

Direct Dark Matter search with the CRESST-III experiment

Status and prospects

XVIII International Conference on Topics in Astroparticle and Underground Physics 2023

Margarita Kaznacheeva | Technical University of Munich

margarita.kaznacheeva@tum.de

On behalf of the CRESST Collaboration

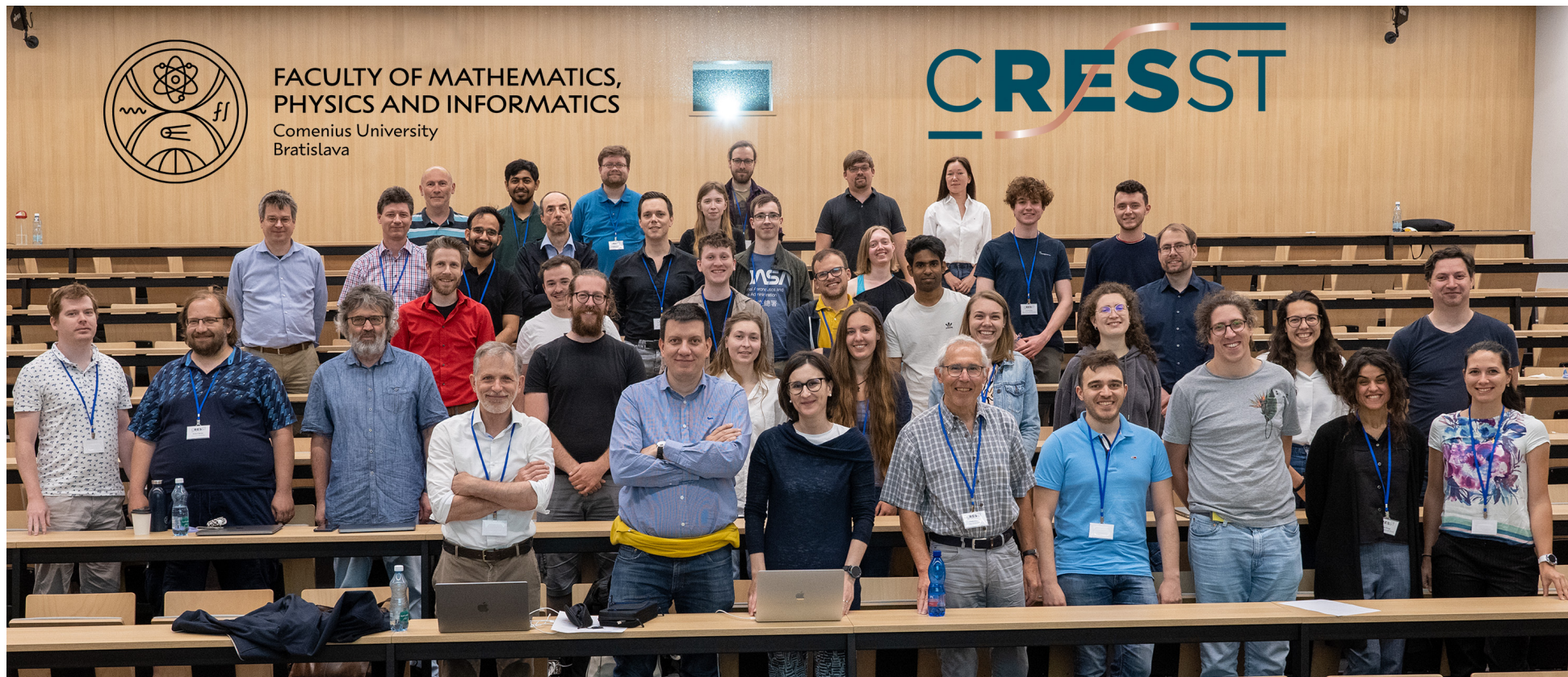


Cryogenic **Rare Event Search**
with Superconducting Thermometers



CRESST collaboration

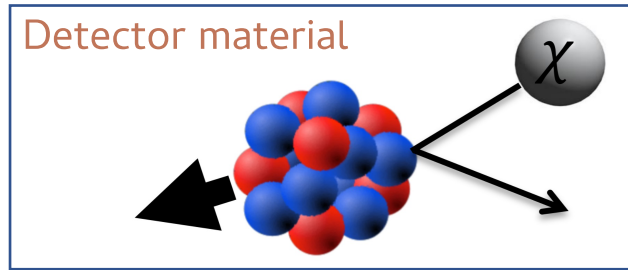
~60 people from 8 institutions from 5 countries in Europe



June 2023, CRESST collaboration meeting in Bratislava

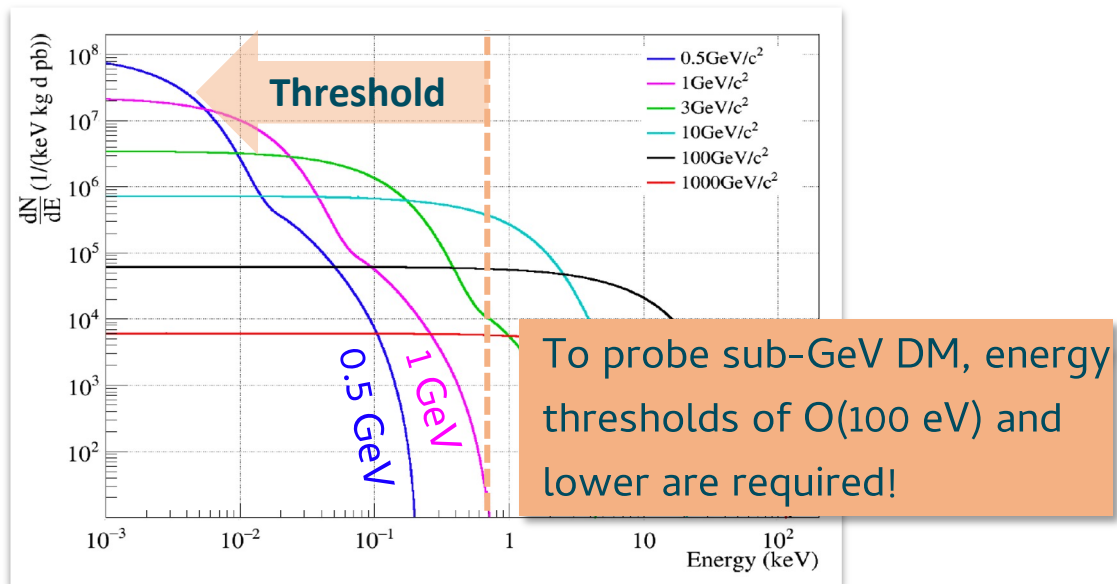
The CRESST experiment searches for sub-GeV DM particles.

Interaction:



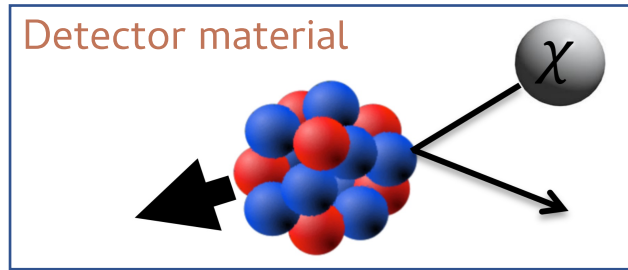
Signal: nuclear recoil energy

Expected differential energy spectrum in CaWO₄



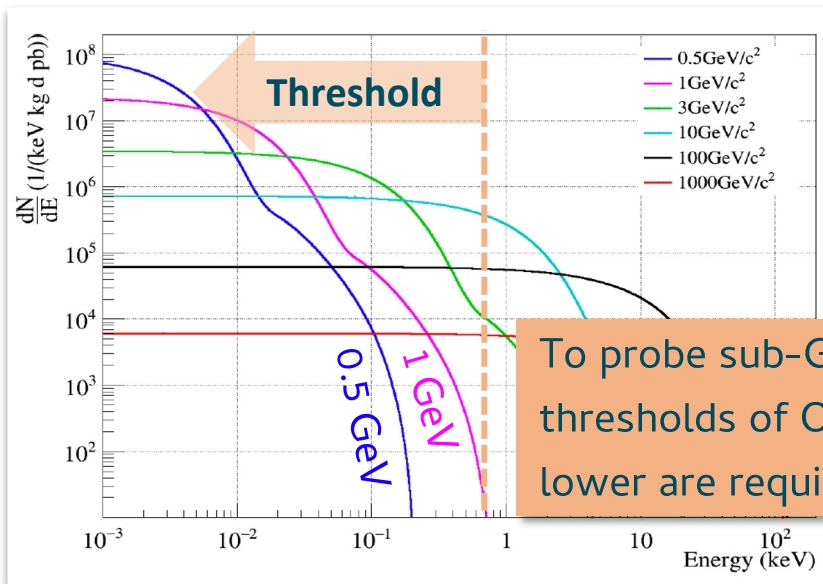
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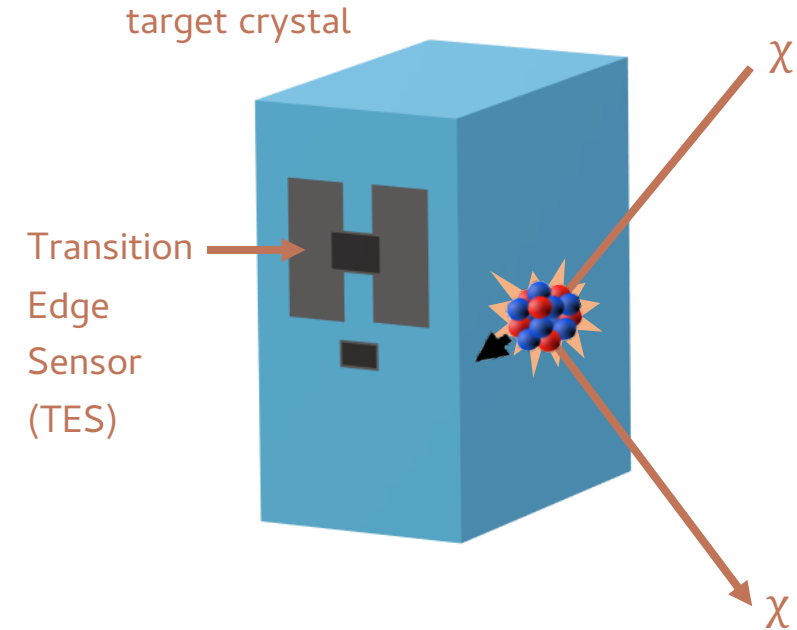
Signal: nuclear recoil energy

Expected differential energy spectrum in CaWO₄



To probe sub-GeV DM, energy thresholds of O(100 eV) and lower are required!

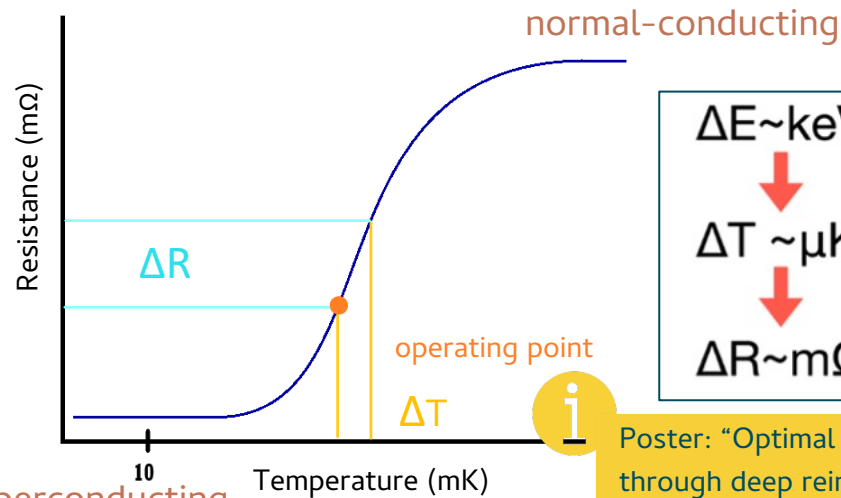
Detector:



Target: various crystal materials (CaWO₄, Al₂O₃, LiAlO₂, Si)

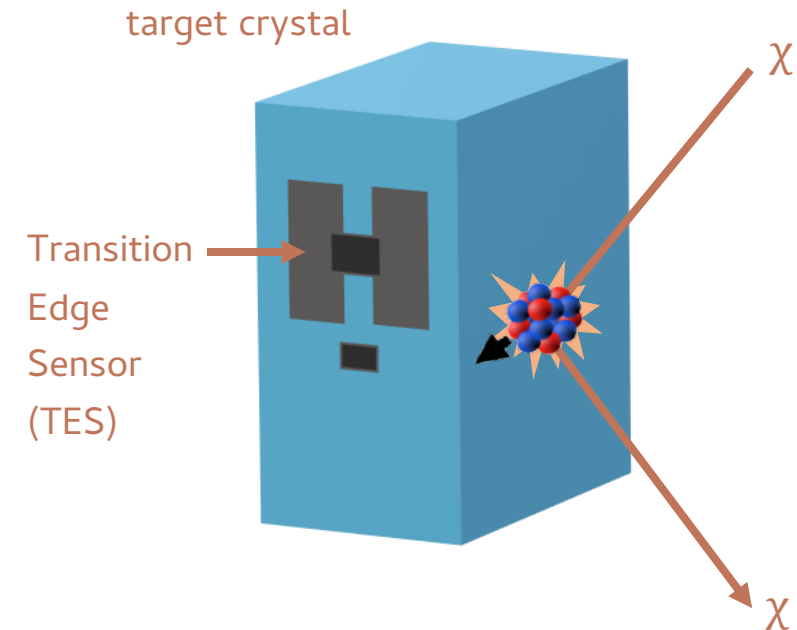
Small energy deposition in the crystal leads to a measurable TES resistance change.

Detector:



$$\begin{aligned} \Delta E &\sim \text{keV} \\ \downarrow \\ \Delta T &\sim \mu\text{K} \\ \downarrow \\ \Delta R &\sim \text{m}\Omega \end{aligned}$$

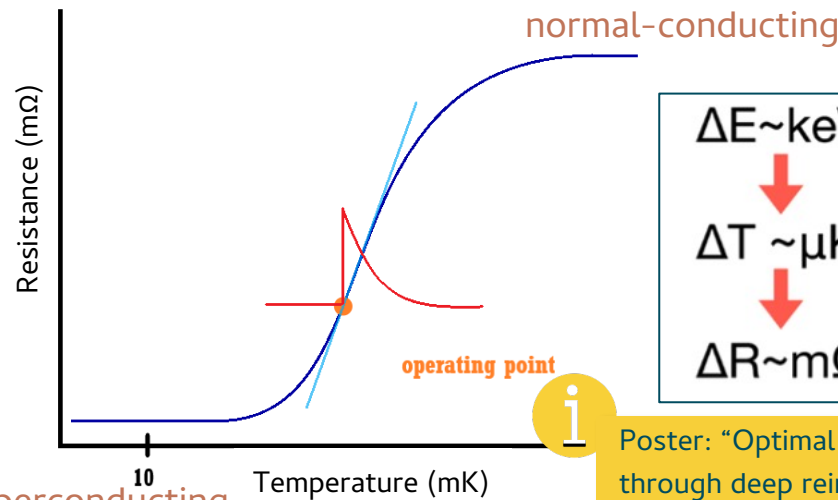
Poster: "Optimal operation of cryogenic calorimeter through deep reinforcement learning" by Felix Wagner



Target: various crystal materials (CaWO_4 , Al_2O_3 , LiAlO_2 , Si)
Sensor: W-TES at 15 mK

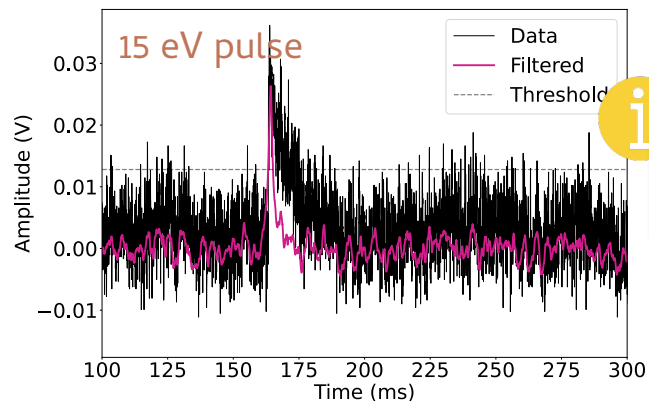
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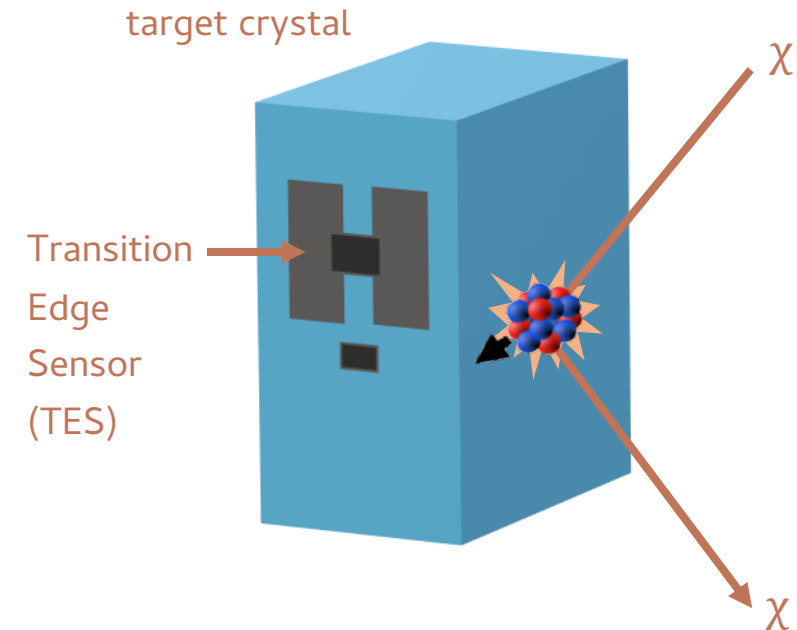
Poster: "Optimal operation of cryogenic calorimeter through deep reinforcement learning" by Felix Wagner



Pulse height is a measure of energy deposition.

Poster: "Optimum Filter Analysis in CRESST-III" by Lena Meyer

Energy calibration with X-rays from a ^{55}Fe source@6 keV

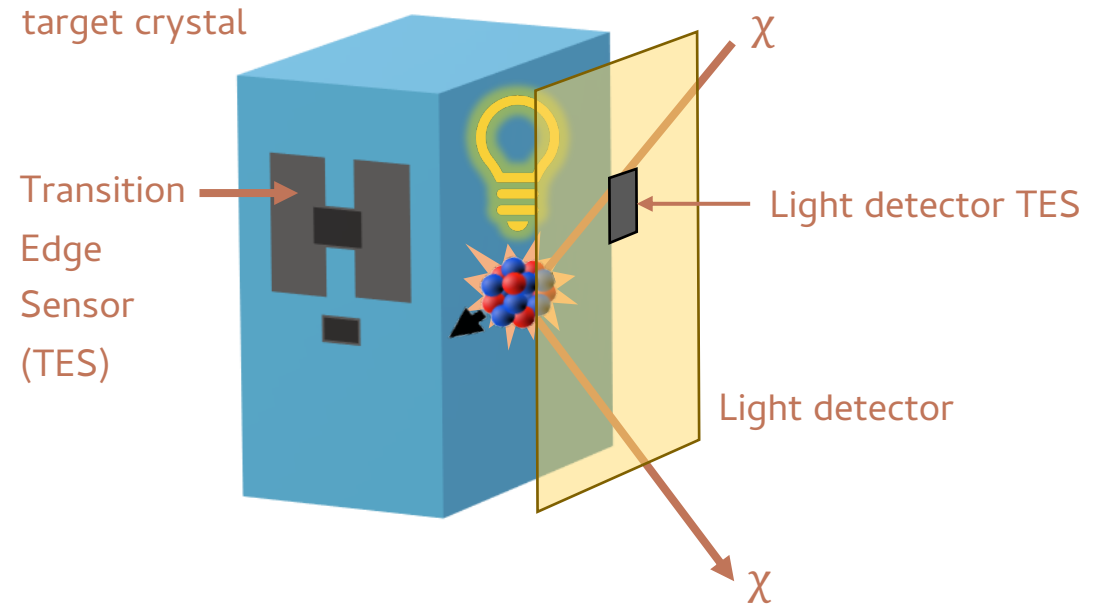
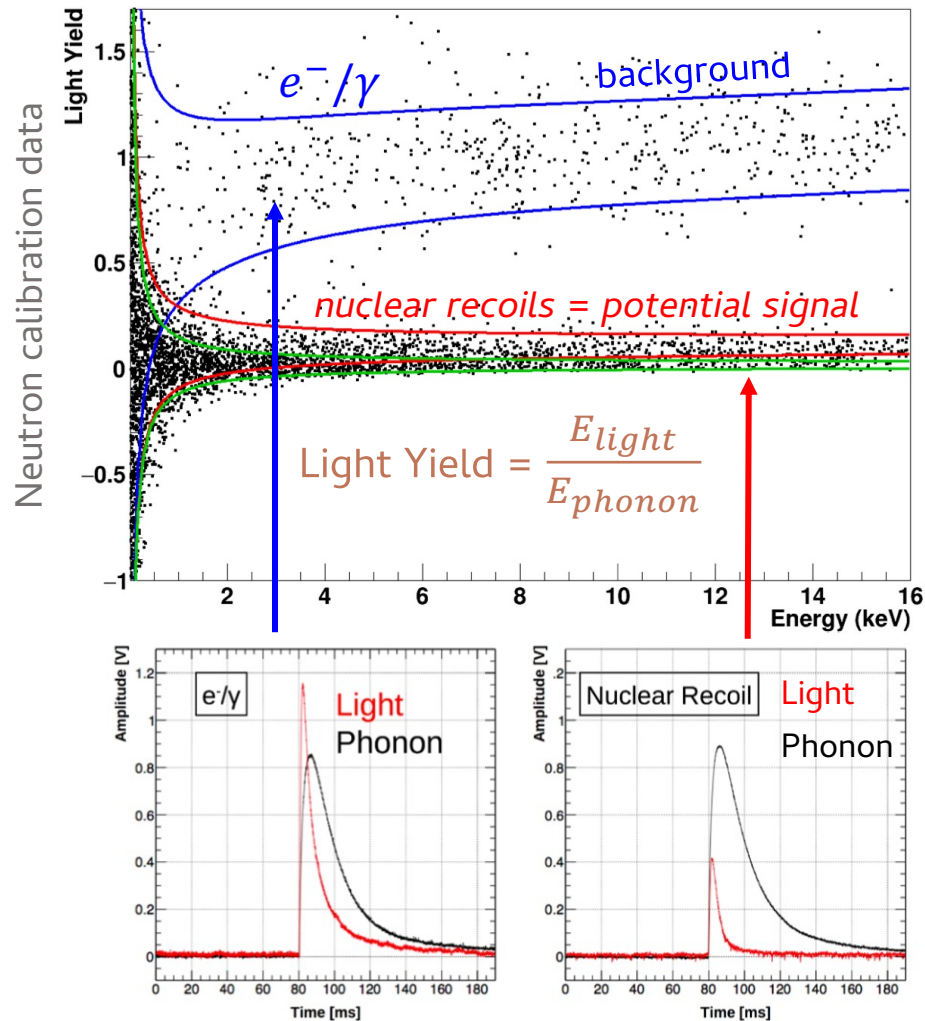


Target: various crystal materials (CaWO_4 , Al_2O_3 , LiAlO_2 , Si)

Sensor: W-TES at 15 mK

Energy threshold of O(10 eV) ✓

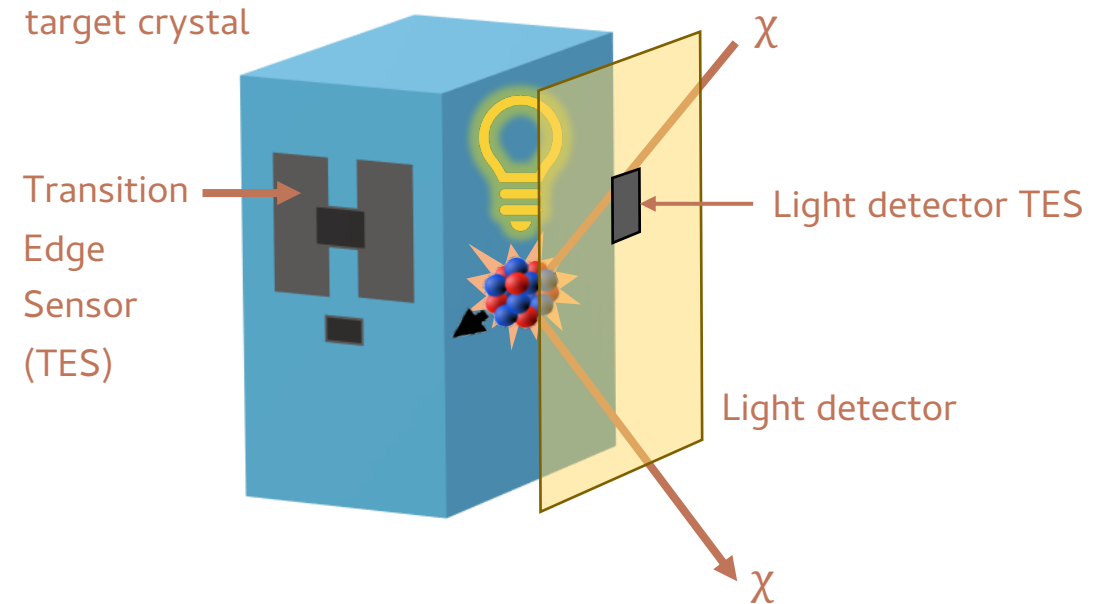
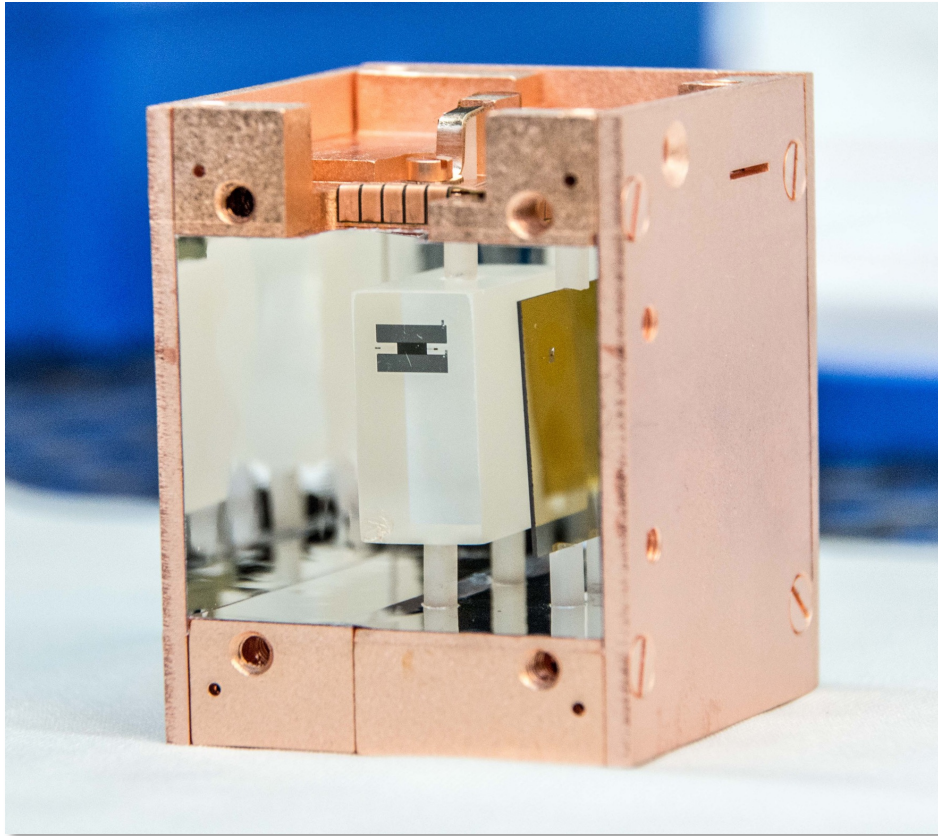
Additional light signal allows to discriminate electron and nuclear recoils to reduce background signal.



For scintillating target materials (e.g. CaWO_4), a small part of the deposited energy is released as **light**.

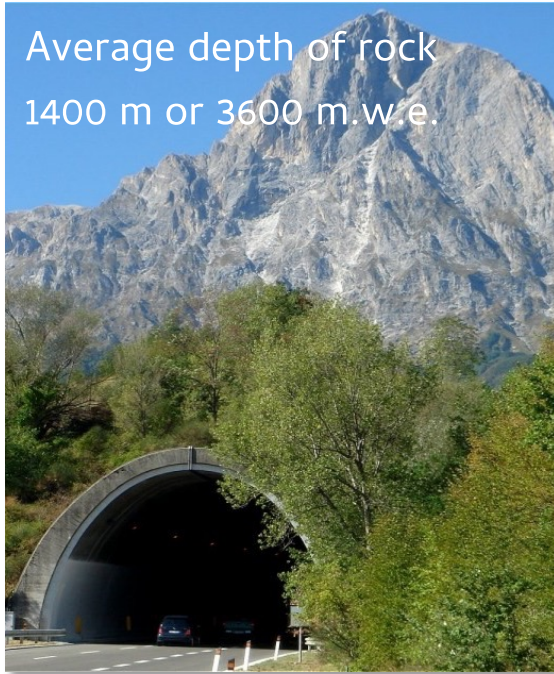
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CRESST-III detector module



For scintillating target materials (e.g. CaWO_4), a small part of the deposited energy is released as *light*.

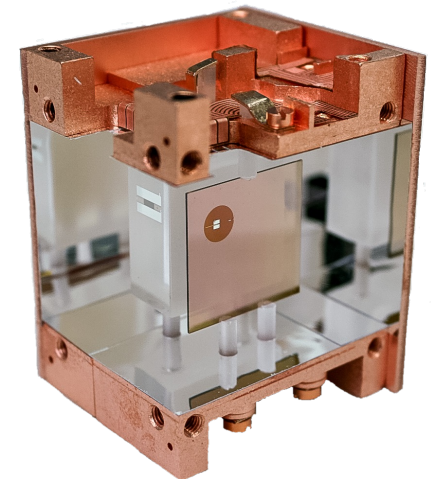
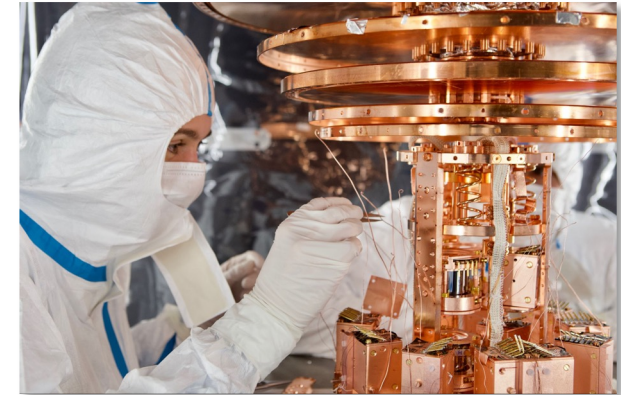
The CRESST experiment is located in the deep underground laboratory LNGS in Italy that provides excellent shielding against cosmic radiation.



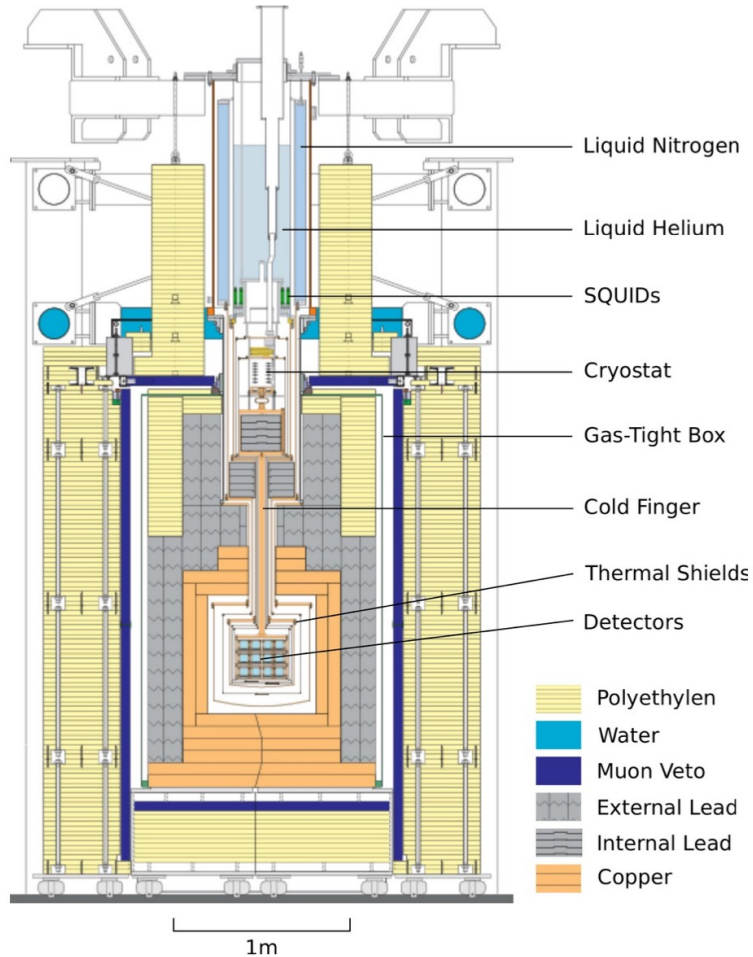
Average depth of rock
1400 m or 3600 m.w.e.



Muon flux: $3 \cdot 10^{-8}/(\text{cm}^2 \text{ s})$
Neutrons: $4 \cdot 10^{-6}/(\text{cm}^2 \text{ s})$



Additional layers of shielding materials and an active muon veto further reduce backgrounds.



A detailed Geant4-based background model is continuously adapted and improved.

i Talk: "Background modeling and simulation of calibration source for the dark matter search experiment CRESST" by Samir Banik (Tue)

i Talk: "Extension of the Geant4 based Electromagnetic Background Model of CRESST-II and CRESST-III" by Jens Burkhart (Tue)

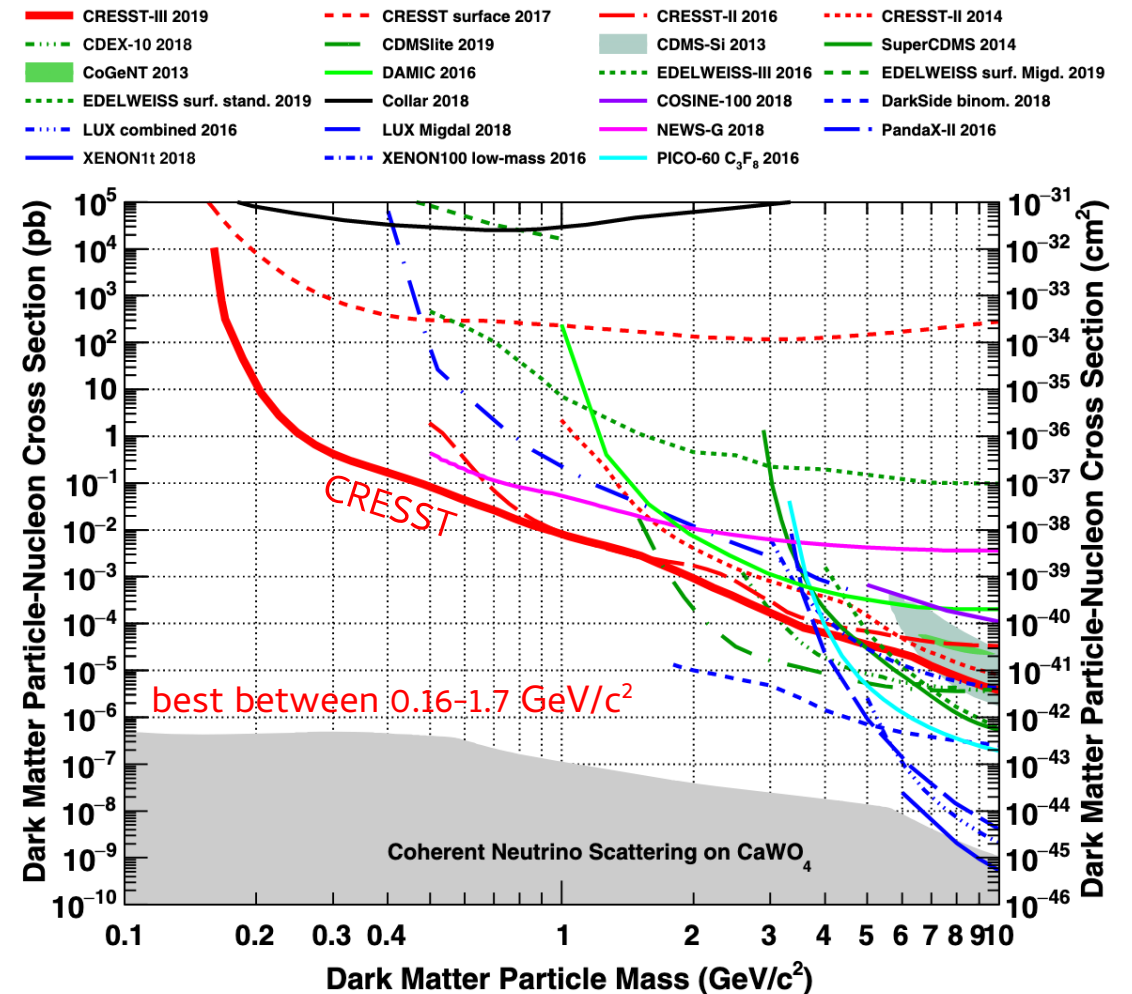
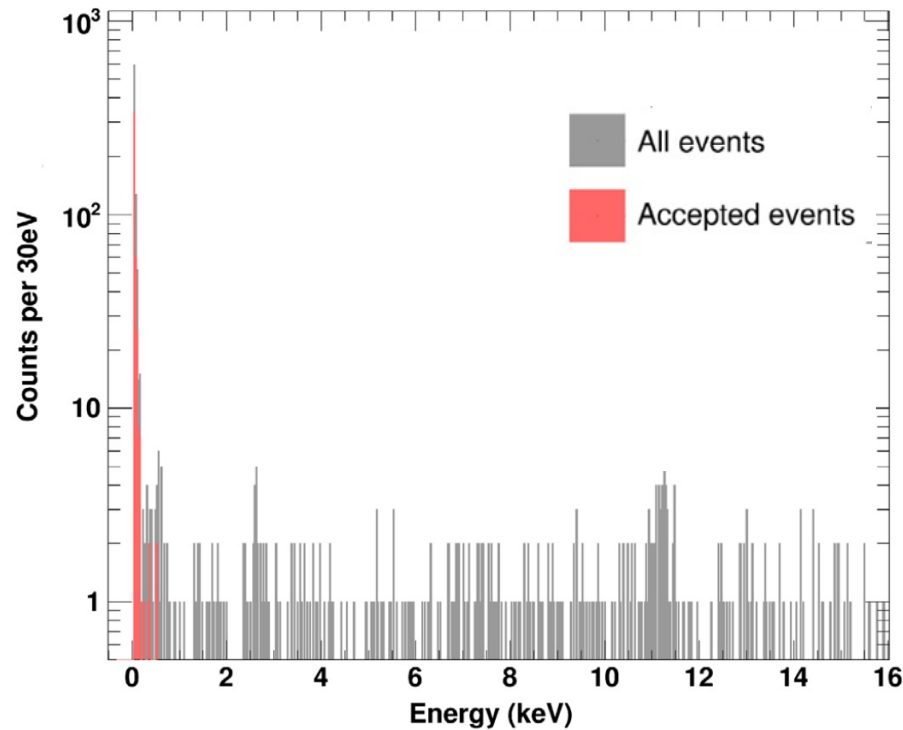
i Poster: "Geant4 simulations of the influence of contamination and roughness of the detector surface on background spectra in CRESST" by Christoph Gruner

First results from CRESST-III (2019): leading limit for sub-GeV DM with a 30 eV energy threshold detector.

Detector A: 23.6 g CaWO₄

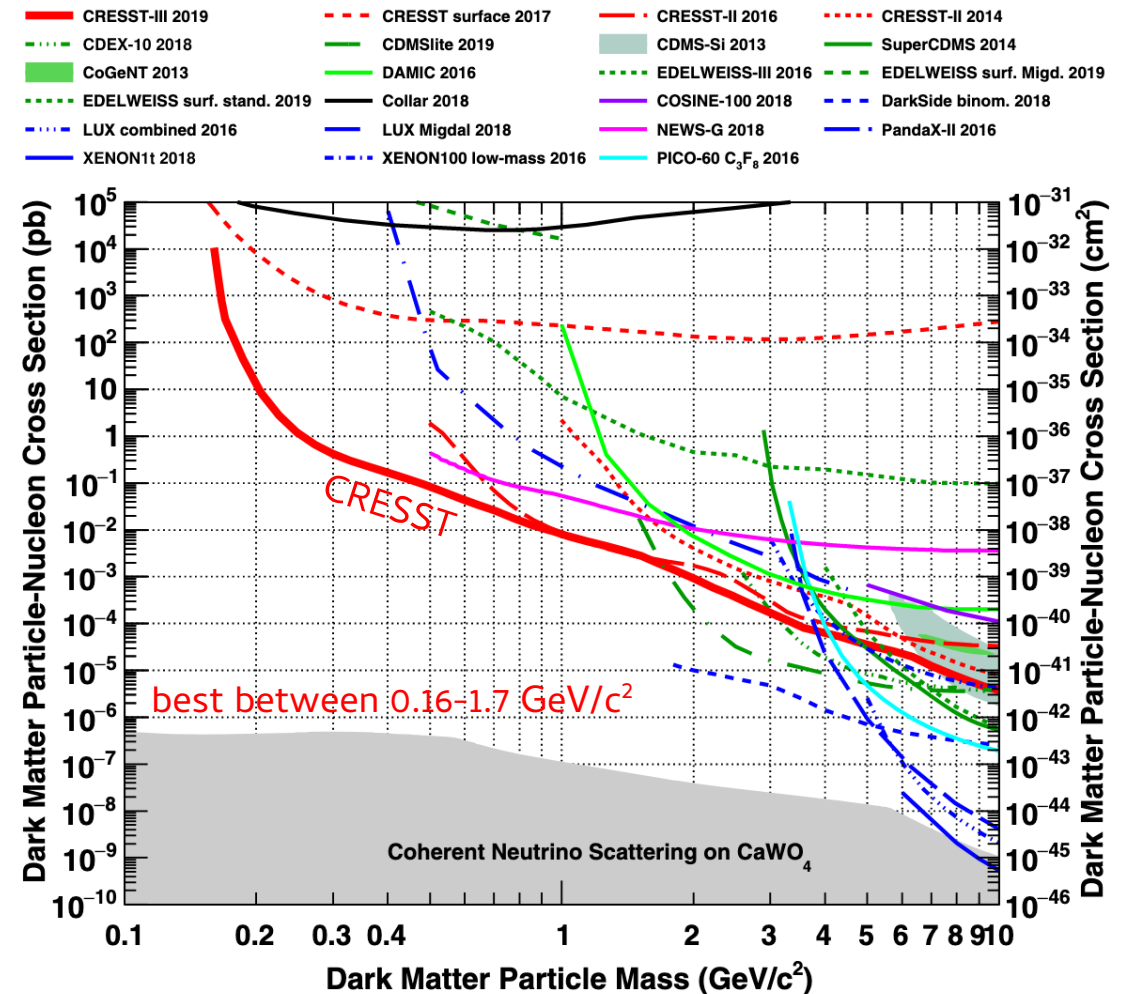
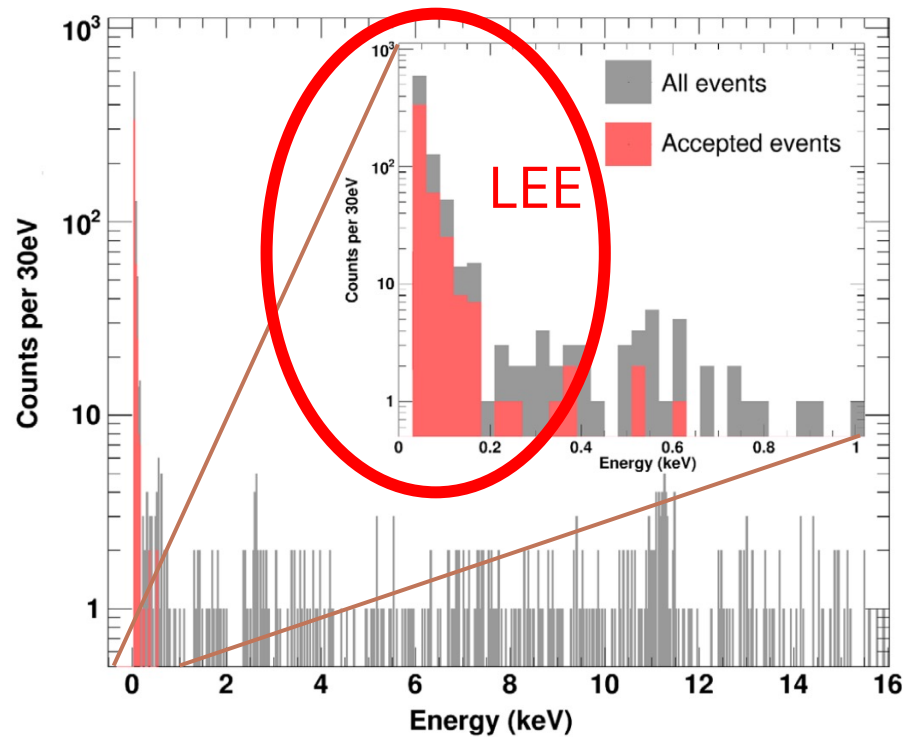
Exposure: 5.698 kg · days

Data-taking period: Oct 2016 – Jan 2018 Energy threshold: 30.1 eV



Sharply rising energy spectrum below 200 eV - **low energy excess (LEE)** - is observed in CRESST-III detectors and limits the sensitivity to sub-GeV DM.

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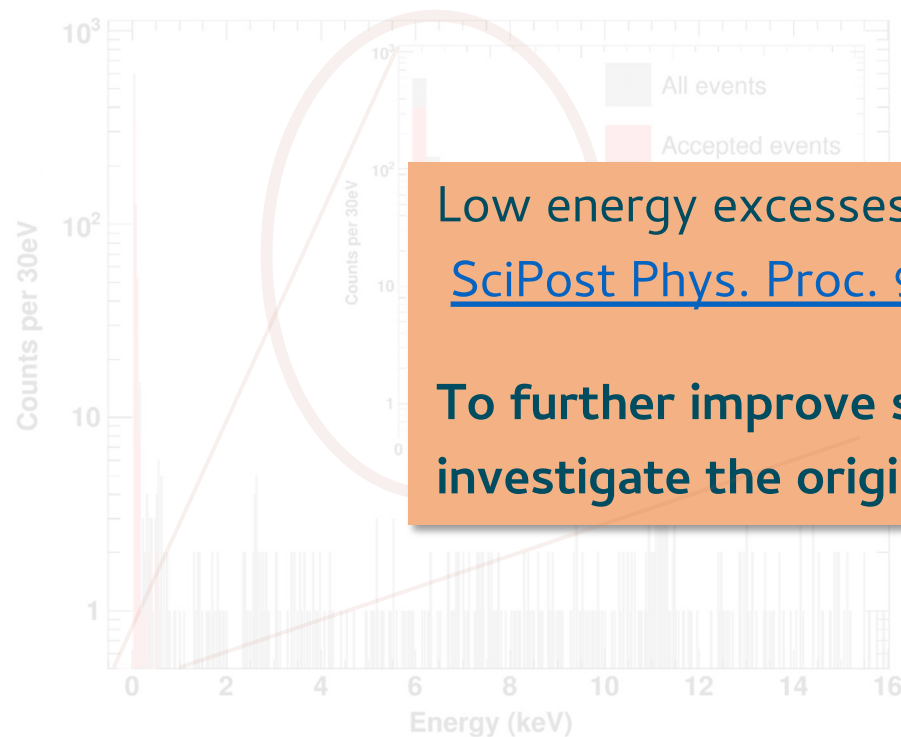


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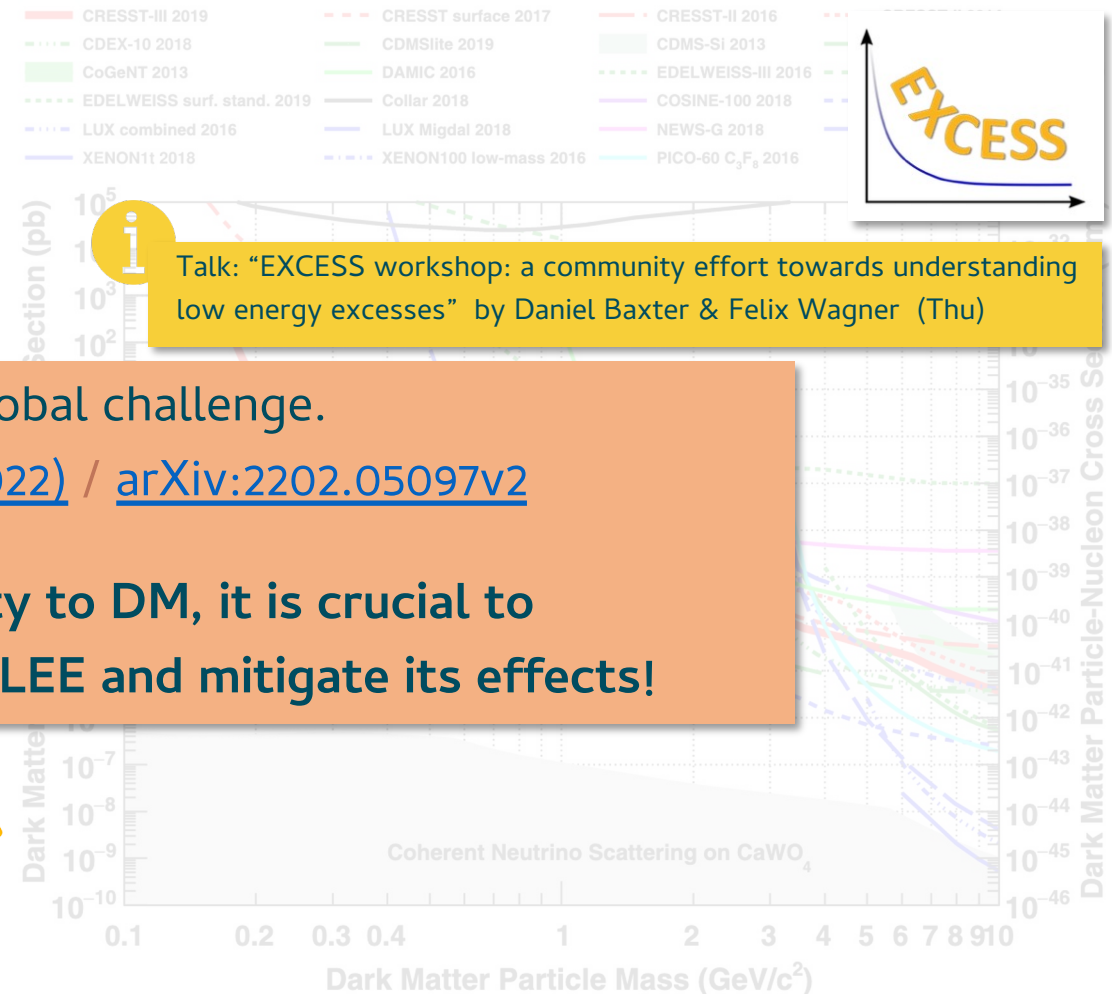
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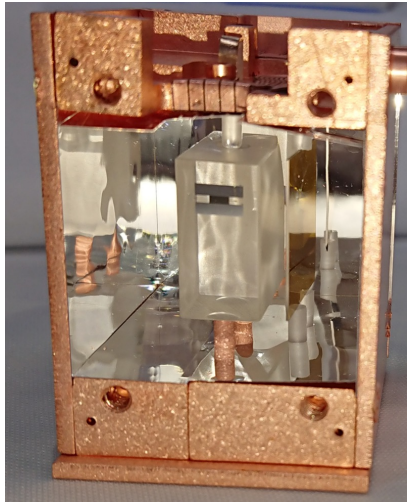
Low energy excesses are a global challenge.

[SciPost Phys. Proc. 9, 001 \(2022\)](#) / [arXiv:2202.05097v2](#)

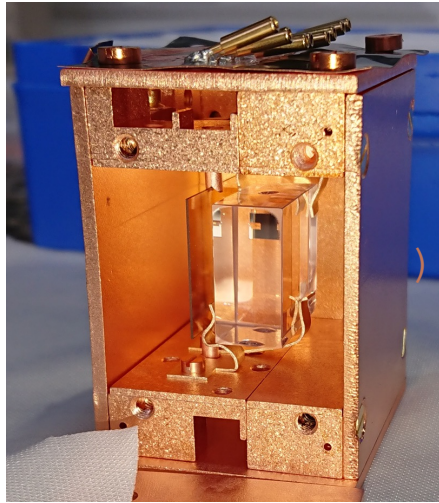
To further improve sensitivity to DM, it is crucial to investigate the origin of the LEE and mitigate its effects!



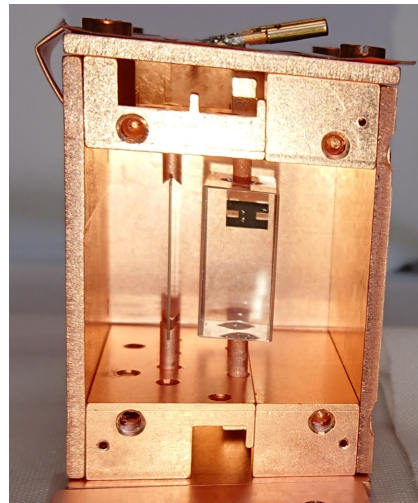
Multiple design modifications were applied in the current data-taking campaign to test ideas about the LEE origin.



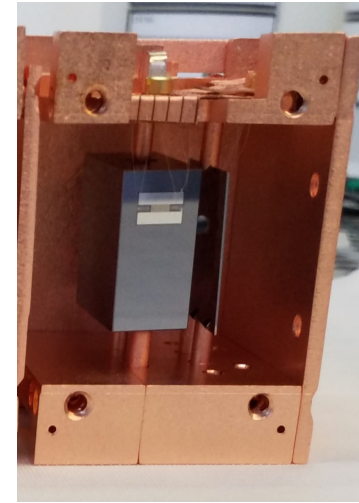
CaWO₄ grown at TUM



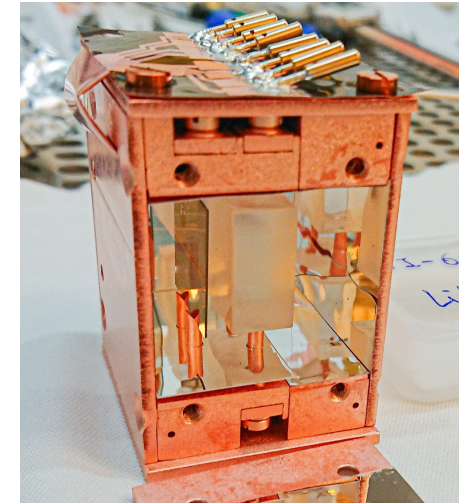
Commercially grown CaWO₄



Al₂O₃



Si



LiAlO₂

Current measurement campaign started in November 2020 and is now at the last stage.

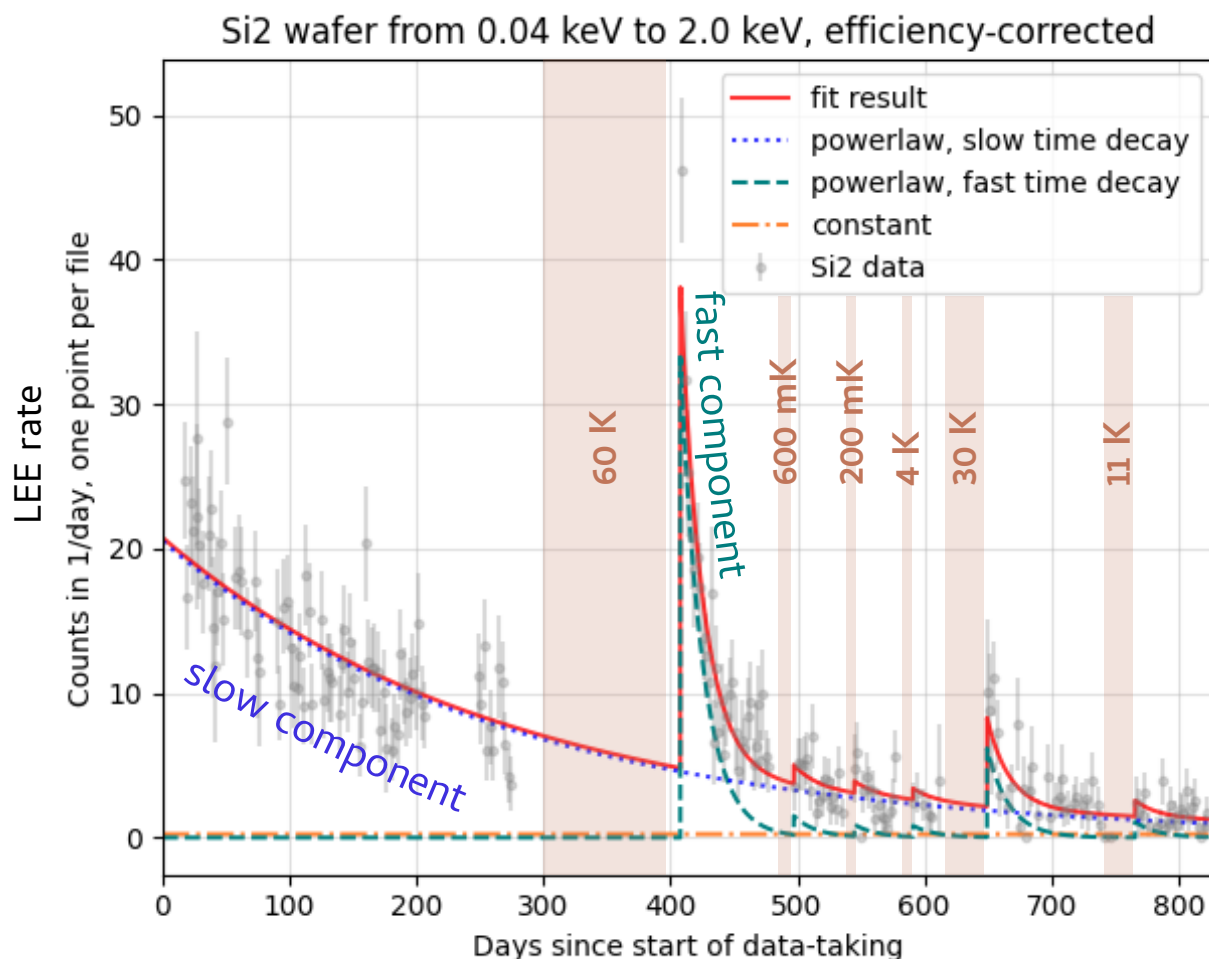
- Various target materials: CaWO₄, Al₂O₃, LiAlO₂, Si
- Different holding structures (sticks, clamps)
- Remove scintillating parts (foil, sticks, scintillating crystals)

All thresholds are at O(10 eV)

LEE is observed in all detectors

LEE rate is decaying with time.

Excess rate increases after warming up the cryostat to O(10K).



1. Continuous decaying trend since the start of the data taking O(100 days) → **slow component**

2. An additional component of about 15 days appears after warming up to O(10K) → **fast component**

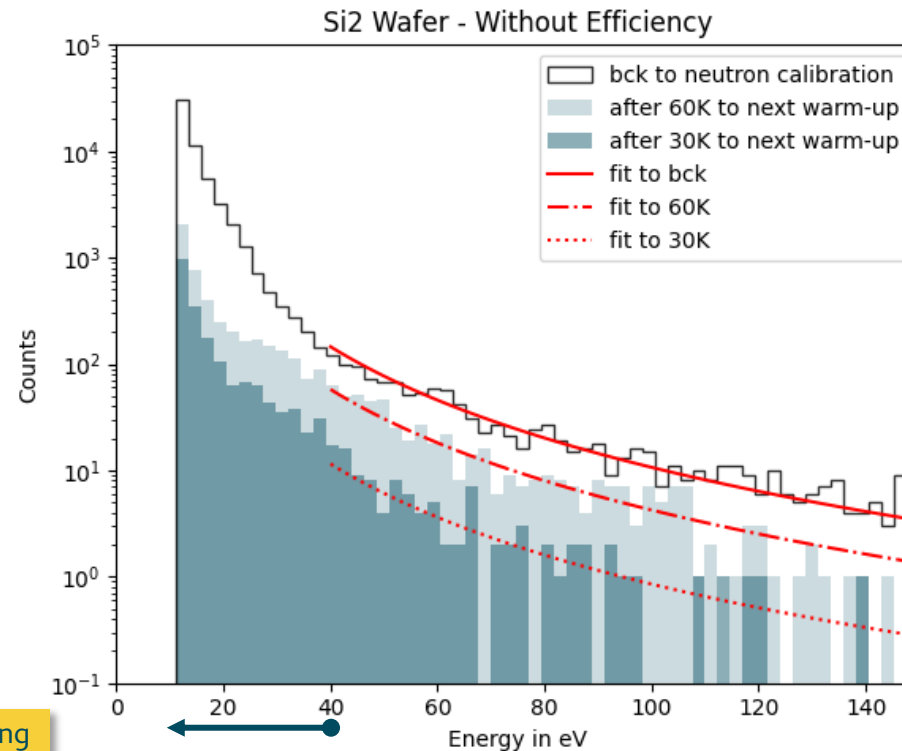
Rate increase after warm-ups excludes external and internal radioactivity as well as DM as a major origin of the LEE.

For all detectors, **spectral shape above 40 eV:**

- is well described by a **single power law**

- has two time components.

$$N(E, t) = C + E^{-\epsilon} (N_{slow} \cdot e^{-t/\tau_{slow}} + N_{fast} \cdot e^{-t/\tau_{fast}})$$



Extension of the model is needed
at lower energies <40 eV.

i

Poster: "Description of the low energy excess in CRESST using two-dimensional unbinned likelihood fits" by Sarah Kuckuk

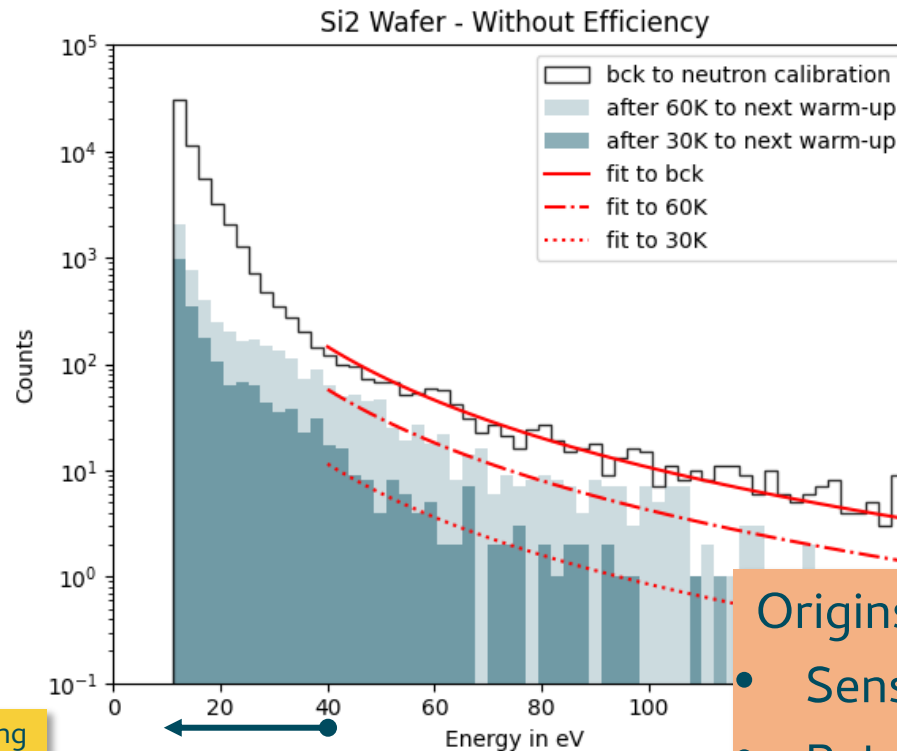
Talk at the Excess@TAUP workshop last Saturday: "The low energy excess in CRESST-III" by Dominik Fuchs

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Origins under investigation:

- Sensor related events
- Relaxation of holding-induced stress
- Intrinsic crystal effects

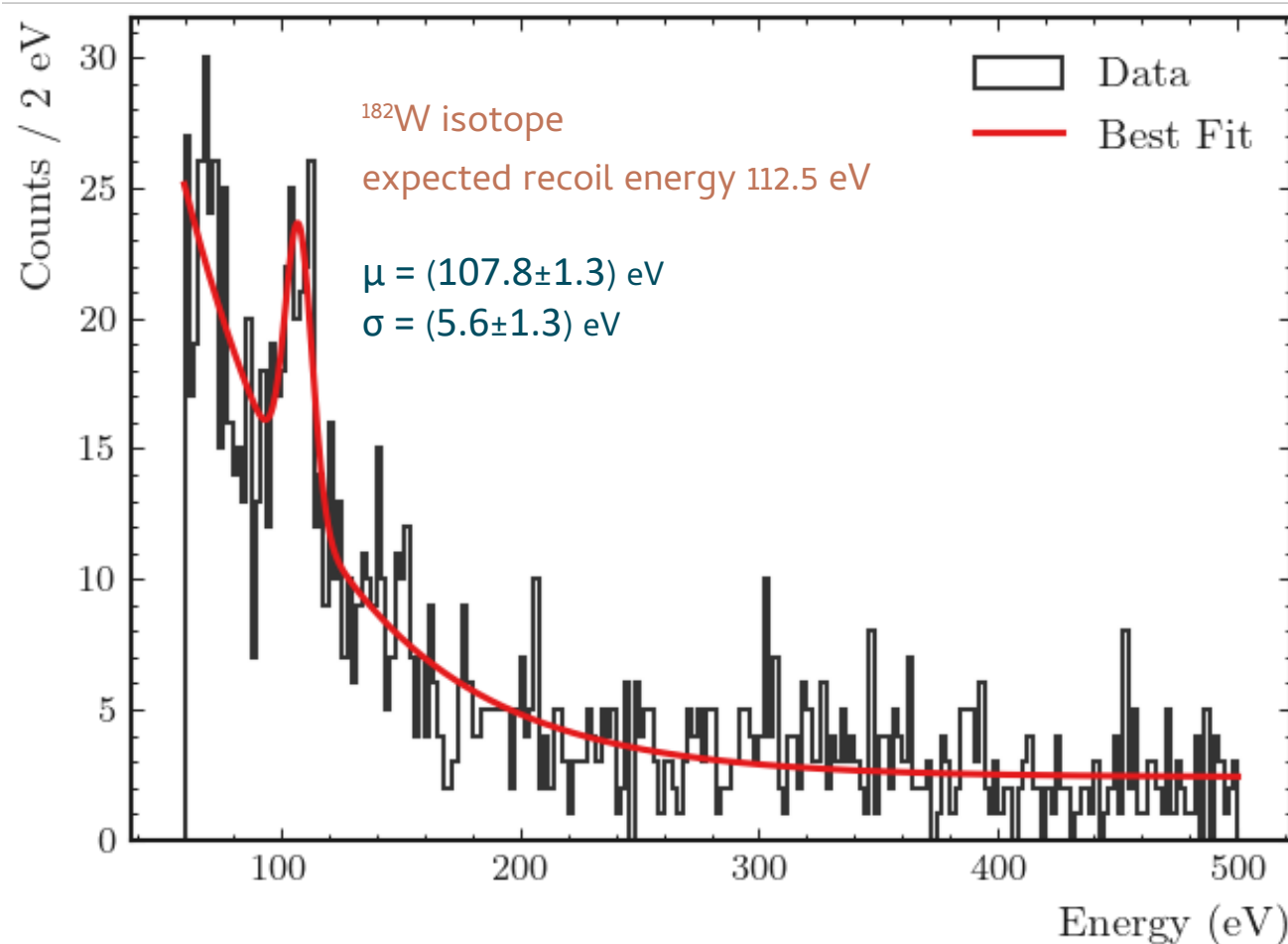
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Talk at the Excess@TAUP workshop last Saturday: "The low energy excess in CRESST-III" by Dominik Fuchs

Observation of a mono-energetic nuclear recoil peak at **112 eV** in three CRESST-III CaWO_4 detectors -> **new calibration technique**

Neutron calibration data (AmBe source)



Thermal neutron capture

→ de-excitation with a single γ

→ mono-energetic nuclear recoil

Proposed as a low-energy calibration method for cryogenic detectors (CRAB collaboration).

[L. Thulliez et al 2021 JINST 16 P07032](#) / [arXiv:2011.13803](#)

Observation of the nuclear recoil peaks around the predicted value in multiple detectors

1. confirms CRESST energy calibration approach,
2. provides a potential for **precise energy calibration for nuclear recoils at 100 eV scale.**

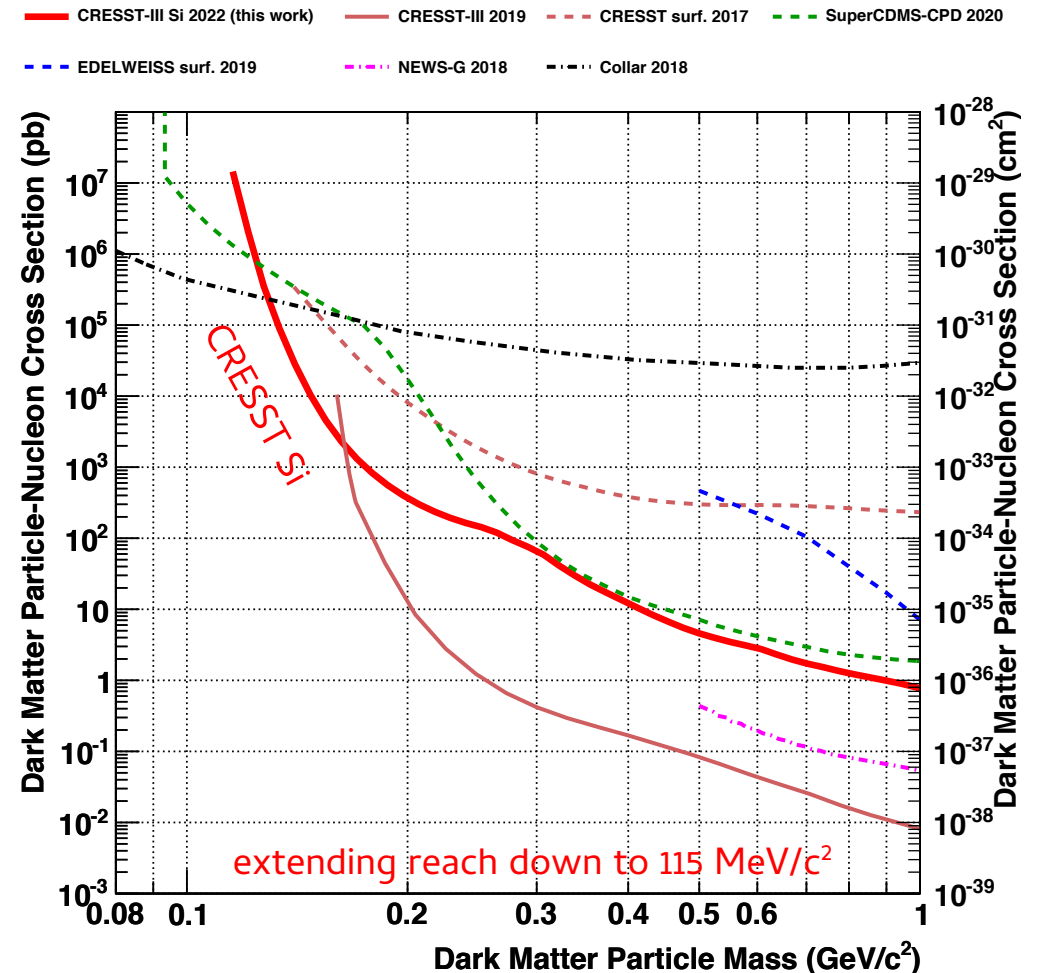
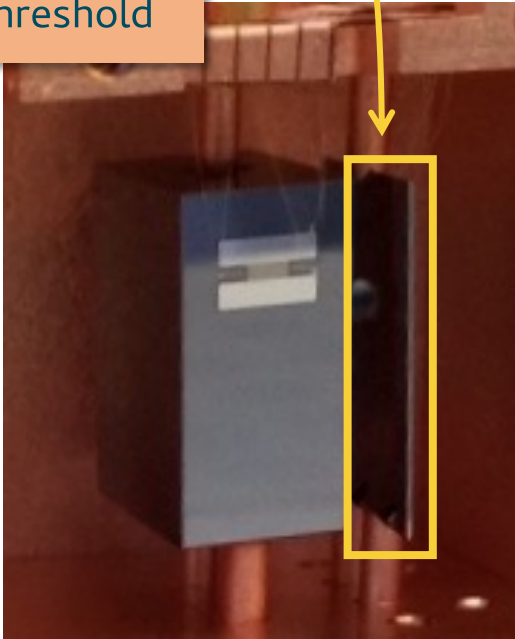
DM results with different target materials

Si wafer detector: energy threshold of 10 eV extends sensitivity to DM particles with the mass down to 115 MeV/c².

Si wafer detector: 0.35 g Exposure: 55.06 g · days
 Data-taking period: Nov 2020 – Aug 2021 Energy threshold: **10.0 eV**

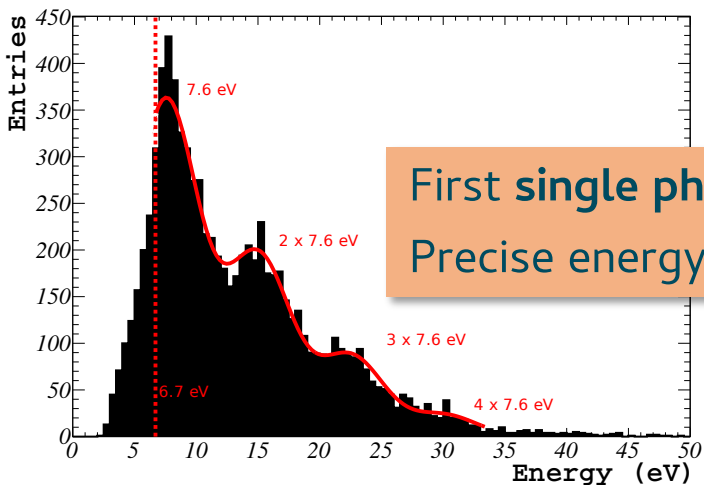
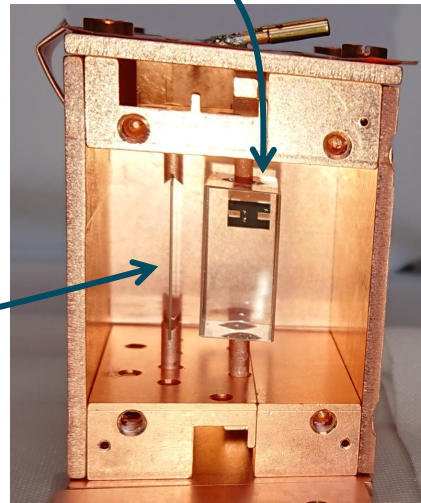
Thin wafer detector is a target
 Bulky detector is a veto to remove coincidence events.

10.0 ± 0.2 eV energy threshold



Silicon-on-sapphire wafer detector achieves an **energy threshold** < 10 eV and reaches sensitivity to DM mass < 100 MeV/c².

Al₂O₃ bulky crystal is exposed to a ⁵⁵Fe calibration source
→ luminescence emission at 7.6 eV
→ single photon detection by a facing wafer detector @ $n \times 7.6$ eV



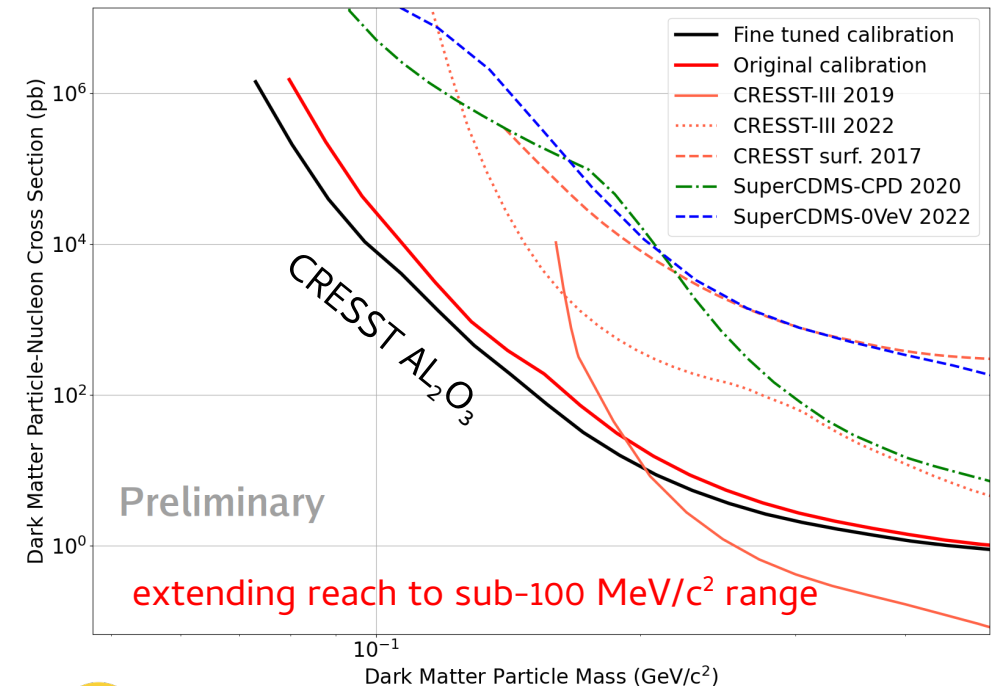
First single photons observation in CRESST!
Precise energy calibration at eV-scale!

Al₂O₃ wafer detector: 0.6 g

Data-taking period: Nov 2020 – Aug 2021

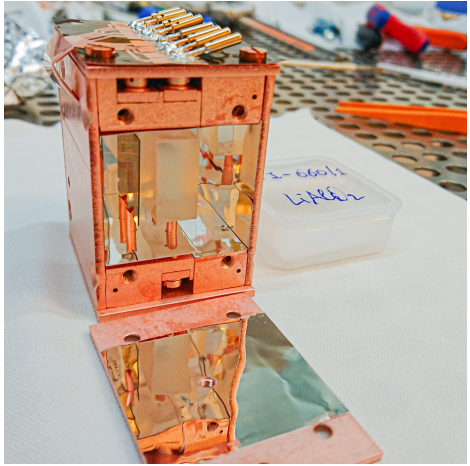
Exposure: 0.14 kg · days

Energy threshold: **6.7 eV**



Talk: “First observation of single photons in a CRESST detector and new Dark Matter exclusion limits” by Dominik Fuchs (Mon)

Well performing LiAlO_2 detector allowed to greatly improve the limits for **spin-dependent** DM-nucleon interactions.



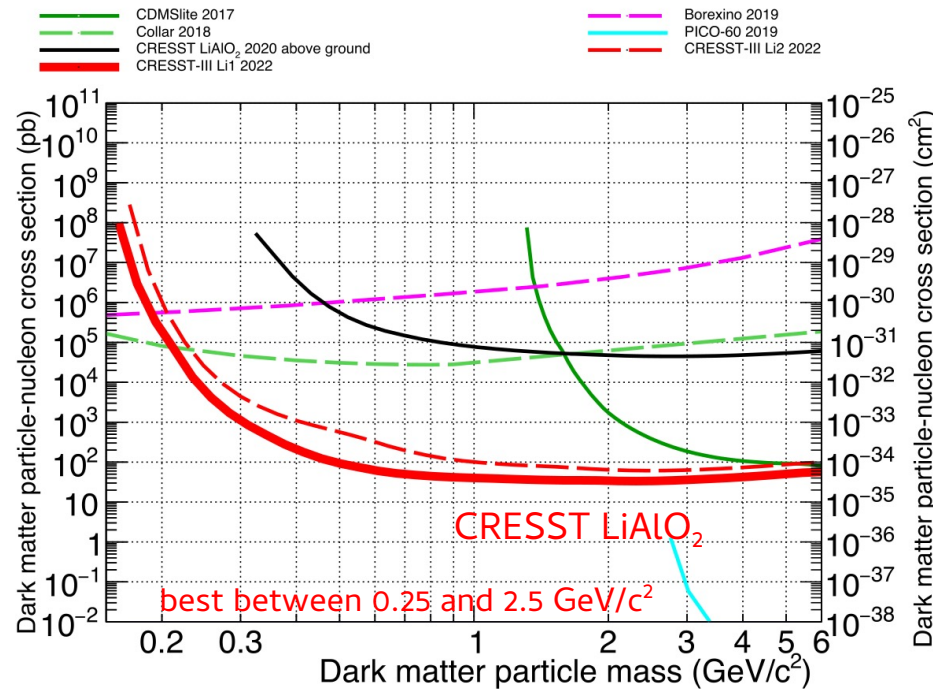
Li1 detector: 10.5 g LiAlO_2

Data-taking period: Feb 2021 – Aug 2021

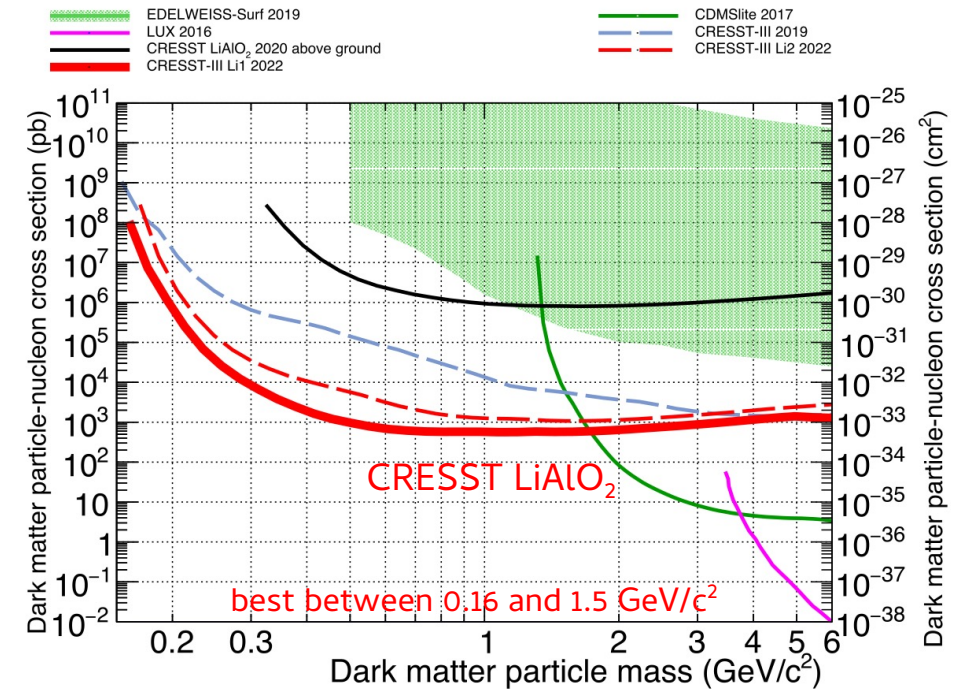
Exposure: 1.161 kg · days

Energy threshold: 83.6 eV

Proton



Neutron



CRESST technique allows to use various target materials to test different interaction channels!

i

Poster: “First Results on ^{17}O Enrichment of CaWO_4 Crystals for Spin-dependent DM search with CRESST” by Angelina Kinast

Future: New CRESST detectors

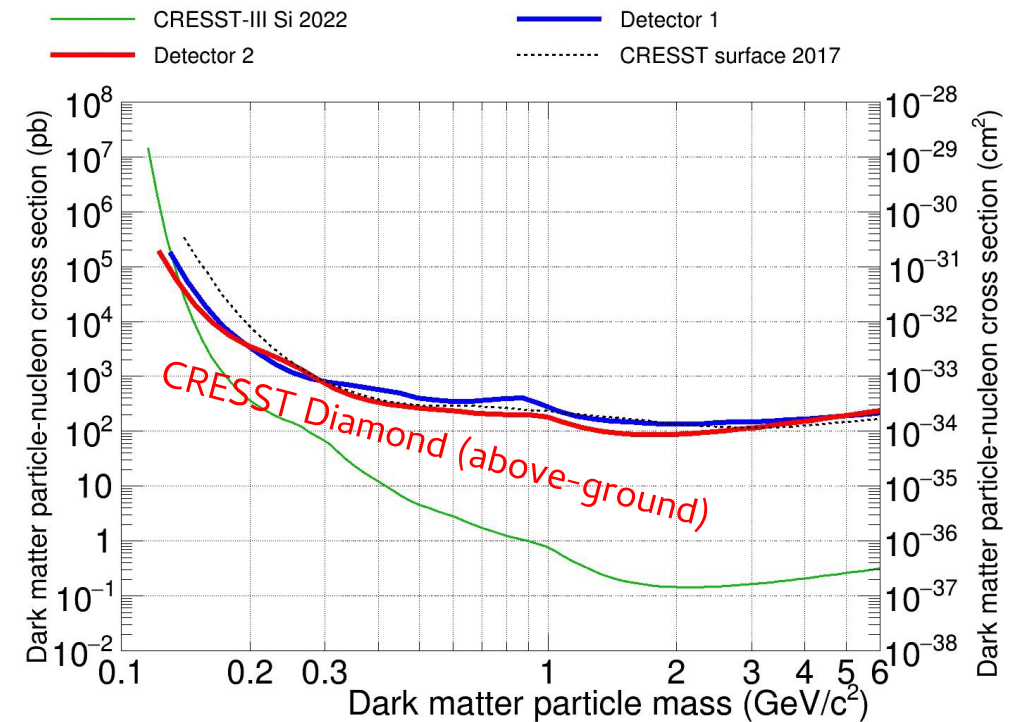
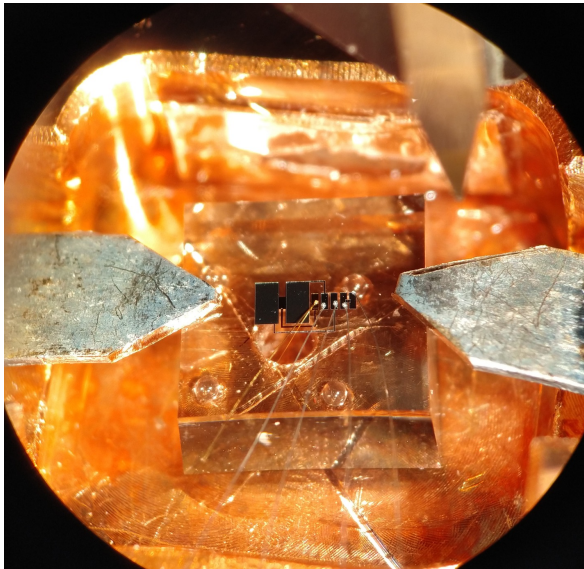
New material: diamond crystal operated as a cryogenic detector showed a high potential for sub-GeV DM search.

Diamond detector: 0.175 g

Exposure: 0.27 g · days

37 h of stable data above-ground

Nuclear recoil threshold: 16.8 eV

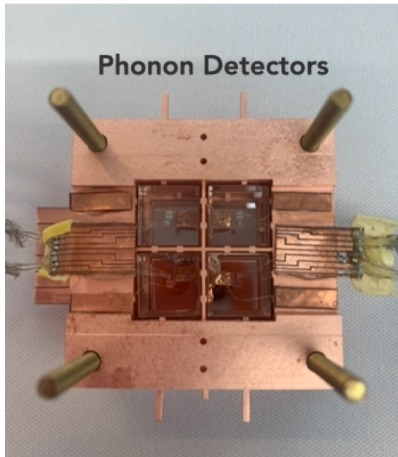


i Poster: "A low-threshold diamond cryogenic detector for sub-GeV Dark Matter searches" by Anna Bertolini

Detectors with new module designs to study the LEE origin will be installed at LNGS in the next run starting at the end of 2023.

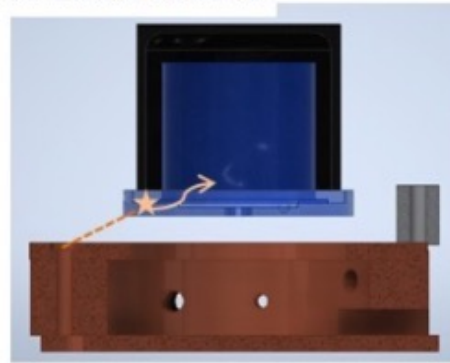
Holding structures

novel stress-free holding structure
-> LEE reduction?



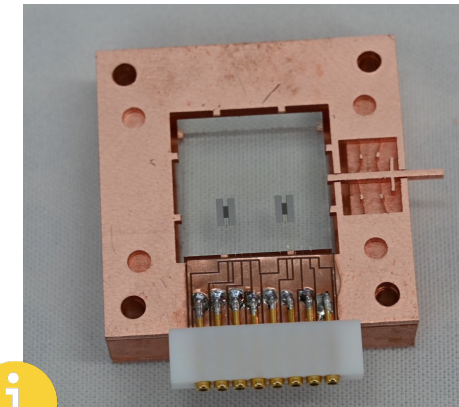
instrumented holders
-> discriminate events transmitted through the holding structure

• transmitted events



Double TES read-out

-> discriminate events in the absorber from events in the TES films or material interfaces



i

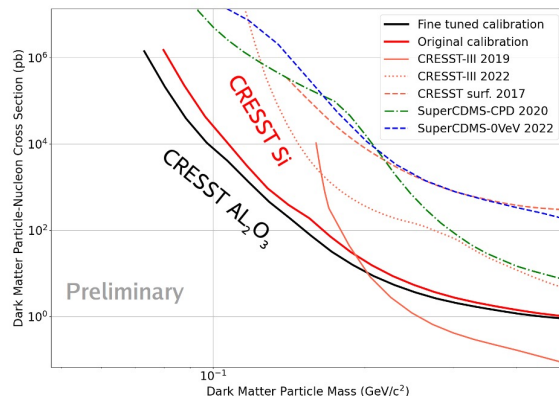
Talk: "Results of doubleTES detectors" by Francesca Pucci at the Excess@TAUP workshop last Saturday

More features

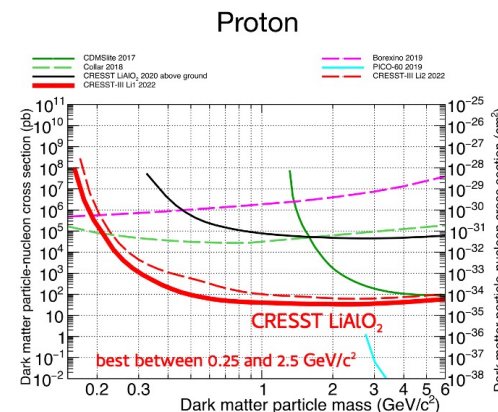
- maximize sensitivity to dark matter recoils
- 4π veto

Above-ground measurements show very promising performance!

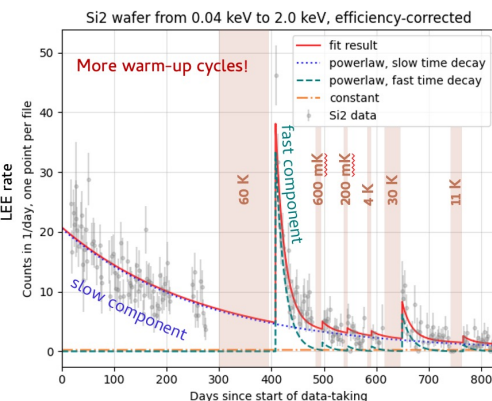
Conclusions & Outlook



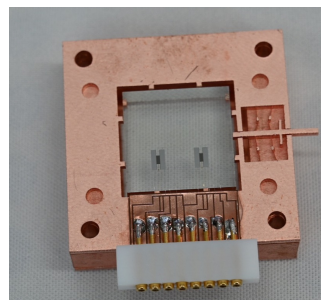
Reach to DM masses $<100 \text{ MeV}/c^2$
with low threshold detectors



Improved limits for spin-dependent
interactions with LiAlO_2 target.

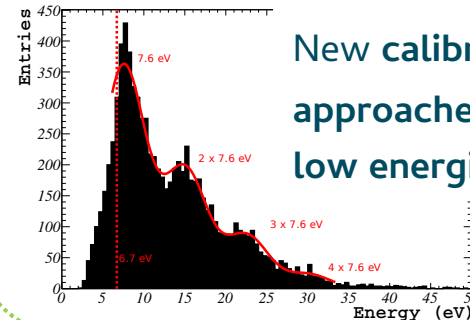
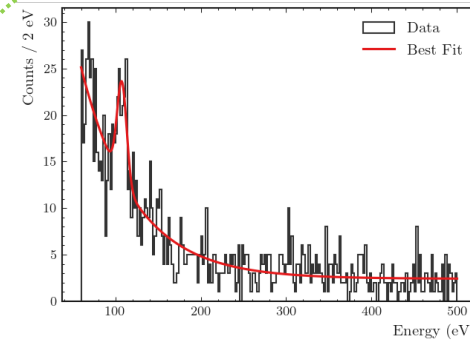


Strong hints towards solid state
physics effects playing a major
role for low energy excess.



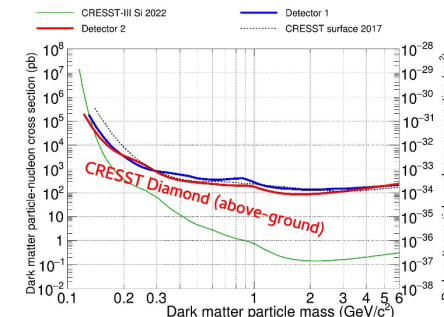
Coming to
LNGS late 2023

Novel detector module designs to
study sensor-related events and the
impact of the stress from holders



New calibration
approaches at
low energies

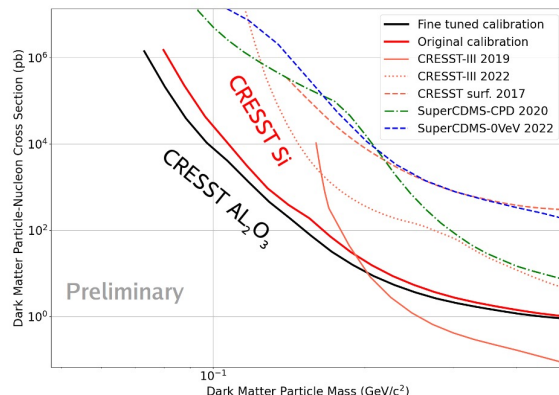
New target material:
**diamond detectors for
sub-GeV DM searches.**



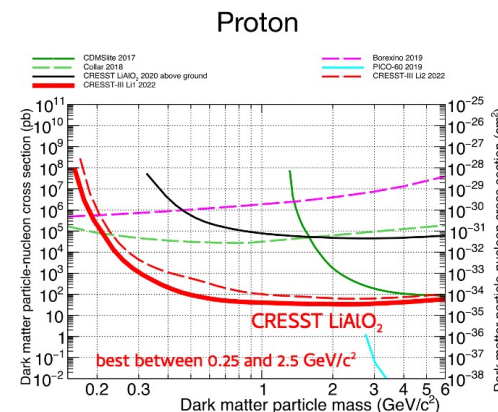
A major CRESST setup upgrade to 288 read-
out channels is in preparation:

- SQUIDs and wiring are already produced
- new DAQ and bias electronics are designed
- goal: installation at LNGS in 2024

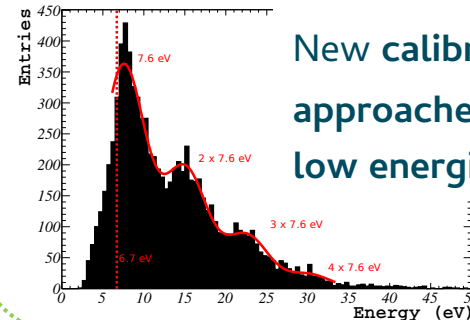
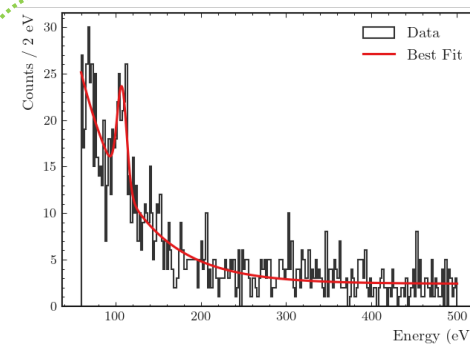
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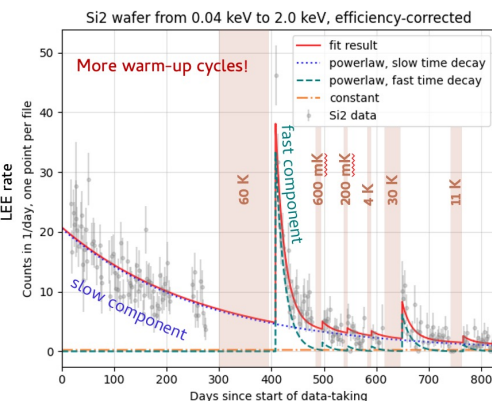
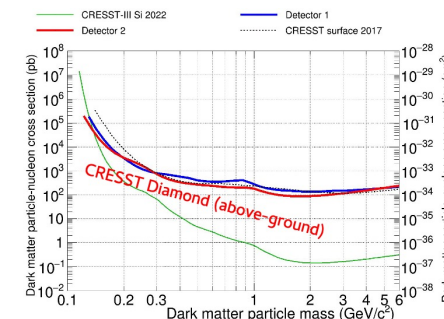


Improved limits for spin-dependent
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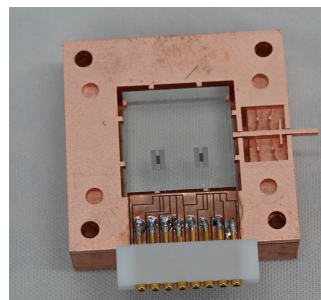
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LNGS late 2023



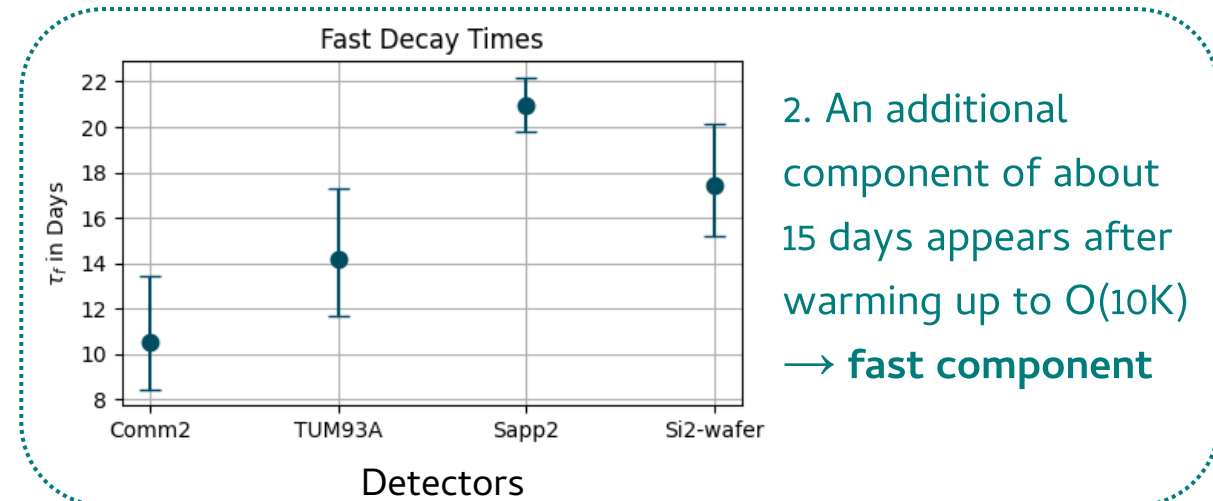
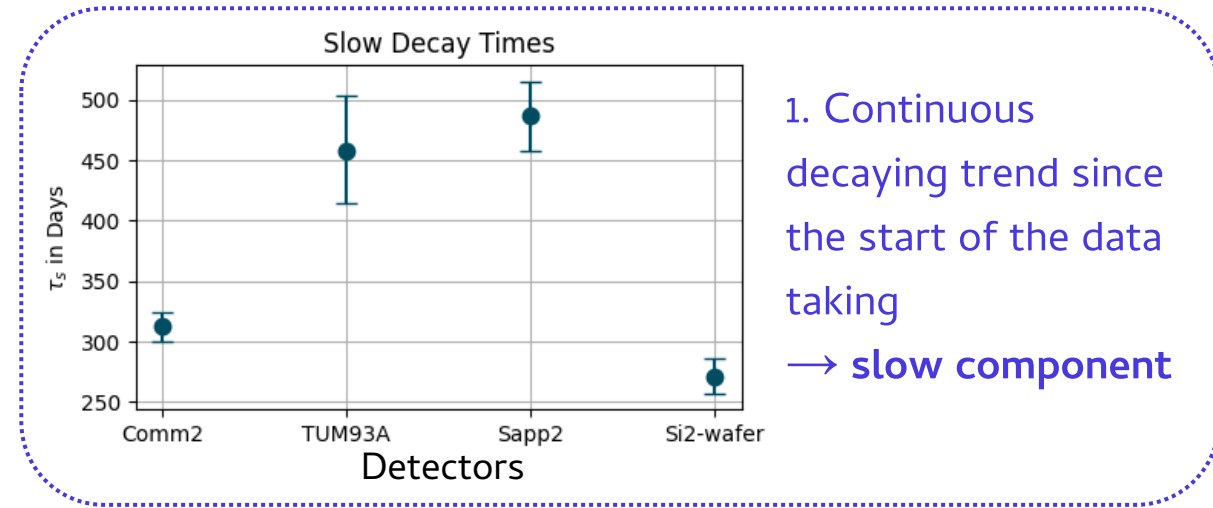
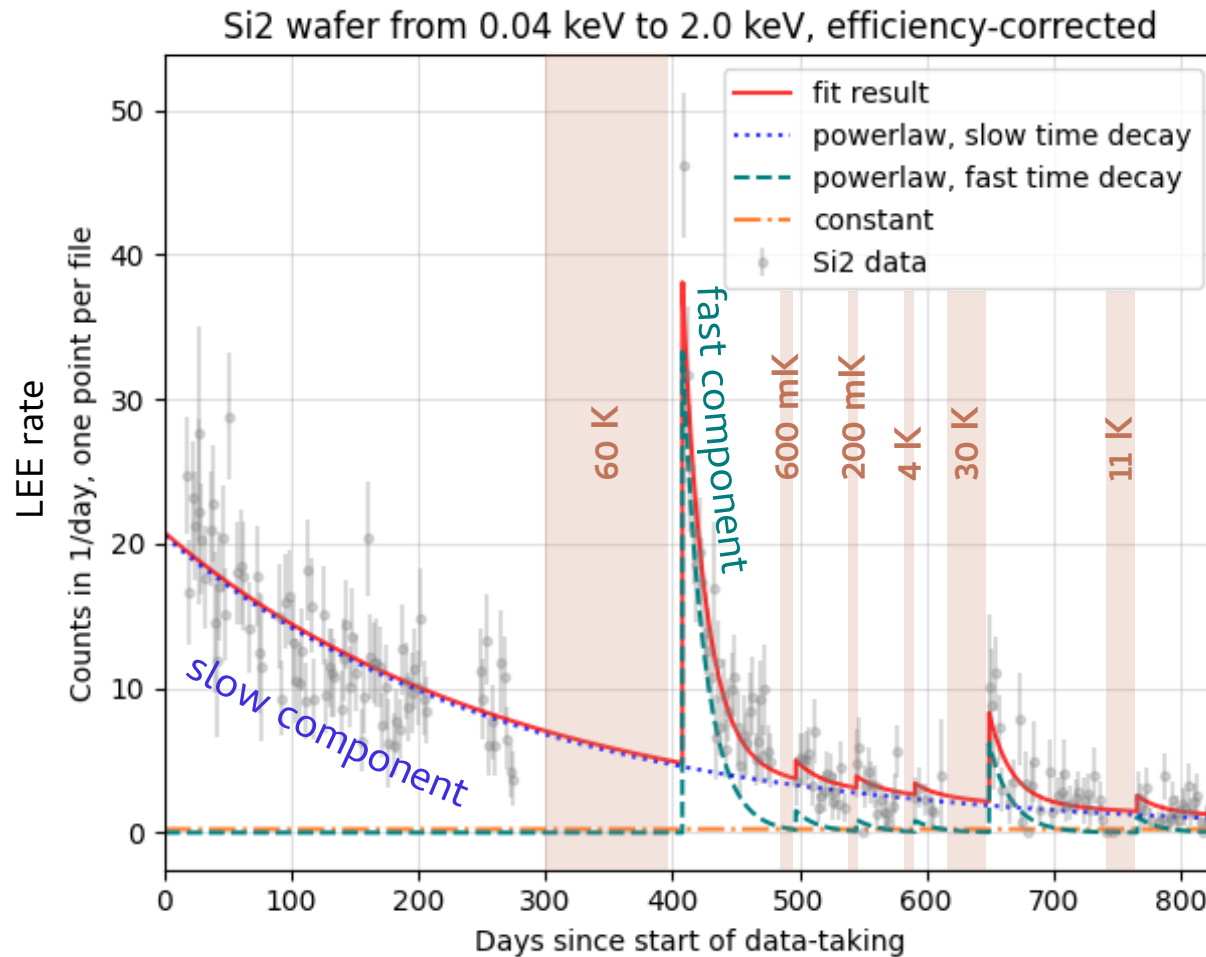
Novel detector module designs to
study sensor-related events and the
impact of the stress from holders

**A major CRESST setup upgrade to 288 read-
out channels is in preparation:**

- SQUIDs and wiring are already produced
- new DAQ and bias electronics are designed
- goal: installation at LNGS in 2024

Back up

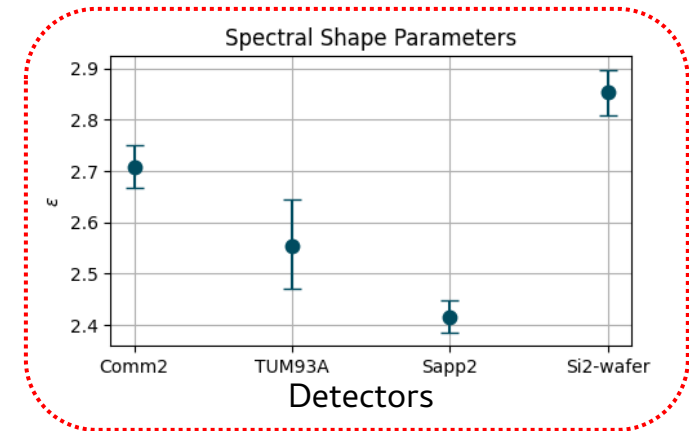
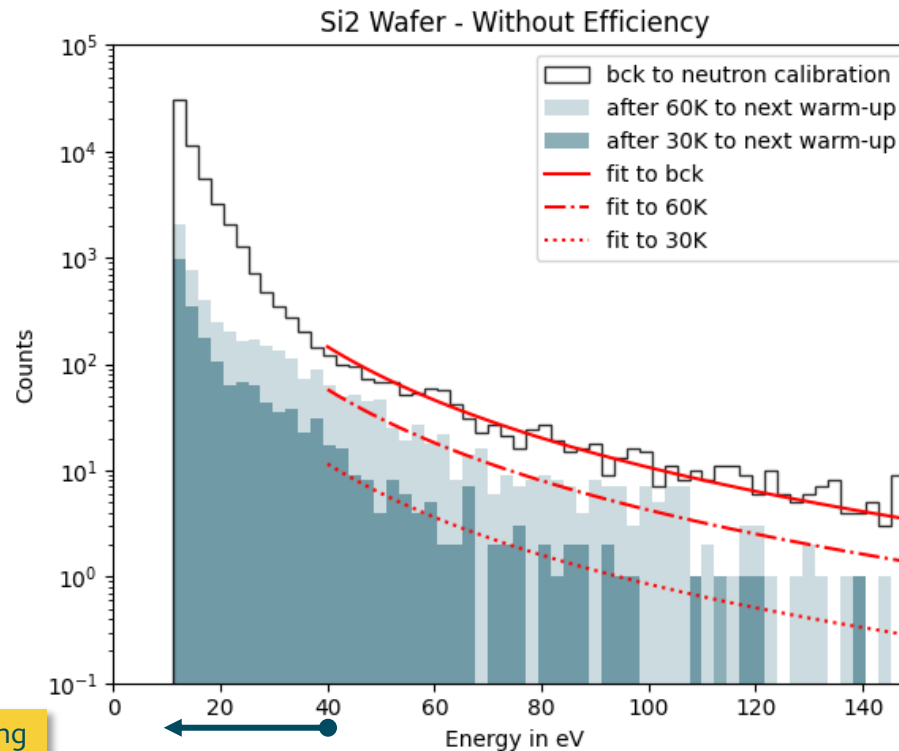
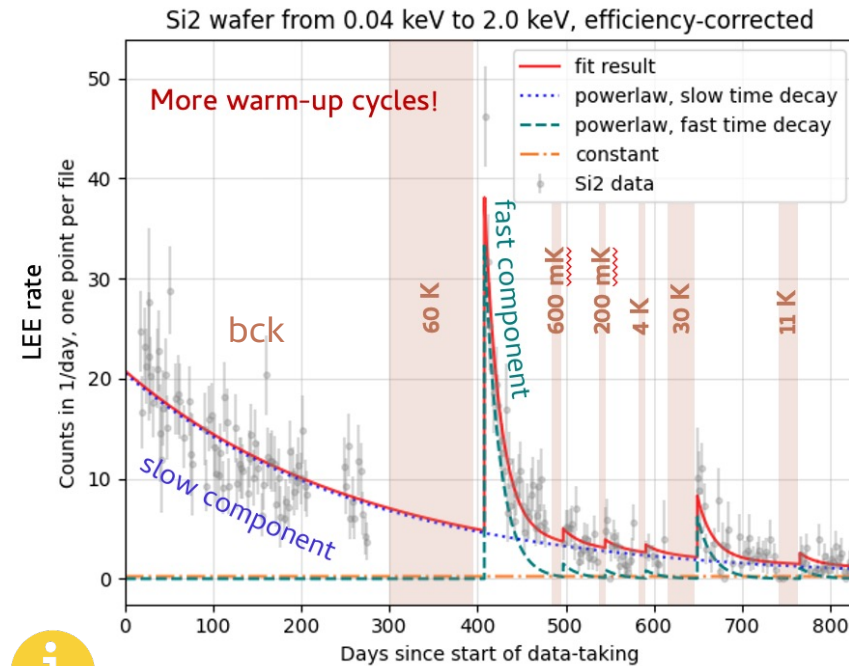
Excess rate increases after warming up the cryostat to O(10K).



For all detectors, **spectral shape above 40 eV:**

- is well described by one **power law**
- has two time components.

$$N(E, t) = C + E^{-\epsilon} (N_{slow} \cdot e^{-t/\tau_{slow}} + N_{fast} \cdot e^{-t/\tau_{fast}})$$



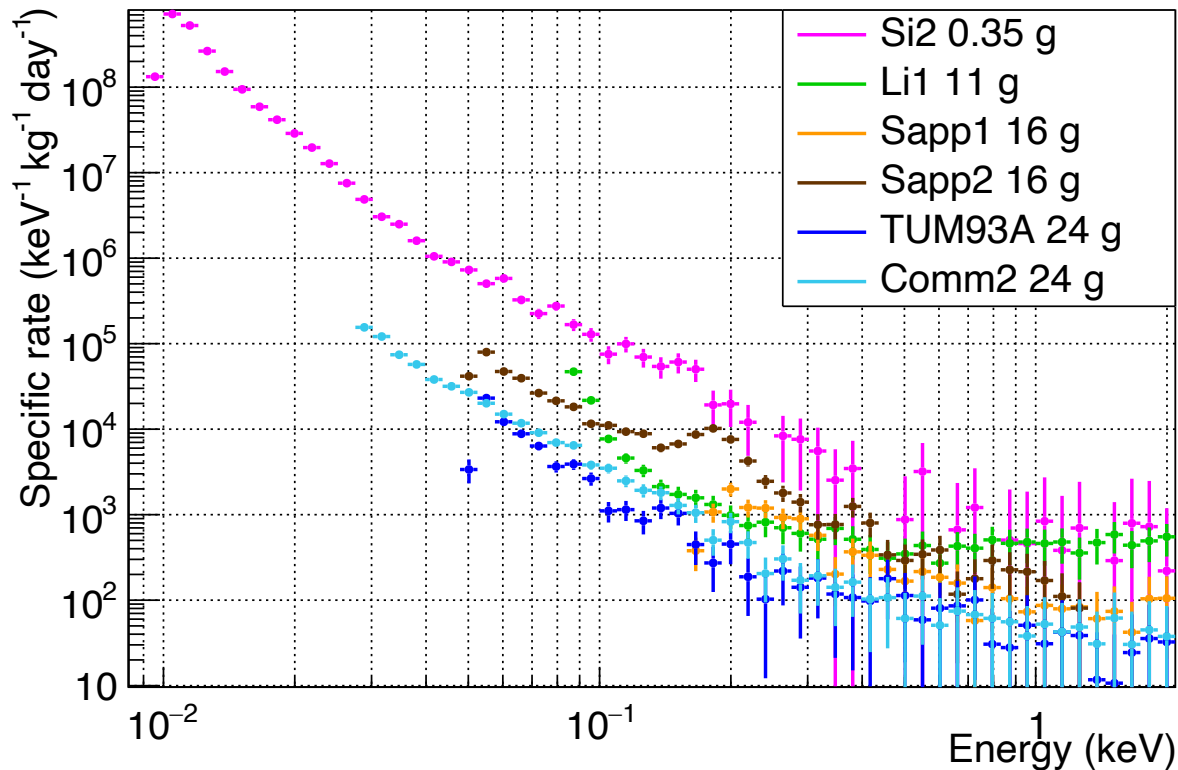
Extension of the model is needed
at lower energies <40 eV.

Poster: "Description of the low energy excess in CRESST using two-dimensional unbinned likelihood fits" by Sarah Kuckuk

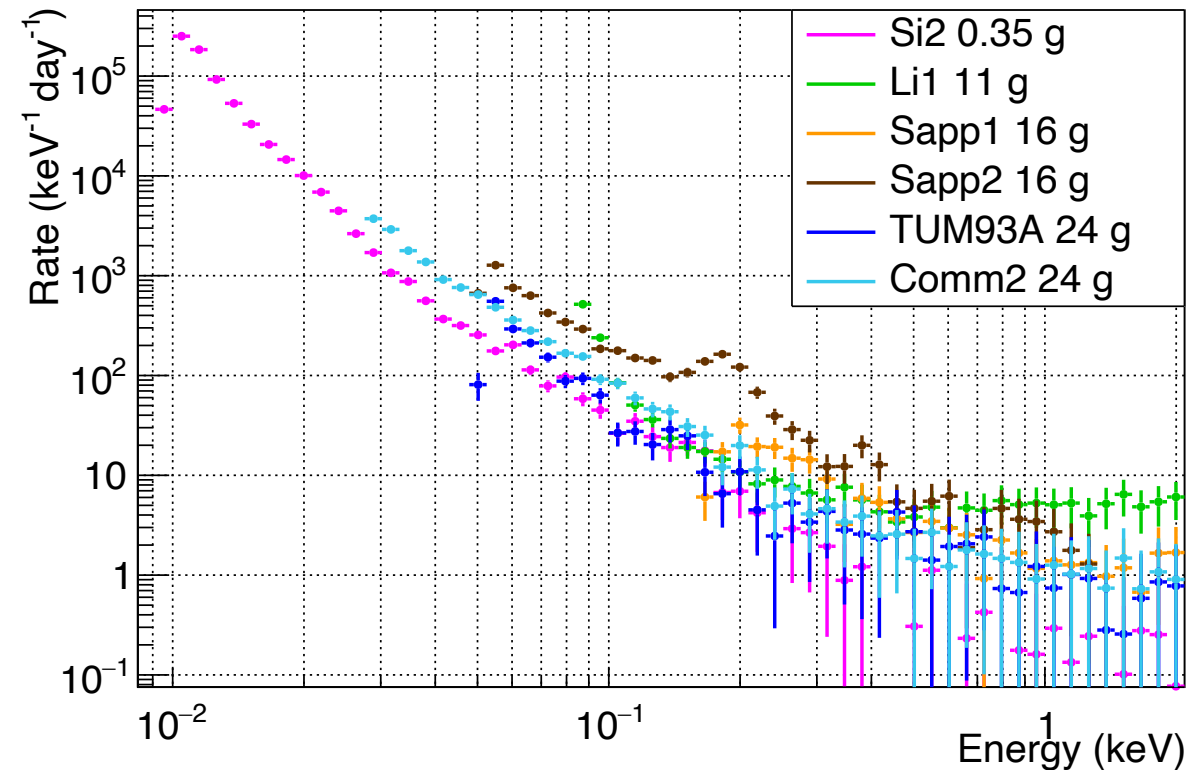
Talk at the Excess@TAUP workshop last Saturday: "The low energy excess in CRESST-III" by Dominik Fuchs

LEE is observed in all detector materials and its rate does not scale with the absorber mass.

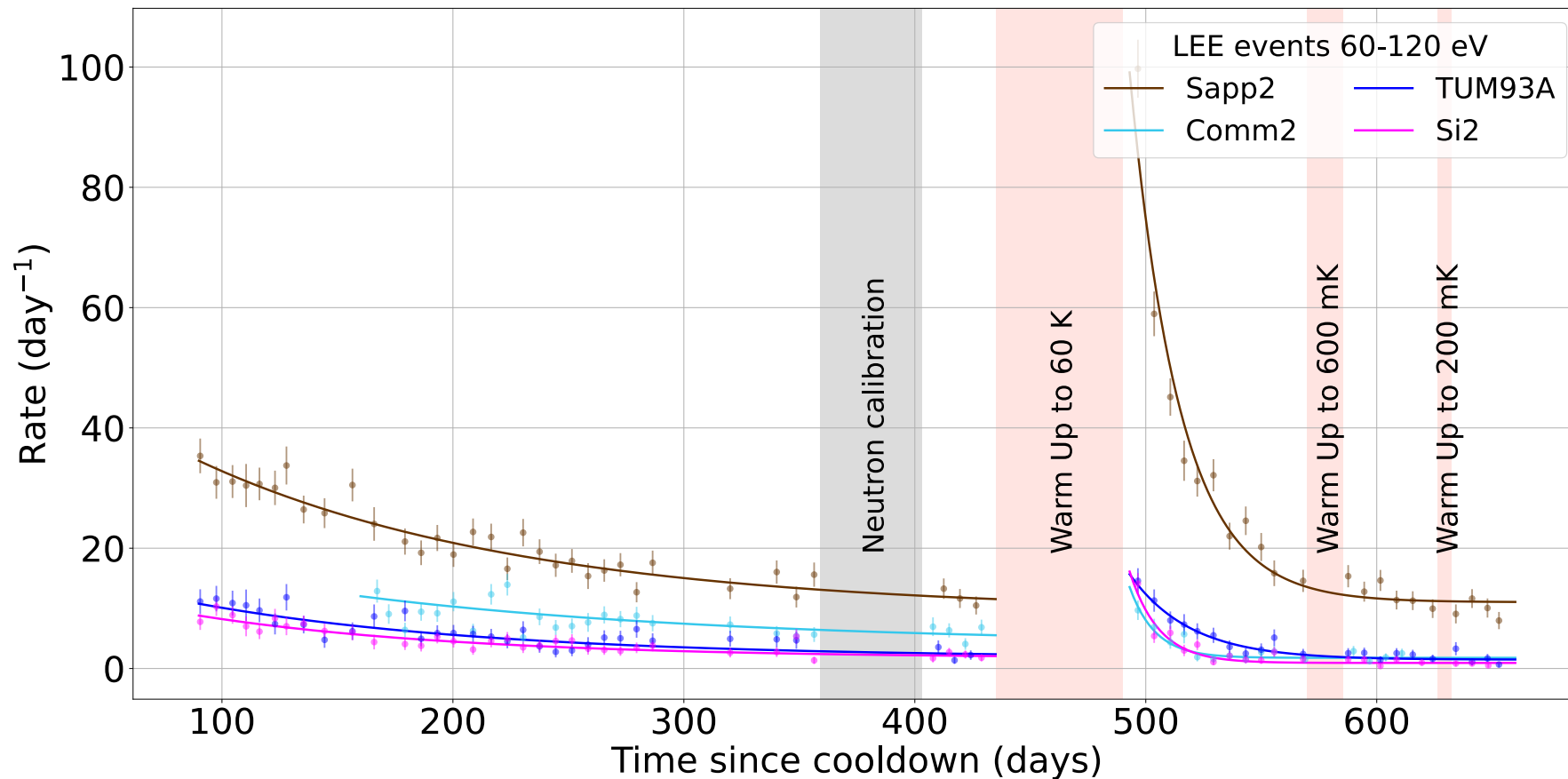
Scaled with the absorber crystal mass



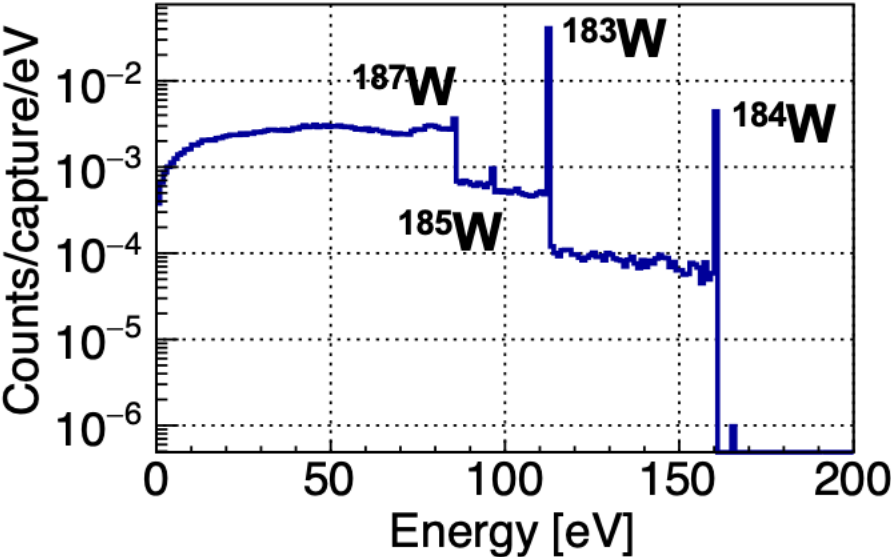
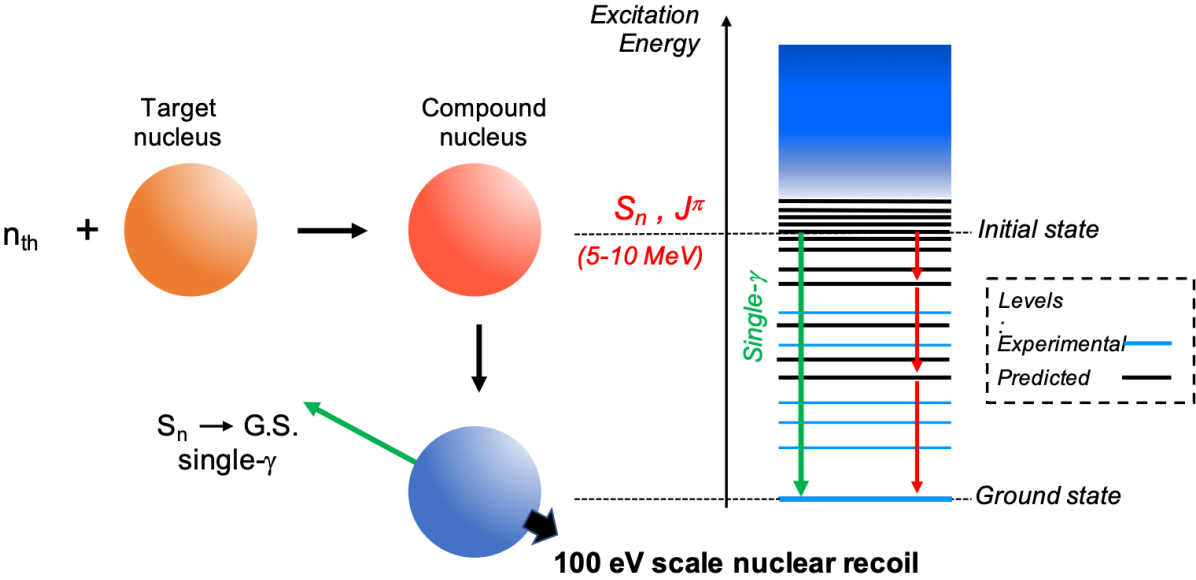
NOT scaled with the absorber crystal mass



Time evolution of LEE rates (60-120 eV)



CRAB (Calibrated nuclear Recoils for Accurate Bolometry)



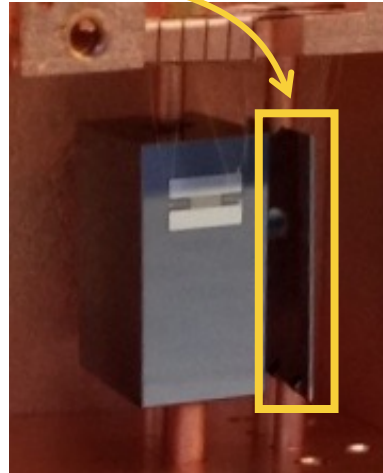
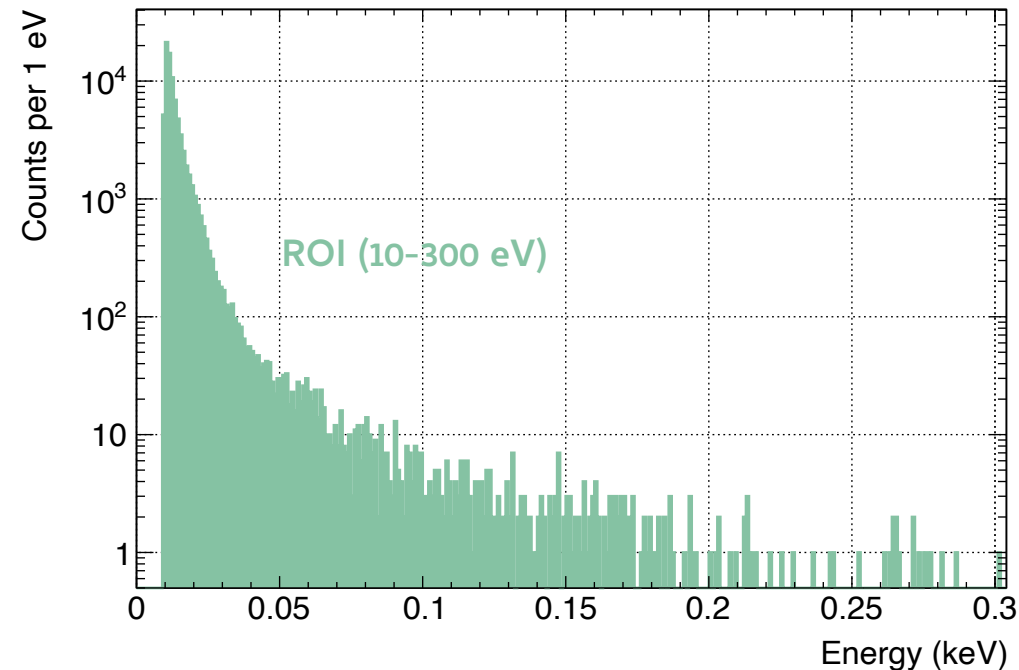
Target nucleus (A)			Compound nucleus (A+1)			
Isotope	Y_{ab} [24] (%)	$\sigma_{n,\gamma}$ [25] (barn)	S_n [26] (keV)	I_γ^s [26, 27] (%)	Recoil (eV)	FoM (a.u.)
^{182}W	26.50	20.32	6191	13.94	112.5	7506
^{183}W	14.31	9.87	7411	5.83	160.3	823
^{184}W	30.64	1.63	5754	1.48	96.1	74
^{186}W	28.43	37.89	5467	0.26	85.8	280
^{70}Ge	20.53	3.05	7416	1.95	416.2	122
^{72}Ge	27.45	0.89	6783	0.0	338.7	0
^{73}Ge	7.76	14.70	10196	0.0	754.9	0
^{74}Ge	36.52	0.52	6506	2.83	303.2	54
^{76}Ge	7.74	0.15	6073	0.0	257.3	0

L. Thulliez et al 2021 JINST 16 P07032

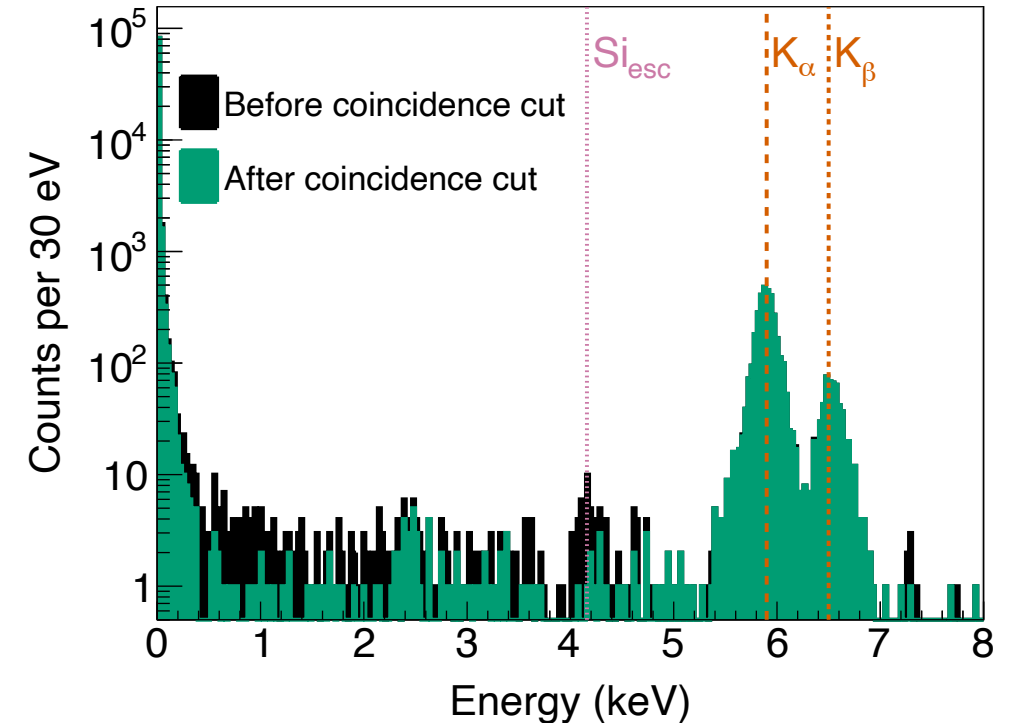
Si wafer detector: energy spectrum

Si wafer detector: 0.35 g Exposure: 55.06 g · days
 Data-taking period: Nov 2020 – Aug 2021 Energy threshold: **10.0 eV**

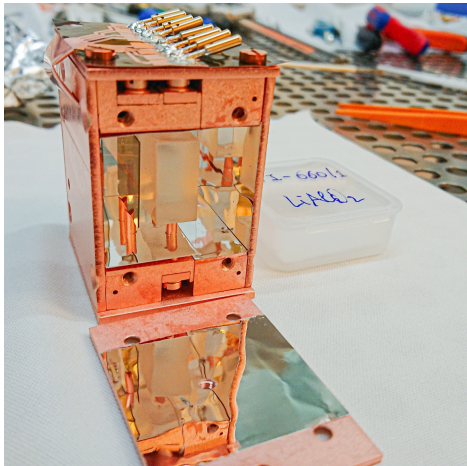
Thin wafer detector is a target
 Bulky detector is a veto to remove coincidence events.



Only 1% of the LEE events are in coincidence with any signal in the bulk detector!



Well performing LiAlO_2 detector allowed to greatly improve the limits for spin-dependent DM-nucleon interactions.

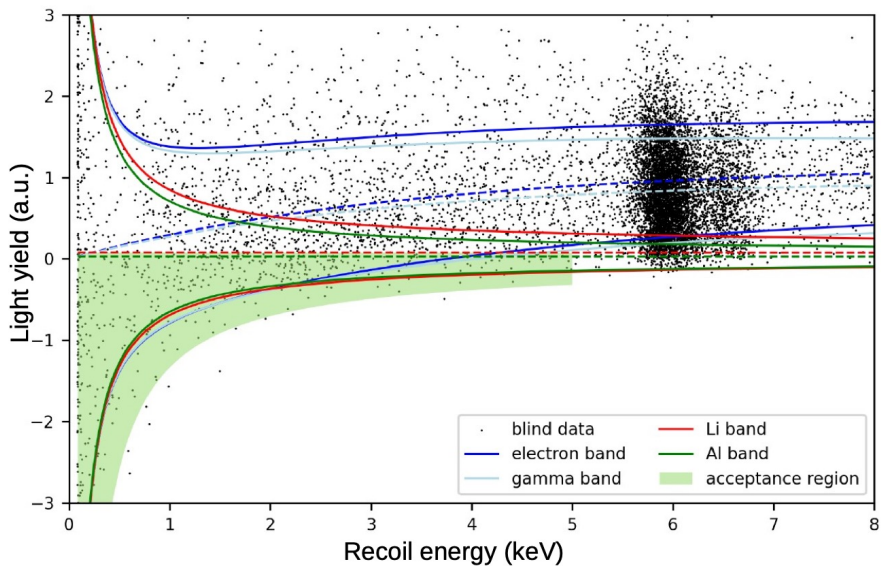


Li1 detector (Run36): 10.5 g LiAlO_2
 Data taking period: Feb 2021 – Aug 2021
 Exposure: 1.161 kg · days
 Nuclear recoil threshold: 83.6 eV

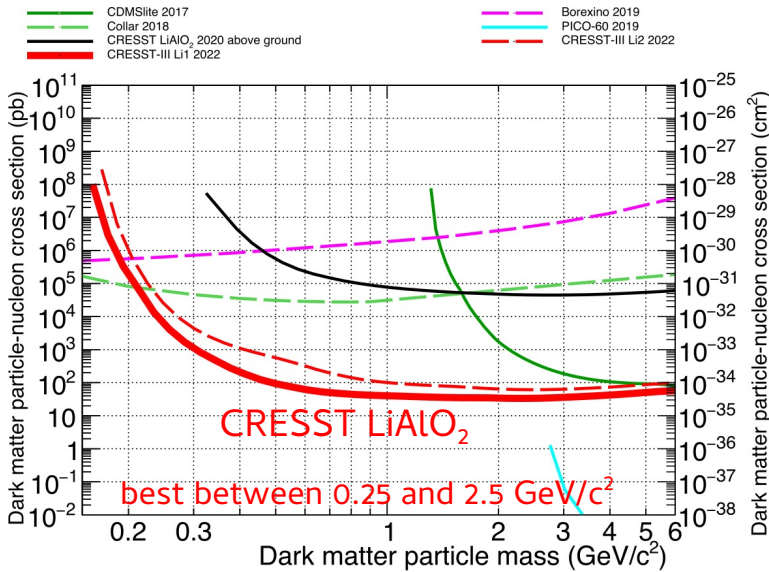
$$\sigma_0^{SD} \propto \mu_N^2 \cdot \frac{J_N + 1}{J_N} \cdot [a_p \cdot \langle S^p \rangle + a_n \cdot \langle S^n \rangle]^2$$

Isotopes sensitive to SD interactions:

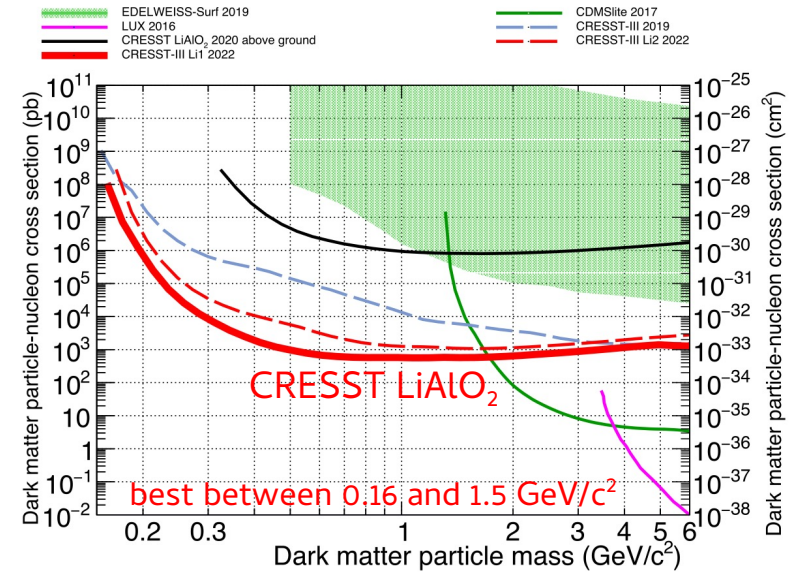
Isotope	$\langle S_p \rangle$	$\langle S_n \rangle$
${}^6\text{Li}$	0.472	0.472
${}^7\text{Li}$	0.497	---
${}^{27}\text{Al}$	0.343	0.0296



Proton



Neutron



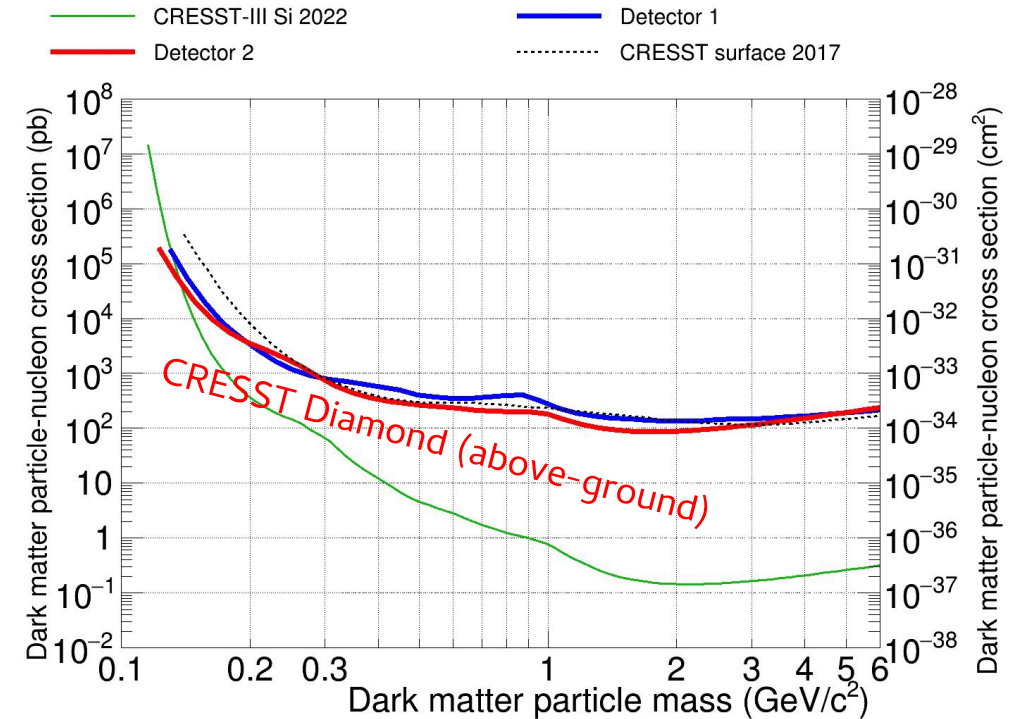
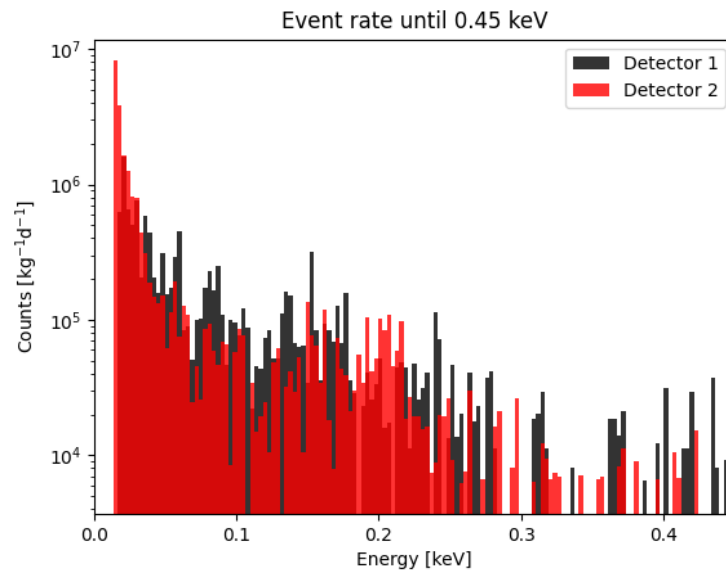
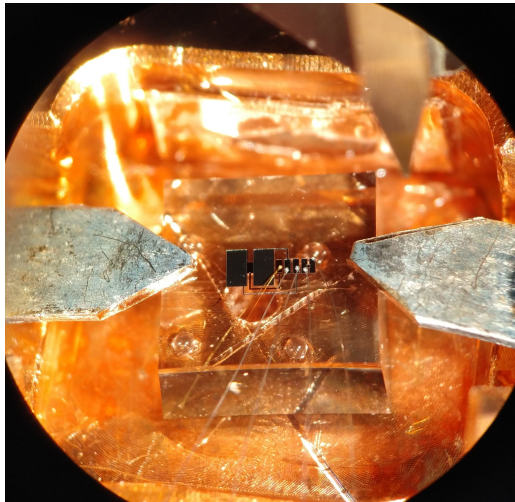
New material: **diamond crystal** operated as a cryogenic detector showed a high potential for sub-GeV DM search.

Diamond detector: 0.175 g

Exposure: 0.27 g · days

37 h of stable data **above-ground**

Nuclear recoil threshold: **16.8 eV**



Why diamond?

- High Debye temperature
- Light nucleus

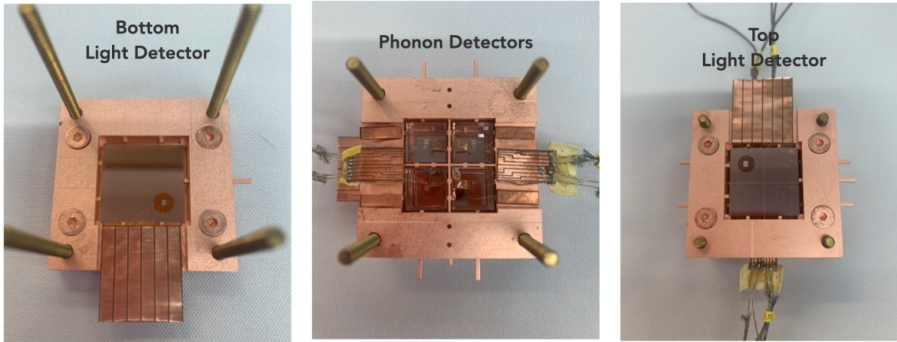
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Poster: "A low-threshold diamond cryogenic detector for sub-GeV Dark Matter searches" by Anna Bertolini

New detector module designs to study the LEE origin

Cm³ module (4 cubes, 2 light detectors):

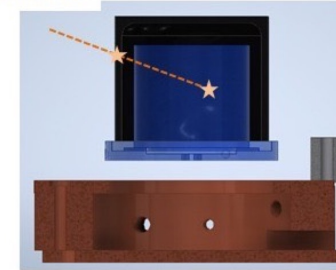
- maximize sensitivity to dark matter recoils
- novel stress-free holding structure



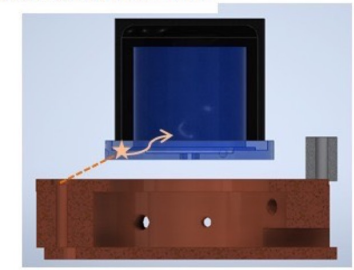
MiniBeaker module :

- instrumented Si beaker -> 4π veto
- active holders -> discriminate events transmitted through the holding structure

• 4π veto



• transmitted events

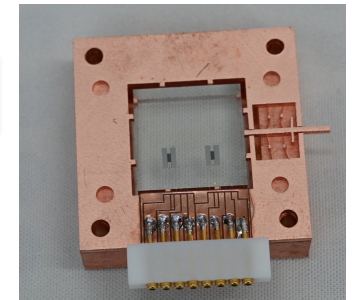


Double TES module:

- novel stress-free holding structure
- discriminate events in the absorber from events in the TES films
- TES-crystal interfaces studies



Talk at the Excess workshop last Saturday by Francesca Pucci



First results from CRESST-III: leading limit for sub-GeV DM with a 30 eV energy threshold detector.

Detector A (Run34) – 23.6 g CaWO₄ Exposure 5.698 kg · days
Data taking period Oct 2016 – Jan 2018 Energy threshold **30.1 eV**

