Sudakov Shoulder Resummation in Thrust and Heavy Jet Mass

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Based on

arXiv:2205.05702 (PRD106.074011)

with Arindam Bhattacharya (Harvard) and Xiaoyuan Zhang (Harvard)

arXiv:2306.xxxxx

with Arindam Bhattacharya (Harvard), Xiaoyuan Zhang (Harvard), Iain Stewart (MIT) and Johannes Michel (MIT)

#### Motivation: HJM is an outlier

Salam and Wicke 2001 (hep-ph/0102343)



Secondly fits for the heavy-jet mass (a very non-inclusive variable) lead to values for  $\alpha_s$ which are about  $10\%$  smaller than for inclusive variables like the thrust or the mean jet mass. This needs to be understood. It could be due to a difference in the behaviour of the perturbation series at higher orders. But in appendix  $\overline{D}$  there is evidence from Monte Carlo

> Chien and Schwartz 2010 (arXiv:1005.1644) NNNLL resummation with NNLO matching



#### Motivation: fixed order pertubation theory

• data for thrust seems matches shape of NNLO theory better than HJM in the far tail





Fits used range  $0.8 < \rho < 0.18$  [Dissertori et al 0712.0327] • Only uses 25% of data bins!

#### What is different about thrust and HJM?

#### 1. Different perturbative behavior: Sudakov Shoulders (this talk)



Could resummation of the Sudakov shoulder improve the theory prediction?

#### What is different about thrust and HJM?

#### 2. Different power corrections (not this talk)



Thurst only involves 3 vectors

• Insensitive to mass scheme

$$
E = |\vec{p}| \text{ or } E = \sqrt{\vec{p}^2 + m_\pi^2}
$$



$$
\rho=\frac{1}{Q^2}\max(M_L^2,M_R^2)
$$

Heavy jet mass involves 4-vectors

Sensitive to mass scheme

Nason & Zanderighi (2301.03607) Nason & Zanderighi (2301.03607)



positive power corrections everywhere



#### Sudakov Shoulders

- Catani and Webber (hep-ph/9710333)
	- C parameter has a right-shoulder
	- discontinuity at C=0.75



- Seymour (hep-ph/9707338)
	- Jet shape has a discontinuity at r=1



#### Sudakov Shoulders

Shoulders are a **generic feature** when range of observable changes order by order

LO:  $C < \frac{3}{4}$ ,  $p, \tau < \frac{1}{3}$ 

• Maximum value at unique trijet configuration

$$
p_2 = \frac{Q}{3}(1, 0, \frac{\sqrt{3}}{2}, -\frac{1}{2})
$$
  

$$
p_1 = \frac{Q}{3}(1, 0, 0, 1)
$$
  

$$
p_3 = \frac{Q}{3}(1, 0, -\frac{\sqrt{3}}{2}, -\frac{1}{2})
$$

- Matrix element is finte at this endpoint
	- $|M| \sim$  constnat for  $\rho \sim \frac{1}{3}$
- Phase space zero volume at endpoint
	- Phase space volume vanishes linearly as  $\rho \rightarrow \frac{1}{3}$
- Cross section vanishes linearly as endpoint approached
	- Leads to kink (discontinuity in first derivative)



### Where do logs come from?

NLO: 4 partons

 $p_1$  $p_3$ Allee  $p_4$  $p_2$ 

- 5 dimensional phse space
	- Choose 5 energies, angles, invariants

 $\cong |{\cal M}_0|^2 \frac{\alpha_s}{4\pi} C_F^2 r \ln^2 r \ \ \twoheadleftarrow$ 

• Compute cross section in small r region

 $T \equiv \max_{\vec{n}} \frac{\sum_j |\vec{p}_j \cdot \vec{n}|}{\sum_i |\vec{p}_j|}$ Thrust axis only region that • 7 possible options for thrust axis contributes for  $r = \frac{1}{3} - \rho = 0.01$ 6  $5<sup>5</sup>$  $\overline{4}$  $T_1$  max  $T_2$  max  $\theta$  3  $T_3$  max  $T_A$  max  $T_{12}$  max  $\overline{2}$  $T_{13}$  max  $T_{14}$  max  $0.26$   $0.28$   $0.30$   $0.32$   $0.34$   $0.36$   $0.38$   $0.40$ cross section Phase space closes off as  $r \rightarrow 0$  $I\sim |{\cal M}_0|^2\frac{\alpha_s}{4\pi}C_F^2\int_0^r\frac{ds_{34}}{s_{34}}\int_{\frac{9}{4} s_{34}}^1\frac{dz}{z}\int_{\frac{1}{3}-r}^{\frac{1}{3}+2r}ds_{234}$ 

Large shoulder logs

## Factorization

1. Soft and/or collinear emissions turn LO partons into jets

2. Derive constraint relating kinematics to ρ or τ



Consider the case of three massive partons (collinear radiation only)

• Phase space is 2d: described by s and t or  $s_{12}=(p_1+p_2)^2$  and  $r = \frac{1}{3}-\rho$ 



Suppose thrust axis points in 1 direction (other cases by permuation)

 $T \equiv \max_{\vec{r}} \frac{2}{\vec{r}}$ Two inequalities ( $T_1 > T_2$  and  $T_1 > T_3$ ), combined together into

$$
\frac{1}{3}-r+2m_1^2-2m_3^2
$$

- For  $m_1$  =  $m_2$  =  $m_3$  = 0 get ¼-r <  $s_{12}$  < ¼+ 2r so phase space cuts off as r $\rightarrow$ 0
- With masses, phsae space is zero unless

$$
m_1^2 < r + m_2^2 + m_3^2 \,
$$

3. Constraint turns double logs for jet mass into shoulder logs

#### Factorization

- 1. Soft and/or collinear emissions turn LO partons into jets
- 2. Derive constraint relating kinematics to ρ or τ
	- Leading power constraint for HJM is

 $m_1^2 < r + m_2^2 + m_3^2$ 

- Solutions for  $r > 0$  and  $r < 0$
- Large logs in both left and right shoulder
- 3. Constraint turns double logs for jet mass into shoulder logs

• Leading power constraint for thrust is

 $\begin{split} \end{split}$   $\begin{split} t < m_1^2 + m_2^2 + m_3^2\ 0 \end{split}$  Only solutions for t > 0

- $t \equiv \tau \frac{1}{3}$ 
	- $\cdot$  t < 0 no phase space limits
		- not sensitive to emissions
	- Only right shoulder



$$
\int_{\frac{1}{3}-r}^{\frac{1}{3}+2r} ds_{12} \int_{0}^{r} dm_{1}^{2} \frac{d\sigma}{dm^{2}} \approx \int_{\frac{1}{3}-r}^{\frac{1}{3}+2r} ds_{12} \int_{0}^{r} dm_{1}^{2} \frac{\ln m_{1}^{2}}{m_{1}^{2}} = r \ln^{2} r
$$
\nPhase space closes off as  $r \to 0$  cross section\n\n• generates  $r$  prefactor\n• generates  $s$  udakov logs

#### Soft radiation

Including soft and collienar radiation constraint is

$$
\mathbf{W^{\text{NN}}}^{\text{max}} + 2p_1k_1 + 2v_2k_2 + 2v_3k_3 < r + m_2^2 + 2p_2k_2 + m_3^2 + 2p_3k_3 + 2v_1k_1
$$

 $t < m_1^2 + m_2^2 + m_3^2 + 2p_1k_1 + 2p_2k_2 + 2p_3k_3 + 2v_1k_1 + 2v_2'k_2 + 2v_3'k_3$ 



#### 1-loop soft function

 $n_1$ 







$$
I_1 = \mathcal{N}\left[4\kappa - \frac{4}{3}\pi \ln 2 + \epsilon c_3\right], \quad I_2 = \mathcal{N}\left[-2\kappa + 2\pi \ln 2 + \epsilon c_4\right]
$$

$$
I_3 = \mathcal{N}\left[4\kappa + \frac{4}{3}\pi \ln 2 + \epsilon c_5\right], \quad I_4 = \mathcal{N}\left[-2\kappa + 2\pi \ln 2 + \epsilon c_6\right]
$$

• Gieseking's constant (transcendality 2 number)

$$
\kappa = \text{Im Li}_2 \left( e^{\frac{i\pi}{3}} \right) = 1.014 \dots
$$
  
• Like Catalan  $C = \text{Im Li}_2 \left( e^{\frac{i\pi}{2}} \right) = 0.915 \dots$   
• Or  $-\frac{\pi^2}{12} = \text{Li}_2(e^{i\pi})$ 

For NNLL resummation we need the constants. Some of these we were only above to get numerically

$$
\kappa = \Im \left[ \text{Li}_2 \left( e^{\frac{i\pi}{3}} \right) \right], \quad c_1 = \Im \left[ \text{Li}_3 \left( \frac{i}{\sqrt{3}} \right) \right], \quad c_2 = \Im \left[ \text{Li}_3 \left( 1 + i\sqrt{3} \right) \right], \quad c_3 = -1.89958
$$
\n
$$
c_4 = -3.83452, \quad c_5 = -12.3488, \quad c_6 = 5.56704, \quad c_7 = 15.9482, \quad c_8 = 3.52263
$$

 $I_{4}$ 

#### 1-loop soft function

Combine integrals together to produces trijet hemisphere soft function

$$
q_H
$$
\n
$$
S(q_L, q_H) = \int d^6 q_i S_6(q_i) \delta(q_L - q_1 - q_2 - q_3) \delta(q_H - q_1 - q_2 - q_3)
$$
\n
$$
k_1 \underbrace{\sum_{k_1 \text{ times } k_2 \text{ times } k_1 \text{ times } k_2 \text{ times } k_
$$

#### Resummed distribution

HJM, ρ < Ψ<sub>3</sub> 
$$
\frac{1}{\sigma_1} \frac{d\sigma_g}{dr} = \frac{\Pi_g(\partial_{\eta_\ell}, \partial_{\eta_h})}{\Pi_g(\partial_{\eta_\ell}, \partial_{\eta_h})} \left(\frac{rQ}{\mu_s}\right)^{\eta_\ell} \left(\frac{rQ}{\mu_s}\right)^{\eta_h} \frac{e^{-\gamma_E(\eta_\ell + \eta_h)}}{\Gamma(2 + \eta_\ell + \eta_h)} \frac{\sin(\pi \eta_\ell)}{\sin(\pi (\eta_\ell + \eta_h))} -2C_A \left(\frac{\alpha_s}{4\pi}\right) \ln r
$$
\nHJM, ρ > Λ<sub>3</sub> 
$$
\frac{1}{\sigma_1} \frac{d\sigma_g}{ds} = \frac{\Pi_g(\partial_{\eta_\ell}, \partial_{\eta_h})}{\Pi_g(\partial_{\eta_\ell}, \partial_{\eta_h})} s \left(\frac{sQ}{\mu_s}\right)^{\eta_\ell} \left(\frac{sQ}{\mu_s}\right)^{\eta_h} \frac{e^{-\gamma_E(\eta_\ell + \eta_h)}}{\Gamma(2 + \eta_\ell + \eta_h)} \frac{\sin(\pi \eta_h)}{\sin(\pi (\eta_\ell + \eta_h))} -4C_F \left(\frac{\alpha_s}{4\pi}\right) \ln r
$$
\nThrust, τ > Υ<sub>3</sub> 
$$
\frac{1}{\sigma_1} \frac{d\sigma_g}{dt} = \frac{\Pi_g(\partial_{\eta_\ell}, \partial_{\eta_h}) t}{\pi} \left(\frac{tQ}{\mu_s}\right)^{\eta_\ell} \left(\frac{tQ}{\mu_s}\right)^{\eta_h} \frac{e^{-\gamma_E(\eta_\ell + \eta_h)}}{\Gamma(2 + \eta_\ell + \eta_h)}
$$
\n
$$
\frac{e^{-\gamma_E(\eta_\ell + \eta_h)}}{\Gamma(2 + \eta_\ell + \eta_h)}
$$
\n
$$
\frac{\sin(\pi \eta_h)}{\sin(\pi (\eta_\ell + \eta_h))} -2C_A \left(\frac{\alpha_s}{4\pi}\right) \ln r
$$
\n
$$
\frac{1}{\sigma_1} \frac{d\sigma_g}{dt} = \frac{\Pi_g(\partial_{\eta_\ell}, \partial_{\eta_h}) t}{\pi} \left(\frac{tQ}{\mu_s}\right)^{\eta_\ell} \left(\frac{tQ}{\mu_s}\right)^{\eta_h} \frac{e^{-\gamma_E(\eta_\ell + \eta_h)}}{\Gamma(2 + \eta_\ell + \eta_h)}
$$
\n
$$
\frac{e^{-\
$$

Check fixed order expansion

$$
\frac{1}{\sigma_1} \frac{d\sigma^{\text{sub}}}{dr} = \frac{\alpha_s}{4\pi} \left\{ -\frac{1}{2} (2C_F + C_A) \Gamma_0 r \ln^2 r + \left[ (C_A + 2C_F) \Gamma_0 + \gamma_{jg} + 2\gamma_{jq} + 2\gamma_{sg} + 4\gamma_{sq} \right] r \ln r \right\}
$$
\n
$$
= \frac{\alpha_s}{4\pi} \left\{ -2(2C_F + C_A) r \ln^2 r + \left[ 2C_F \left( 1 + 4 \ln \frac{4}{3} \right) + C_A \left( \frac{1}{3} + 4 \ln \frac{4}{3} \right) + \frac{4}{3} n_f T_F \right] r \ln r \right\} + C_0 (\alpha_s) + C_1 (\alpha_s) r
$$

#### Check fixed order expansion

Expanding to order  $\alpha_s$  for the HJM left shoulder

$$
\frac{1}{\sigma_1} \frac{d\sigma^{\text{sub}}}{dr} = \frac{\alpha_s}{4\pi} \left\{ -\frac{1}{2} (2C_F + C_A) \Gamma_0 r \ln^2 r + \left[ (C_A + 2C_F) \Gamma_0 + \gamma_{jg} + 2\gamma_{jq} + 2\gamma_{sg} + 4\gamma_{sq} \right] r \ln r \right\}
$$
\n
$$
= \frac{\alpha_s}{4\pi} \left\{ -2(2C_F + C_A) r \ln^2 r + \left[ 2C_F \left( 1 + 4 \ln \frac{4}{3} \right) + C_A \left( \frac{1}{3} + 4 \ln \frac{4}{3} \right) + \frac{4}{3} n_f T_F \right] r \ln r \right\} + c_0(\alpha_s) + c_1(\alpha_s) r
$$

And similarly (with differnet constants) for the HJM right shoulder and thrust

Excellent agreement with event 2 (NLO fixed order)

constant and linear term not predicted (more on this soon)



used cutoff =  $10^{-12}$ , 12 trillion events for event2

#### Resummed distribution



#### UV divergences

Factorization formula is of the form

$$
\frac{1}{\sigma_\text{LO}}\frac{d\sigma_g}{d\rho}=\Pi_g\,\int_0^\infty dm_\ell^2\int_0^\infty dm_h^2\,\,\widetilde{sj}_g\cdot\left(\frac{Qm_l^2}{\mu_{s\ell}}\right)^a\left(\frac{Qm_h^2}{\mu_{sh}}\right)^b\frac{1}{m_l^2m_h^2}\frac{e^{-\gamma_E(a+b)}}{\Gamma(a)\Gamma(b)}\Bigg|_{\substack{a=\eta_\ell\\b=\eta_h}}\times\,(r+m_h^2-m_\ell^2)\theta(r+m_h^2-m_\ell^2)
$$

- can have  $r \ll 1$ ,  $m_h^2 \gg 1$ ,  $m_\ell^2 \gg 1$
- integral is linearly UV divergent in this region for  $0 < a,b < 1$

Simplest to **take two more derivatives** and consider

$$
f(r) \equiv \frac{1}{\Gamma(a)} \frac{1}{\Gamma(b)} \int_0^\infty dx \int_0^\infty dy \ x^{a-1} y^{b-1} \delta(r+y-x)
$$
  
= 
$$
\frac{1}{\Gamma(a+b)} \left[ r^{a+b-1} \frac{\sin(\pi a)}{\sin(\pi(a+b))} \theta(r) + (-r)^{a+b-1} \frac{\sin(\pi b)}{\sin(\pi(a+b))} \theta(-r) \right]
$$

- Now UV finite
	- Will need to integrate twice and set integration constants  $c_0 + c_1 r$
- Formula is valid for  $p < \frac{1}{3}$  (r>0) as well as  $p > \frac{1}{3}$  (r<0)
- Still has Sudakov Landau pole at *a+b=1,2,3*,...

#### Position vs momentum space

In momentum space, distribution is complicated and non-analytic

$$
f(r) \equiv \frac{1}{\Gamma(a)} \frac{1}{\Gamma(b)} \int_0^\infty dx \int_0^\infty dy \ x^{a-1} y^{b-1} \delta(r+y-x)
$$
  
= 
$$
\frac{1}{\Gamma(a+b)} \left[ r^{a+b-1} \frac{\sin(\pi a)}{\sin(\pi(a+b))} \theta(r) + (-r)^{a+b-1} \frac{\sin(\pi b)}{\sin(\pi(a+b))} \theta(-r) \right]
$$

In **position space**, distribution is remarkably simple

$$
\tilde{f}(z) = \int_{-\infty}^{\infty} dr f(r) e^{izr} = (-iz)^a (iz)^b
$$

difficult to compute most carefully track analytic continuation

**No longer has Sudakov landau poles** at a+b = 1,2,3...

- In **Laplace space** distribution is **not simple**
	- For 1-sided Laplace transform, need to flip sign in exponent to make integral convergent

$$
\mathcal{L}[f](\nu) = \int_0^\infty dr \ e^{-\nu r} f(r) + \int_{-\infty}^0 dr \ e^{\nu r} f(r) = \nu^{-a-b} \frac{\sin(\pi a) + \sin(\pi b)}{\sin(\pi(a+b))}
$$

• Still has Sudakov Landau pole

# $Γ<sub>0</sub>$  approximation

To undestand better what is happening, consider the " $\Gamma_0$  approximation"

- set all  $\gamma_j = 0$  and  $\beta = 0$
- keep only leading cusp anomalous dimension  $\Gamma_0$

#### Momentum space

$$
\frac{1}{\sigma_{LO}} \frac{d^3 \sigma_g}{d\rho^3} = e^{-2\hat{\alpha}C_F \ln^2 \frac{\mu_h}{\mu_{jh}} - 2\hat{\alpha}C_F \ln^2 \frac{\mu_{sh}}{\mu_{jh}} - \hat{\alpha}C_A \ln^2 \frac{\mu_h}{\mu_{j\ell}} - \hat{\alpha}C_A \ln^2 \frac{\mu_{s\ell}}{\mu_{j\ell}} \times \frac{e^{-\gamma_E(a+b)}}{\Gamma(a+b)}
$$
\n
$$
\times \left[ \frac{1}{r} \left( \frac{rQ}{\mu_{s\ell}} \right)^a \left( \frac{rQ}{\mu_{sh}} \right)^b \frac{\sin(\pi a)}{\sin(\pi(a+b))} \theta(r) + \frac{1}{(-r)} \left( \frac{-rQ}{\mu_{s\ell}} \right)^a \left( \frac{-rQ}{\mu_{sh}} \right)^b \frac{\sin(\pi b)}{\sin(\pi(a+b))} \theta(-r) \right]
$$
\n
$$
\mu_{s\ell} = \mu_{sh} = |r|Q \qquad \text{canonical}
$$
\n
$$
\mu_{s\ell}^2 = \mu_{jh}^2 = Q\mu_{s\ell} = Q\mu_{sh}
$$
\n
$$
\left( \frac{1}{\sigma_{LO}} \frac{d^3 \sigma}{d\rho^3} \right) = \frac{1}{r} e^{-\Gamma_0 \frac{1}{2} C_A \hat{\alpha} \ln^2 |r| - \Gamma_0 C_F \ln^2 |r|} \frac{e^{-\gamma_E(a+b)}}{\Gamma(a+b)} \left[ \frac{\sin(\pi a)}{\sin(\pi(a+b))} \theta(r) - \frac{\sin(\pi b)}{\sin(\pi(a+b))} \theta(-r) \right]
$$
\n
$$
a = -C_A \hat{\alpha} \Gamma_0 \ln |r|
$$
\n
$$
b = -2C_F \hat{\alpha} \Gamma_0 \ln |r|
$$

#### • Sudakov Landau poles are there

- Not of the double-logarithmic form  $exp[ L g_1(\alpha L) ]$ 
	- Includes problematic subleading terms

$$
\tilde{\sigma}(z) = \int_{-\infty}^{\infty} dr \ e^{izr} \ \frac{1}{\sigma_{LO}} \frac{d^3 \sigma}{d\rho^3}
$$
\n
$$
= e^{-2\hat{\alpha}C_F \ln^2 \frac{\mu_h}{\mu_{jh}} - 2\hat{\alpha}C_F \ln^2 \frac{\mu_{sh}}{\mu_{jh}} - \hat{\alpha}C_A \ln^2 \frac{\mu_h}{\mu_{j\ell}} - \hat{\alpha}C_A \ln^2 \frac{\mu_{s\ell}}{\mu_{j\ell}} \left( -iz \frac{\mu_{s\ell}e^{\gamma_E}}{Q} \right)^{-a} \left( iz \frac{\mu_{sh}e^{\gamma_E}}{Q} \right)^{-b}
$$
\ncanonical complex

\n
$$
\mu_{s\ell} = i \frac{Qe^{-\gamma_E}}{z}, \quad \mu_{sh} = -i \frac{Qe^{-\gamma_E}}{z}
$$
\n
$$
\tilde{\sigma}(z) = \exp \left[ -\frac{1}{2} \hat{\alpha} C_A \Gamma_0 \ln^2(-ize^{\gamma_E}) - \hat{\alpha} C_F \Gamma_0 \ln^2(ize^{\gamma_E}) \right]
$$

Position space

- Sudakov Landau poles are gone
- Exactly of the double-logarithmic form  $exp[ L g_1(\alpha L) ]$ 
	- Problematic subleading terms are absent

# $Γ<sub>0</sub>$  approximation



## Matching to the dijet region



We need to match bewteen the shoulder region and the dijet region

- want pure shoulder for  $p > 0.25$
- fade to dijet by  $p < 0.15$

## Matching to the dijet region

Turn off resummation in shoulder and dijet region by using ρ dependent profile functions for soft and jet scales:



#### Scale variations



## Final result



- Matched result is
	- 20% higher than NLO
	- More important than NNLO correction
- Pure dijet + NLO is 5% to 20% larger for  $0.15 < \rho < 0.3$

things go nuts because denominator is zero

#### **Conclusions**

- Sudakov shoulders are large logarithms associated with perturbative phase space boundaries
	- Heavy jet mass has a left Sudakov shoulder ( $ρ$  < <sup>1</sup>/<sub>3</sub>) and a right sudakov shoulder ( $ρ$  > <sup>1</sup>/<sub>3</sub>)
		- large logs extend down to  $p = 0.15$  where  $\alpha_s$  fits are done
	- Thrust only has a right shoulder  $(\tau > \frac{1}{3})$ 
		- shoulder resummation is not critical for  $\alpha_s$  fits for thrust



- Factorization theorem for the Sudakov shoulder allows for systematic resummation
	- Scales must be set in **position space** before Fourier transforming to avoid spurious Sudakov Landau poles
	- We resummed the HJM shoulder to NNLL, matching to NNLL dijet resummation and NLO fixed order



#### • **Effect is signifcant**

• 5% to 100% larger than dijet+NLO for 0.15 < ρ < 0.3

#### Next steps

- Extend matching to NNLO and  $N^3$ LL dijet
- Study **power corrections**
	- Factorization formula valid in the trijet region gives operator definition
	- How to interpolate power corrections from dijet to shoulder region?
	- How many non-perturbative parameters are needed?
- $0.40$  Extract  $\alpha_s$  from e<sup>+</sup>e<sup>-</sup> data