

The nEXO Experiment

Searching for Neutrinoless Double Beta Decay in ^{136}Xe

G Adhikari et al. (nEXO Collaboration), 2022 J. Phys. G: Nucl. Part. Phys. 49 015104

Soud Al Kharusi, McGill University
(on behalf of the nEXO Collaboration)

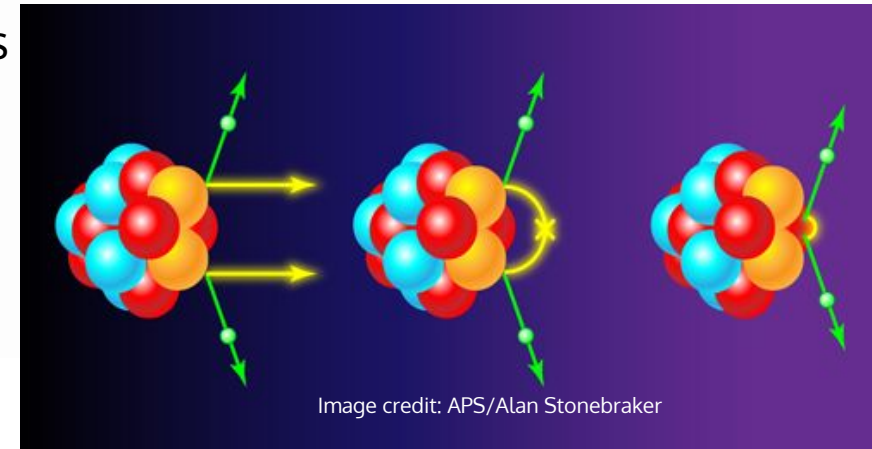
Lake Louise Winter Institute

Feb. 22nd 2023

soud.alkharusi@mail.mcgill.ca

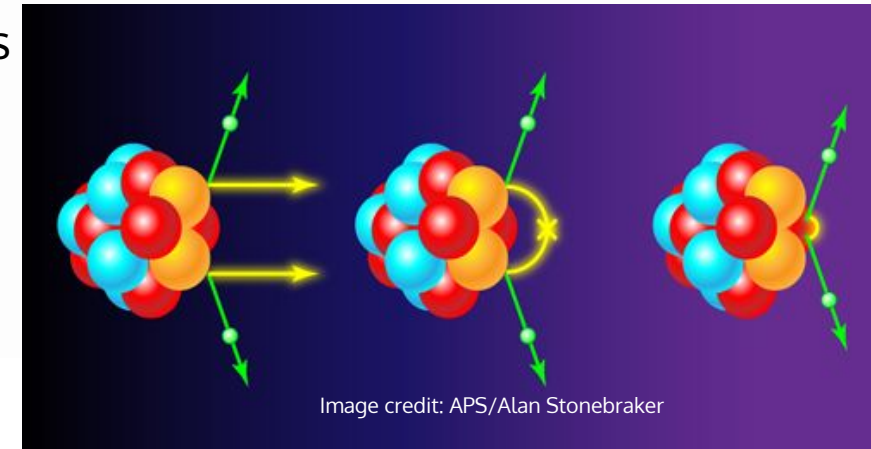
Why Search for $0\nu\beta\beta$?

- The discovery of neutrino mass from oscillation experiments provides **new pathways to mass generation** in the neutrino sector
 - Dirac vs Majorana masses
 - feeble couplings to Higgs field vs seesaw mechanisms
- Neutrinoless double beta decay ($0\nu\beta\beta$) **exploits the nucleus as a virtual environment to probe high energy physics processes**
- Implications for **matter-antimatter asymmetry problem**



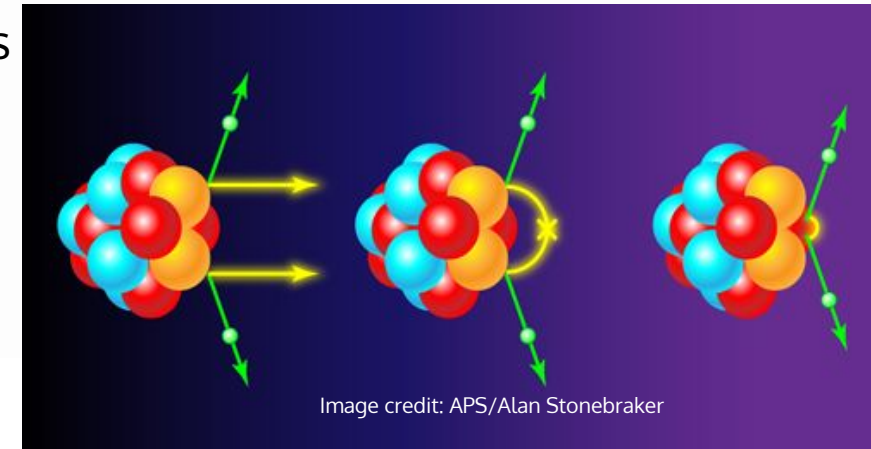
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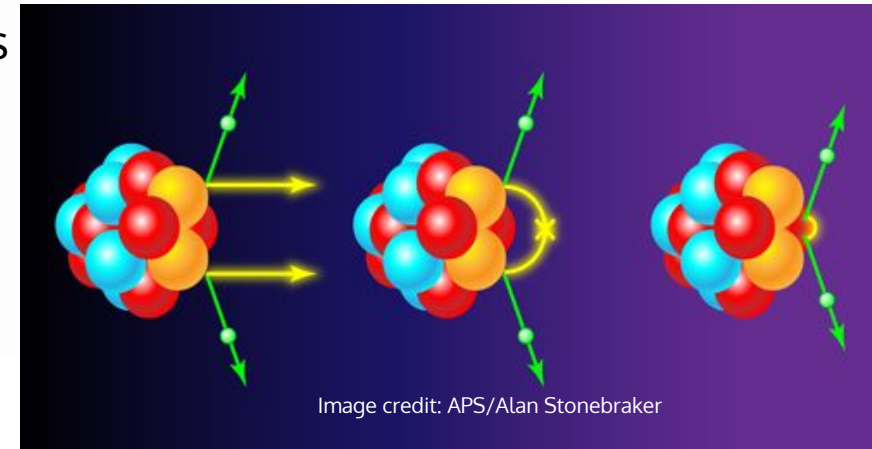
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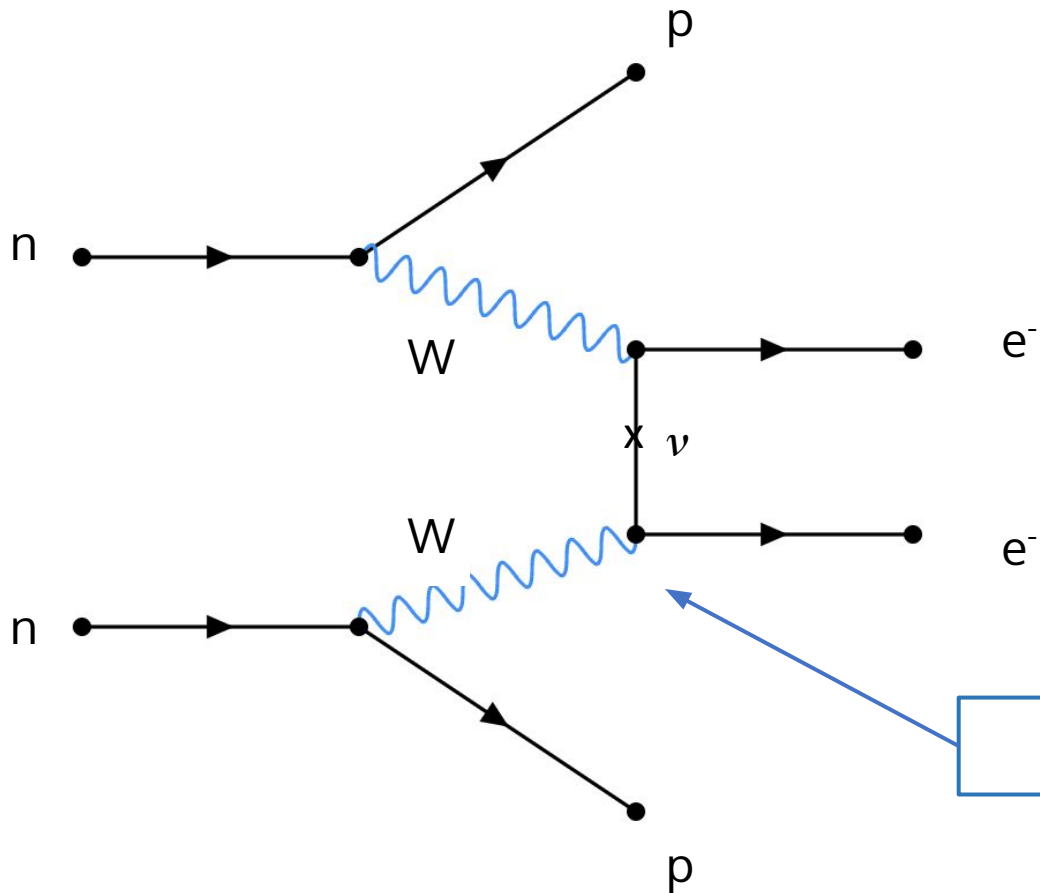
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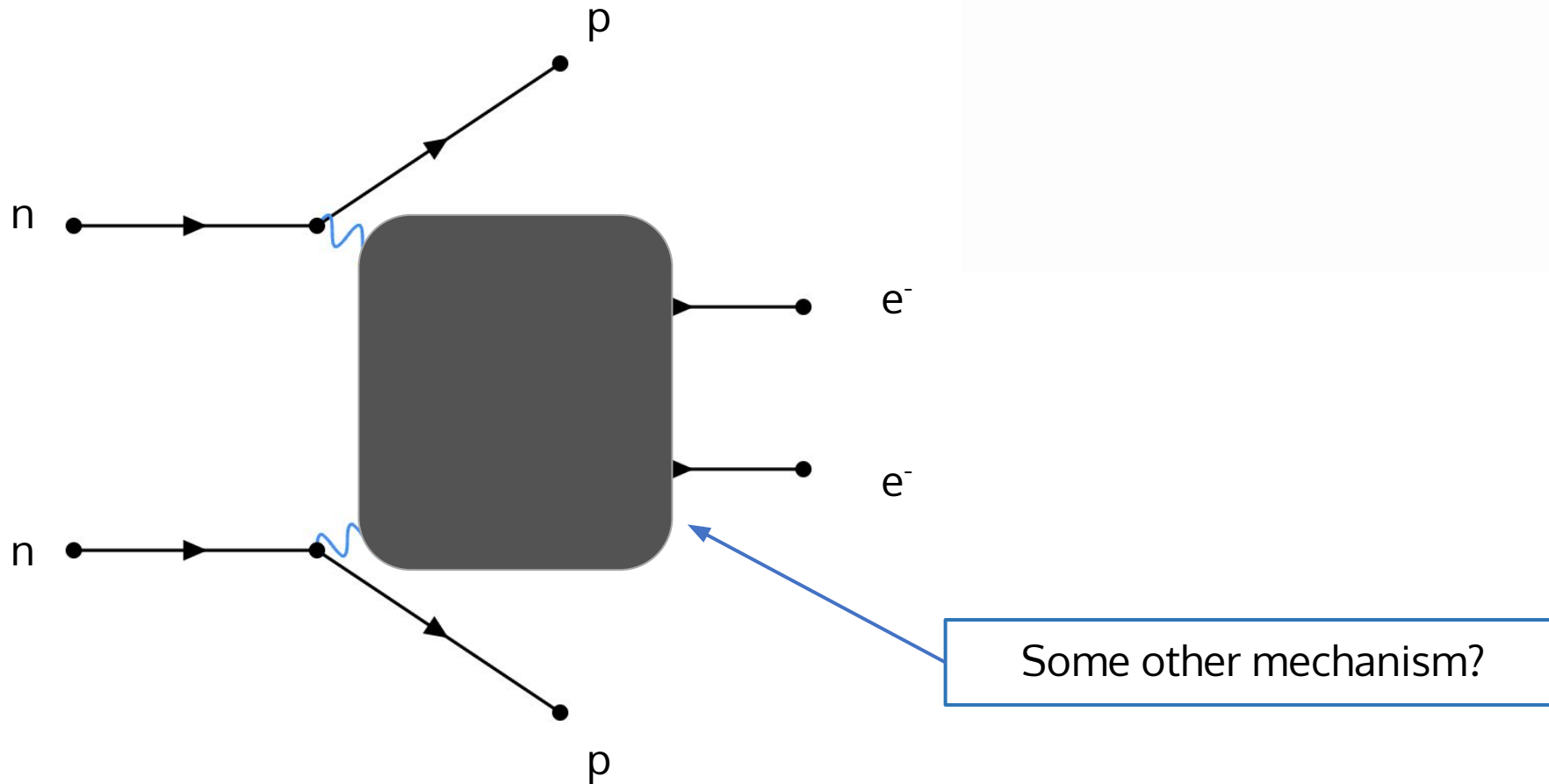
Searches for neutrinoless double beta decay ($0\nu\beta\beta$) are searches for Lepton Number Violation & Physics Beyond the Standard Model

How would $0\nu\beta\beta$ even work?



Light Majorana neutrino exchange?

How would $0\nu\beta\beta$ even work?



Searching for $0\nu\beta\beta$: The Real Motivation

Particle physics community searching for physics beyond the standard model



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Regardless of what mechanism $0\nu\beta\beta$ proceeds by, it always implies new physics

([Schechter, and Valle. Phys. Rev. D 25.11 \(1982\): 2951.](#) "black box theorem")

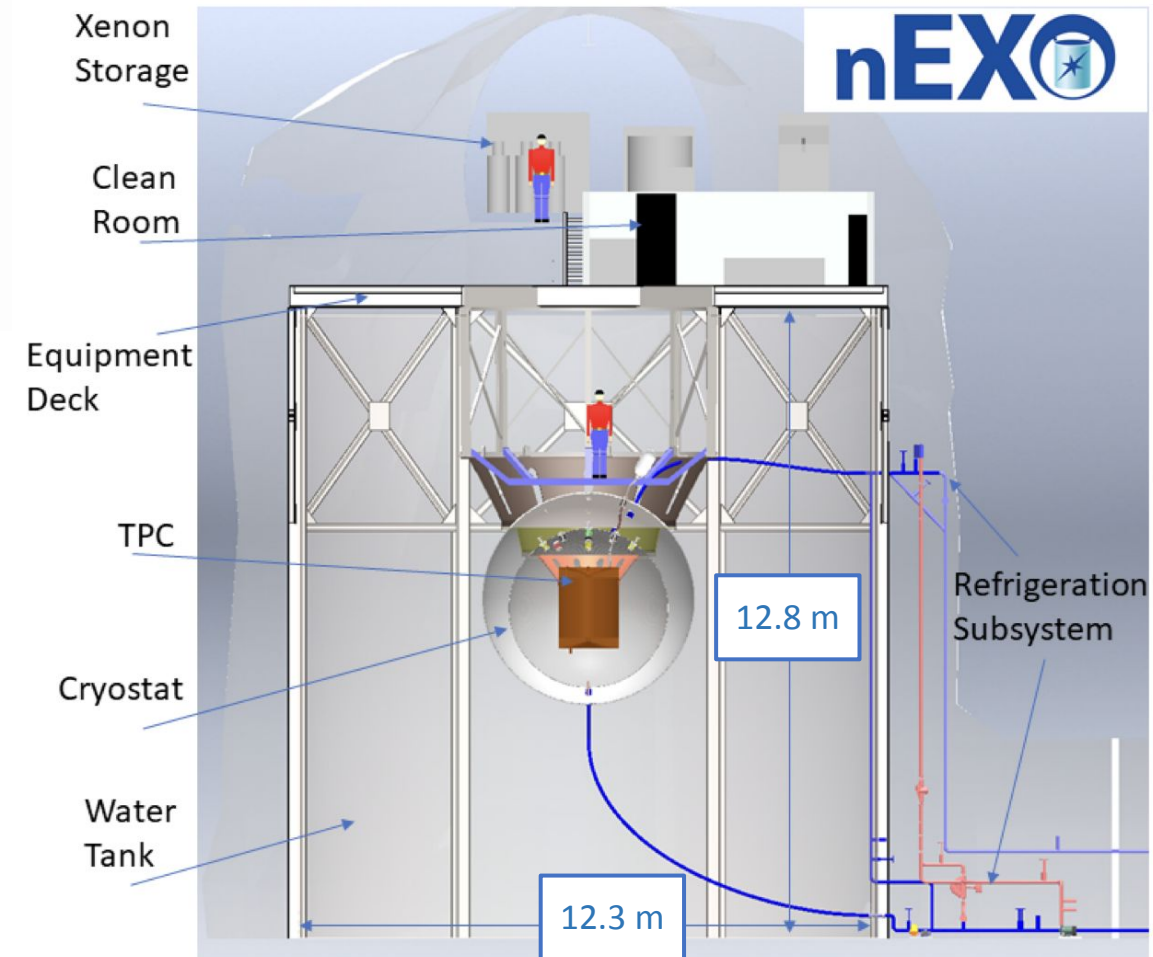
What is nEXO?

nEXO is a **proposed experiment searching for $0\nu\beta\beta$** , following successes of EXO-200

currently in the conceptual design stage:

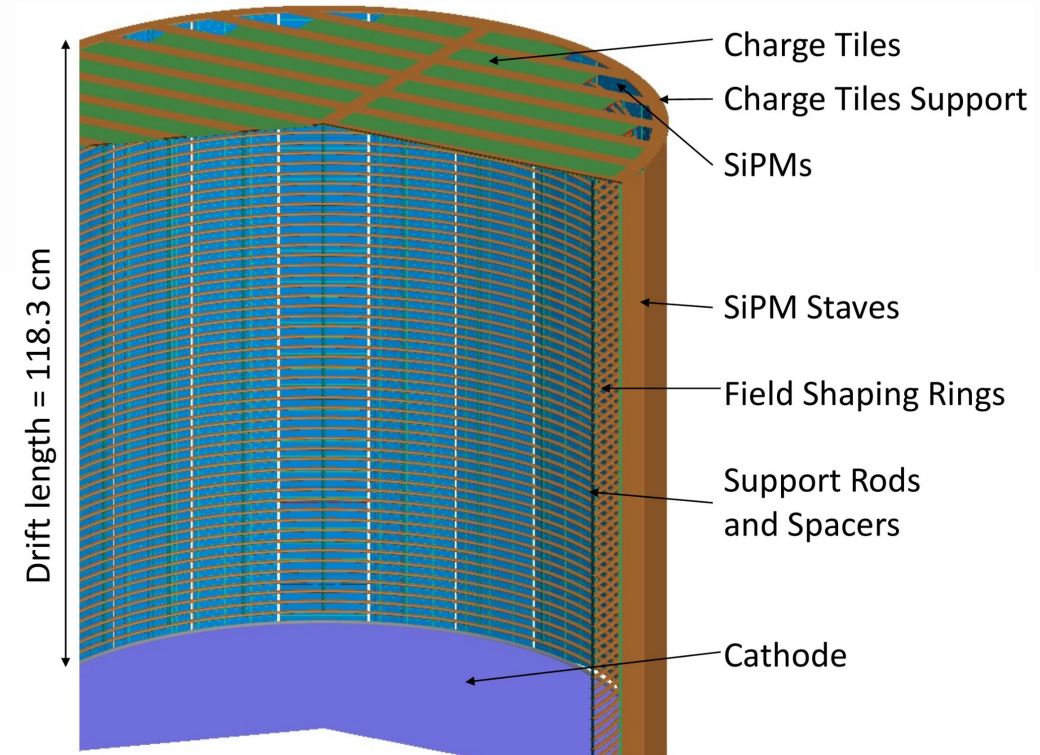
funding for the nEXO project from U.S. DoE has started to flow!

- 5-tonne single-phase liquid xenon Time Projection Chamber (LXe TPC)
- LXe is **enriched to 90% in the target isotope, ^{136}Xe**
- Extensive radio-assay program
 - ultra low backgrounds **validated by EXO-200 data**



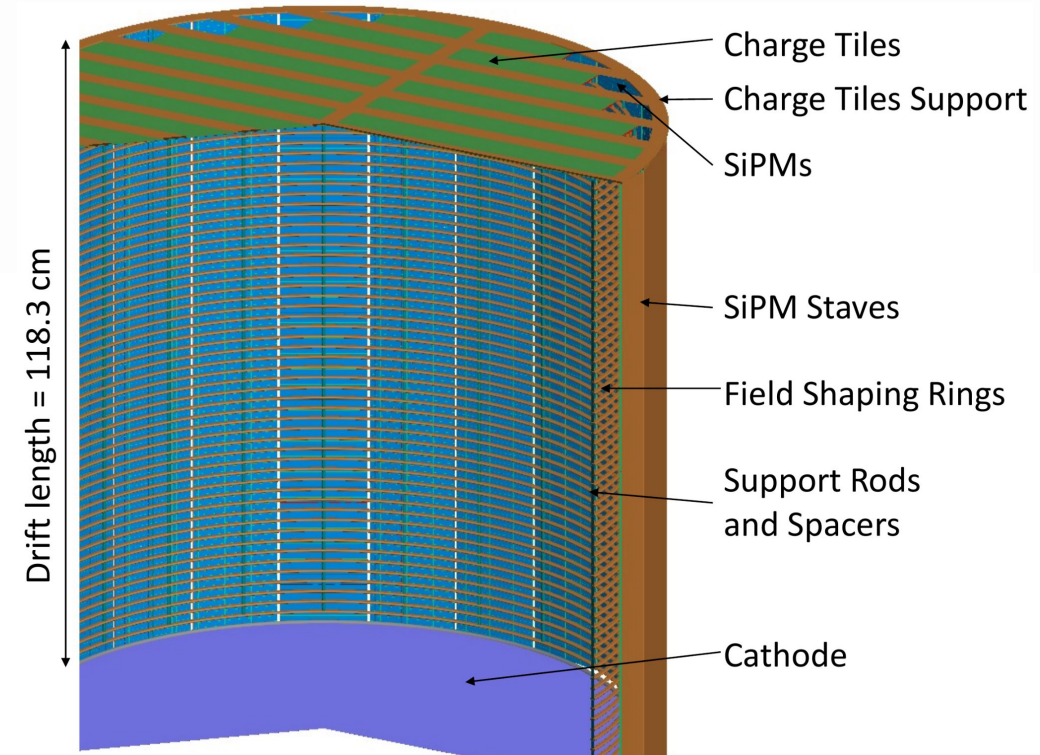
nEXO: Distinguishing Features

- Homogeneous, dense, liquid detector medium with high-Z nucleus
 - online purification
 - self-shielding of γ radiation
 - scalability
- Multiparameter Analysis
- Possibility to tag daughter nucleus
- Possibility for control run in case of discovery



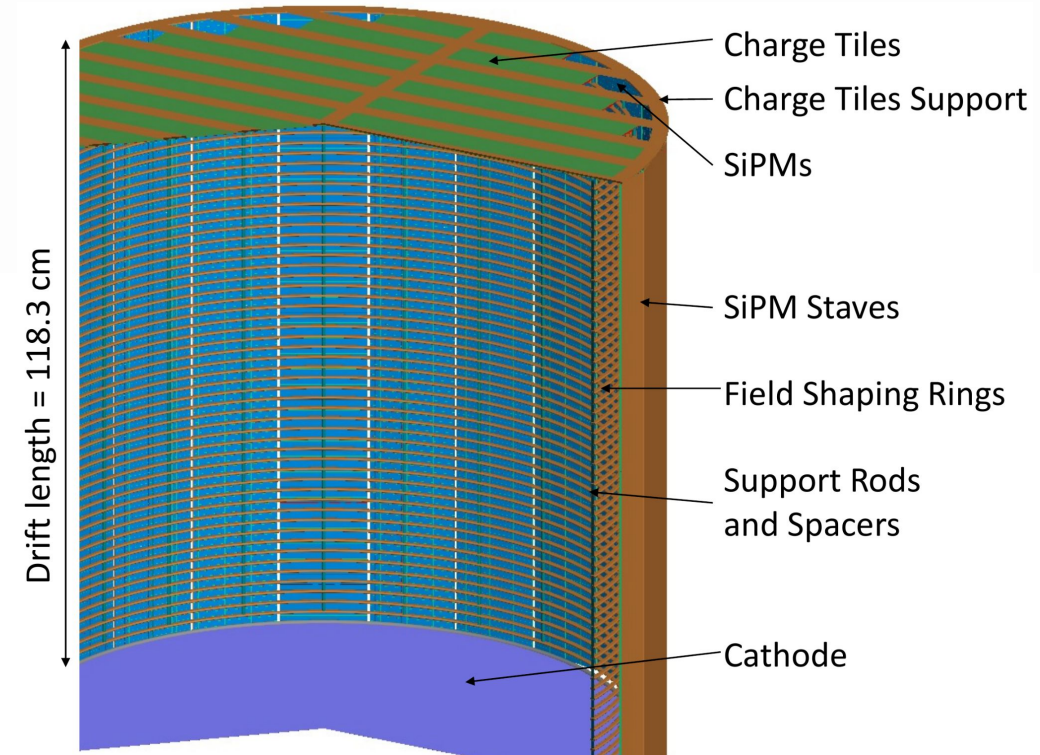
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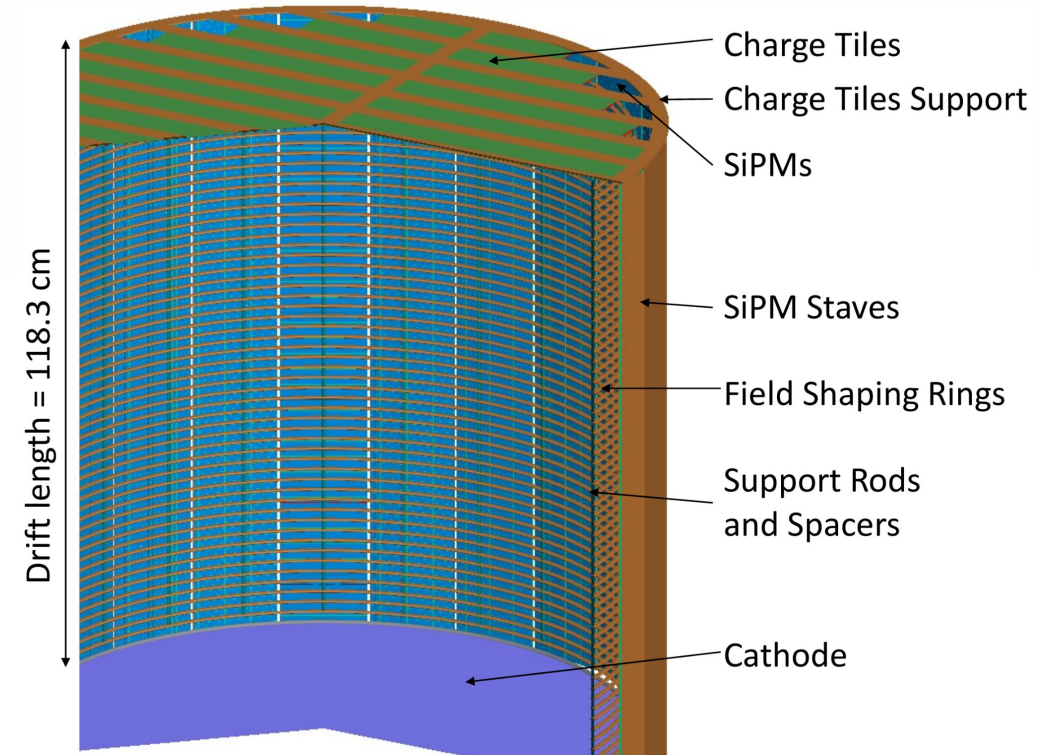
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 - [Nature 569, no. 7755 \(2019\): 203-207](#)
- Possibility for control run in case of discovery



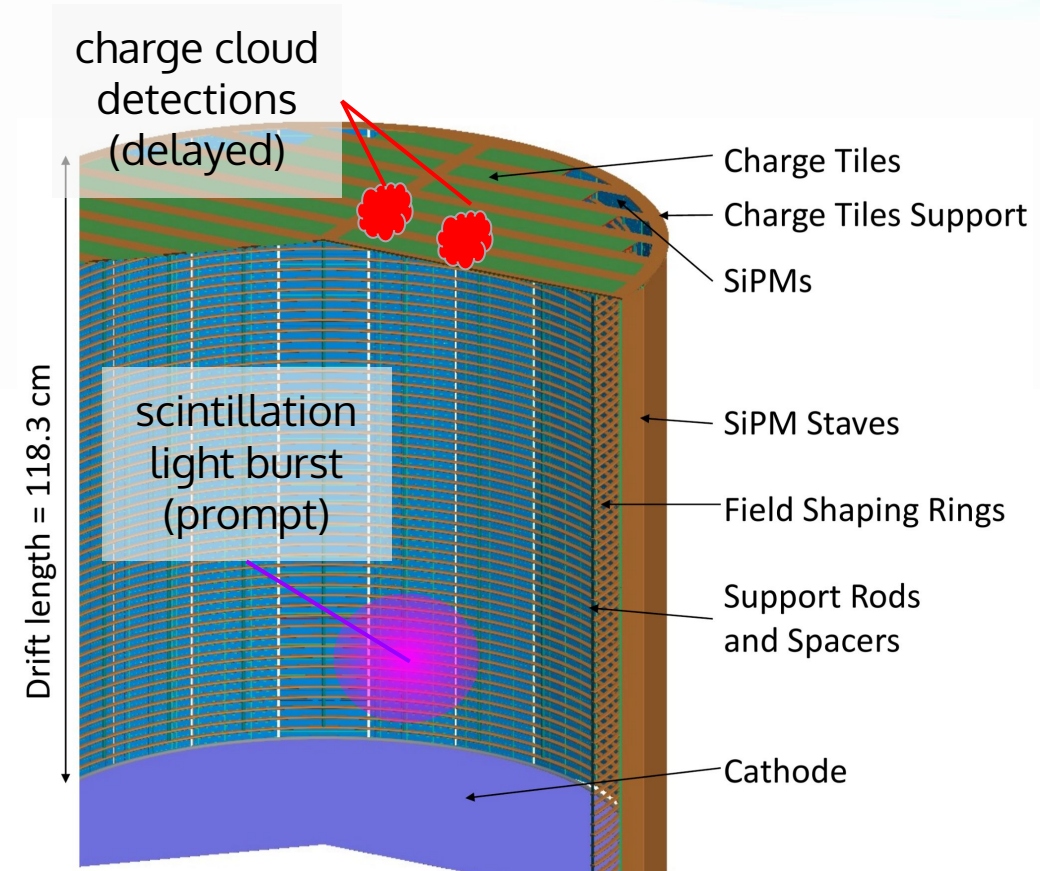
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- Possibility for control run in case of discovery
 - use unenriched xenon & repeat the experiment!



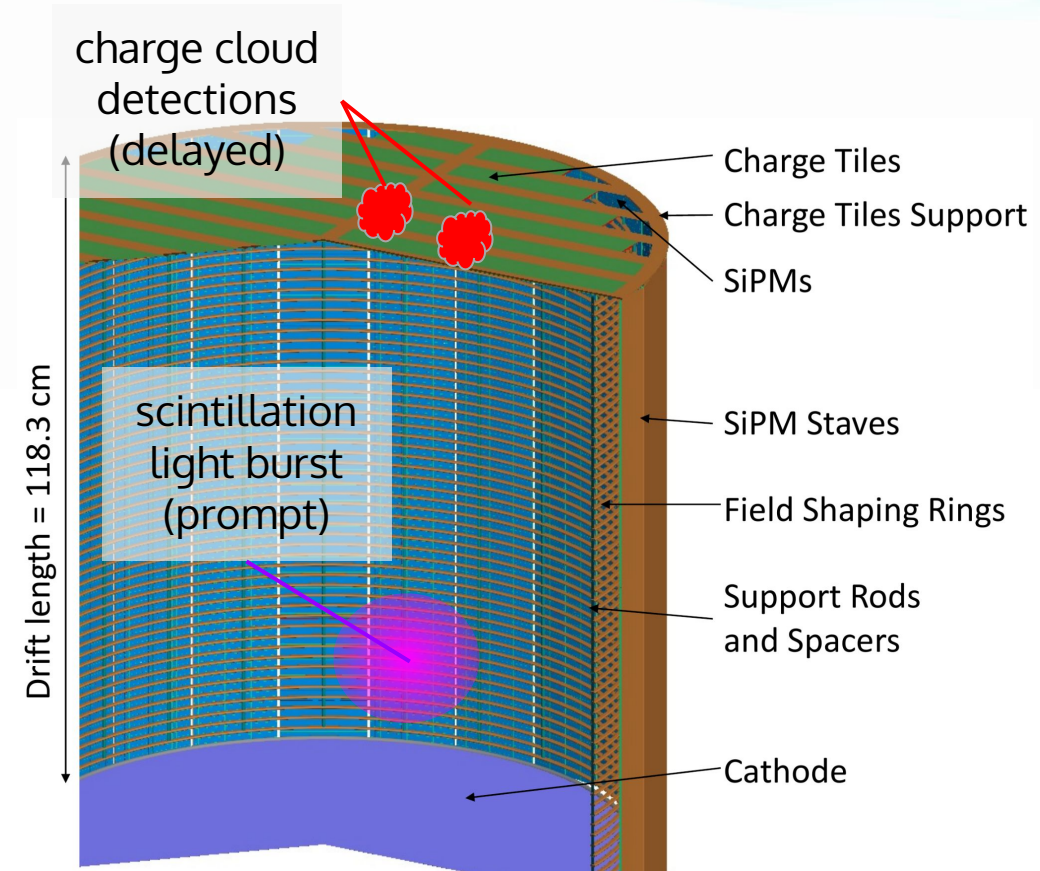
How does nEXO work?

- Energy deposits in the LXe liberate electrons, ionize the surrounding liquid
 - excited dimers of Xe release ~175 nm scintillation light
 - ionization clouds drift to segmented anode in applied E-field
- Combination of light + charge readout gives us...



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- Combination of light + charge readout gives us:
 - Better energy resolution
 - Better spatial positioning (localization)
 - Topological discriminator between α , β and γ events

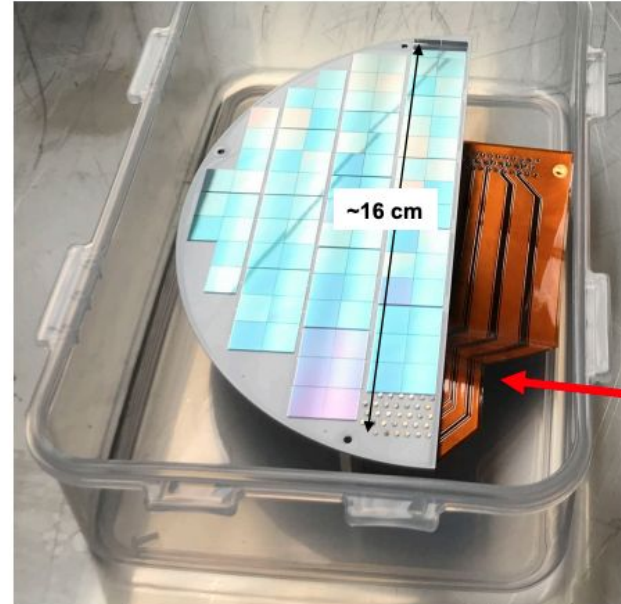


Hardware?

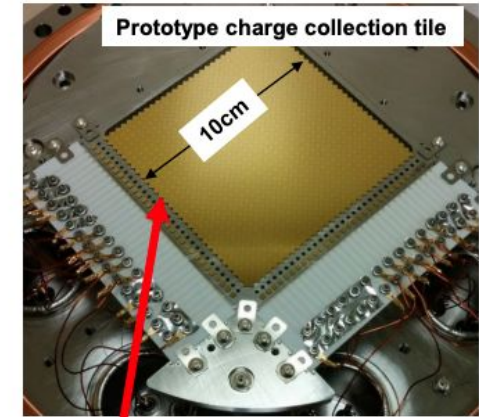
An Active R&D Program

- Basic principles and backgrounds validated and measured by EXO-200
- Technological upgrades and further background reductions give nEXO a $>250x$ projected improvement over EXO-200 sensitivity to $0\nu\beta\beta$ half-life with only $\sim 25x$ increase in mass!

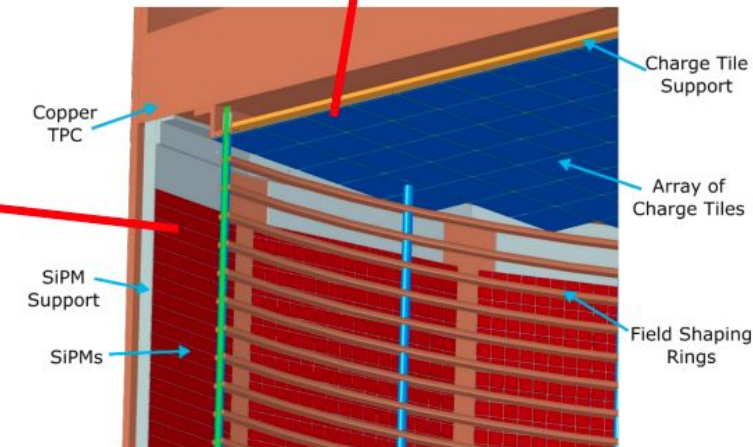
At the core of the TPC are Light and Charge collection devices



Prototype VUV SiPM array (FBK)



Prototype charge collection tile

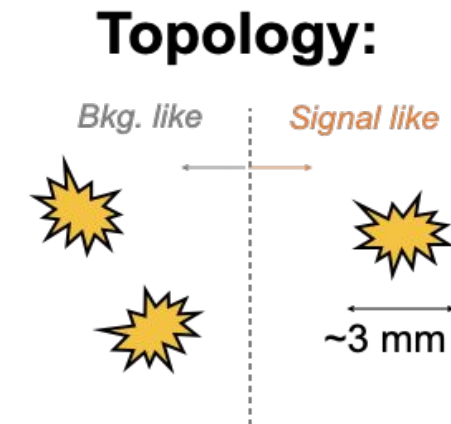
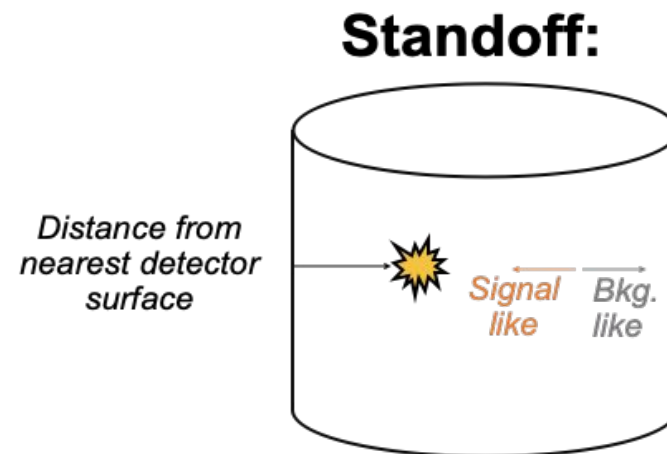
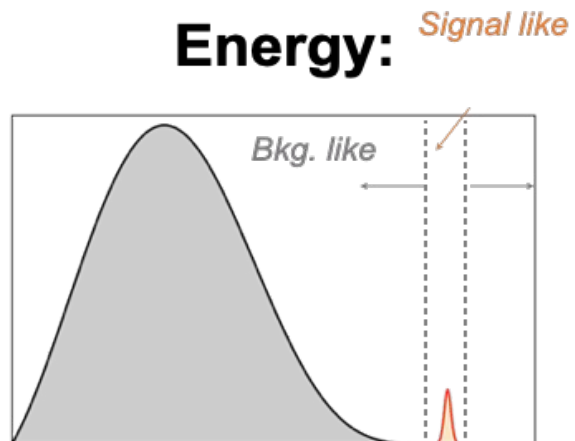


Multiparameter Analysis

nEXO is not a counting experiment

Three high-level variables:

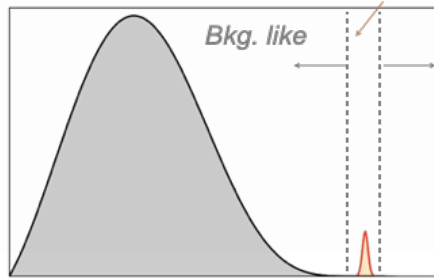
- ~1% **Energy** resolution at $Q_{\beta\beta}$
- **Standoff** distance to detector components (precise event localization, depth in xenon)
- **Topology** score (DNN): single- and multi-site discrimination (β -like vs γ -like event separation)



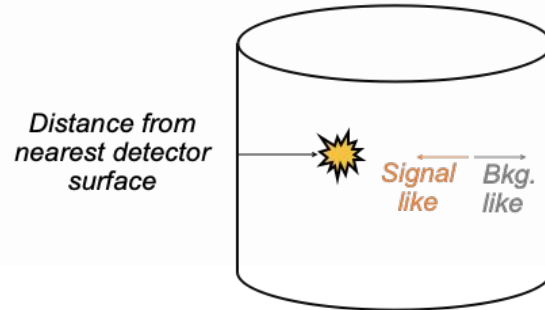
Multiparameter Analysis

A 3D Parameter Space

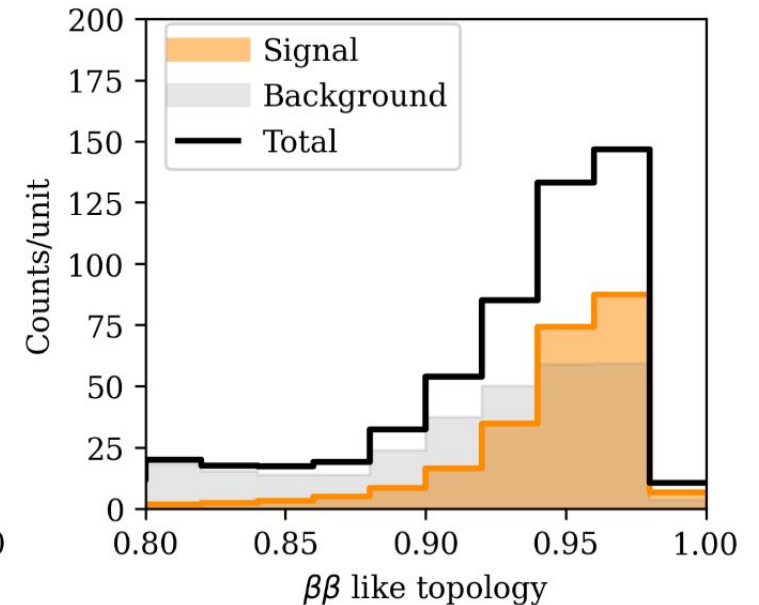
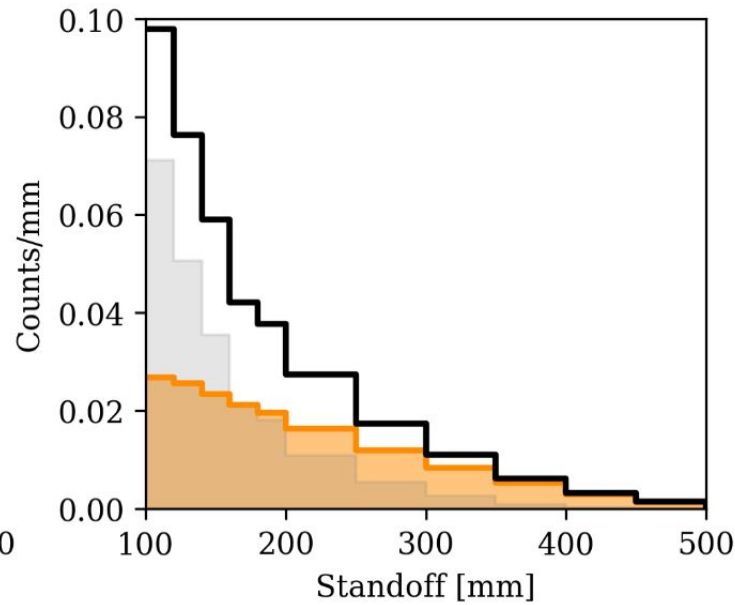
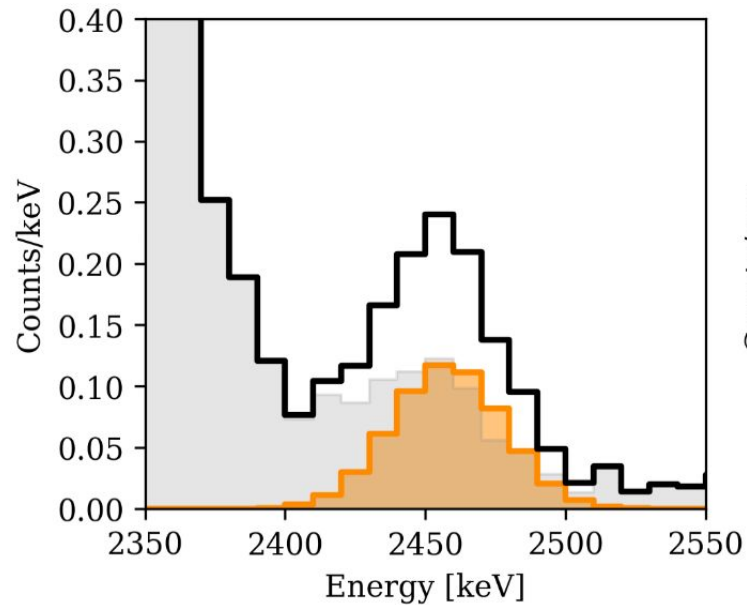
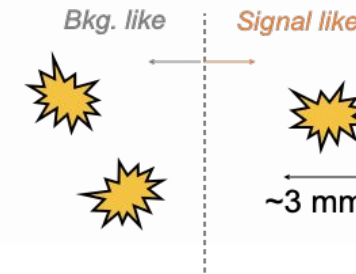
Energy: *Signal like*



Standoff:



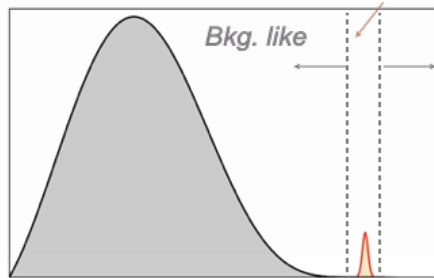
Topology:



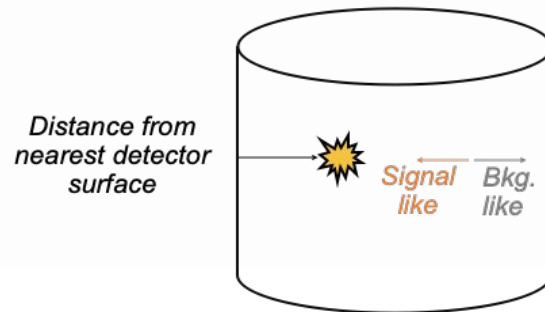
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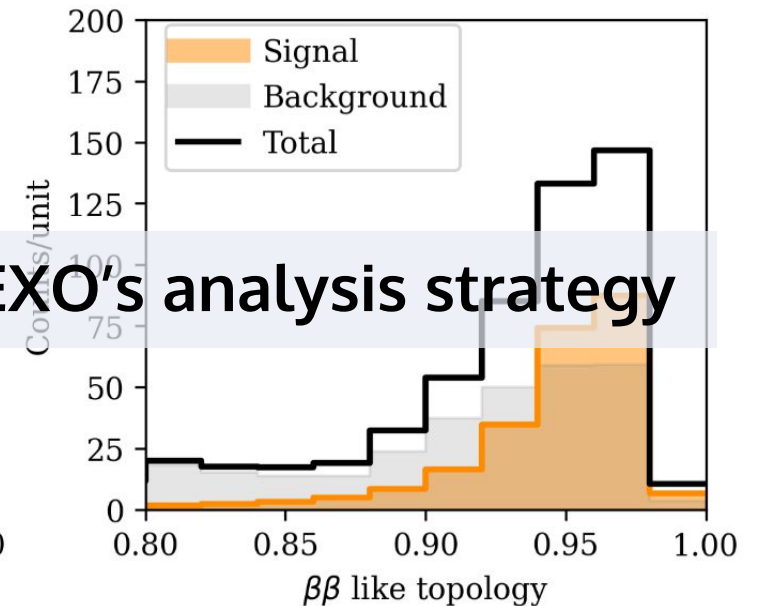
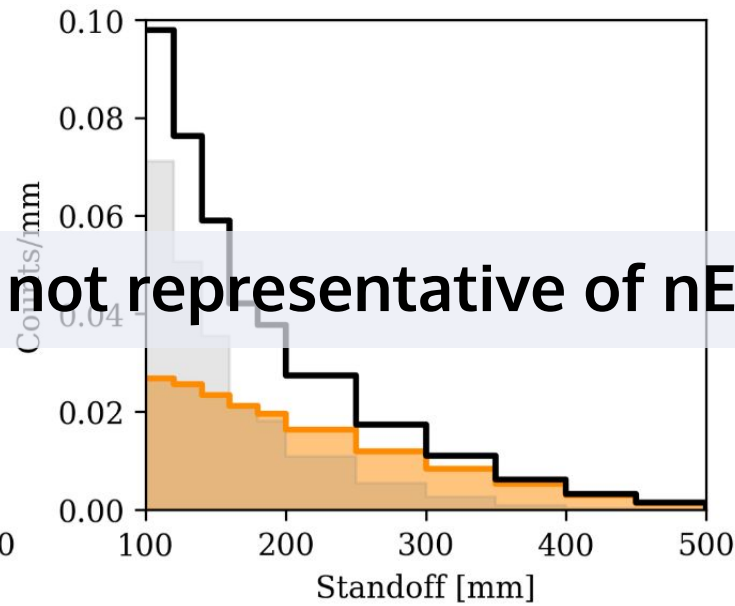
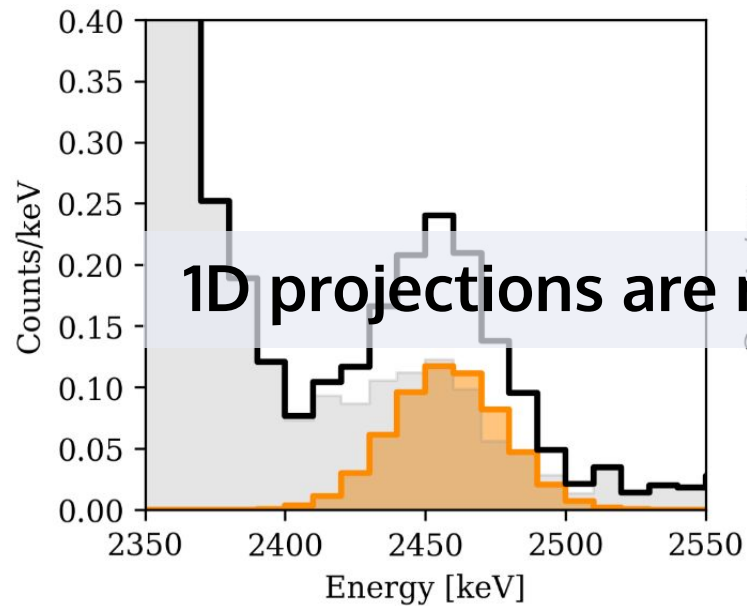
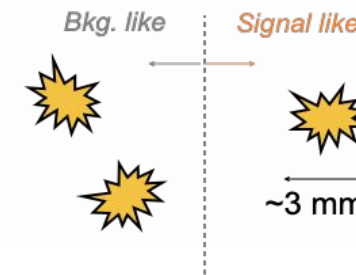
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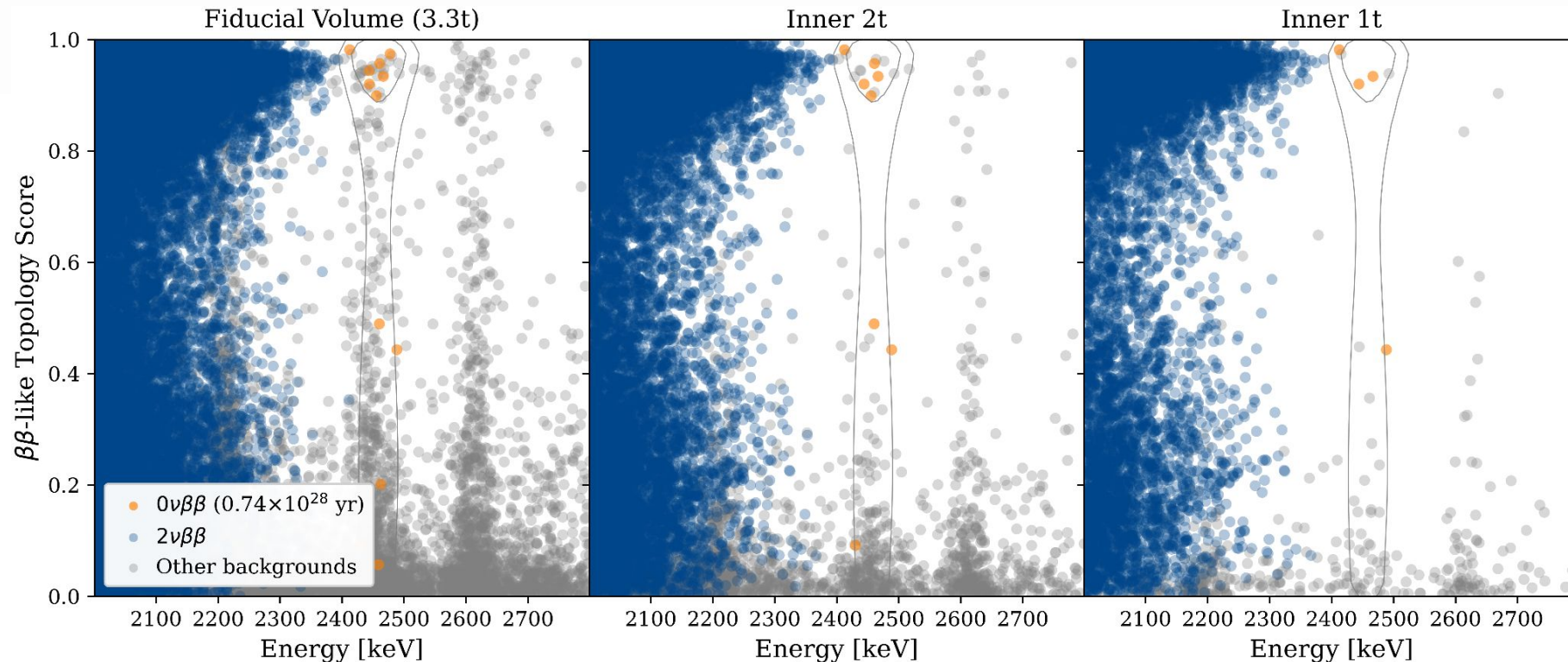


1D projections are not representative of nEXO's analysis strategy

Multiparameter Analysis

What will nEXO data look like?

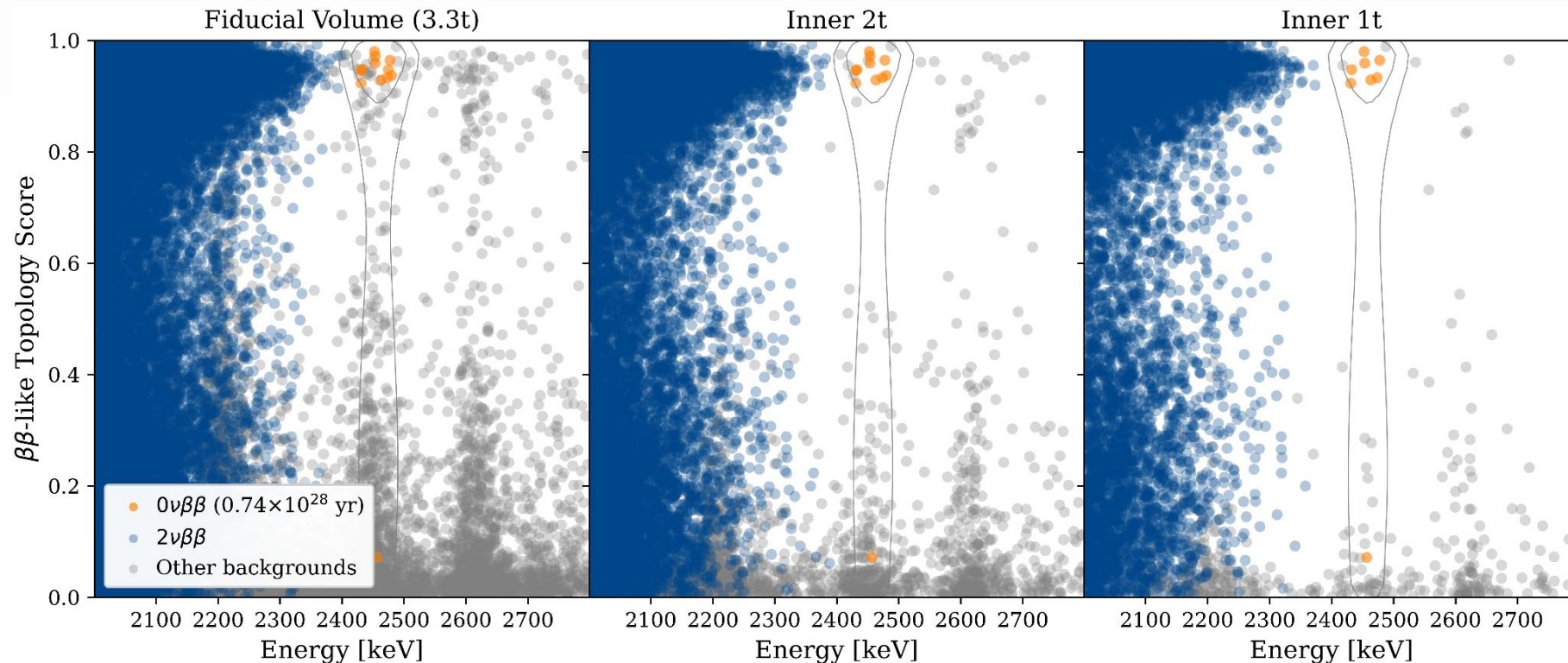
- 1 and 2σ contours on signal ($0\nu\beta\beta$)
- Deeper in the LXe, backgrounds are quieter, signal dominates... but we **use all the LXe in analysis!**
- Below: realizations of nEXO 10 yr dataset at 0.74×10^{28} yr half life for $0\nu\beta\beta$ in ^{136}Xe (3σ discovery)



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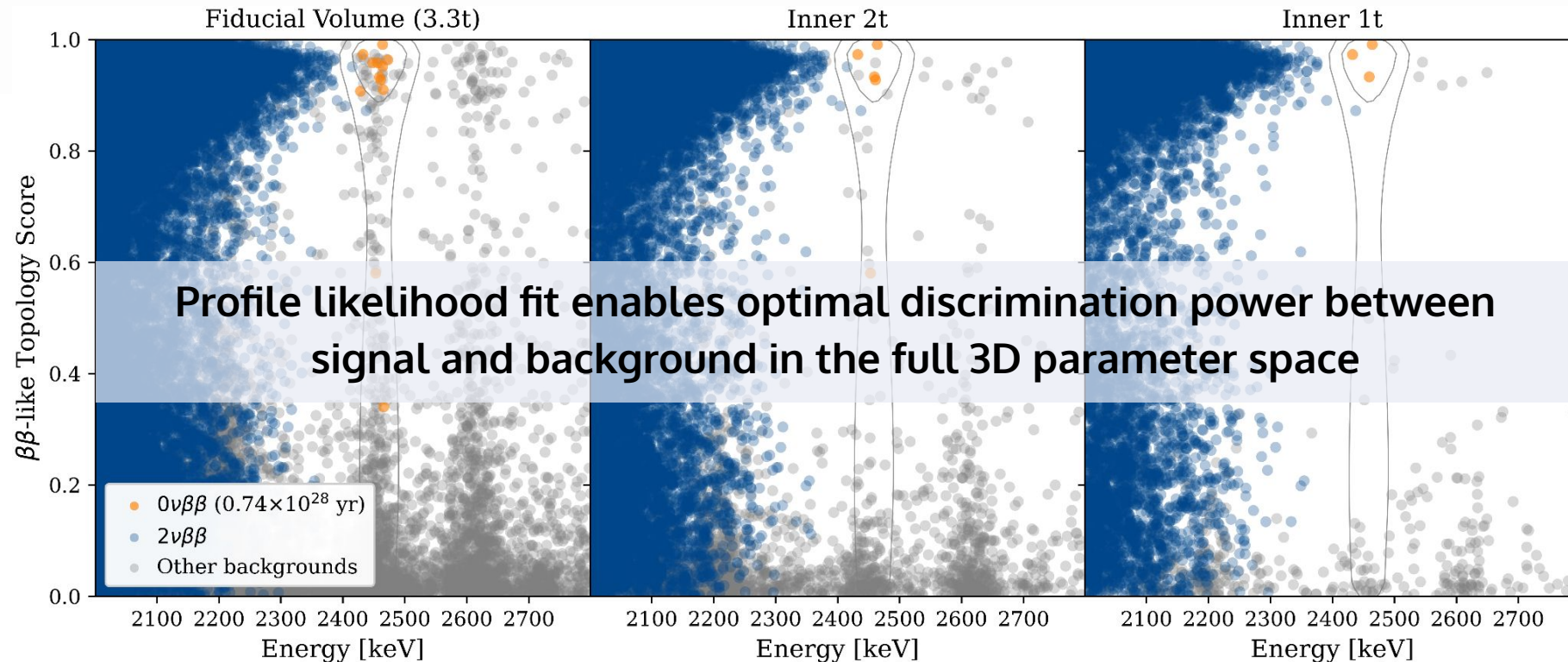
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Multiparameter Analysis

3D profile likelihood fit: ultimate test of $0\nu\beta\beta$ hypothesis

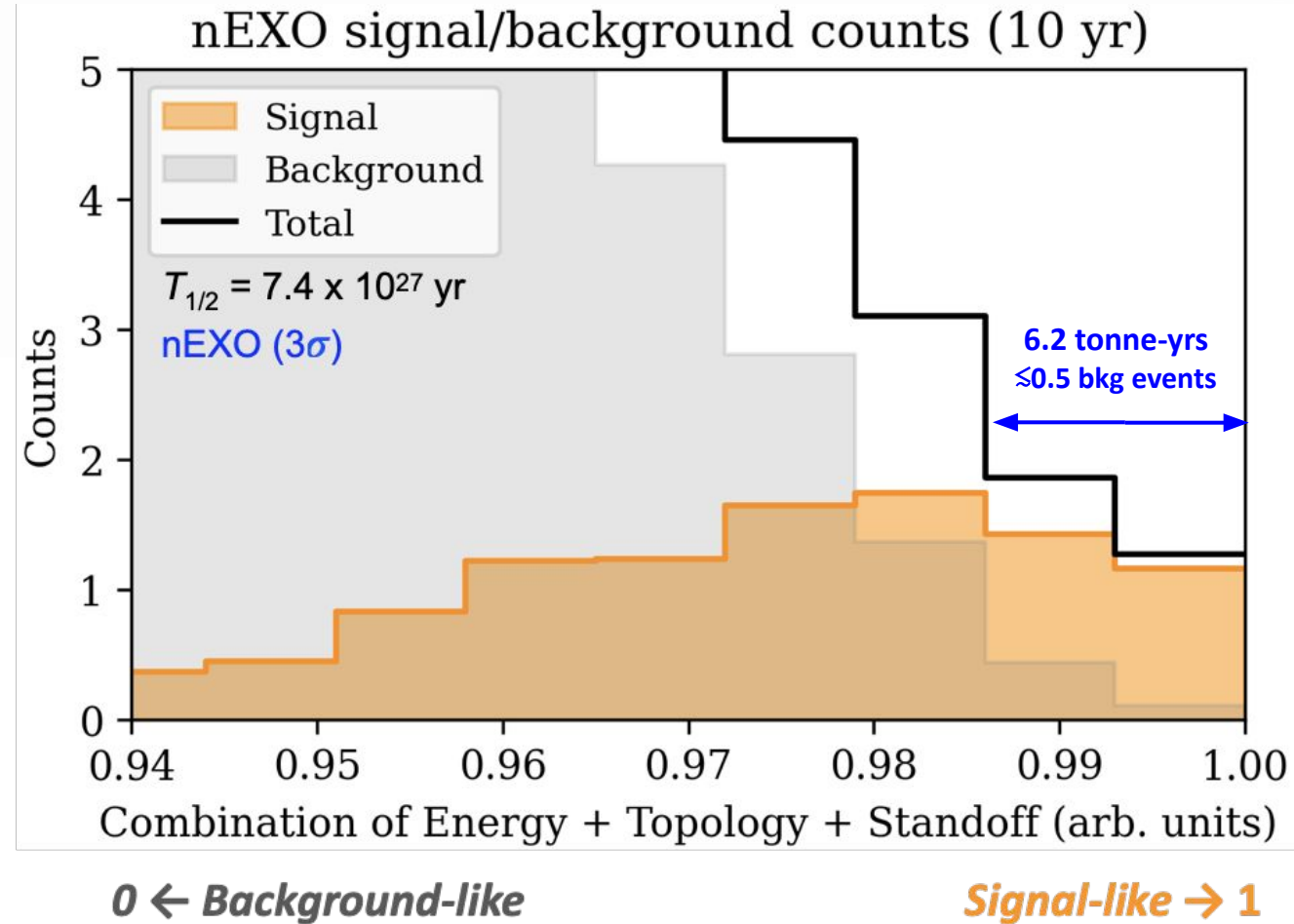
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Multiparameter Analysis

3D → 1D visualization

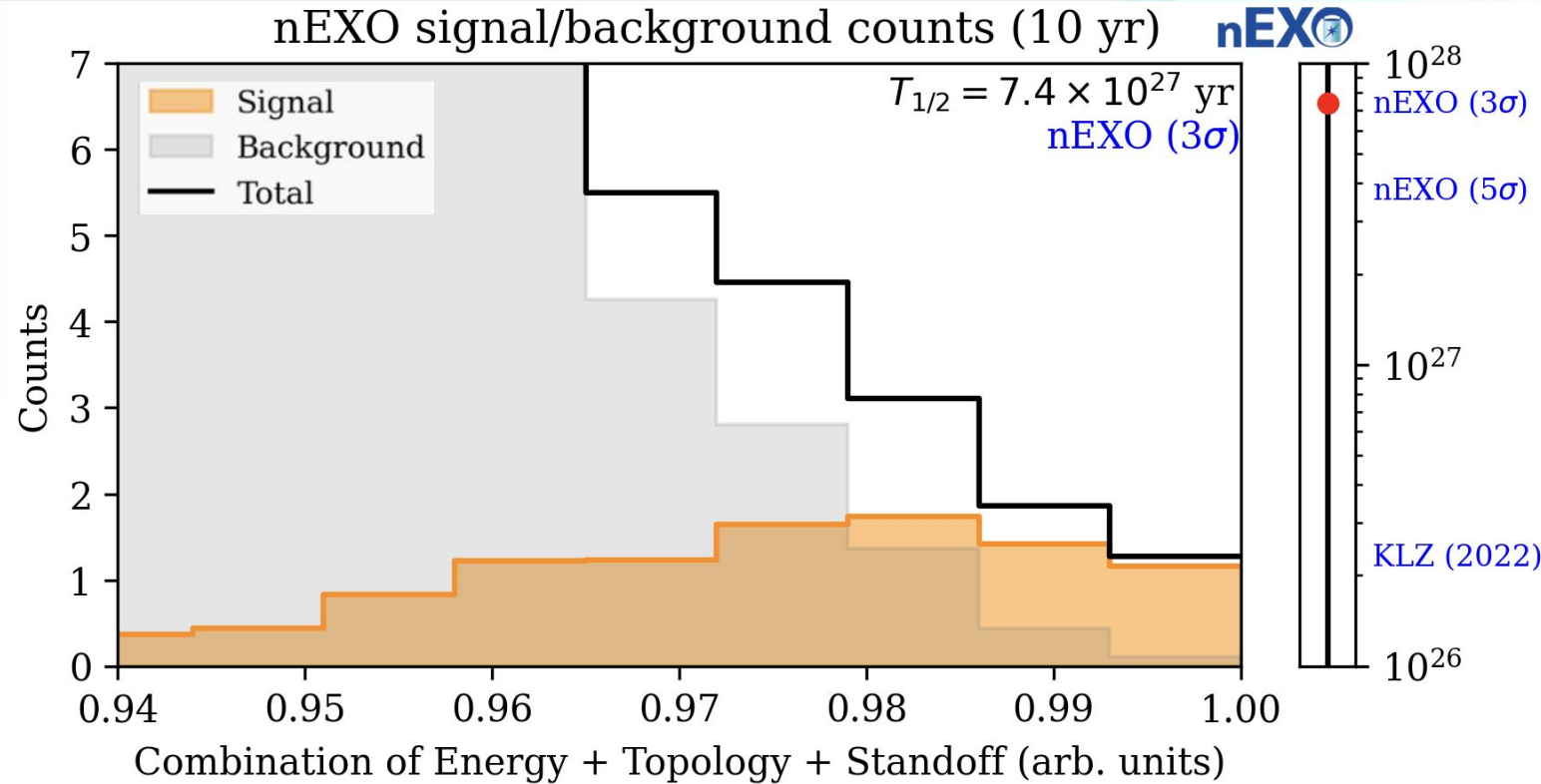
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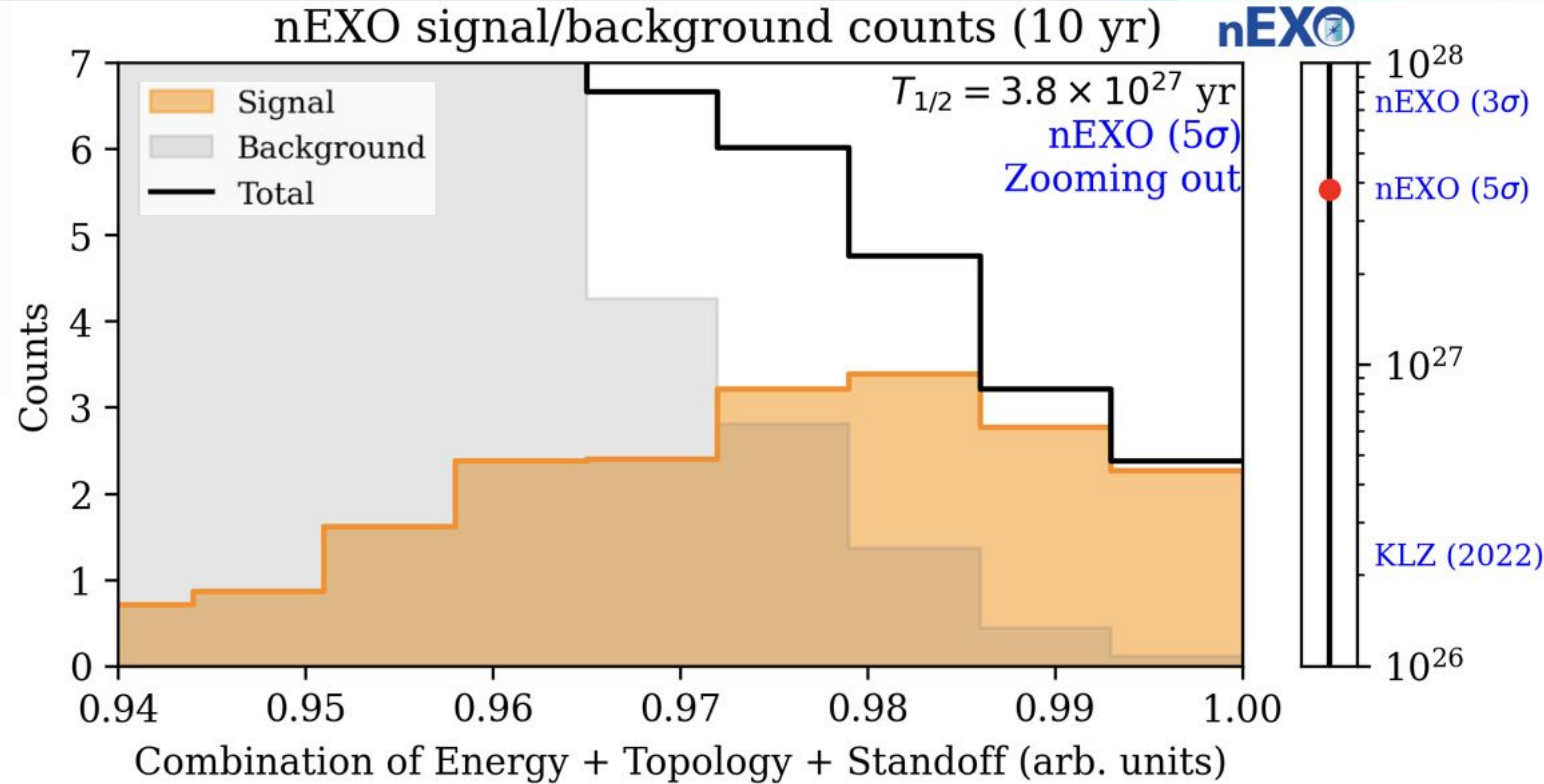
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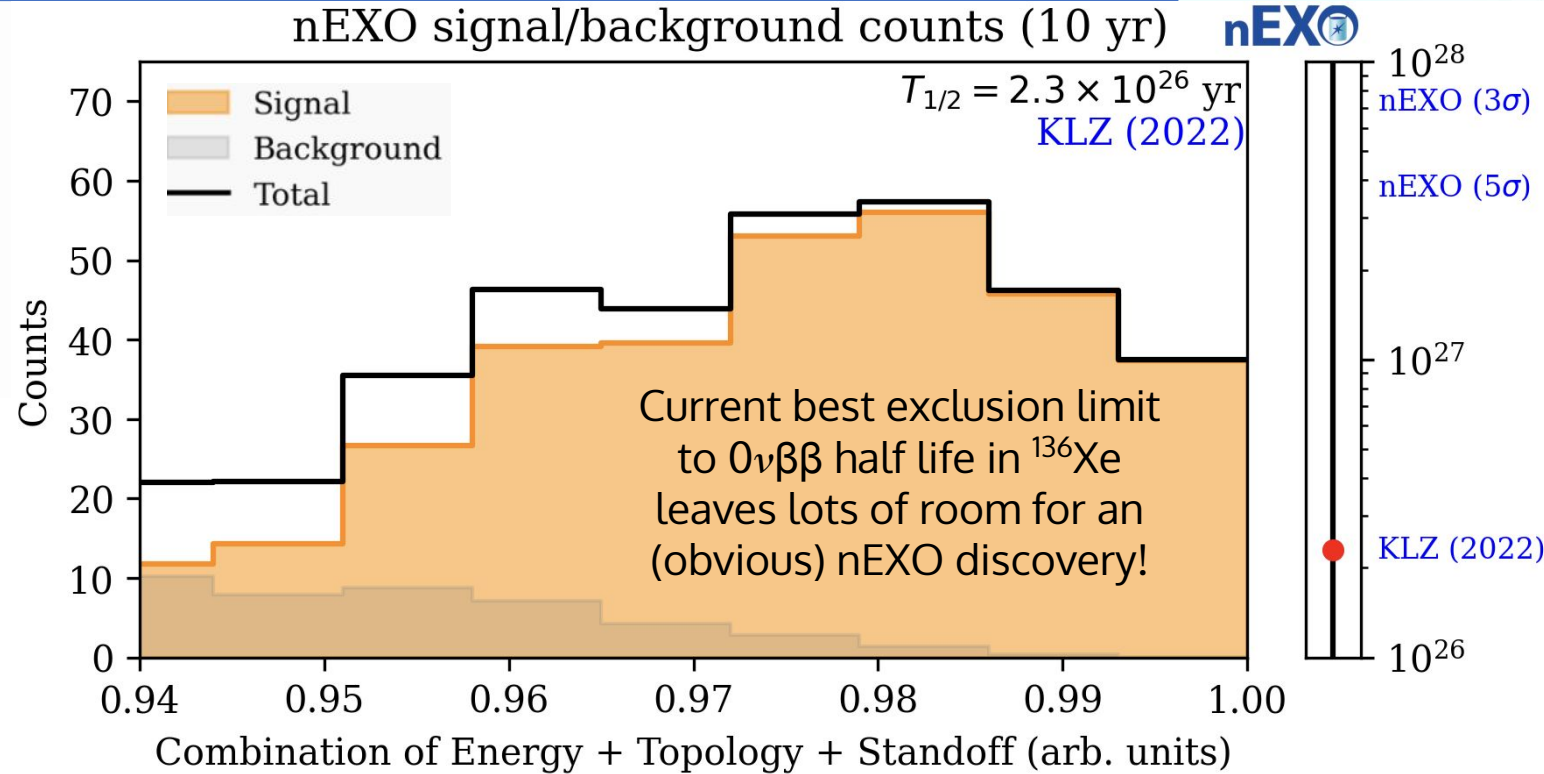
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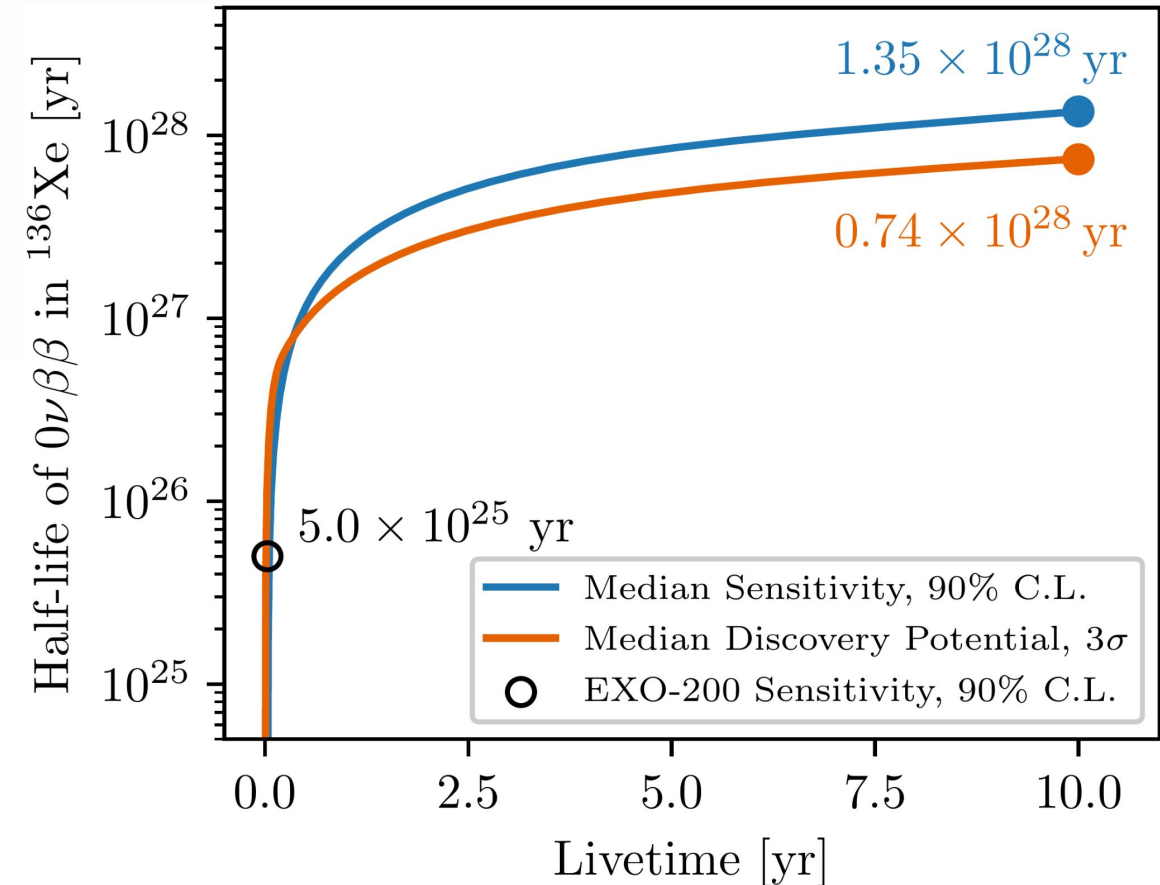
KLZ exclusion limit: <https://doi.org/10.1103/PhysRevLett.130.051801>

nEXO Sensitivity

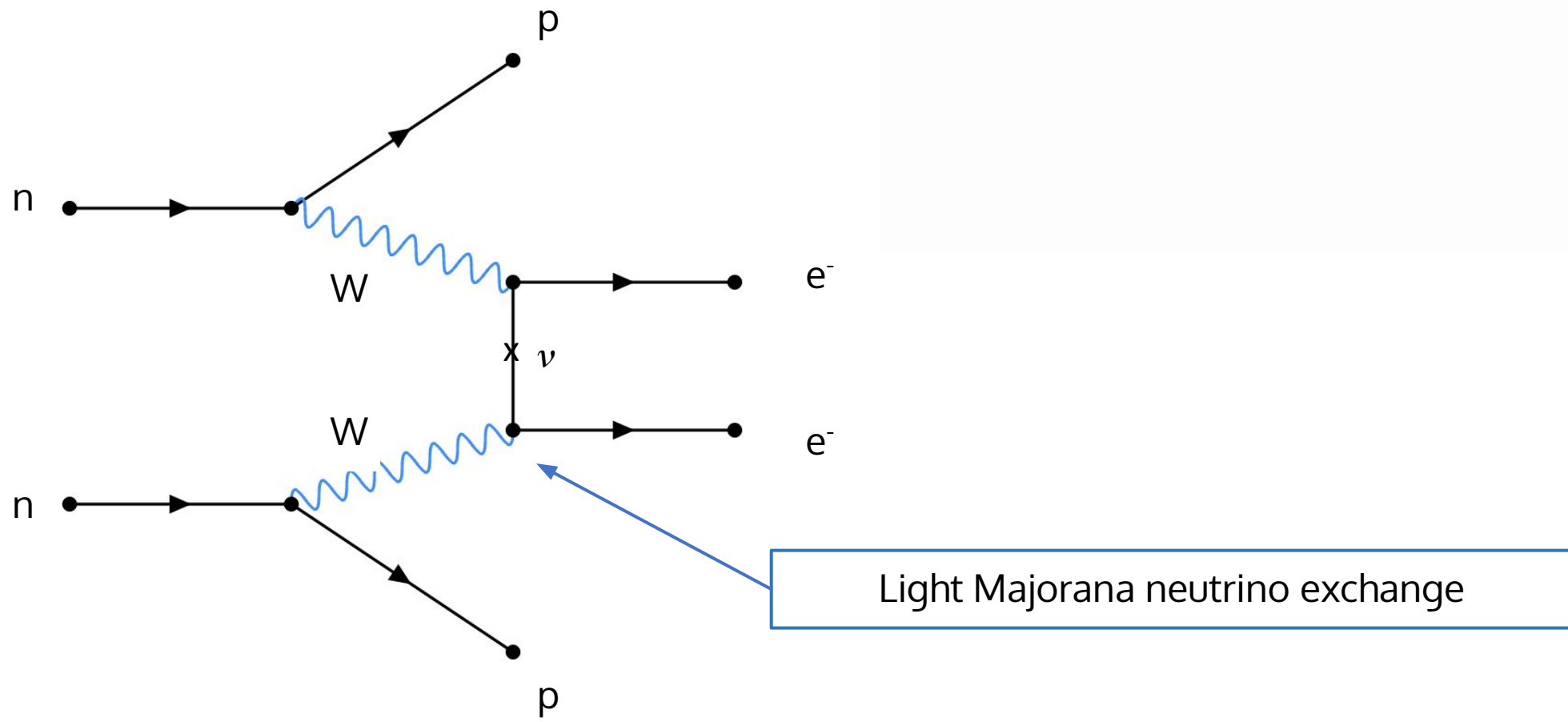


Ultimate Goal of 1.35×10^{28} yr half life

- In 6.5 years of data, nEXO will reach a exclusion sensitivity to $0\nu\beta\beta$ half life in xenon $>10^{28}$ years (90% C.L.)
 - Age of the universe $\times 10^{18}$!



A Neutrino Mass Measurement?



nEXO Sensitivity

Neutrino Mass Measurement

- **Half lives of $0\nu\beta\beta$ correspond to an effective Majorana mass of the electron neutrino $\langle m_{\beta\beta} \rangle$**
 - combination of 3 neutrino mass states
 - **Assumes dominant process for $0\nu\beta\beta$ is light-Majorana neutrino exchange**
- $\langle m_{\beta\beta} \rangle$ is isotope-independent

$$\langle m_{\beta\beta} \rangle = \left| \sum_{i=1}^3 U_{ei}^2 m_i \right|$$

nEXO Sensitivity

Neutrino Mass Measurement

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 - combination of 3 neutrino mass states
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- $\langle m_{\beta\beta} \rangle$ is isotope-independent
 - Depends on your choice nuclear matrix element (NME) when converting from a half life measurement to neutrino mass, **NME is least constrained theoretical parameter below**
 - Complex nuclear/particle physics could change $\langle m_{\beta\beta} \rangle \rightarrow$ **we need to search for $0\nu\beta\beta$ in multiple isotopes**

$$\langle m_{\beta\beta} \rangle = \left| \sum_{i=1}^3 U_{ei}^2 m_i \right|$$

$$\left(T_{1/2}^{0\nu} \right)^{-1} = \frac{\langle m_{\beta\beta} \rangle^2}{m_e^2} G^{0\nu} g_A^4 |M^{0\nu}|^2$$

Phase space factor

Axial coupling, $g_A = 1.27$

Nuclear Matrix Element (NME)

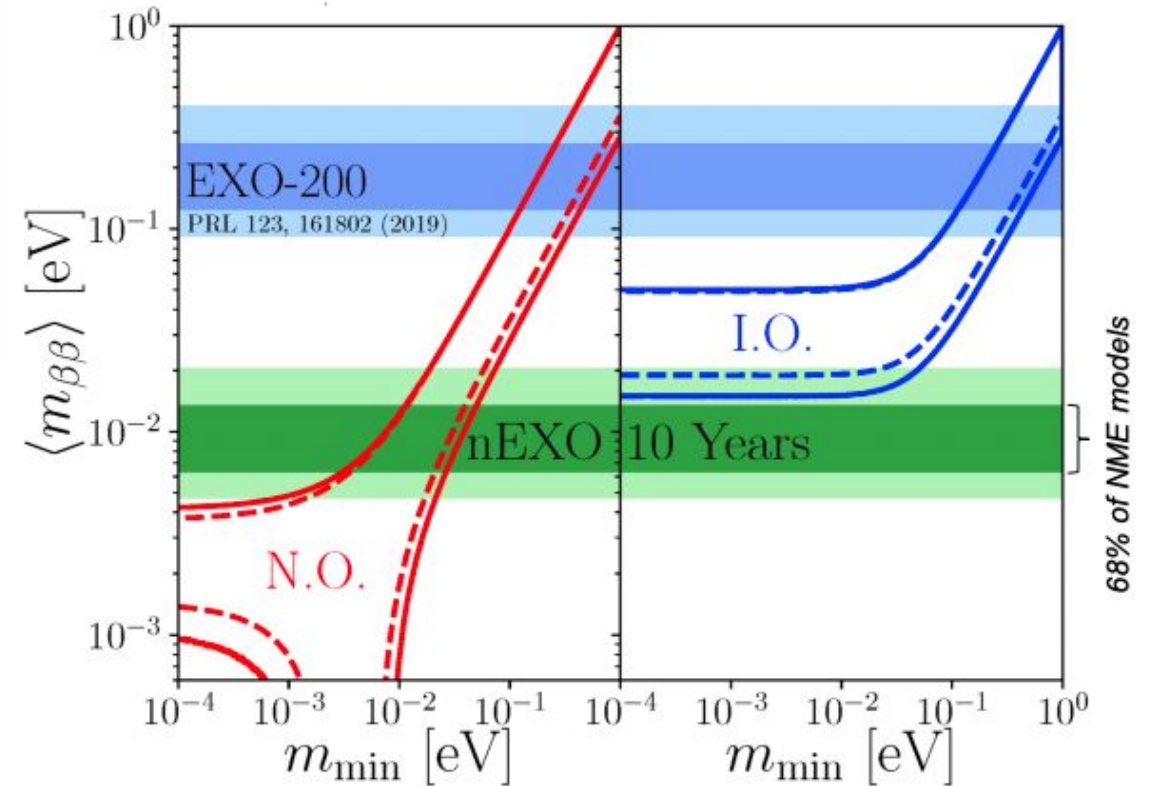
J. Kotila and F. Iachello, Phys Rev C 85, 034316 (2012)

nEXO Sensitivity



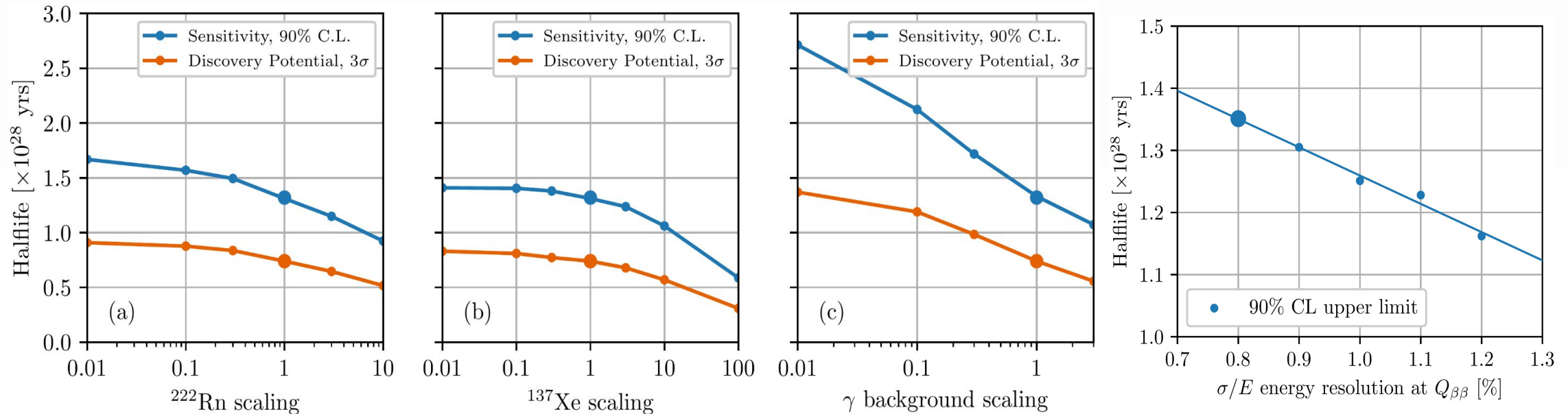
Neutrino Mass Measurement

- In 6.5 years, nEXO will reach a sensitivity to $0\nu\beta\beta$ half life in xenon $>10^{28}$ years
 - Age of the universe $\times 10^{18}$!
- Effective Majorana mass of the neutrino $\lesssim 8$ meV; excludes inverted mass ordering parameter space



nEXO Sensitivity Robustness

- Well studied response to fluctuations in background model, energy resolution, ...

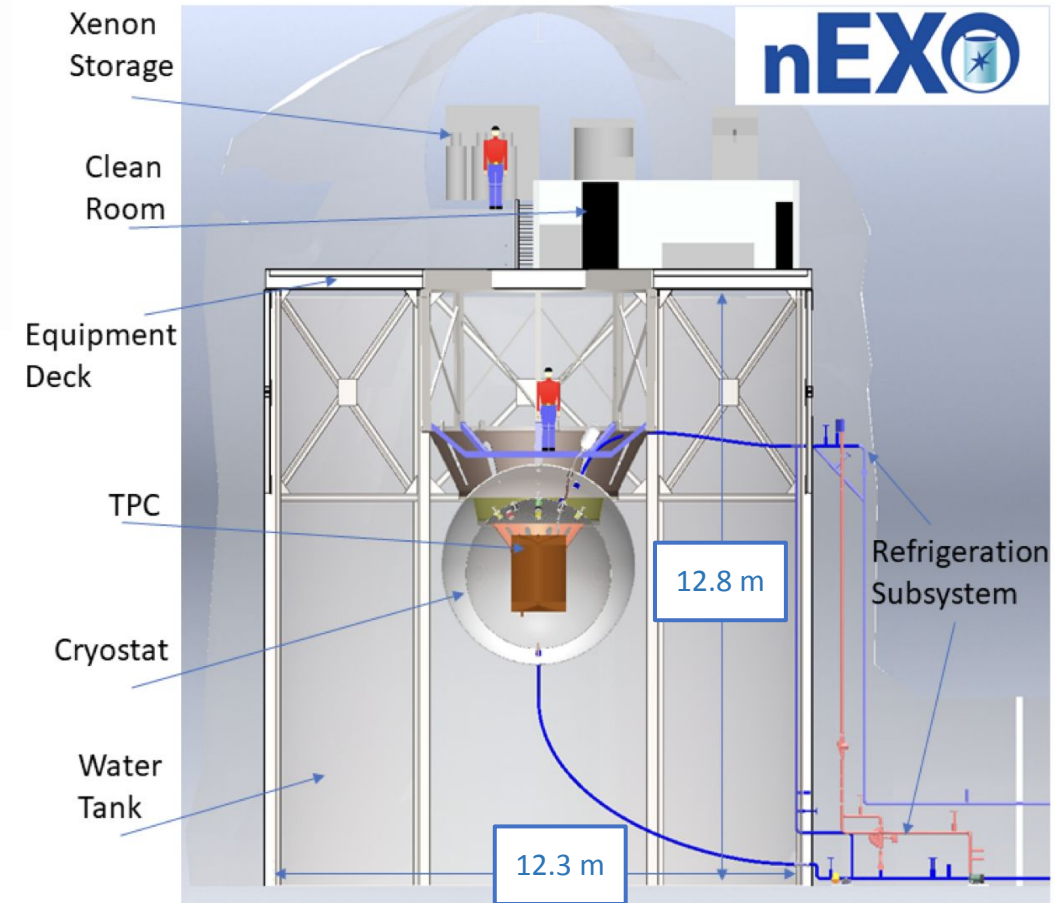


Confidence in the sensitivity estimate arises from a detailed conservative model with measured input parameters

G Adhikari et al. (nEXO Collaboration), 2022 J. Phys. G: Nucl. Part. Phys. 49 015104

Summary

- nEXO is searching for Lepton Number Violation via $0\nu\beta\beta$
- nEXO utilizes a 5 tonne single-phase LXe TPC
 - LXe is purifiable in-situ
 - LXe is self-shielding against γ backgrounds
 - Multiparameter analysis provides robustness to unknown backgrounds, and background fluctuations
 - Scalable technology & repeatable experiment
 - Possibility for "Ba-tagging" upgrade
- nEXO's sensitivity to $0\nu\beta\beta$ half life in ^{136}Xe is 1.35×10^{28} yr
 - $\lesssim 8$ meV effective Majorana mass of the neutrino



Thank you!

Ask me about nEXO Diversity Equity & Inclusion Activities:

- Mentorship program
- Climate surveys
- Outreach

Follow us!

@nEXOexperiment



soud.alkharusi@mail.mcgill.ca

<https://www.physics.mcgill.ca/~soudal/>

Thank you!

nEXO Publications:

2022:

- Performance of novel VUV-sensitive Silicon Photo-Multipliers for nEXO
- Development of a ^{127}Xe calibration source for nEXO

2021:

- nEXO: neutrinoless double beta decay search beyond 1028 year half-life sensitivity
- Reflectivity of VUV-sensitive silicon photomultipliers in liquid Xenon
- Event reconstruction in a liquid xenon Time Projection Chamber with an optically-open field cage

2020:

- Reflectance of Silicon Photomultipliers at Vacuum Ultraviolet Wavelengths
- Measurements of electron transport in liquid and gas Xenon using a laser-driven photocathode

2019:

- Characterization of the Hamamatsu VUV4 MPPCs for nEXO
- Simulation of charge readout with segmented tiles in nEXO

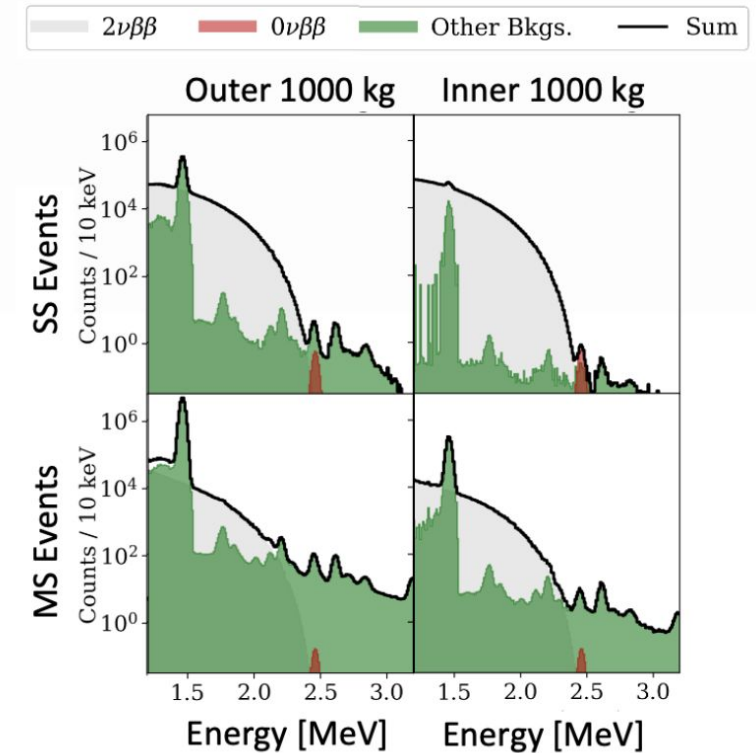
2018

- nEXO pre-conceptual design report



Unknown external background?

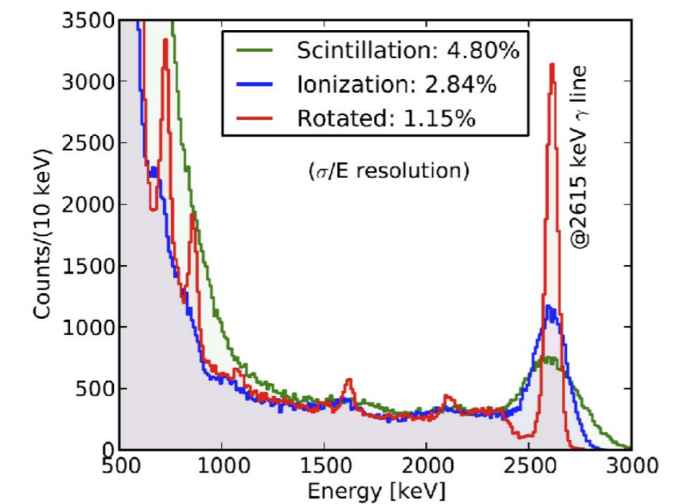
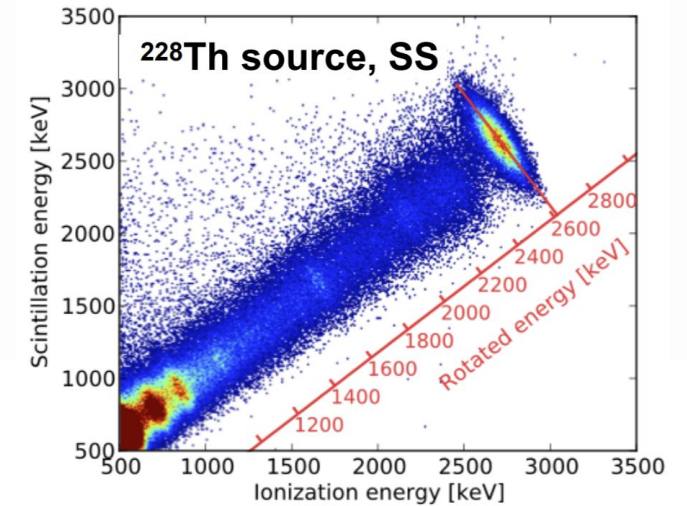
If an unknown decay were strong enough to produce as many SS events in the inner 3000 kg as a 3σ discovery at a half-life of 5.7×10^{27} yr, this decay would produce 271 counts in the MS outer volume, enough to rule out the expected background model at $p < 0.00001$.



[Phys. Rev.C 97, 065503 \(2018\)](#)

Rotated energy scale

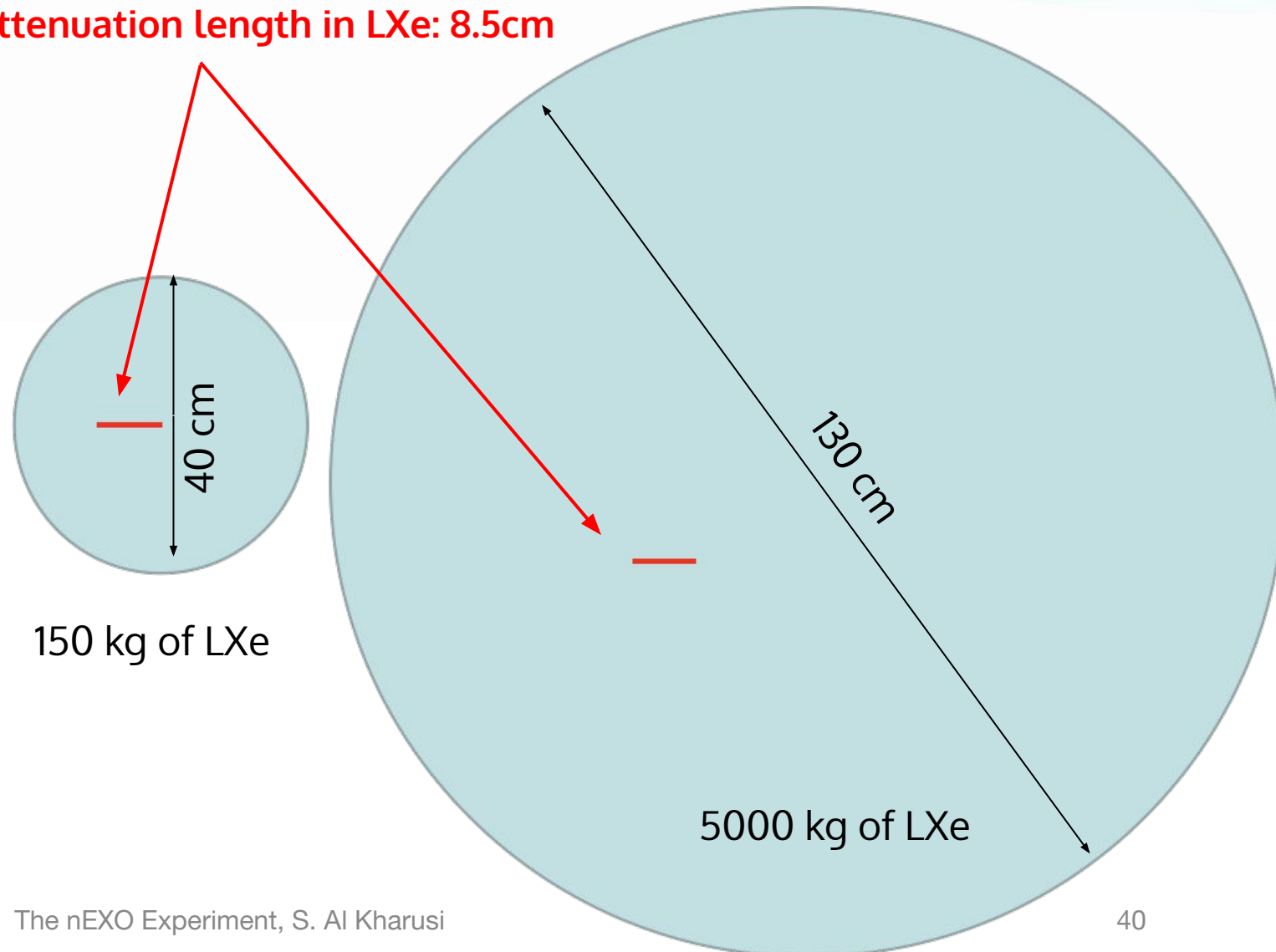
- LXe rotated energy (exploiting anticorrelation in charge and light) allows for optimization of energy resolution
 - [Conti, E., et al. "Correlated fluctuations between luminescence and ionization in liquid xenon." Phys. Rev. B 68.5 \(2003\): 054201.](#)
- 2022: LZ [achieved <0.7% energy resolution](#) in LXe!



LXe TPC Scalability (1/2)

2.5MeV γ ray attenuation length in LXe: 8.5cm

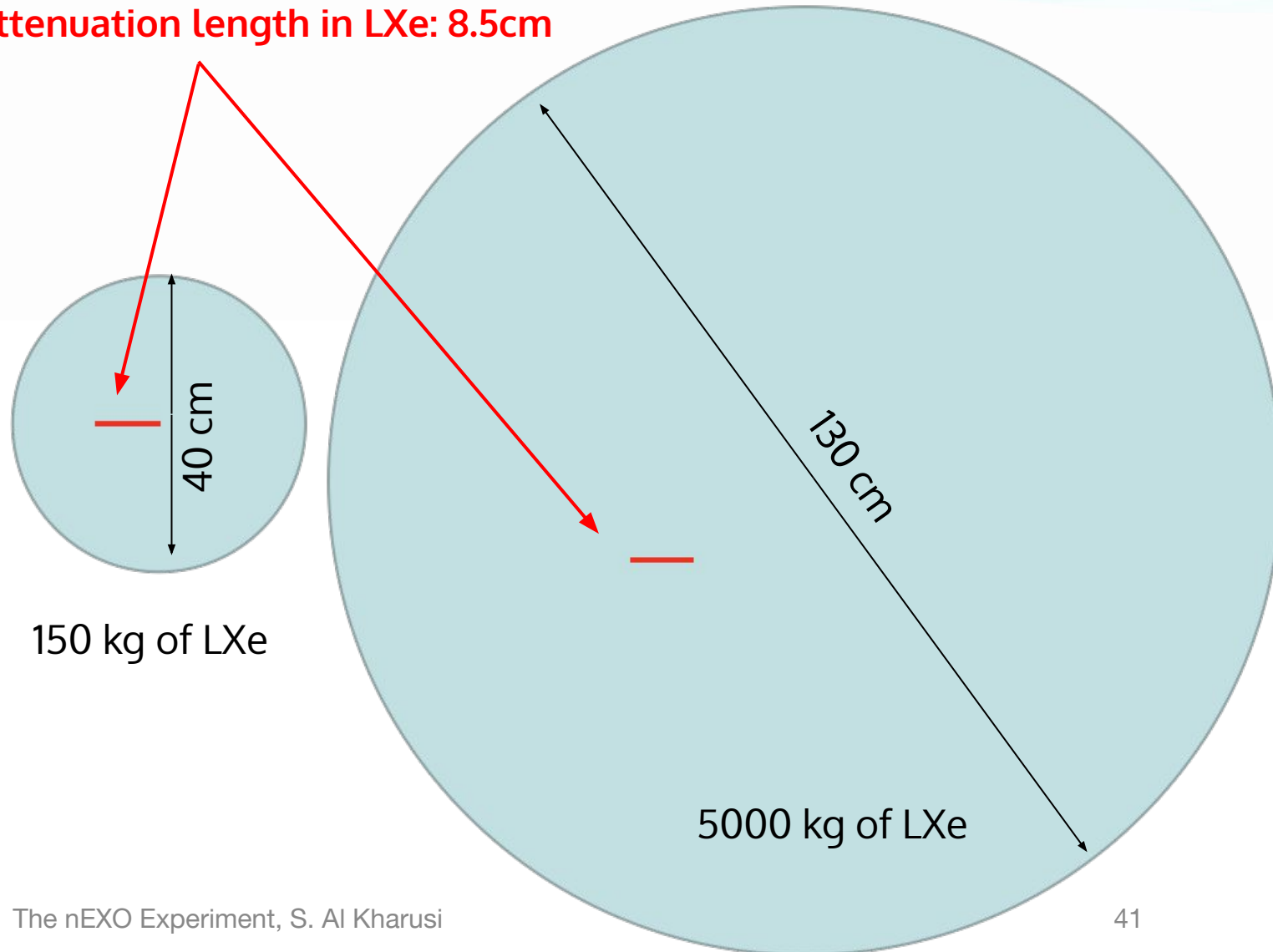
- LXe is self shielding: larger TPC means **better attenuation of gammas**, and even **better constraints on fluctuations to backgrounds**.
- Any potential $0\nu\beta\beta$ signal would have to **not be anomalous in all 3 high level distributions**: Energy, Standoff, and Topology.
 - Due to gamma attenuation lengths \ll detector scale, this improves with larger masses



LXe TPC Scalability (2/2)

2.5MeV γ ray attenuation length in LXe: 8.5cm

- Going from 5 tonne to 100 tonne would require LXe TPCs of size scale = $1.3 \cdot (100/5)^{1/3} \sim 3.5$ m
- We know how to make liquid noble TPCs even larger (see DUNE)



Beyond $0\nu\beta\beta$ discovery?

- If $0\nu\beta\beta$ is discovered in any isotope, we would want to explore **what mechanism is producing the decay**
 - We would do this by measuring the **energy and angular distributions** of the two emitted electrons in $0\nu\beta\beta$ events
 - Straightforward in an enriched gaseous xenon TPC
 - Design constraints set by half life measurements in an LXe TPC (e.g. nEXO)
- $0\nu\beta\beta$ decay mechanisms change the value of $\langle m_{\beta\beta} \rangle$, and probe couplings to BSM physics
- Discovering $0\nu\beta\beta$ and exploring it in multiple isotopes is key
 - Nuclear physics is hard, and extracting BSM physics couplings without multiple isotopes confirming $0\nu\beta\beta$, half lives, mechanisms etc... will be difficult

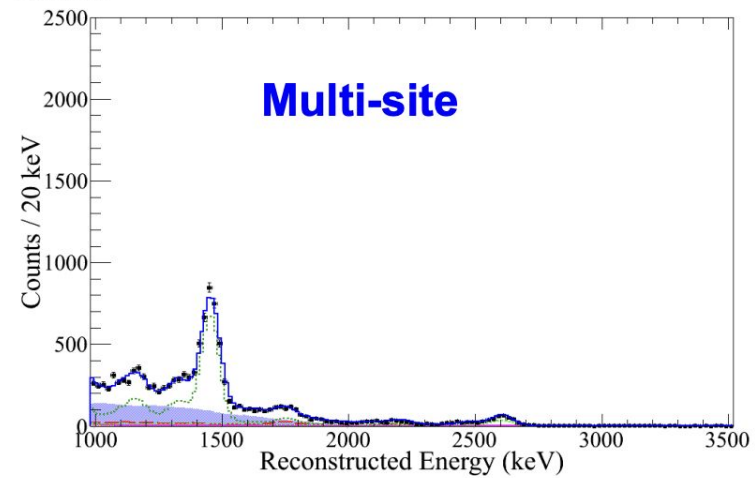
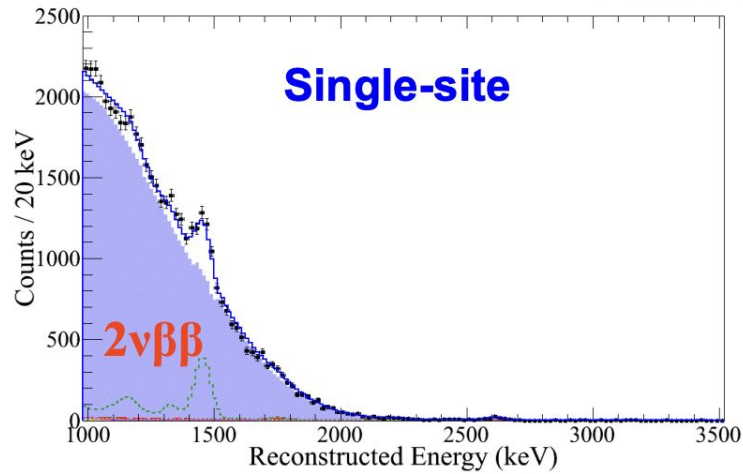
Multiparameter Analysis

EXO-200 Validation

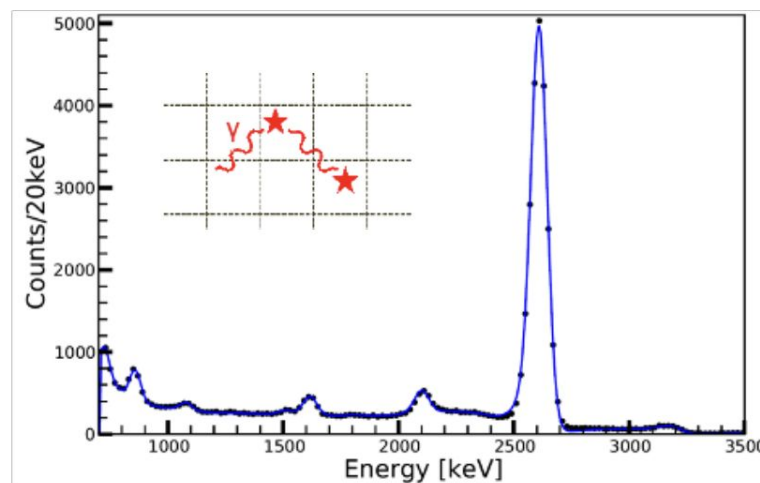
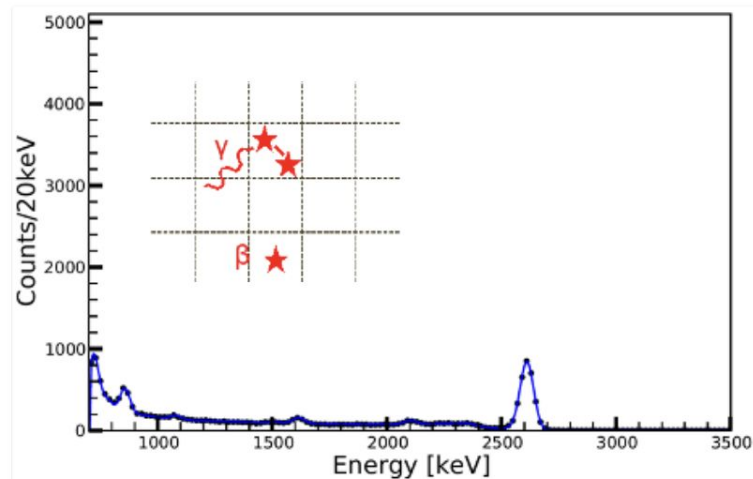


EXO-200 data

Low background
data



^{228}Th calibration
source



Multiparameter Analysis

nEXO is not a counting experiment

- $\sim 1\%$ energy resolution at $Q_{\beta\beta}$
- Topology scoring: single- and multi-site discrimination (β -like vs γ -like event separation)
- standoff distance to detector components (precise event localization, depth in xenon)

