

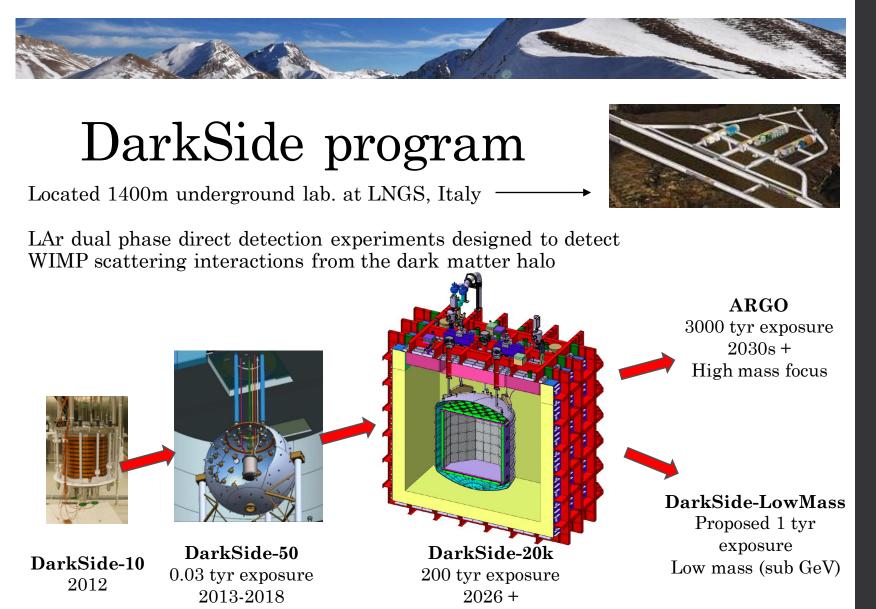


DarkSide-50/-20k experiments –

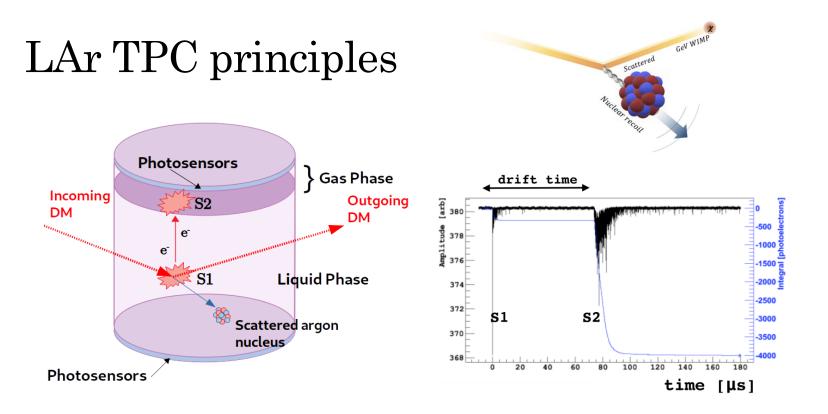
recent results and prospects

Ellen Sandford, on behalf of the DarkSide Collaboration Lake Louise Winter Institute

21st February 2023

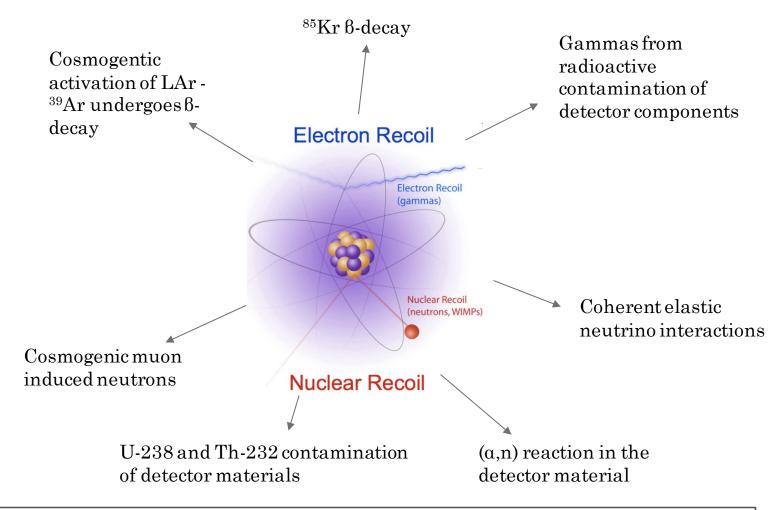


Multiple collaborations joined together in 2017 to form the Global Argon DM Collaboration (GADMC) for future LAr-based dark matter detectors



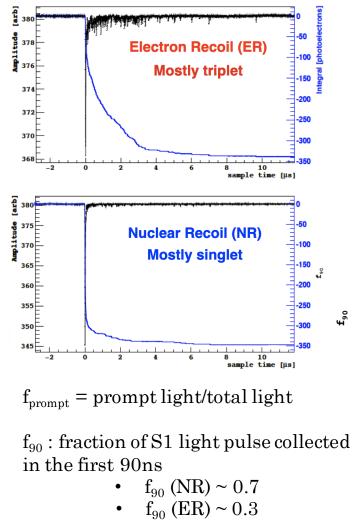
- DM-nucleus interaction creates initial scintillation pulse (S1) in LAr
- Electric field drifts ionisation electrons up to top of TPC
- Electrons are accelerated across gas phase to produce secondary ionisation pulse (S2)
- These are seen by photosensors on top and bottom of the TPC
- Use the S1-S2 time difference to measure Z position of event, and the S2 light distribution to determine XY position: 3D reconstruction

Important backgrounds for DarkSide



DarkSide-20k aim: <0.1 background events in the total 200tyr exposure

Pulse shape discrimination

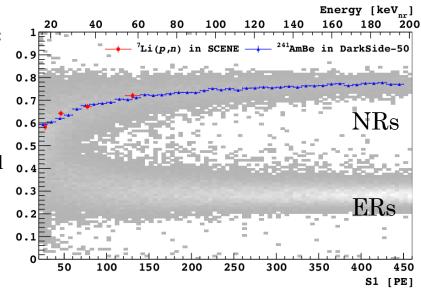


Argon dimers decay with two different decay constants:

 $tau_{\rm fast}\,{\sim}\,6$ ns and $tau_{\rm slow}\,{\sim}\,1600 \text{ns}$

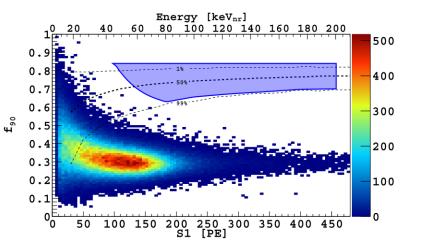
Electron recoil and nuclear recoil events excite a different fraction of singlet and triplet state decays.

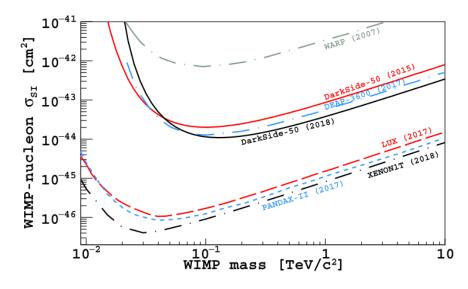
This can be utilised to discriminate between WIMP NR events and ER backgrounds.



DarkSide-50 high mass results

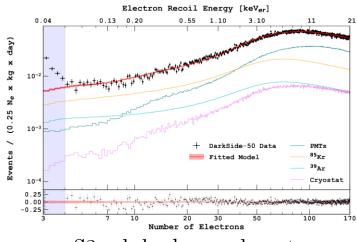
- Total exposure: 166,600 kg days, using 532 days of data
- Demonstrated 1400x reduction in ³⁹Ar backgrounds using underground argon vs atmospheric
- Reject ER background events using PSD to a level of one in 1.5x10⁷
- This allowed the blind DM search to be **background free**





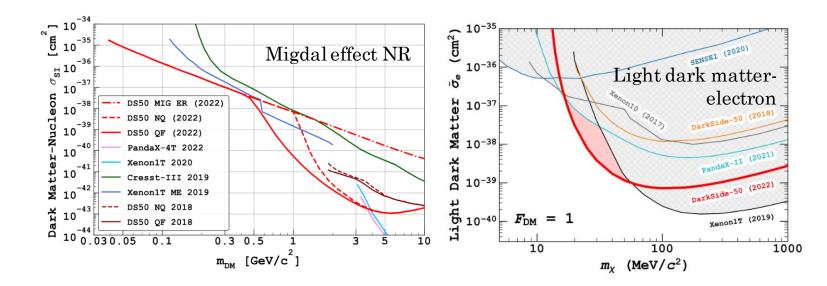
DarkSide-50 low mass results

arXiv:2207.11966 arXiv:2207.11967 arXiv:2207.11968



S2 only background spectra

- Using only the S2 signal we can access much lower energy events
- Analysis threshold : 4e- (0.6keVnr)
- Migdal effect: additional ER energy from electron cloud can push NR events above detector threshold
- Migdal analysis gives sensitivity down to 0.04 GeV most stringent limit in this mass region (below ~3 GeV)
- Also placed exclusion limits on a range of models with electron final states



DarkSide-20k design

UAr Compresso

ProtoDUNE Cryostat UAr Cryogenics

UAr Transportation Skid and Storage

Nested design:

Ti vessel Top OP

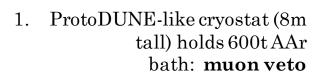
vPDU

TPC Barrel

Calibration

Pipe Bottom OP

Bottom OP with PDUs



2. Stainless steel vessel holds 12t UAr: **neutron veto**

3. Inner TPC contains 50t UAr

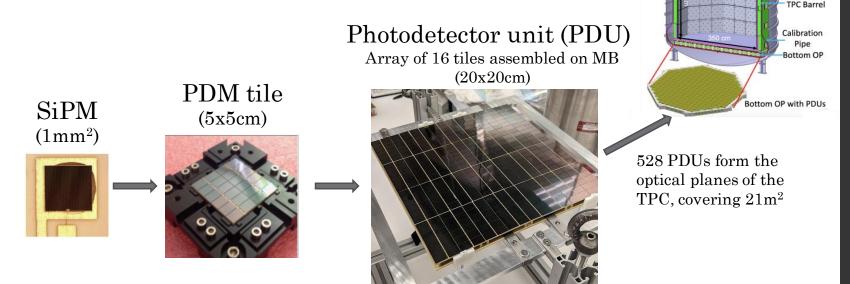


Hall C @ LNGS

Cryogenics Support Structure

- Octagonal TPC with inner diameter 350cm
- Instrumented with SiPM sensors on top and bottom optical plane: 21m² coverage
- Drift field: 2.8kV/cm
- Walls coated with Reflector + PEN (WLS)

DarkSide-20k SiPM sensors



Why SiPMs?:

 $Higher \, photon \, detection \, efficiency, \, lower \, radioactivity, \, lower \, cost, \, operated \, with \, lower \, bias$

Challenges:

- Small area need to produce large area arrays of SiPM sensors
- High electronic noise
- Higher dark rate and correlated noise

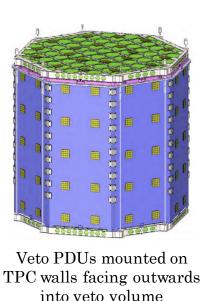
<u>Result:</u>

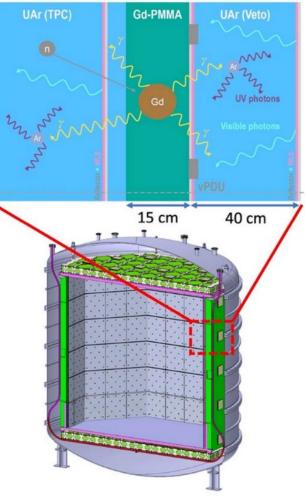
DarkSide have successfully developed state of the art large-area PDU sensors which meet and exceed the performance requirements of the detector, including having single photon resolution.

vesse

DS20k Background reduction: NRs

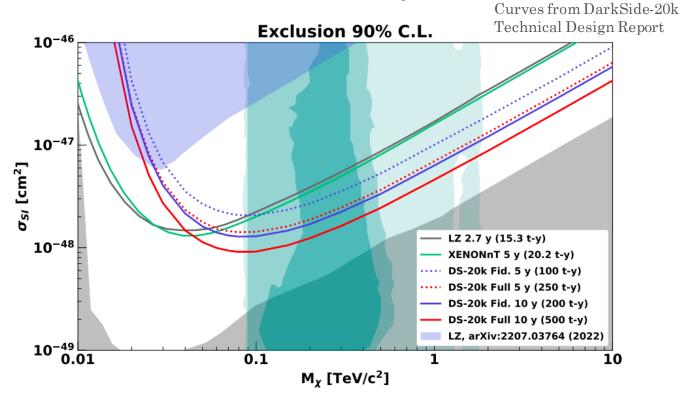
- Reduce radiogenic neutron backgrounds by tagging them in the veto volume (between TPC and stainless steel vessel)
- Neutron veto : UAr volume instrumented with 5m² of vPDUs (single phase)





Gd-loaded acrylic layer (10cm thick) between the TPC and veto is used to capture neutrons and release high energy gammas (shower up to 8 MeV)

DarkSide-20k sensitivity



- DarkSide-20k expected to be the most sensitive dark matter search in the next decade
- Projected sensitivity based on background free search (using fiducial volume), and using the full active volume, compared to xenon-based detectors
- Neutrino backgrounds become important: 3.2 events in 200 tyr
- DS20k will additionally be able to carry out S2 only searches to reach much lower masses than shown here (as seen from DarkSide-50)

Conclusions

- DarkSide-50 was able to carry out a background free DM search, setting an upper limit on the dak matter cross-section of 1e-44cm² at 100GeV
- DarkSide-50 can set world-leading exclusion limits in mass region down to 0.04 GeV using S2 only analysis and the Migdal effect
- DarkSide-20k R&D complete, and construction phase is starting, including SiPM production, construction of cryostat, UAr production
- Data taking is expected from 2026
- Expected sensitivity at 100 GeV: ~1e-48 cm², approaching the neutrino floor.
- We have highlighted some of the design aspects which contribute to DarkSide-20k sensitivity, including the muon and neutron veto system, SiPM photosensors, underground argon

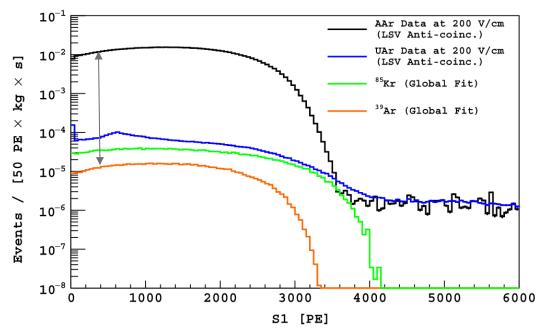
Backup slides



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Underground argon

Phys. Rev. D 93, 081101 (2016)



39Ar is produced by cosmic rays in the atmosphere -> beta-decay

- Activity in atmospheric argon: ~1 Bq/kg
- Activity in underground argon: 0.73 ± 0.10 mBq/kg

DarkSide-50 demonstrated 1400x reduction in 39Ar backgrounds using underground argon

Underground argon for DS20k

1.

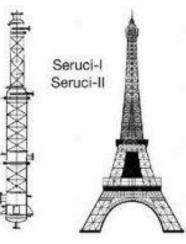
- Underground argon is extracted from industrial CO2 wells in Colorado, USA
- Extraction plant • ready to ship
- Extract 330 • kg/day
- Purity is 99.99%

3. Qualification: DArT

- Single phase LAr detector which can measure the depletion of 39Ar • in Canfranc, Spain
- 1.42 kg of UAr, installed inside the ArDM detector •

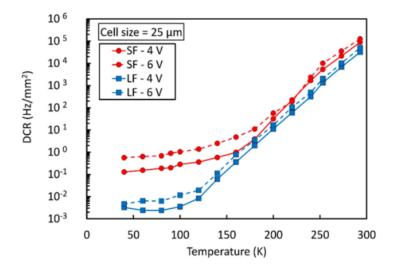
Extraction of argon: Urania 2. Purification of argon: Aria

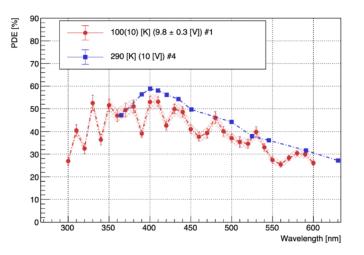
- Argon is transported from Colorado to Sardinia, Italy
- At the coal mine, a distillation column purifies the argon at a rate of 1t/day
- Reduction of 39Ar • by a factor of 10
- Chemical and • isotopic purification

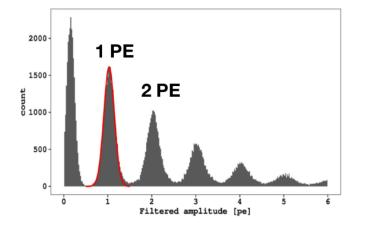


Ellen Sandford, DarkSide status and recent results, Lake Louise Winter Institute 2023

SiPM sensor testing





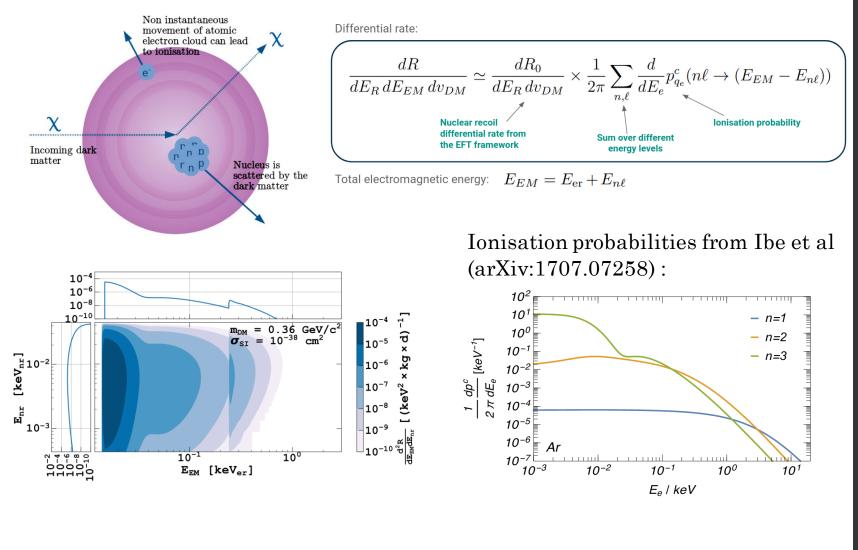


parameter	spec required	spec achieved
PDE @ 420 nm	> 40%	> 42%
DCR (87 K)	250 Hz / tile	~ 20 Hz / tile
correlated noise probabilities (afterpulses, cross talk)	< 50% + 50%	<10% + 35%
SiPM gain	> 1E6	> 1E6
SNR after ARMA filter	> 8	> 15
time resolution	~ 10 ns	~15 ns

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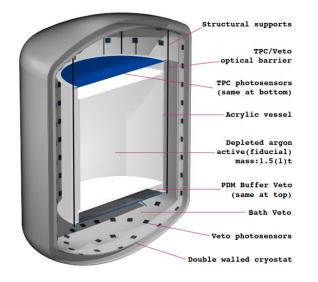
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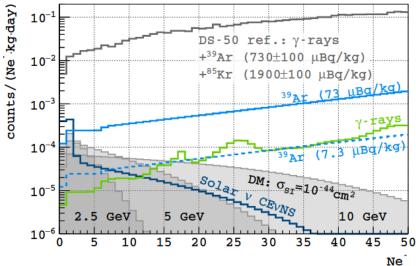
Migdal effect



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DarkSide-LowMass





- Proposed small (1 tonne scale) S2-only dedicated detector
- Additional depletion of UAr to reduce 39Ar background
- R&D to reduce spurious • electron backgrounds
- Expected to reach neutrino floor within 1 year of running
- Aim to use a 2e- threshold

DarkSide-LowMass

