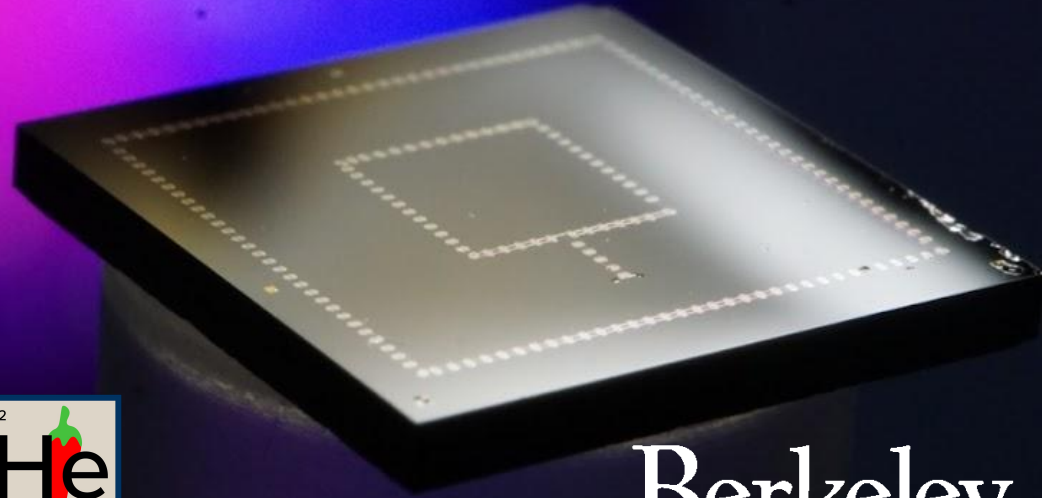


# SPICE: A Search for Light (MeV - GeV) Dark Matter Using Polar Crystal Calorimeters

Roger K. Romani for the  
SPICE/HeRALD Collaboration

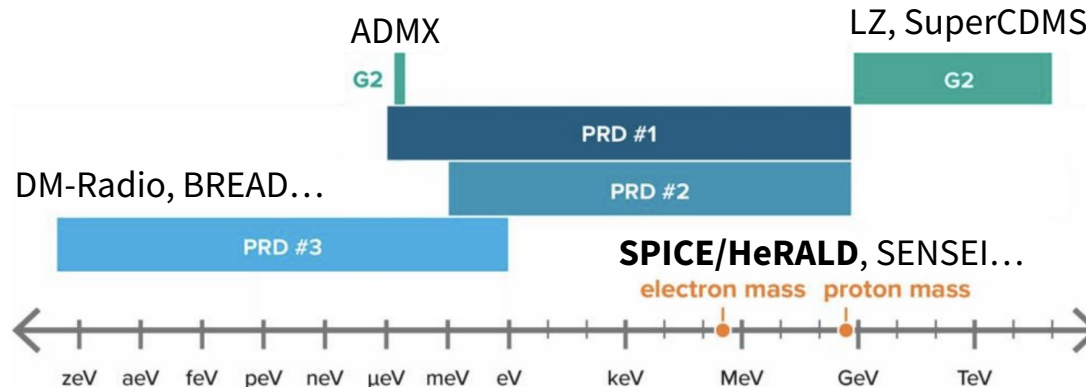


**Berkeley**  
UNIVERSITY OF CALIFORNIA

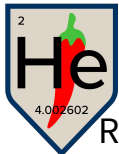
# The Search for DM Direct Detection 30 Years On

Still no sign of DM in the laboratory!

- WIMPs ( $\sim 10$  GeV -  $\sim 100$  TeV): mature technologies, big collaborations, nearing limits
- Axions/ALPs (less than 1 eV): rapid development of many good ideas
- **“Light DM” ( $\sim 100$  keV-  $\sim 1$  GeV): exciting new technologies, new motivations**

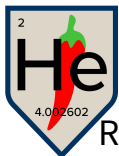
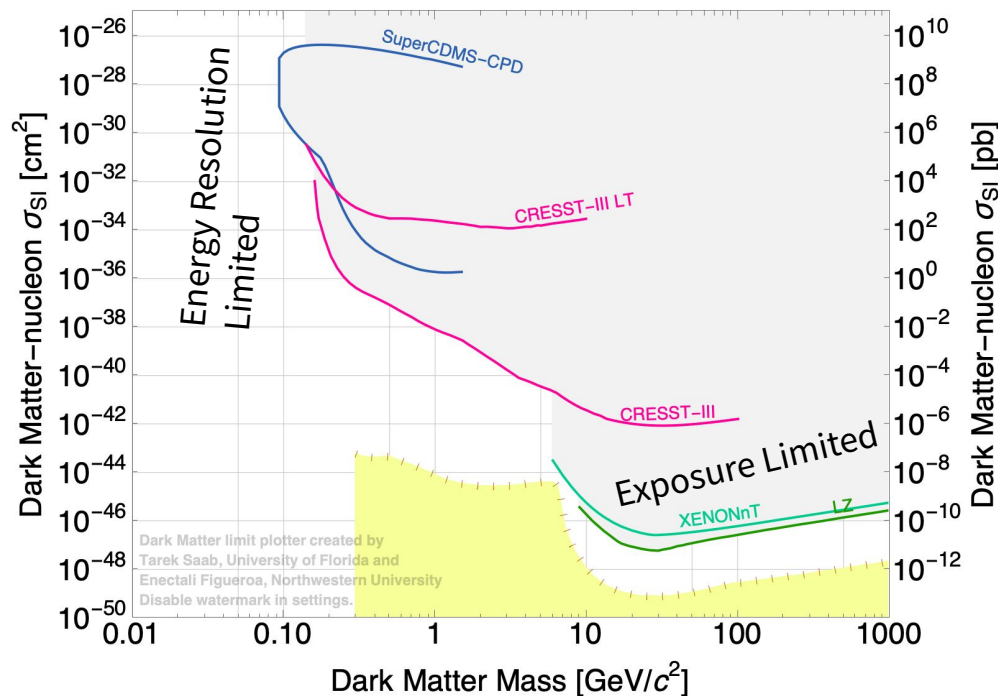


“Basic Research Needs for Dark Matter Small Projects New Initiatives”, 2018



# General Light DM Direct Detection Design Drivers

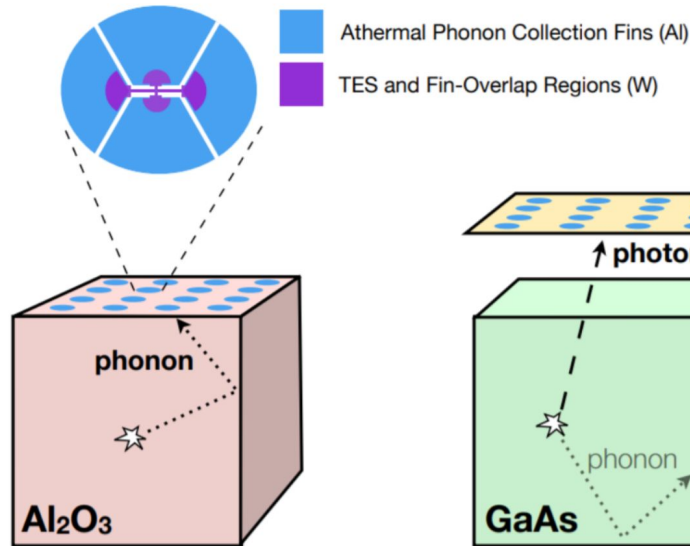
- As  $M_{\text{DM}}$  goes down,  $n_{\text{DM}}$  goes up given constant  $\rho_{\text{DM}}$ 
  - **Don't need huge exposure** to get to interesting cross sections
- As  $M_{\text{DM}}$  goes down,  $E_{\text{dep}}$  goes down as well
  - Main challenge of light DM direct detection: **detect very small energy depositions**
  - True even for different models, materials



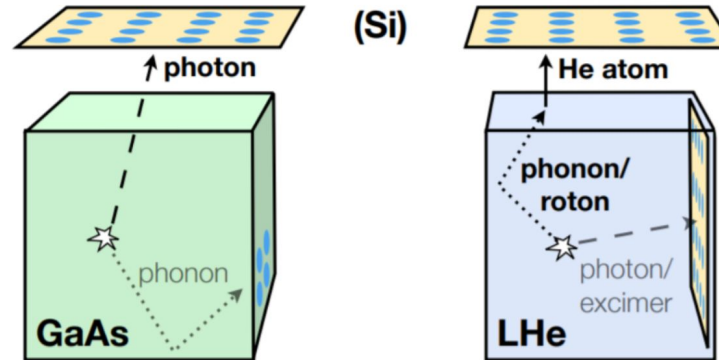
# SPICE/HeRALD\*: A New Light DM Collaboration

Different direct detection mediums unified by a **Transition Edge Sensor** based readout

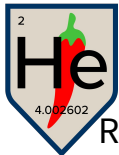
## SPICE



## HeRALD



\*also called  
TESSERACT

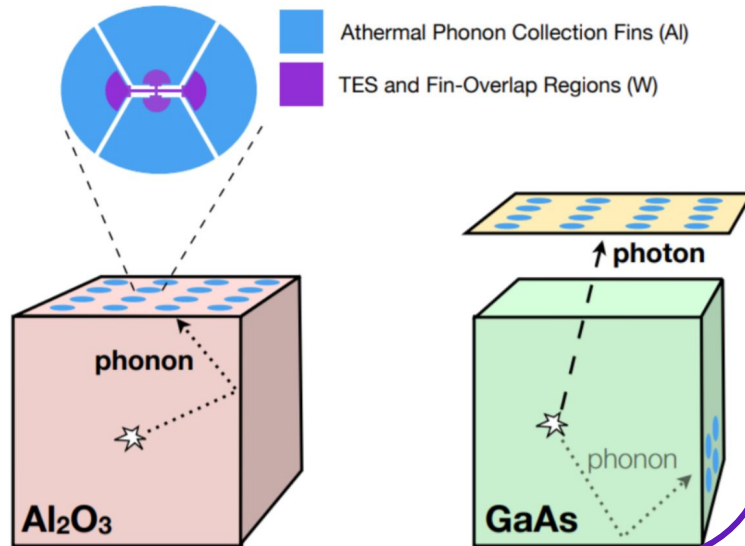


# SPICE/HeRALD: A New Light DM Collaboration

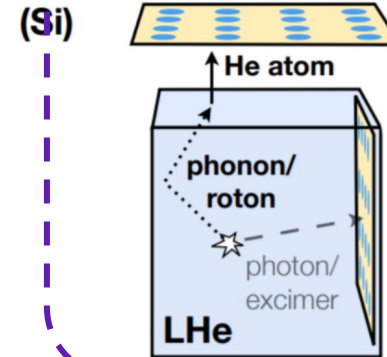
Different direct detection mediums unified by a **Transition Edge Sensor** based readout

This Talk

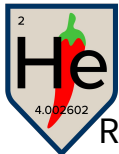
## SPICE



## HeRALD



Stick  
around  
for Prof.  
Scott  
Hertel's  
talk!



# TES Based Calorimetry Basics

Al Fins

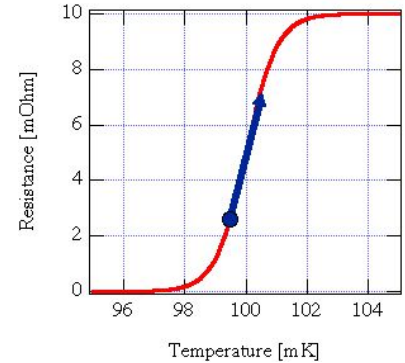
Phonons break Cooper pairs in fins, make quasiparticles

W TES

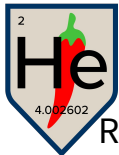
DM interaction

Phonons

Detector Crystal



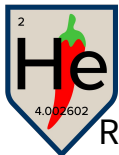
**\*all at ~10 mK in dilution refrigerator**



# Materials

# Different Models, Different Materials of Choice

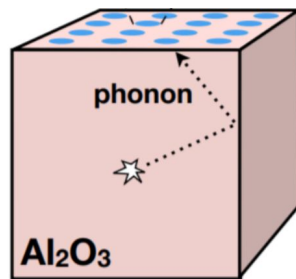
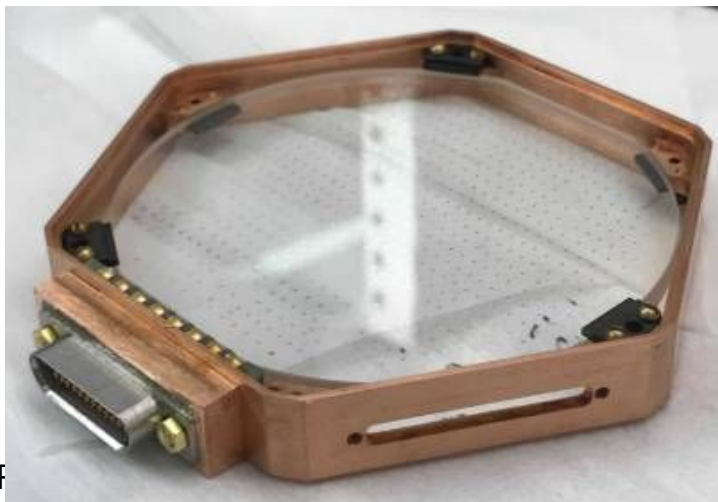
Light dark photon mediator (Sec. III, Fig. 1)			
Detection channel	Quantity to maximize to reach ...		Best materials
	... lower $m_\chi$	... lower $\bar{\sigma}_e$	
(Optical) phonons	$\omega_O^{-1}$ (Eq. (24))	quality factor $Q$ defined in Eq. (27)	SiO <sub>2</sub> , Al <sub>2</sub> O <sub>3</sub> , CaWO <sub>4</sub>
Electron transitions	$E_g^{-1}$ (Eq. (28))	depends on details of electron wavefunctions	InSb, Si
Nuclear recoils	$(A\omega_{\min})^{-1}$ (Eq. (29))	$(Z/A)^2 \omega_{\min}^{-1}$ (Eq. (31))	diamond, LiF
Hadrophilic scalar mediator (Sec. IV, Figs. 2, 3)			
Detection channel	Quantity to maximize to reach ...		Best materials
	... lower $m_\chi$	... lower $\bar{\sigma}_n$	
(Acoustic) phonons	$c_s/\omega_{\min}$ (Eq. (36))	Light mediator: $\omega_{\min}^{-1}$ (Eq. (35))	diamond, SiO <sub>2</sub>
		Heavy mediator: $c_s^{-1}$ or $\omega_{\text{ph}}^{-1}$ or $A\omega_{\text{ph}}$ depending on $m_\chi$ (Eqs. (37), (38), (39))	all complementary
Nuclear recoils	$(A\omega_{\min})^{-1}$ (Eq. (29))	Light mediator: $\omega_{\min}^{-1}$ (Eq. (40))	diamond, LiF
		Heavy mediator: $A$ (Eq. (43))	CsI, Pb compounds



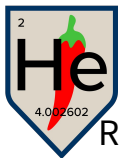
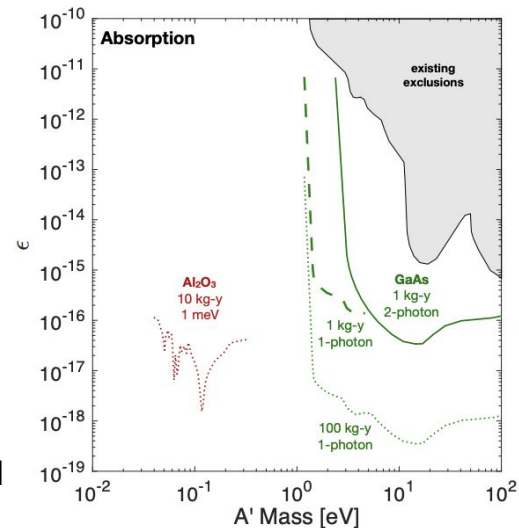
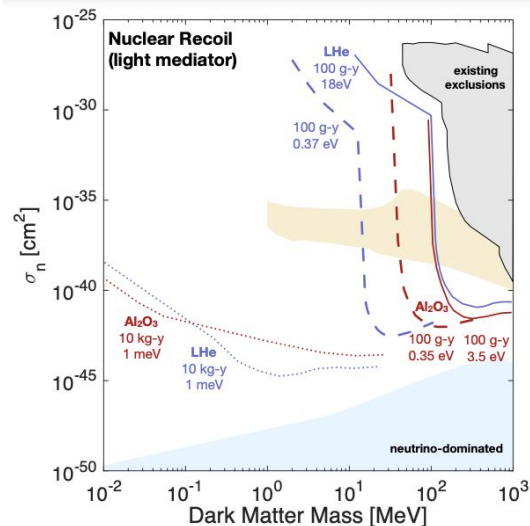


# Sapphire as a SPICE Target

- Low mass oxygen nuclei as NRDM scattering target
- Polar unit cell: optical phonons down to 100s of meV
  - Dark photon coupling due to differently charged nuclei
- Prototype detector has been run

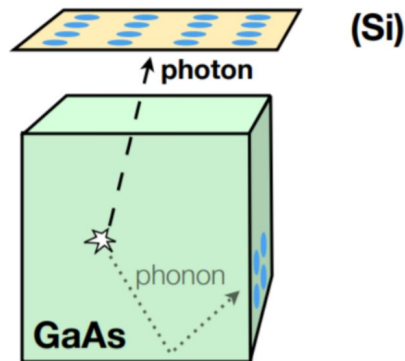
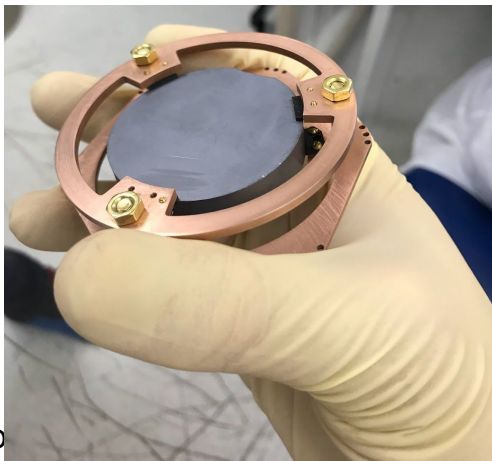


TESSERACT  
SNOWMASS LOI

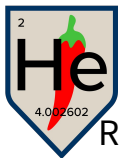
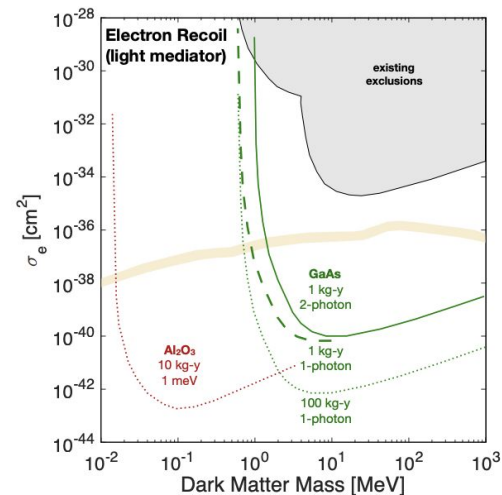
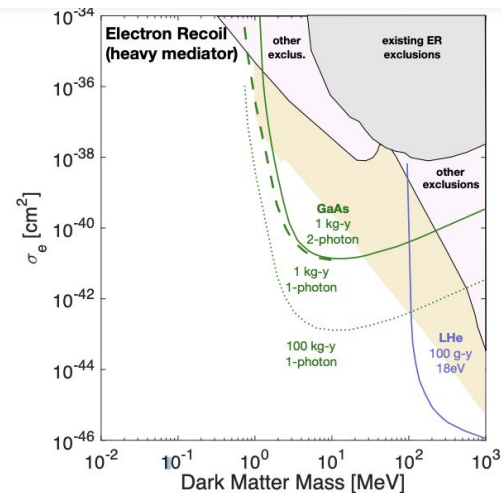


# GaAs as a SPICE Target

- Polar crystal: coupling to dark photons
- Scintillation + phonon signal allows for NR/ER discrimination down to eV scale signals
- GaAs scintillation yield being measured

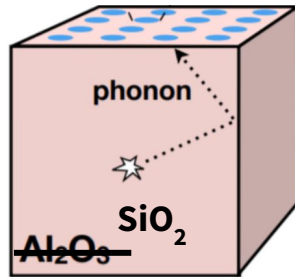


TESSERACT  
SNOWMASS LOI



# SiO<sub>2</sub> as a SPICE Target

- Excellent coupling to dark photons, high “quality factor”
  - See arXiv: 1910.10716
- TESs on SiO<sub>2</sub> substrate tested

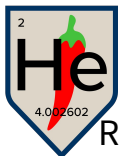
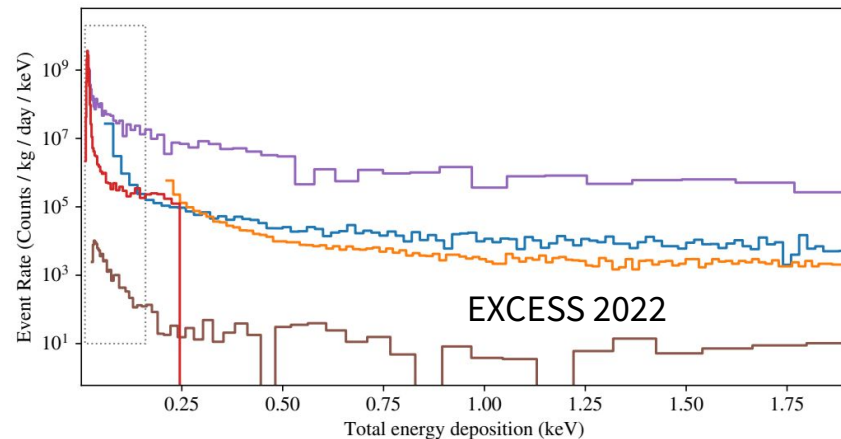
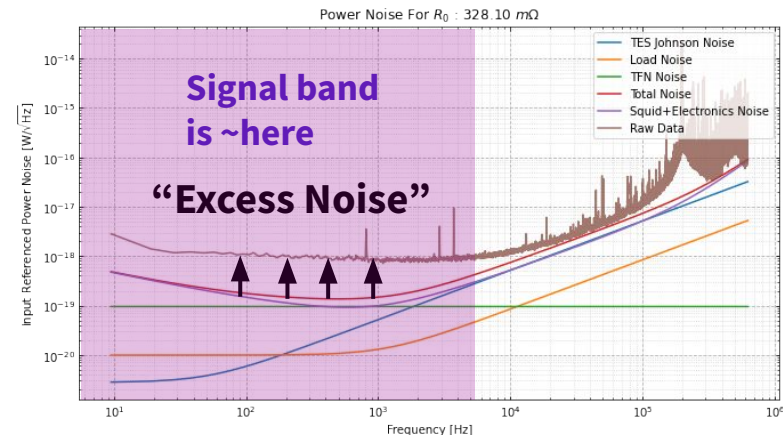


# TES Readout Technology

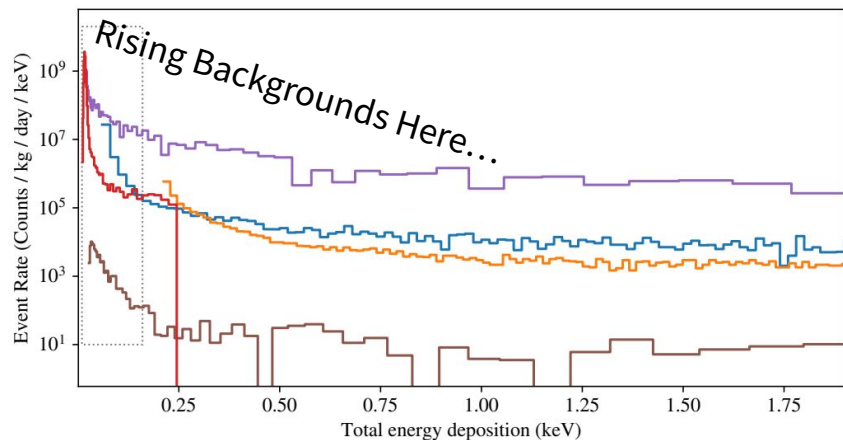
# Two Key Design Drivers:

- Good energy resolution
  - Excess noise limits resolution
- Low background
  - Excess events limits reach

Don't understand either!



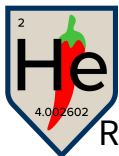
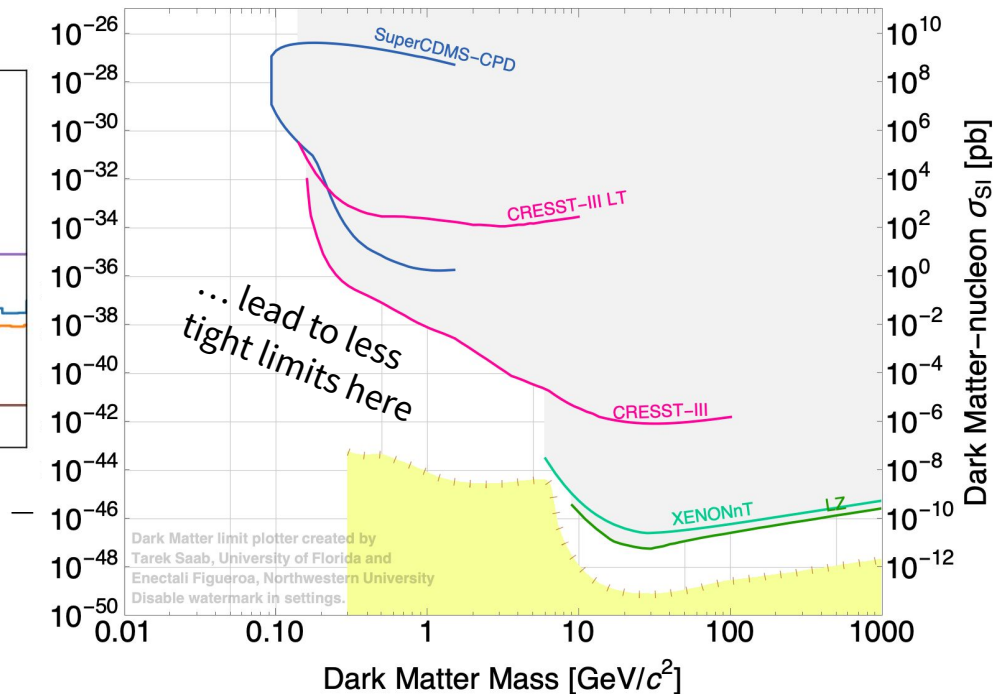
# Low Energy Backgrounds: “LEE”



EXCESS 2022

See EXCESS 2023 Review  
by Dan Baxter

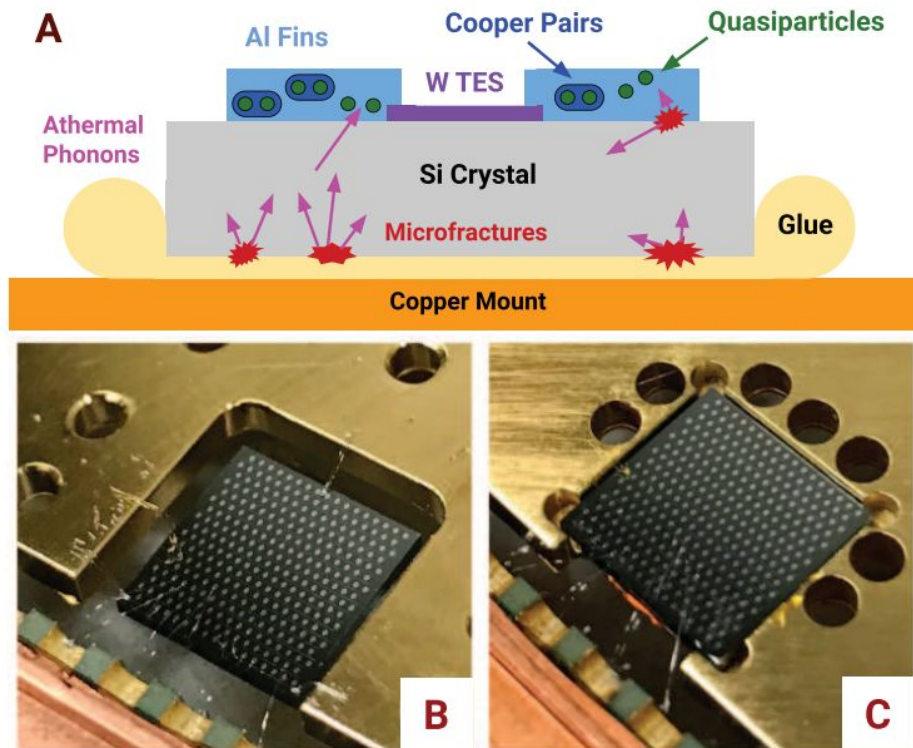
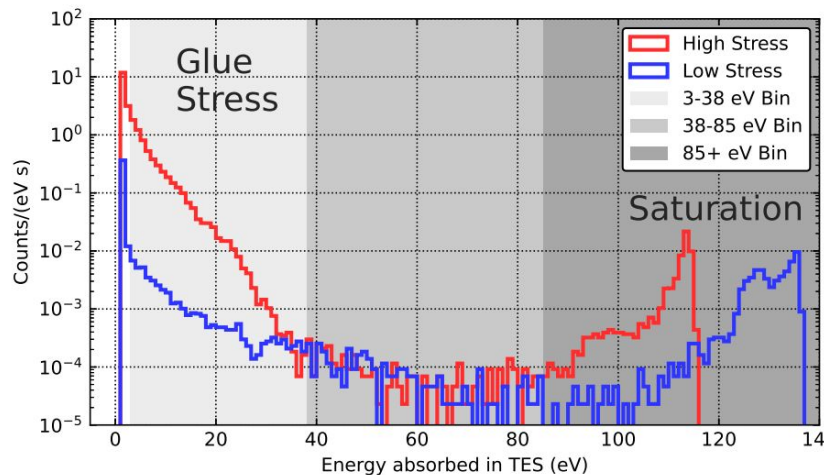
R. K. Romani





# Stress Causes LEE-Like Events!

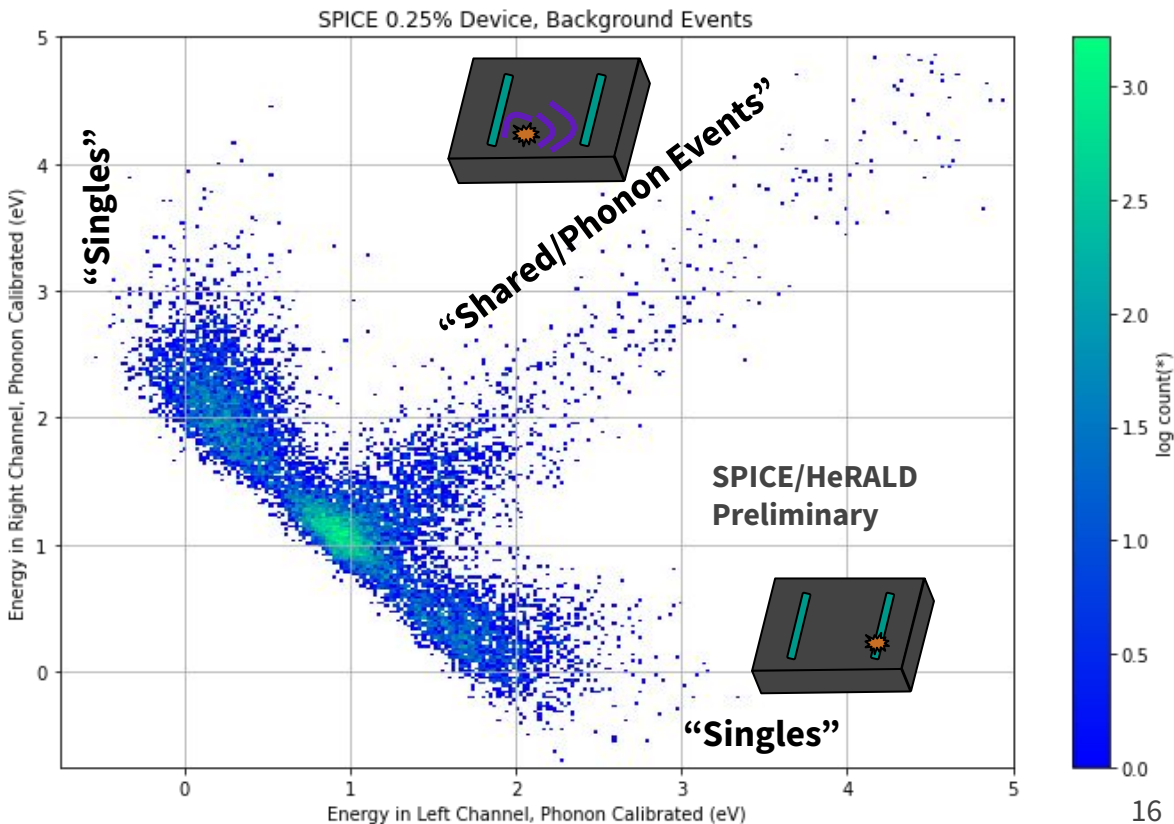
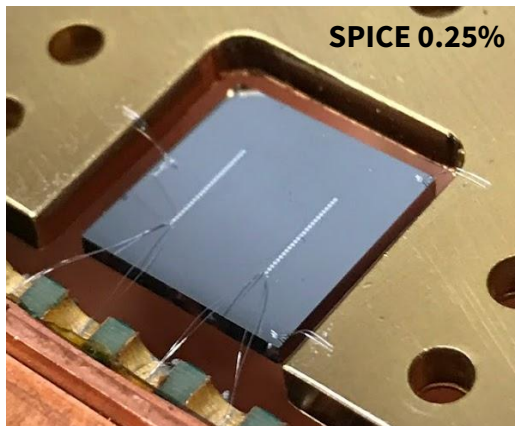
- Compare: high/low stress (hanging/glued) detectors
- Found: stress causes LEE-like events!
- Another source... films?



# Films Cause LEE? Two Channel Devices

Two components of LEE:

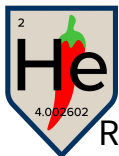
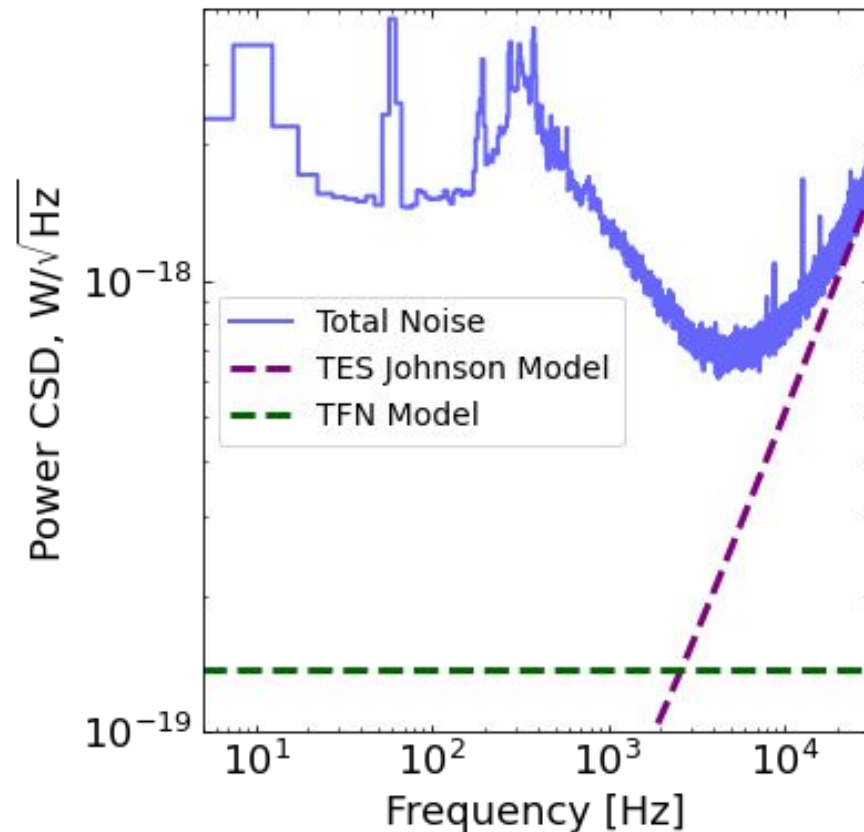
- “Shared” events: phonon pulse shape, partitioning
- “Singles:” single channel partitioning, faster pulse
  - **Film events!**





# Excess Noise

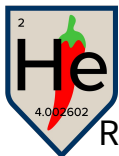
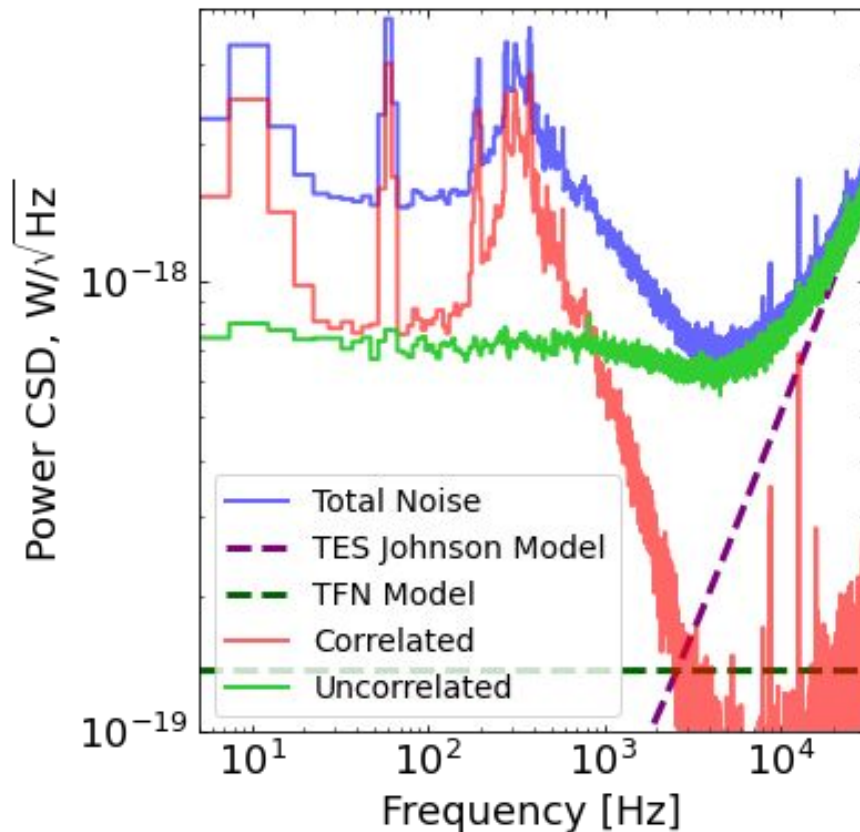
- Signal band: completely dominated by excess noise
  - Can't improve by lowering  $T_c$ , shrinking TESs...
- Split into noise in one sensor, shared between two sensors...



# Excess Noise

- Signal band: completely dominated by excess noise
  - Can't improve by lowering  $T_c$ , shrinking TESs...
- Correlated: consistent with excess very small phonon pulses
- Uncorrelated: consistent with fast events right in TESs

Excess noise: sub-threshold shared/single events?

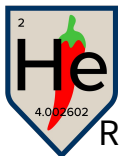


# Everything Goes Down Over Time!

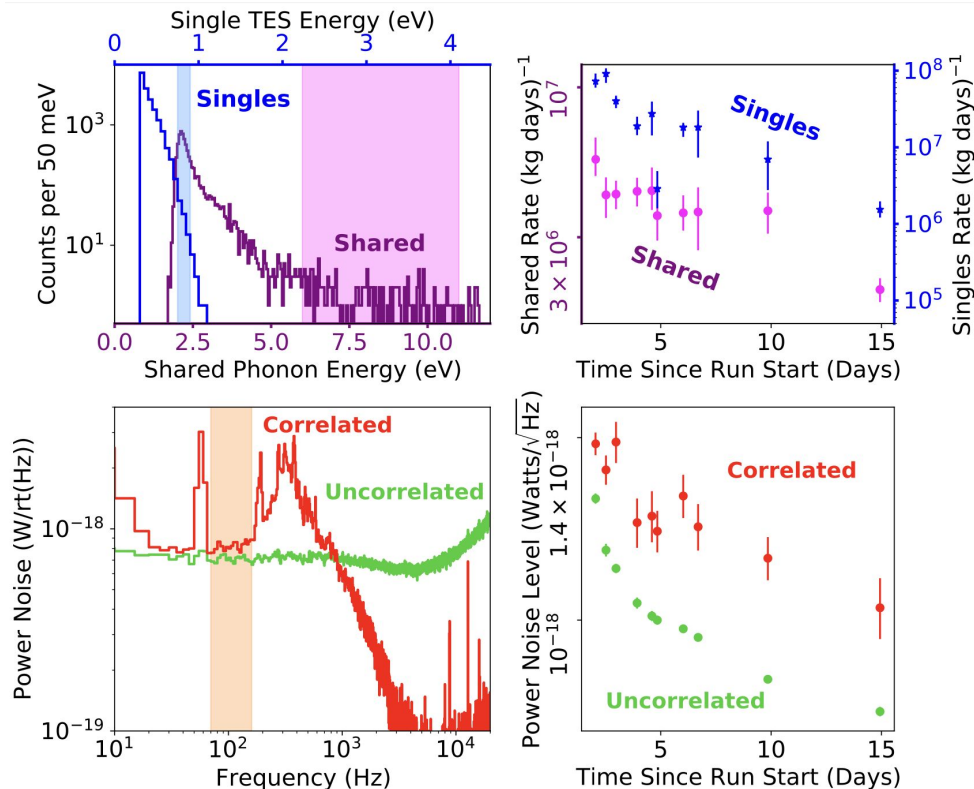
**Relaxation (of stress?) causing both problems: excess events + excess noise**

- Excess noise limiting mass reach
- Excess events limiting cross section reach/backgrounds

Solve this problem, low mass DM searches are open for business!



R. K. Romani



# SPICE/HeRALD: An Exciting Light DM Program

- A suite of materials, with advantages for different model and readouts
- Cutting edge calorimeter R&D, making strong progress towards solving LEE
- An exciting near-term program of DM limits expected

