



**SAPIENZA**  
UNIVERSITÀ DI ROMA



Istituto Nazionale di Fisica Nucleare

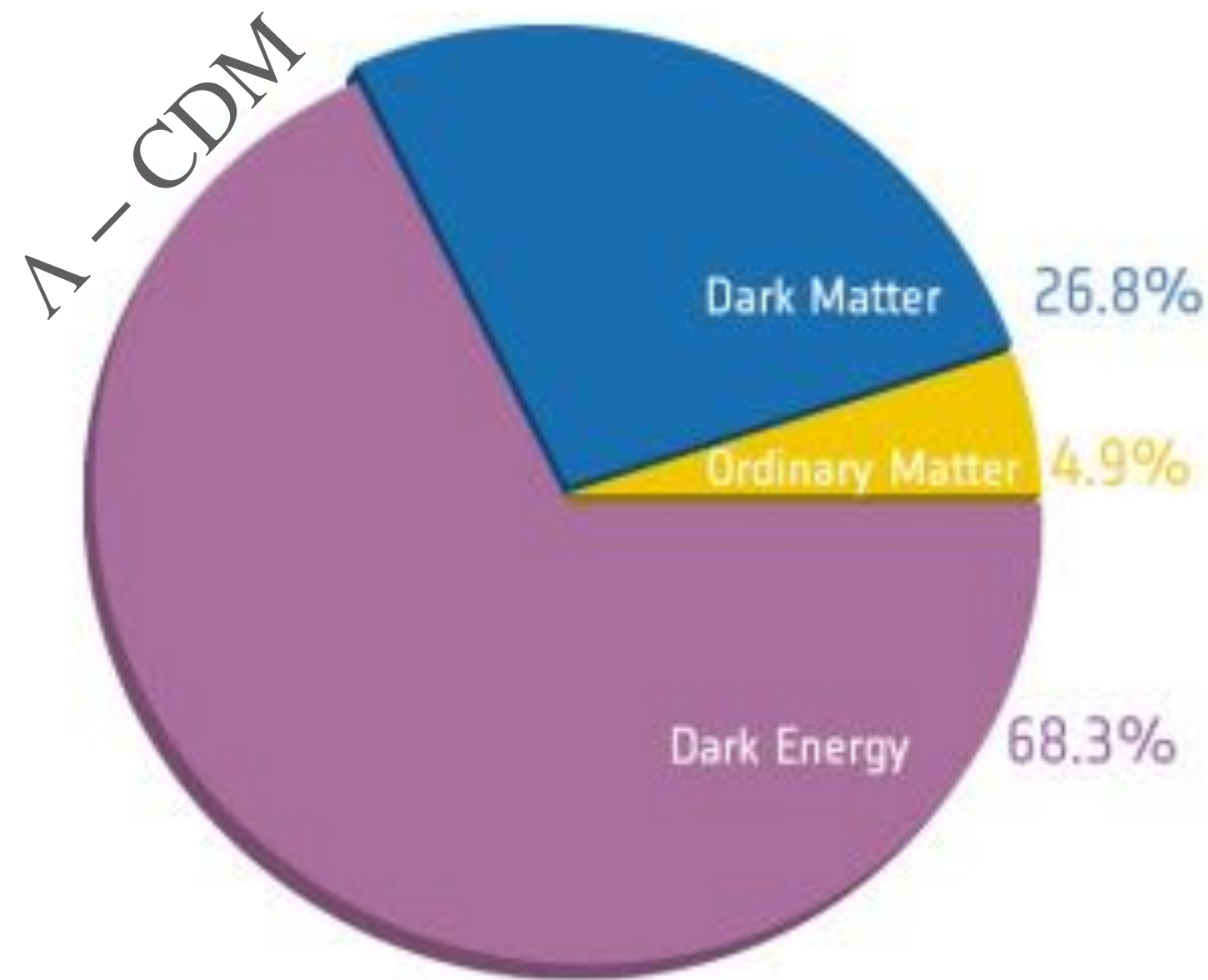
# New results from BULLKID

**Marco Vignati, on behalf of the BULLKID coll.  
Taup, Vienna, 31 August 2023**



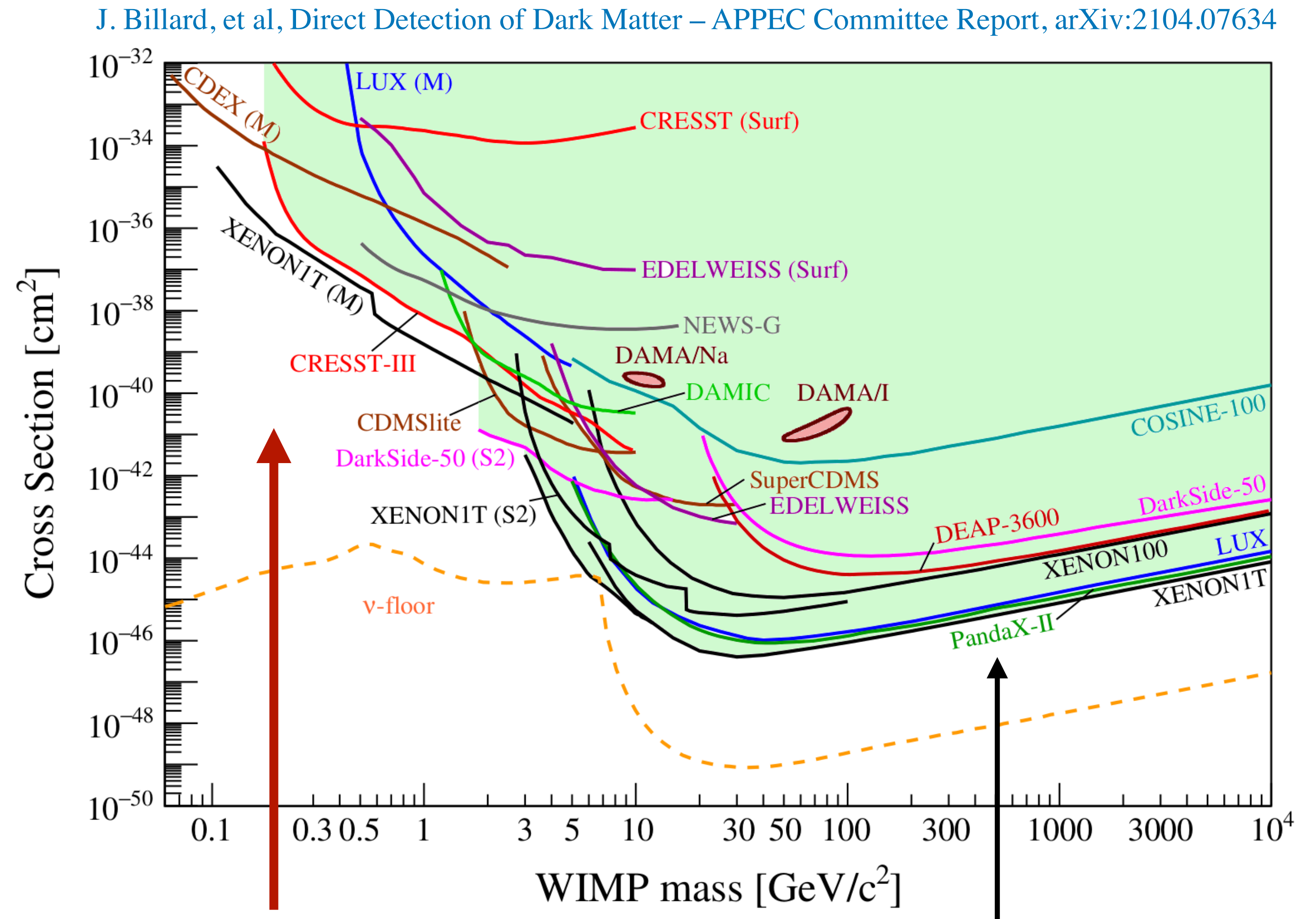


# Dark Matter - direct search

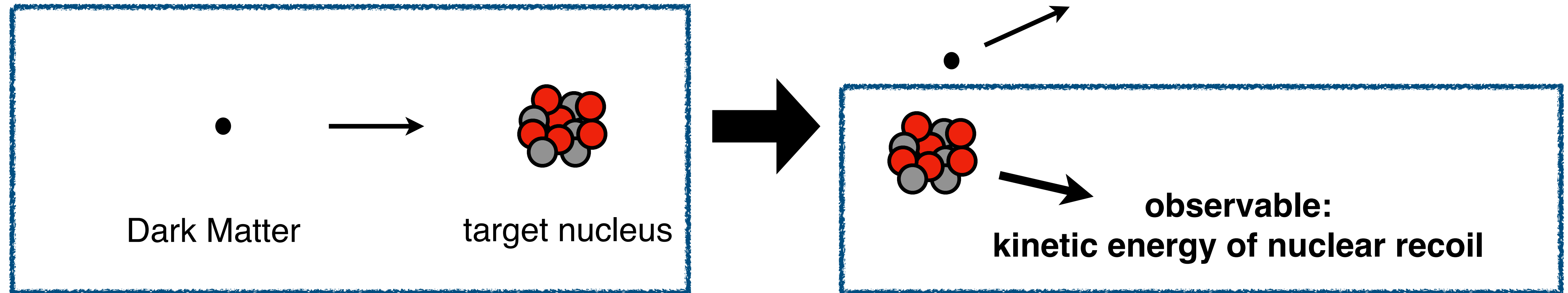


What is the Dark Matter made of?

- primordial black holes?
- $\mu\text{eV}/c^2$  -  $\text{eV}/c^2$  axion-like waves?
- $\text{MeV}/c^2$  -  $\text{TeV}/c^2$  WIMP-like particles?



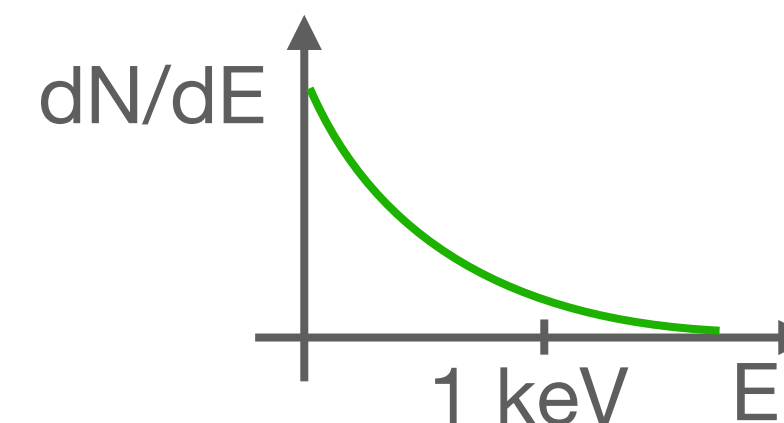
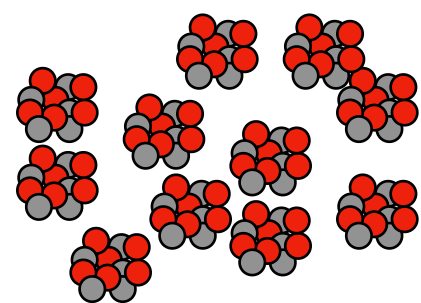
# Direct dark matter search below 1 GeV/c<sup>2</sup>



cross section  $\sigma < 10^{-40} \text{ cm}^2$

energy  $< 1 \text{ keV}$

large number of targets  
 $O(1 \text{ kg})$



low-energy threshold  
 $O(100 \text{ eV})$

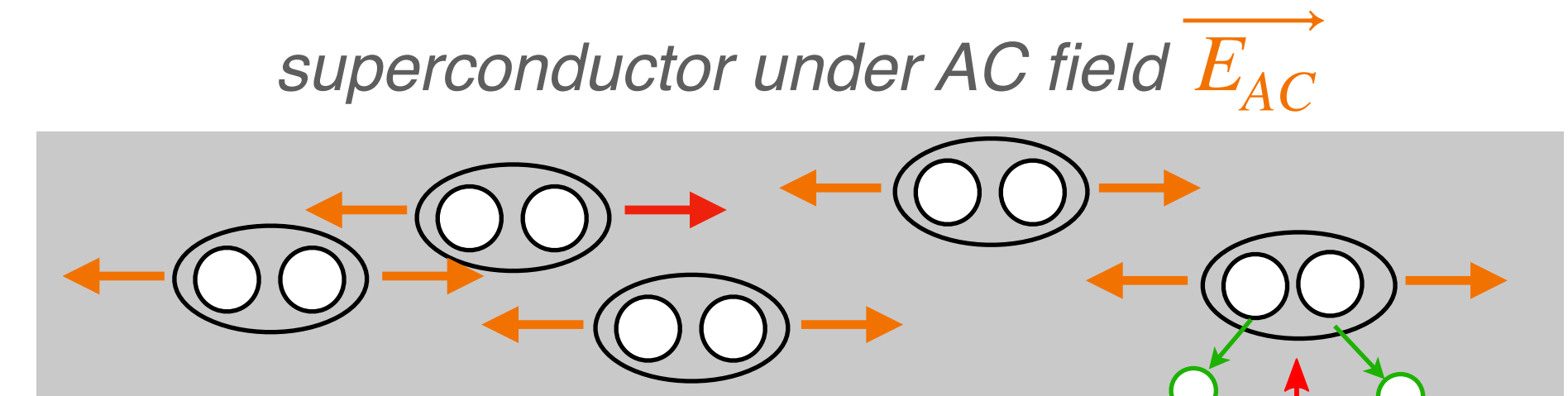
**Difficult with Low-T detectors**

**Motivation for Low-T detectors**

# Kinetic Inductance Detectors

## AC superconductivity

- Electrons bound into Cooper pairs (no dissipation)
- High quality factors ( $Q \sim 10^4 - 10^6$ )
- Inertia from the mass of pairs (*kinetic inductance*,  $L_k$ )



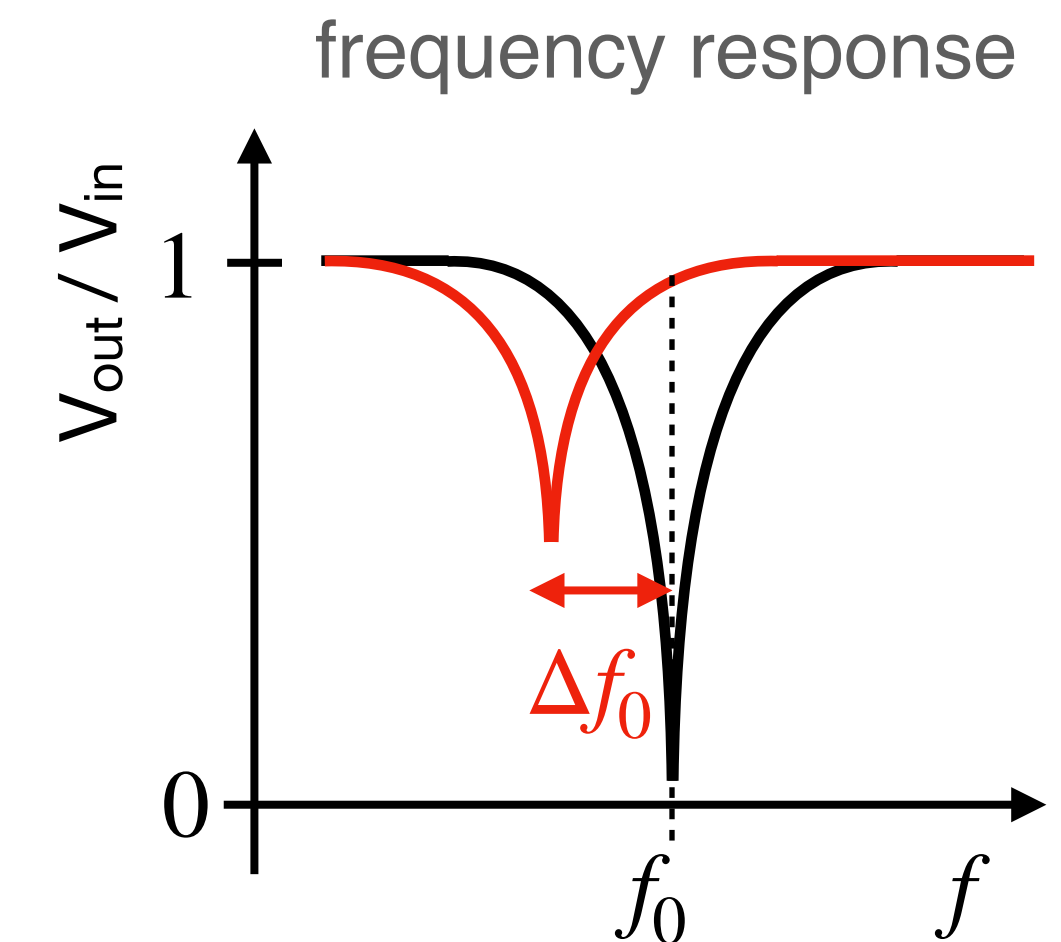
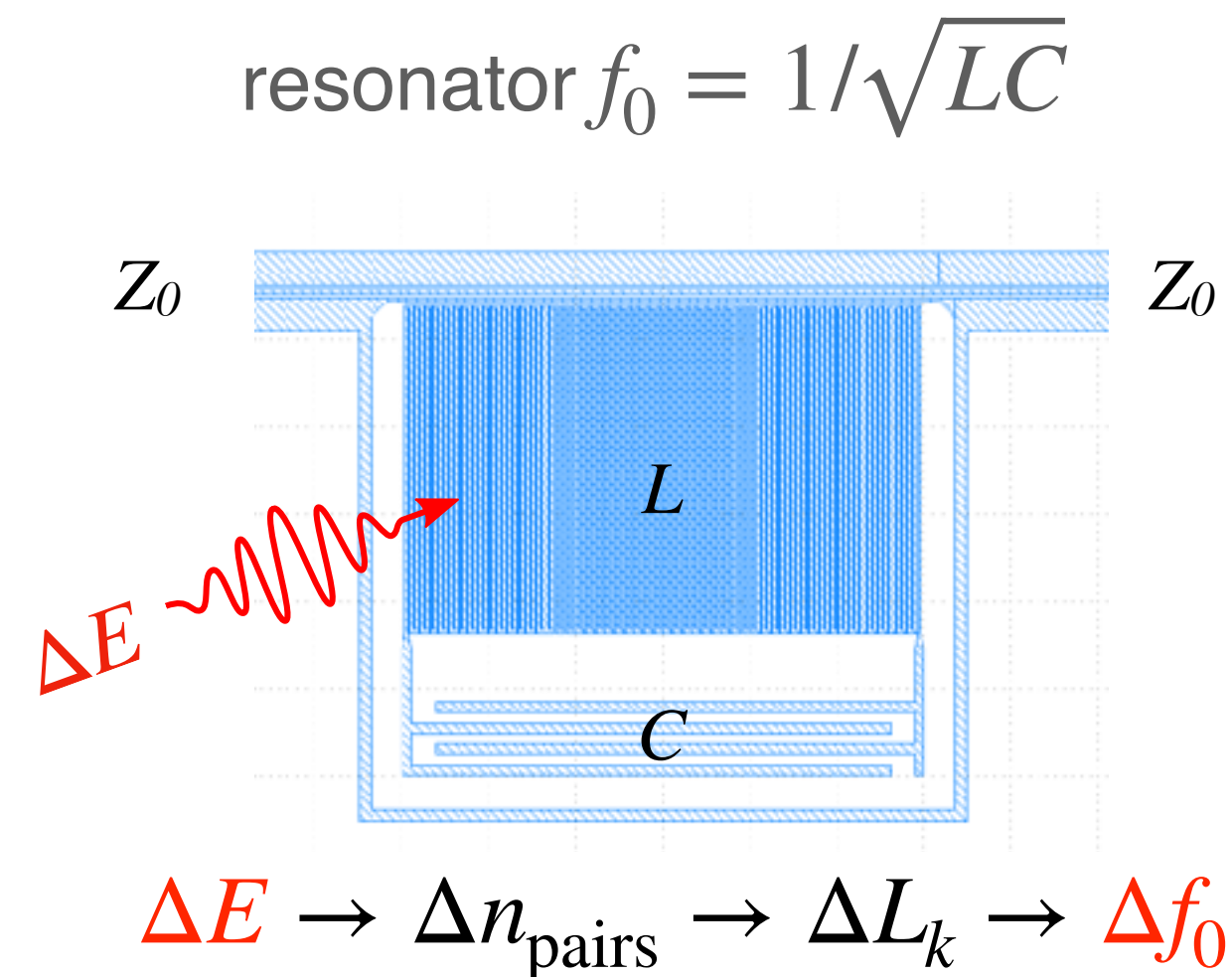
$$L_k = \frac{m_e}{2 e^2 n_{\text{pairs}}}$$

photon or phonon  
absorption  
 $\Delta n_{\text{pairs}} \rightarrow \Delta L_k$

## Kinetic Inductance Detector (KID)

- Superconductor at  $T < 200$  mK (Al)
- LC resonator

Invented by J. Zmuidzinas and his group at Caltech in 2003 for astrophysical applications



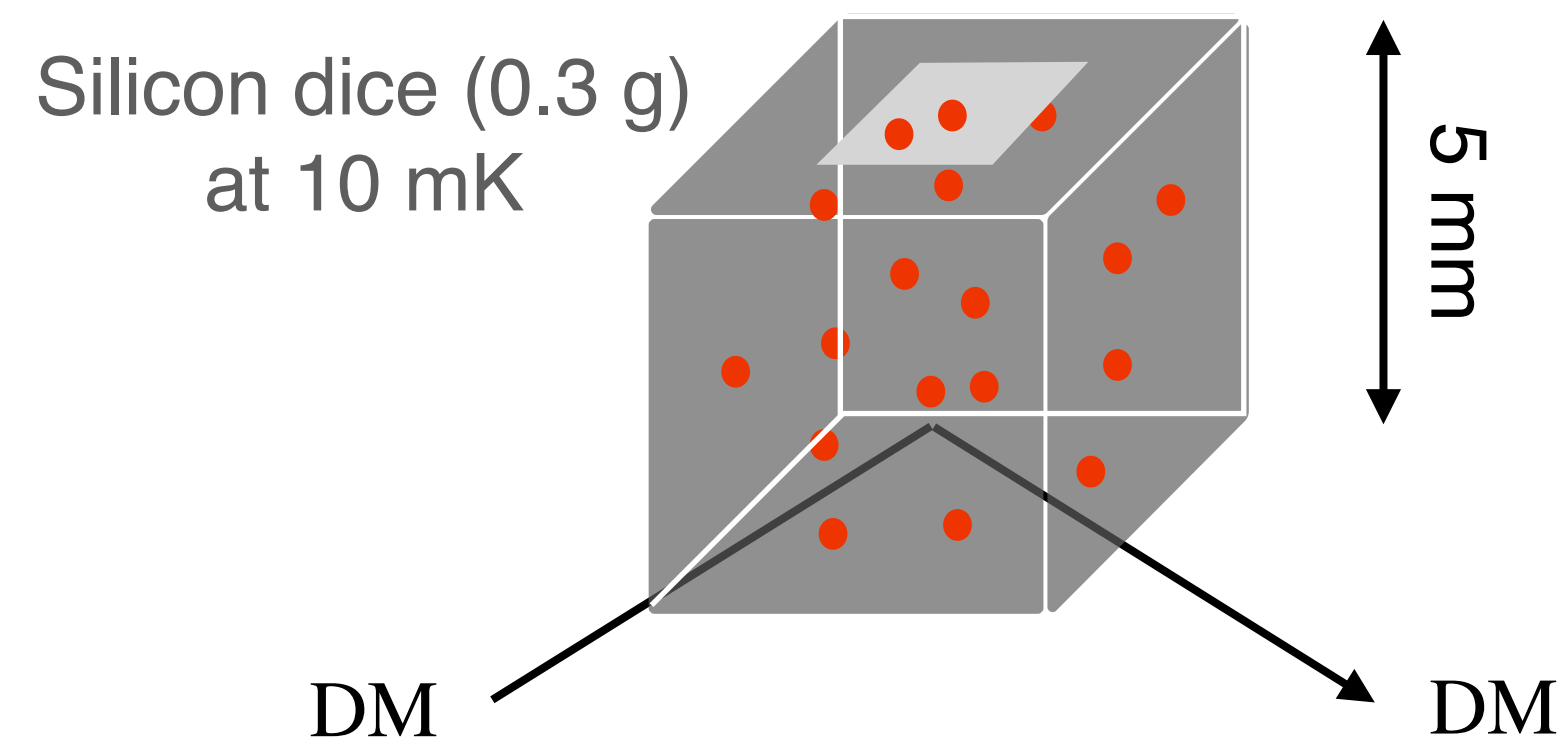


# The BULLKID phonon detector array

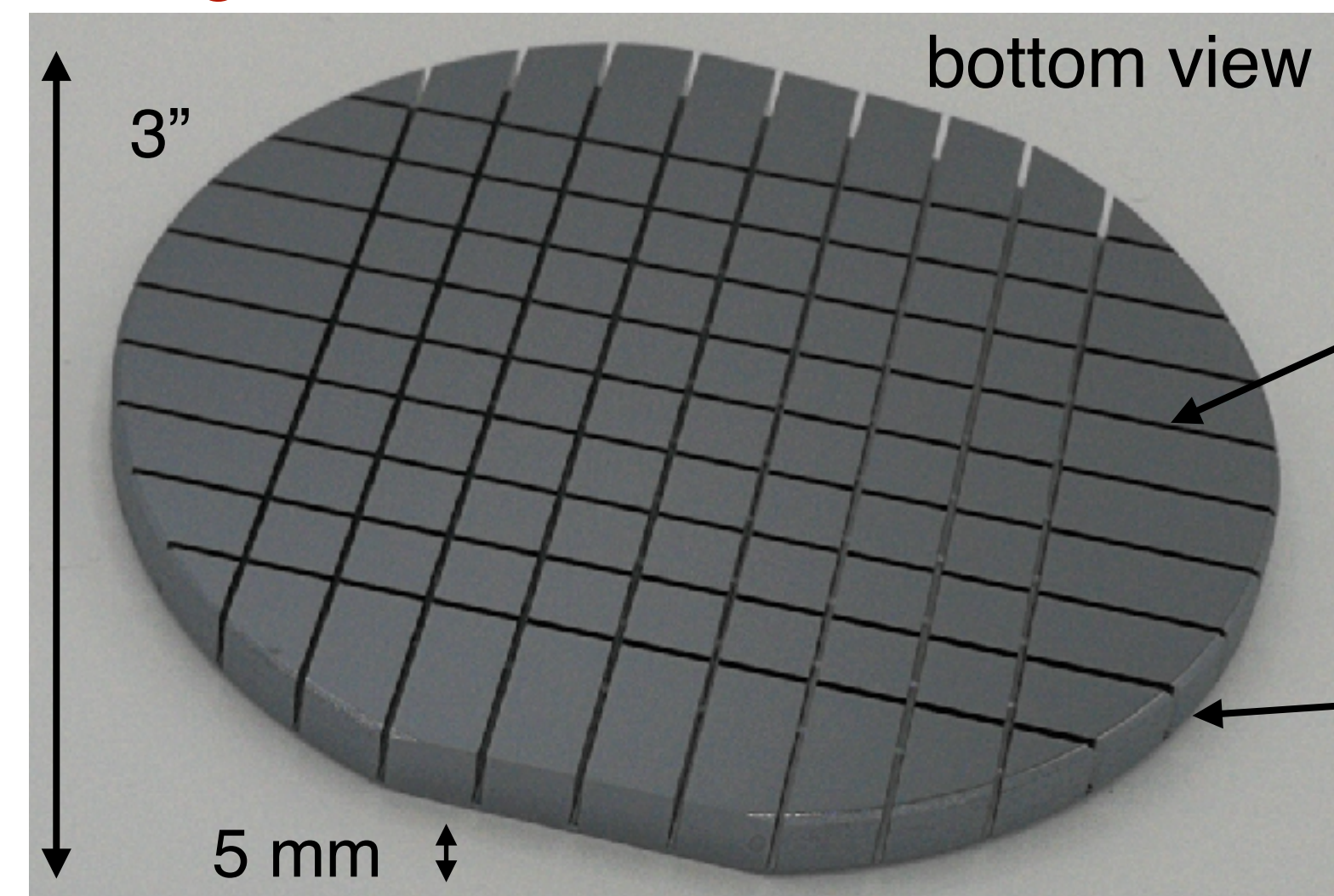
## Phonon mediation

detect phonons created by nuclear recoils  
in a silicon dice

KID ( $\sim 2 \times 2 \text{ mm}^2 \times 50 \text{ nm}$ ,  $0.5 \text{ }\mu\text{g}$ )



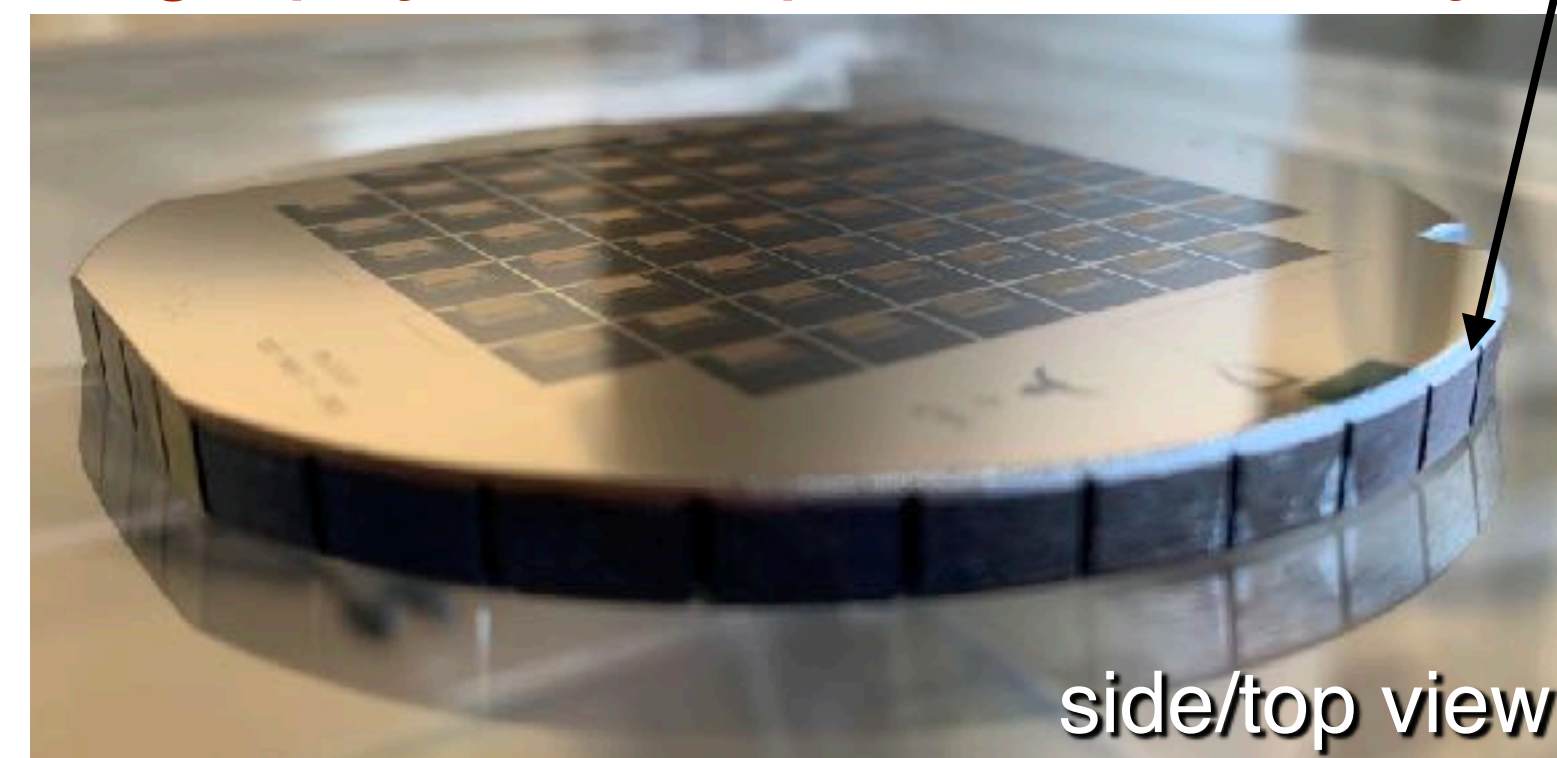
## carving of dices in a thick silicon wafer



4.5 mm deep grooves  
- 6 mm pitch  
- chemical etching

0.5 mm thick common disk:  
- holds the structure  
- hosts the KIDs

## lithography of multiplexed KID array



## KID array

- 60 nm aluminum film
- 60 KIDs lithography

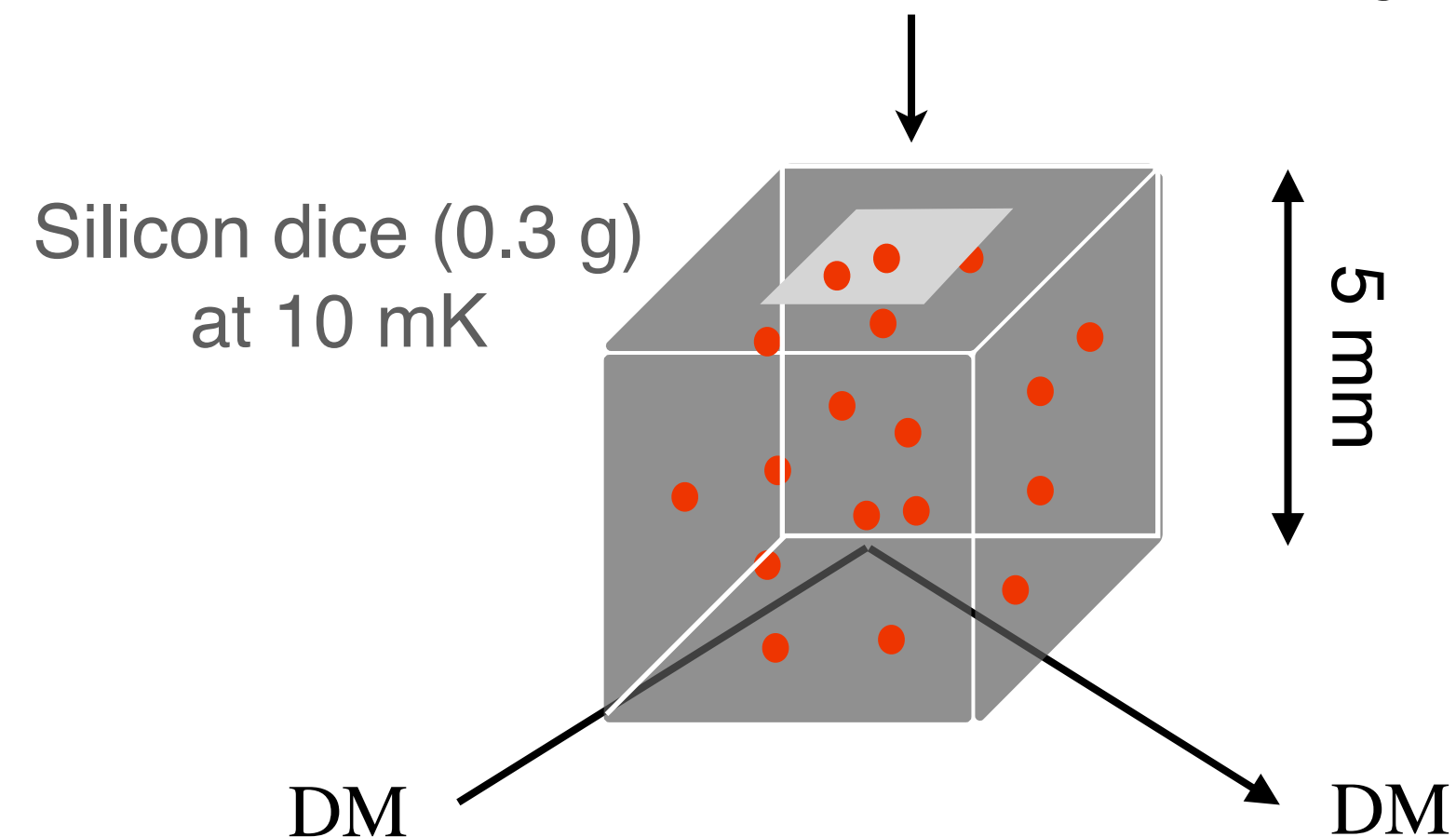


# The BULLKID phonon detector array

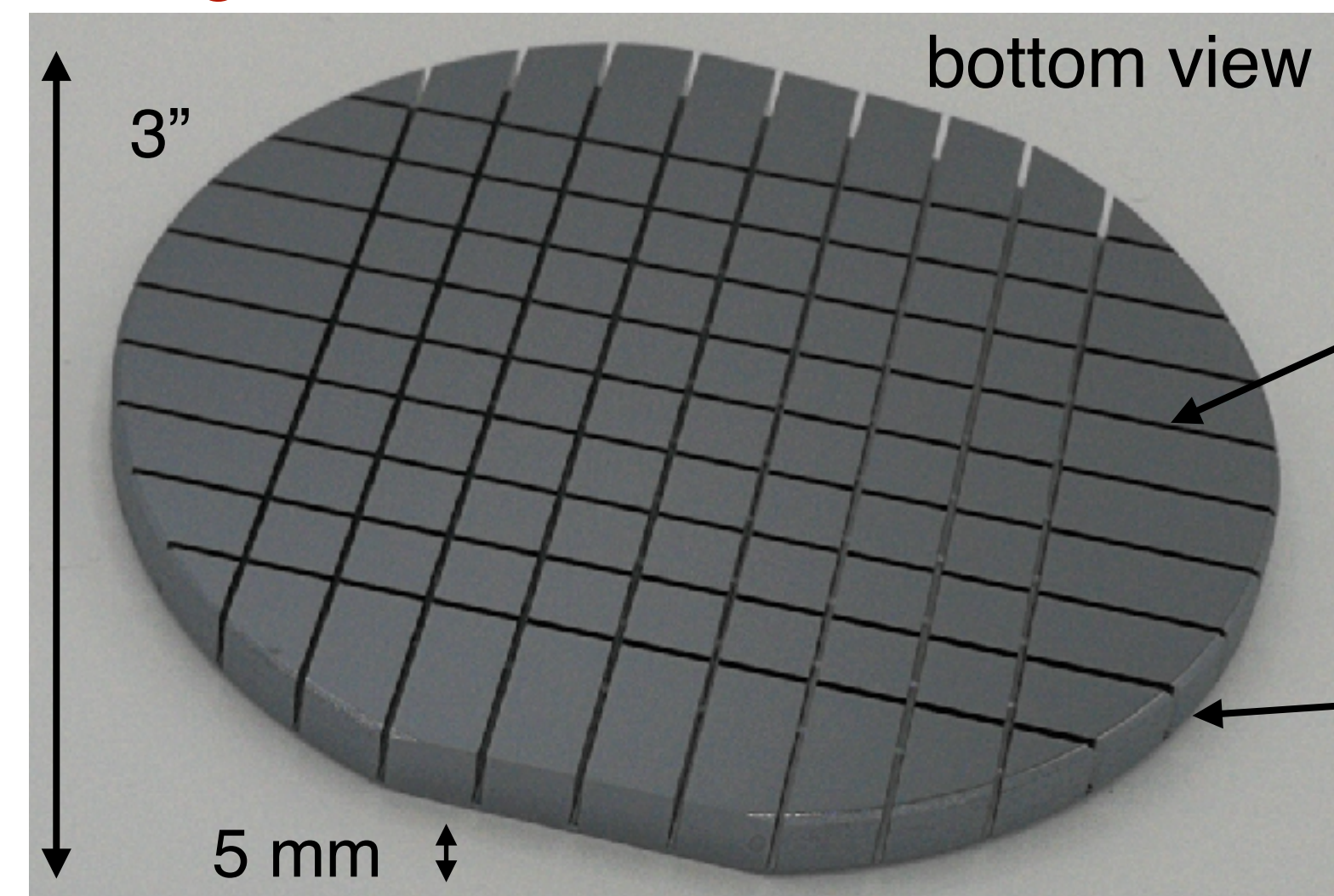
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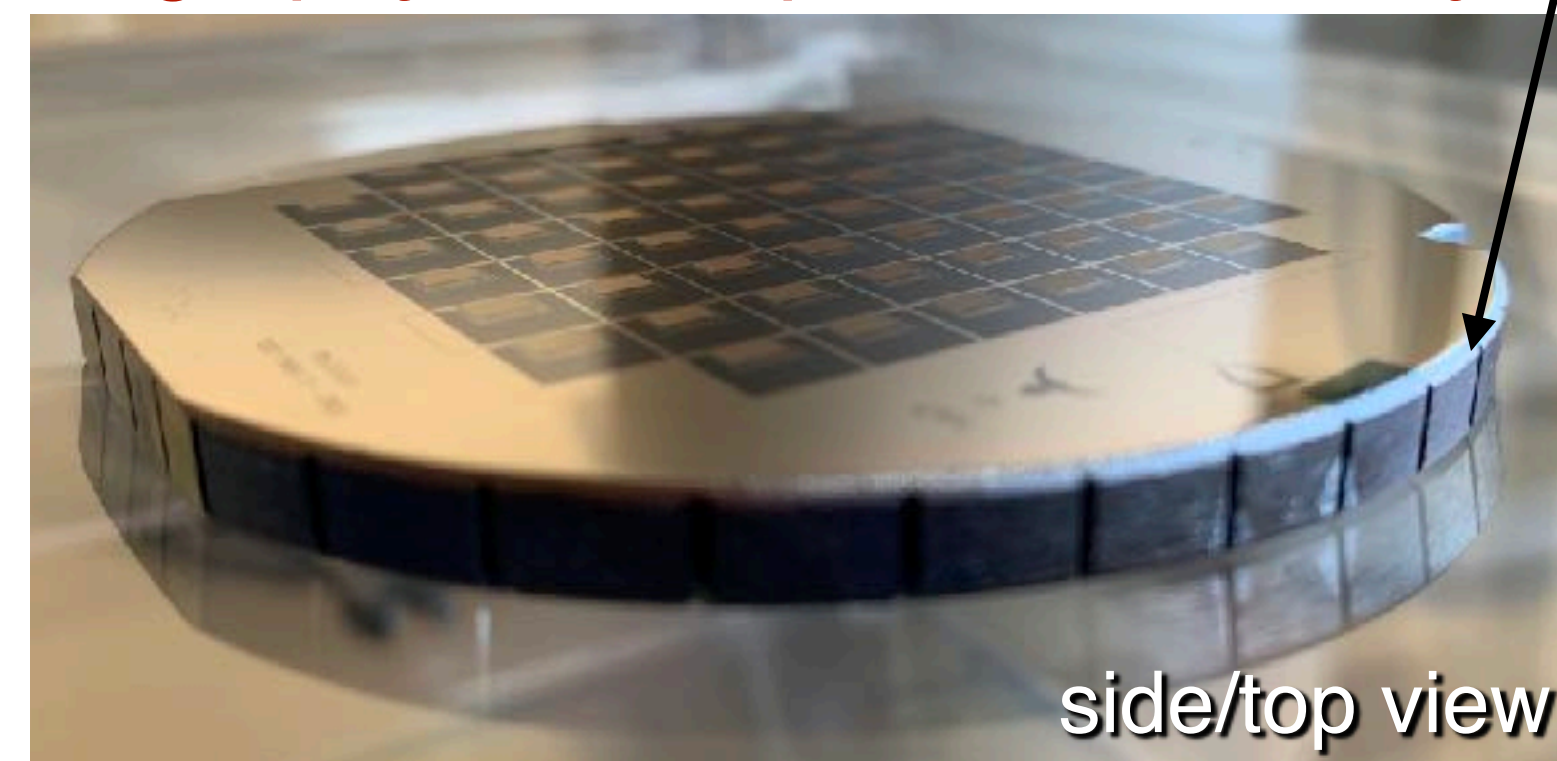
## carving of dices in a thick silicon wafer



4.5 mm deep grooves  
- 6 mm pitch  
- chemical etching

0.5 mm thick common disk:  
- holds the structure  
- hosts the KIDs

## lithography of multiplexed KID array



## KID array

- 60 nm aluminum film
- 60 KIDs lithography

✓ 60 detectors in 1

Fully multiplexed  
(single readout line)



# BULLKID: the team



## Sapienza University & INFN Rome

*A. Ahmad, **A. Cruciani**, M. del Gallo Roccagiovine,  
D. Delicato, G. Del Castello, M. Giammei,  
D. Maiello, V. Pettinacci, **M. Vignati**;*



## University & INFN Ferrara

*L. Bandiera, V. Guidi, **A. Mazzolari**, M. Romagnoni,  
M. Tamisari;*



## Institut Néel - CNRS

*M. Calvo, D. Delicato, **A. Monfardini**;*

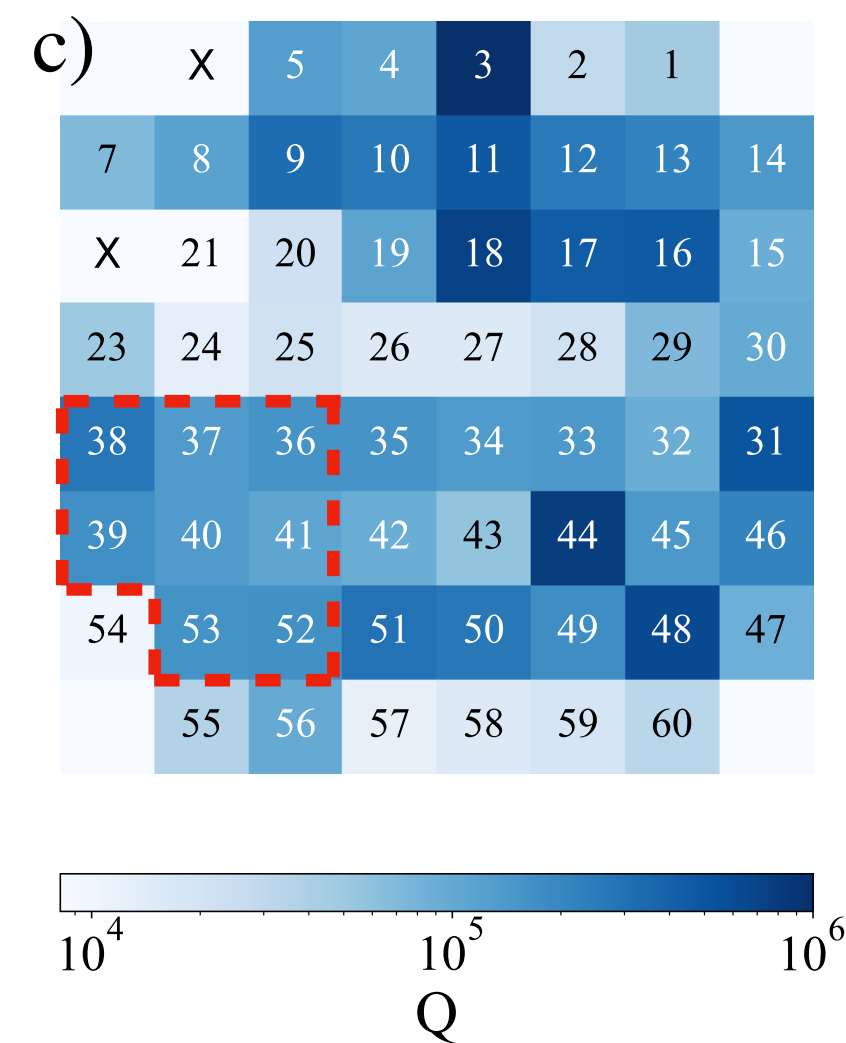
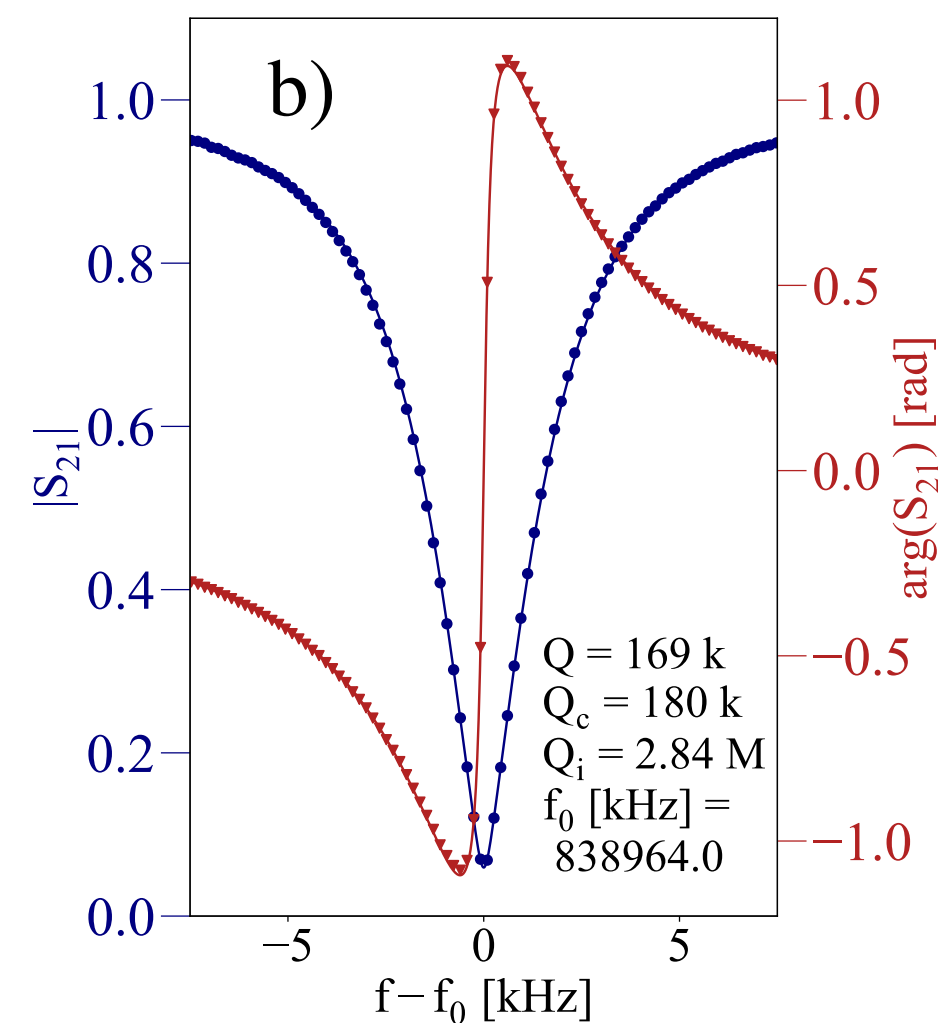
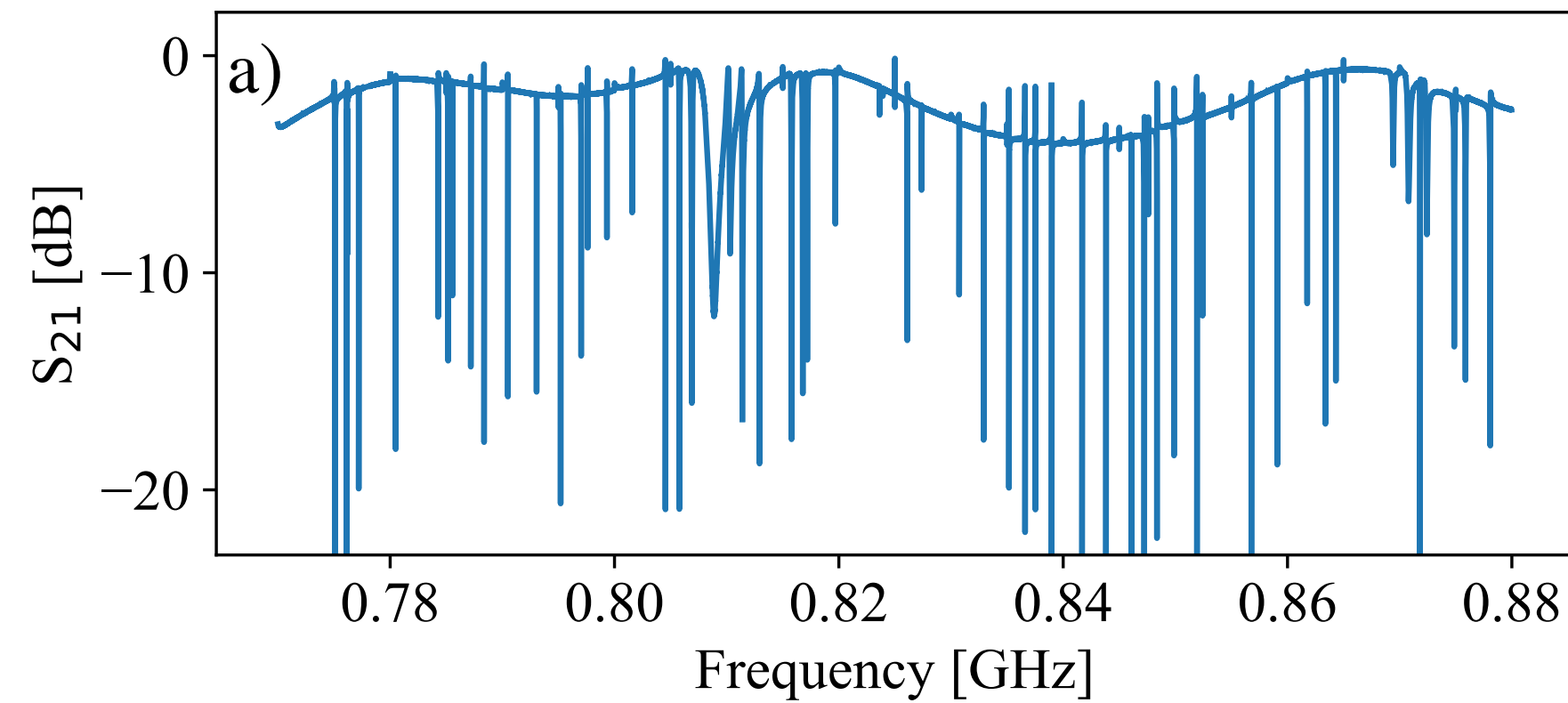


**BULLKID**

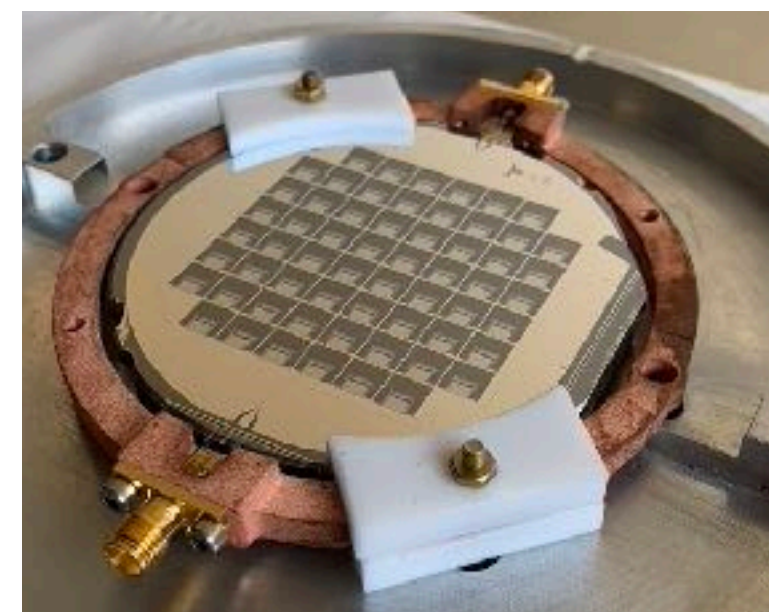
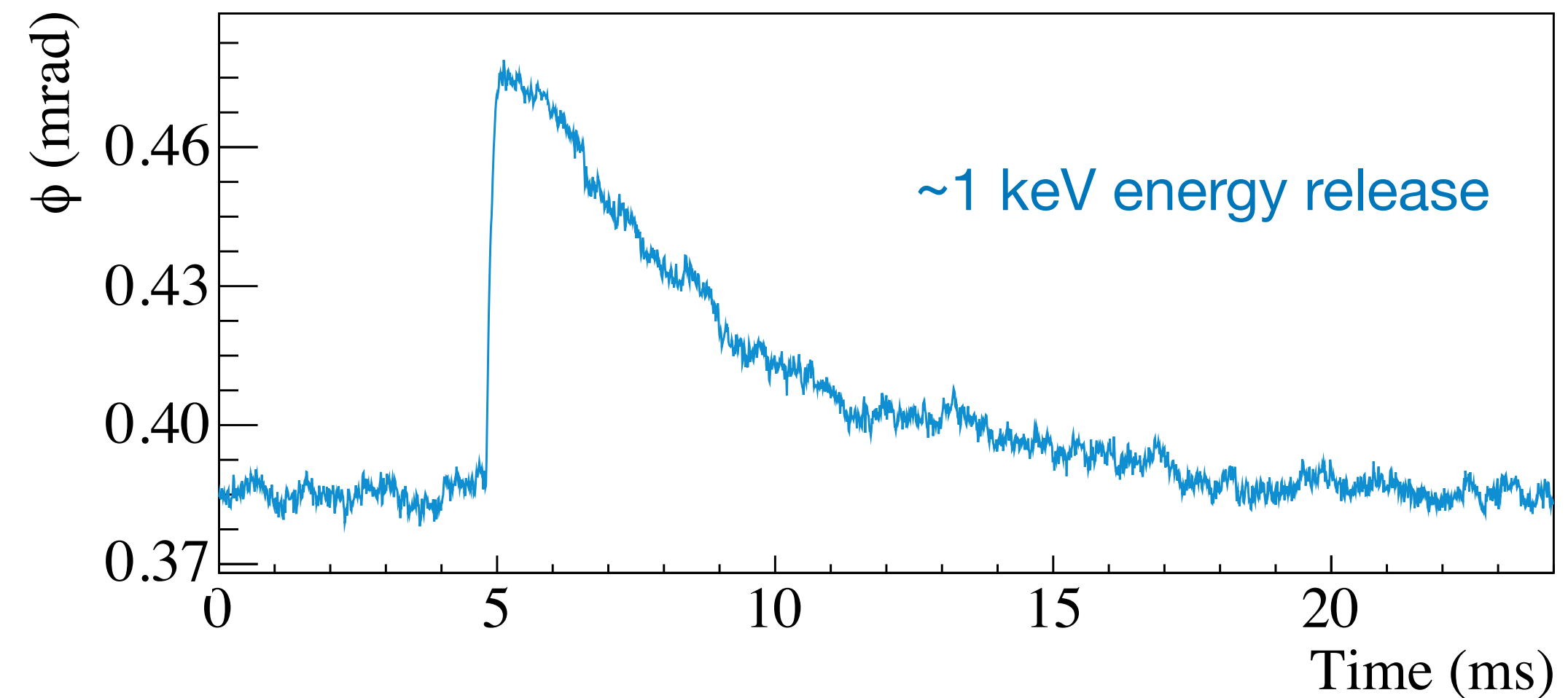
*funded by the INFN and by Sapienza U.*

# First results (Summer '22)

Frequency scan of the KID array



Particle interaction in a dice



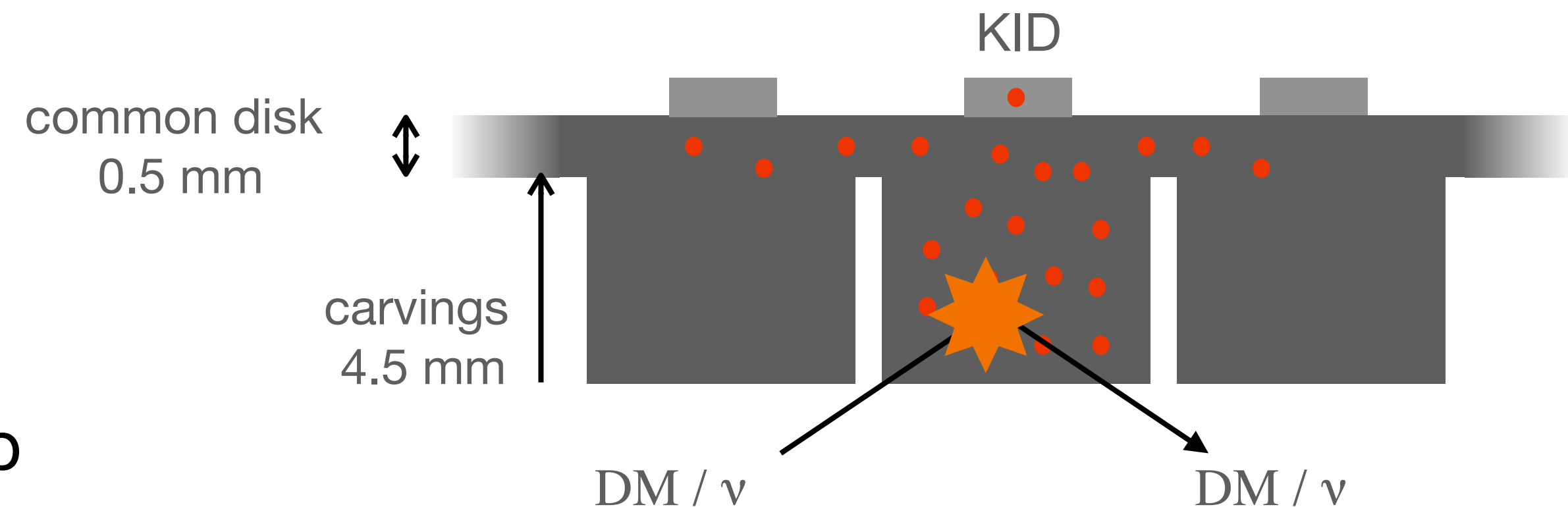
- + ) 58/60 KIDs alive
- + ) RMS@0 eV:  **$26 \pm 7 \text{ eV}$**
- ) Response not uniform

[A. Cruciani, et al, Appl. Phys. Lett. 121, 213504 \(2022\)](#)

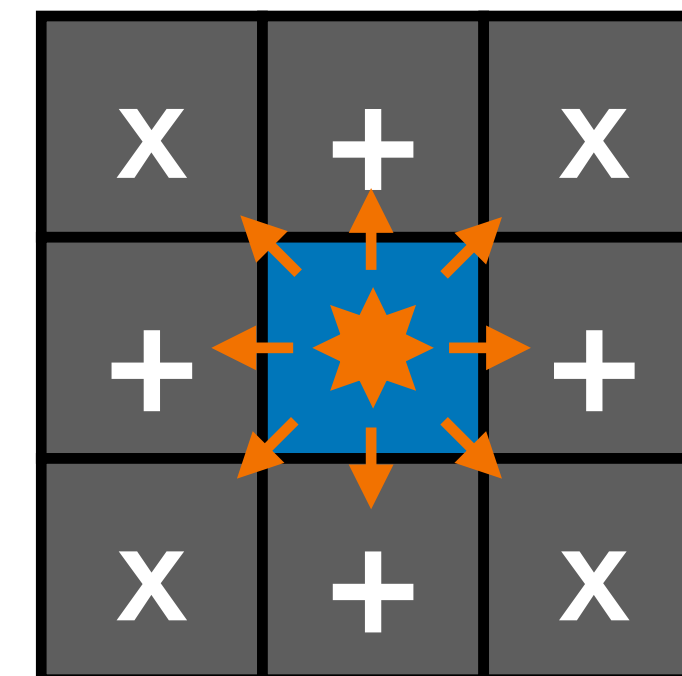


# Phonon leakage

- Phonons generated by interactions
  - 40% absorbed by the KID
  - the rest leaks in nearby voxels or decays below the KID aluminum gap



- Measured energy leakage relative to central voxel:
  - $(14 \pm 3) \%$  in each “+” voxel
  - $(5 \pm 1) \%$  in each “x” voxel



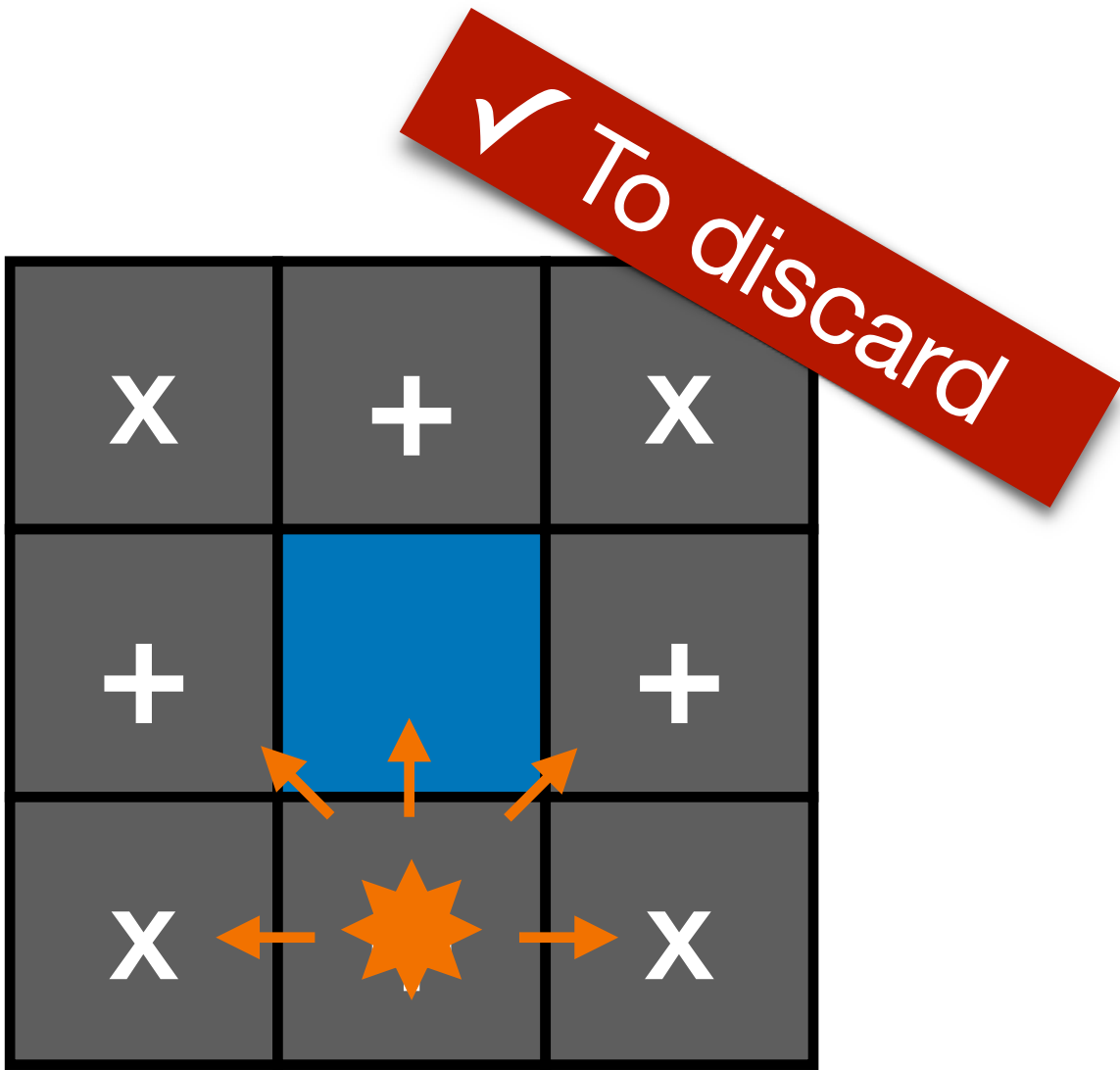
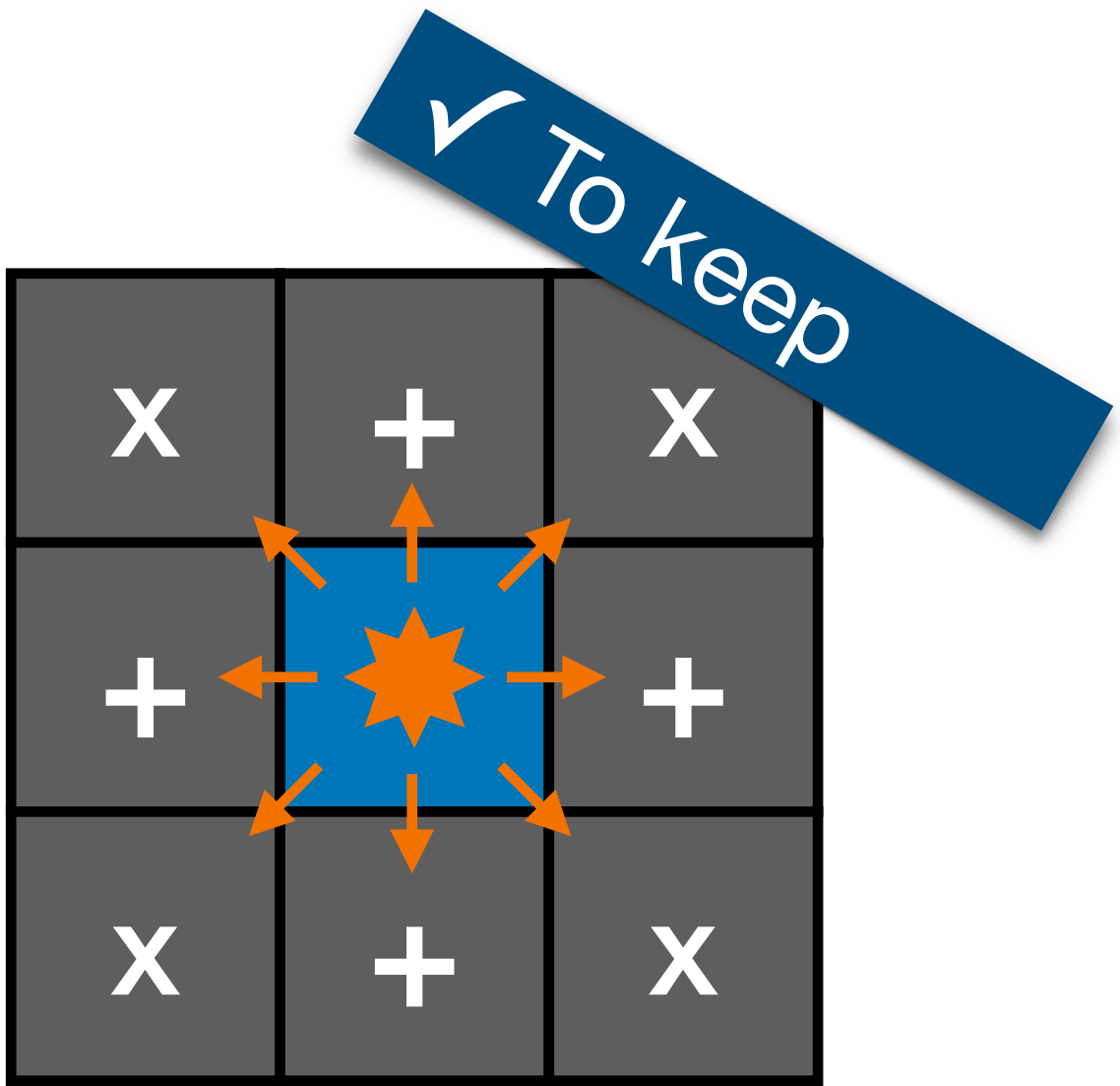
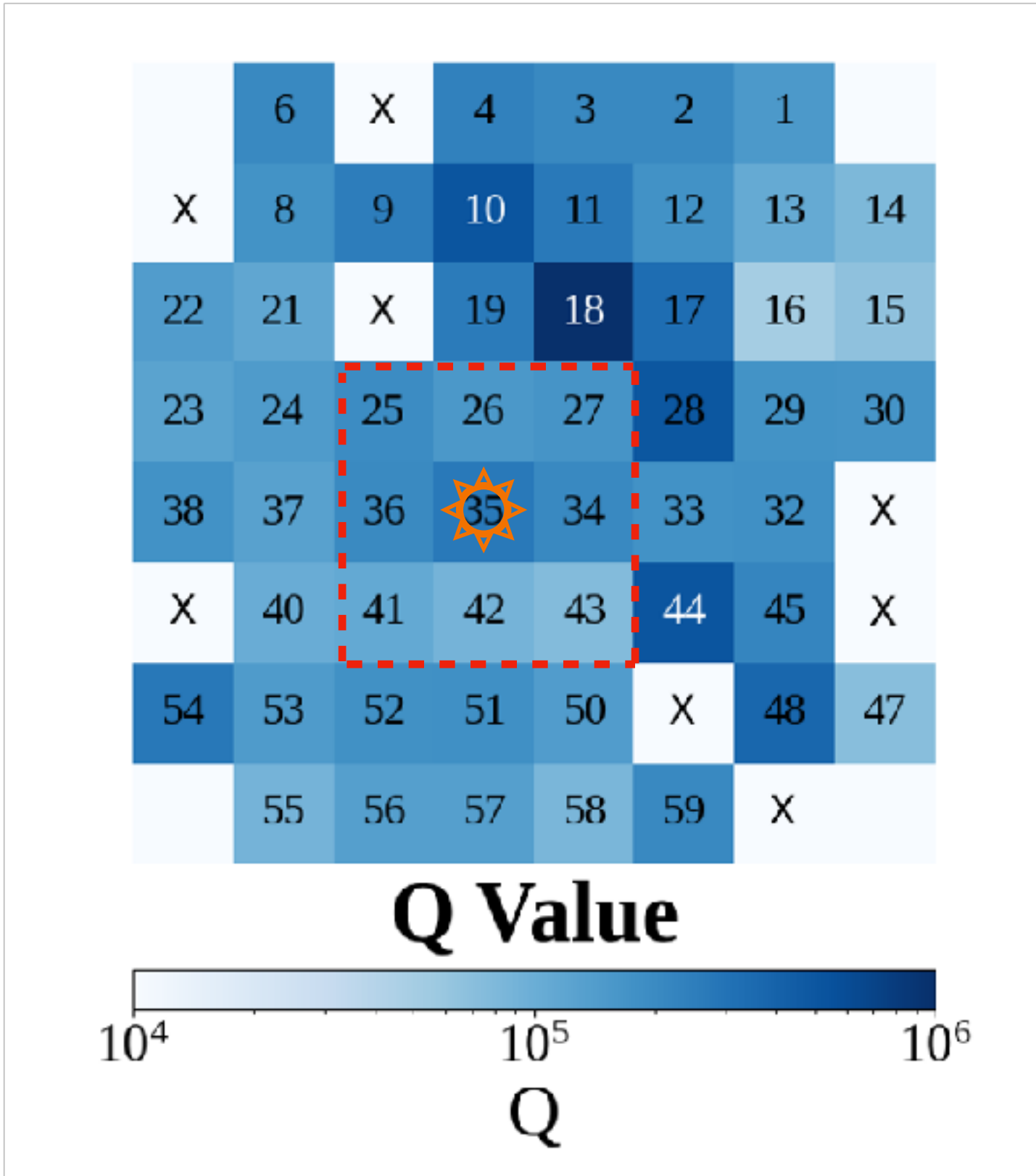
This effect reduces the phonon focusing on the KID but **it can be exploited to reconstruct the interaction voxel**



# New: combined analysis of a 9-dice cluster

Measurement of the energy spectrum of the central voxel

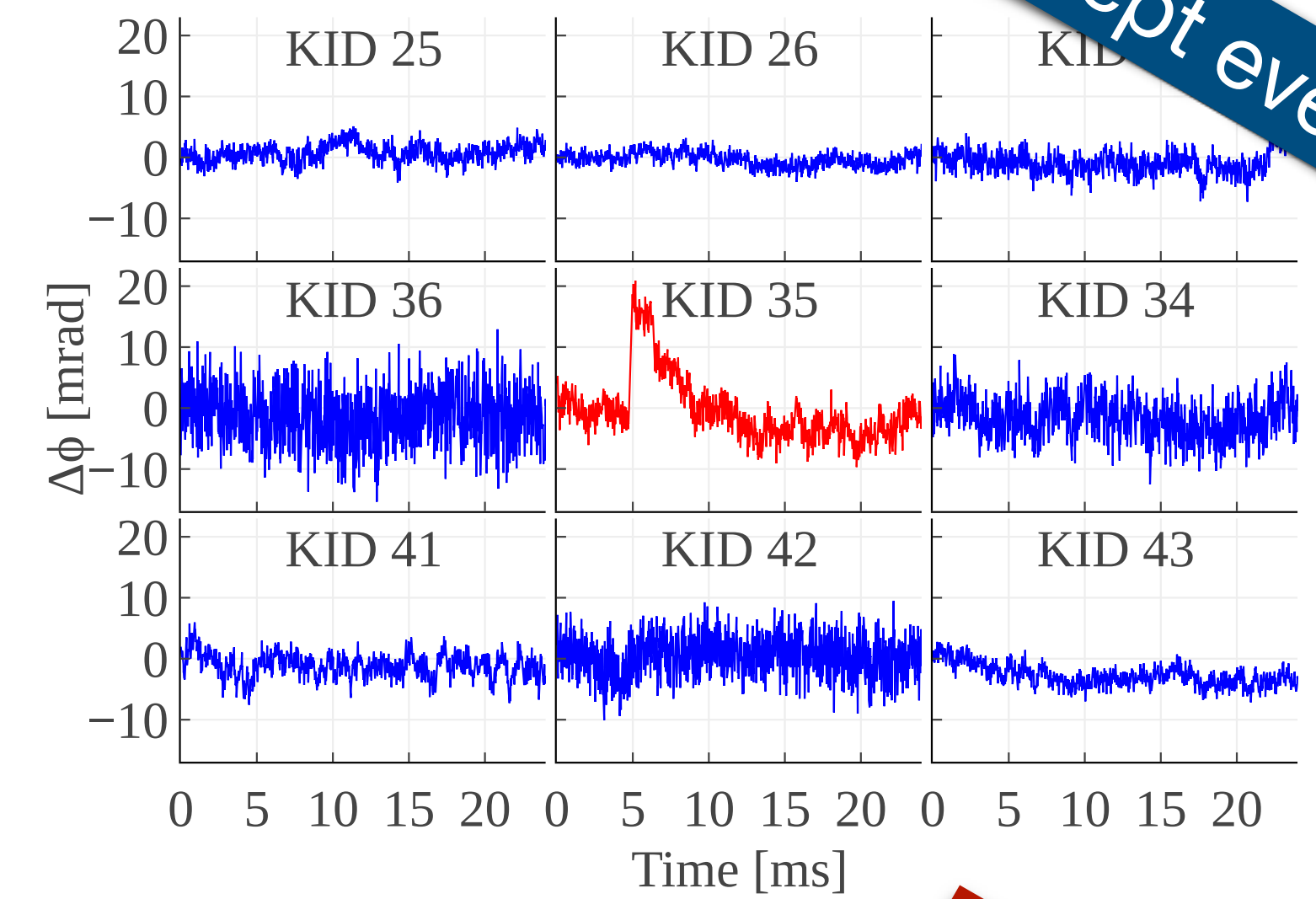
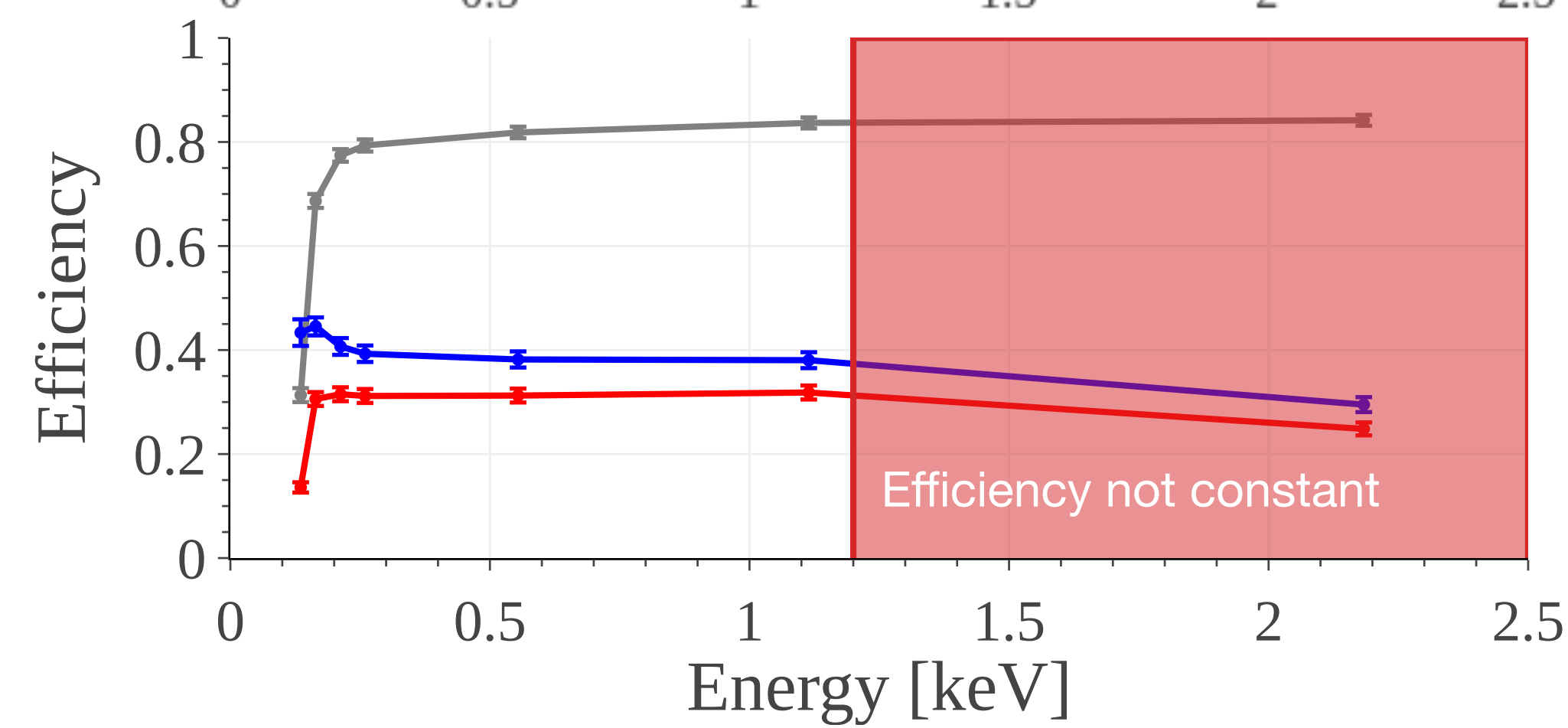
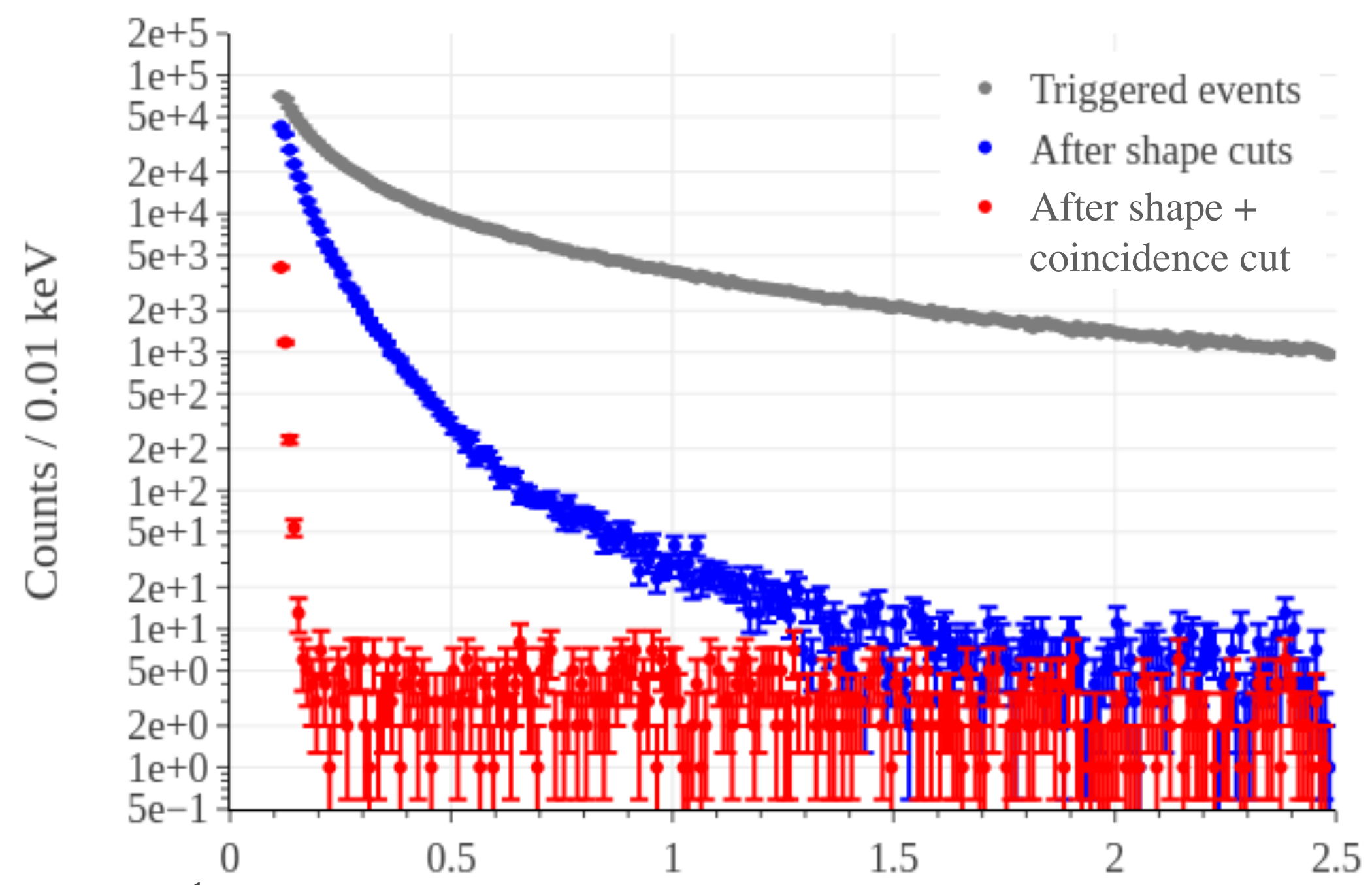
Use the 8 external voxels as “veto” exploiting the phonon leakage



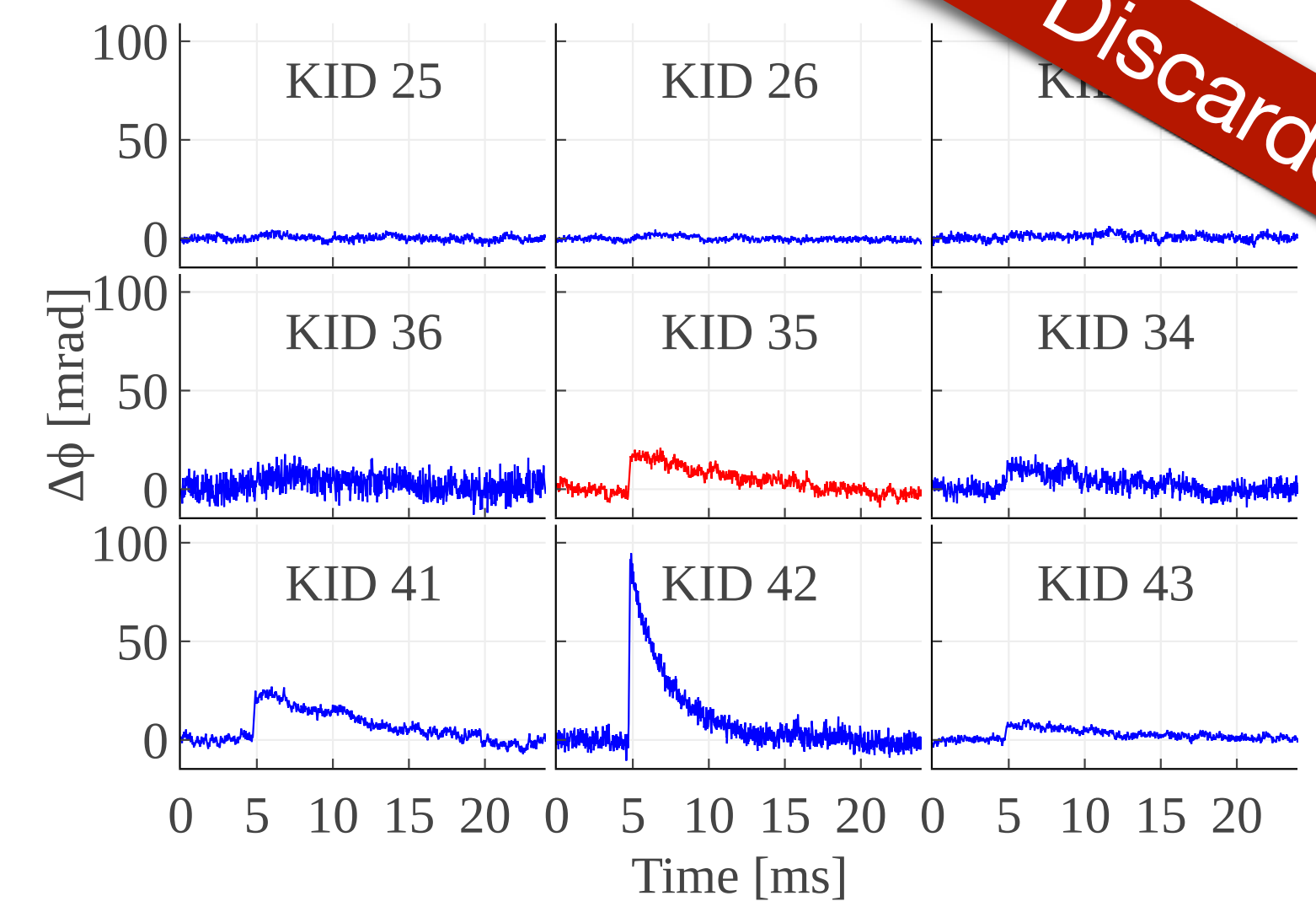


# Background: shape + coincidence cut

✓ Kept event



✓ Discarded

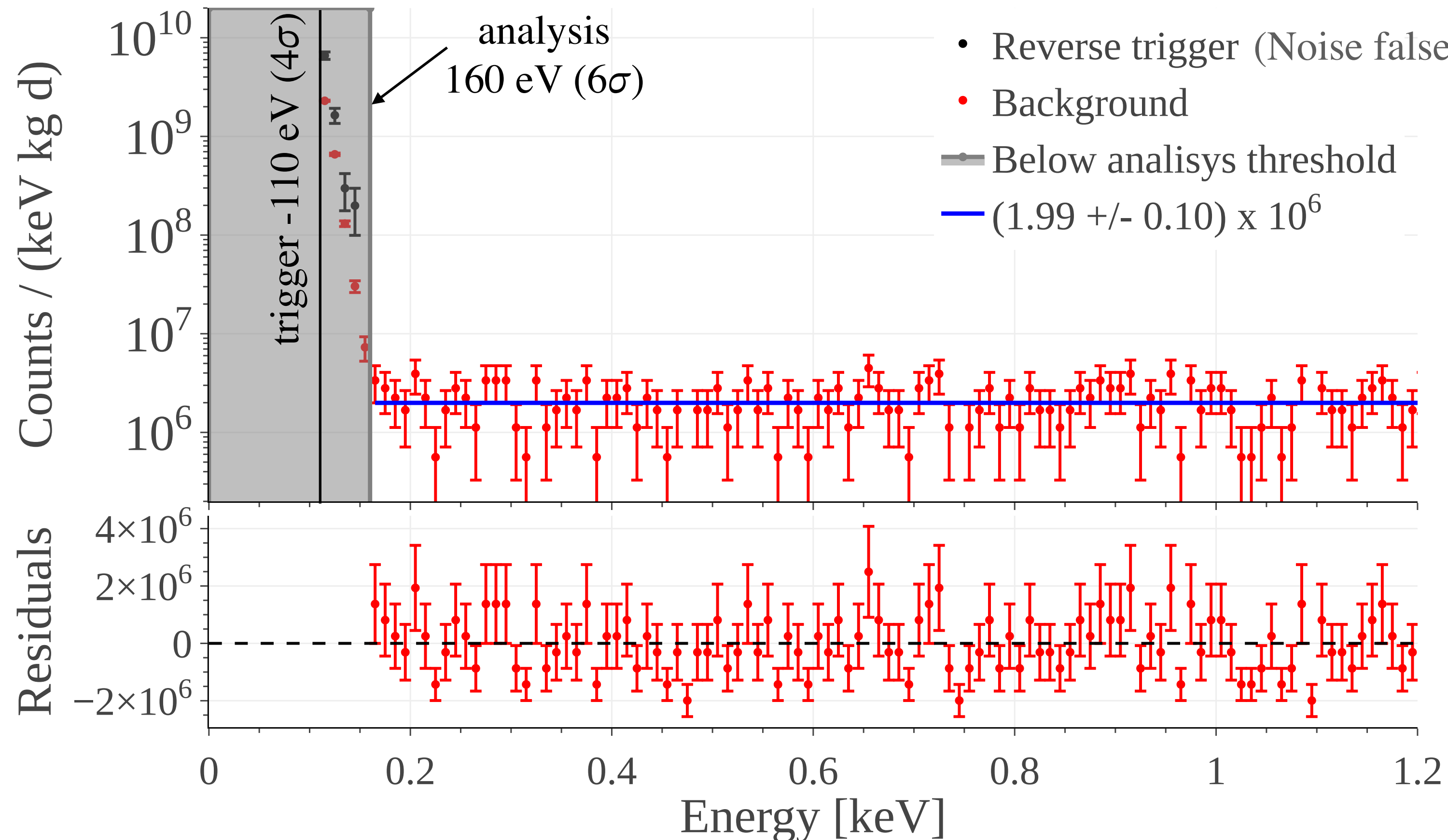




# Background: result

Above ground lab, no shield, 39 live hours

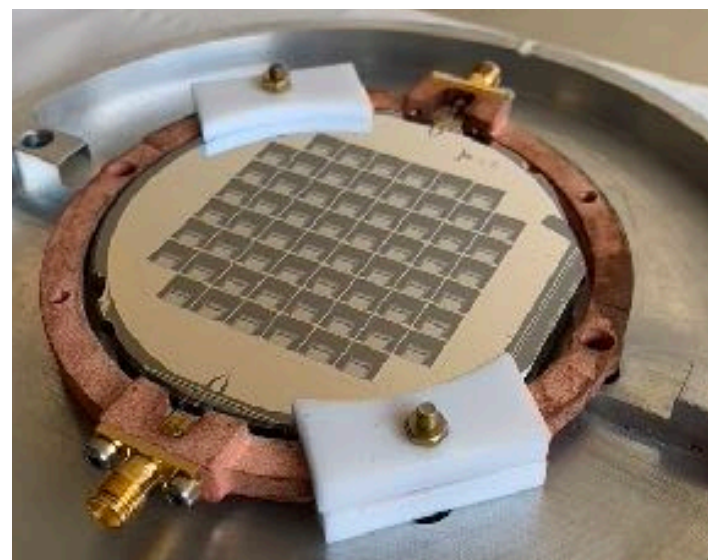
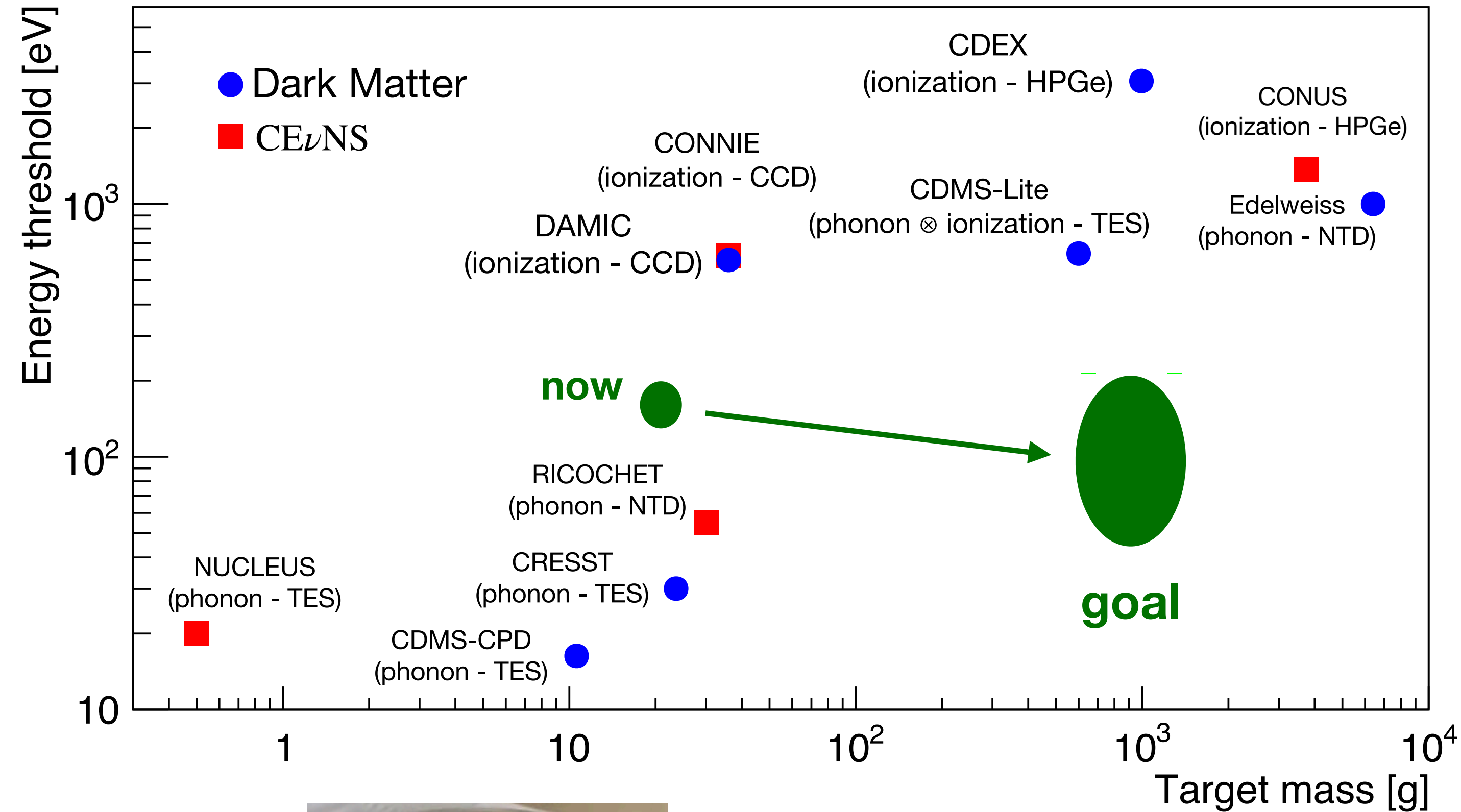
D. Delicato et al,  
[arXiv:2308.14399](https://arxiv.org/abs/2308.14399)



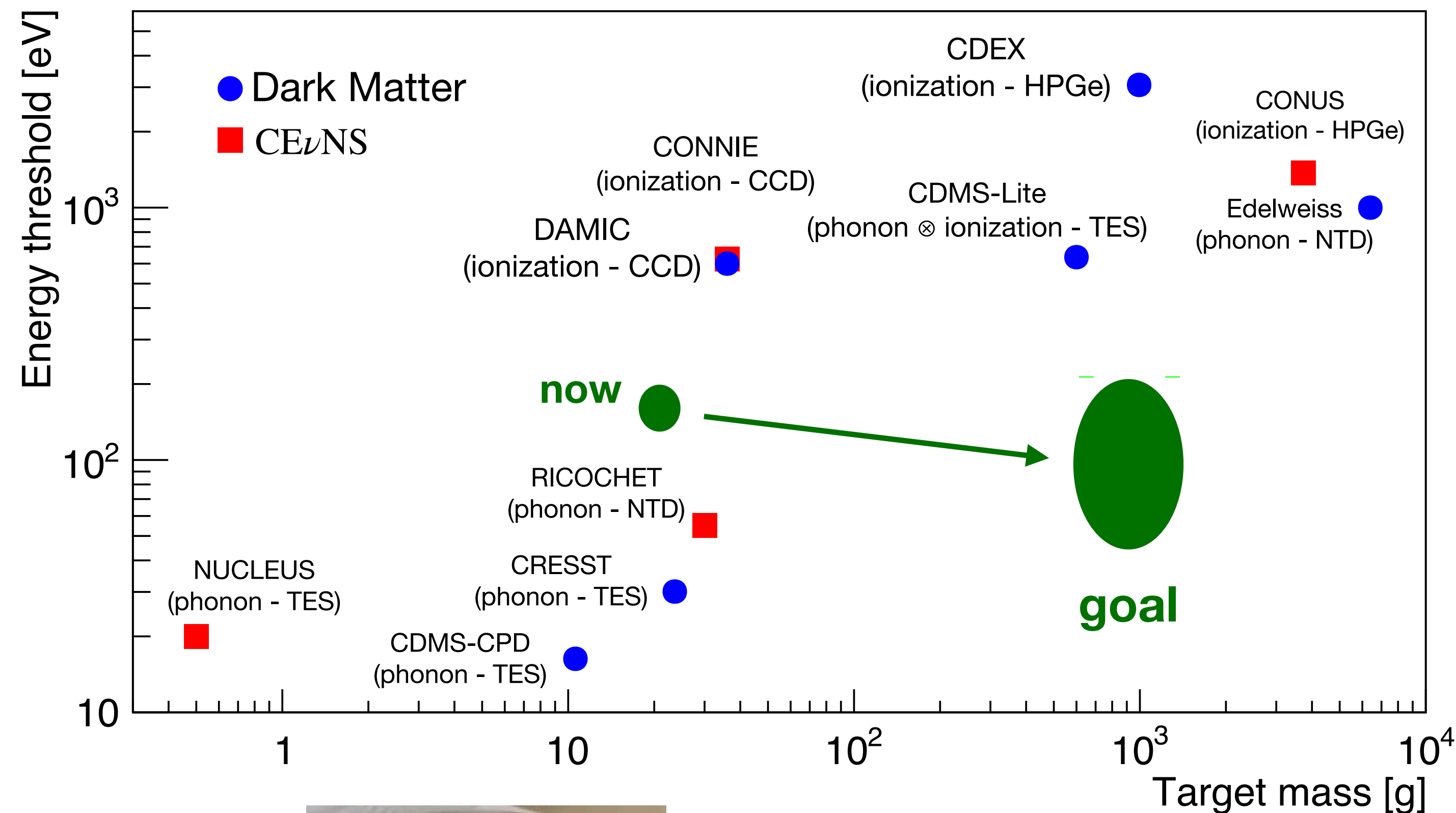
The excess above trigger threshold is compatible with noise false positives.  
Background is flat above analysis threshold.



# Mass and threshold improvement

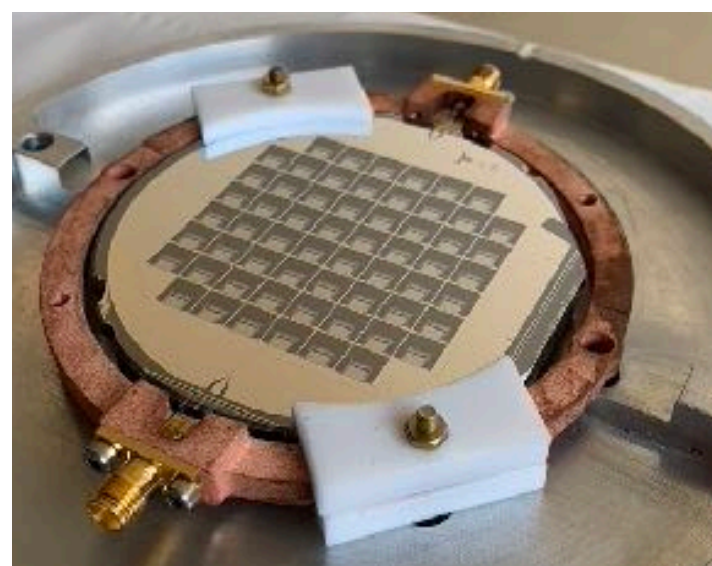
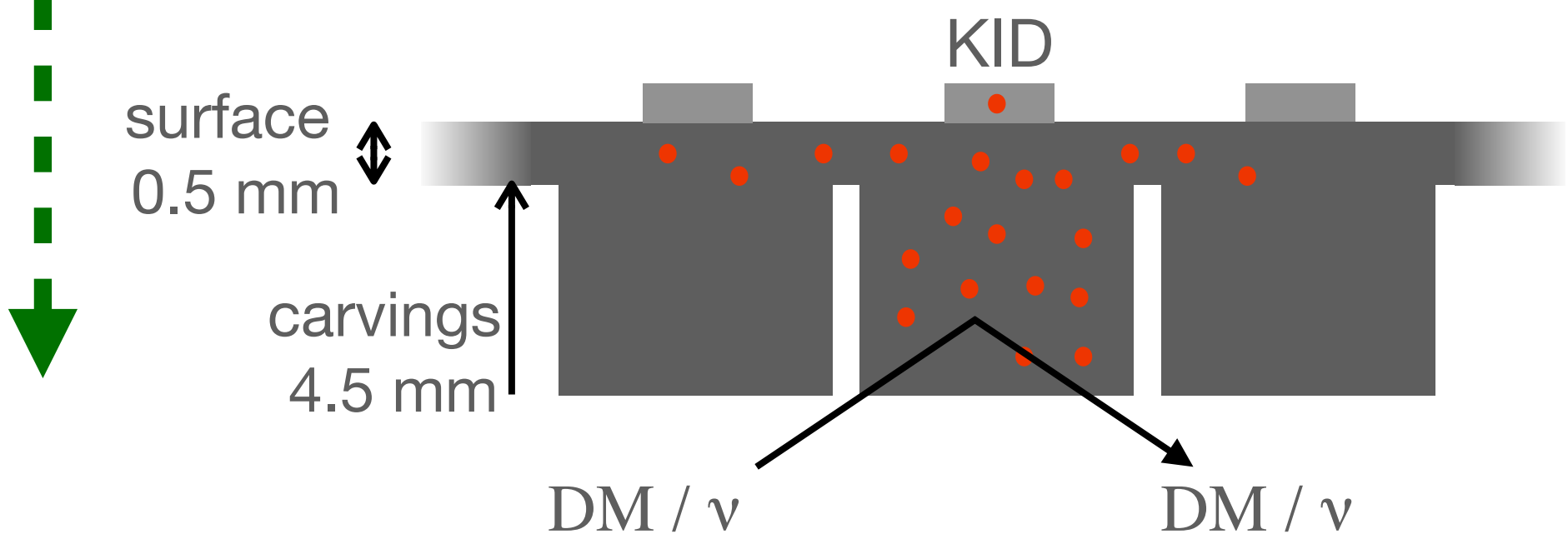


# Mass and threshold improvement



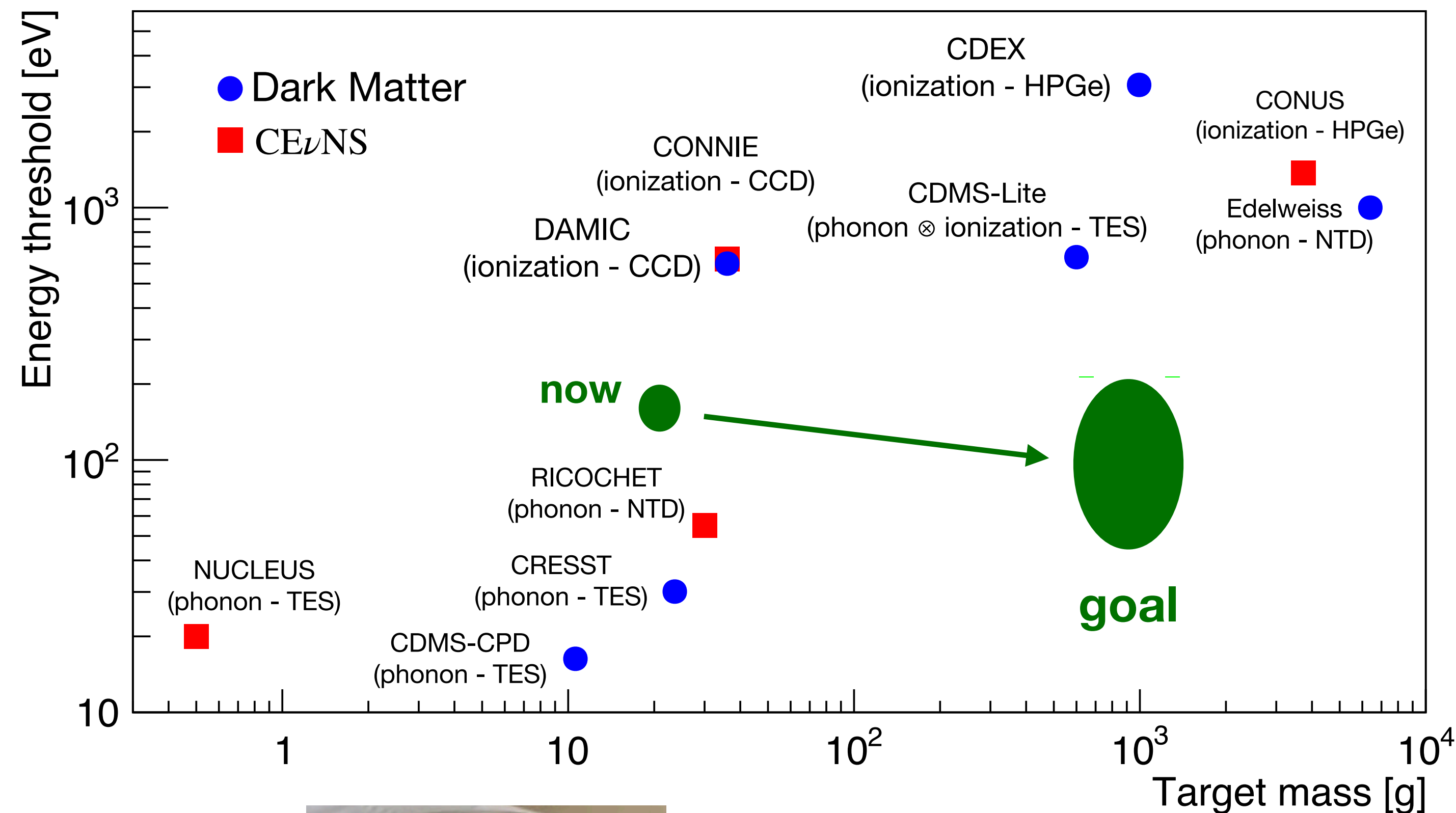
Threshold (ongoing R&D):

1. Replace Al with Al-Ti-Al KIDs - 5x inductance
2. Deeper carvings for higher phonon focussing



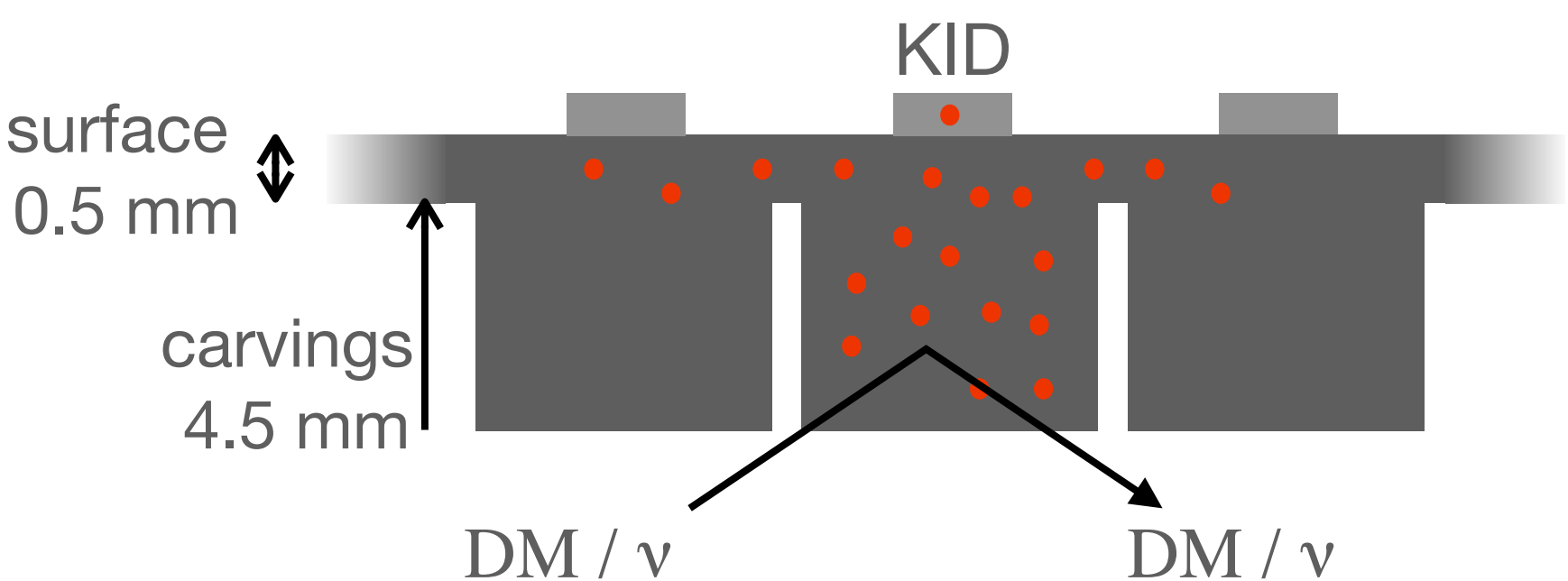


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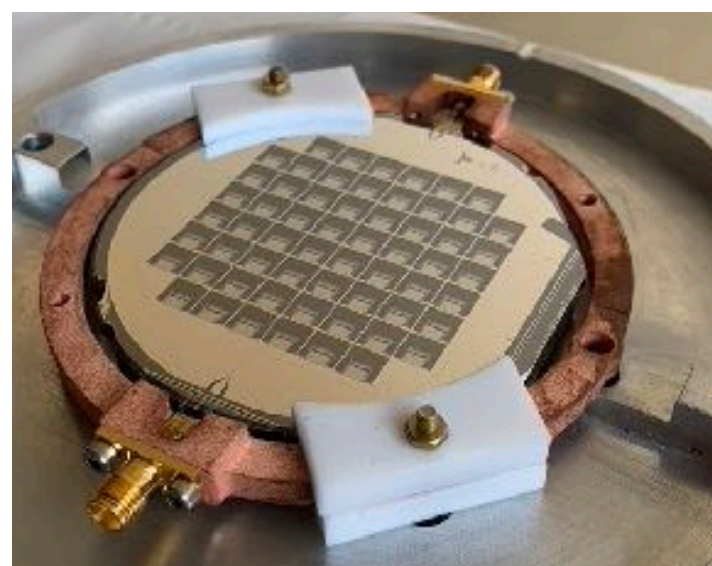


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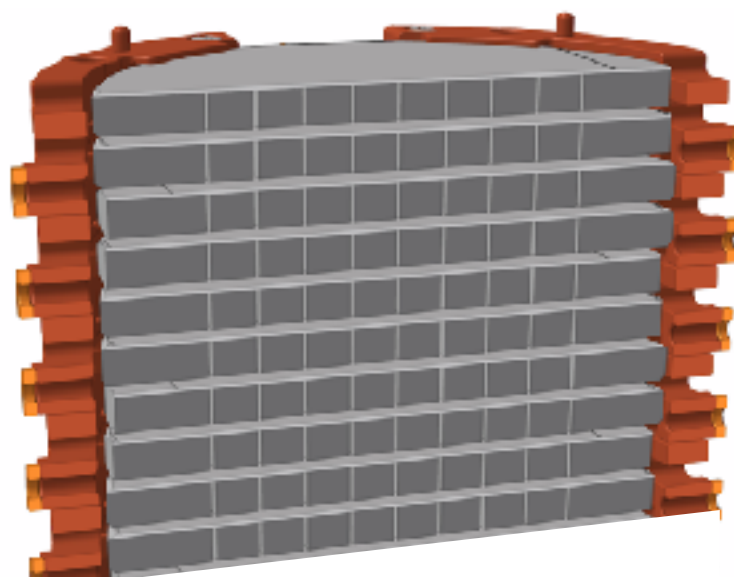
1. Replace Al with Al-Ti-Al KIDs - 5x inductance
2. Deeper carvings for higher phonon focussing



3D copper printing at LNGS  
3 wafer holder prototype



Mass:  
from 3" to 4" wafers  
stack of wafers





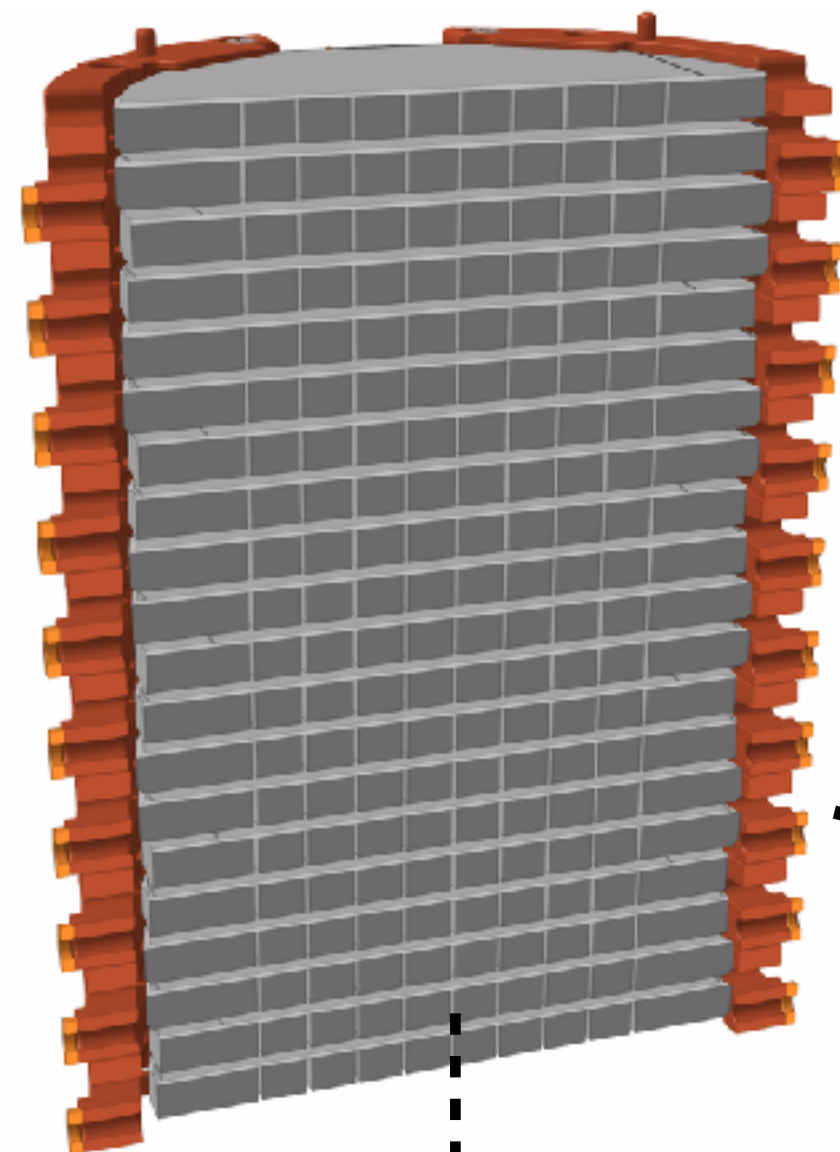
# Impact on Dark Matter search

Nuclear recoil detector with:

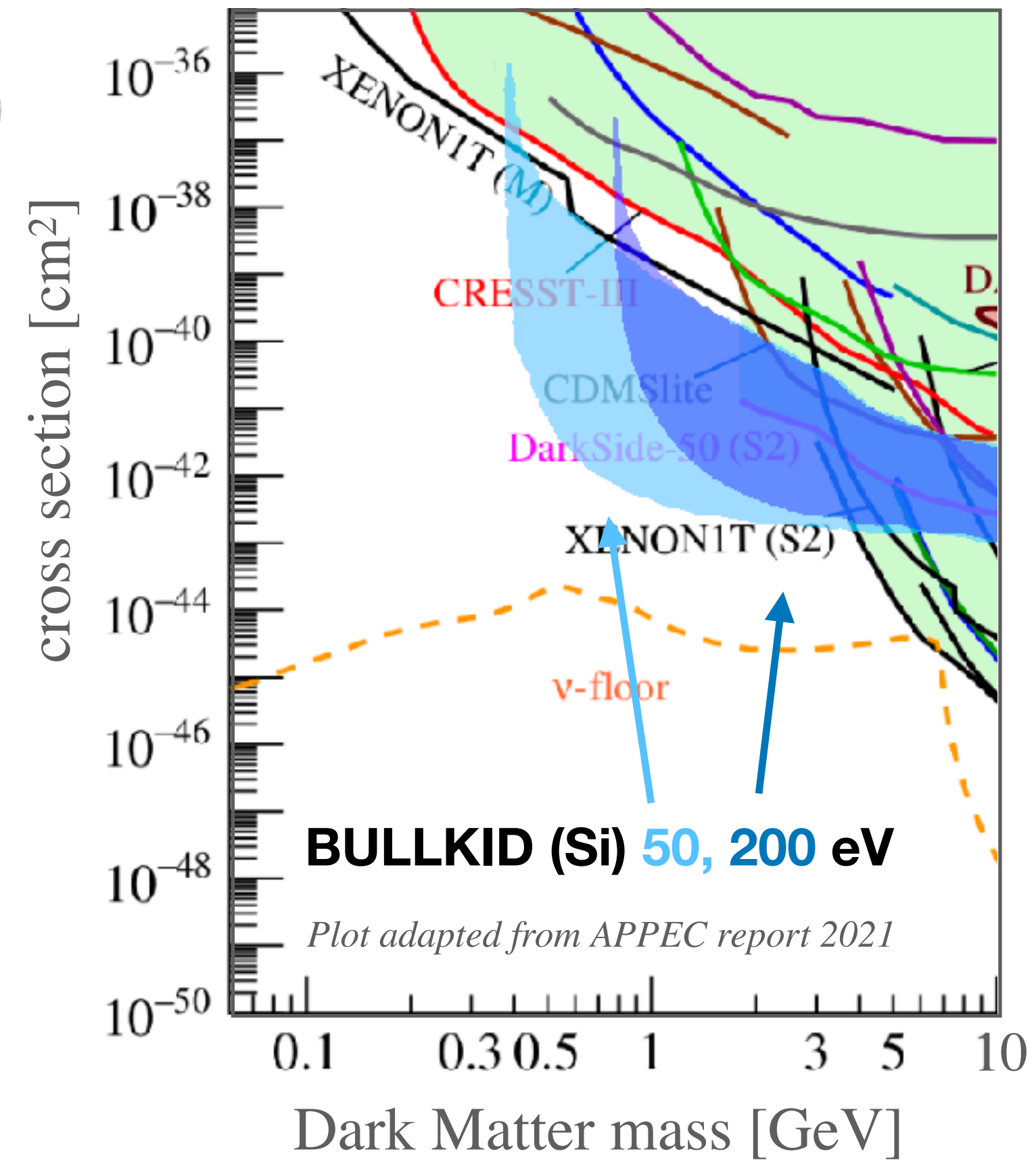
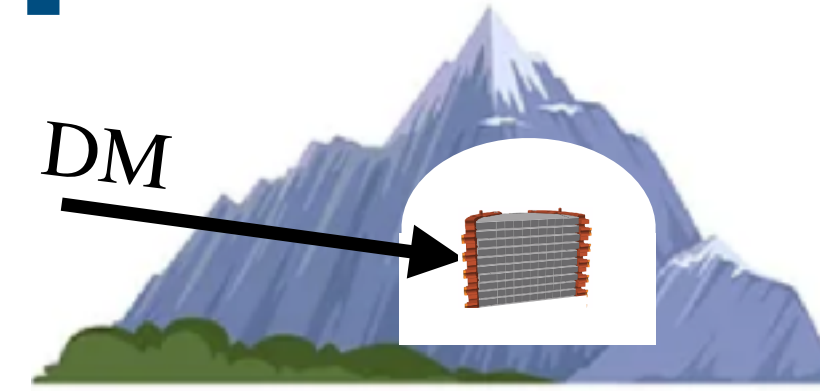
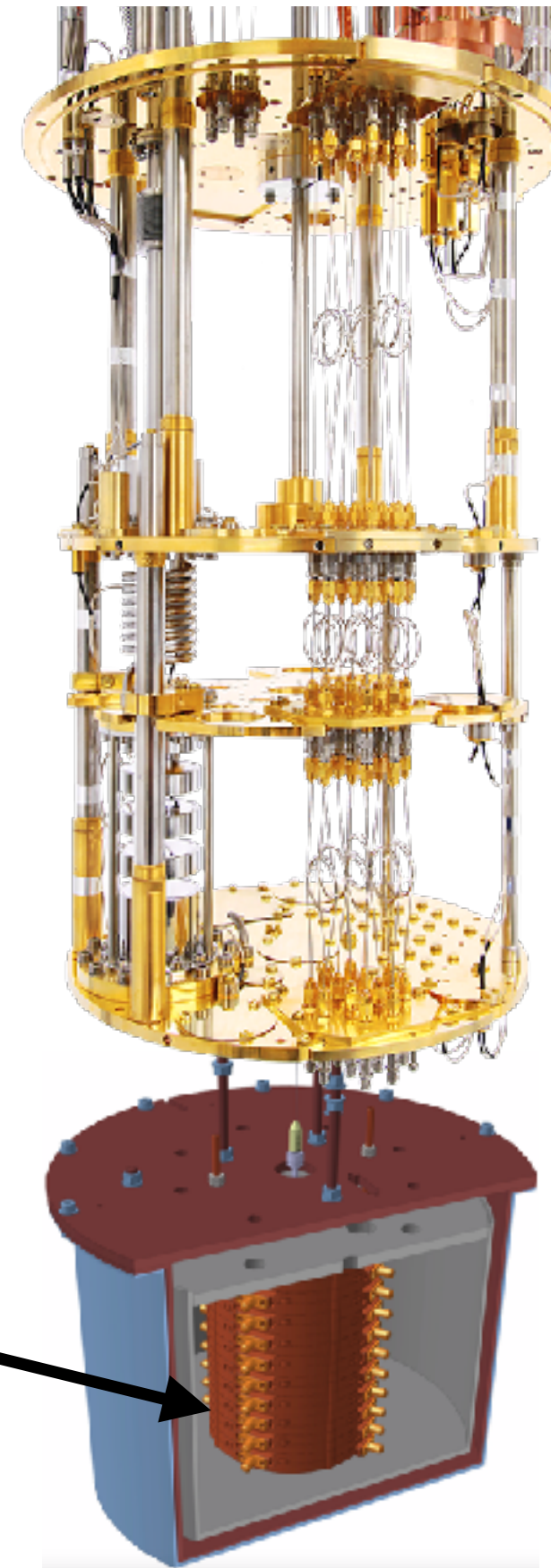
- ✓ 15 (4") or 30 (3") BULLKIDs (2000 voxels)
- ✓ 0.6 kg of silicon target
- ✓ 200 ÷ 50 eV threshold (160 eV demonstrated)

Unique features for bkg. suppression:

- ✓ No inert material in detector volume
- ✓ fully active
- ✓ fiducialization



✓ scalable



bands range  
same background  
as CRESST

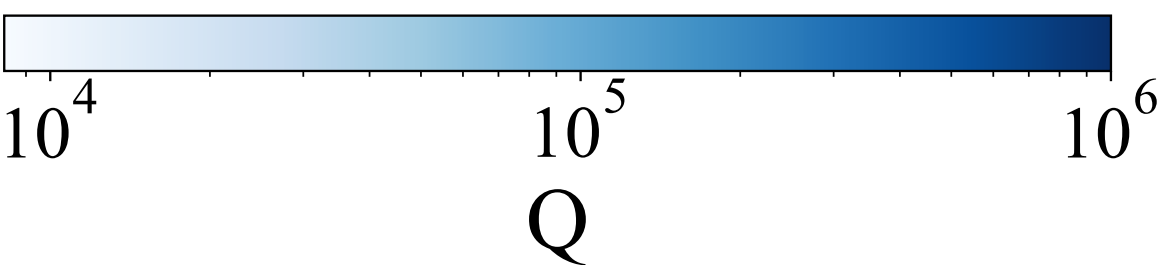
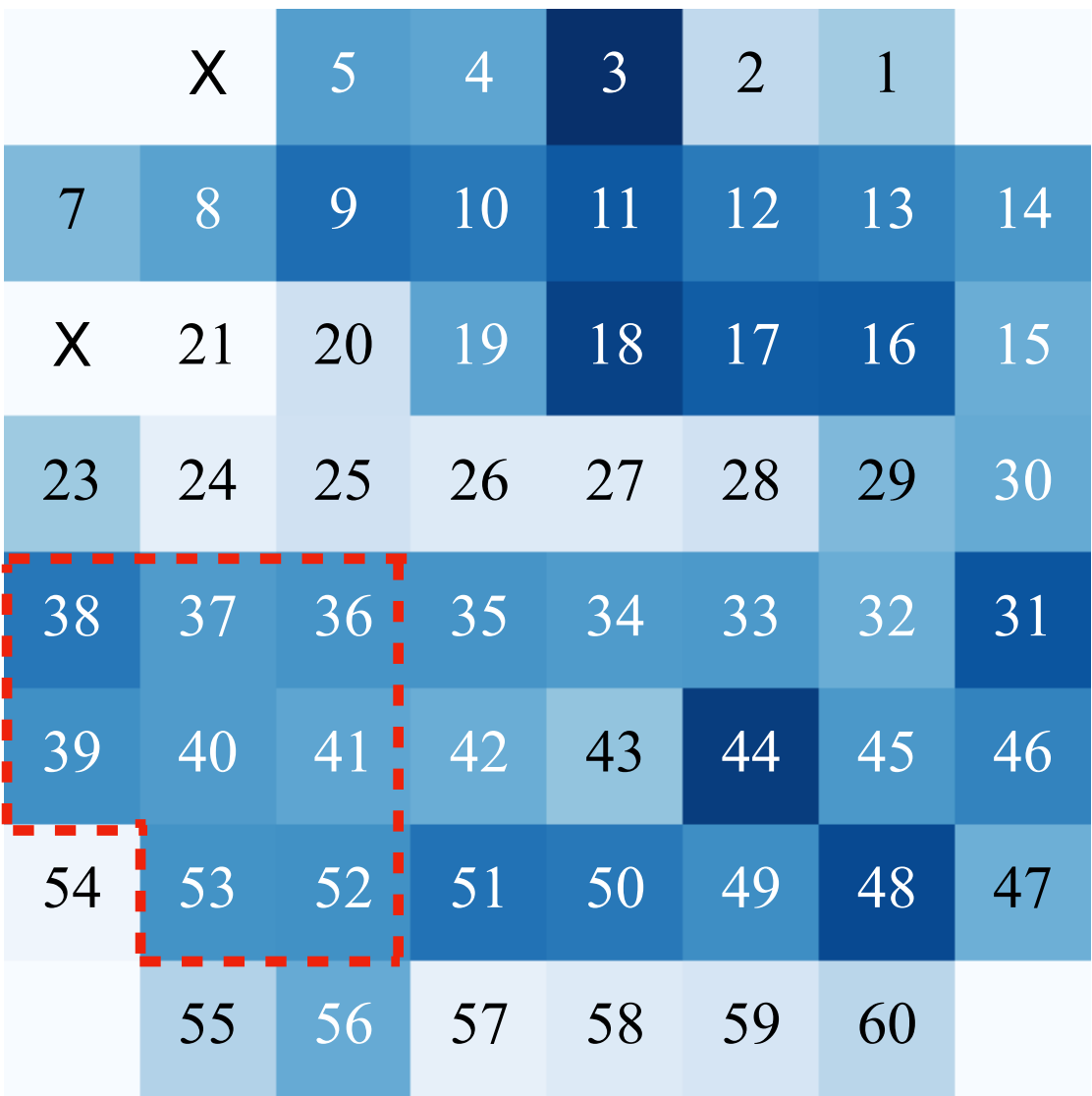
↓  
zero background



# Backup slides

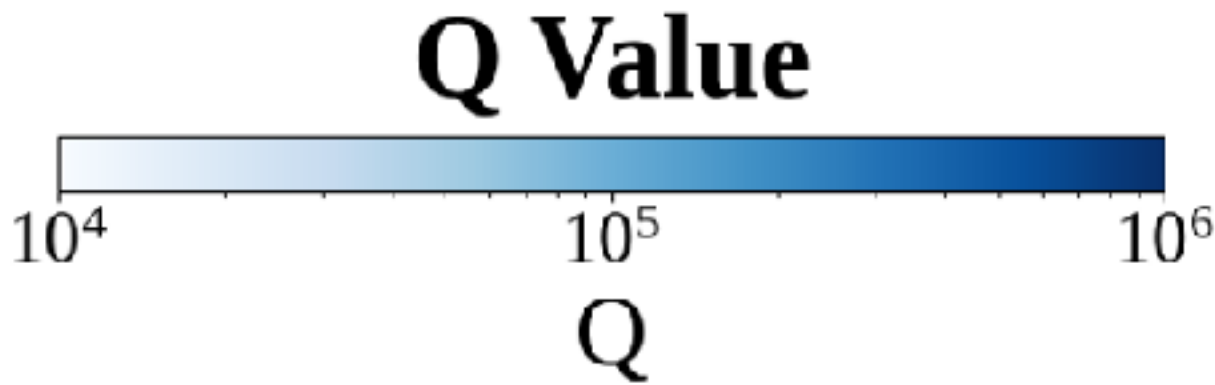
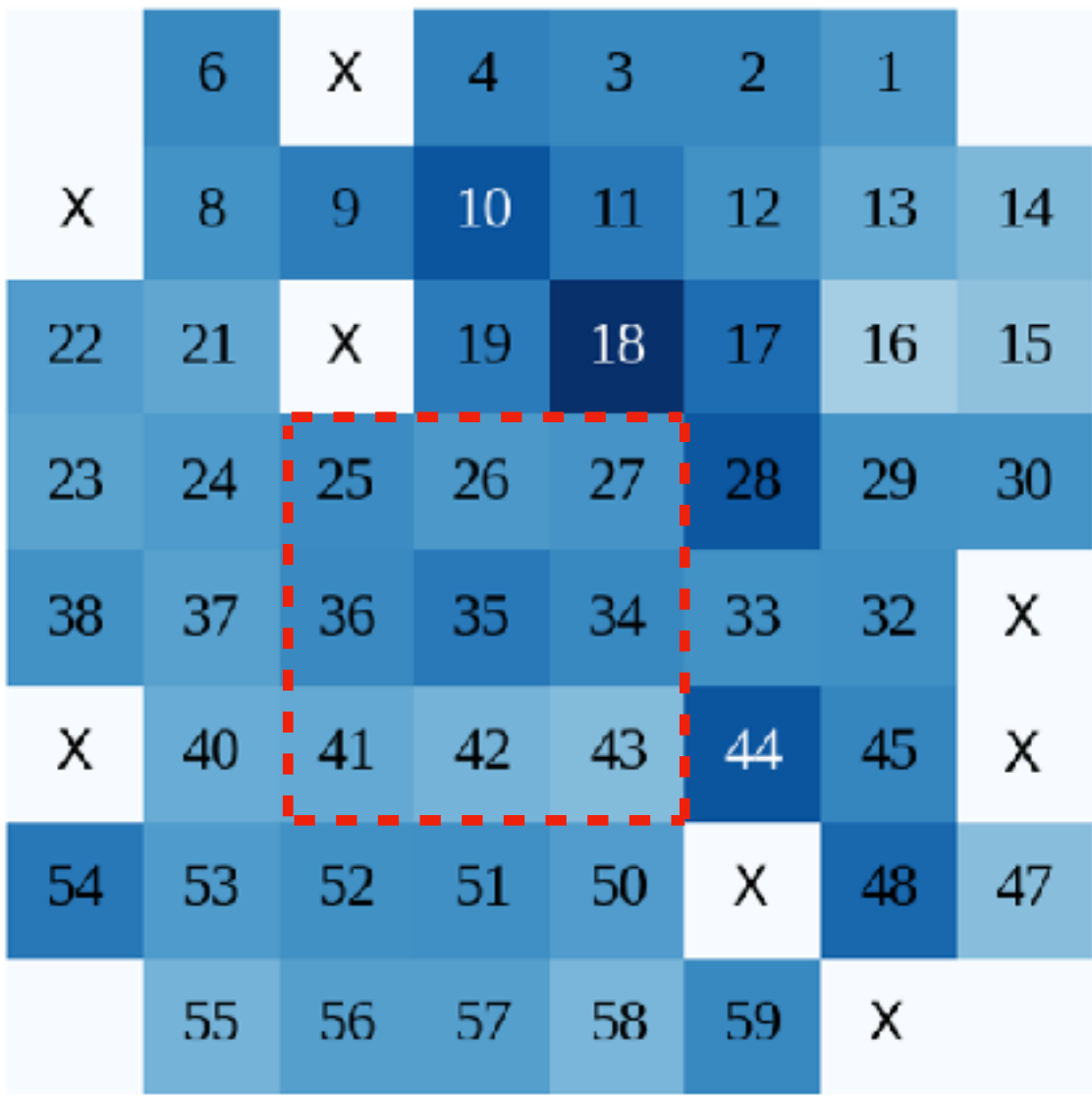
# Improvement of uniformity

First version of the array

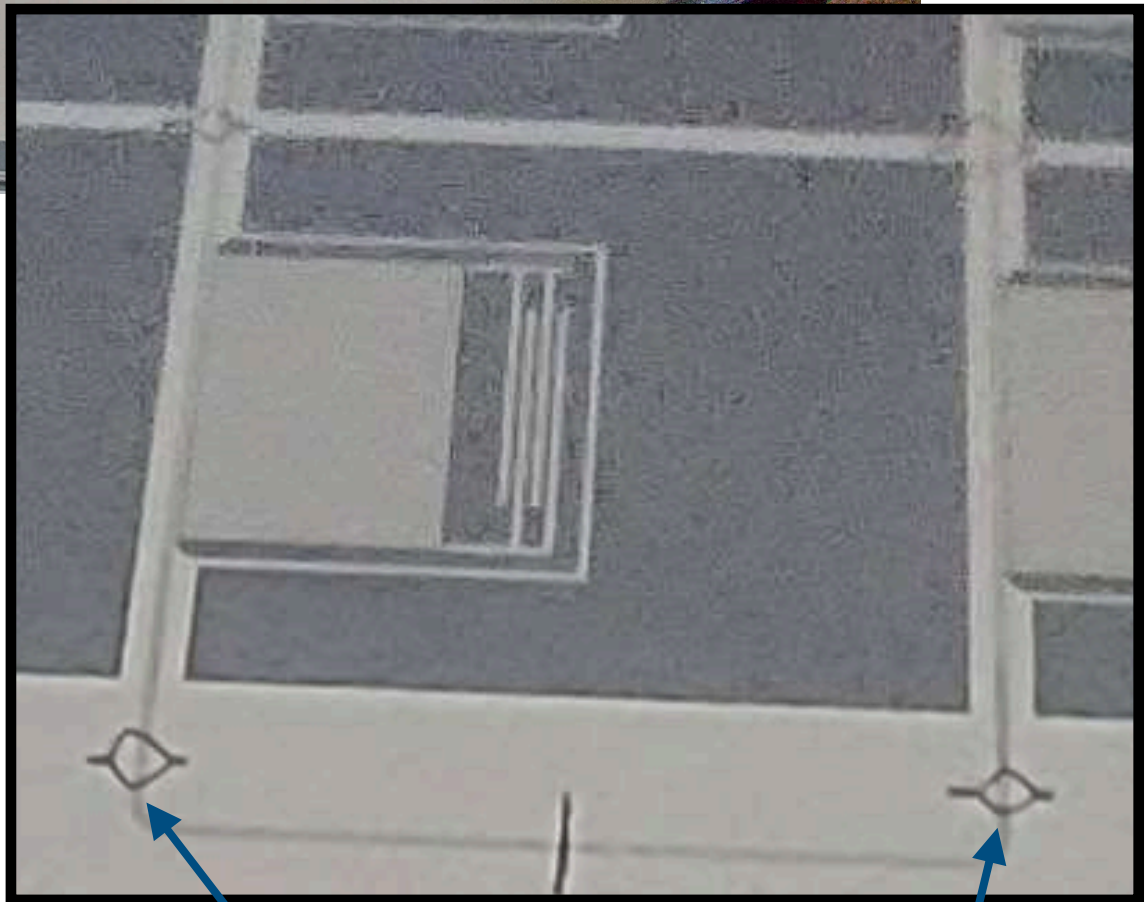
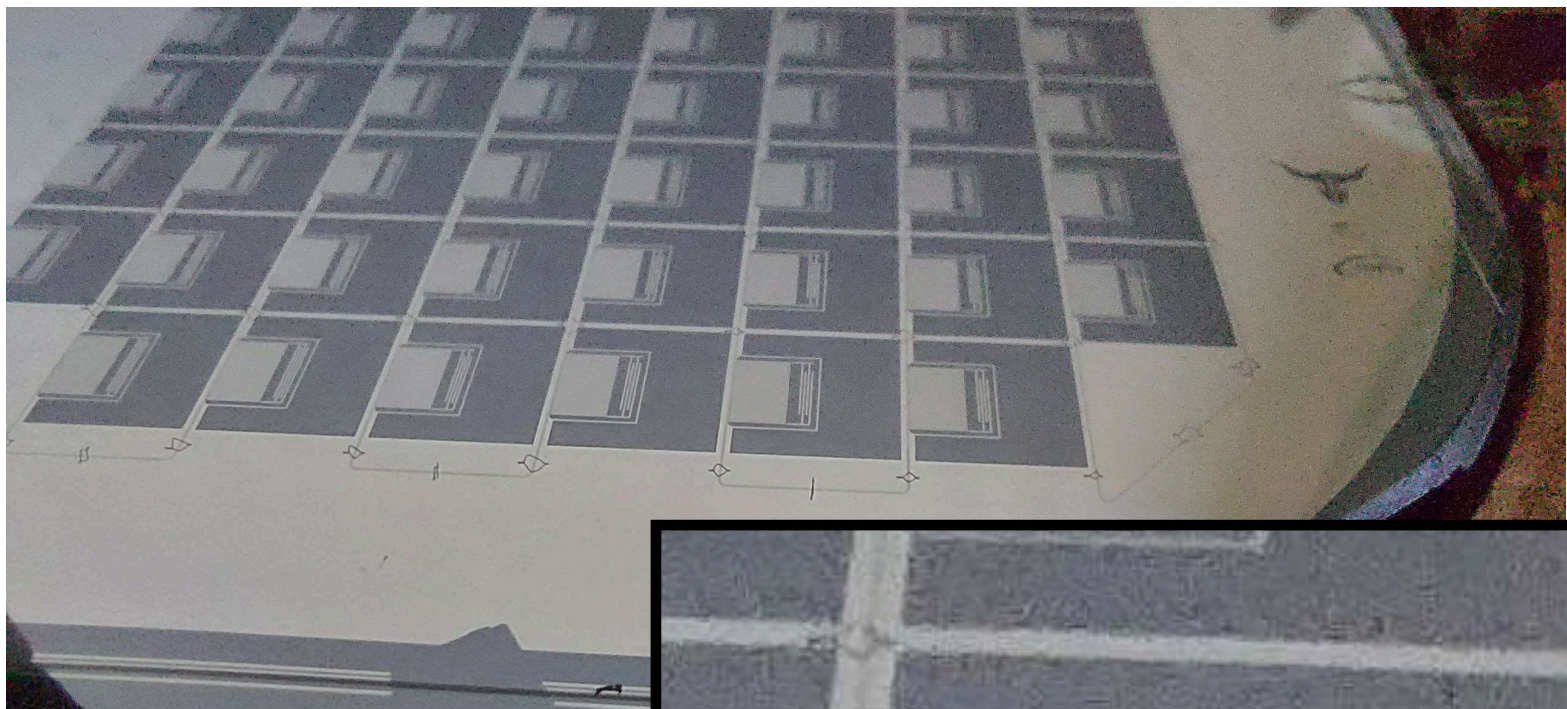


- + ) 58/60 KIDs alive
- ) Response not uniform

Same array with improved grounding



- + ) All KIDs with  $Q \sim 10^5$  (optimal sensitivity)
- ) Some resonator lost during operations

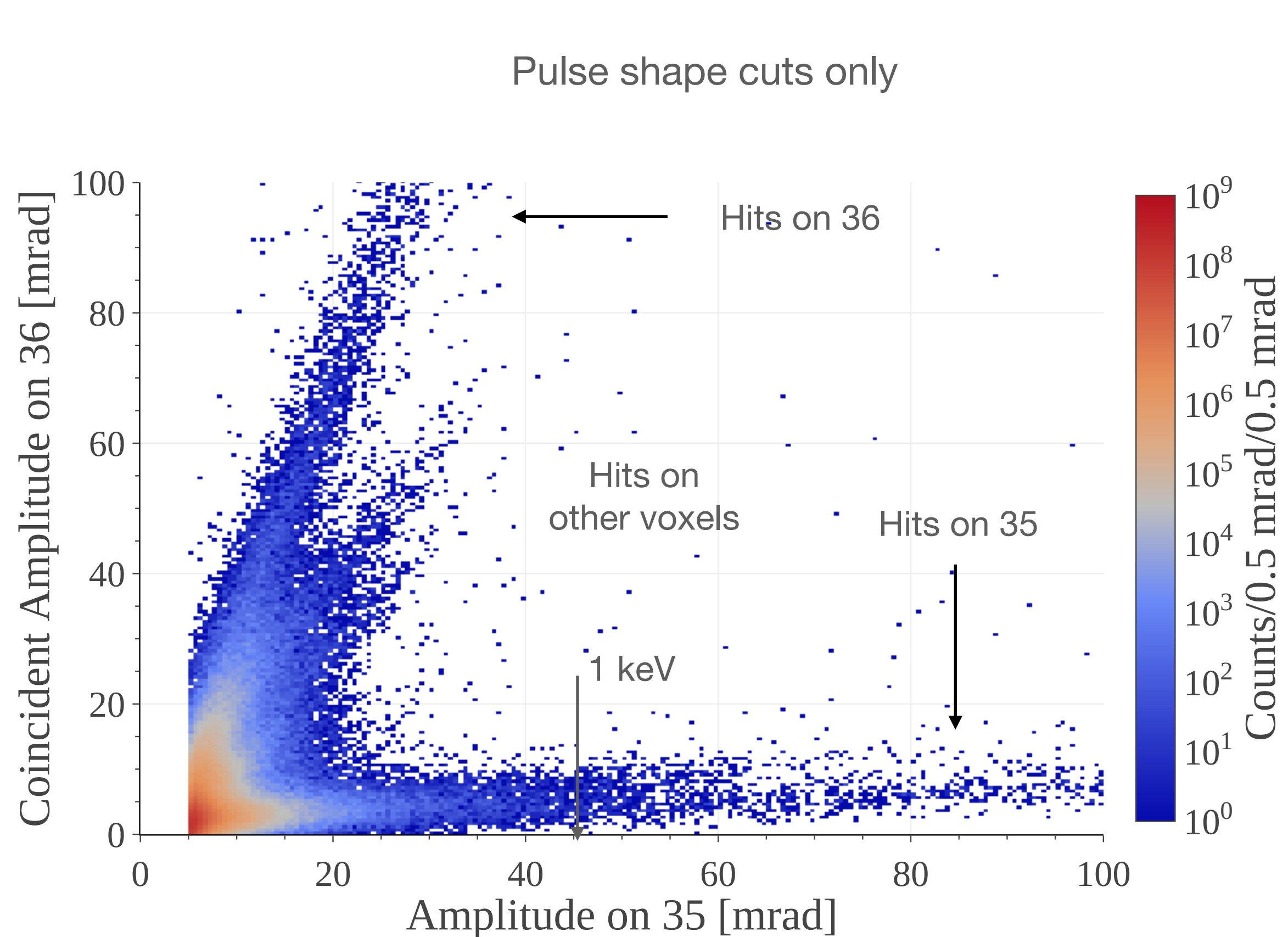


Airbridges connecting GND planes



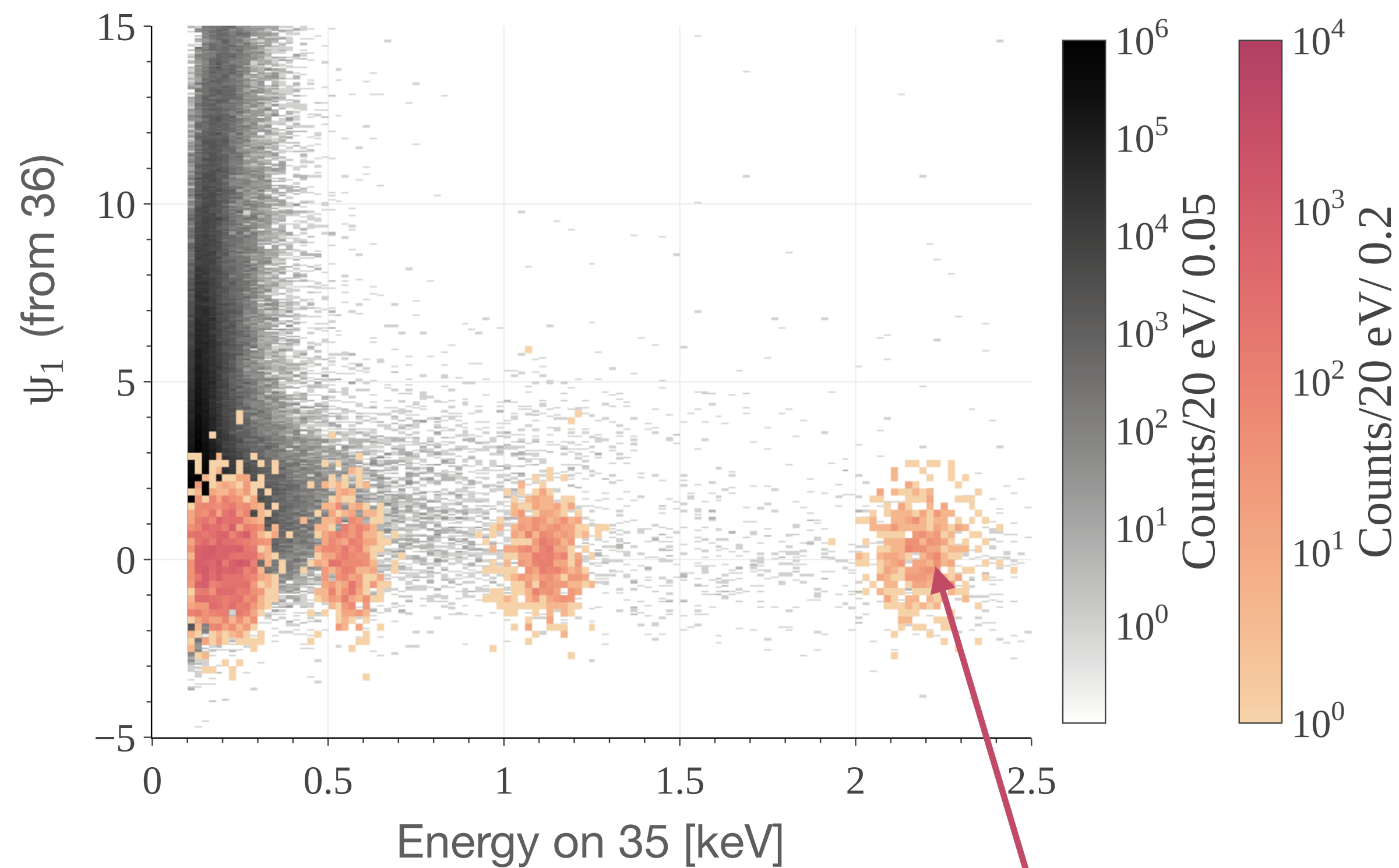
# Phonon leakage, selection

25	26	27
36	35	34
41	42	43



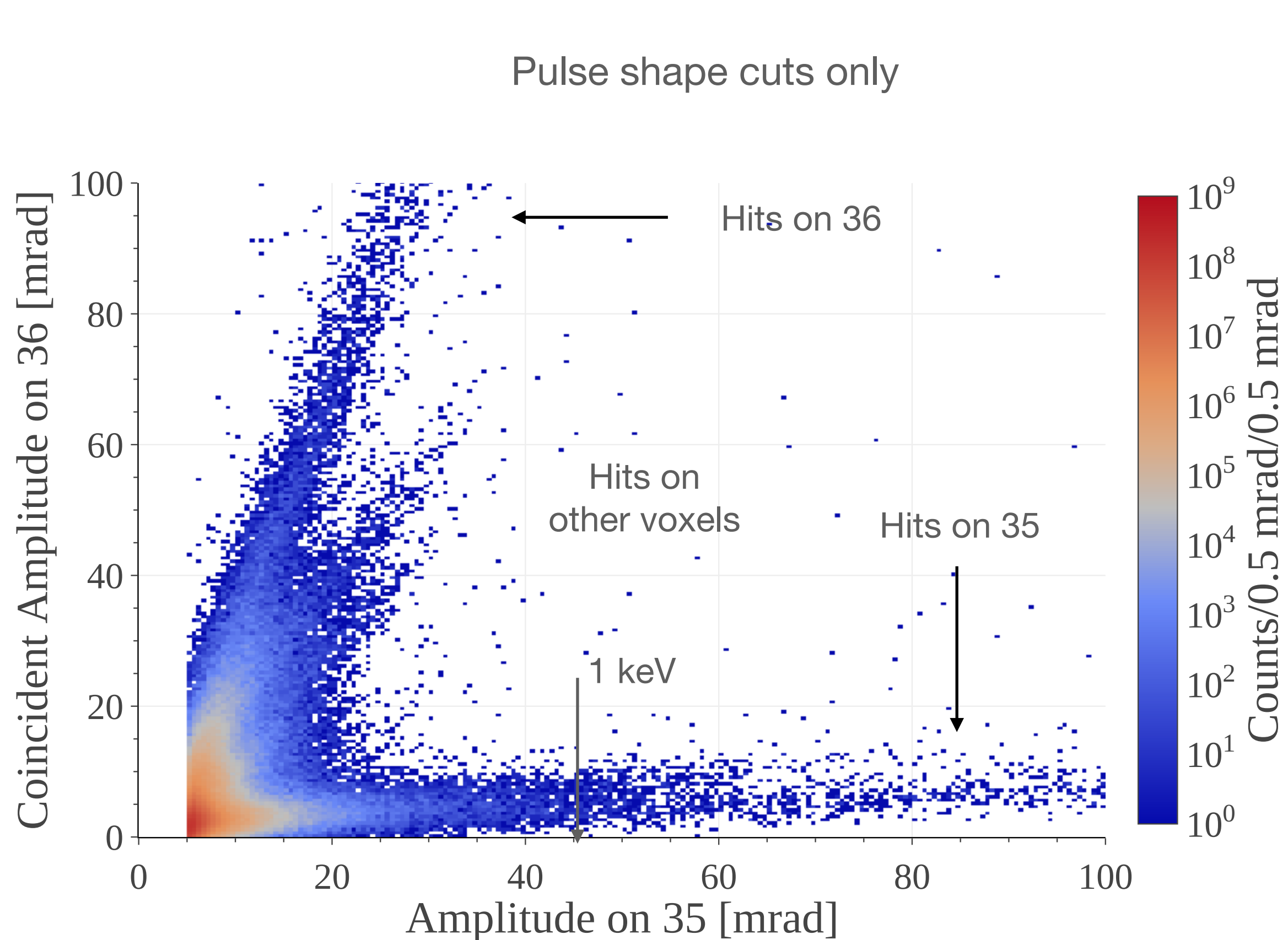
$$\psi_n = \frac{A_n - A \cdot r_n}{\sqrt{\sigma_{0,n}^2 + r_n^2 \cdot \sigma_0^2}}$$

$r_n$ : expected leakage



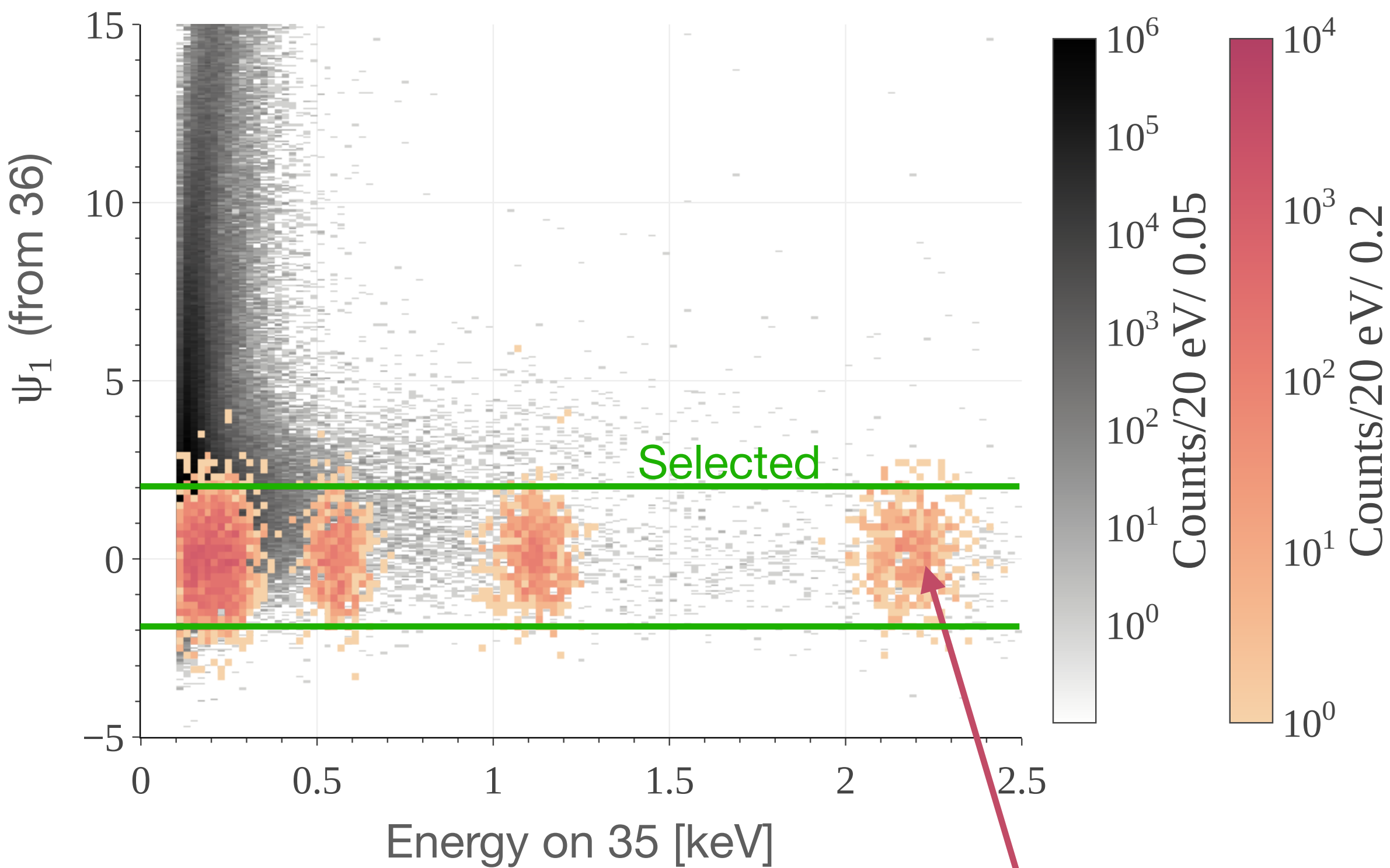
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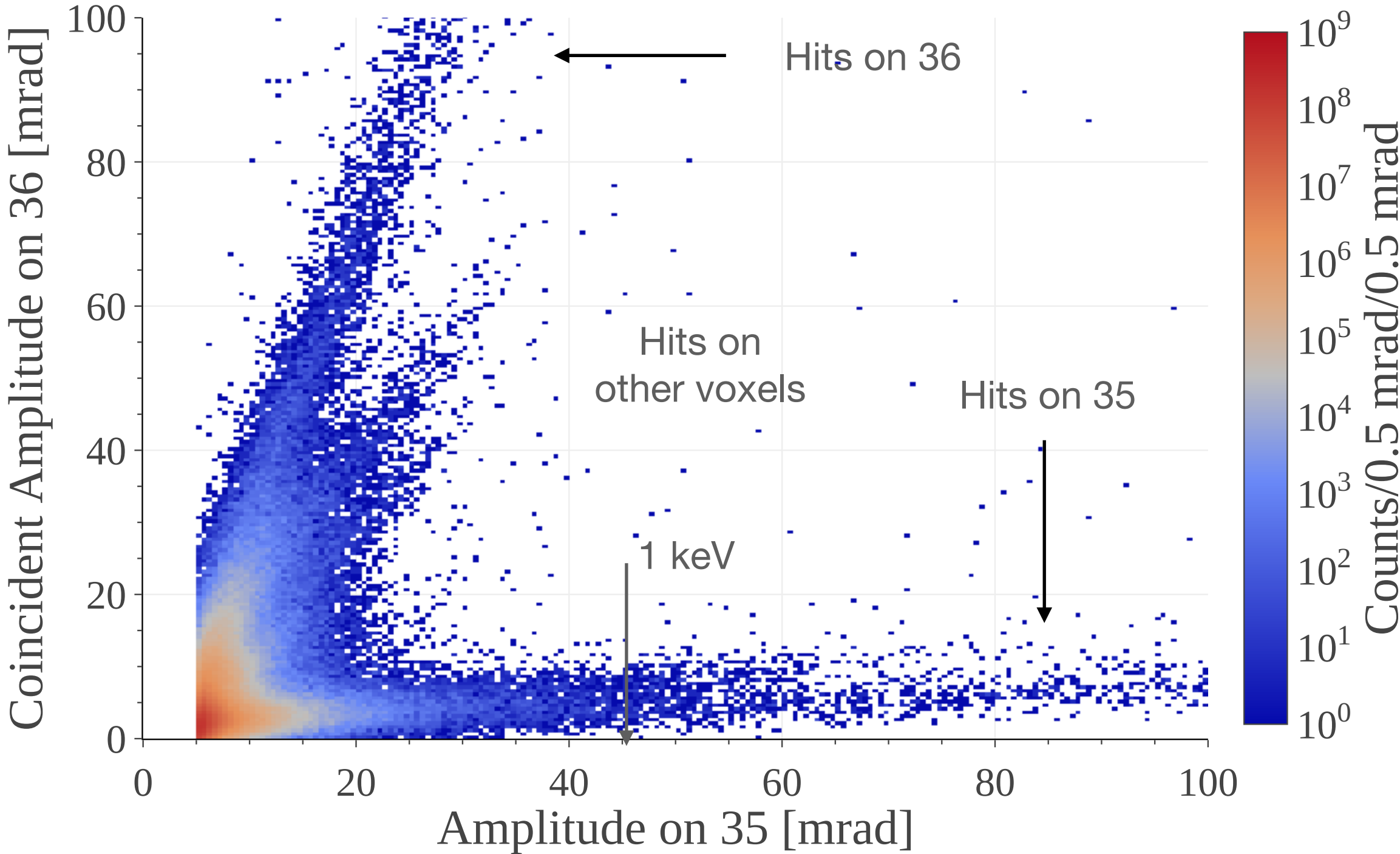




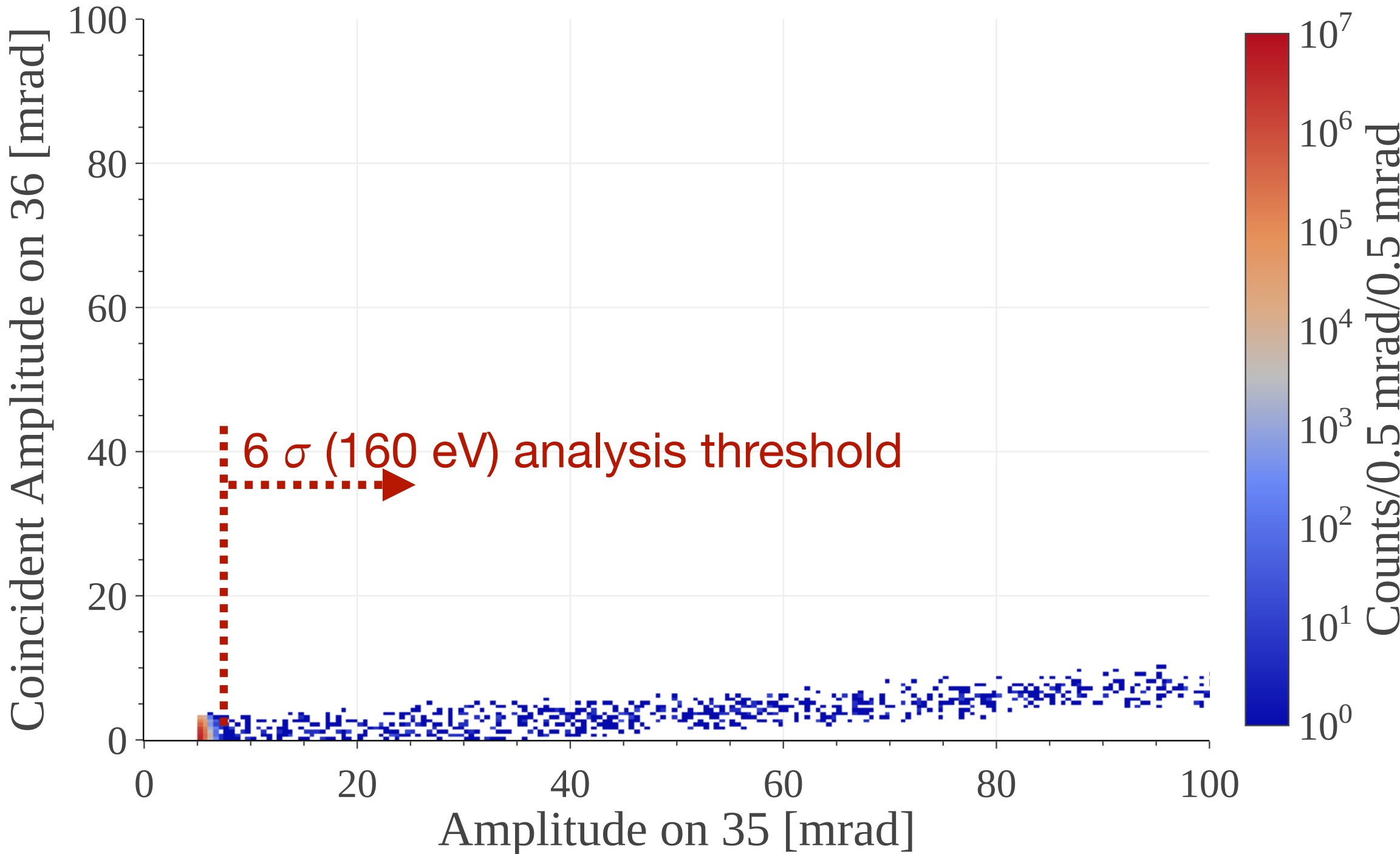
# Phonon leakage, selection

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Pulse shape cuts only



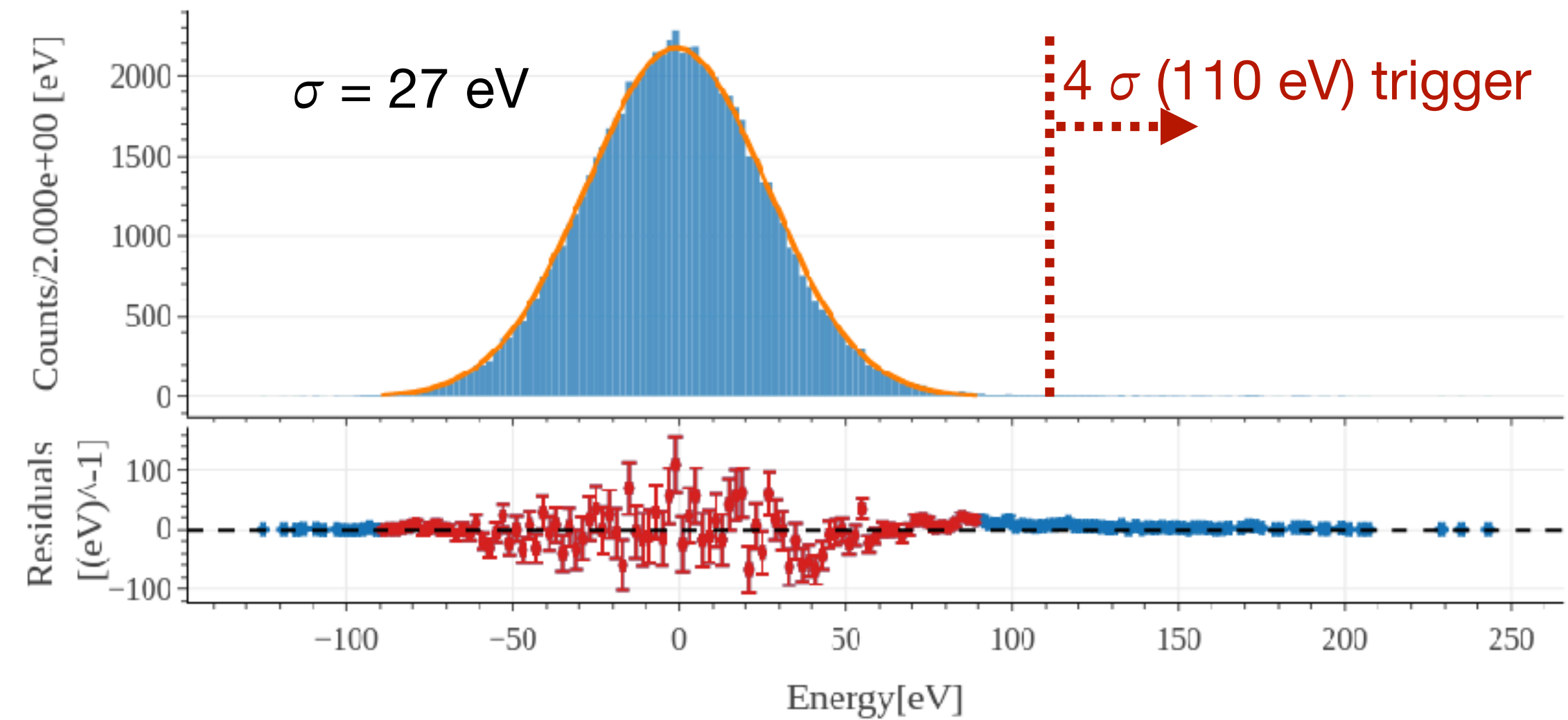
Pulse shape cuts  
+ coincident amplitude cut on 8 neighbours



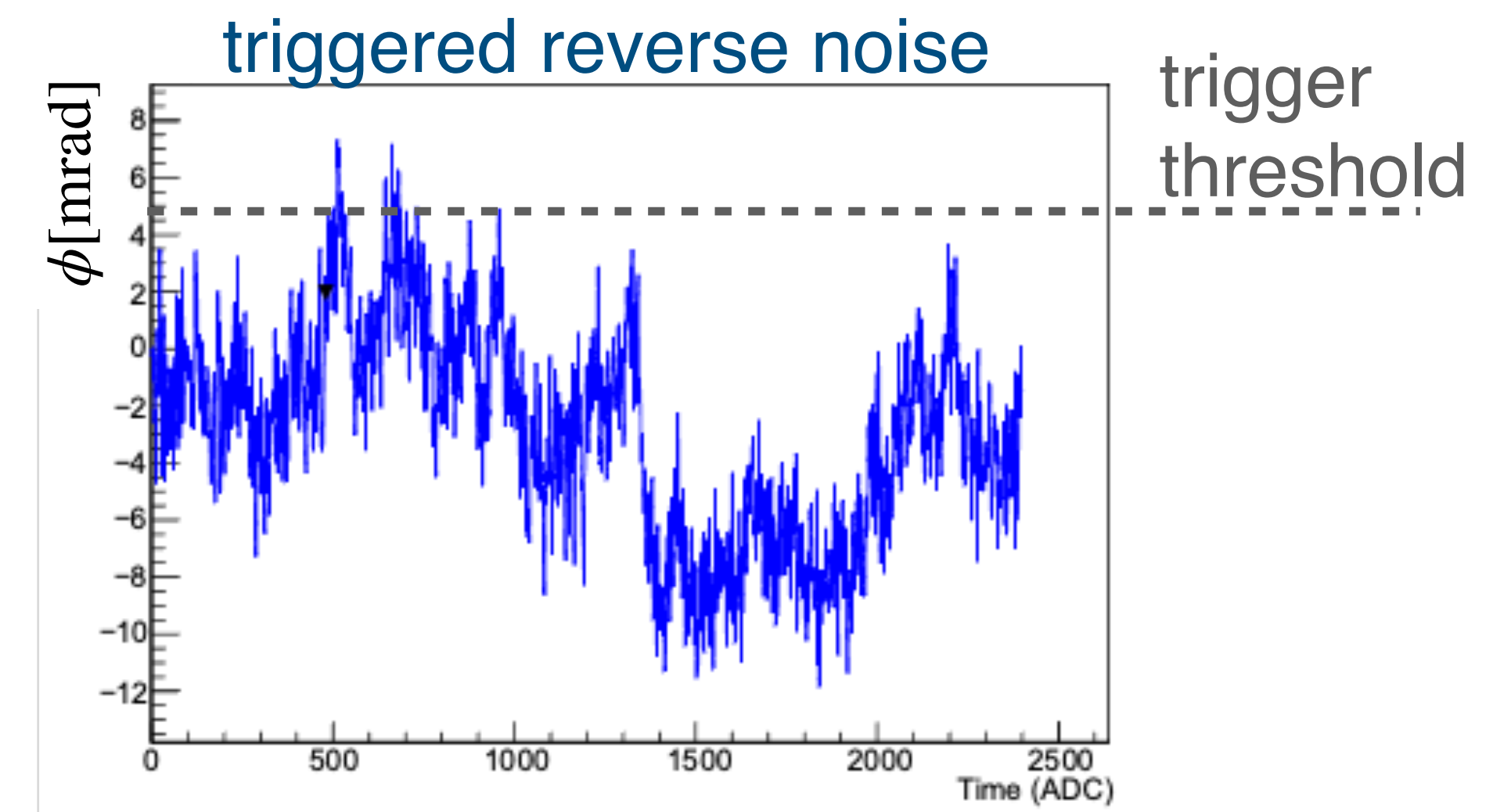
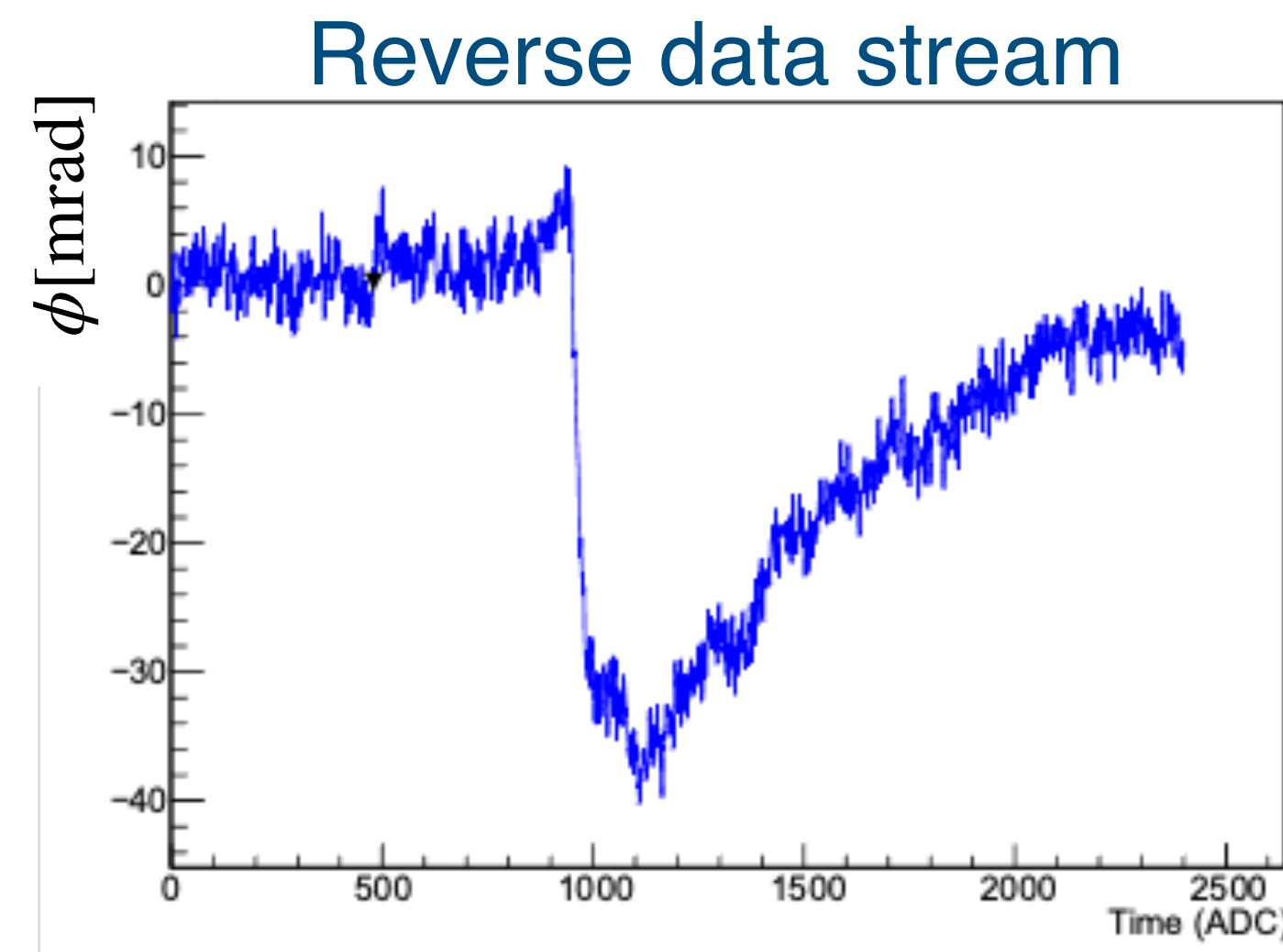
# Study of the energy threshold

Energy threshold usually set at 5 noise  $\sigma$

- Reduces the rate of of false triggers
- We choose  $4\sigma$  for the trigger

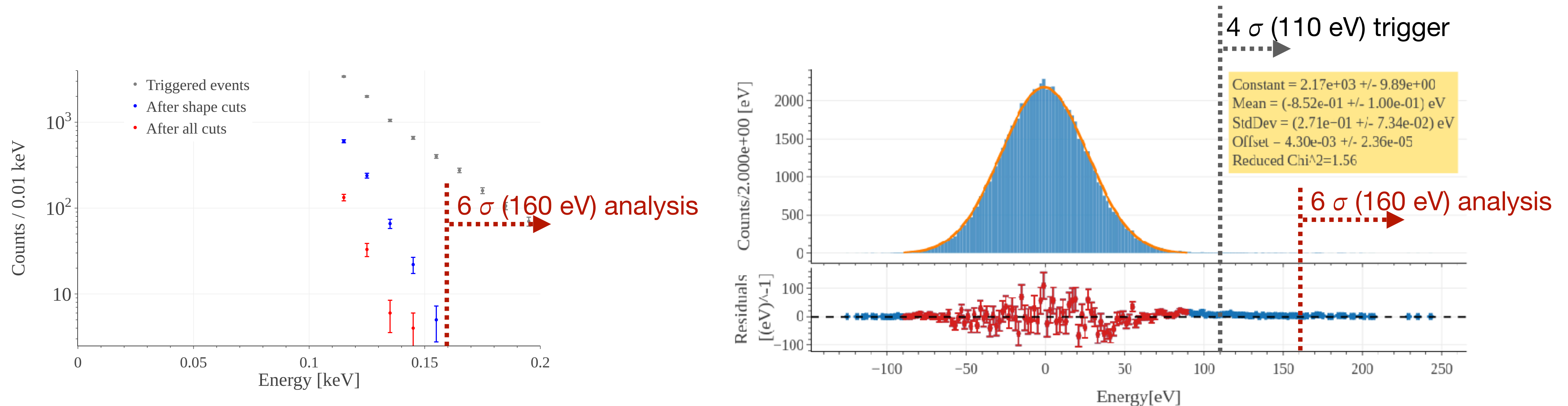


Statistics of false triggers from **reverse data stream** (minus sign on samples)





# Reverse trigger spectrum (30 mins live time)



From the spectrum of false positive reverse triggers surviving the cuts  
we set the analysis threshold to  $6\sigma$

# Towards the experiment

## MC Simulations

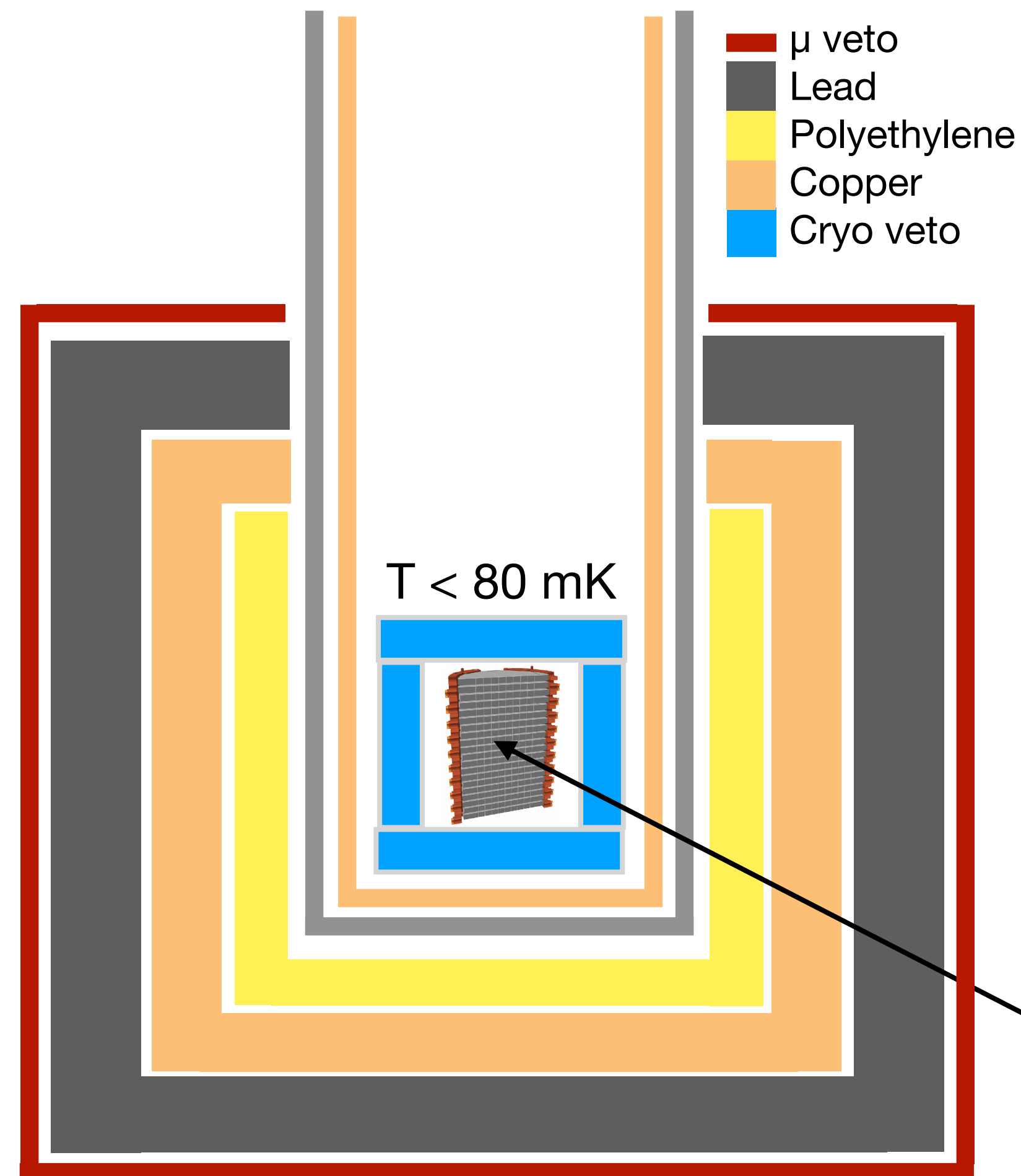
Design of the apparatus  
Definition of required radiopurity

## Apparatus

Cryostat outer shielding (PE, Pb, ...)  
Inner shielding  
Outer muon veto (scint. panels)?  
Cryo-veto around the BULLKIDs?  
(BGO + Light detector?)

## Energy calibration

Not possible with fibers:  
neutron recoils (a la CRAB)?  
Cs or Co Compton ?



## Underground cryo-infrastructure

Dilution refrigerator with  $T < 80 \text{ mK}$

## RF Readout

~20 RF lines,  
SDR boards with sync,  
trigger logic (clusters)

## DAQ

Data handling  
Data storage

## Data analysis

2k pixel,  
cluster analysis



**erc DANA E**  
stack of wafers

**Collaboration is forming and is open to contributions**