



Effective Spin foam models

An overview

Seth Kurankyi Asante, FSU Jena

*[Bianca Dittrich, Hal Haggard, Sebastian Steinhaus, José Diogo Simão,
Alexander Jercher, José Padua-Argüelles, Taylor Brysiewicz*]*

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Fort Lauderdale, Florida, USA

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Emmy
Noether-
Programm

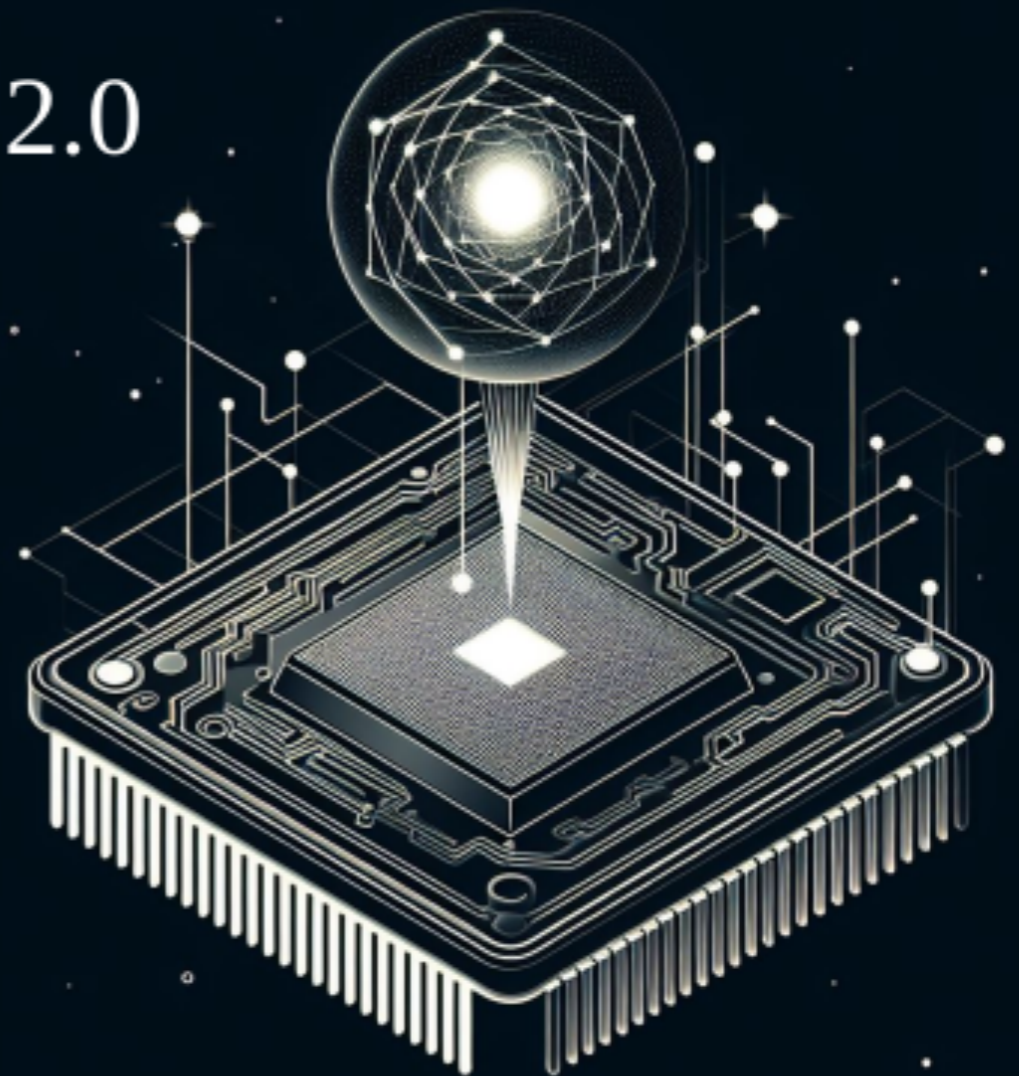
DFG Deutsche
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Quantum Gravity on the Computer 2.0

September 9 - 13

FSU Jena



What can we learn from (discrete) gravitational path integrals?

Quantum geometry

Quantum gravity

Semi-classical approximations

Topological transitions

Black hole dynamics

Early universe Cosmology applications

Renormalization/Continuum limit

Quantum field theory in curved space time

What are 'effective spin foam models'?

They are defined as discrete geometric path integrals (sums) for quantum gravity

- maintain the 'key dynamical principles' of quantum geometry à la LQG

They are **effective** because:

- they provide spin foam models with efficient and computable dynamics
- they provide a general family of models by imposing constraints differently

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They can provide many insights into discrete path integral models

- easy construction of Lorentzian and Euclidean models
- applications to semi-classical geometries, cosmology, ...
- avenue to study **continuum limit** of discrete models

Outline

❖ **Constructing effective spin foams**

- ⦿ Discrete quantum geometries

❖ **Testing the model**

- ⦿ Semi-classical analysis

❖ **Cosmology applications**

- ⦿ Mini superspace models

❖ **Continuum limit**

- ⦿ First Steps: Perturbative



Effective Spin foams

Constructing the models

[SKA, Dittrich, Haggard, Padua-Argüelles, Brysiewicz]
arXiv: 2004.07013, 2011.14468, 2104.00485, 2402.17080

[Ashtekar, Rovelli, Smolin, Thiemann, Lewandowski, Perez, Bianchi, Freidel, Corichi, Dittrich, Varadarajan, Livine, Bonzom...]

many more ...

Construct a simple spin foam model where:

- area variables fundamental

- enlarged space of (discrete) length geometries

support from: LQG, black hole entropy, thermodynamics, generalized geometry, geometric entanglement, strings

[Bekenstein, Ryu, Takayanagi, Cattaneo, Perez, Jacobson, Headrich, Zweibach, Schuller, Wohlfahrt,...]

[Rovelli, Smolin, Ashtekar, Lewandowski, Corichi, Wieland, Freidel, Geiller, Pranzetti...]

- spectra for area operators

Space-like areas

$$a_S = \gamma \ell_P^2 \sqrt{j(j+1)} \sim \gamma \ell_P^2 j, \quad j \in \mathbb{N}/2$$

Time-like areas

$$a_T = \ell_P^2 n, \quad n \in \mathbb{N}/2$$

The Construction

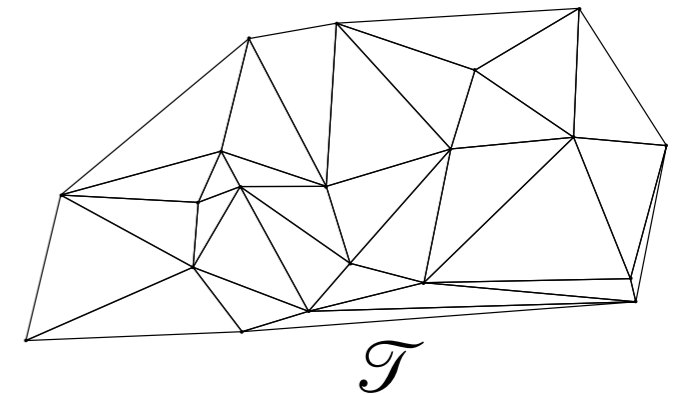
Area Regge Calculus

Simple action (4D) [Regge '61; Rovelli '93; Mäkelä '94, Barrett, Roček, Williams '97, SKA, Dittrich, Haggard '18...]

Area Regge action:
$$S_{\text{ARC}} = - \sum_{t \in \text{bulk}} a_t \epsilon_t(a_{t'}) - \sum_{t \in \text{bdry}} a_t \psi_t(a_{t'})$$

discrete GR action for area configurations

relevant for LQG and spin foam models [Rovelli '93; Barrett et al...]



Classical dynamics: $\epsilon_t(a_{t'}) = 0$ does **not** reproduce discrete GR dynamics

vanishing curvature + non-shape matching

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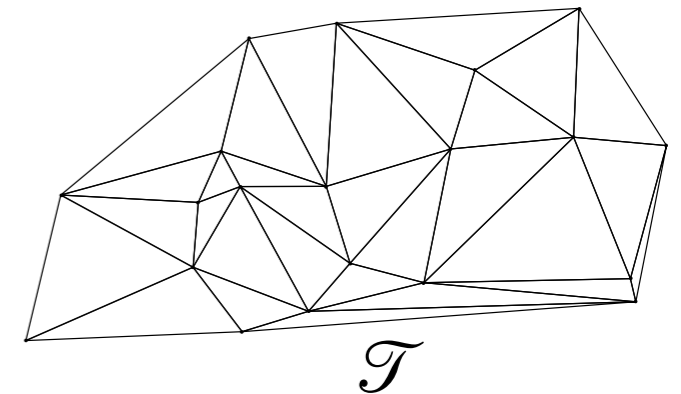
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Preliminary state-sum model:
$$Z = \sum_{\{a\}} \mu(a) \exp(i S_{\text{ARC}}(a))$$

discrete areas

Add constraints to reproduce discrete GR [Dittrich, Speziale '07]

The Construction

Constrain area variables

Area-length constraints set of polynomial equations

$$\left\{ a_t^2 = \frac{1}{16}(4l_1^2 l_2^2 - (l_1^2 + l_2^2 - l_3^2)^2), \quad \forall t \right\} \quad \text{cf. (part of) simplicity constraints}$$

[Heron of Alexandria, AD 70]

The Construction

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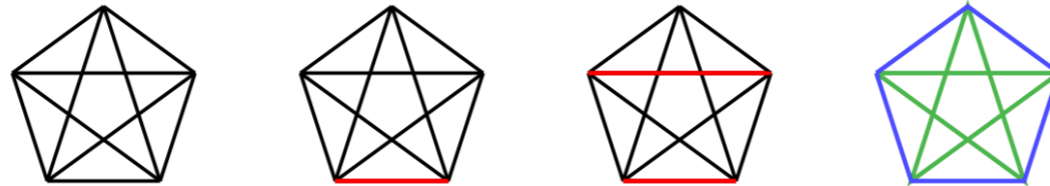
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Homotopy continuation (tools from numerical algebraic geometry)

[Brysiewicz, SKA '24]

A **4-simplex** with 10 areas has **up to 64 length** configurations



Galois group

$$G = \mathbb{Z}^2 \wr \mathbb{S}^{32}$$

\implies not solvable

Generically, area-length constraints has **no/few** solutions

Diophantine equations for discrete areas

[Dittrich, Haggard, SKA '21]

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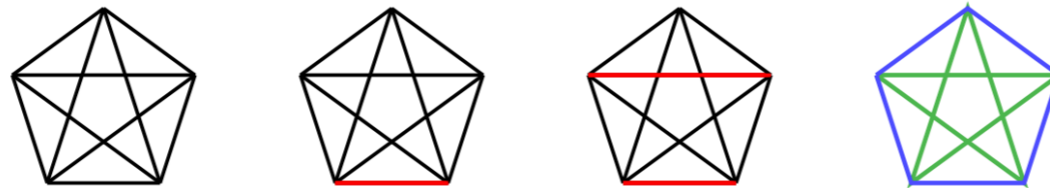
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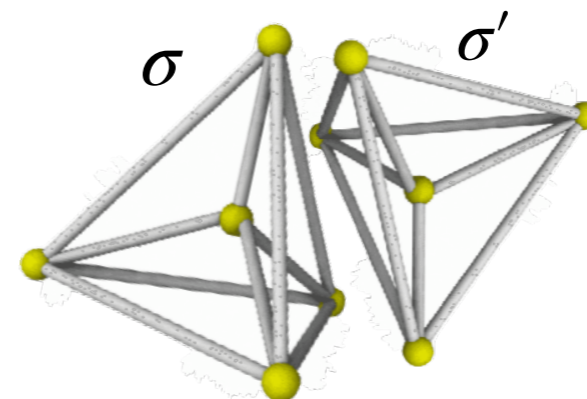
Diophantine equations for discrete areas

[Dittrich, Haggard, SKA '21]

Constraints between neighbouring simplices

$$\Phi_{e_i}^{\tau, \sigma}(a_t) = \Phi_{e_i}^{\tau, \sigma'}(a_t), \quad i = 1, 2$$

match pair of 3d dihedral angles



The Construction

Weak implementation of constraints

Forced between

[courtesy of: *Hal Haggard*]

Scylla: Reduce too much density of states

Charybdis: Impose dynamics that doesn't match GR



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Charybdis: Impose dynamics that doesn't match GR



(localized geometric constraints)

$$\mathcal{C}_i^\tau := \phi_{e_i}^\tau - \Phi_{e_i}^{\tau,\sigma}(a_t) = 0$$

(second-class constraints)

$$\{\mathcal{C}_i^\tau, \mathcal{C}_j^\tau\} = \gamma (9/2) \text{Vol}_\tau$$

[*Dittrich, Speziale, Ryan, Haggard,...*]
[*Kapovich-Millson*]

γ - an anomaly parameter

Impose constraints 'weakly': as strongly as allowed by uncertainty relation

[SF: *Engle-Perriera-Rovelli-Livine, Perez*]

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Use coherent states:

'Integrate out' ϕ^τ variables

coupling between σ and σ'

$$|K(\phi^\tau, \Phi_{e_i}^{\tau,\sigma})\rangle$$



$$G_\tau = \langle K_{\Phi_{e_i}^{\tau,\sigma}} | K_{\Phi_{e_i}^{\tau,\sigma'}} \rangle$$

[Livine, Speziale '17]

[Steinhaus, Simão, SKA '22]

ansatz $\sim \mathcal{N}_k \exp\left(-\frac{\mathcal{C}^2}{4\Sigma^2(j)}\right)$

The Construction

Effective spin foam model

Combine simple amplitude and impose constraints ‘weakly’

[Dittrich, Haggard, Padua-Argüelles, SKA]

Effective spin foam models are discrete geometrical path integrals for quantum gravity

$$Z_{\text{ESF}} = \sum_{\{a_t\}} \mu(a) \exp(i S_{\text{ARC}}(a)) \prod_{\tau} G_{\tau}^{\sigma, \sigma'}(a) \prod_{\sigma} \Theta_{\sigma}^{\text{tr}}(a)$$

simplex inequalities

$\mu(a)$ - measure on space of discrete areas

Oscillatory behaviour controlled by area Regge action

Weak imposition of constraints by Gaussian terms localized on tetrahedra

Spin foam amplitudes may be cast into a similar form

[Steinhaus, Simão, SKA '22]

(have to integrate normal vectors)

Features of the model

Regge calculus Gluing terms

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[Dittrich, Haggard, Padua-Argüelles, SKA]

◆ 'Effective' dynamics of quantum geometries

keep dynamic principles of LQG and spin foam models

◆ Computationally very efficient

fast numerical computations cf. to BF and EPRL models, restricted spin foam models

[Speziale, Dona, Sarno, Gozzini, Frisono; Han, Liu, Qu, Huang; Bahr, Steinhaus, Simão, SKA...]

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◆ Related to spin foam models for higher gauge group

$Z \propto \cos(S_{\text{Regge}})$ [Baez, Girelli, Pfeiffer, Popescu; Baratin, Freidel; Miković, Vojinović; Dittrich, Girelli, Riello, Tsimiklis, SKA]

◆ Control: can test many interesting features

[3D SFs: Simão ‘24, Jercher, Steinhaus, Simão]

Simple construction of Lorentzian model: allows spacelike and timelike area configurations

Cosmology applications*

Lorentzian geometries

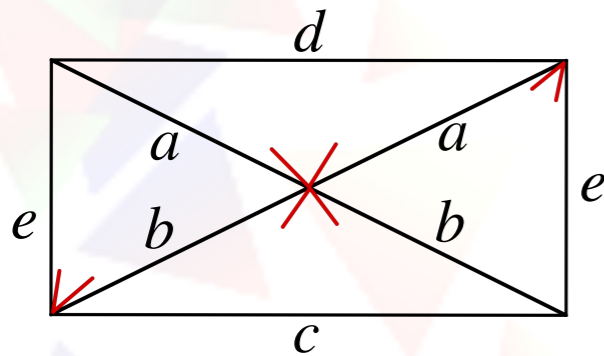
Configurations

Configurations can be grouped into two sets: Regular and Irregular
according to light cone structure of faces

2D Examples (Hinge causality)

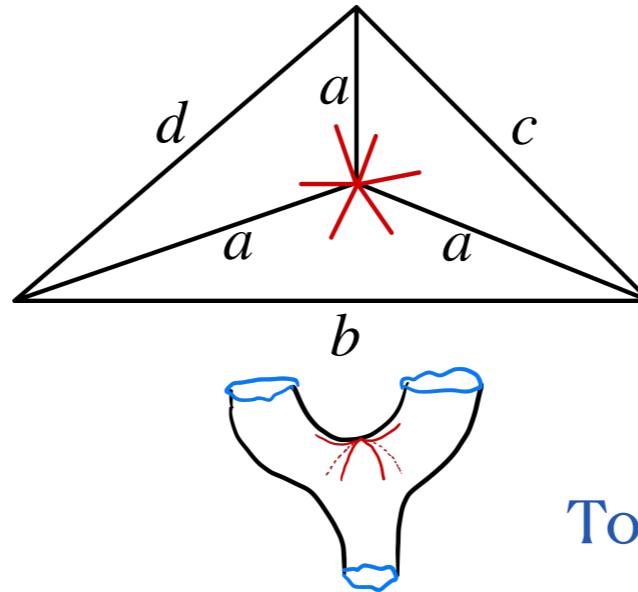
[Louko-Sorkin '95, Sorkin '19]

Regular configuration

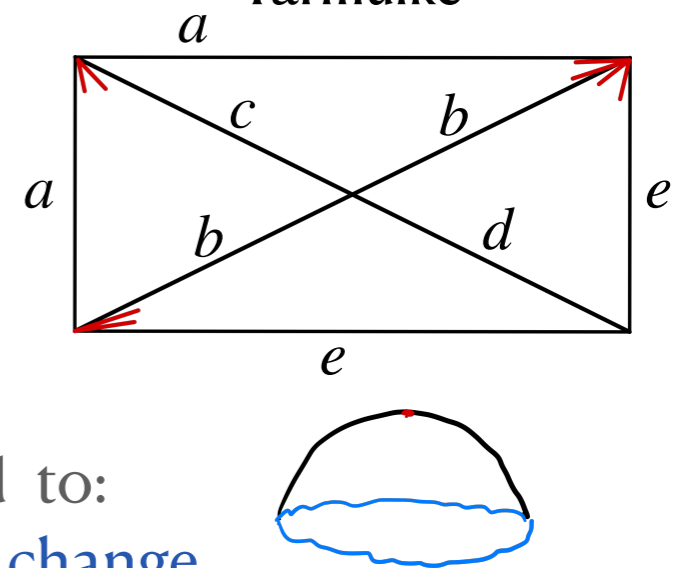


Irregular configurations

Trouser-like



Yarmulke



related to:
Topology change

[Sorkin'19, Dittrich, Padua-Argüelles, SKA '21]

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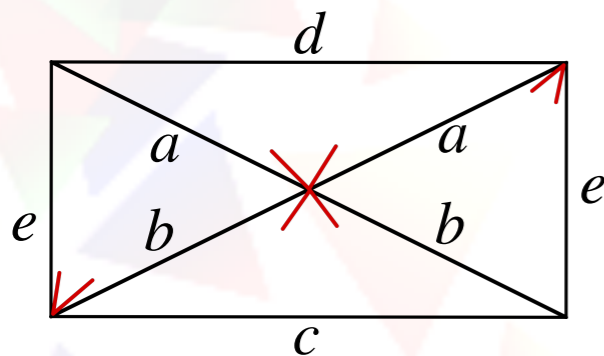
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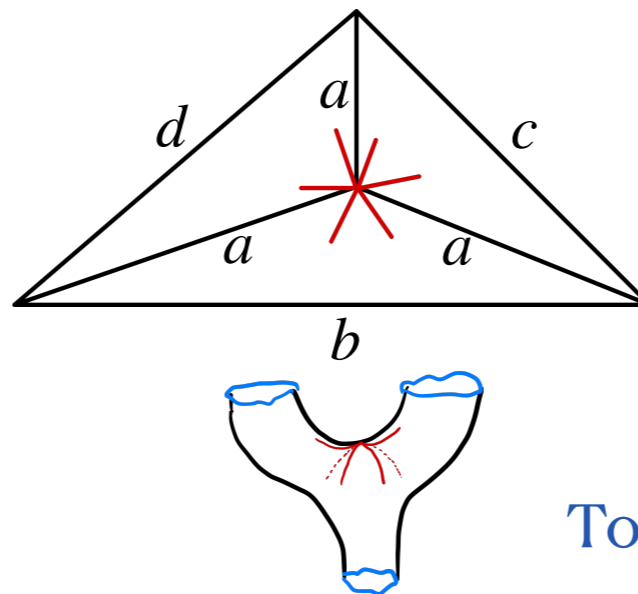
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Regular configuration

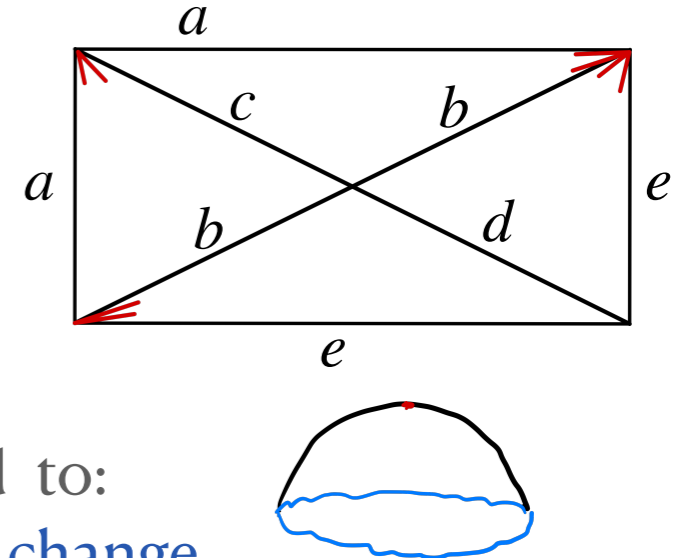


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Higher Dimensions: Other causality conditions **Edge causality, Vertex Causality**

[Jordan, Loll '13]

[Borgolte, SKA wip]

Irregular configurations leads to complex valued Regge action [Sorkin'19, Dittrich, Padua-Argüelles, SKA '21]

Convergence

Techniques to deal with complex amplitudes

Methods for treating complex amplitudes: Applications to Lorentzian path integrals

Picard-Lefschetz, Holomorphic-gradient flow [Han, Wan, Huang, Liu, Qu: Jia: Dittrich, Padua-Argüelles, SKA '22]

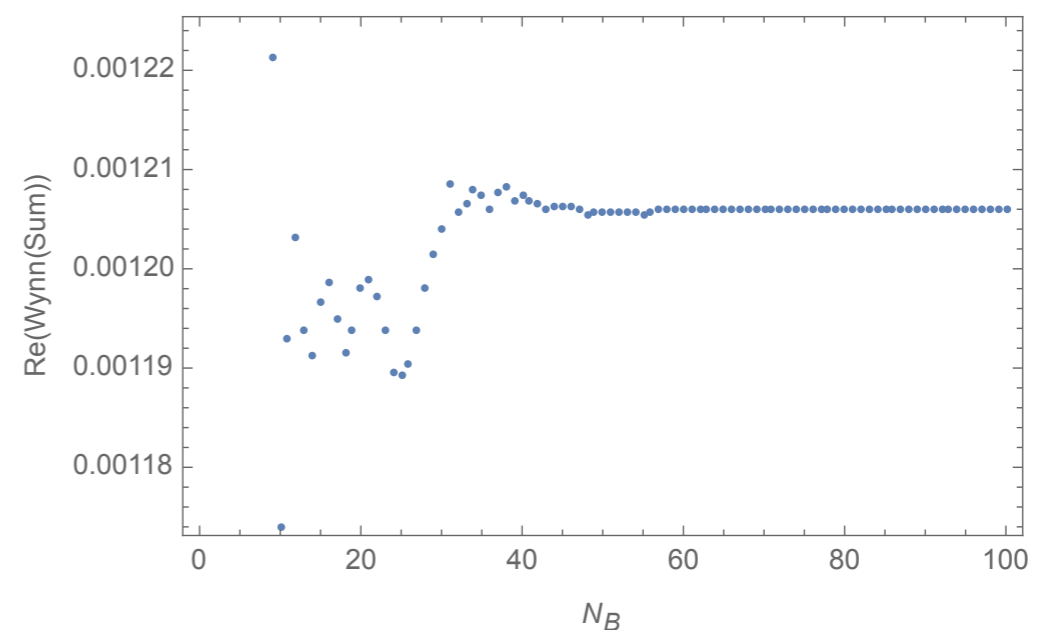
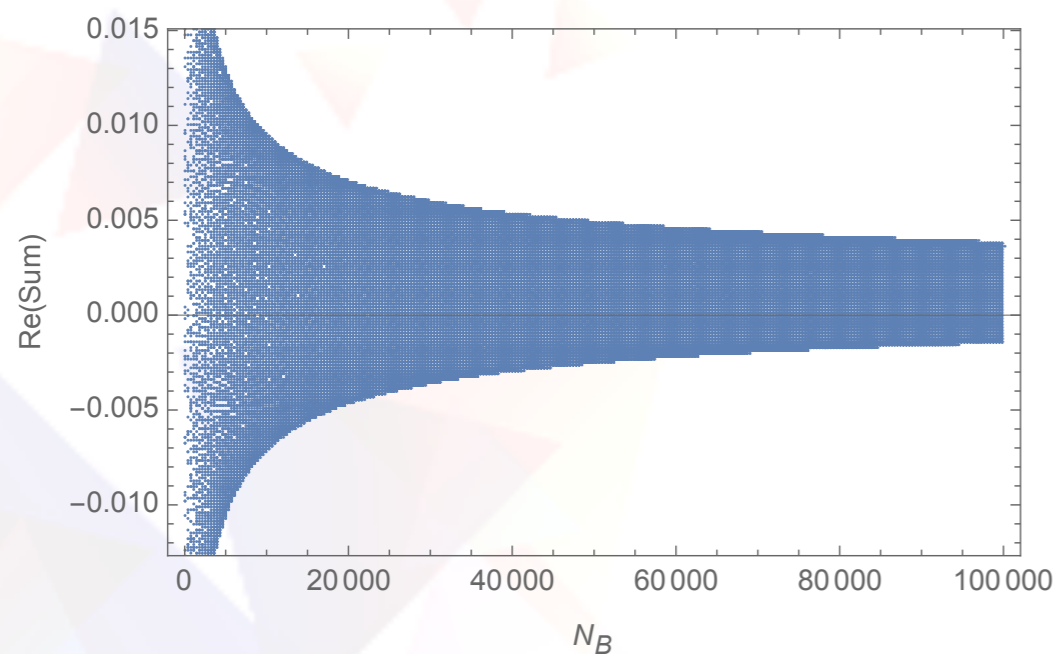
Tensor network techniques [Bahr, Steinhaus, Dittrich, Cunningham, Mizera, Kaminski, Martin-Benito]

Quantum simulations, machine learning techniques

Monte Carlo techniques [Dona, Frisoni '23; Steinhaus '24]

Acceleration operators [EPRL: Speziale, Dona, Sarno, Gozzini, Frisoni] [ESF: Dittrich, Padua-Argüelles '23]

Shanks transform, Wynn epsilon algorithm, Aitken's delta process



[plots by: Dittrich, Padua-Argüelles]



Effective Spin foams

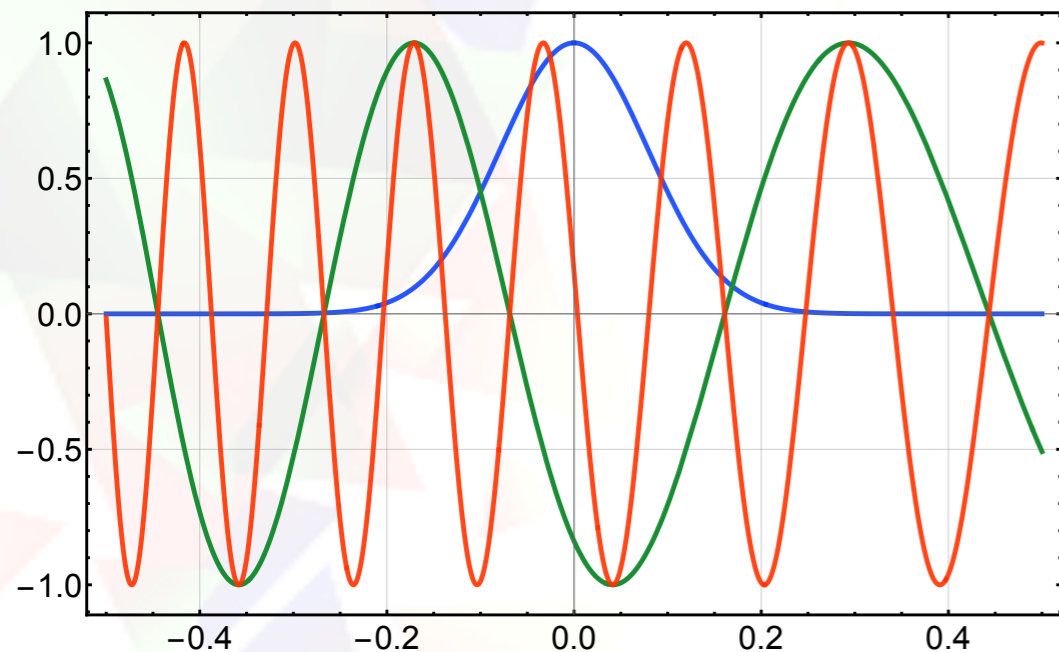
Testing the models

[SKA, Dittrich, Haggard, Padua-Argüelles]
arXiv: [2004.07013](#), [2011.14468](#), [2104.00485](#)

Discrete gravity dynamics

Semi-classical

Weakly imposed constraints



$$Z_{\text{ESF}} = \sum_{\{a_t\}} \mu(a) \exp(i S_{\text{ARC}}(a)) \prod_{\tau} G_{\tau}^{\sigma, \sigma'}(a) \prod_{\sigma} \Theta_{\sigma}^{\text{tr}}(a)$$

Oscillations

Gaussians
peaked on constraints

Semi-classical limit:

Few oscillations over Gaussian needed

$$\gamma \sqrt{a_t} \text{curv}_t \lesssim \mathcal{O}(1)$$

[SKA, Dittrich, Haggard]

[SF: Han 13]

[Dittrich, Haggard, SKA '20] [SF: Han, Huang, Liu, Qu '21]

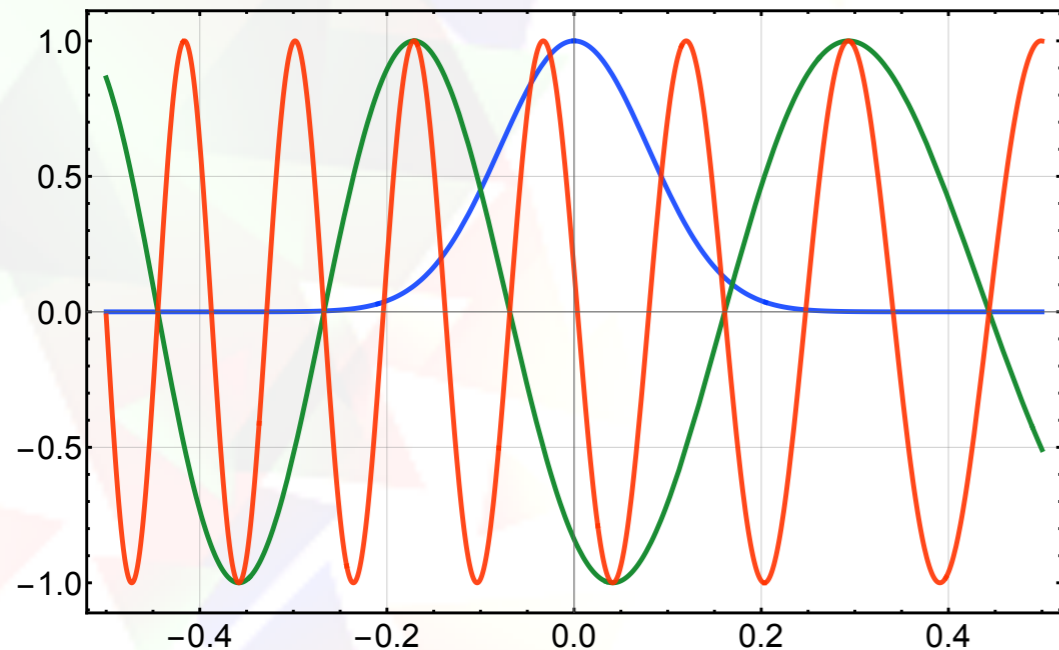
Alternative idea: 'true' critical points are complex

Imaginary part of saddle point controlled by γ has to be small

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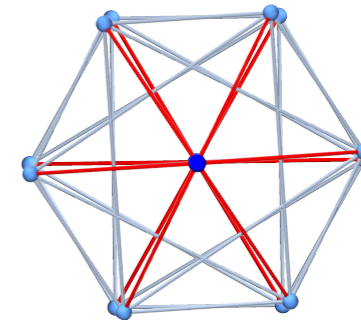
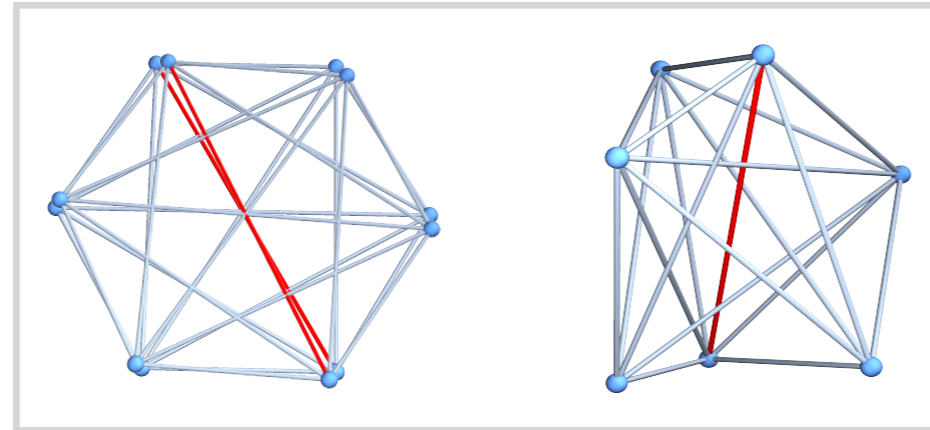
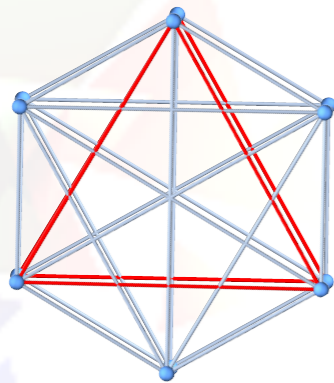
What are some effects of **discrete area spectrum** and **'weakly' imposed constraints**?

Testing ESF model

Numerical Tests

- ◆ Early results from explicit evaluations **Non-perturbative tests**

Several examples of discrete geometries with curvature



explicit sum over discrete areas, implement weakened constraints

Compute expectation values of bulk variables

testing discrete EOMs

- ▶ Recover discrete gravity dynamics for certain range of parameters

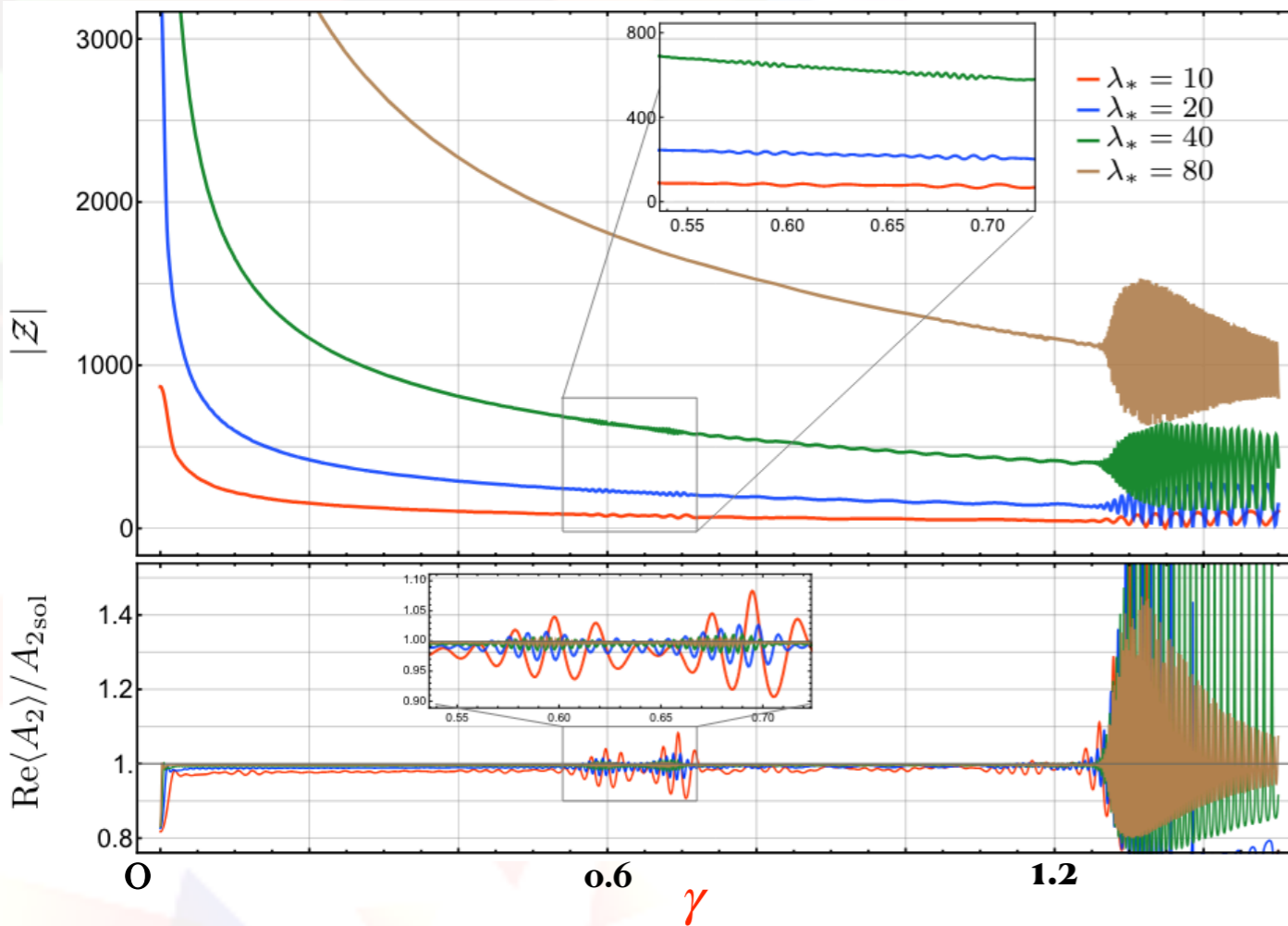
interesting effects beyond saddle point evaluation

larger acceptable range for γ than expected

Numerical results

Bulk-Edge

Small curvature



Oscillations due to interplay between discrete areas and constraint imposition

Nice surprises:

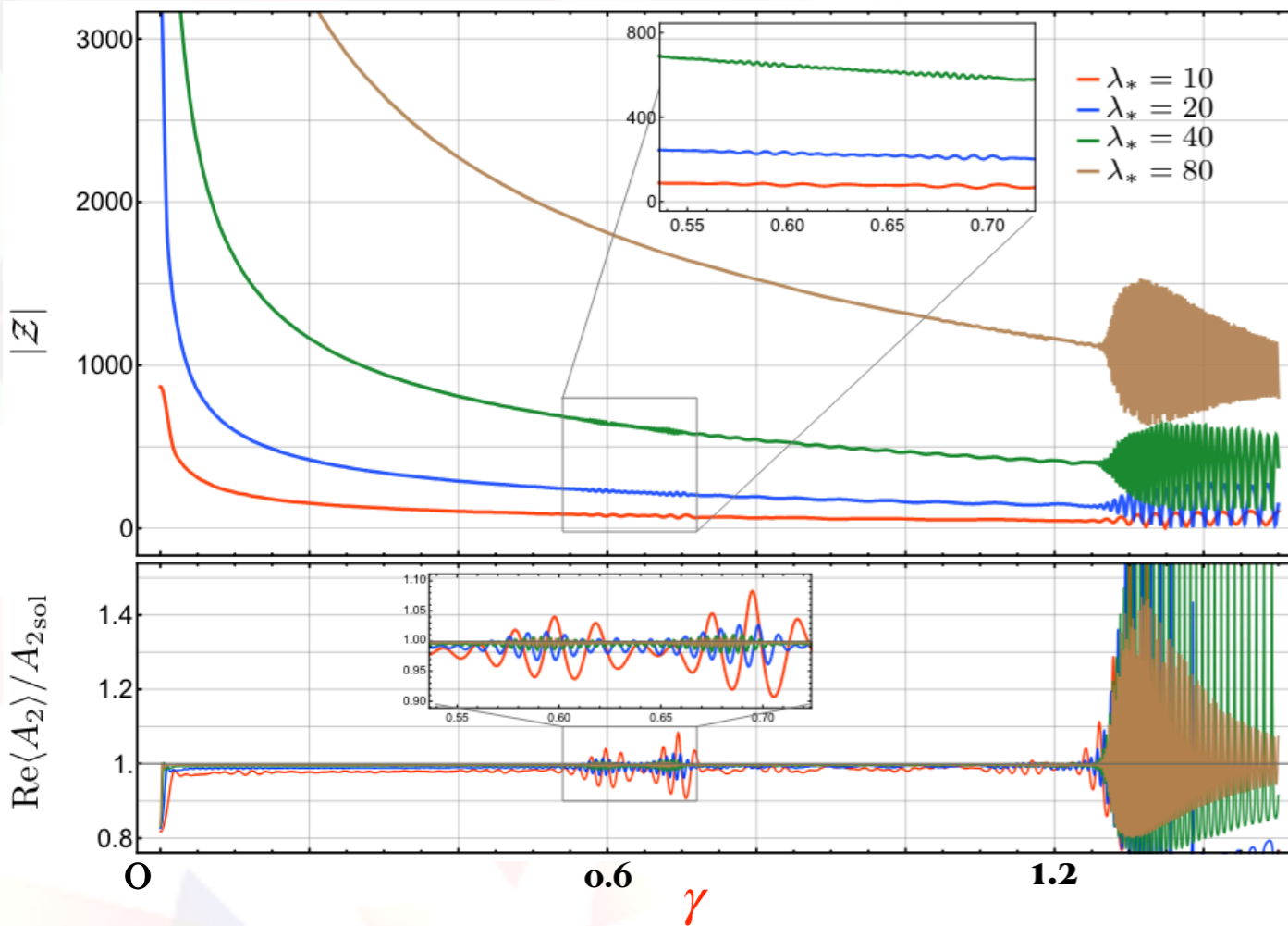
Large acceptable range for γ for the expectation values

Better matching of classical EOMs for large boundary scales

Numerical results

Bulk-Edge

Small curvature



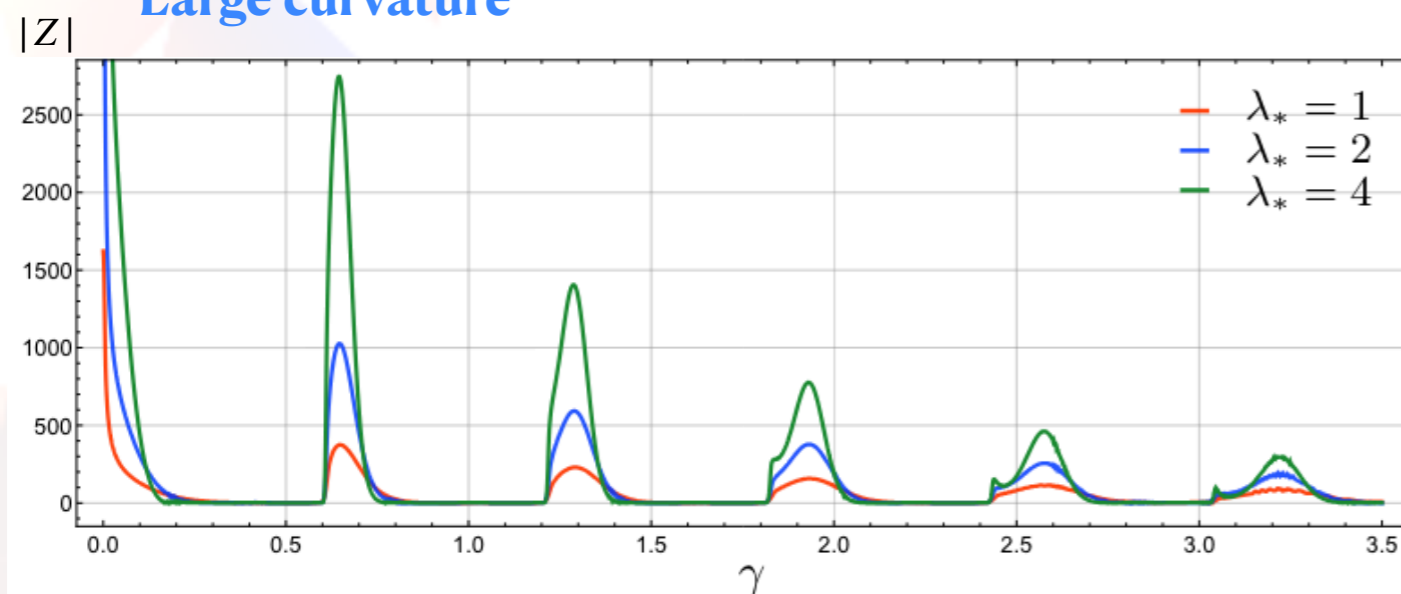
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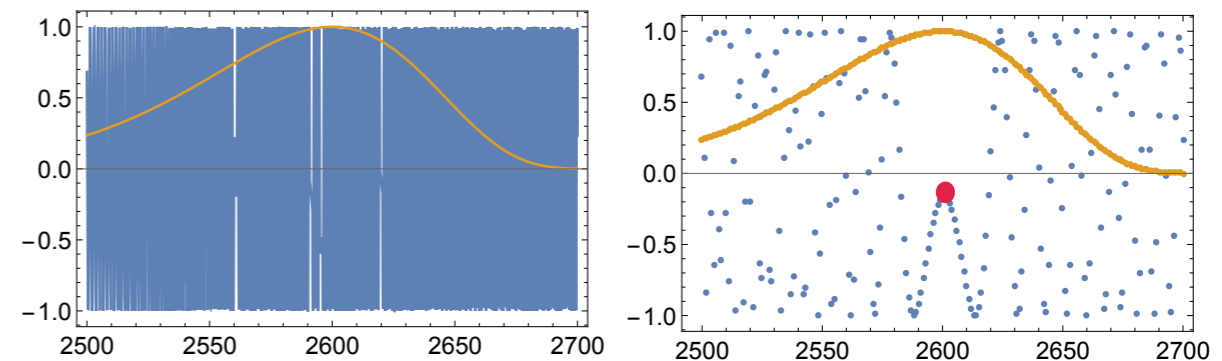
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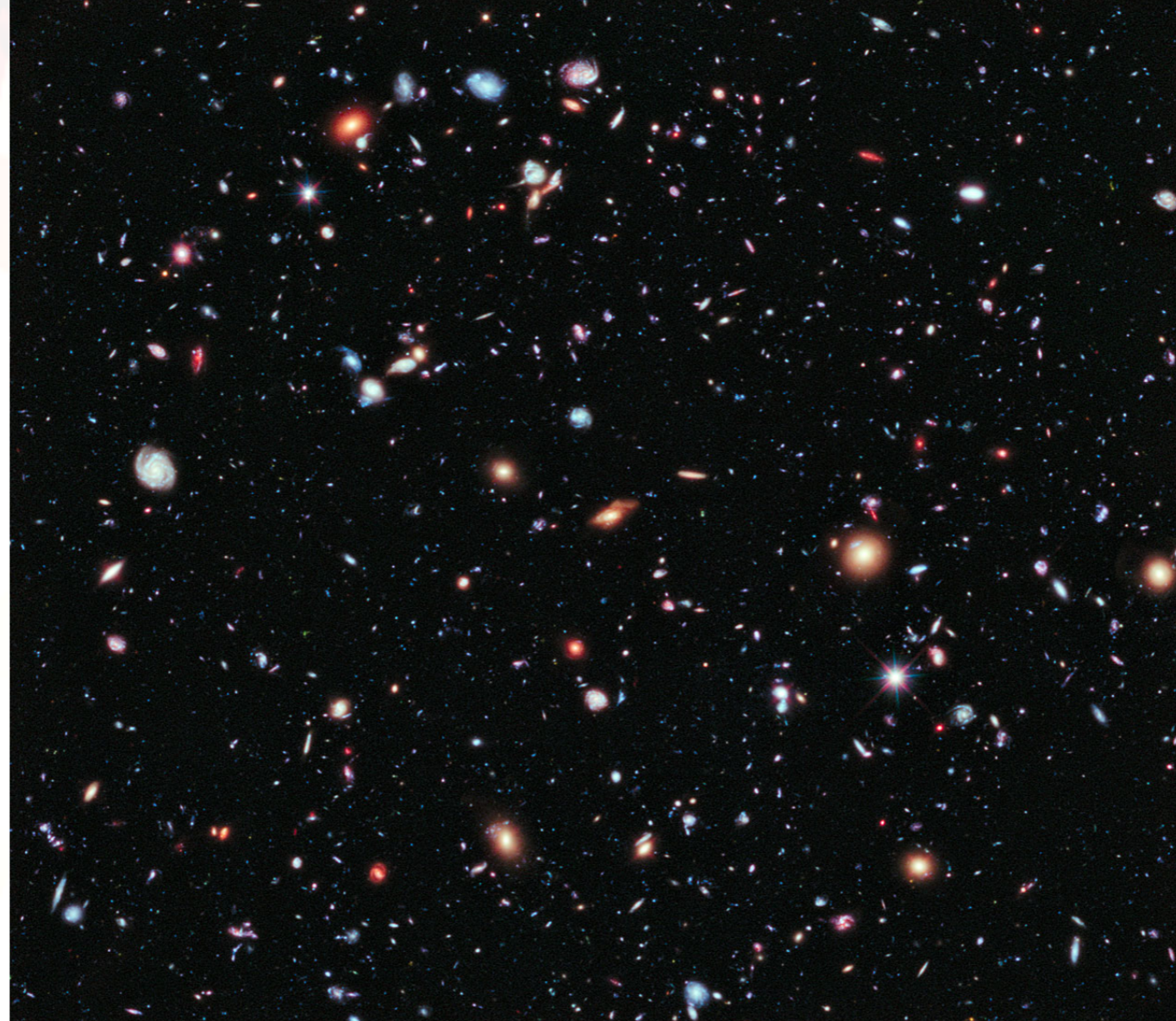
Better matching of classical EOMs for large boundary scales

Large curvature



Peaks due to discretization effects: pseudo-saddle points





Credit: NASA, ESA, HUDF09 Team

Effective Spin foams

Cosmology applications

[SKA, Dittrich, Padua-Argüelles, Gielen, Schander, Steinhaus, Jercher]
arXiv: [2109.00875](#), [2112.15387](#), [2306.06012](#), [2312.11639](#)

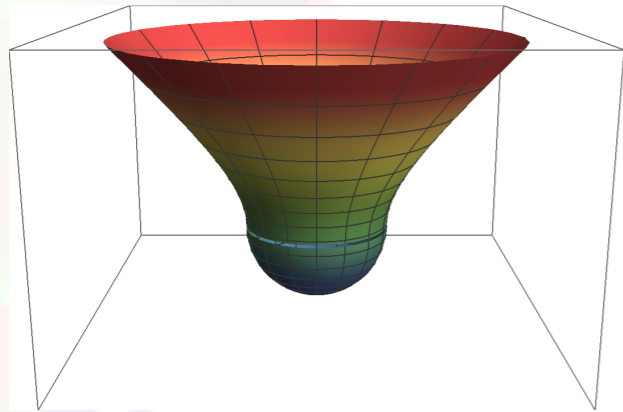
Cosmology Applications

Mini-superspace models

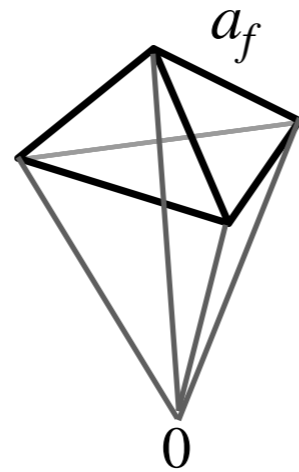
[Hartle, Hawking, Feldbrugge, Lehnert, Turok, Dorronsoro, Halliwell, Hertog, Jansen,....]

[Williams, Lui, Collins, Dittrich, Gielen, Schander, Vidotto, Gozzini, Frisoni,....]

No boundary proposal

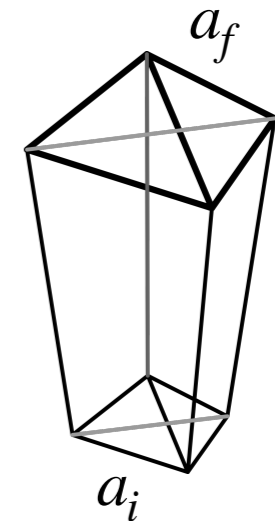
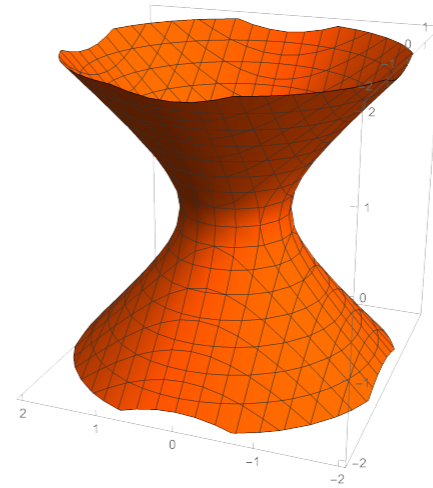


[Dittrich, Gielen, Schander '21]



partial discretization

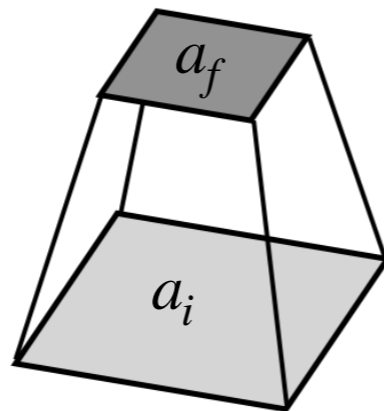
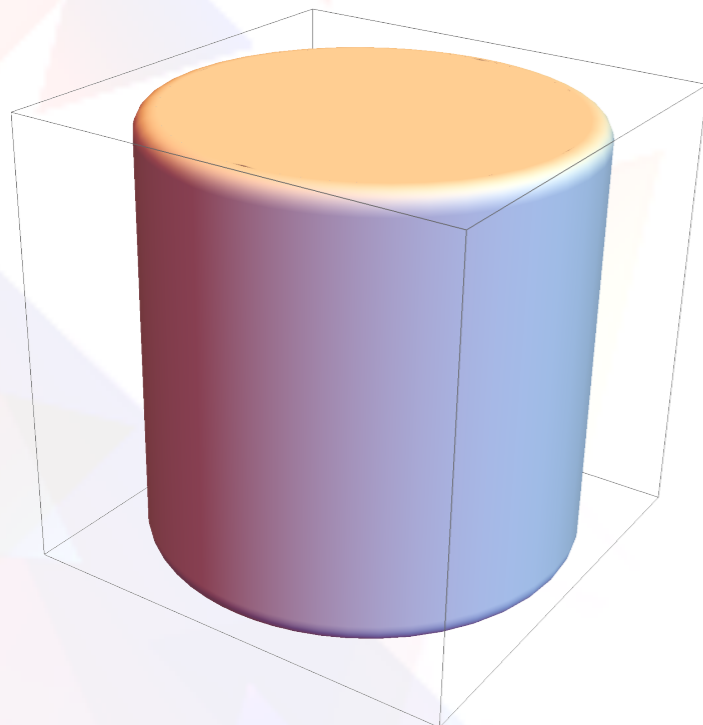
deSitter cosmological spacetime



partial discretization

[Dittrich, Padua-Argüelles '23]

Spatially flat cosmology [Steinhaus, Jercher]



Alexander Jercher's talk!

Spin foam cosmology

[Han, Liu, Qu, Vidotto, Zhang]

Cosmology Applications

[Dittrich, Padua-Argüelles '23]

Effective spin foam model

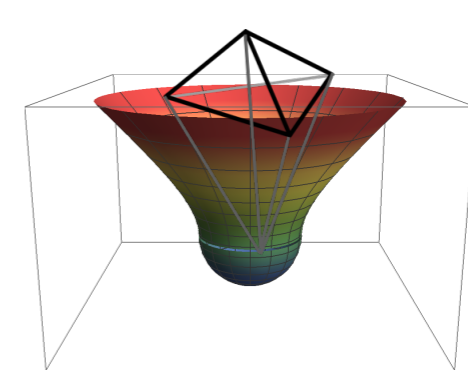
Results from deSitter cosmology

Deal with slowing converging and diverging sums

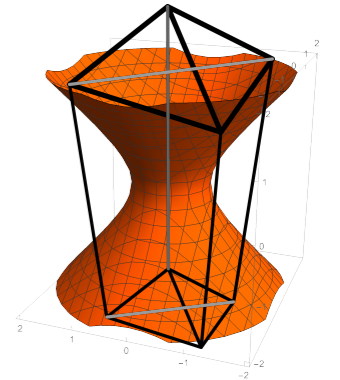
Unbounded sum over lapse variable (**regular** and **irregular** configurations)

Acceleration techniques (Wynn's algorithm): Speeds up convergence of discrete sums

- works well for actions linear in summation variable (also for spin foams)



Ball model

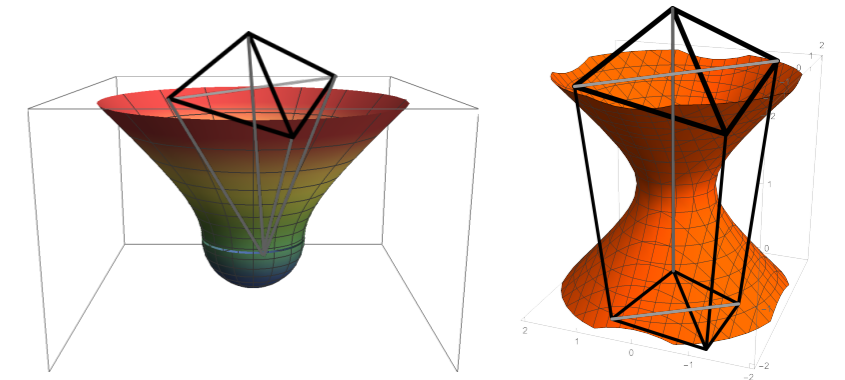


Shell model

Cosmology Applications

[Dittrich, Padua-Argüelles '23]

Effective spin foam model



Ball model

Shell model

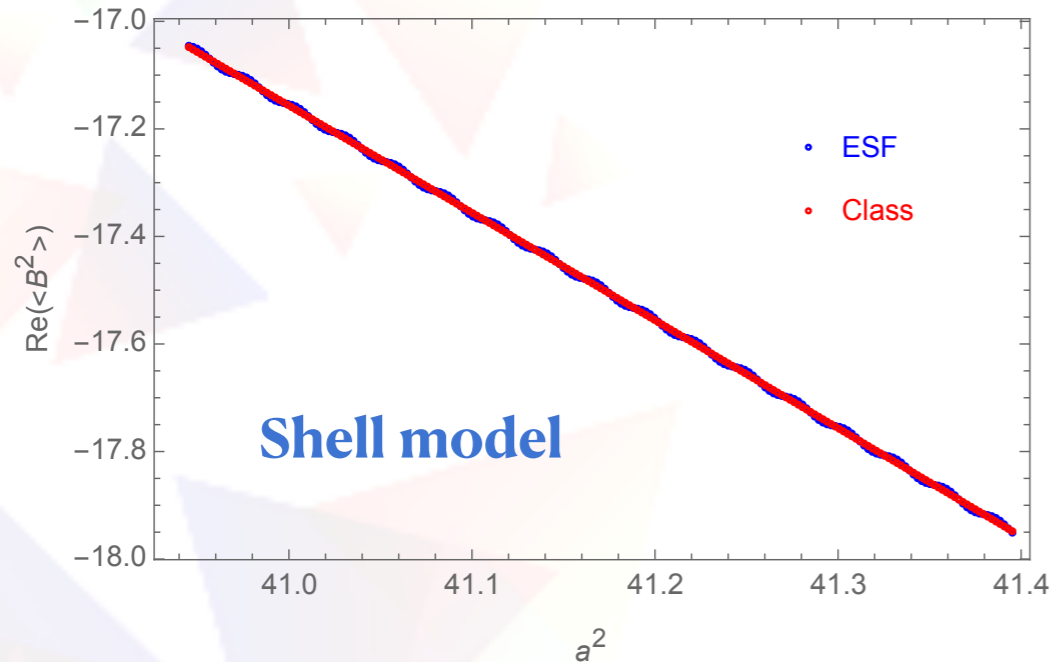
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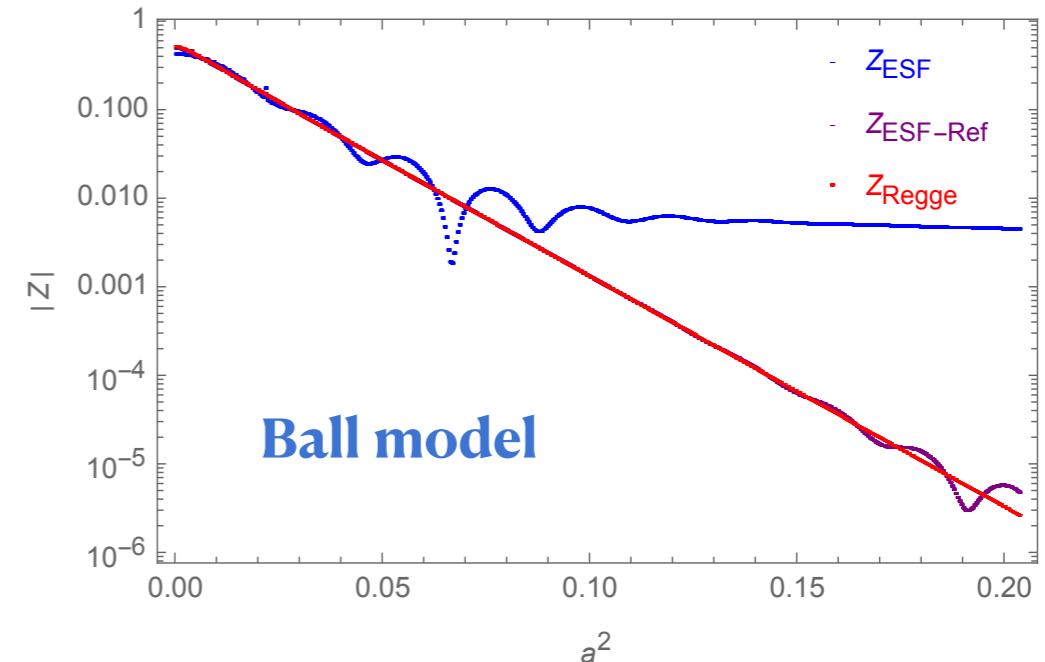
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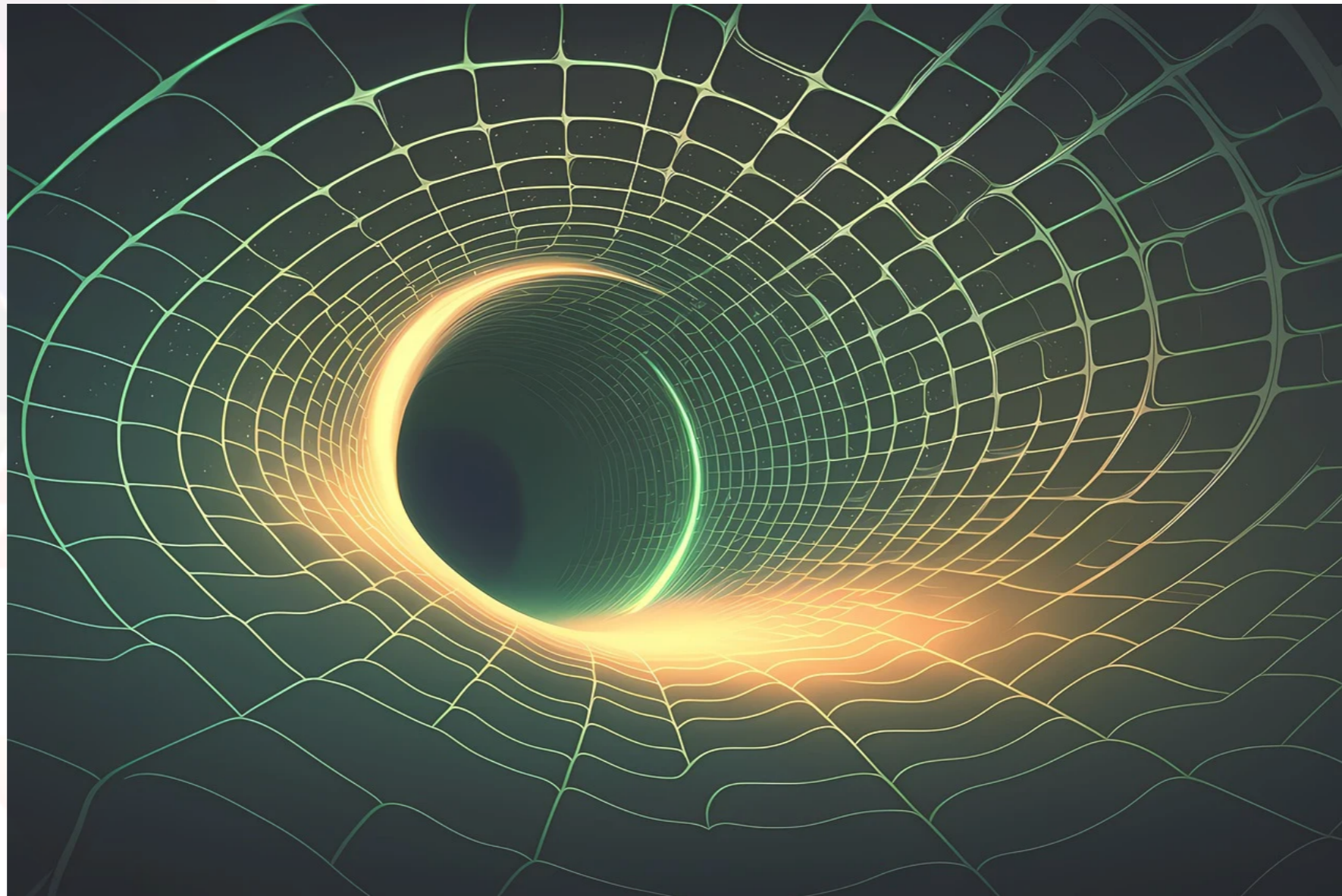


Shell model



Ball model

Partition function and expectation values sensitive to discrete area spectrum



Credit: Fzalai

Effective Spin foams

Continuum limit

[Dittrich, Kogios, Borissova, Krasnov, Steinhaus, SKA]

arXiv: 2105.10808, 2203.02409, 2207.03307, 2211.09578, 2312.13935

Continuum limit

First Steps

[Dittrich '21, Dittrich-Kogios '22]

Action:
$$S = S_{\text{ARC}}(a) - i \sum_{\tau} \ln G_{\tau}(a)$$

[Handbook: *Dittrich, Steinhaus, SKA*]

First attempts: Linearize area Regge calculus around a flat background on hyper cubic lattice(s)

Interesting Results

[Dittrich '21, Dittrich-Kogios '22]

Linearized continuum limit

Area Regge Calculus ~ GR + Weyl²

leading order in lattice derivatives

$$\partial_k^2$$

sub-leading

$$\partial_k^4$$

- Also holds starting from effective spin foams action

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Related work (Area metric formulations)

♦ Modified Plebanski actions [Borissova, Dittrich '22]

♦ Continuum area-metric actions [Borissova, Dittrich, Krasnov '23, Borissova, Ho '24]

Mechanism Compute Hessian around flat background

Quantum Regge Calculus *[Williams, Roček '81]*

Hessian
$$S^{(2)} = \sum_t \frac{\partial a_t}{\partial l_e} \frac{\partial \epsilon_t}{\partial l_{e'}} \delta l_e \delta l_{e'}$$

Quantum: Area Regge Calculus

$$S^{(2)} = \frac{\partial \epsilon_t}{\partial a_{t'}} \delta a_t \delta a_{t'} \quad \text{[Dittrich, Haggard, SKA '18]}$$

Strategy: Organize variables according to order of lattice derivatives in Hessian

Leading order contributions results from variables forming an **area metric**

$$\begin{array}{l} 20 \text{ area metric variables per point} \\ = \quad h \quad (10) \quad + \quad \chi \quad (10) \\ \text{trace part} \quad \quad \quad \text{trace-free part} \end{array}$$

Integrate Integrate out non-metric variables \longrightarrow Effective action for metric variables

$$S_{\text{eff}} = h \cdot (H_{hh} - H_{h\chi} H_{\chi\chi}^{-1} H_{\chi h}) \cdot h$$

Summary & outlook

What can we learn from (discrete) gravitational path integrals?

Using 'effective spin foam models'

Quantum geometry

Quantum gravity

Semi-classical approximations

Topological transitions

Black hole dynamics

Early universe Cosmology applications

Quantum field theory in curved space time

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THANK YOU!

