











# How to detect a QCD axion with MAD MAX

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On behalf of the MADMAX collaboration















Motivation for MADMAX

#### Latest news:

Run at CERN and monitoring system

New cryostat

MADMAX new magnet update

#### Next steps:

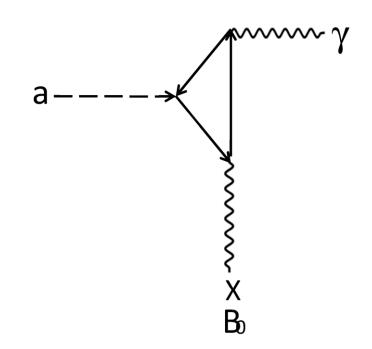
Cold operation of current prototype

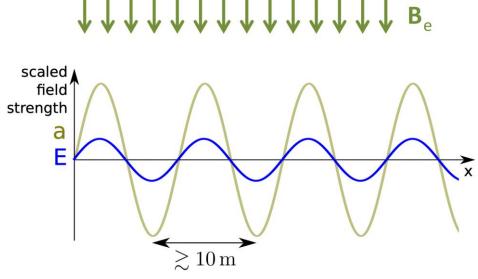
The next prototype

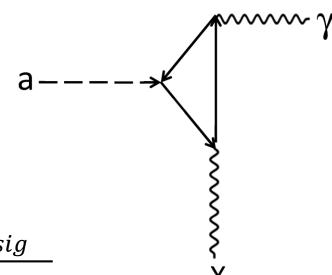
Prototype cryostat

#### Idea:

 Induce inverse Primakoff effect in a strong external B field



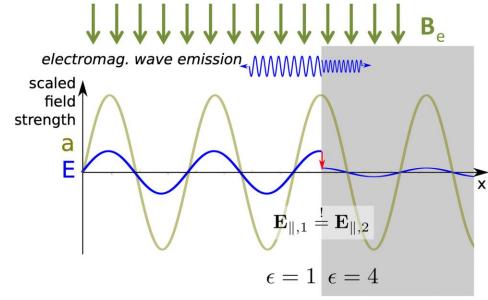


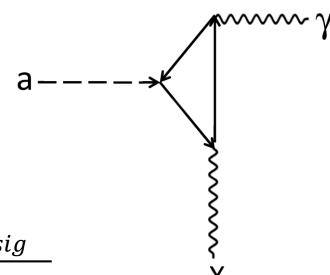


#### Idea:

- Induce inverse Primakoff effect in a strong external B field
- 2) Boost the signal using spatiallyperiodic dielectric discontinuities

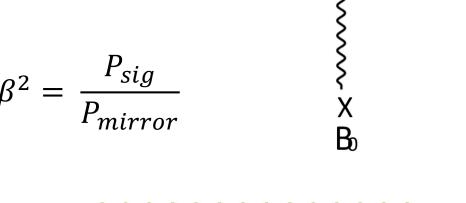
$$\beta^2 = \frac{P_{sig}}{P_{mirror}}$$

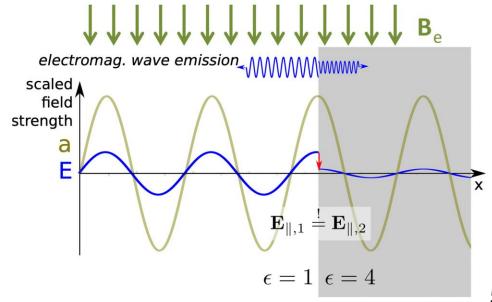




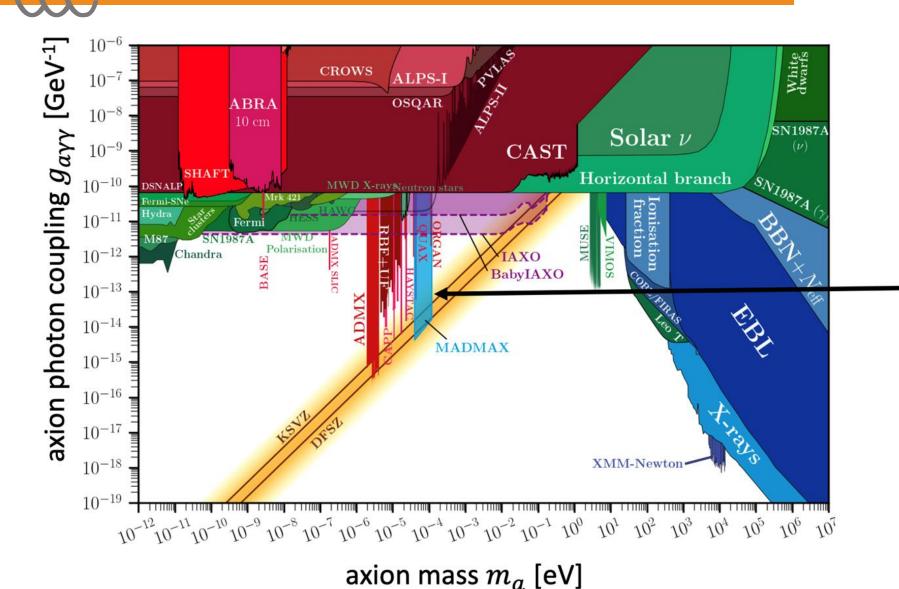
#### Idea:

- Induce inverse Primakoff effect in a strong external B field
- 2) Boost the signal using spatiallyperiodic dielectric discontinuities
- Reduce thermal background with cryogenics
- 4) Analyze boosted signal





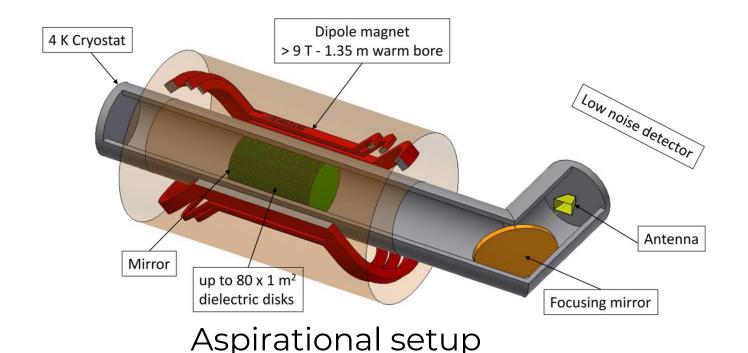
### Status of axion experiments

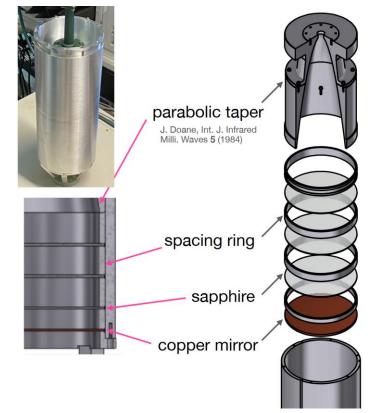




Aspirational setup

Probing the QCD axion between 40-400  $\mu$ eV



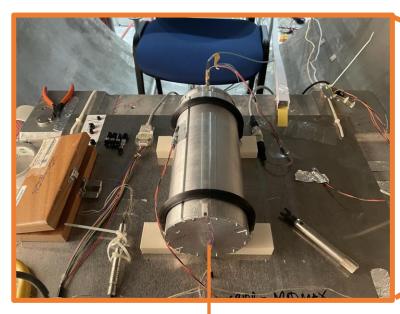


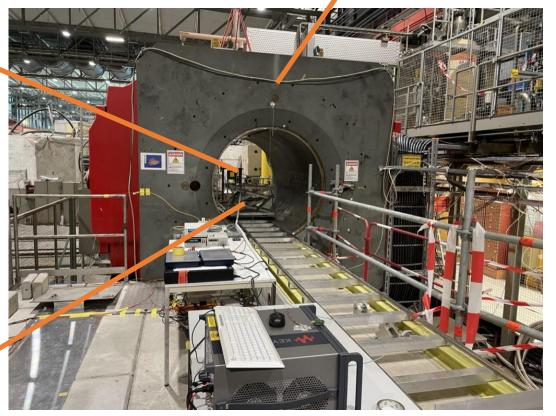
Prototype: CB-100

$$g_{a\gamma} = 0.2 m_a \left(\frac{50000}{\beta^2}\right)^{\frac{1}{2}} \left(\frac{10\text{T}}{B_e}\right) \left(\frac{1\text{m}^2}{A}\right)^{\frac{1}{2}} \left(\frac{SNR}{5}\right)^{\frac{1}{2}} \left(\frac{T_{\text{sys}}}{4\text{K}}\right) \left(\frac{1.8 \text{ days}}{\tau}\right)^{\frac{1}{4}} \left(\frac{\Delta \nu}{20 \text{ kHz}}\right) \left(\frac{300 \text{ MeV cm}^{-3}}{\rho}\right)^{\frac{1}{2}}$$

### **CERN runs**

Morpurgo magnet: 1.6 T dipole field

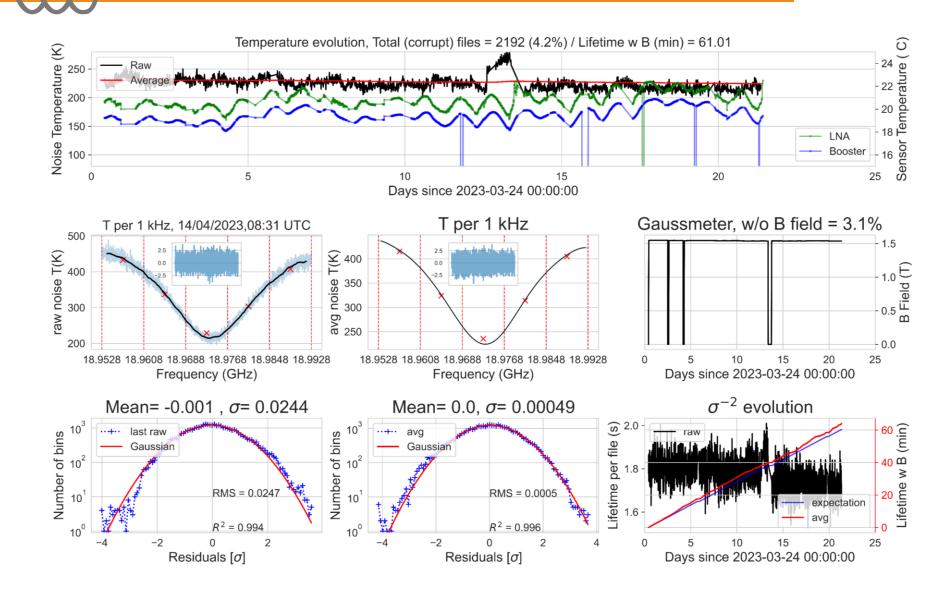




- CB-100 at room temperature
- ~10 hours integration time

Data analysis ongoing

### Monitoring system



# 24/7 experiment monitoring

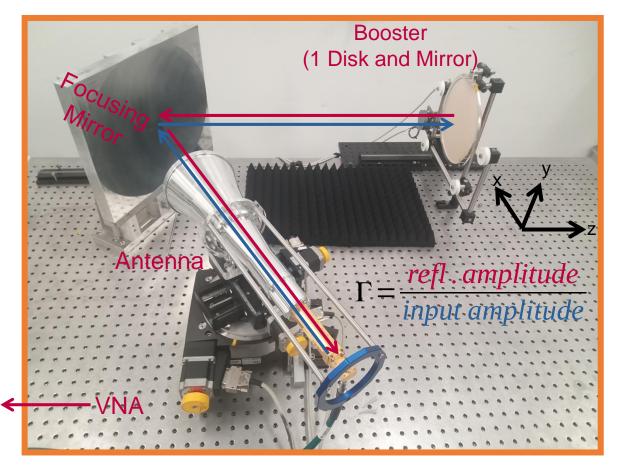
- Noise gaussianity
- Frequency dependent linearity
- Magnetic field
- Overheating
- Allan variance and total integration time
- Shifters incorporated
- E-mail alarms for urgent action (B field off, DAQ frozen, etc.)

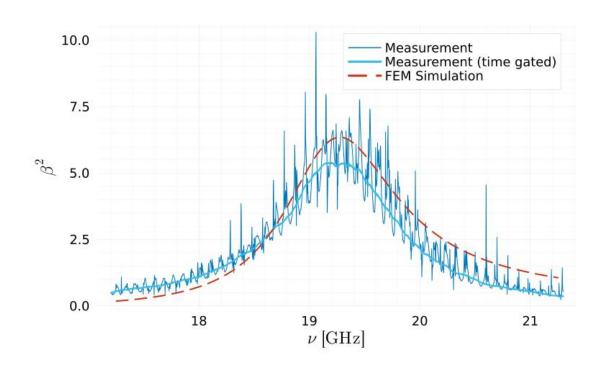
### **Boost factor determination**

Before: Only data-tuned simulations

Additional setup to determine  $\beta^2$  by direct measurement of the field

Now: measurement also possible via bead-pull method



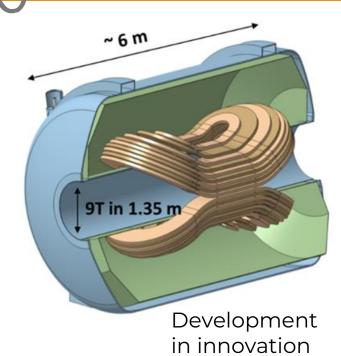


J. Egge (MADMAX): "Reciprocity approach"

<a href="https://iopscience.iop.org/article/10.1088">https://iopscience.iop.org/article/10.1088</a>

/1475-7516/2023/04/064

### MADMAX magnet update





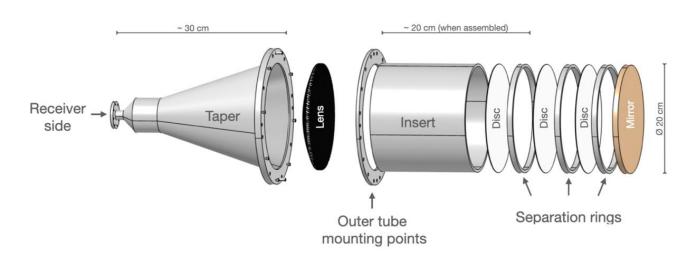




partnership

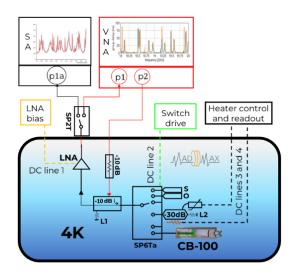
- Dipole Magnet most critical item for fullsize MADMAX
- Cable in conduit conductor (CICC) with a copper matrix
  - **→** production is feasible
- Quench propagation velocity was measured in dedicated setup
  - → Main project risk mitigated: Quench propagation according to requirements for safe operation

# Next steps



### Next prototype: CB-200

Gain in sensitivity of ~40%

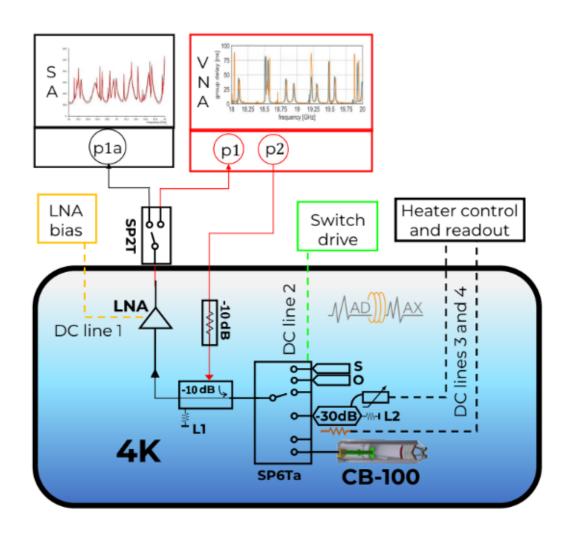


### Cold (4K) run with CB-100

Gain in sensitivity of ~1 order of magnitude

For more information see poster "Towards a cryogenic calibration of a dielectric haloscope for direct dark matter detection" – Juan PA M.

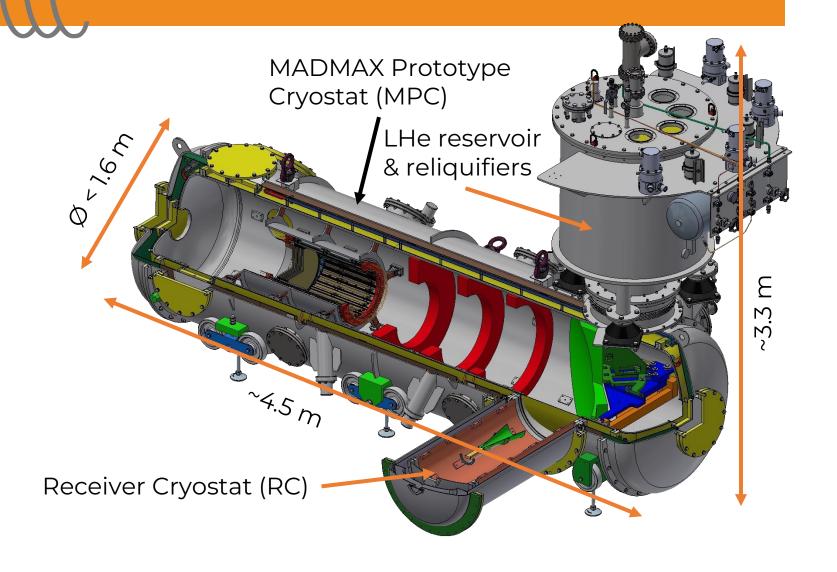
### Cold operation idea



Single thermal cycle semiautomatic calibration.

Maximizes stability and reproducibility of our measurements

### Plans for 2024-2025



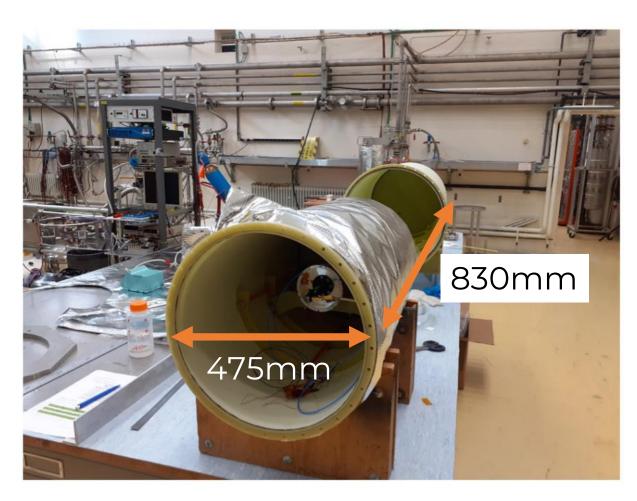
#### Prototype cryostat

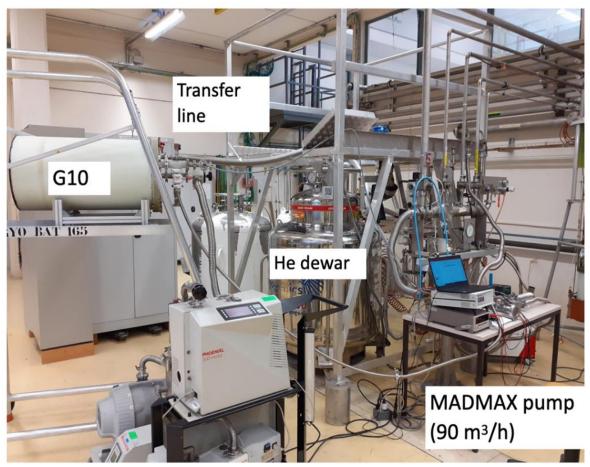
- Delivery expected beginning of 2024
- Commissioning site: Hamburg
- Planned ALP search at CERN in 2025

### G10 cryostat









Stability reached: 24 hours at 10K

# Summary

MADMAX will search for axions between 40-400  $\mu$ eV

First runs with CB-100 at 300K done; data analysis ongoing

Magnet feasibility confirmed

Prototype cryostat soon to be available

G10 cryostat tested and ready to use inside magnet

First cryogenic operation intended for 2024 at CERN

Upgrade of the prototype also planned for 2024