

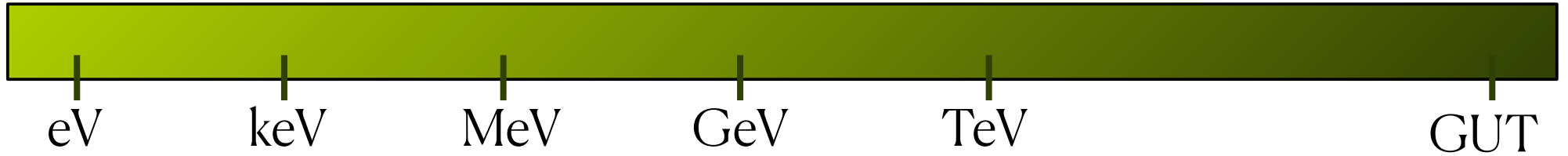
# Short-baseline neutrino anomalies

Pilar Coloma – Instituto de Física Teórica UAM-CSIC

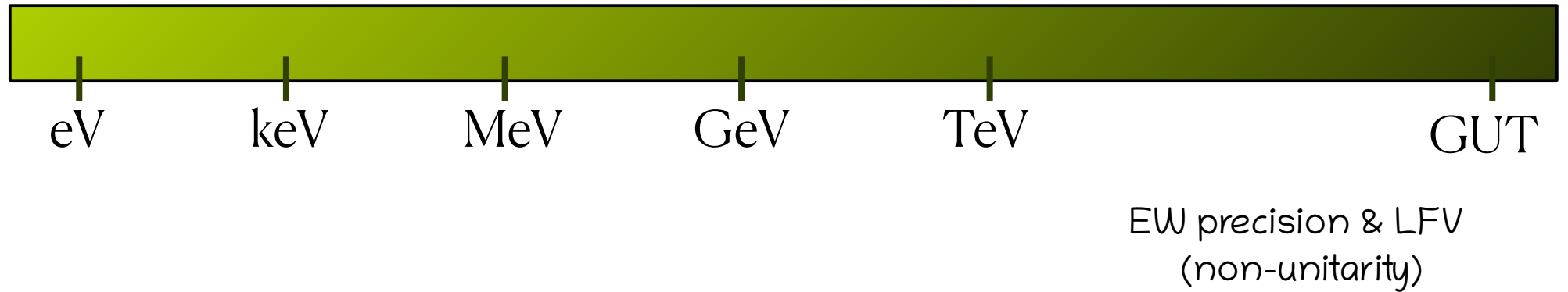
ALPs'23 Workshop, Obergurgl (March 30<sup>th</sup>, 2023)



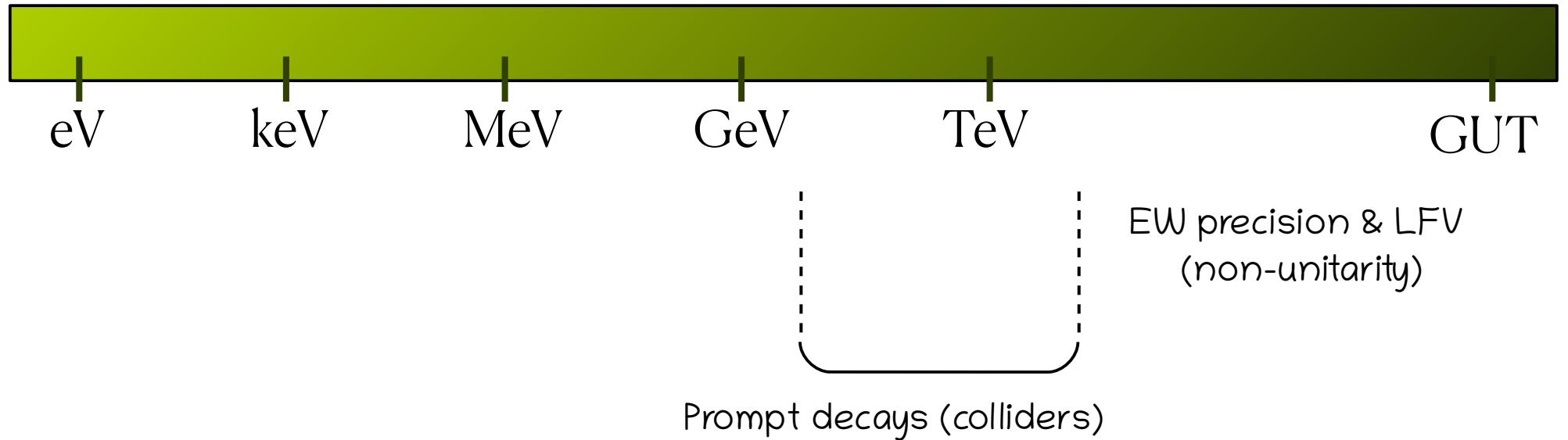
# Sterile neutrino searches



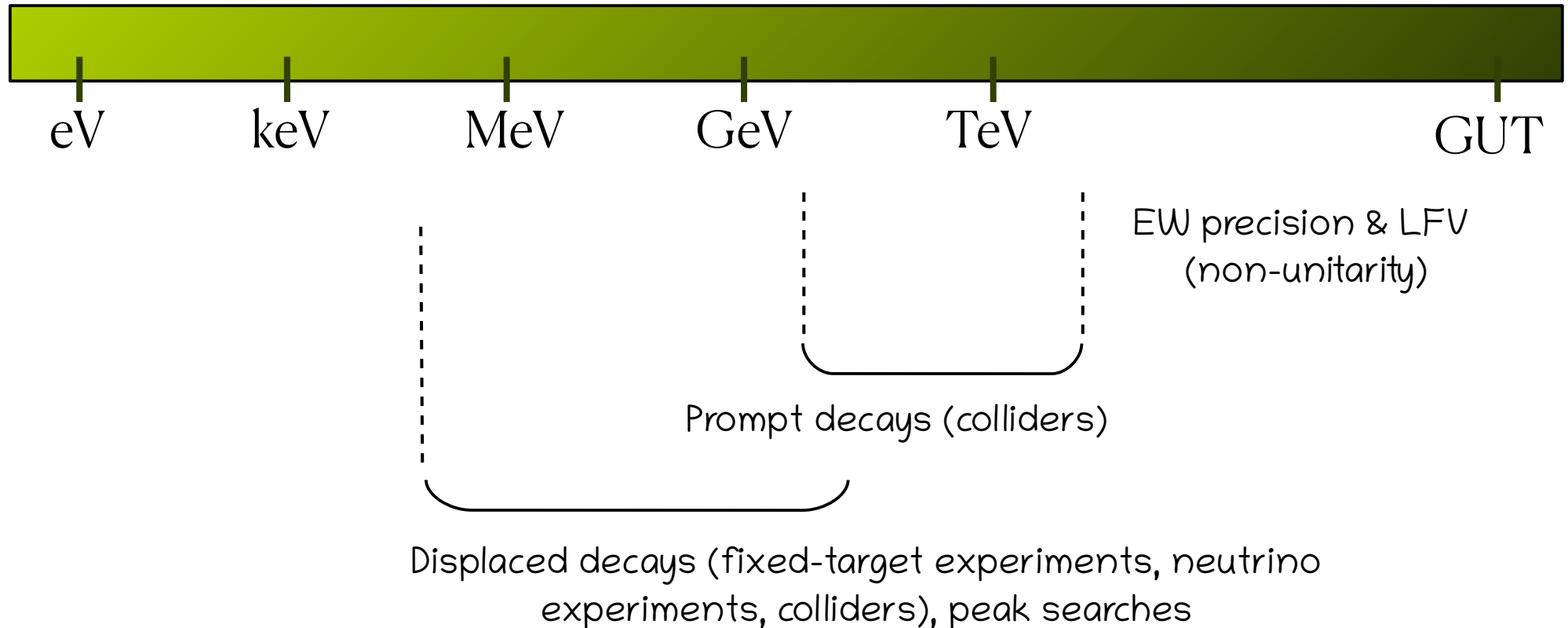
# Sterile neutrino searches



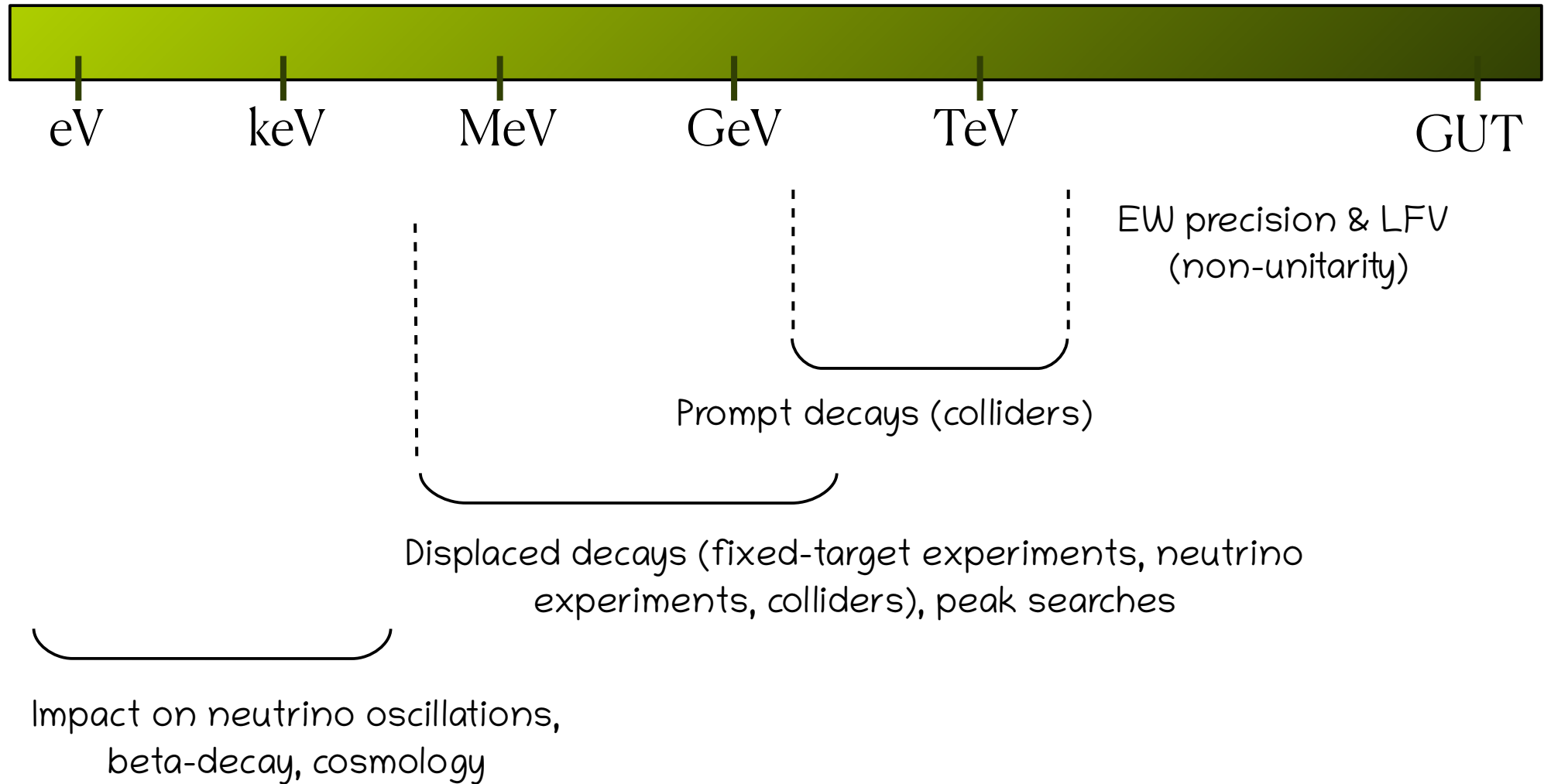
# Sterile neutrino searches



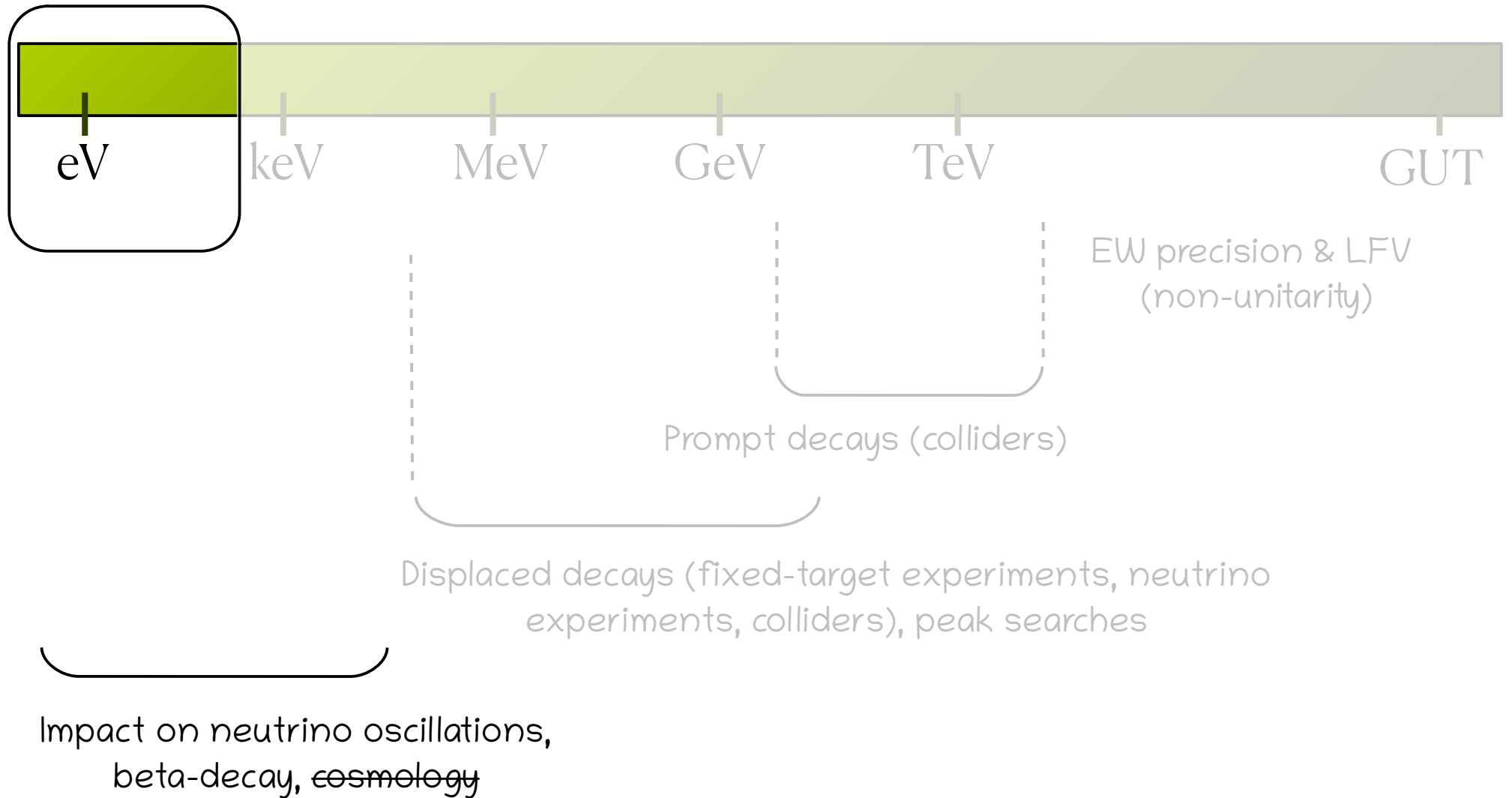
# Sterile neutrino searches



# Sterile neutrino searches



# Sterile neutrino searches



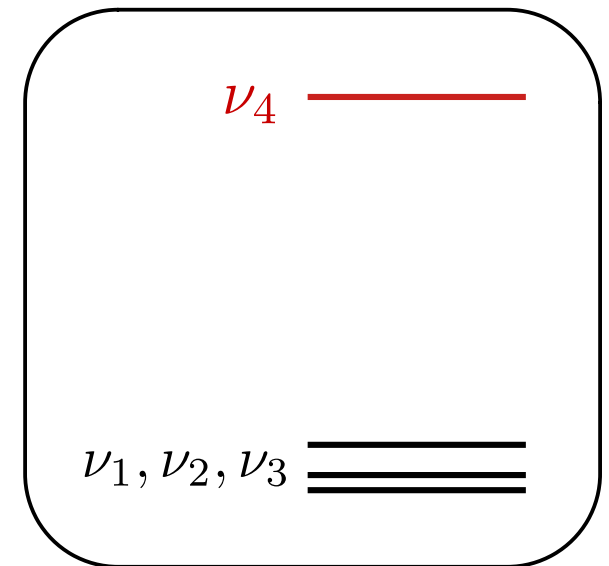
# eV-scale sterile neutrinos

$$i \frac{d}{dt} \Psi_\nu = (U H_0 U^\dagger + V) \Psi_\nu$$

$$U = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} & U_{e4} \\ U_{\mu1} & U_{\mu2} & U_{\mu3} & U_{\mu4} \\ U_{\tau1} & U_{\tau2} & U_{\tau3} & U_{\tau4} \\ U_{s1} & U_{s2} & U_{s3} & U_{s4} \end{pmatrix}$$

Oscillations depend on

$$\frac{\Delta m^2 L}{4E} \simeq 1.27 \frac{\Delta m^2 (\text{eV}) L(\text{m})}{E(\text{MeV})}$$





## $\nu_e$ and $\bar{\nu}_e$ appearance ( $P_{\mu e}$ )

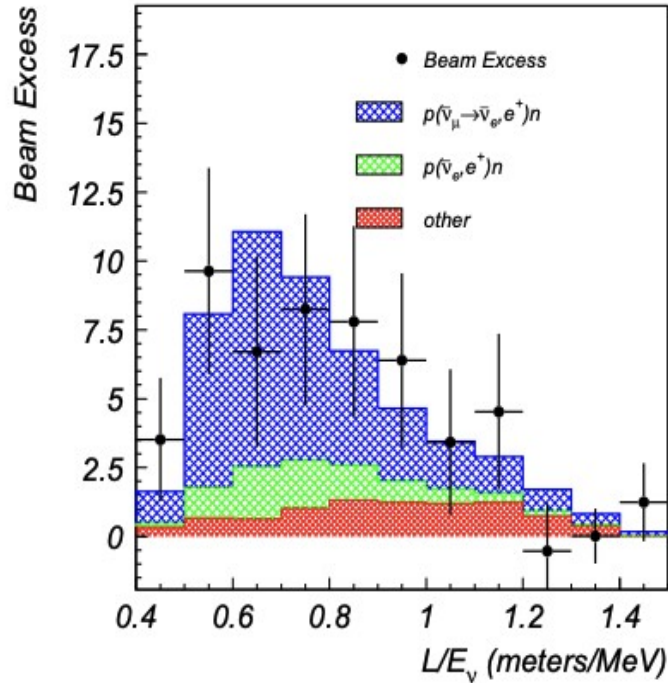
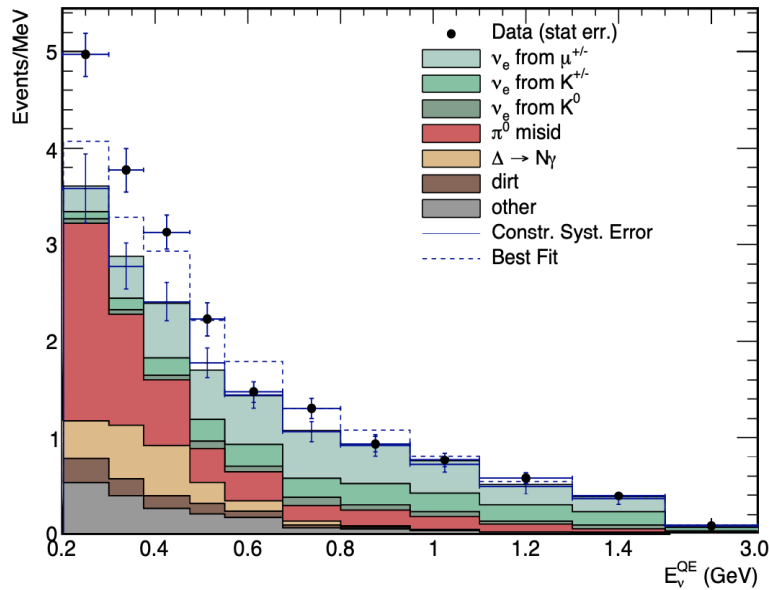
$$P_{\mu e} = \sin^2 2\theta_{\mu e} \sin^2 \left( \frac{\Delta m^2 L}{4E} \right)$$



$$4|U_{e4}|^2|U_{\mu4}|^2$$

LSND  
MiniBooNE  
MicroBooNE

# eV-scale sterile neutrinos

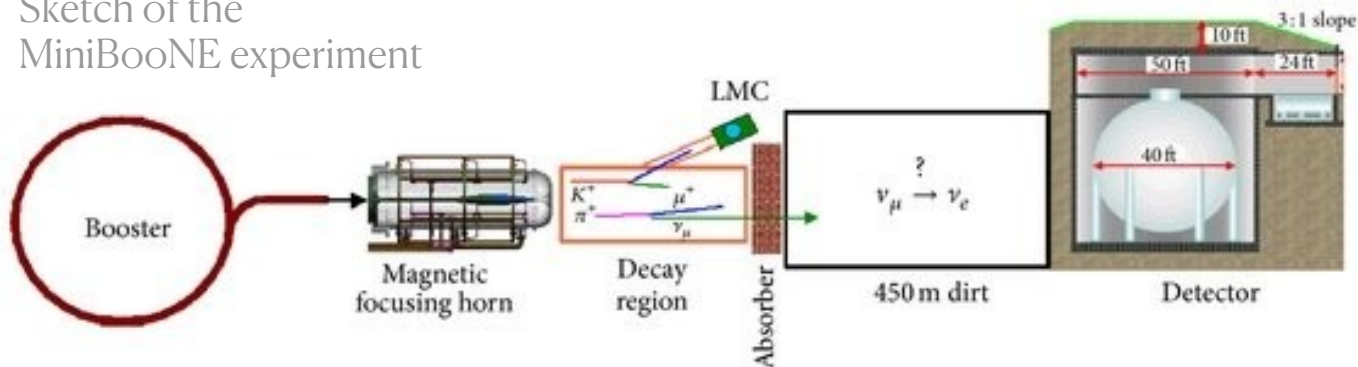


MiniBooNE:  $\sim 4.8\sigma$

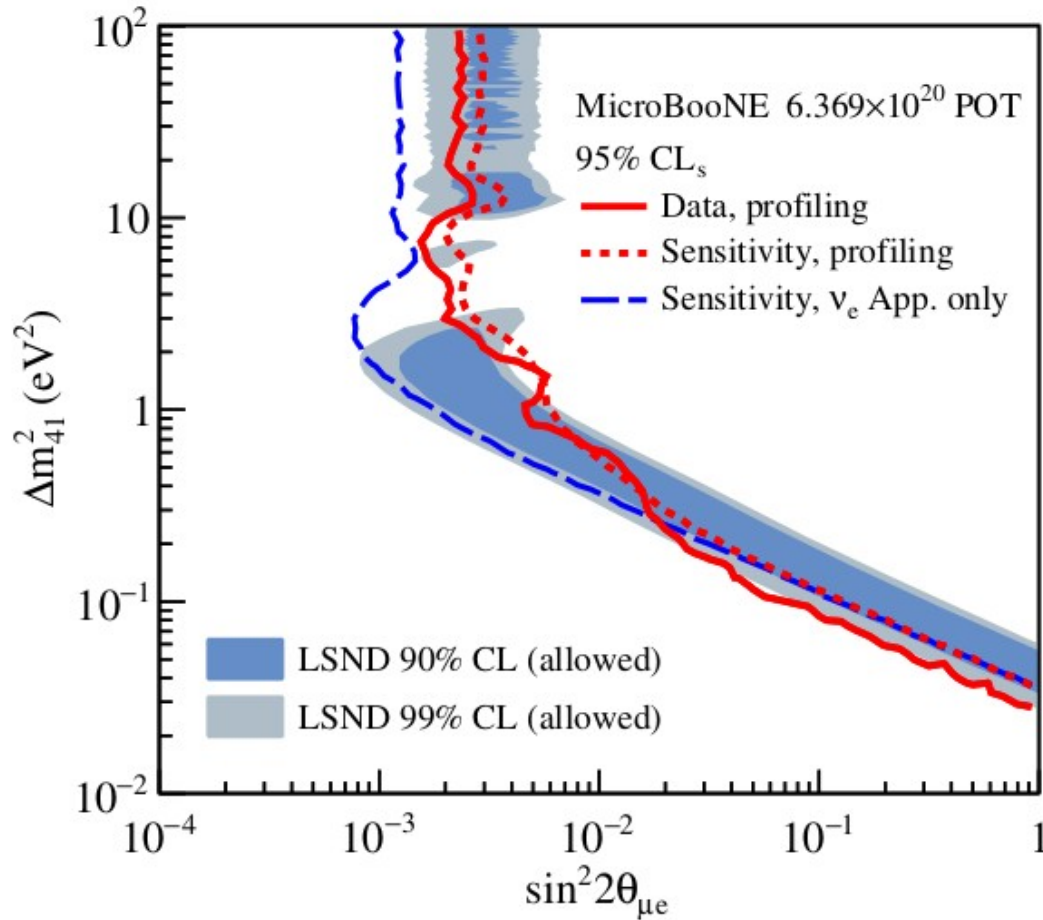
LSND:  $\sim 3.8\sigma$

LSND coll.,  
 hep-ex/0104049  
 MiniBooNE  
 coll.,1805.12028

Sketch of the  
 MiniBooNE experiment

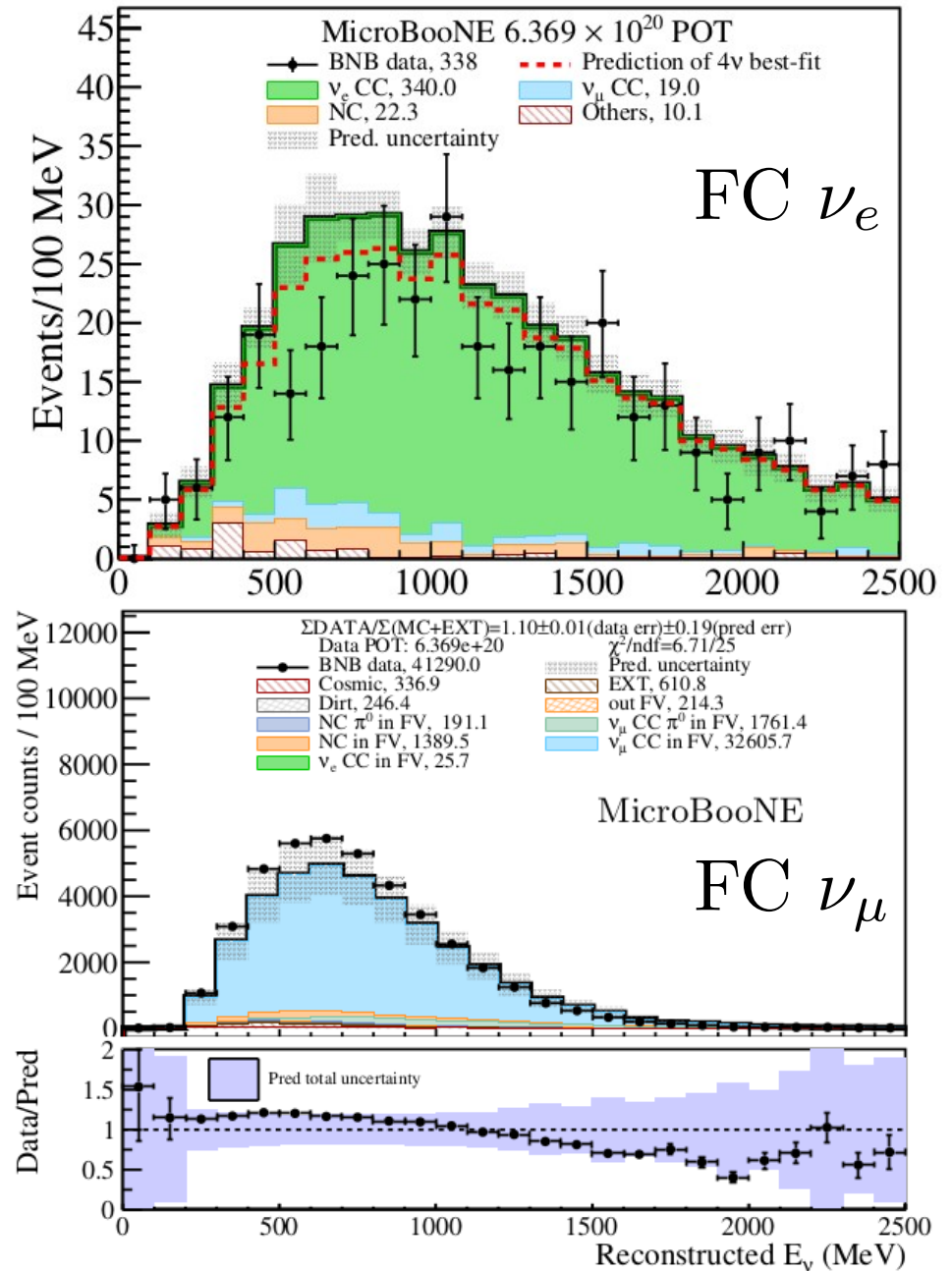


# MicroBooNE




MicroBooNE coll., 2210.10216

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## $\nu_e$ and $\bar{\nu}_e$ disappearance ( $P_{ee}$ )

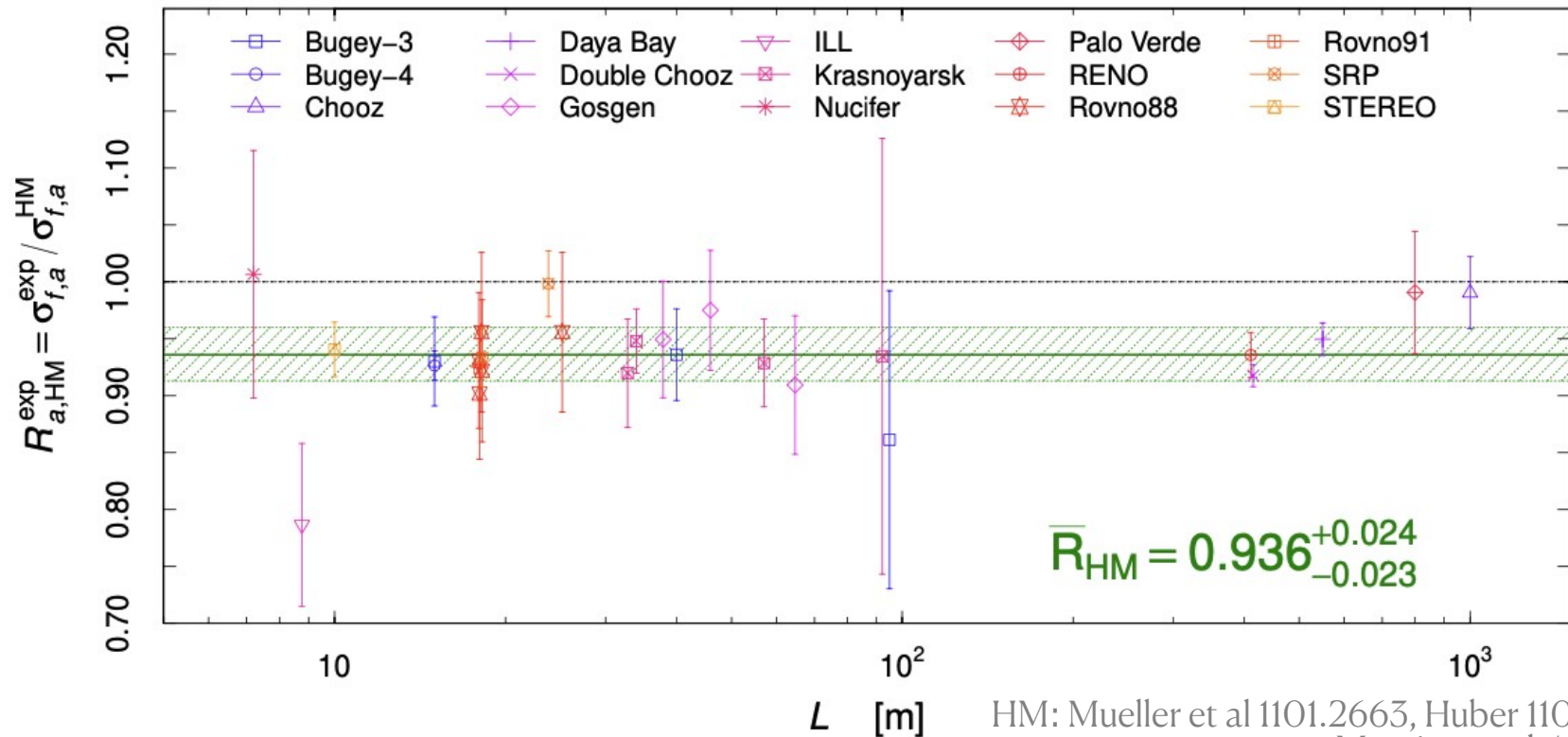
$$P_{ee} \equiv P(\nu_e \rightarrow \nu_e) = 1 - \sin^2 2\theta_{ee} \sin^2 \left( \frac{\Delta m^2 L}{4E} \right)$$


$$4|U_{e4}|^2(1 - |U_{e4}|^2)$$

Reactors  
Gallium experiments  
Solar neutrinos

# Reactor antineutrino anomaly

Situation in 2011:

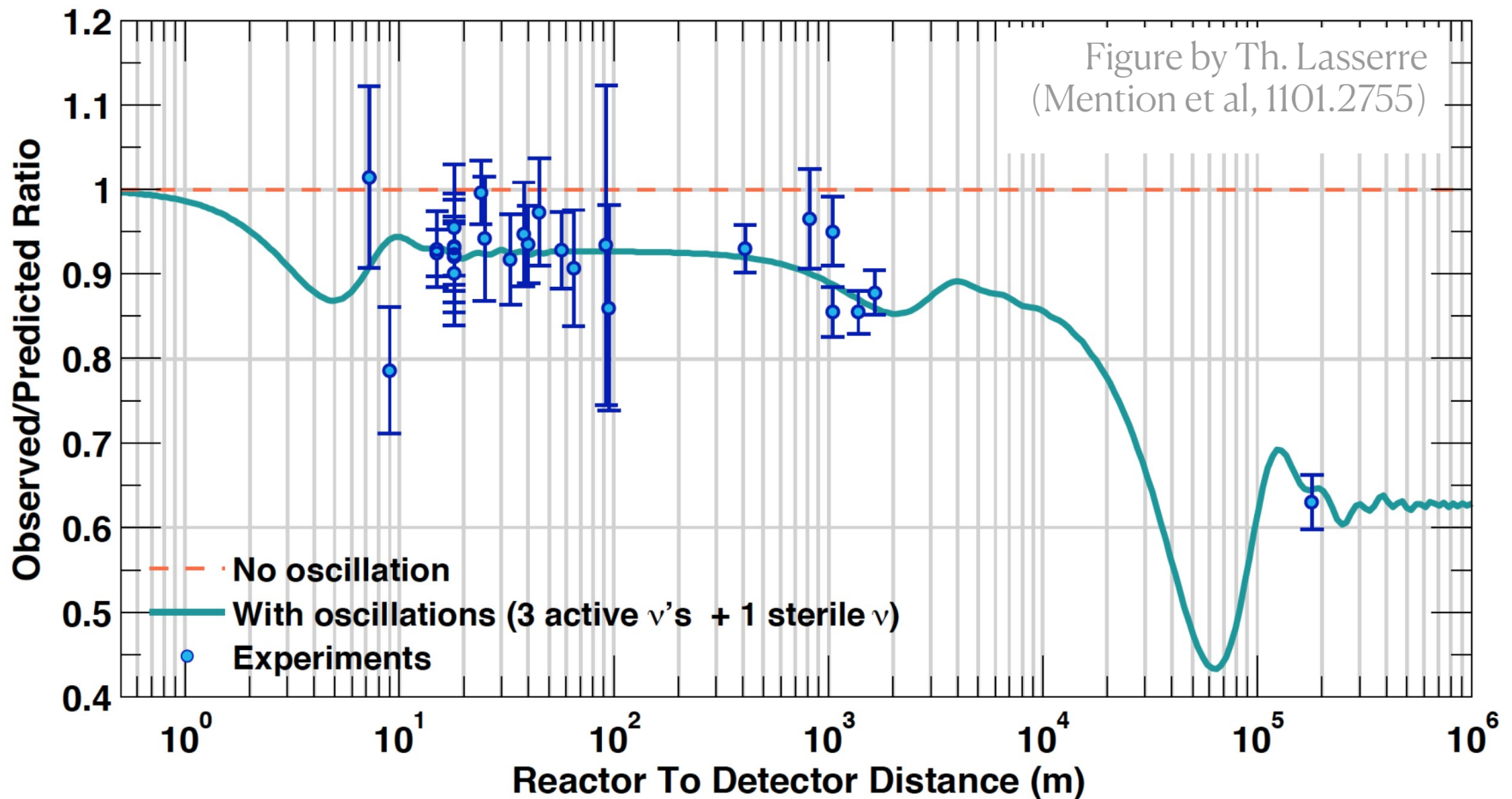


HM: Mueller et al 1101.2663, Huber 1106.0687;  
Mention et al, 1101.2755

Figure from Giunti, Li, Ternes, Xin, 2110.06820

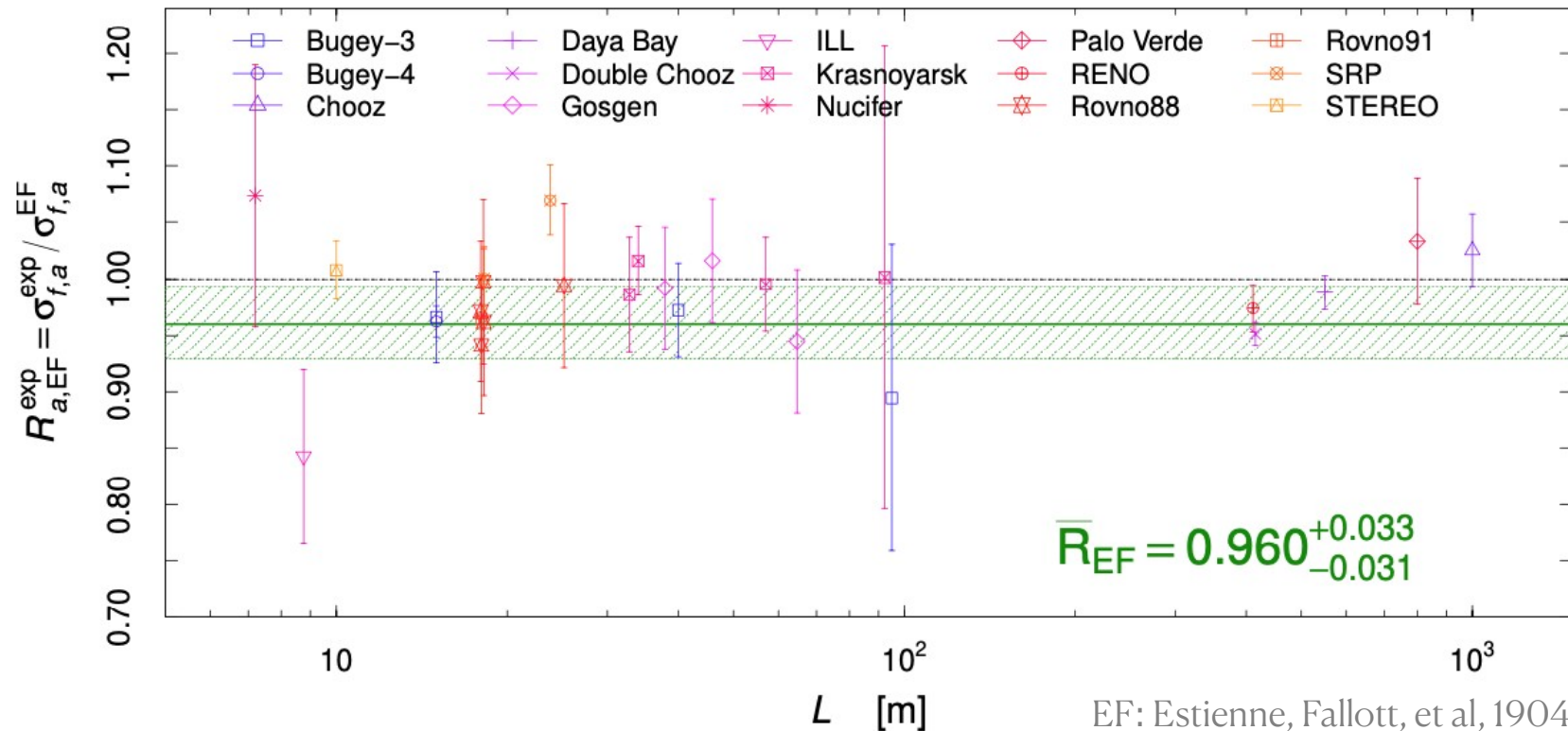
# Reactor antineutrino anomaly

$$P_{ee} = 1 - \sin^2 2\theta_{ee} \sin^2 \left( 1.27 \frac{\Delta m^2 [\text{eV}^2] L [\text{m}]}{E_\nu [\text{MeV}]} \right)$$



# Reactor antineutrino anomaly?

Situation in 2022:

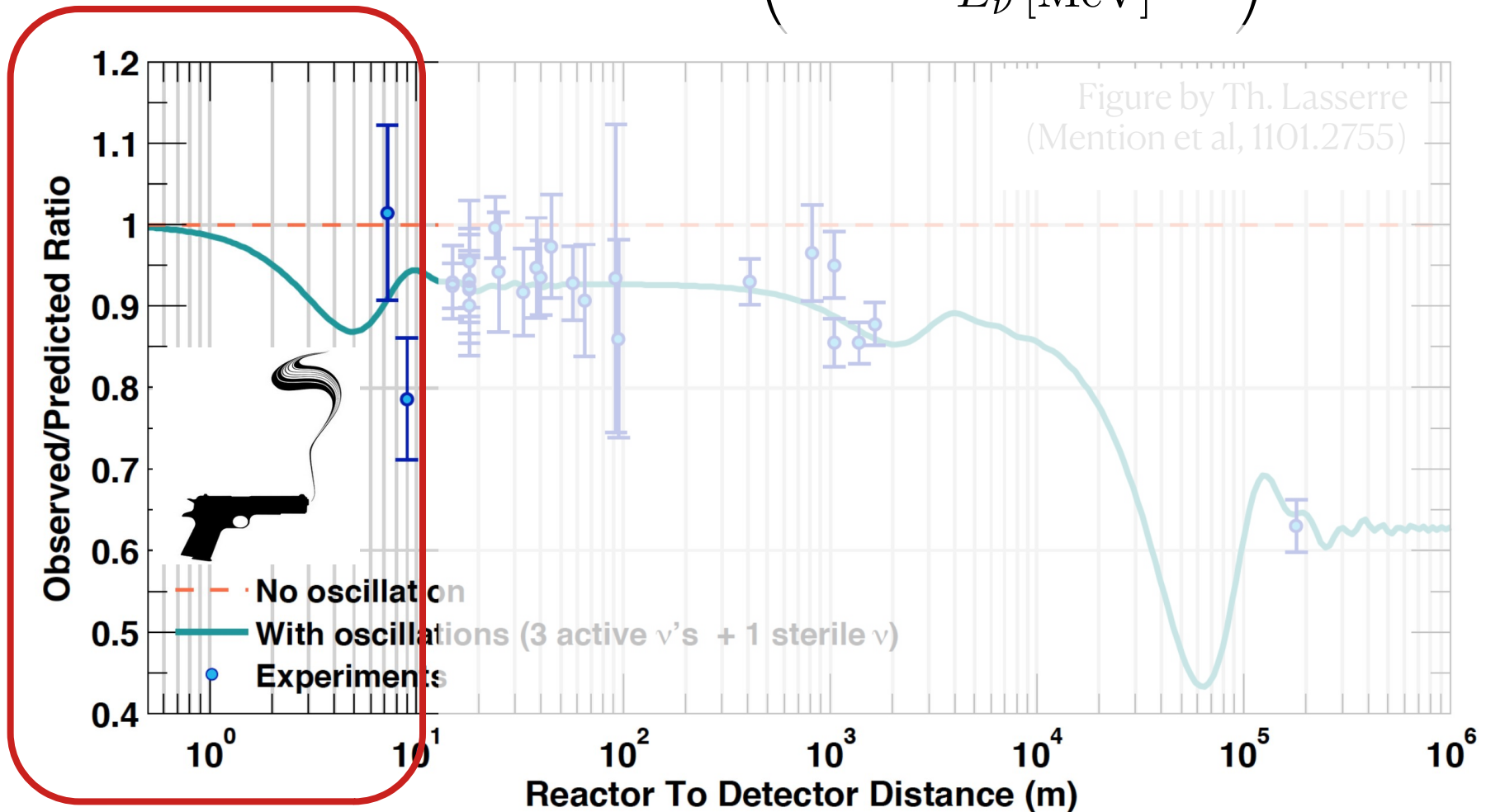


EF: Estienne, Fallott, et al, 1904.09358;

Berryman and Huber, 1909.09267 & 2005.01756  
Figure from Giunti, Li, Ternes, Xin, 2110.06820

# Short-baseline Reactors

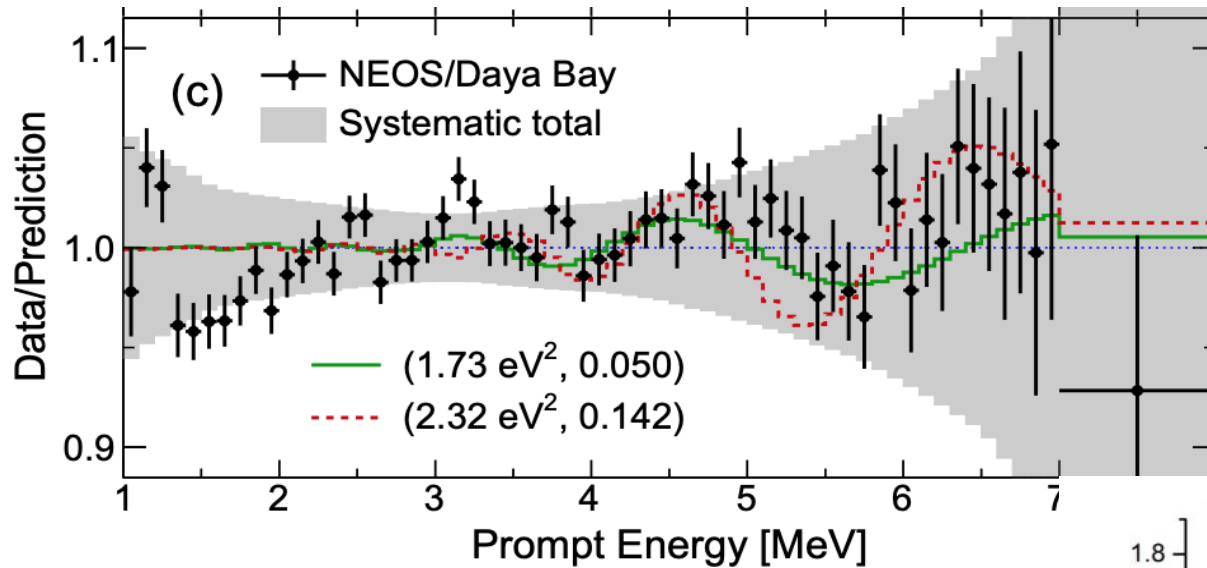
$$P_{ee} = 1 - \sin^2 2\theta_{ee} \sin^2 \left( 1.27 \frac{\Delta m^2 [\text{eV}^2] L [\text{m}]}{E_\nu [\text{MeV}]} \right)$$





# Very short-baseline reactors

NEOS coll., 1610.05134



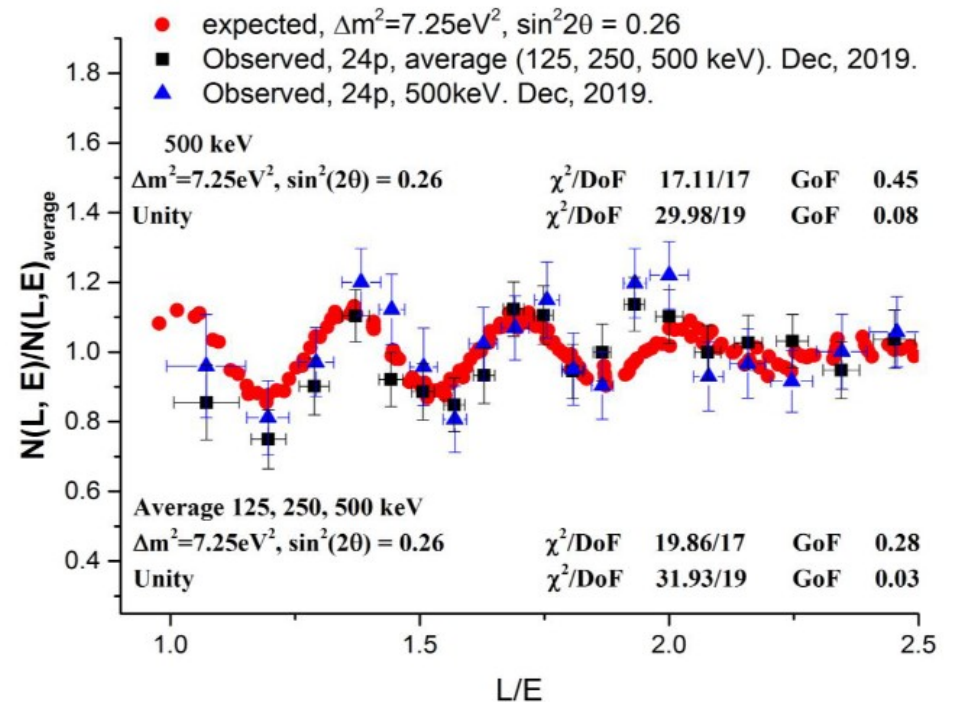
## Quick highlights:

None of the vSBL reactor experiments observed a clear signal at high significance

(**exception:** Neutrino-4 did report a signal at approx.  $3\sigma$ )

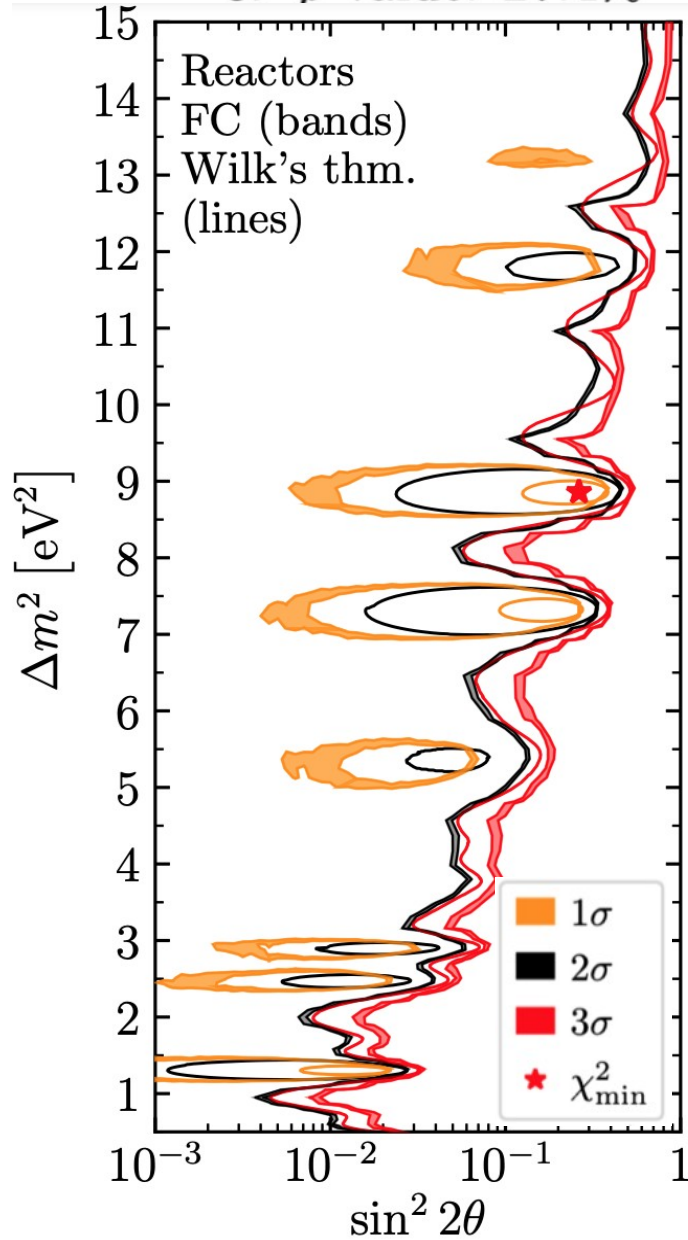
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Neutrino-4 coll., 2005.05301



# Reactors

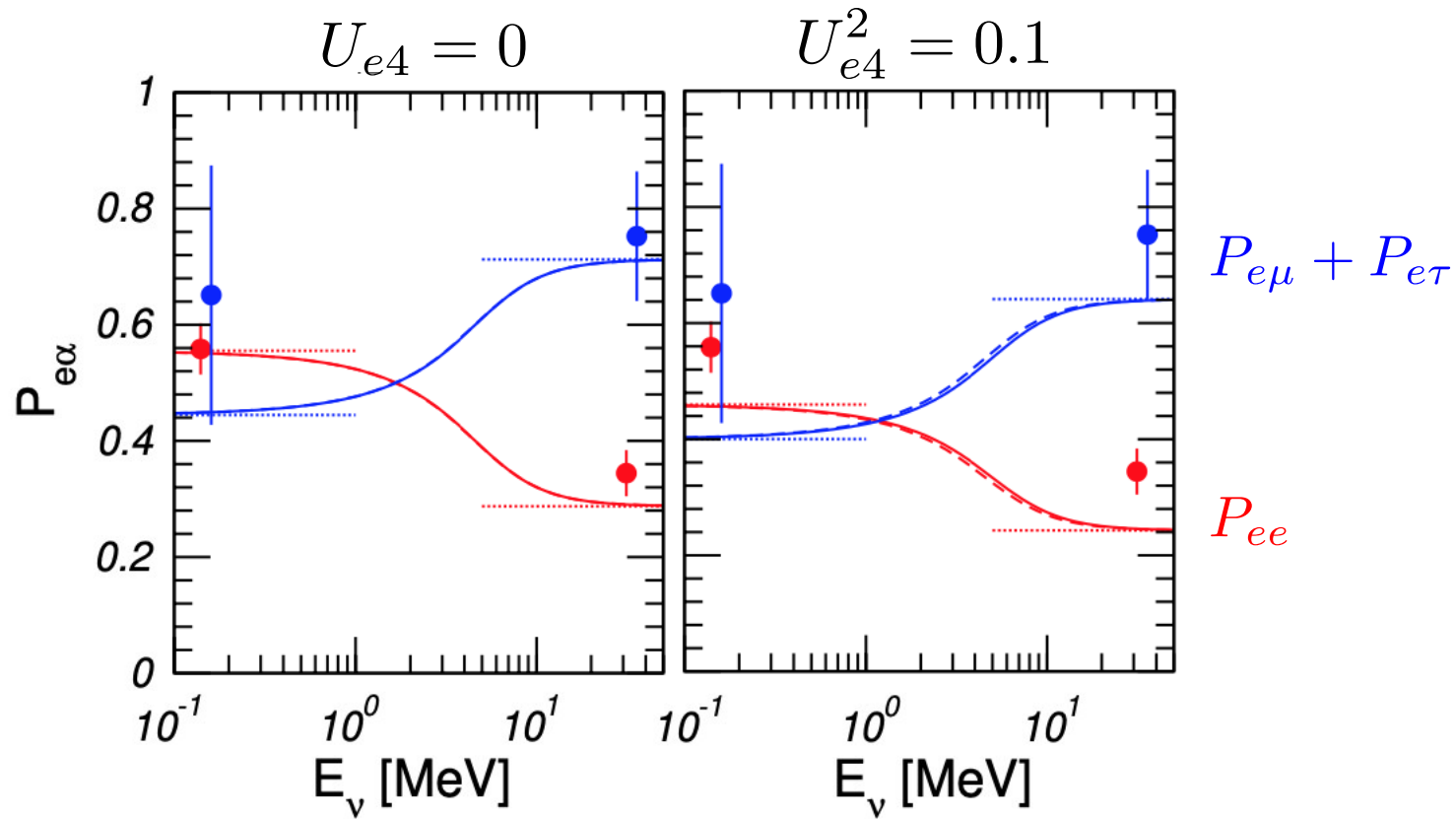
$3\nu$   $p$ -value: 27.4%



Berryman, Coloma, Huber,  
Schwetz, Zhou, 2111.12530

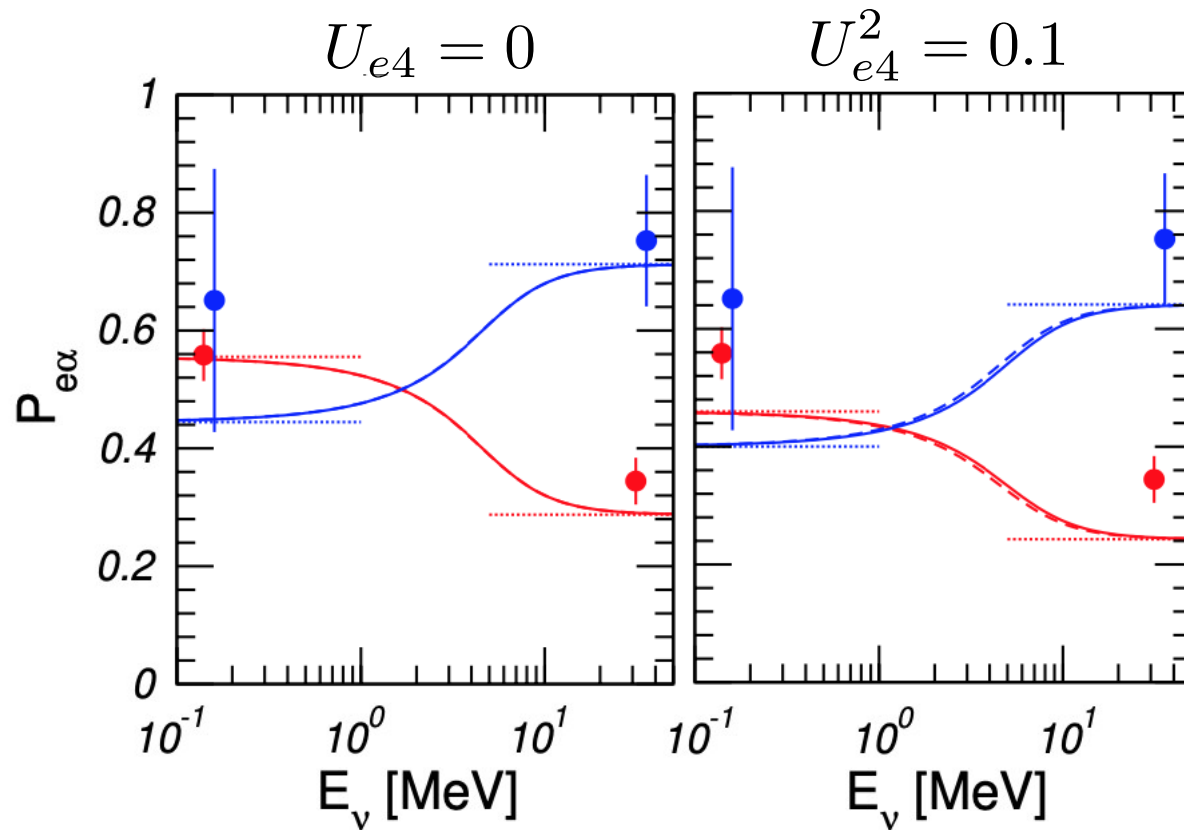
Global fit using data from:  
DANSS (talk at [EPS-HEP 2021](#))  
STEREO, 1912.06582  
PROSPECT, 2006.11210  
Neutrino-4, 2005.05301 (v2)  
NEOS, 1610.05134

# Solar neutrinos



Goldhagen, Maltoni, Reichard, Schwetz, 2109.14898  
(solar analysis as in Esteban et al, 2007.14792)

# Solar neutrinos



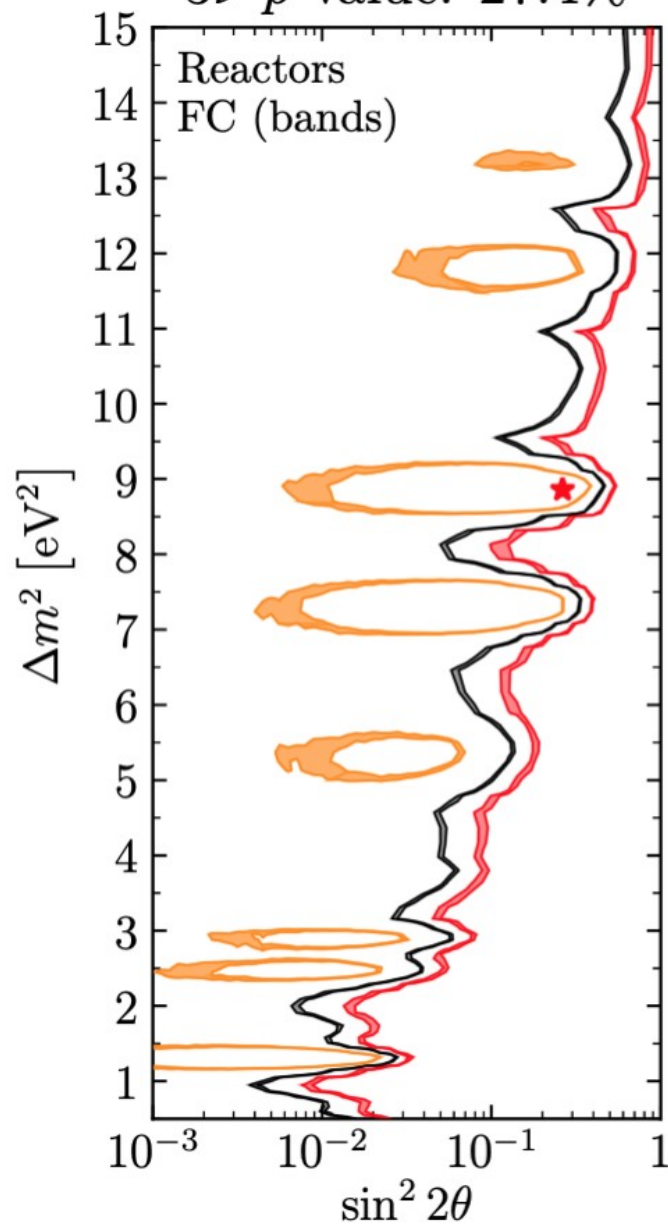
$$|U_{e4}|^2 < 0.168$$

(90% C.L.)

Goldhagen, Maltoni, Reichard, Schwetz, 2109.14898  
(solar analysis as in Esteban et al, 2007.14792)

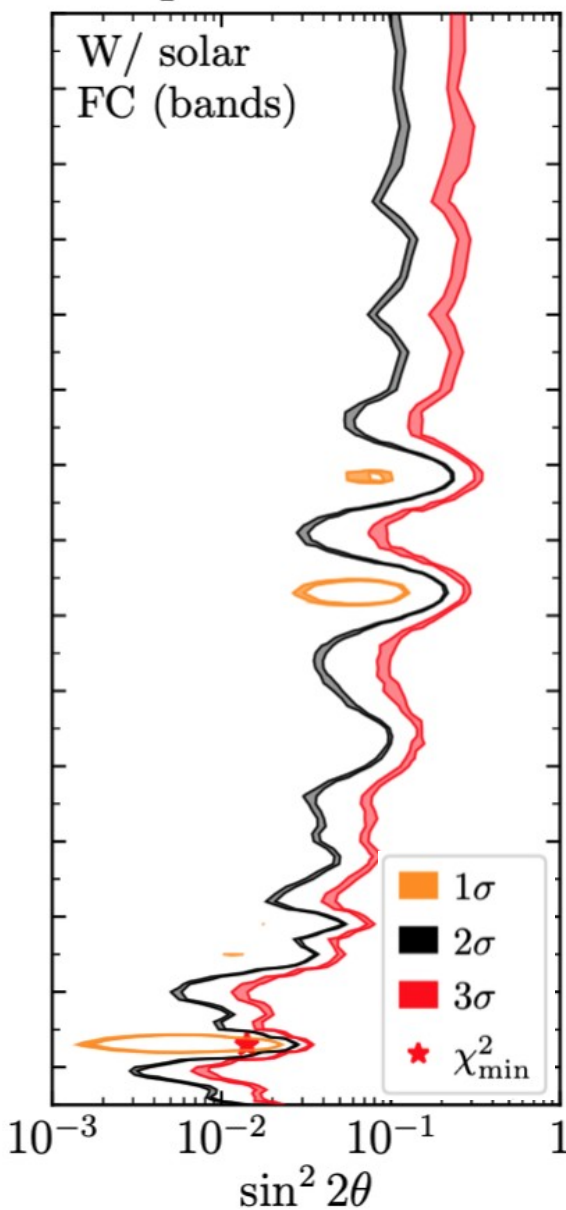
## Reactors

$3\nu$   $p$ -value: 27.4%

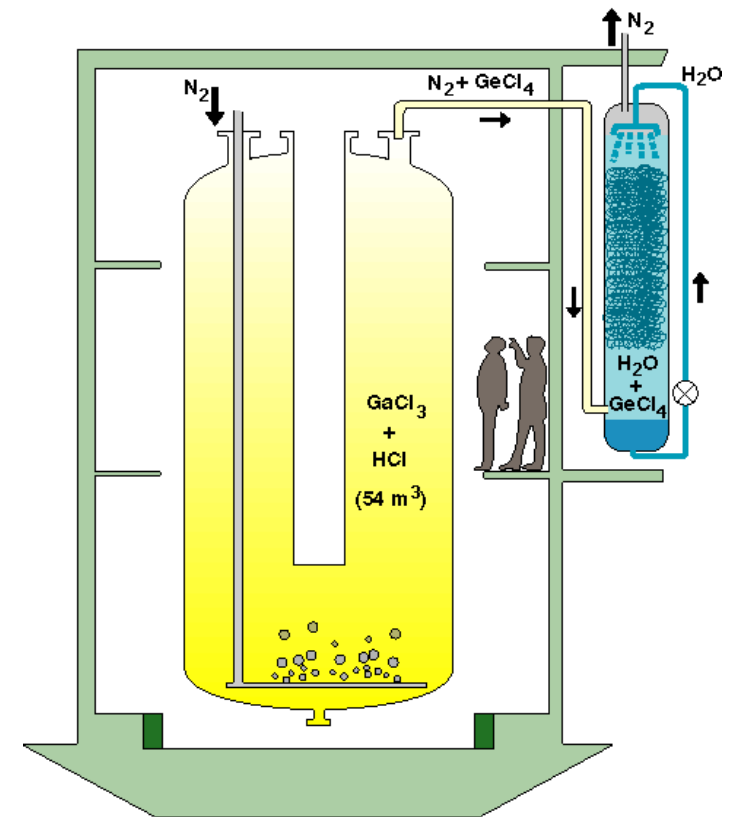
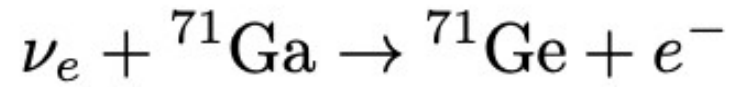
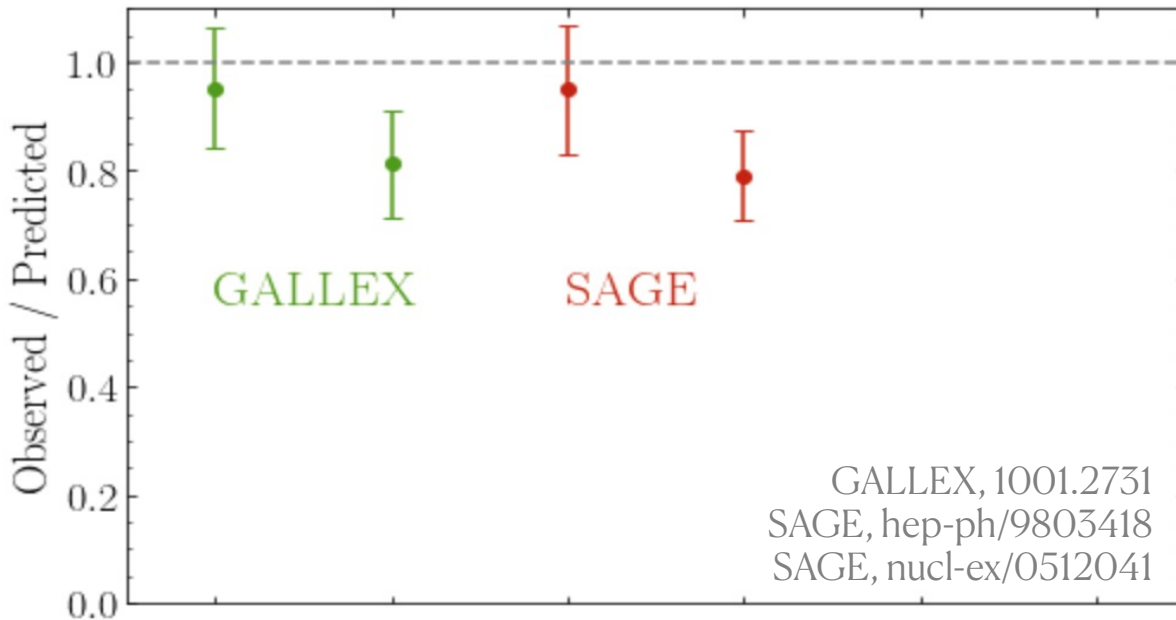


## Reactors + Solar

$3\nu$   $p$ -value: 17.8%



# Gallium experiments



# Gallium experiments

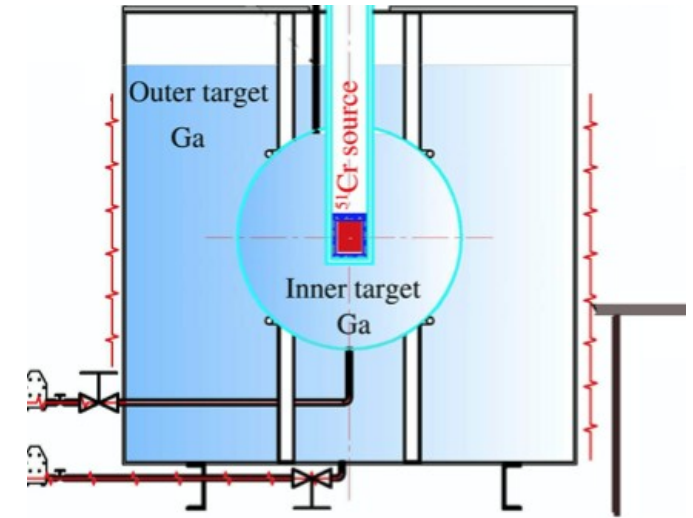
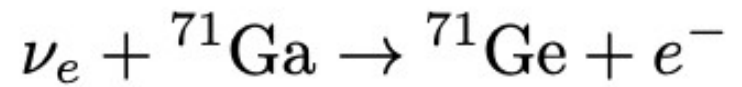
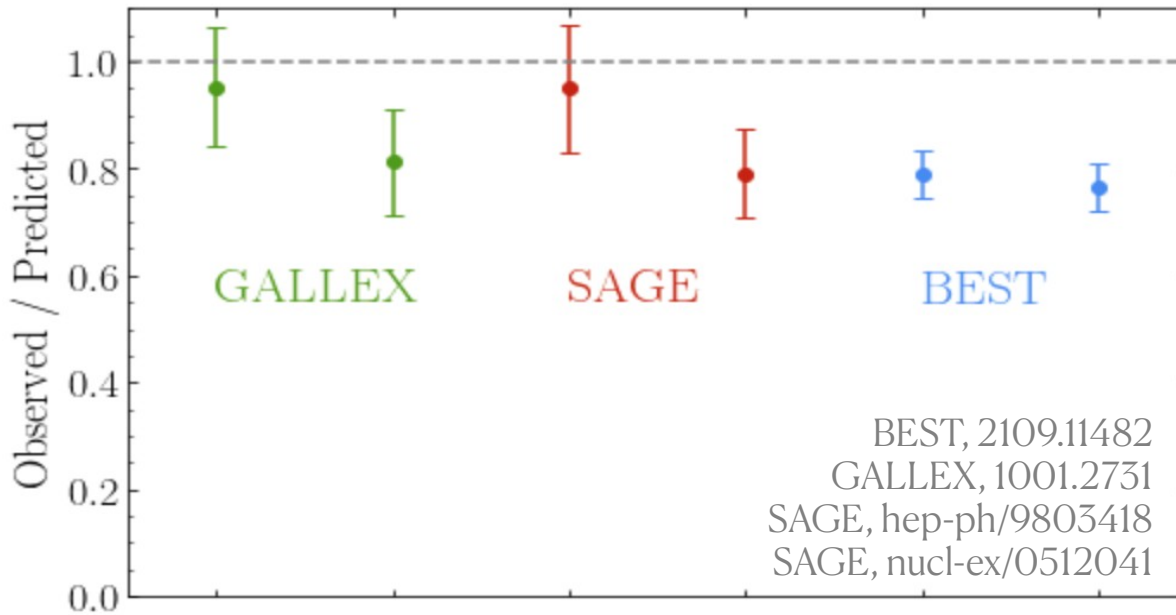
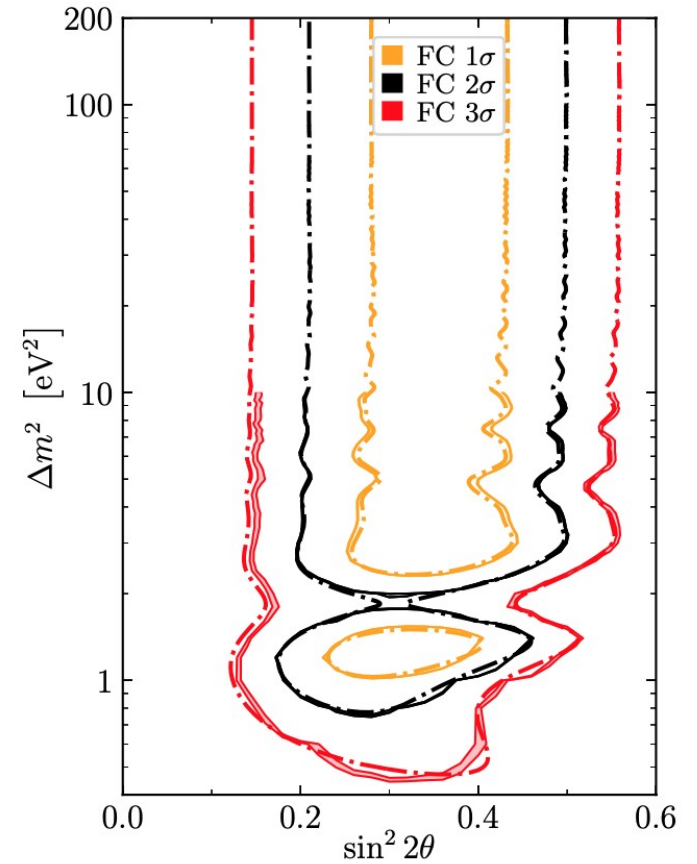
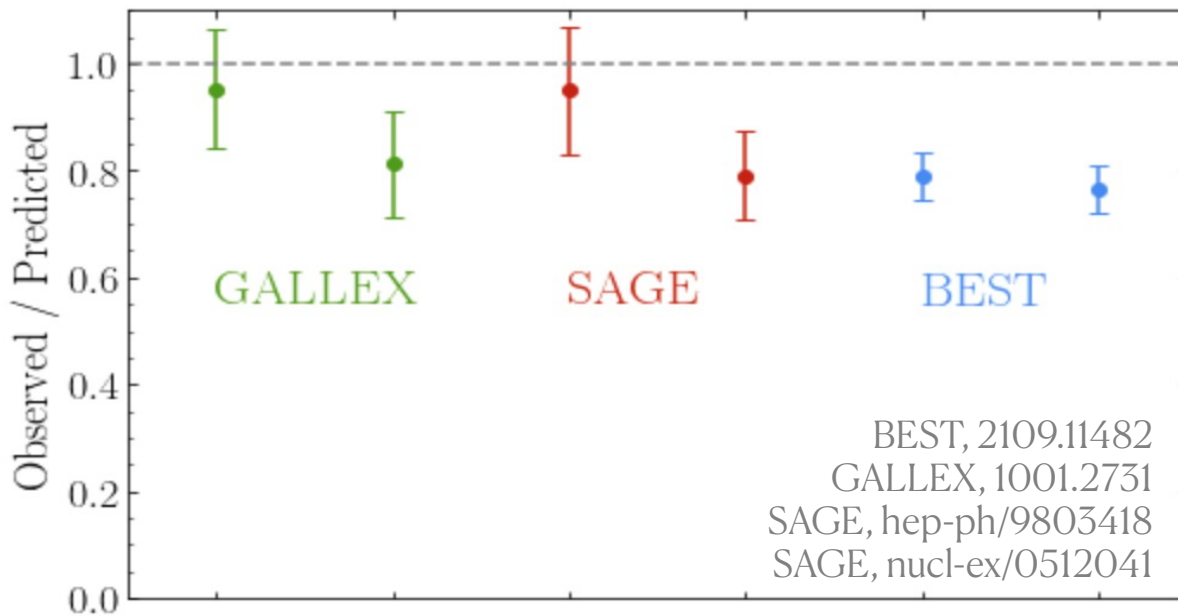


Figure from 2109.11482

# Gallium experiments



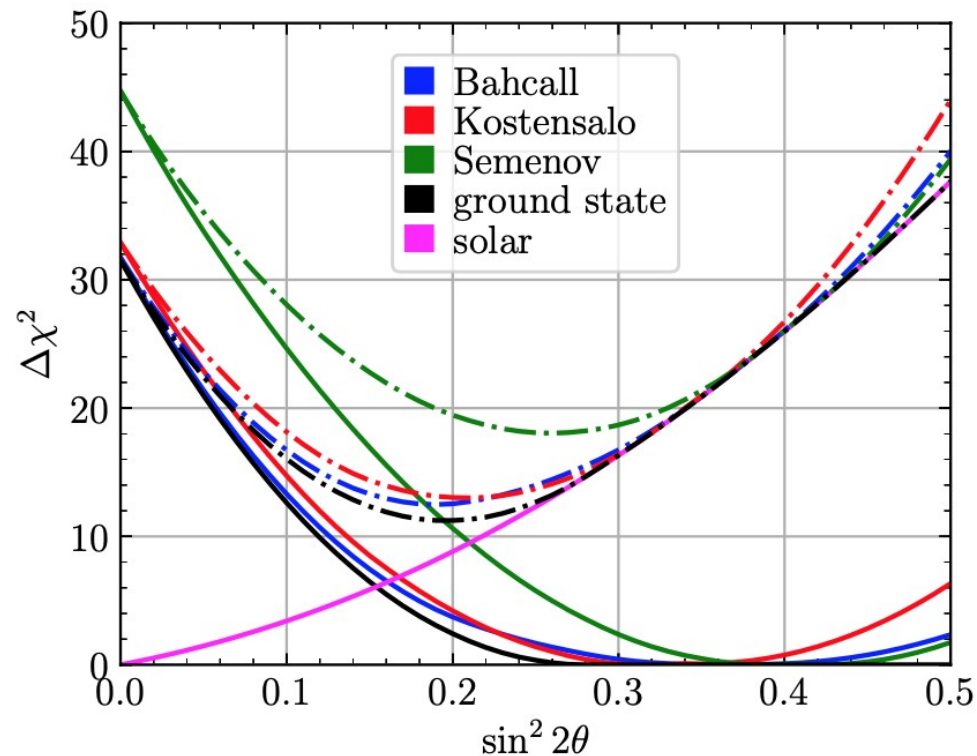
→ From MC simulation:

$$p_0 < 2.7 \times 10^{-8} \text{ (} 5.6\sigma \text{)}$$

Berryman, Coloma, Huber, Schwetz, Zhou, 2111.12530  
 (see also Barinov, Gorbunov, 2109.14654)



# Consistency tests: Solar vs Ga

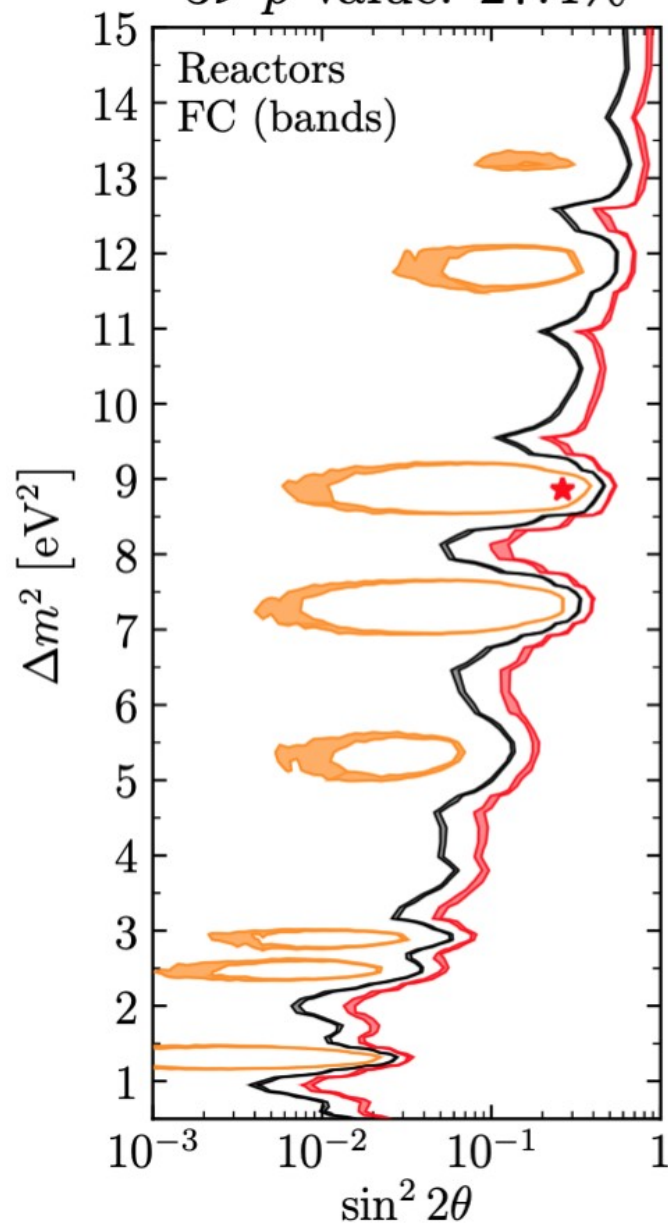


→ Tension above  $3\sigma$ !

Data set	$\chi^2_{\text{PG}}/\text{dof}$	$p^{(W)}$	$\#\sigma^{(W)}$	$p_{\text{b.f.}}$	$\#\sigma_{\text{b.f.}}$
Reactor vs Solar	0.65/1	0.42	0.8	0.39	0.9
Reactor vs Gallium	1.4/2	0.50	0.67	0.62	0.5
Solar vs Gallium	13.0/1	$3.1 \times 10^{-4}$	3.6	$1.6 \times 10^{-3}$	3.2
Reactor vs Solar vs Gallium	15.6/3	$1.4 \times 10^{-3}$	3.2	$5.1 \times 10^{-3}$	2.8

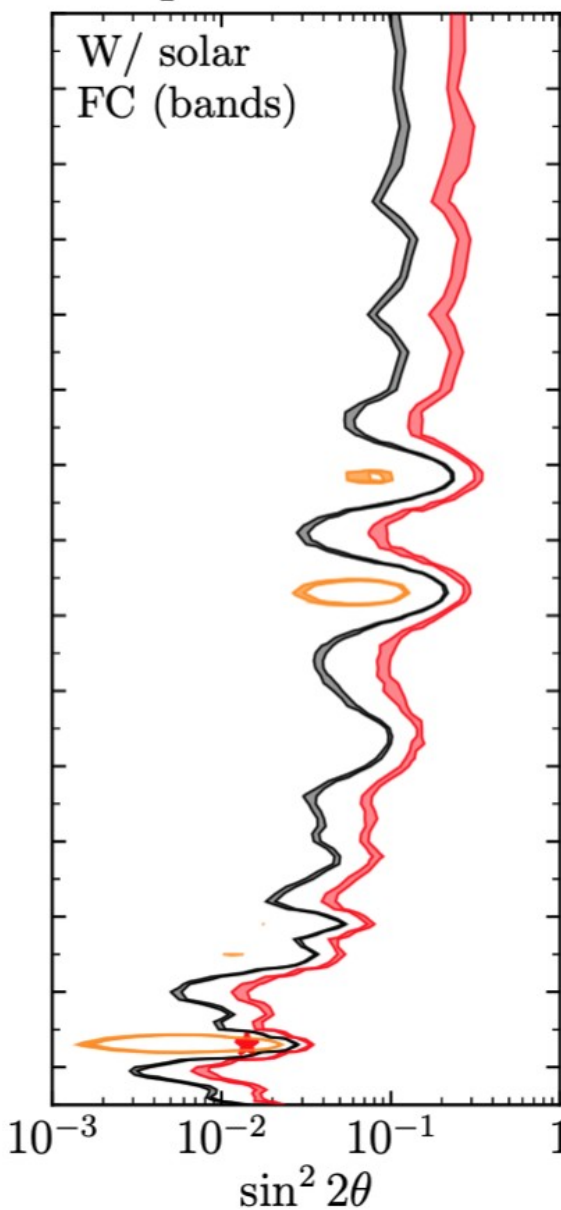
### Reactors

$3\nu$   $p$ -value: 27.4%



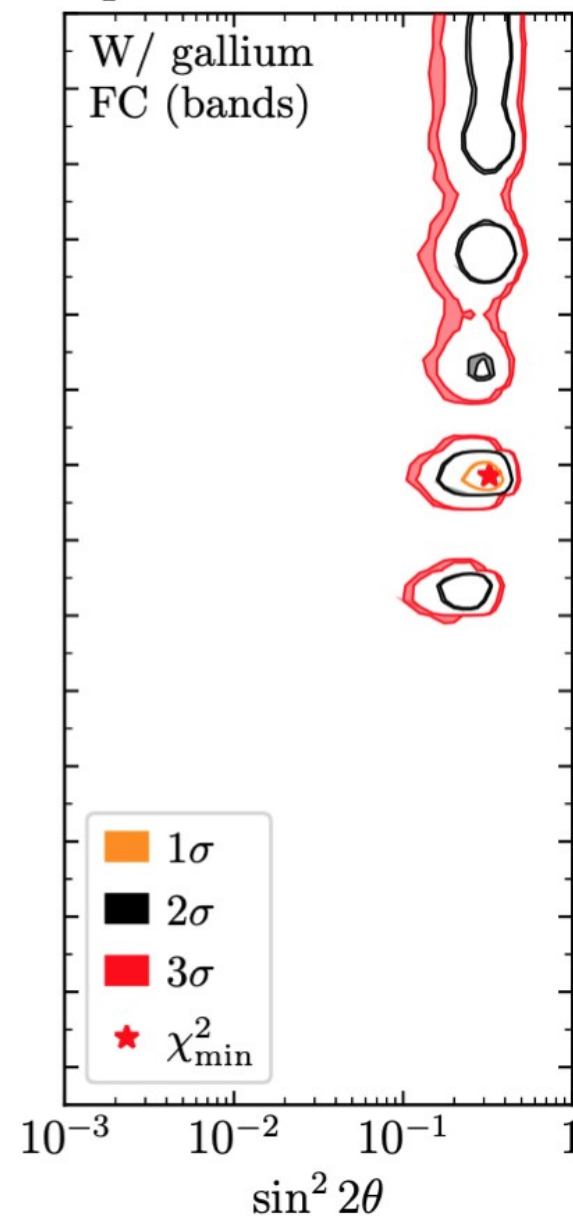
### Reactors + Solar

$3\nu$   $p$ -value: 17.8%



### Reactors + Ga

$3\nu$   $p$ -value:  $< 1.4 \times 10^{-7}$

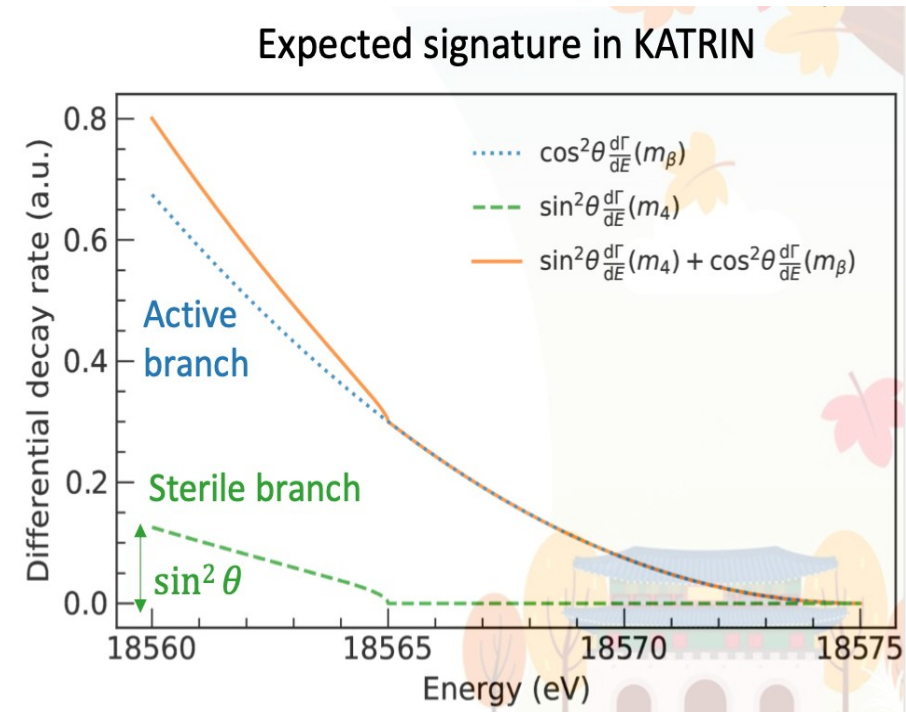


Other probes:  
end-point of beta-decay

$$m_\nu^2 = \sum_{i=1}^4 |U_{ei}|^2 m_i^2$$

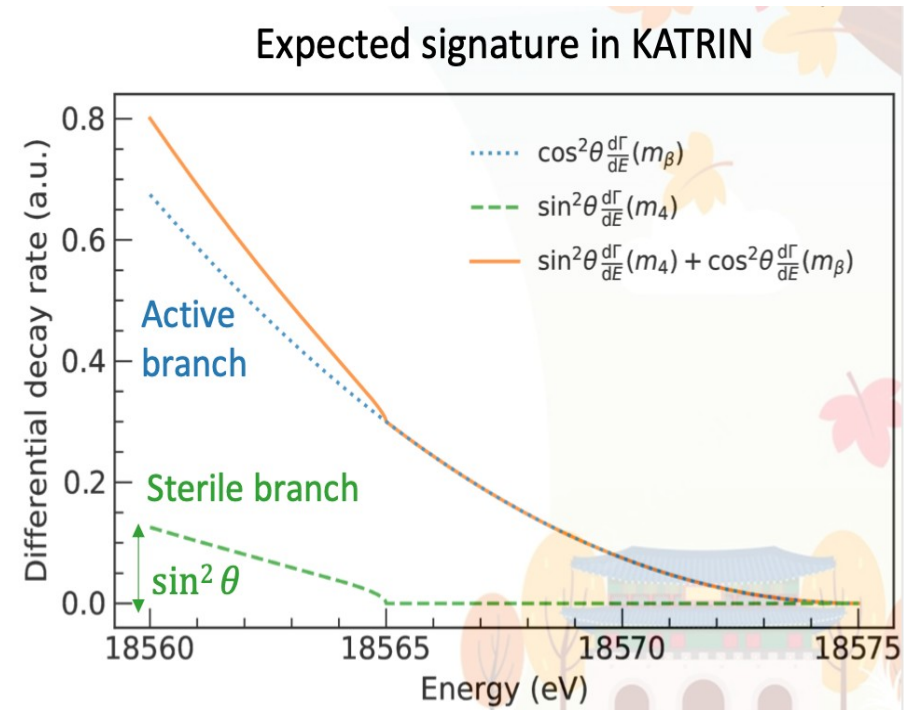
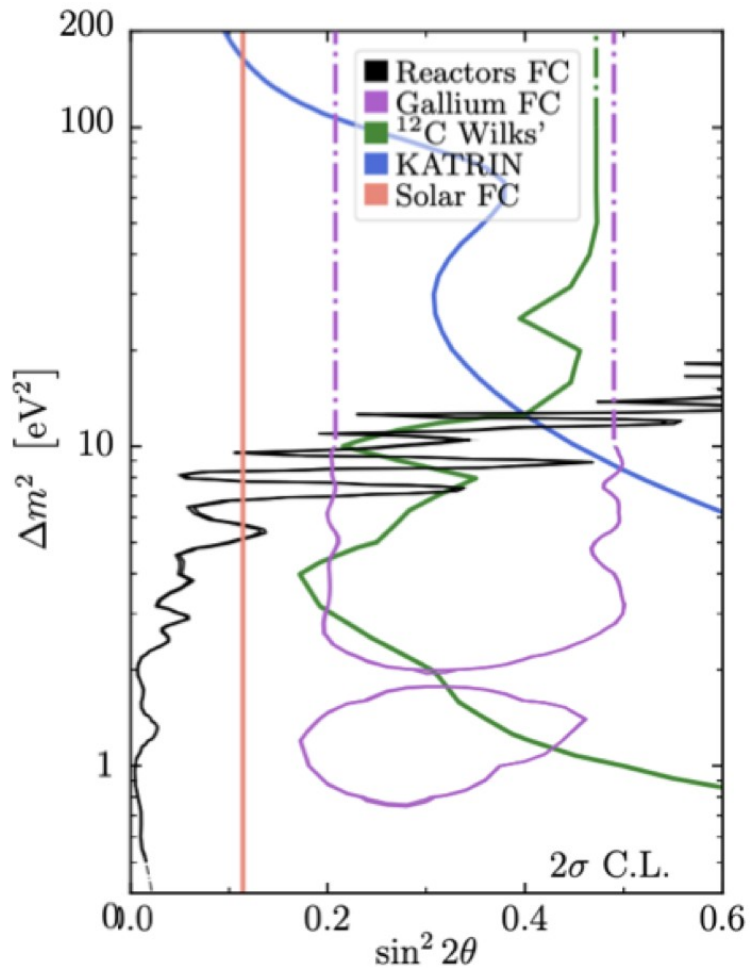
KATRIN

# KATRIN



Riis and Hannestad, 1008.1495;  
Formaggio and Barrett, 1105.1326

# KATRIN



Riis and Hannestad, 1008.1495;  
 Formaggio and Barrett, 1105.1326

Berryman, Coloma, Huber, Schwetz, Zhou,  
 2111.12530

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# Summary

- LSND and MiniBooNE remain unexplained
  - MicroBooNE has shown first results for  $\bar{\nu}+1$  (still limited)
- Reactor flux anomaly: mostly **resolved** (?)
- Reactor spectra: **significance** of hints **reduced** to  $\sim 1\sigma$
- Gallium anomaly (dominated by BEST) is **above  $5\sigma$** 
  - FC does not change this; nor does the cross section
- Global picture
  - Solar + reactor spectra: **compatible**, consistent with  $\bar{\nu}$ -nu ( $1.3\sigma$ )
  - Ga + reactor spectra: **compatible**, favors steriles ( $>5\sigma$ )
  - Ga + solar: in **tension** at  $3\sigma$
  - Ga + reactor flux: in **tension** too...
- In the near future, KATRIN will probe the region at high mass

Thanks!

Work supported by Grants RYC2018-024240-I  
PID2019-108892RB-I00, CEX2020-001007-S



EXCELENCIA  
SEVERO  
OCHOA



# Backup



# Dependence on capture cross sec.

Reference	$\sigma(\text{Cr})$	$\sigma_{\text{g.s.}}(\text{Cr})$	$\sigma(\text{Ar})$	$\sigma_{\text{g.s.}}(\text{Ar})$
Bahcall [56]	$58.1 \pm 2.1$	55.2	$70.0 \pm 4.9$	66.2
Kostensalo <i>et al.</i> [54]	$56.7 \pm 1.0$	$55.3 \pm 0.7$	$68.0 \pm 1.2$	$66.2 \pm 0.9$
Semenov [57]	$59.38 \pm 1.16$	$55.39 \pm 0.19$	$71.69 \pm 1.47$	$66.25 \pm 0.23$

[56] Bahcall, hep-ph/9710491

[54] Kostensalo, Suhonen, Giunti, Srivastava, 1906.10980

[57] Semenov, Phys. Atom. Nucl. 83 (2020) 1549.

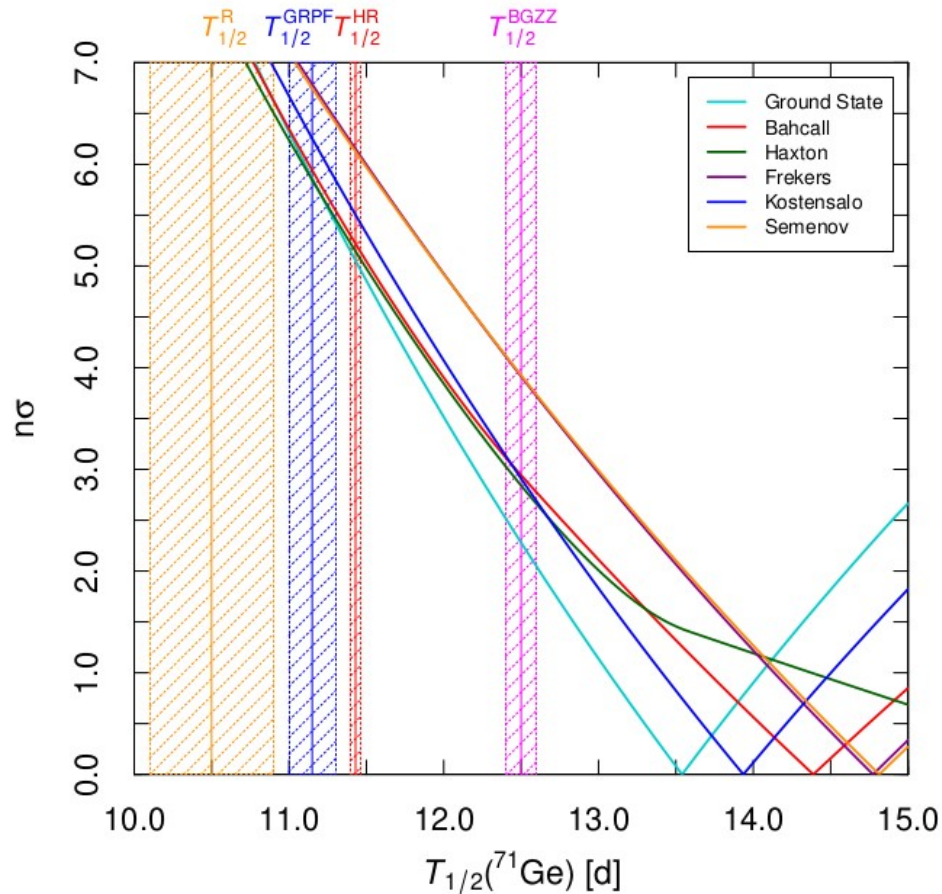
	GALLEX & SAGE		BEST		All gallium combined			
Cross section	$\Delta\chi_{3\nu}^2$	$\#\sigma^{(W)}$	$\Delta\chi_{3\nu}^2$	$\#\sigma^{(W)}$	$\sin^2 2\theta_{\text{min}}$	$\Delta m_{\text{min}}^2$	$\Delta\chi_{3\nu}^2$	$\#\sigma^{(W)}$
Bahcall	3.7	1.4	31.3	5.2	0.35	1.3 eV <sup>2</sup>	31.7	5.3
Kostensalo	4.9	1.7	31.5	5.2	0.32	1.3 eV <sup>2</sup>	32.9	5.4
Semenov	9.4	2.6	42.4	6.2	0.39	1.3 eV <sup>2</sup>	44.7	6.4
Ground state	3.4	1.3	29.7	5.1	0.29	1.3 eV <sup>2</sup>	31.5	5.3

Our default choice →

Berryman, Coloma, Huber, Schwetz, Zhou, 2111.12530  
(see also Barinov, Gorbunov, 2109.14654)

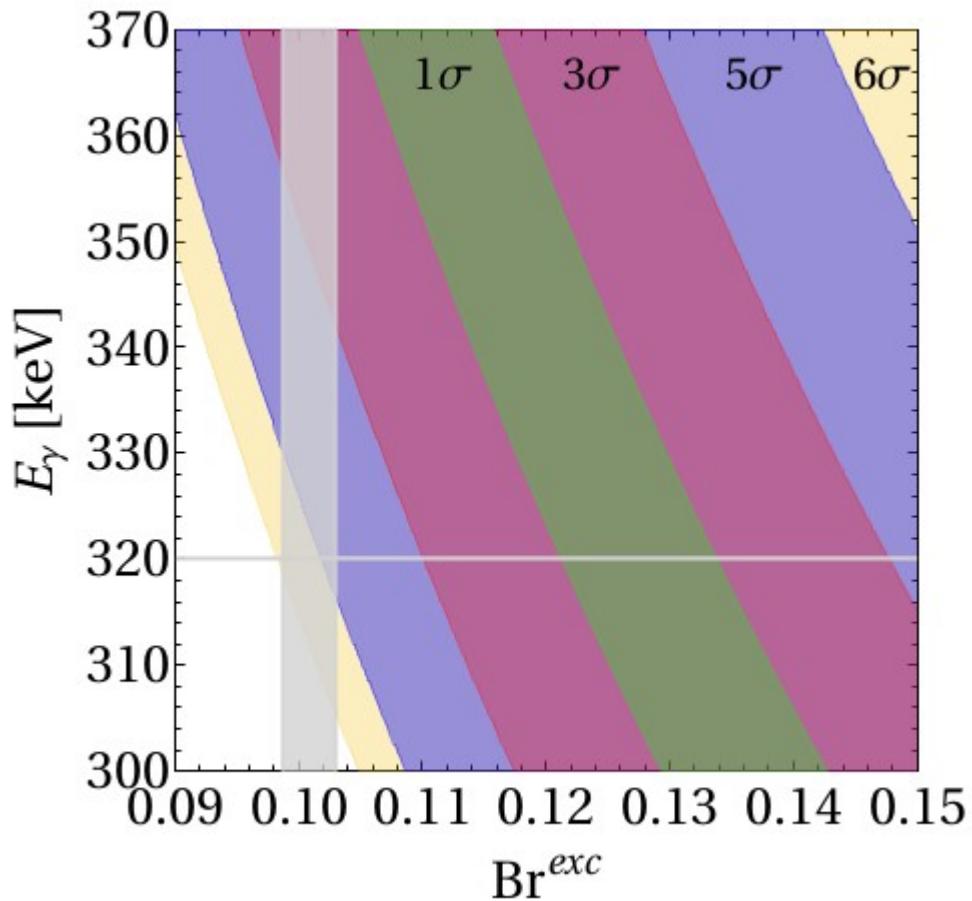
# Dependence on capture cross sec.

$$\sigma_{\text{gs}} = \frac{G_F^2 \cos^2 \vartheta_C}{\pi} g_A^2 \text{BGT}_{\text{gs}} \langle p_e E_e F(Z_{\text{Ge}}, E_e) \rangle = \frac{\pi^2 \ln 2}{m_e^5 f t_{1/2}({}^{71}\text{Ge})} \langle p_e E_e F(Z_{\text{Ge}}, E_e) \rangle$$

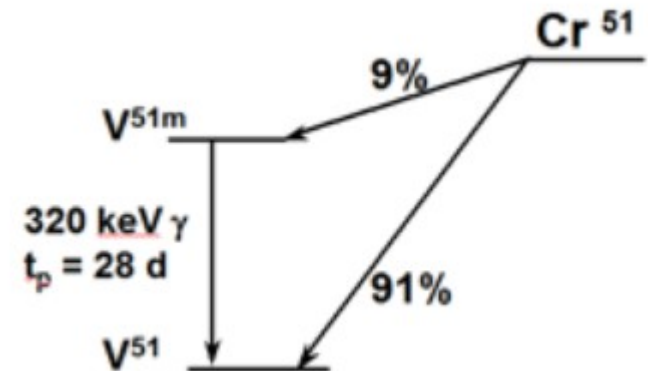


Giunti and Ternes, 2212.09722

# Further scrutiny



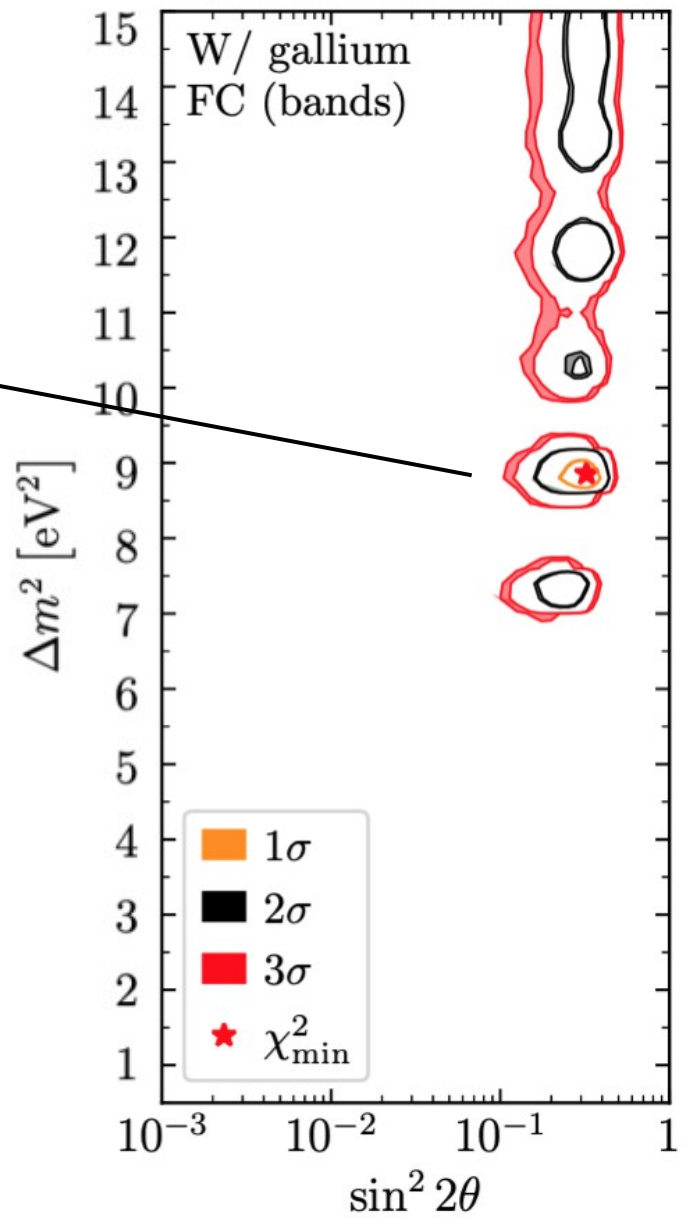
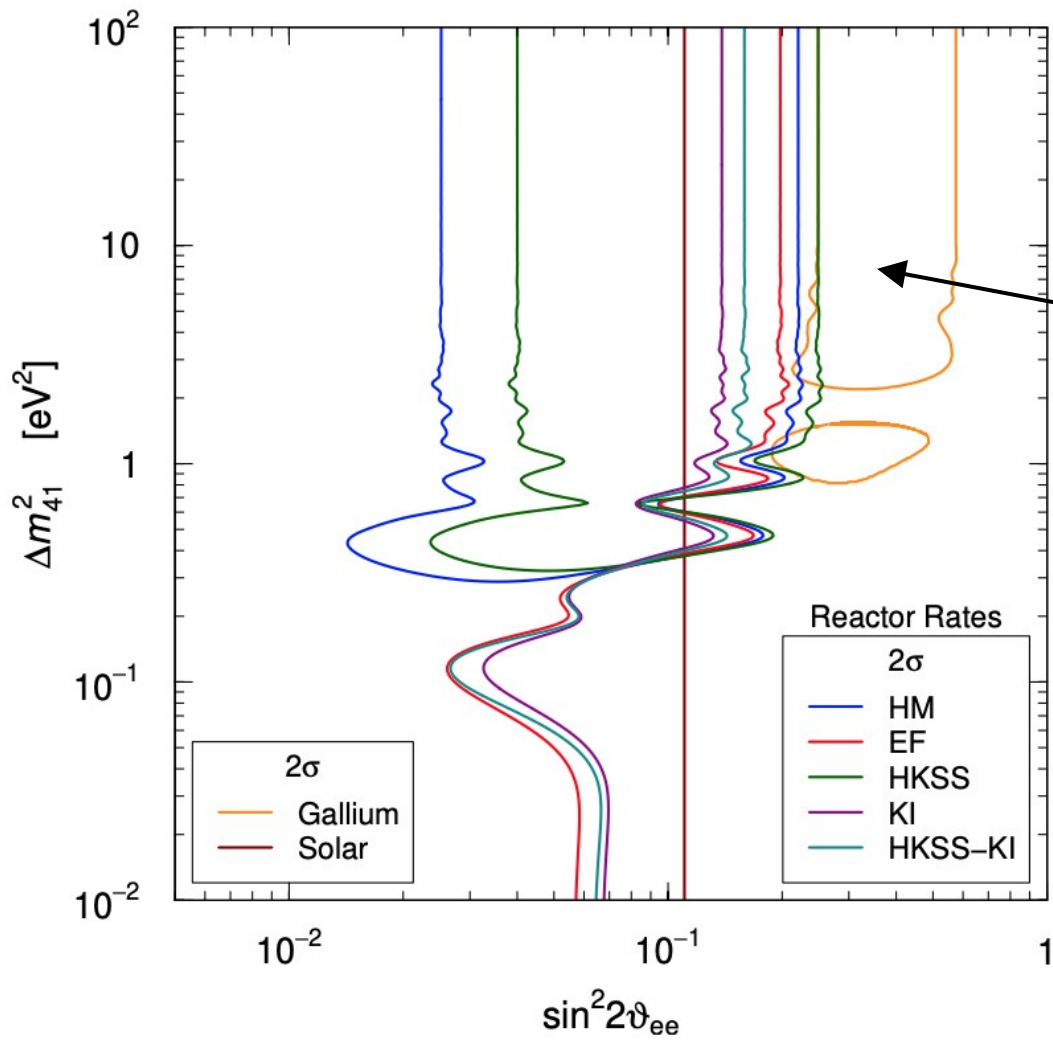
The source calibration is done calorimetrically; however, the BR to the excited state is only 10%...



# Reactors + Ga

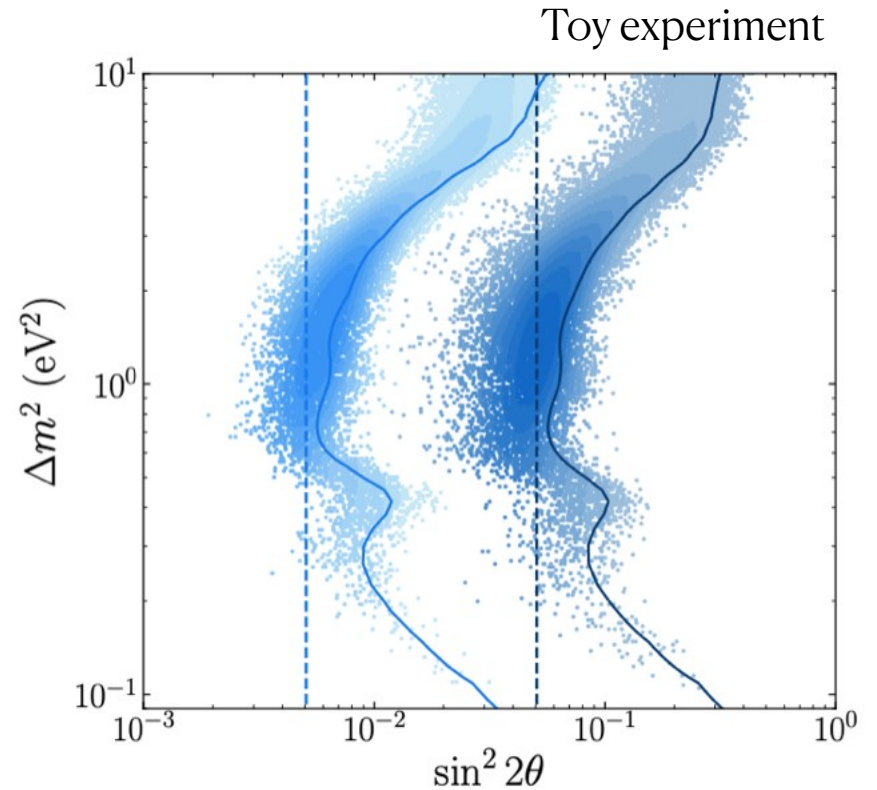
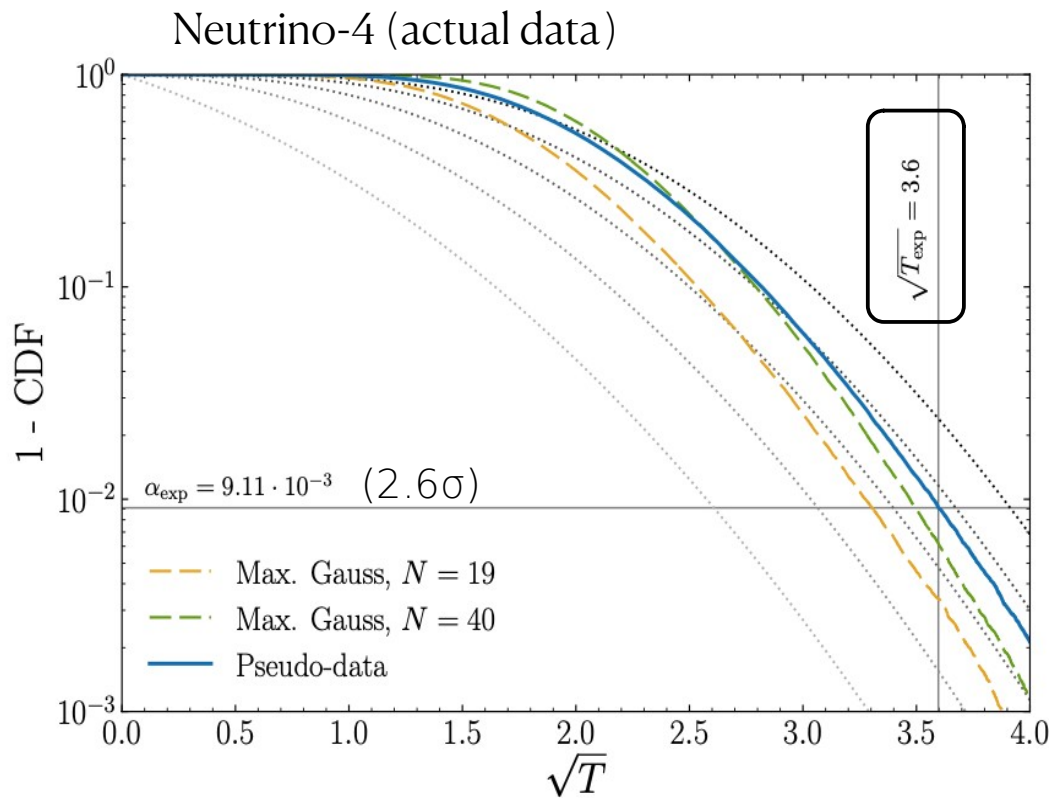
→ Gallium in significant tension with new reactor flux predictions too...

$3\nu$   $p$ -value  $< 1.4 \times 10^{-7}$



Berryman and Huber, 1909.09267 & 2005.01756  
 Figure from Giunti, Li, Ternes, Xin, 2110.06820

# Oscillations or fluctuations?



$$T = \chi_{3\nu}^2 - \chi_{\text{min}}^2$$

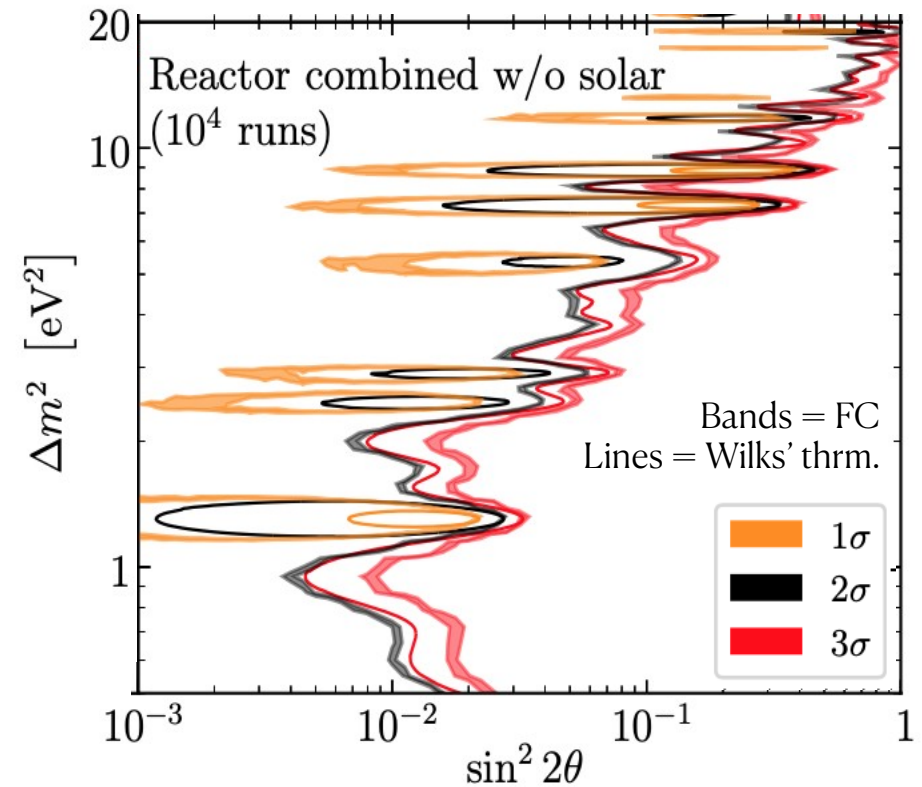
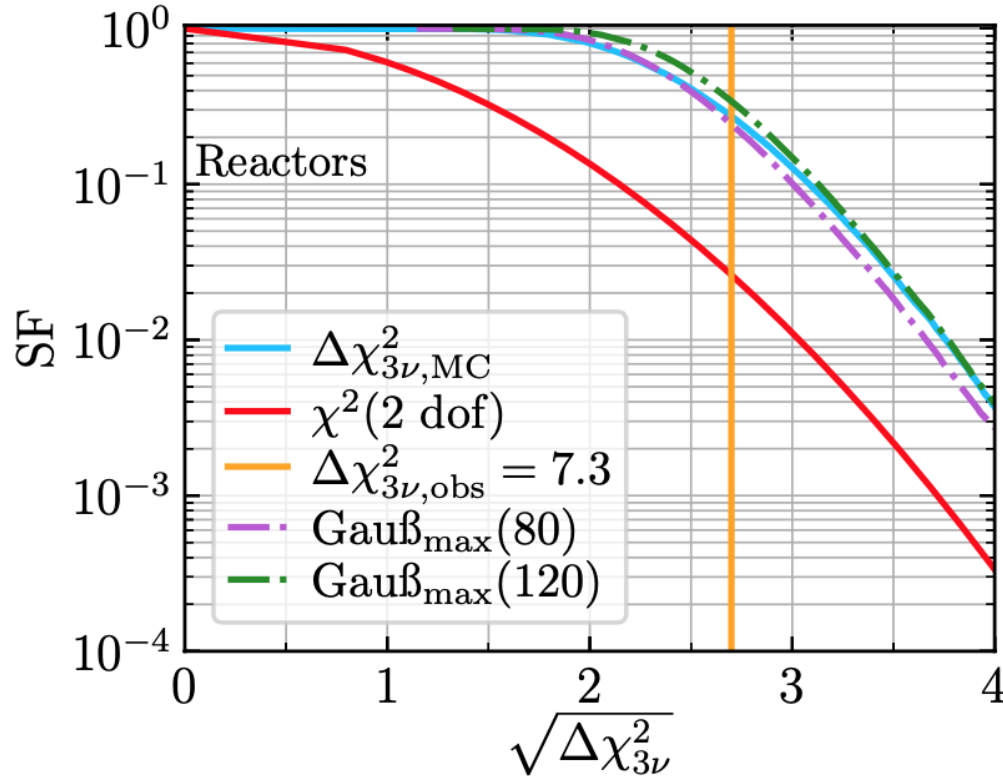
Coloma, Huber, Schwetz, 2008.06083  
 (see also Giunti, 2004.07577; Agostini, Neumair, 1906.11854;  
 PROSPECT and STEREO collaborations, 2006.13147)

# Combined reactors: FC analysis

$$\Delta\chi_{3\nu}^2 = \chi_{3\nu}^2 - \chi_{\min}^2$$

$$\Delta\chi^2(\{\Theta\}) = \chi^2(\{\Theta\}) - \chi_{\min}^2$$

3-nu hypothesis, p-value:  $p_0 = 27.4\%$



Berryman, Coloma, Huber, Schwetz, Zhou, 2111.12530