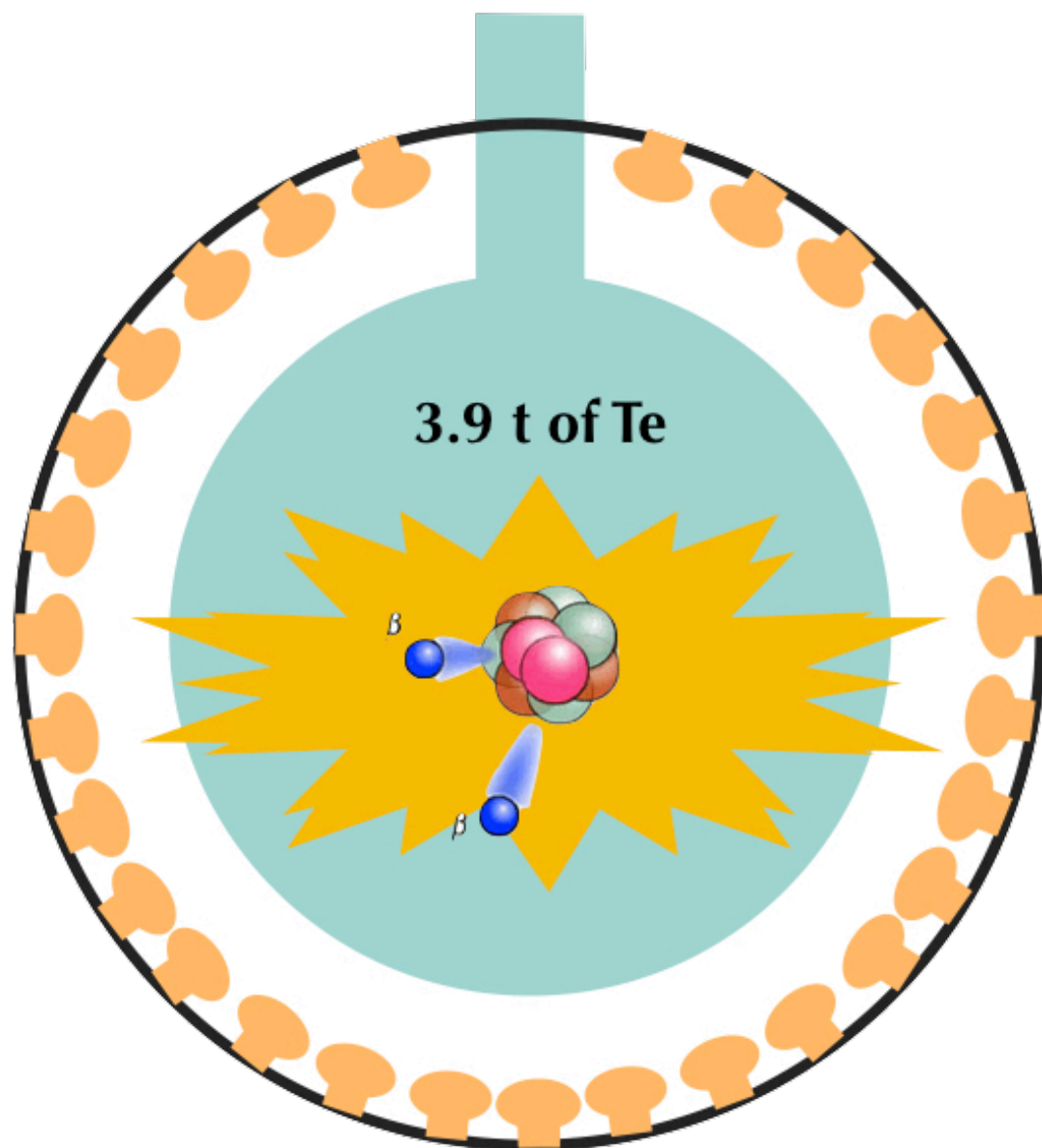
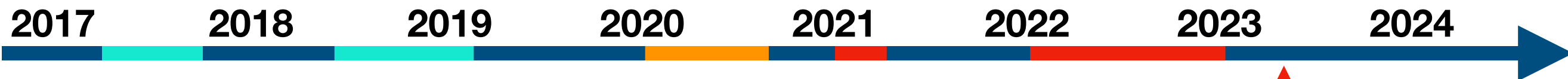


July 2023  
Started to add  
bisMSB to the  
detector

21 July 2023



- ❖ Tracking  $^{210}\text{Po}$  decays
- ❖ BisMSB added at the bottom of the detector (0.5 kg) and started to mix
- ❖ **Clear improvement (~1.5x) in light output**



Fall 2023  
Test batch (~200 kg) of the  
TeA purification plant

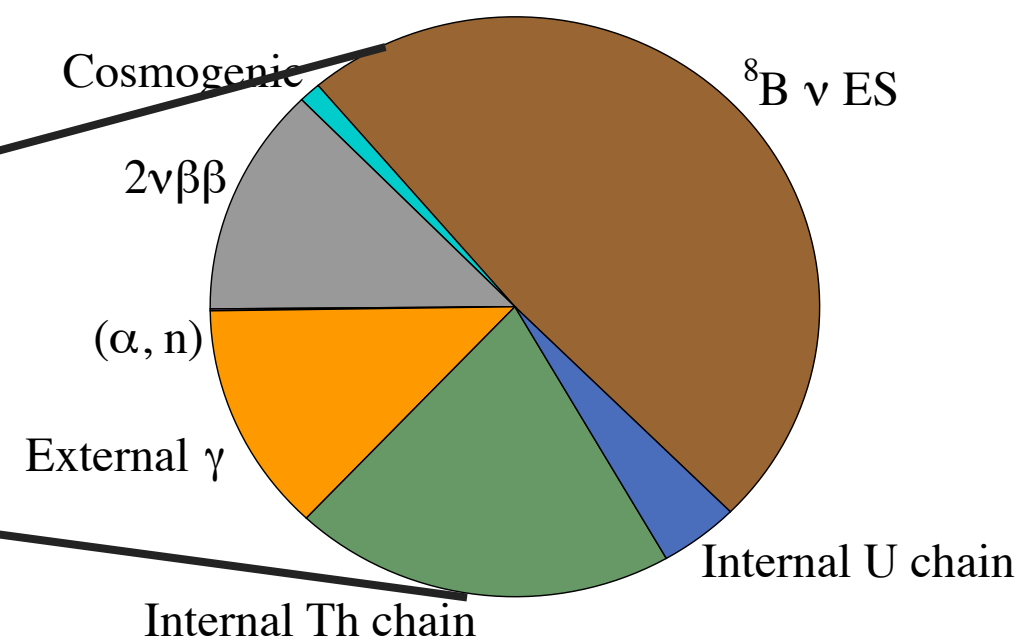
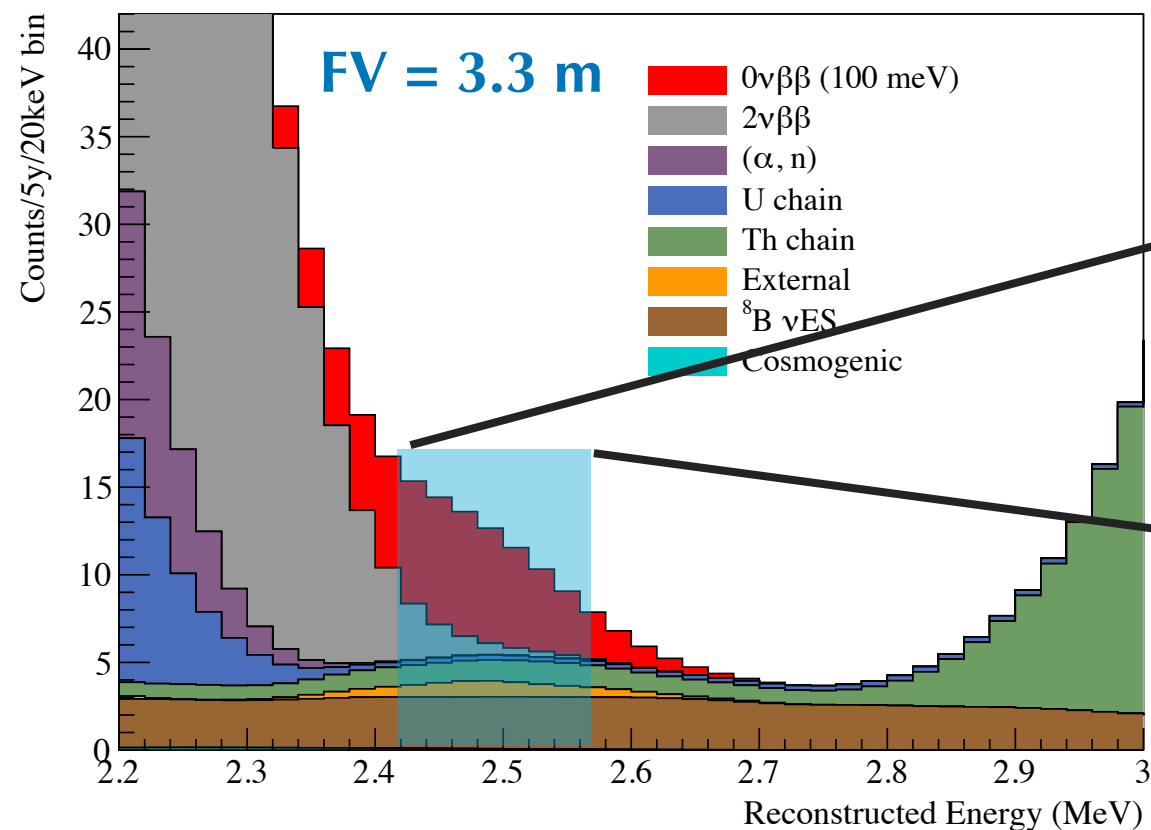
## ❖ Important milestone:

- First full-scale tests of SNO+ Te purification and loading systems
- Samples will be collected for off-site ICP-MS analysis (U/Th)



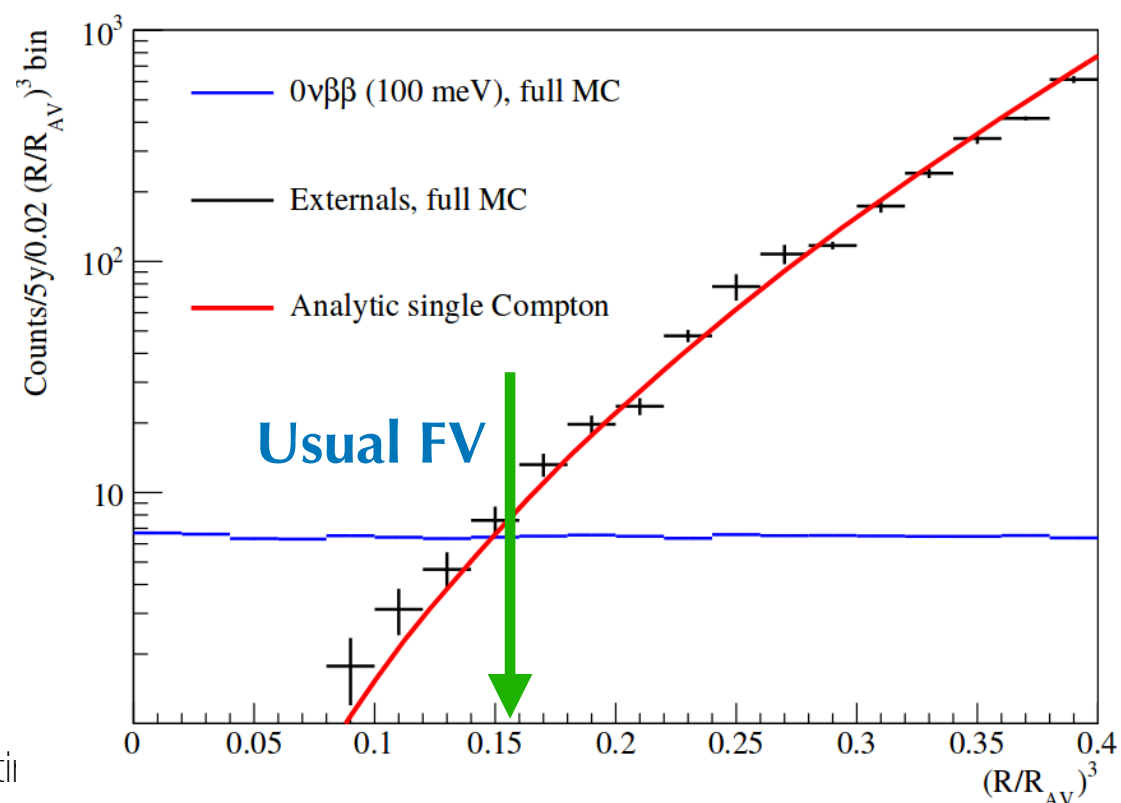


# EXPECTED SIGNAL/BACKGROUND



*Events in the Region Of Interest  
+ Fiducial Volume*

**9.47 events/yr (at nominal  
backgrounds)**

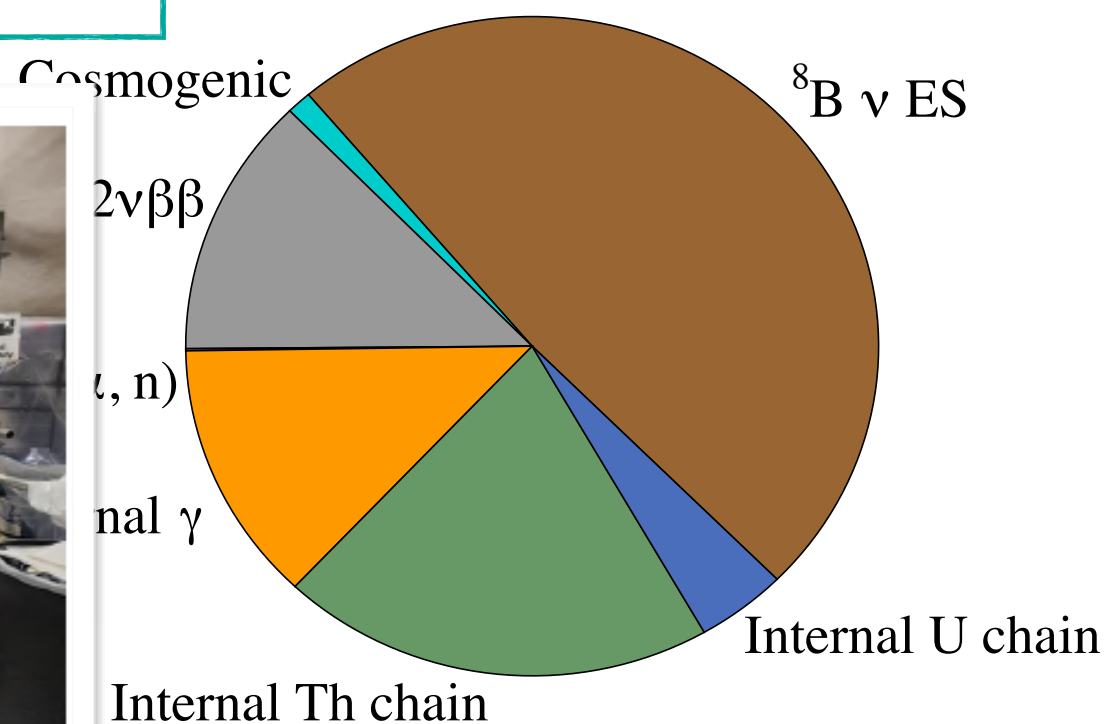




## Cosmogenic:

- \*  $^{60}\text{Co}$ ,  $^{110\text{m}}\text{Ag}$ ,  $^{88}\text{Y}$ ,  $^{22}\text{Na}$
- \* mitigation: purification + “cool-down” UG
- \* First Te UG since 2015
- \* multi-site classifier

## Background breakdown in $0\nu\beta\beta$ ROI



TeA cooling UG since 2015

See Poster

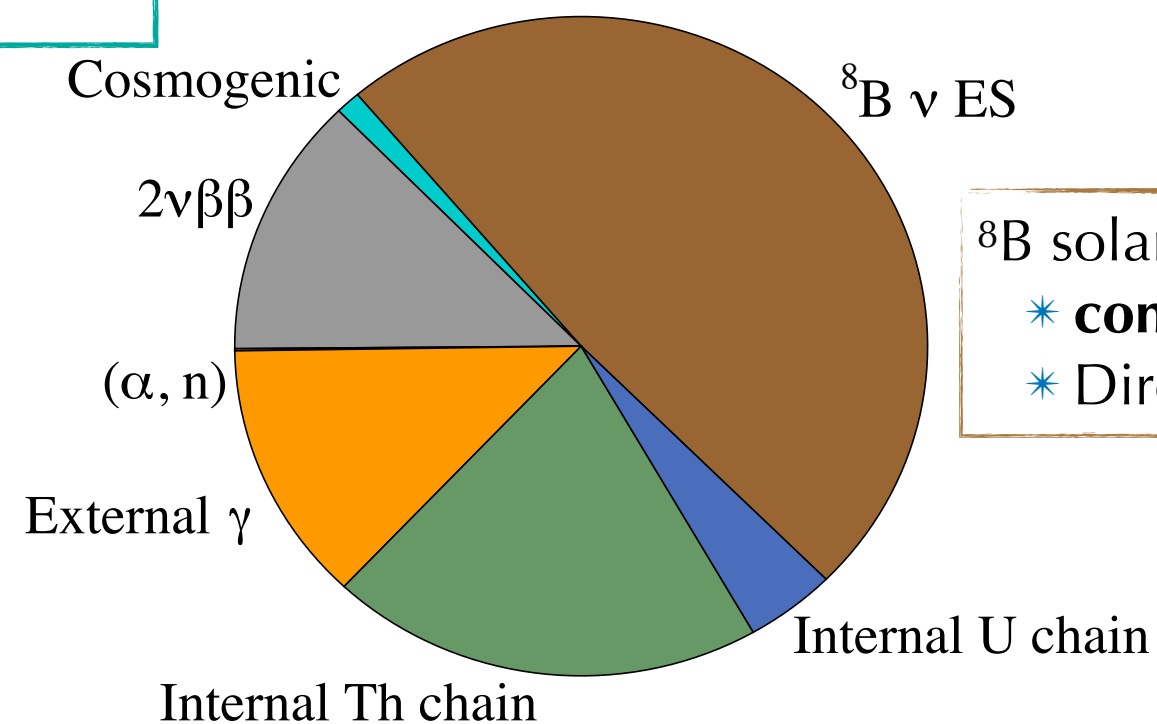
T. Kroupova

Event Reconstruction in the SNO+ Experiment

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## Background breakdown in $0\nu\beta\beta$ ROI



## $^8\text{B}$ solar neutrinos:

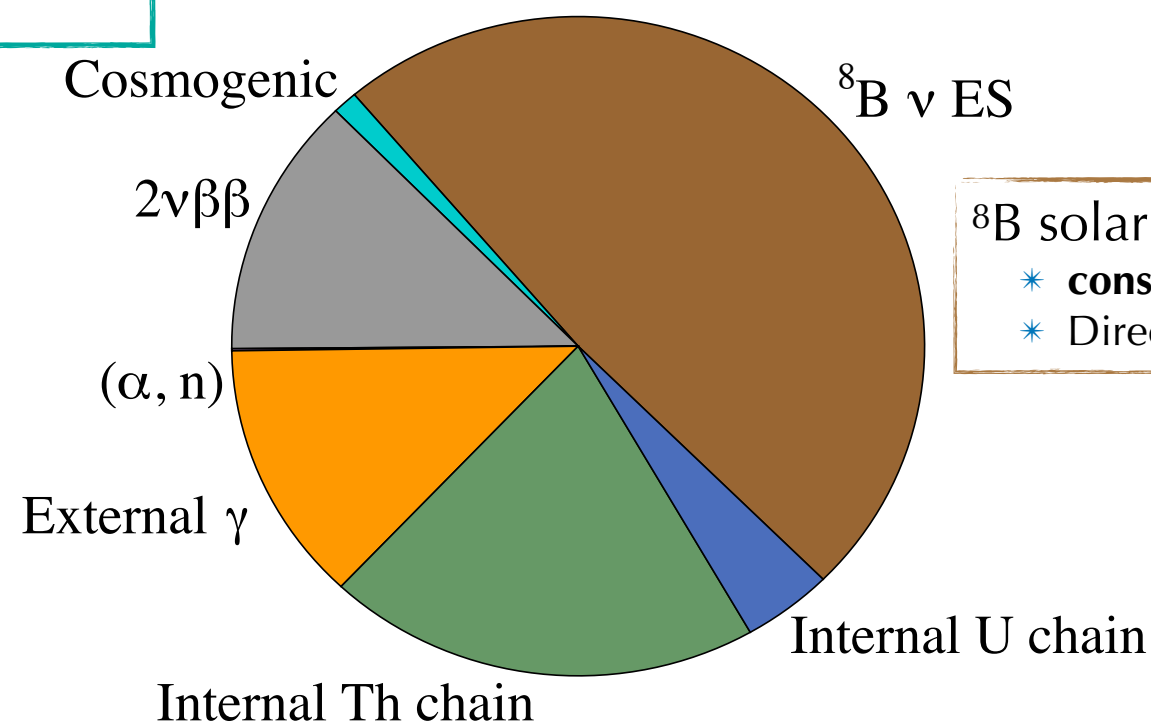
- \* **constrained by SNO/SK data**
- \* Directionality?



## Cosmogenic:

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## Background breakdown in $0\nu\beta\beta$ ROI



## $^8\text{B}$ solar neutrinos:

- \* constrained by SNO/SK data
- \* Directionality?

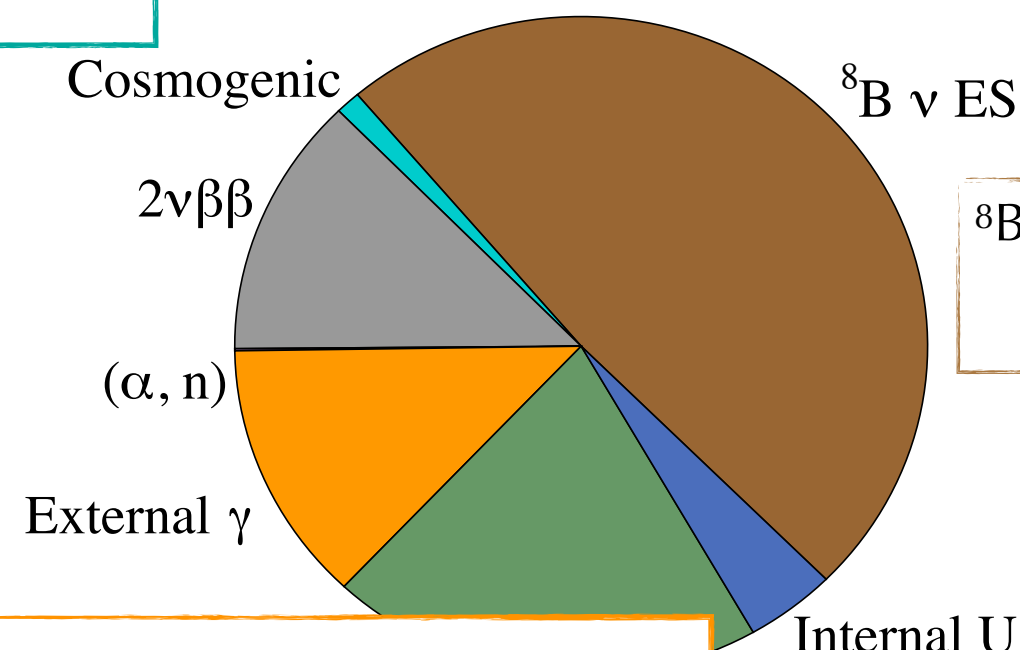
## Internal U/Th chain:

- \*  $^{214}\text{BiPo}$ ,  $^{212}\text{BiPo}$ ,  $^{210}\text{Tl}$
- \* LAB components below requirements for the TeLS phase
- \* Constantly monitoring the contribution from the scintillator before and while adding Te
- \*  $\beta - \alpha$  delayed coincidence tagging + in-window rejection (expected more than 98% rejection on in-window events)

## Background breakdown in $0\nu\beta\beta$ ROI

### Cosmogenic:

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### $^8\text{B}$ solar neutrinos:

- \* **constrained by SNO/SK data**
- \* Directionality?

### External gammas:

- \* from AV, ropes, water, PMTs, mainly  $^{208}\text{Tl}$
- \* **Measured in water phase, below nominal values**
- \* Fiducialization to minimize the leak in ROI

### Internal U chain

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## Background breakdown in $0\nu\beta\beta$ ROI

### Cosmogenic:

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- \* mitigation: purification + “cool-down” UG
- \* First Te UG since 2015
- \* multi-site classifier

Cosmogenic

$2\nu\beta\beta$

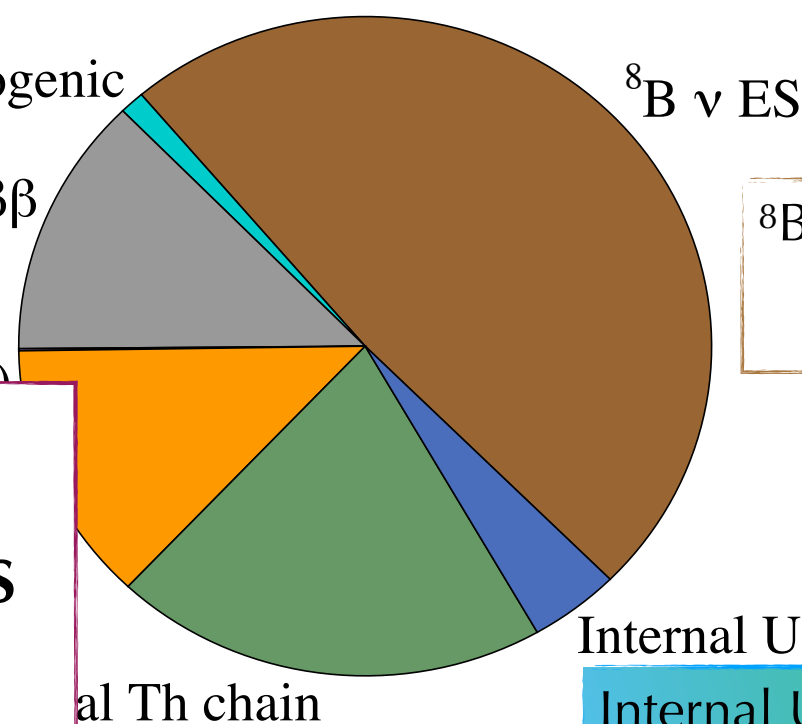
( $\alpha, n$ )

( $\alpha, n$ ):

- \* alpha-capture on  $^{13}\text{C}$  ( $^{18}\text{O}$ )
- \* **Measure the contribution from LS components before Te loading**
- \* delayed coincidence tagging

### External gammas:

- \* from AV, ropes, water, PMTs, mainly  $^{208}\text{Tl}$
- \* **Measured in water phase, below nominal**
- \* Fiducialization to minimize the leak in ROI



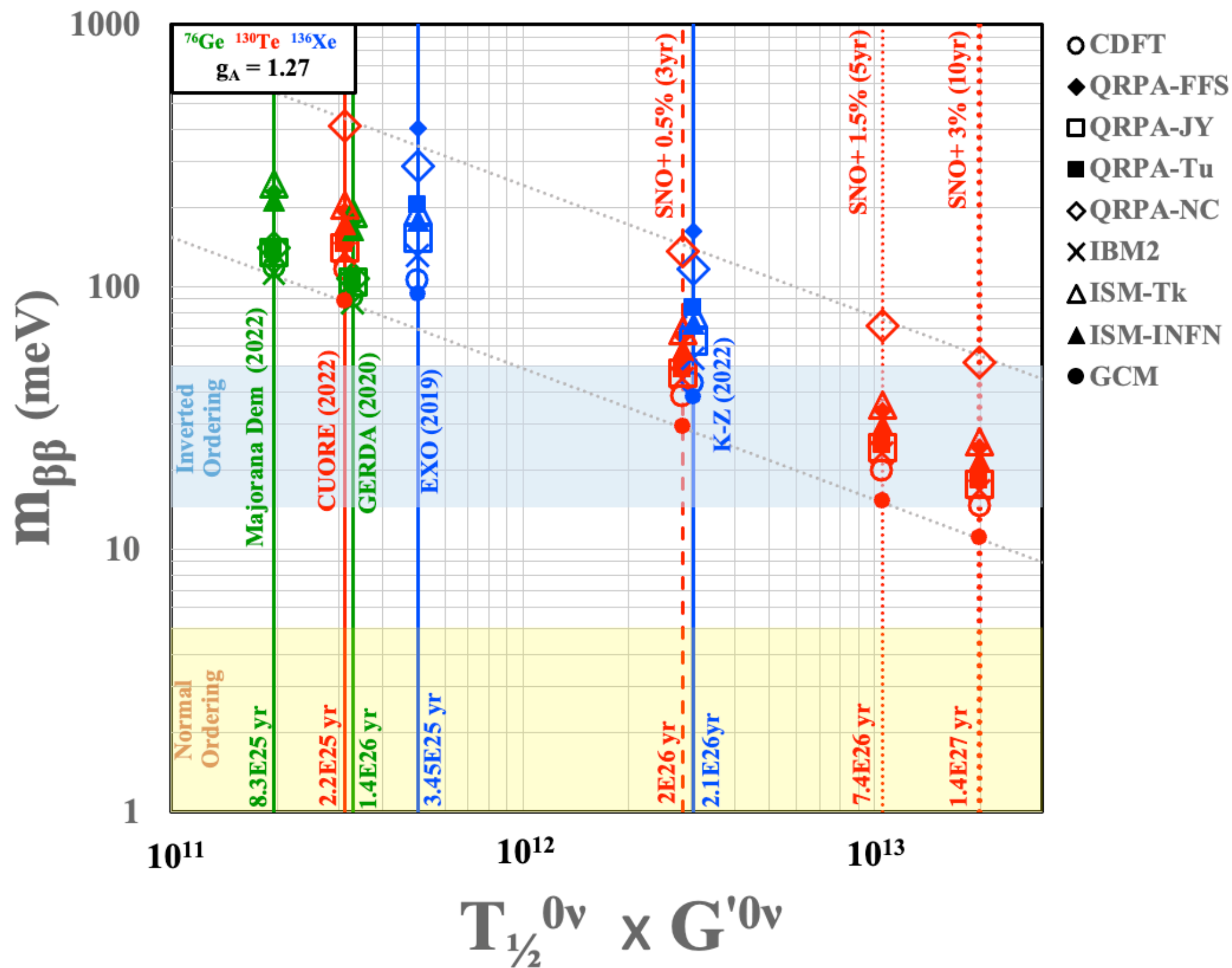
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- \* **constrained by SNO/SK data**
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Internal U chain

### Internal U/Th chain:

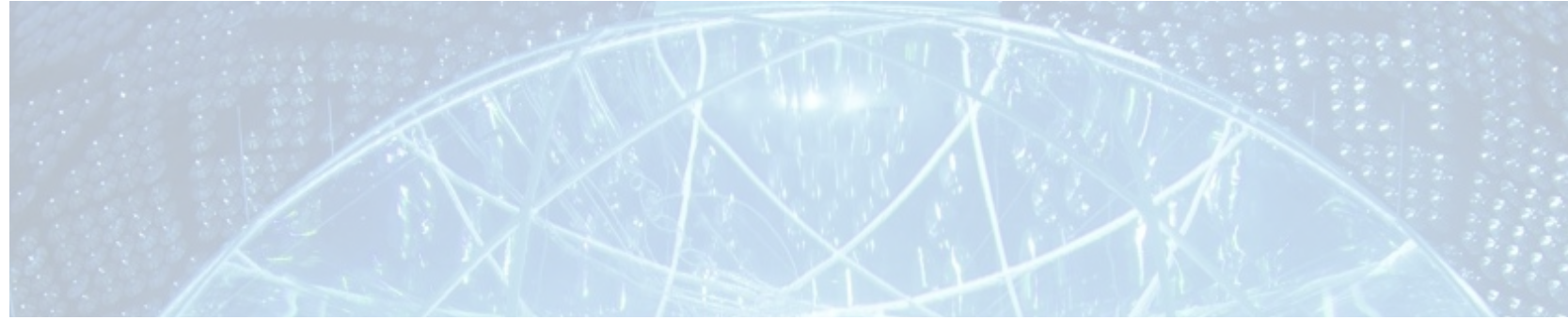
- \*  $^{214}\text{BiPo}$ ,  $^{212}\text{BiPo}$ ,  $^{210}\text{Tl}$
- \* **LAB components below requirements for the TeLS phase**
- \* **Constantly monitoring the contribution from the scintillator before adding Te**
- \*  $\beta - \alpha$  delayed coincidence tagging + in-window rejection





- ▶ SNO+ has successfully completed its scintillator loading and is taking data with 2.2 g/L PPO as of April 2022
- ▶ Initial measurements show radioactive backgrounds at or below the targeted values
- ▶ Many exciting physics publications are expected in the very near future!





LIP Coimbra  
LIP Lisboa



SNOLAB  
TRIUMF  
University of Alberta  
Queen's University  
Laurentian University



TU Dresden



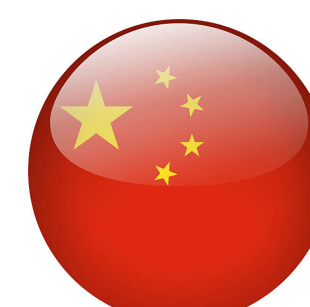
UNAM



Boston University  
BNL  
University of California Berkeley  
LBNL  
University of Chicago  
University of Pennsylvania  
UC Davis

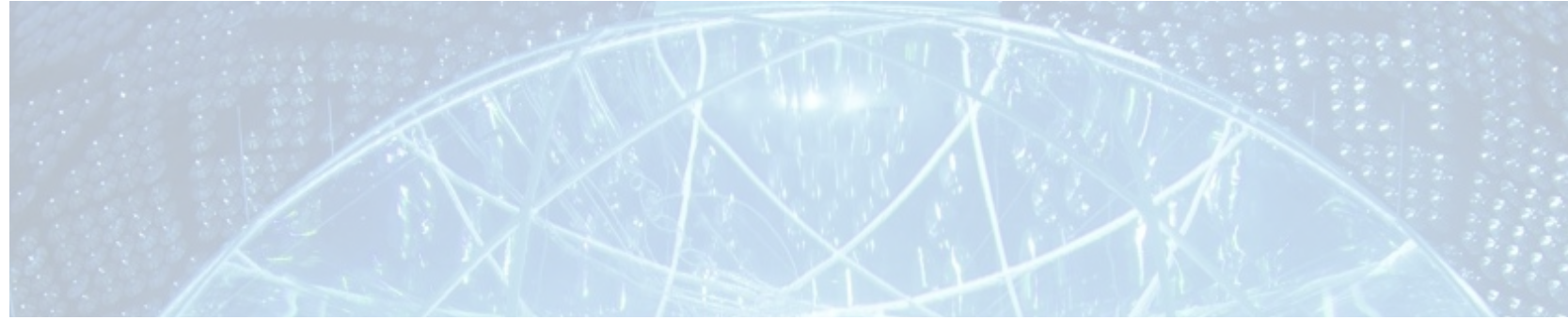


Oxford University  
Kings College London  
University of Liverpool  
University of Sussex  
University of Lancaster



Shandong University





LIP Coimbra  
LIP Lisboa



SNOLAB  
TRIUMF  
University of Alberta  
Queen's University  
Laurentian University



TU Dresden



UNAM



Boston University  
BNL  
University of California Berkeley  
LBNL  
University of Chicago  
University of Pennsylvania  
UC Davis



Oxford University  
Kings College London  
University of Liverpool  
University of Sussex  
University of Lancaster

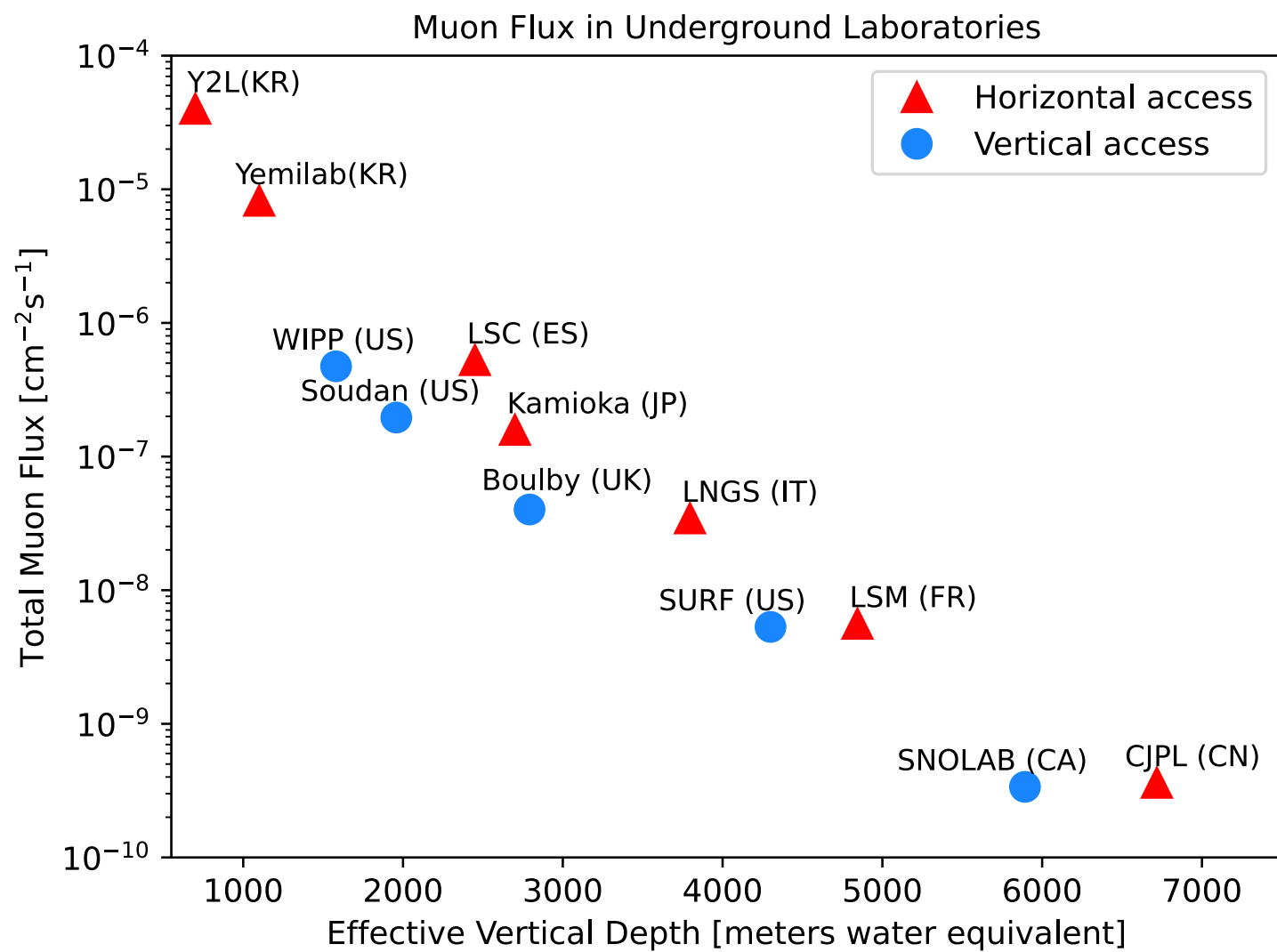


Shandong University

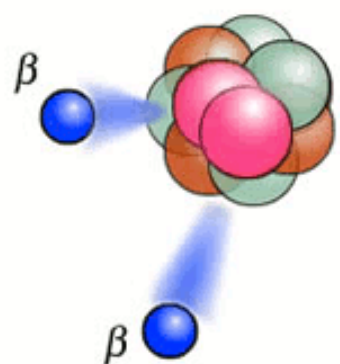
*Back up*



@SNOLAB, Sudbury, Canada



0/1 = not doable/doable in  
water, LS, TeLS phase



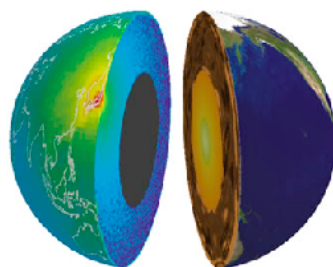
Neutrinoless  
double-beta  
decay of  $^{130}\text{Te}$   
(0,0,1)



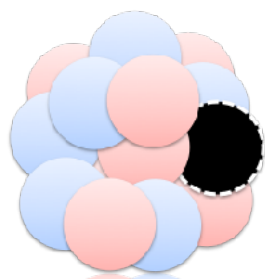
Solar neutrinos (1,1,1)



Reactor antineutrinos  
(1,1,1)



Geo anti-neutrinos  
(0,1,1)

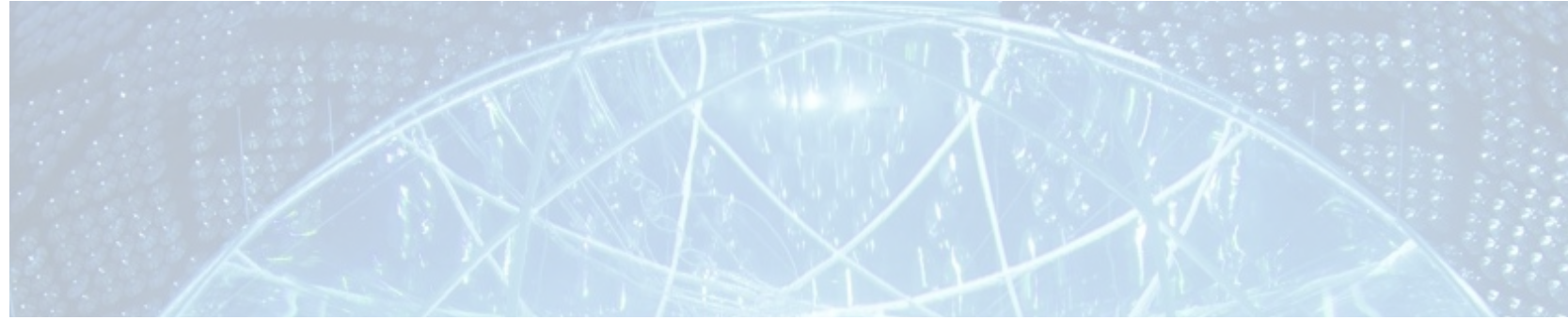


Rare decays and  
interactions  
(1,1,1)



Supernovae  
neutrinos  
(1,1,1)





May - December 2017  
(~115 gold physics days)  
First SNO+ water phase  
*Phys.Rev.D 99, 032008*

October 2018 - June 2019  
(~185 gold physics days)  
Second SNO+ water phase  
*Phys.Rev.D 105, 112012*

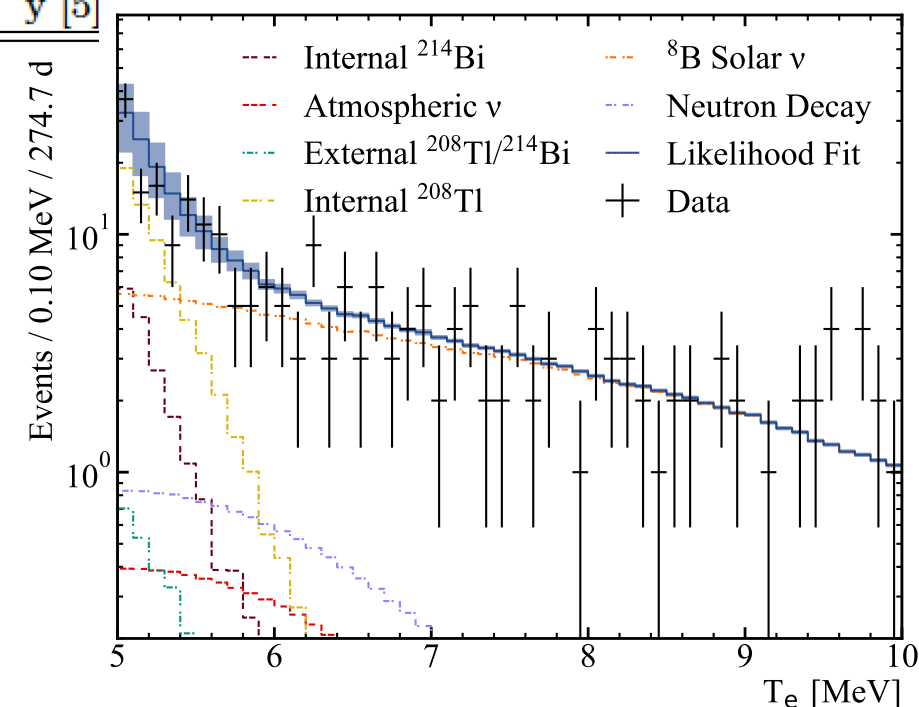
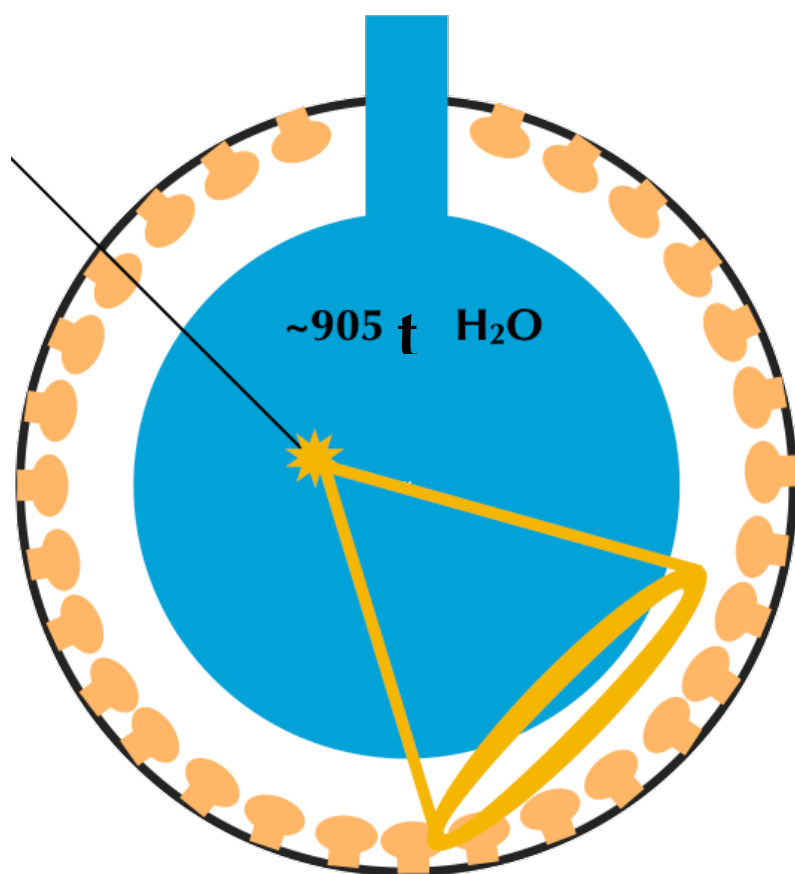
Additional Cover gas  
shielding to reduce Rn  
ingresses in water

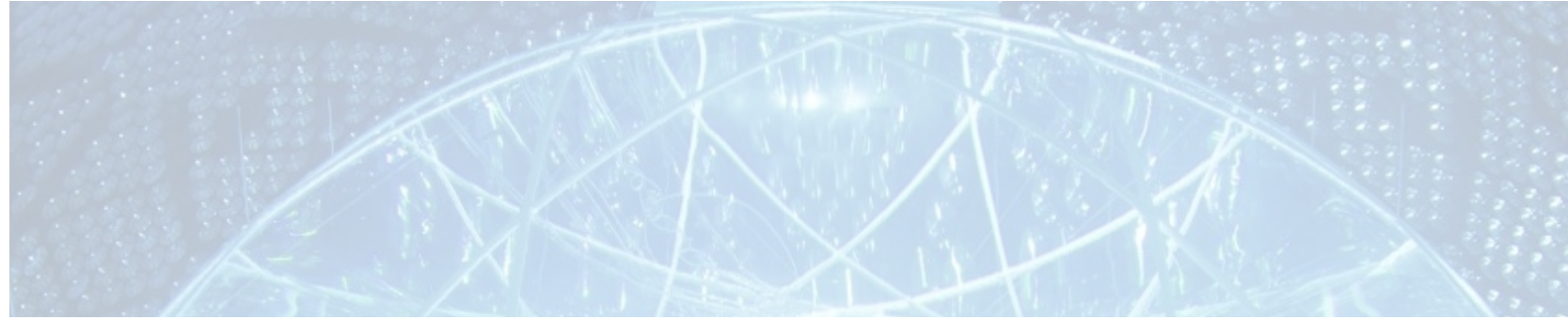
## Major Outcomes

- ✱ Nucleon decay modes into invisible channels

Decay Mode	Partial Lifetime Limit	Existing Limits
n	$9.0 \times 10^{29}$ y	$5.8 \times 10^{29}$ y [5]
p	$9.6 \times 10^{29}$ y	$3.6 \times 10^{29}$ y [6]
pp	$1.1 \times 10^{29}$ y	$4.7 \times 10^{28}$ y [6]
np	$6.0 \times 10^{28}$ y	$2.6 \times 10^{28}$ y [6]
nn	$1.5 \times 10^{28}$ y	$1.4 \times 10^{30}$ y [5]

Best limit  
for all  
channels!





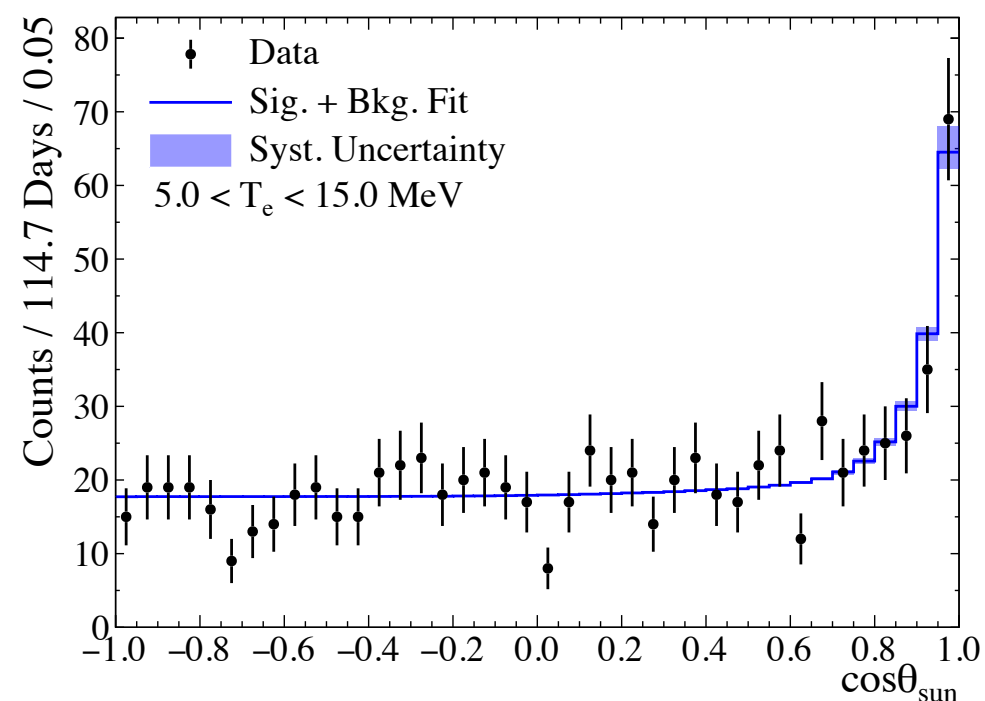
May - December 2017  
(~115 gold physics days)  
First SNO+ water phase  
*Phys.Rev.D 99, 012012*

October 2018 - June 2019  
(~185 gold physics days)  
Second SNO+ water phase  
*In preparation*

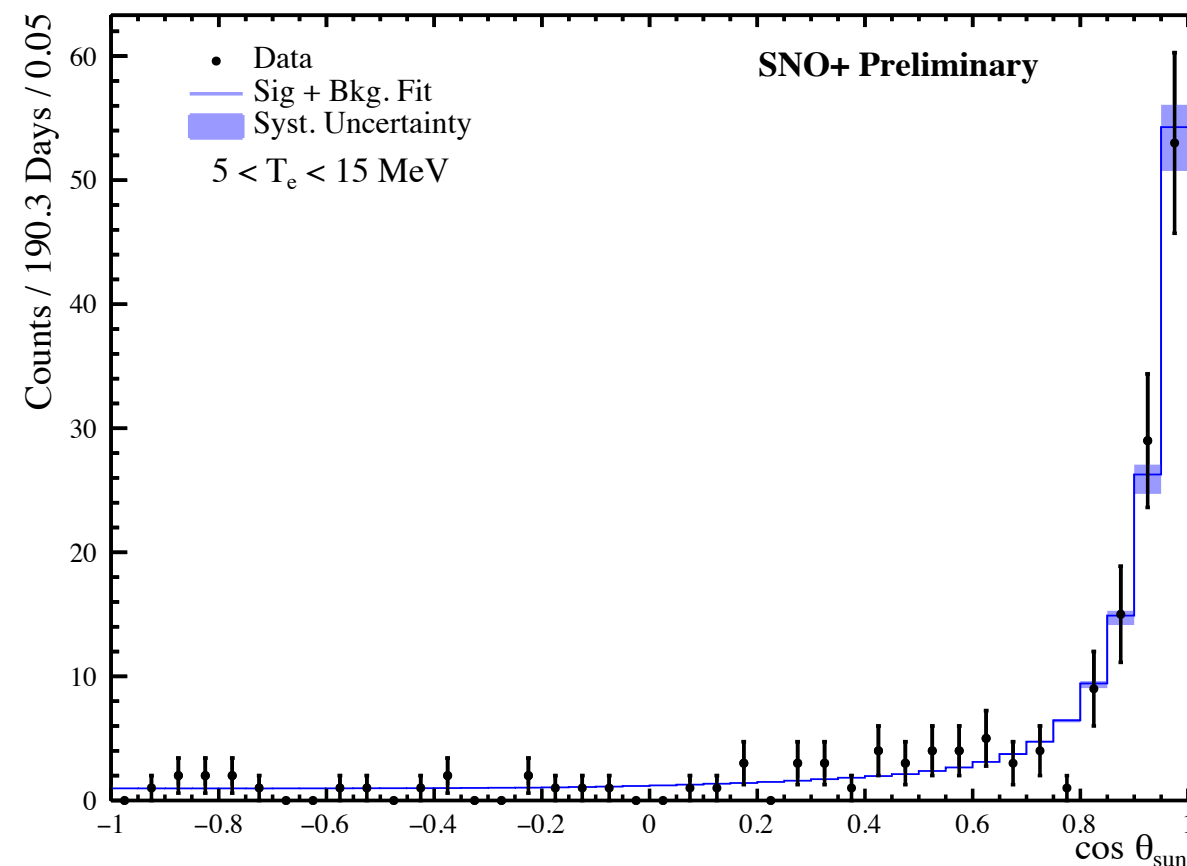
Additional Cover gas  
shielding to reduce Rn  
ingresses in water

## Major Outcomes

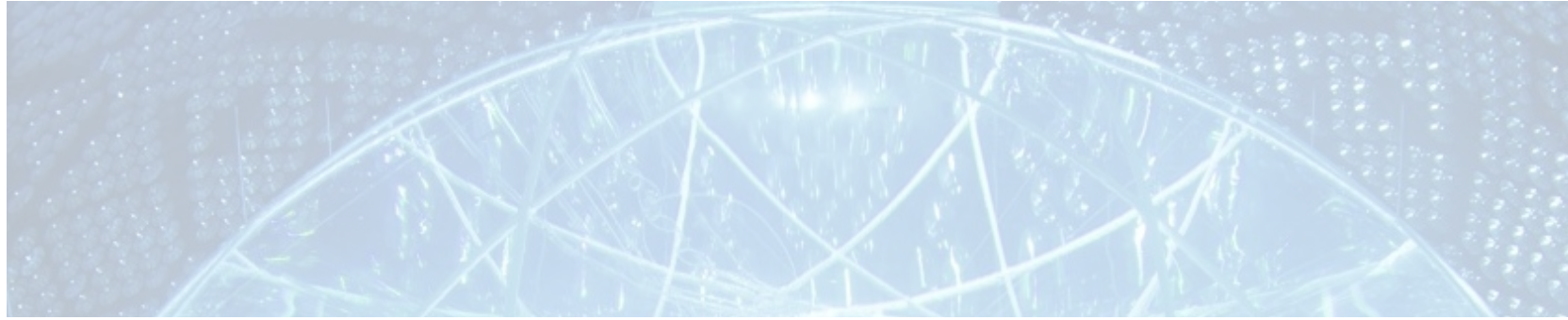
- ✦ Nucleon decay modes into invisible channels
- ✦ Solar neutrinos



Improved background  
reduction

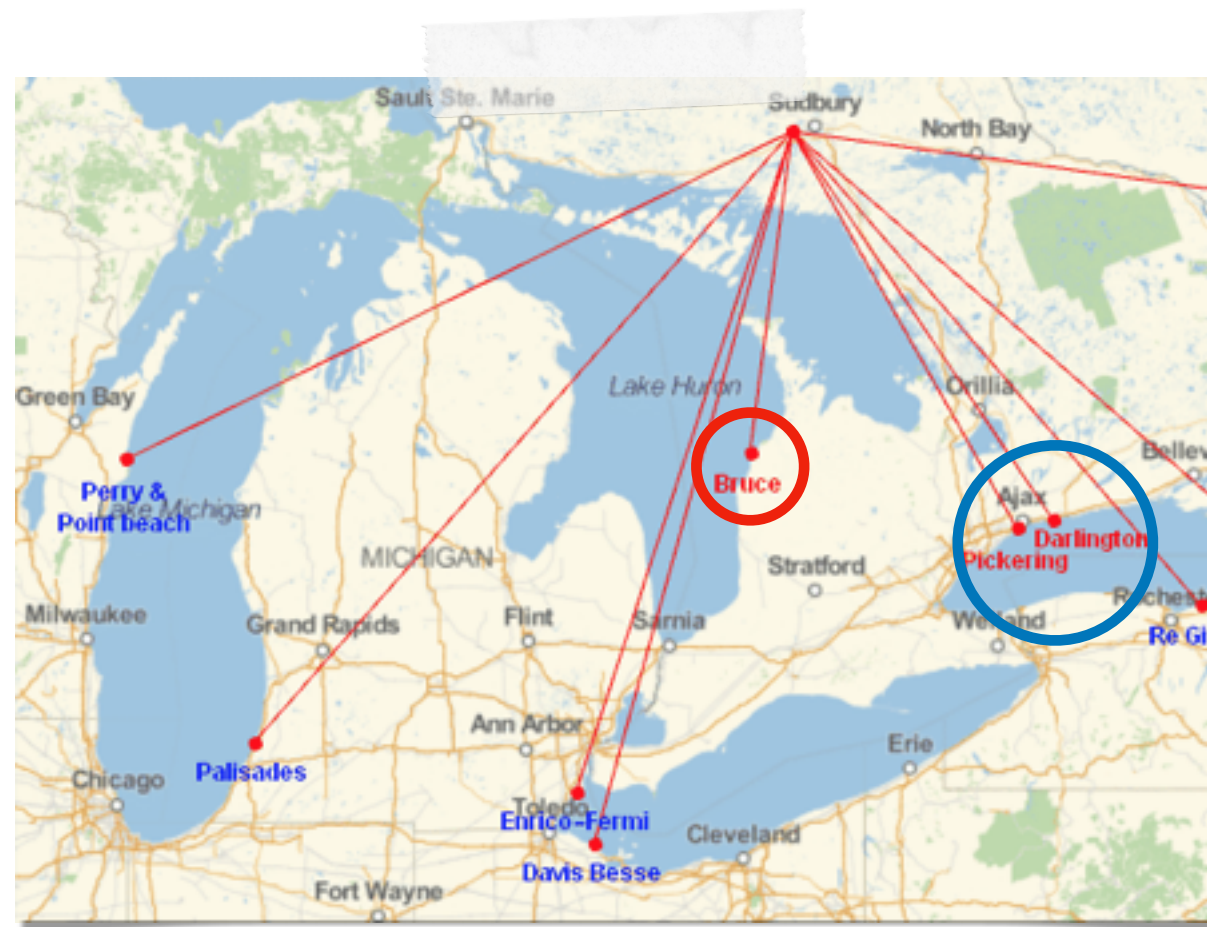
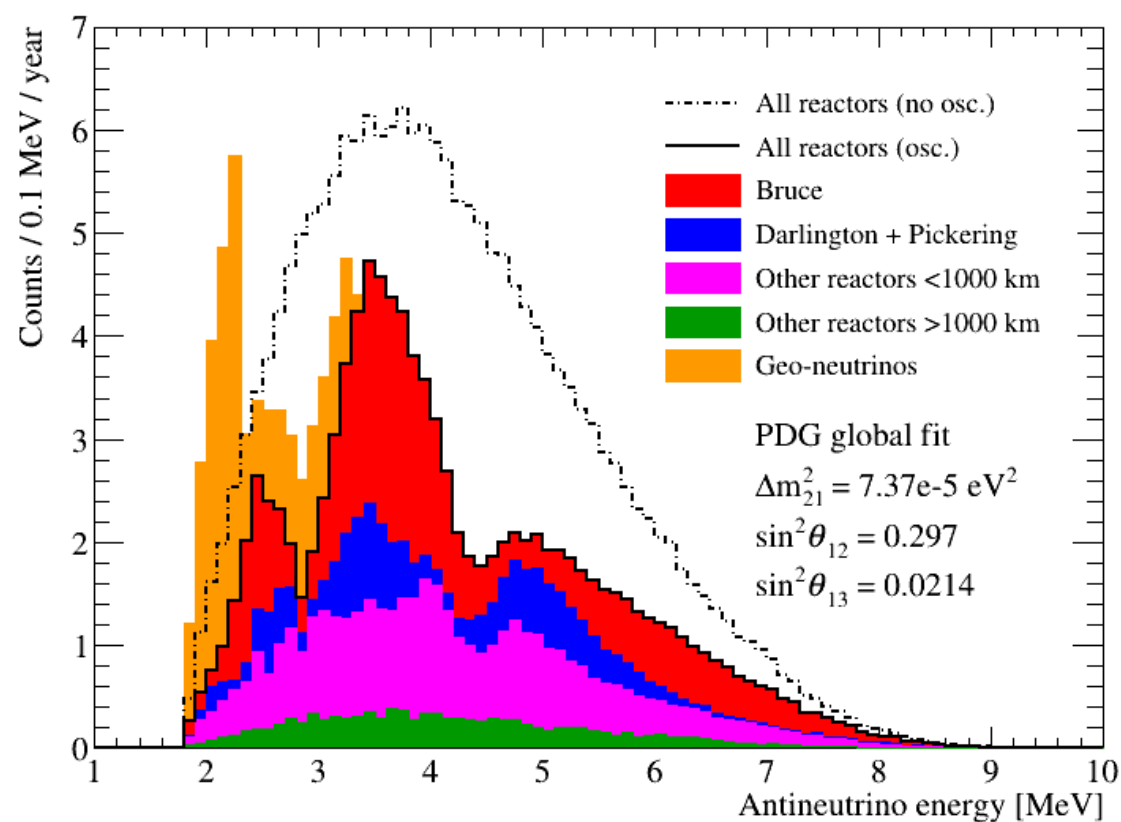






Detection mode = Inverse beta decay

**Reactor anti-neutrinos** = total flux about 20% of KamLAND, but baseline between reactors and SNO+ gives a unique spectral shape distortion



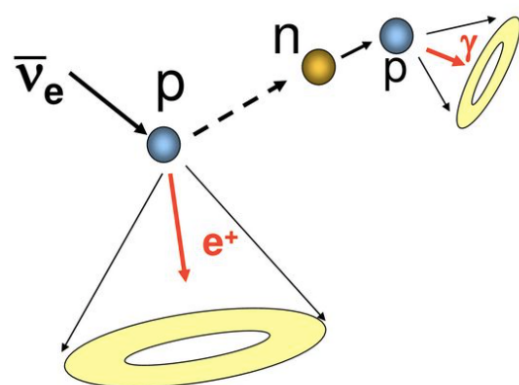
Dashed line: nonoscillated reactor spectrum  
Solid line: geoneutrino spectrum  
**Red:** Bruce reactor at 240km  
**Blue:** Darlington & Pickering reactors at 350km  
**Green+Magenta:** Other reactors



May - December 2017  
(~115 gold physics days)  
First SNO+ water phase

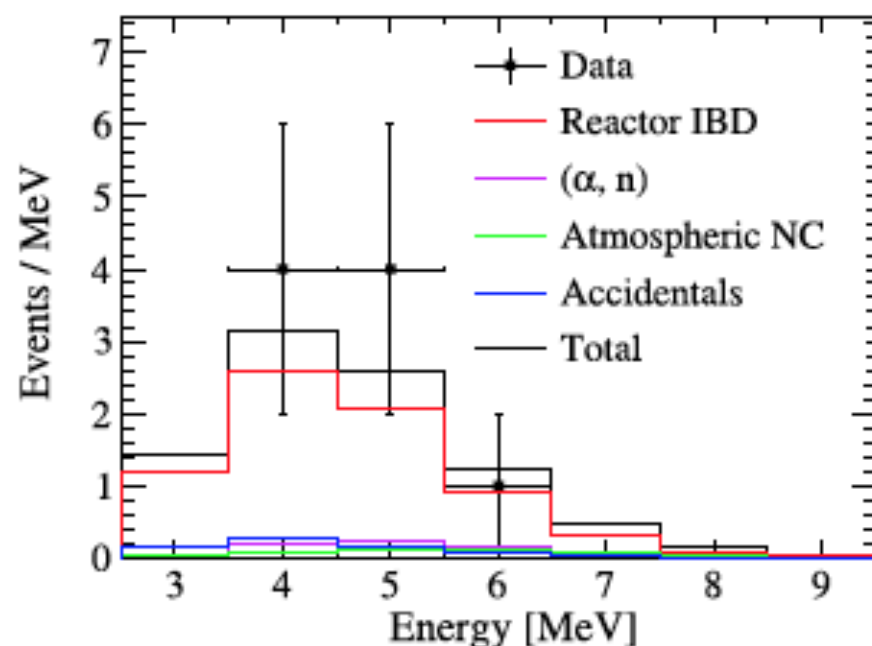
October 2018 - June 2019  
(~185 gold physics days)  
Second SNO+ water phase

Additional Cover gas  
shielding to reduce Rn  
ingresses in water



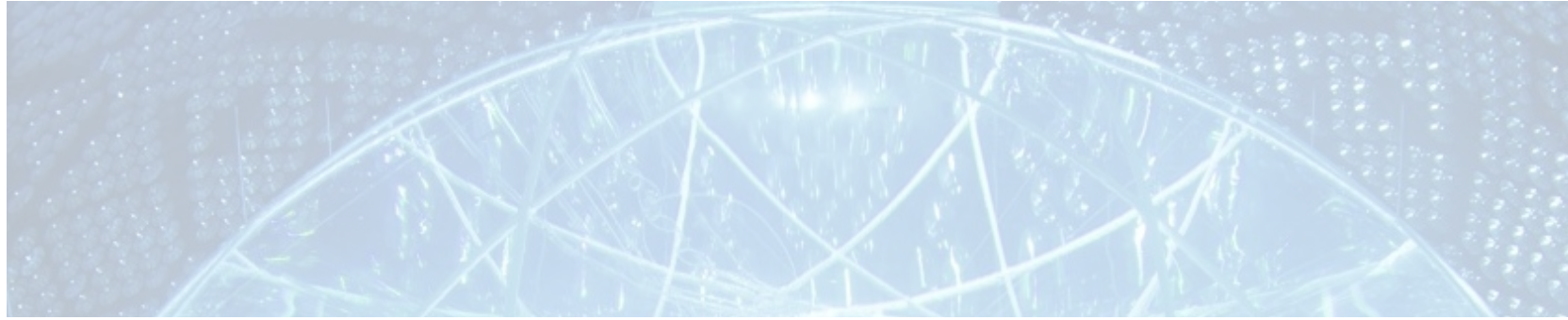
## Major Outcomes

- ❖ Nucleon decay modes into invisible channels
- ❖ Solar neutrinos
- ❖ Reactor antineutrinos
  - ❖ Efficiency for triggering on a neutron:  **$(49.08 \pm 0.39)\%$**  at center *Phys.Rev.C 102, 014002*
  - ❖ **First measurement** of reactor antineutrinos using pure water *Phys.Rev.Lett 130, 091801*  
*PRL Editor's Choice*  
*APS Physics Magazine Highlight*



See talk  
T. Kaptanoglu  
Aug 31, 2023, 3:15PM  
Hörsaal 21 lecture hall

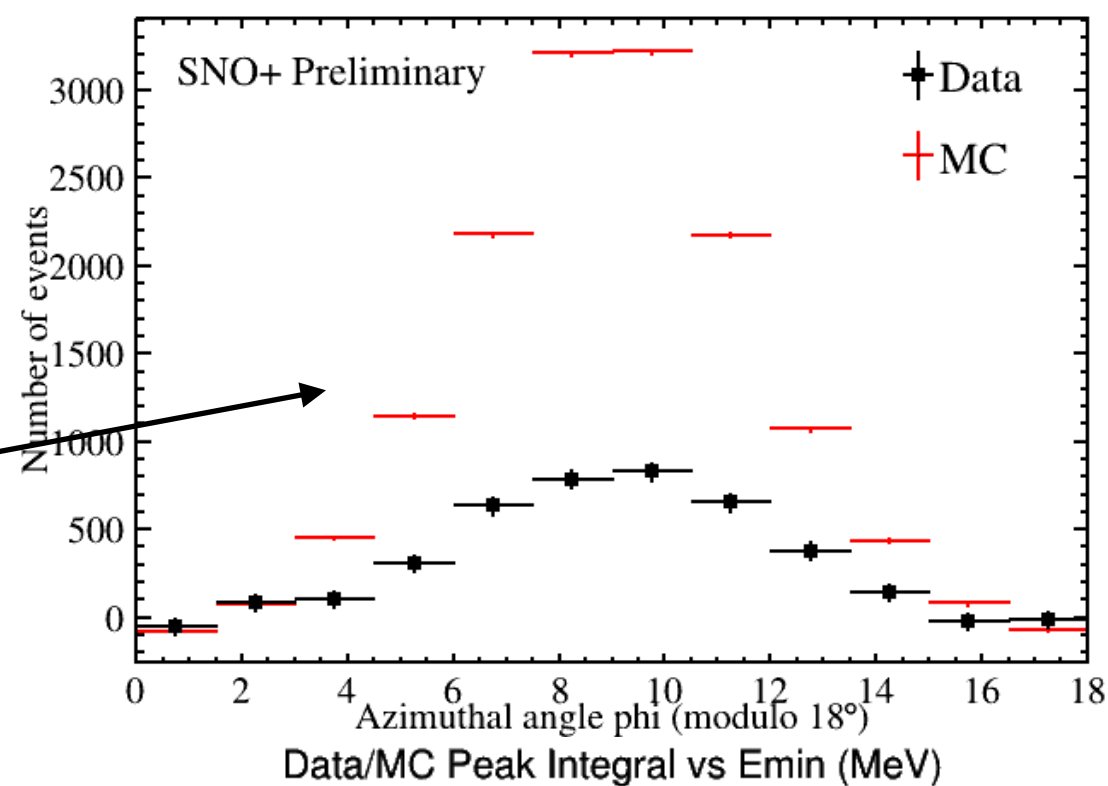
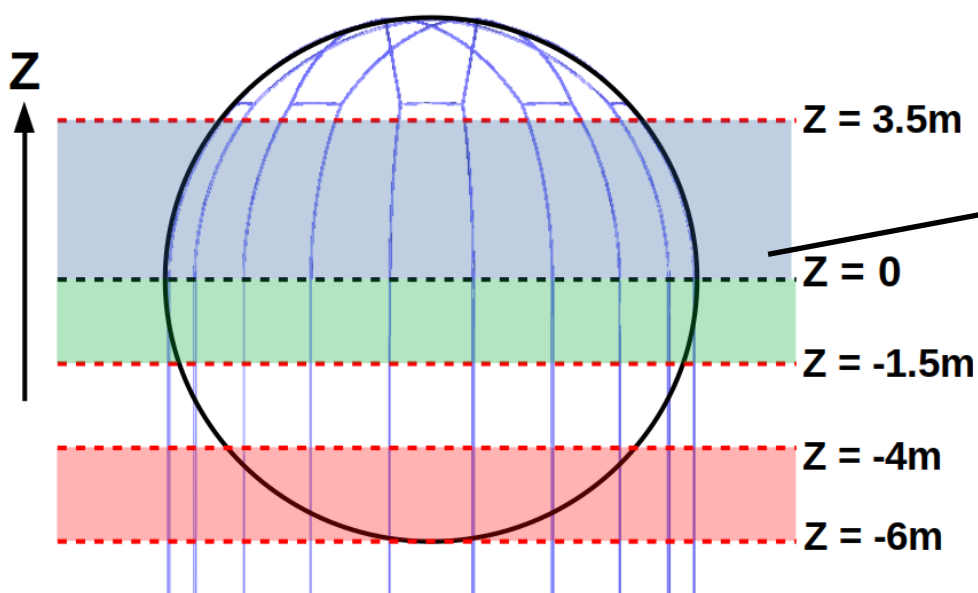




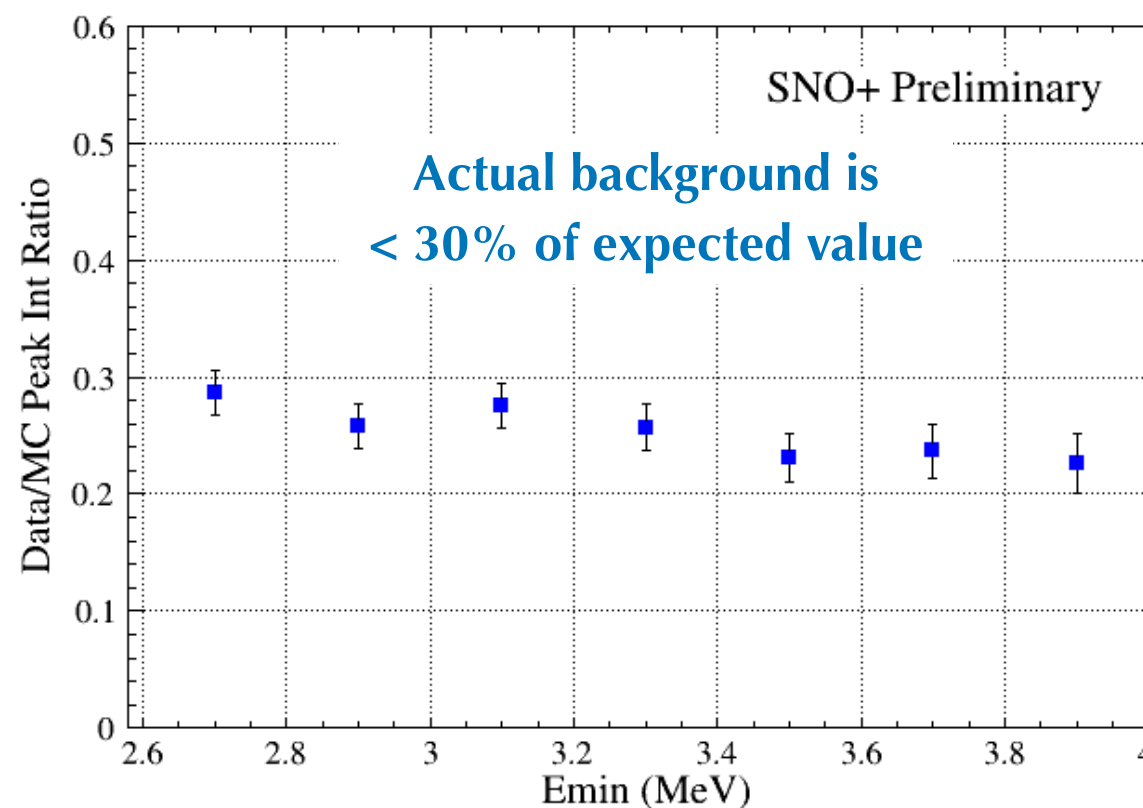
- Measure of the rope's radioactivity.

## Event Selection:

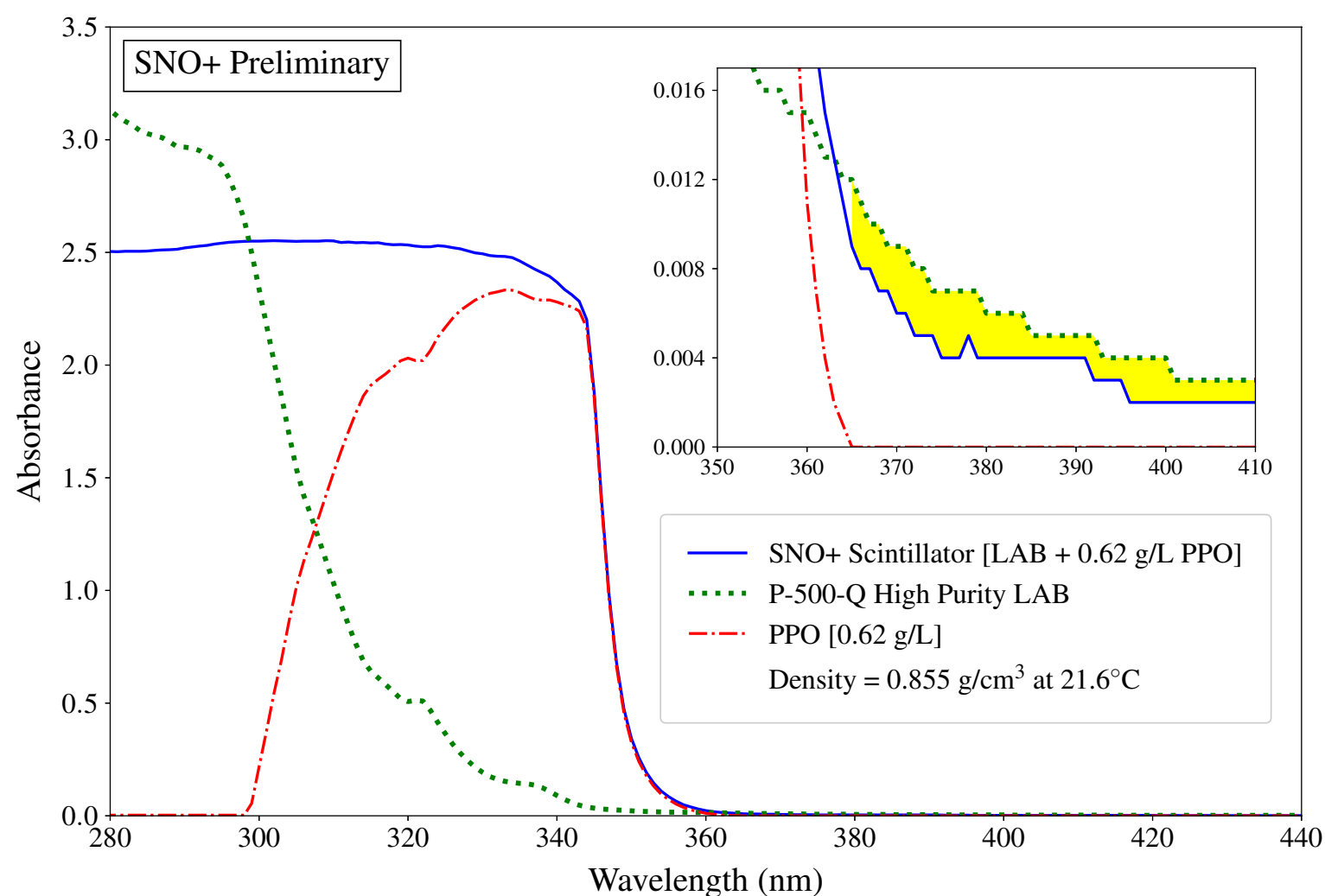
- Box Analysis*



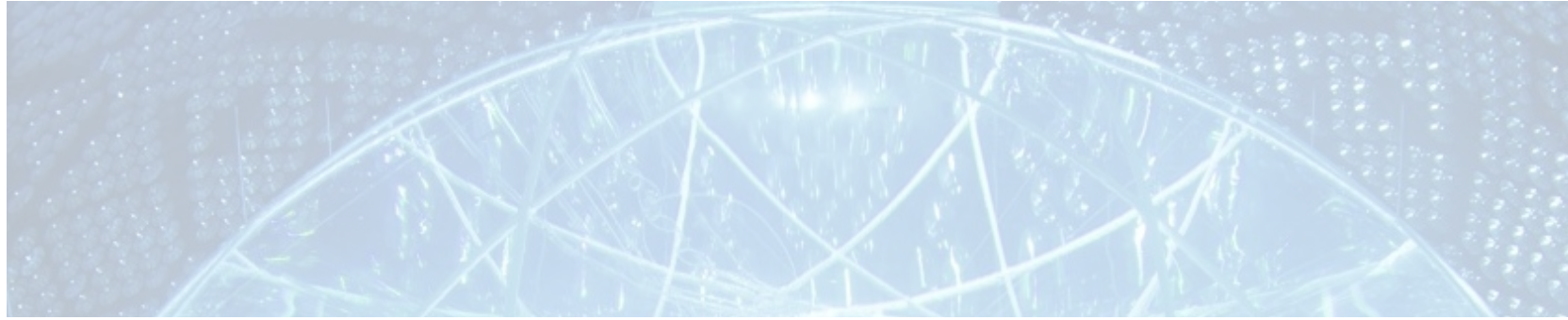
Use the azimuthal dependence ( $\phi$ ) of the ropes  
Other backgrounds are flat in  $\phi$



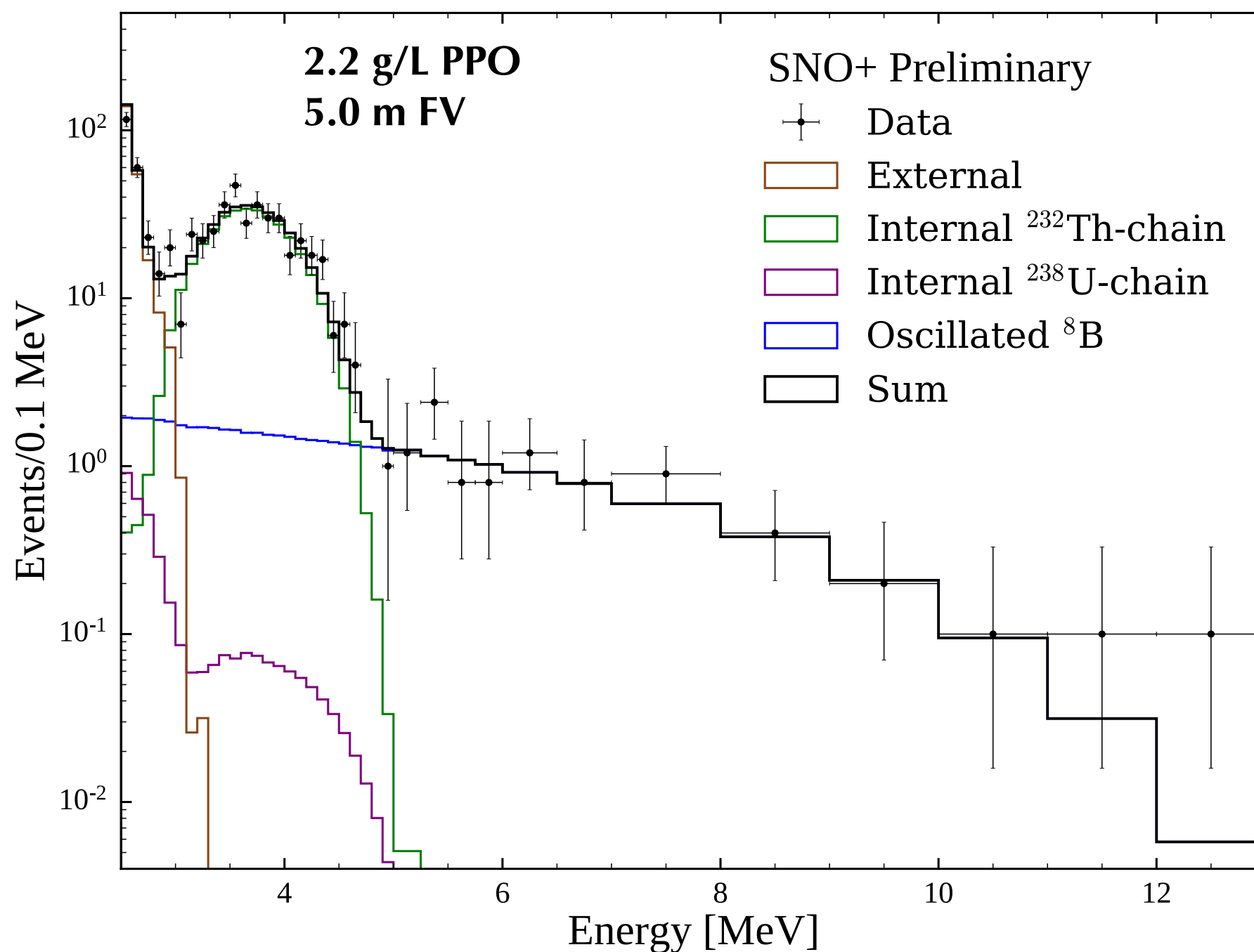
- LAB, Master Solution, and final scintillator assessed for quality hourly during purification plant operation and detector filling
  - Observe excellent clarity above PPO absorption (UV-Vis spectroscopy)

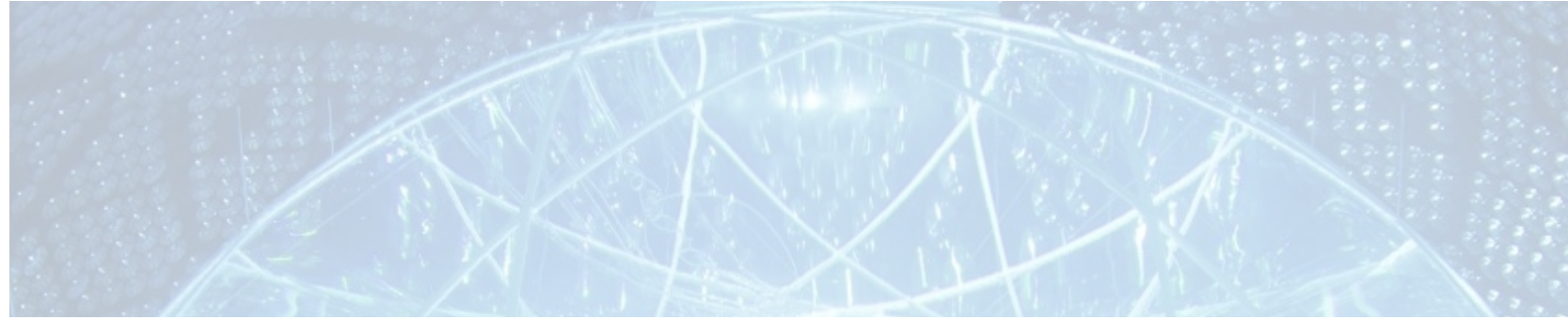






♣ Cross-checked via  $^{208}\text{Tl}$  peak fit = major background for solar  $^8\text{B}$  analysis





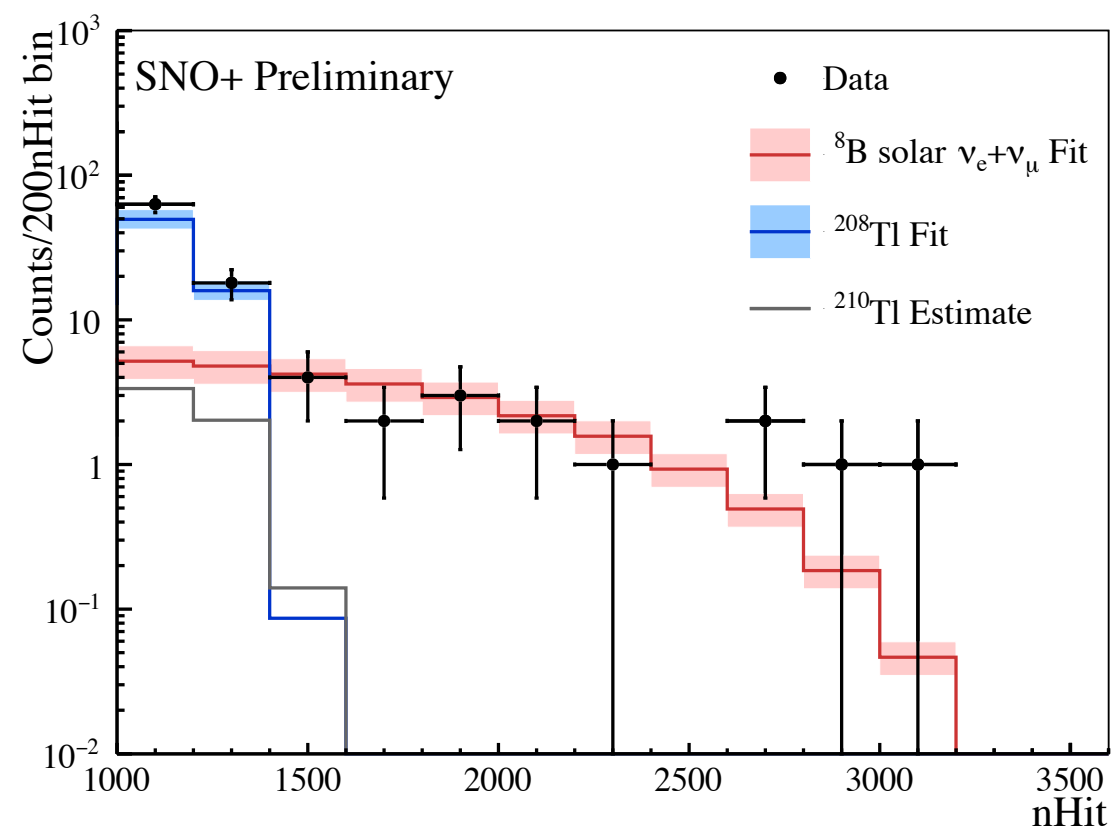
2017 2018 2019 2020 2021 2022 2023

April - October 2020  
(92 gold physics days)

Bonus phase:  
partial fill  
0.6 g/L PPO

## Major Outcomes

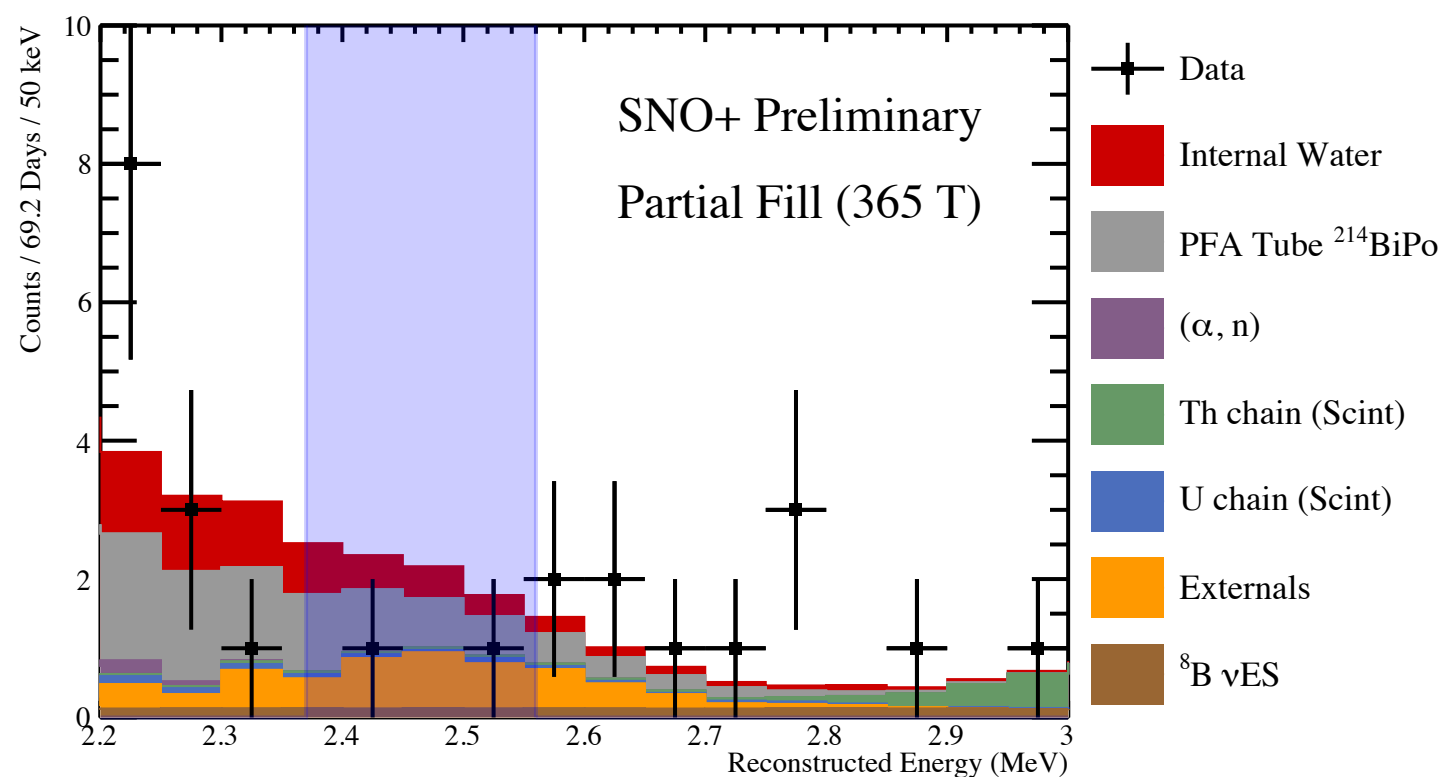
- ❖ Solar neutrinos with 11.2 kt-day during partial fill = PoS-PANIC2021-274



	$\Phi_{8B}$
R = 4.5 m	$6.45 \times 10^6 \text{ }^{+26.2\%}_{-22.4\%} (stat.) \text{ }^{+8.8\%}_{-11.5\%} (syst.) \text{ cm}^{-2} \text{ s}^{-1}$



❖ Preparation for the double-beta decay phase: background and target-out measurement



Background	Expected Counts in Partial Fill ROI
Internal Water	1.8 §
* PFA Tube $^{214}\text{BiPo}$	2.9 §
Externals	2.5
$(\alpha, n)$	0
Th Chain (Scint)	0.1
U Chain (Scint)	0.3
$^8\text{B } \nu\text{ES}$	0.5
<b>Total Backgrounds</b>	<b>8.0</b>

\* tube used to remove water from the detector  
§ not relevant for the full fill phase

June - October 2020  
(~69 gold physics days)

Bonus phase:  
partial fill  
0.6 g/L PPO

## Background breakdown in $0\nu\beta\beta$ ROI

See Poster  
T. Kroupova  
Event Reconstruction in the SNO+ Experiment

### Cosmogenic:

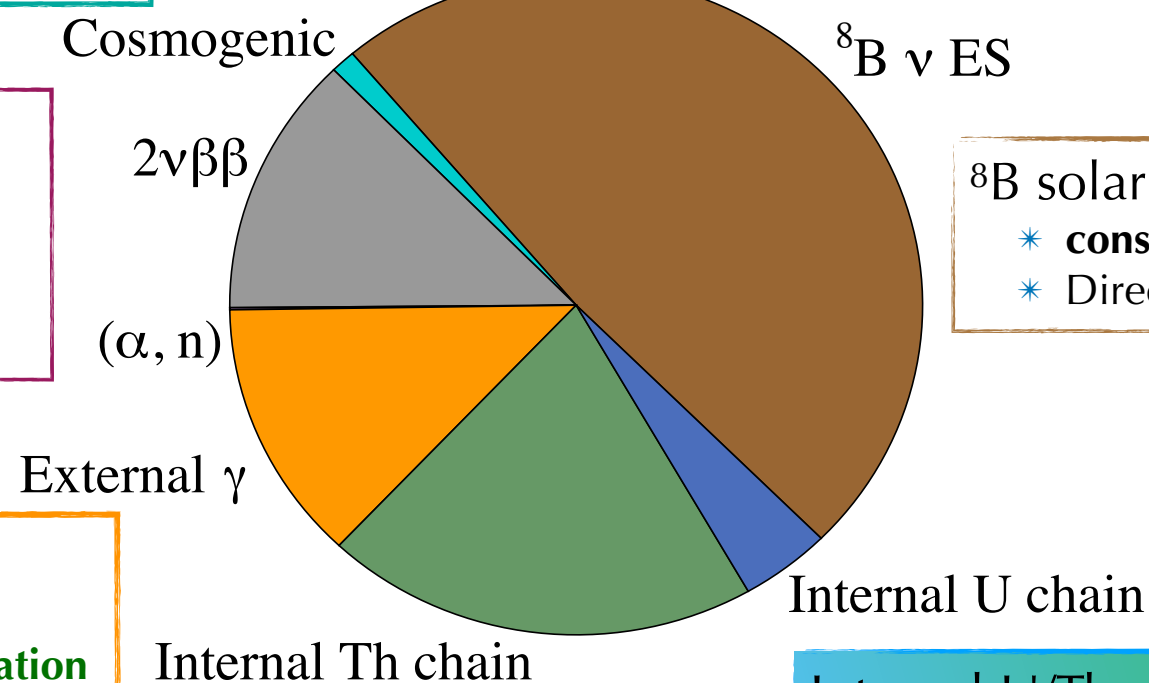
- \*  $^{60}\text{Co}$ ,  $^{110\text{m}}\text{Ag}$ ,  $^{88}\text{Y}$ ,  $^{22}\text{Na}$
- \* mitigation: purification + “cool-down” UG
- \* First Te UG since 2015
- \* multi-site classifier

### (a,n):

- \* alpha-capture on  $^{13}\text{C}/^{18}\text{O}$
- \* **Measure the contribution from LS components before Te loading**
- \* delayed coincidence tagging

### $^8\text{B}$ solar neutrinos:

- \* **constrained by SNO/SK data**
- \* Directionality?



### Internal U/Th chain:

- \*  $^{214}\text{BiPo}$ ,  $^{212}\text{BiPo}$ ,  $^{210}\text{Ti}$
- \* **LAB components below target for the TeLS phase**
- \* **Constantly monitoring the contribution from the scintillator before adding Te**
- \*  $\beta - \alpha$  delayed coincidence tagging + in-window rejection

### Optimised FV and ROI

	$T_{1/2}$ [yr]	$m_{0\nu\beta\beta}$ [meV]
0.5% Te, 5 yr	$2.1 \times 10^{26}$	37 - 89



- Increasing the amount of isotope increases the signal
  - ▶  $^8\text{B}$ -nu solar background (main) remains the same
  - ▶ Improved loading scheme maintains acceptable light yield despite increased absorption
  - ▶ Samples with several % loading have been stable on timescales of years.
  - ▶ Incremental cost  $\sim \$2\text{M} / \text{tonne Te}$

