

Observation of low-lying isomeric states in ¹³⁶Cs: a new avenue for dark matter and solar neutrino detection in xenon detectors





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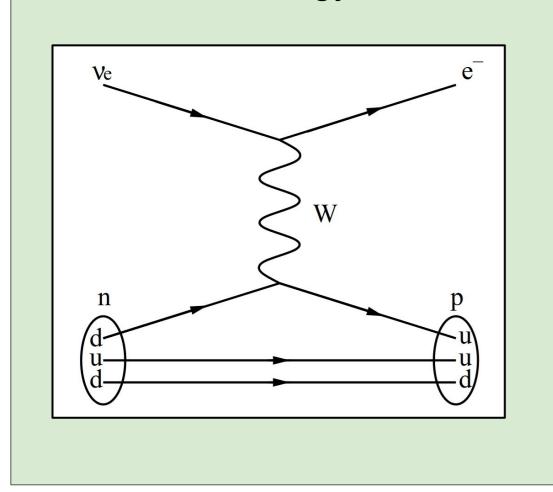


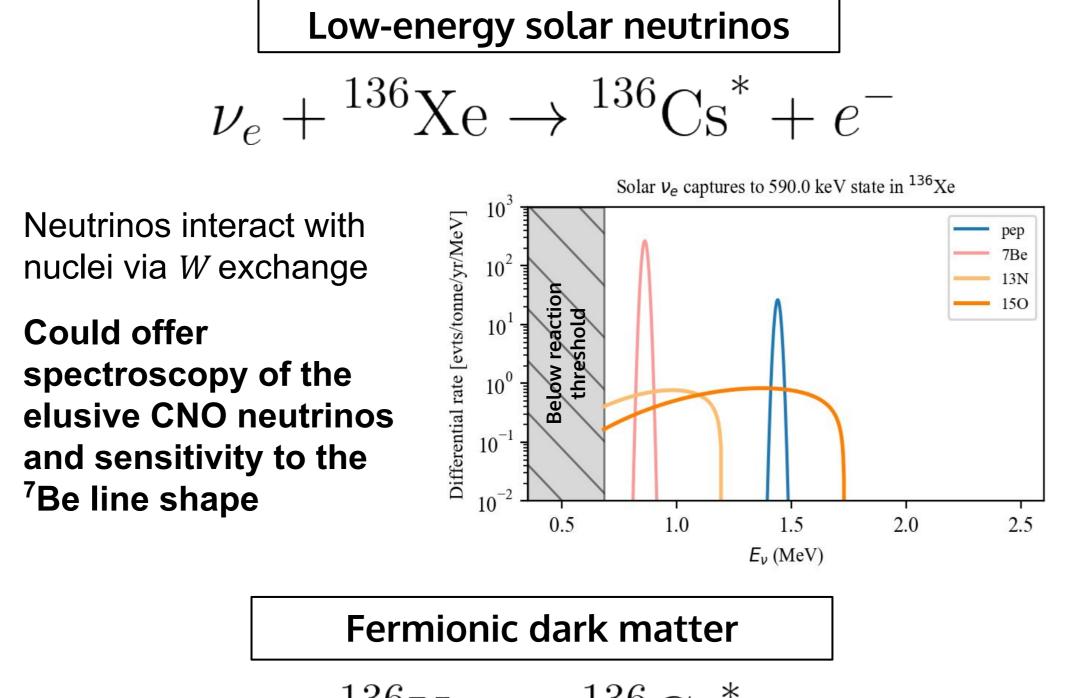
Neutrino / dark matter "capture" on ¹³⁶Xe

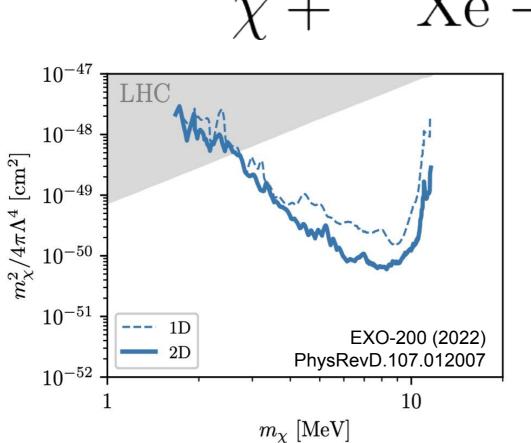
An unexplored channel in Xe detectors

¹³⁶Xe is a promising target for the detection of neutral lepton charged current (CC) interactions on nuclei

- 136Xe is widely deployed in existing and future underground experiments
- $Q = M_{Cs} M_{Xe} = 90.3 \text{ keV}$ \rightarrow Low energy threshold!





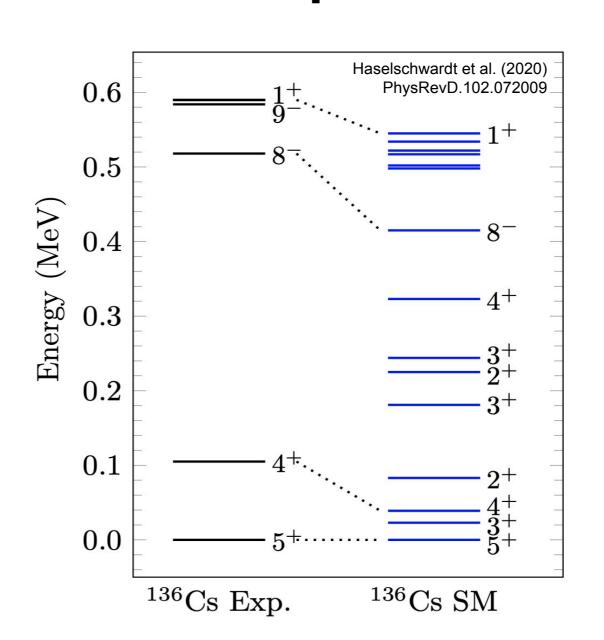


Fermionic dark matter $\chi + {}^{136}\mathrm{Xe} \to {}^{136}\mathrm{Cs}^* + e^-$ Dark matter (DM) particle

interacts via a W'Xenon-based experiments offer world-leading sensitivity to this signal for $M_\chi \sim O(\text{MeV})$

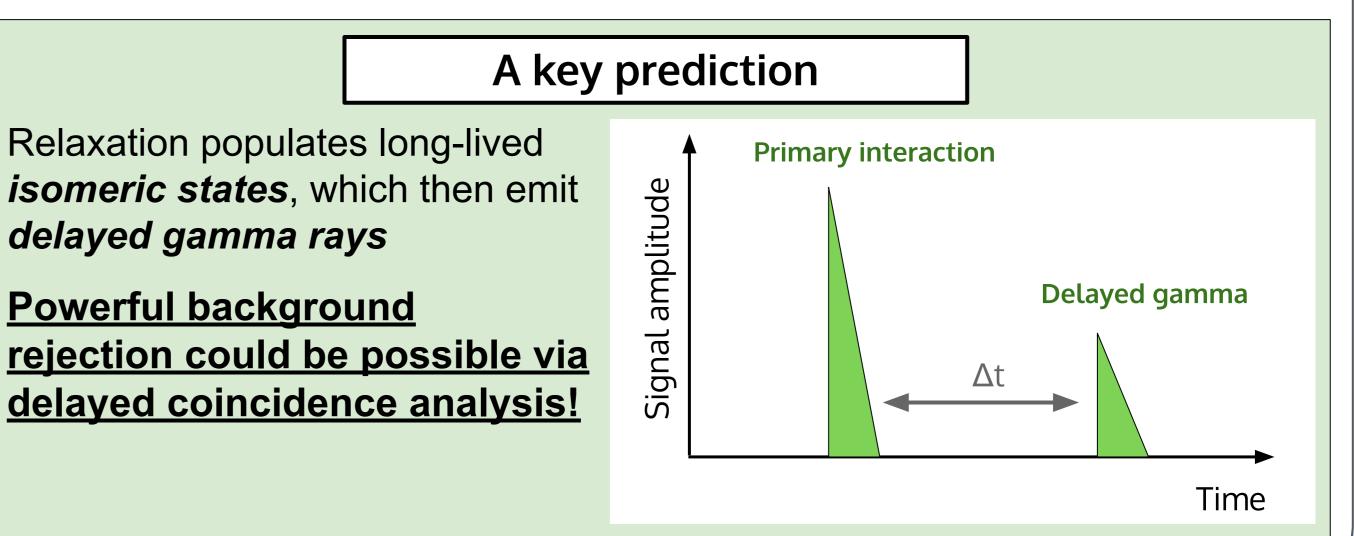
carries lepton number and

The importance of ¹³⁶Cs nuclear structure



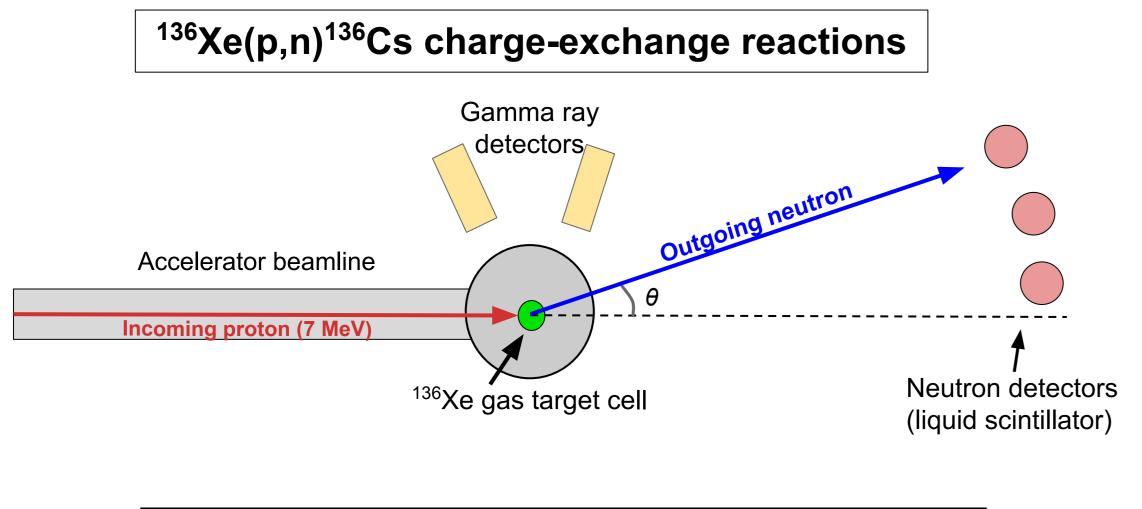
- CC interactions would populate $J^{\pi} = 1^{+}$ excited states in ^{136}Cs
- Relaxation gamma rays need to be accounted for in modeling these events
- Nuclear data for ¹³⁶Cs is sparse, but nuclear shell model calculations predict a complex level scheme

New nuclear structure data is needed!

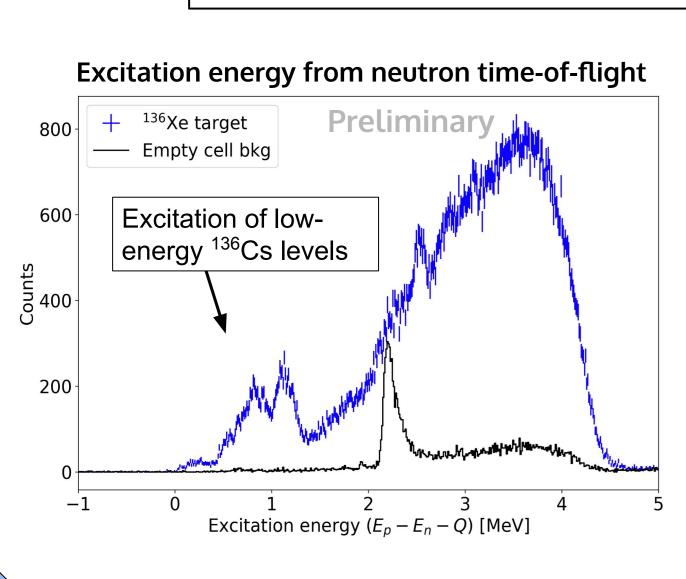


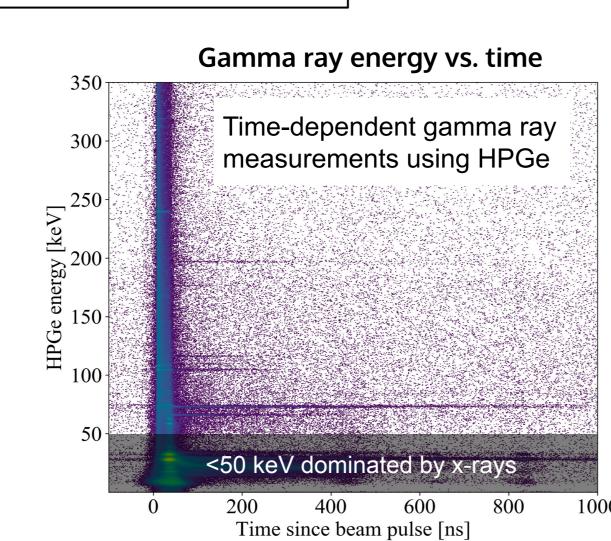
Measuring ¹³⁶Cs at Triangle Universities Nuclear Laboratory (TUNL)

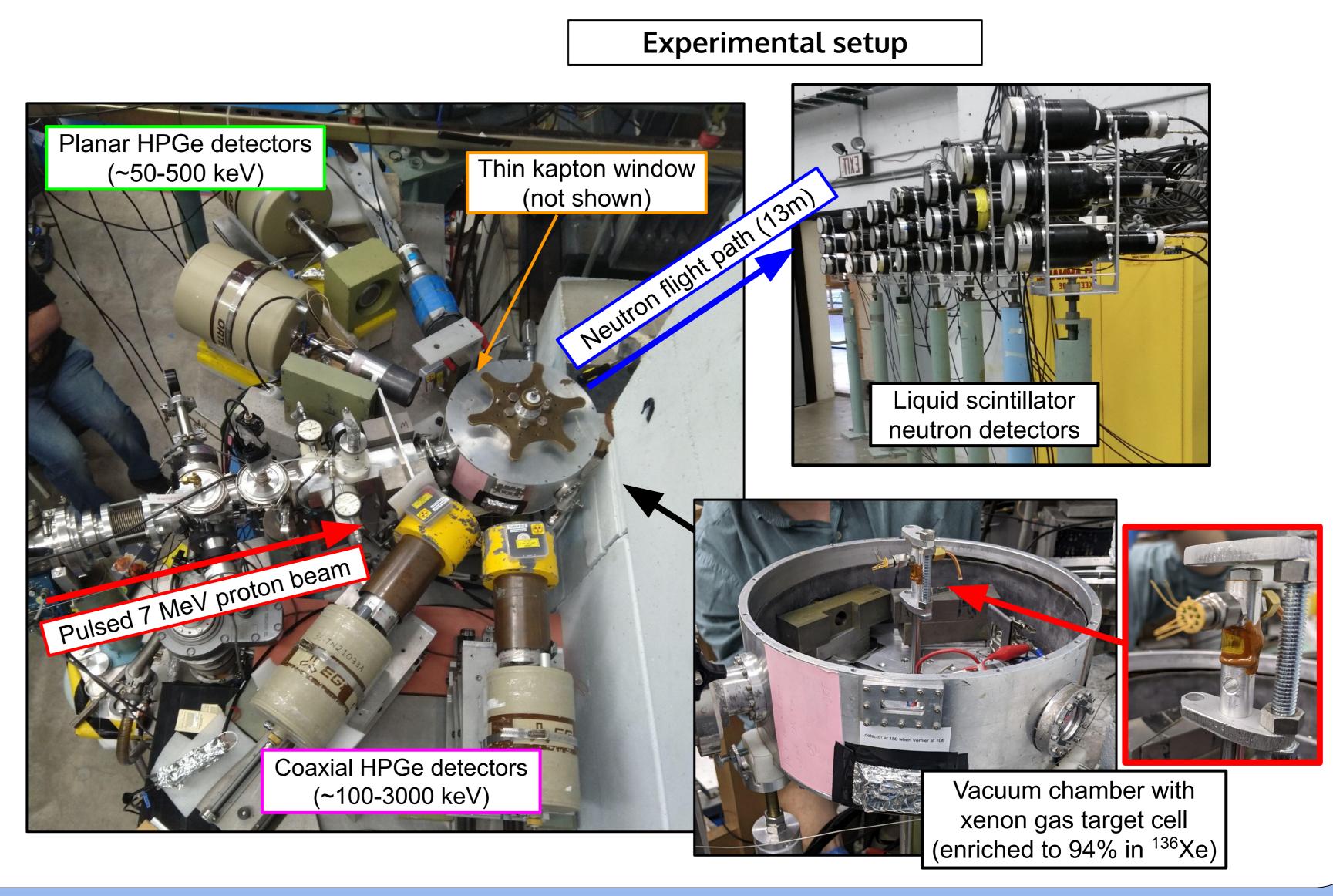




¹³⁶Cs excitation and gamma ray emission





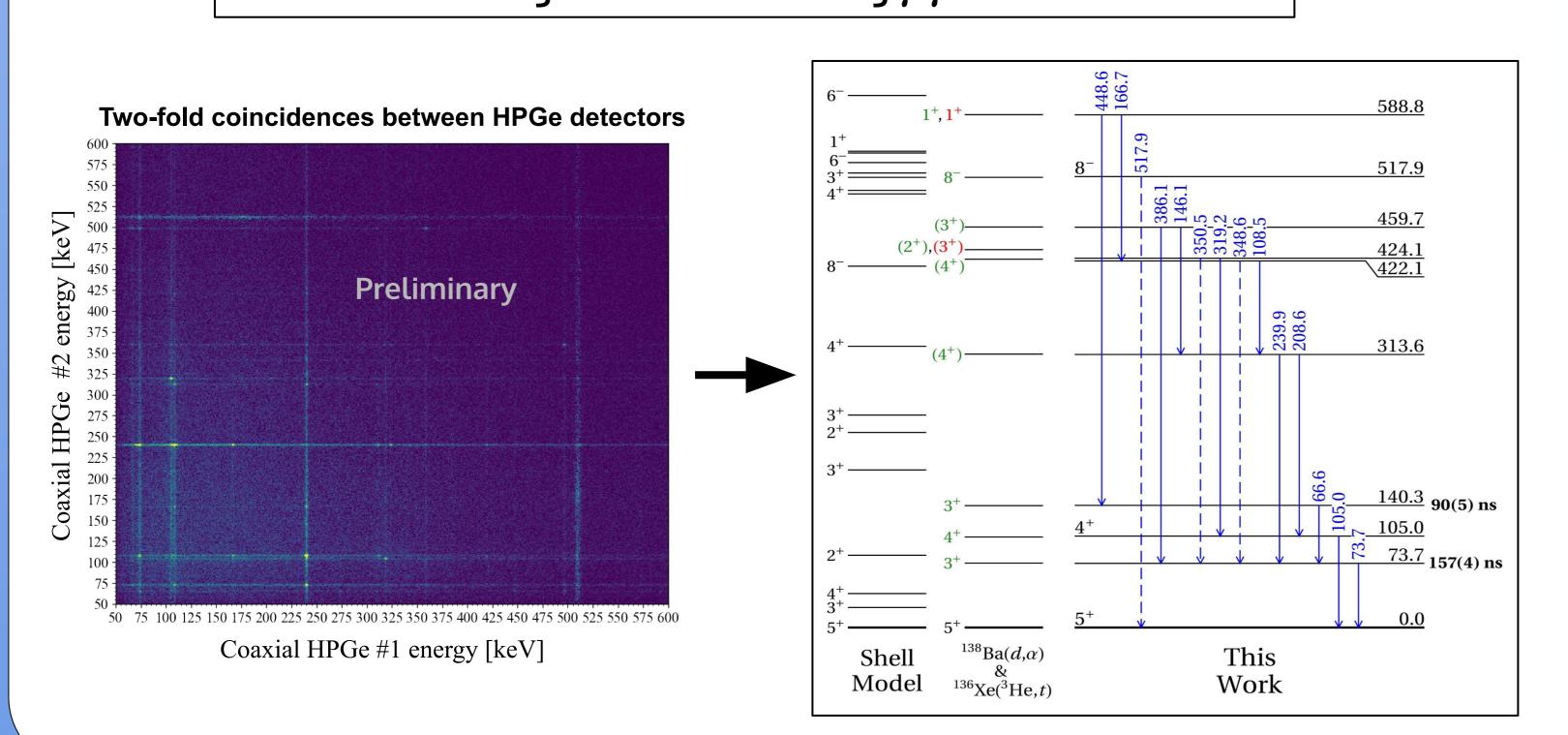


Analysis and results

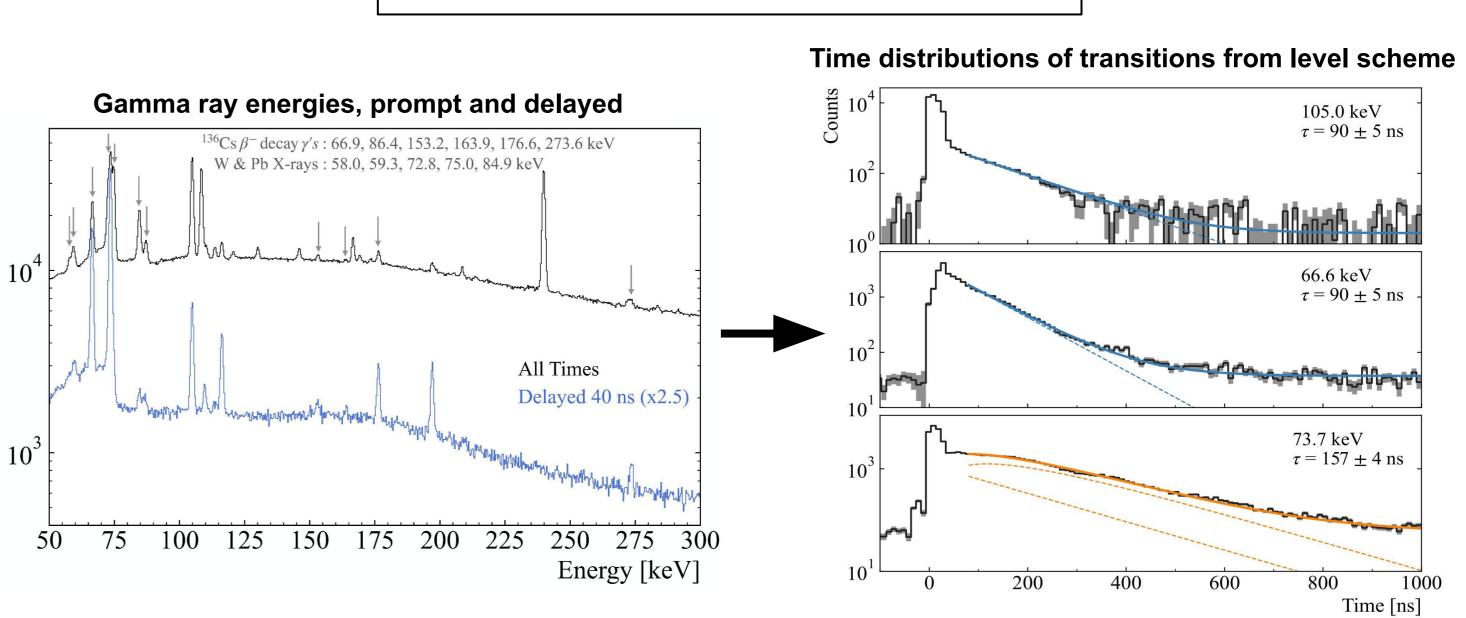
Reconstructed level scheme (below) agrees with charged-particle scattering measurements (PRL.131.052501, PRC.84.051305), but we offer **improved energy resolution** and **gamma ray emission branching fractions**

Combining this with timing information (right), we identify **two isomeric states** in the level scheme: 140.3 keV (τ = 157±4 ns) and 73.7 keV (τ = 90±5 ns)

Constructing level scheme using y-y coincidences



Measurement of lifetimes



Gamma ray emissions from lowest-lying 1⁺ states in ¹³⁶Cs have been measured for the first time

→ Enables robust signal modeling for CC interactions on ¹³⁶Xe
Two isomeric states with lifetimes of O(100) ns have been

Enables powerful background rejection for the detection of solar v's and fermionic DM in current and next-gen Xe experiments

For more info see: Haselschwardt et al., Phys. Rev. Lett. 131 052502 (2023)

identified and measured