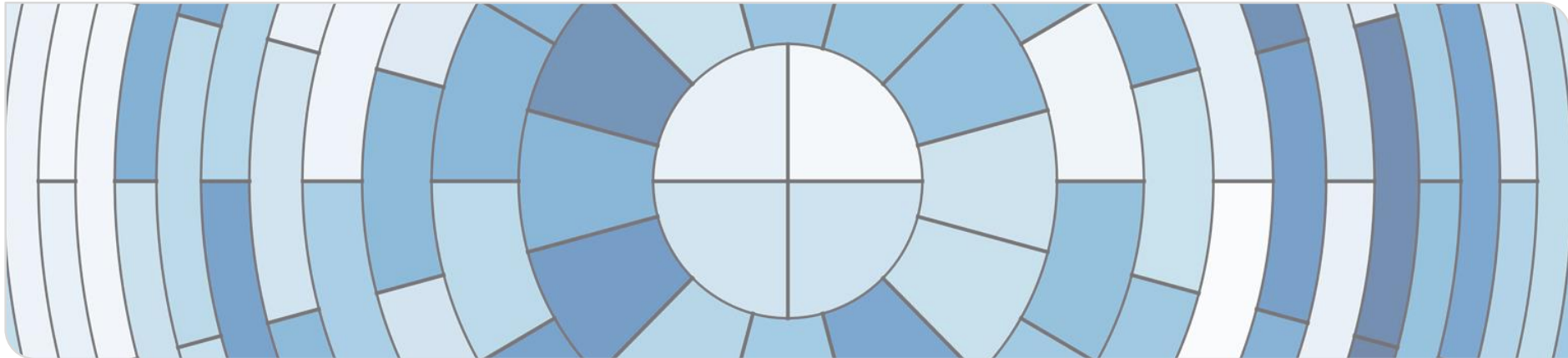




# Latest results from the KATRIN experiment

Fabian Block for the KATRIN collaboration, Lake Louise Winter Institute 2024



# Complementary ways of determining the neutrino mass

## ■ high-precision cosmology

$$\Sigma = \sum_i m_i < 0.06 - 0.14 \text{ eV}$$

Stöcker et al., Phys.Rev.D 103 (2021) 12, 123508  
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model-dependent

(neutrino nature, nuclear matrix elements)

## ■ kinematic neutrino mass measurement

model-independent

(independent of neutrino nature, solely based on kinematic parameters and energy conservation)

# Kinematic neutrino mass measurement (by example of $\beta$ -decay)

- continuous  $\beta$ -decay spectrum described by Fermi's Golden Rule

$$\frac{d\Gamma}{dE} = C \cdot p \cdot (E + m_e) \cdot (E_0 - E) \cdot \sum_{i=1}^3 |U_{ei}^2| \sqrt{(E_0 - E)^2 - m_i^2} \cdot F(E, Z) \cdot \theta(E_0 - E - m_{\nu_i})$$

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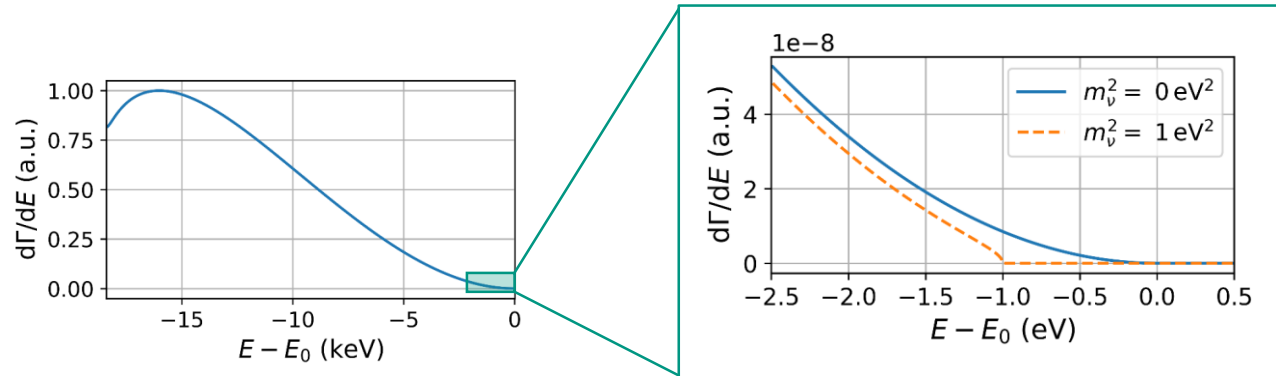


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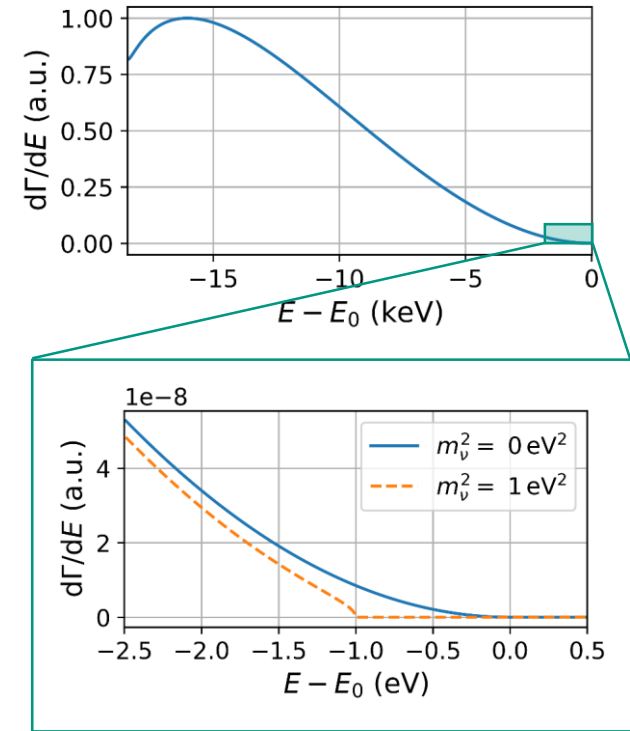
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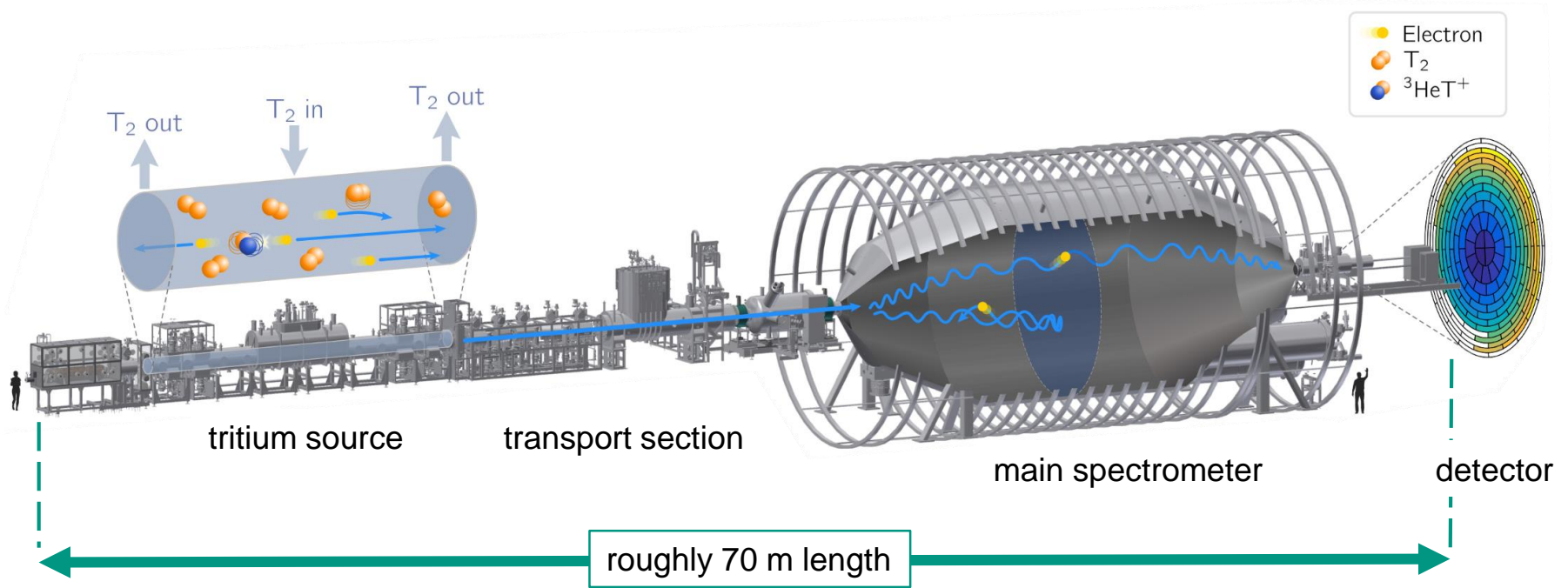


# KATRIN requirements

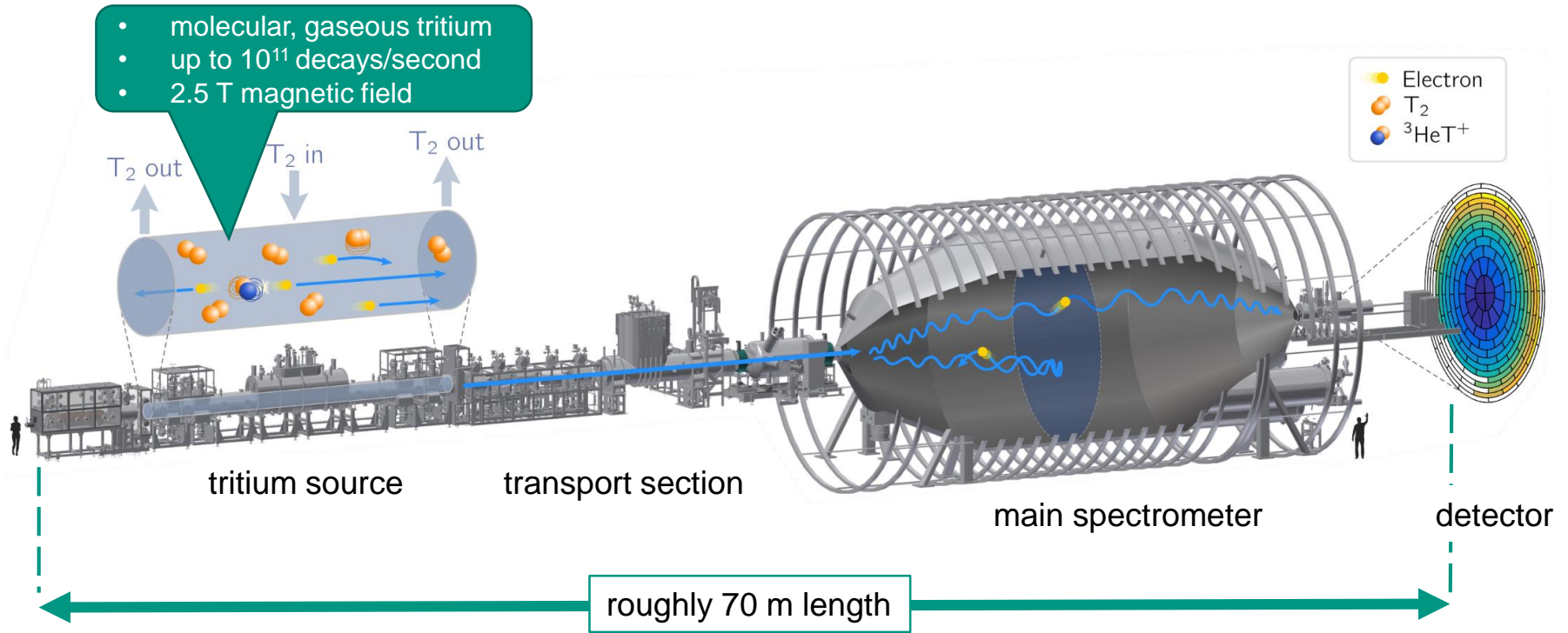
- goal:
  - resolve spectral distortion close to the  $\beta$ -decay endpoint
- requirements:
  - strong  $\beta$ -decay source
  - excellent energy resolution
  - low background level
  - exact understanding of spectrum shape



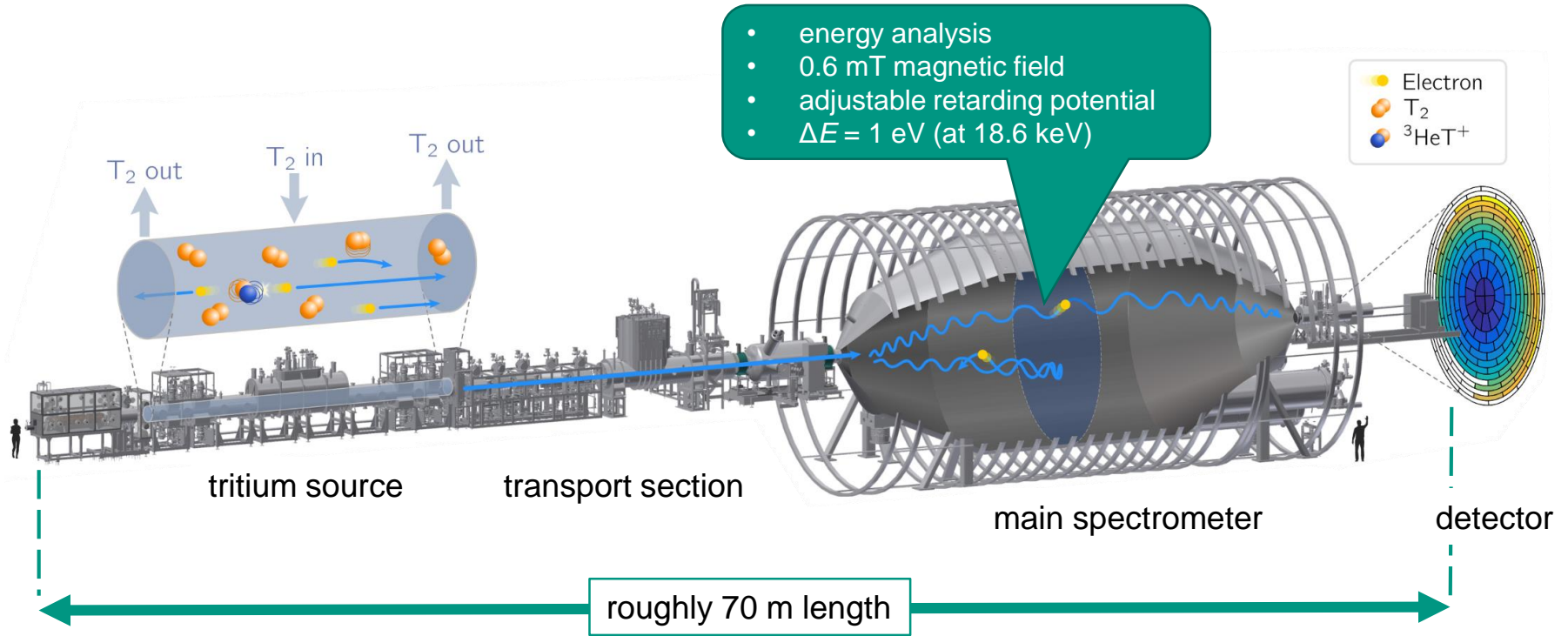
# KATRIN experimental setup



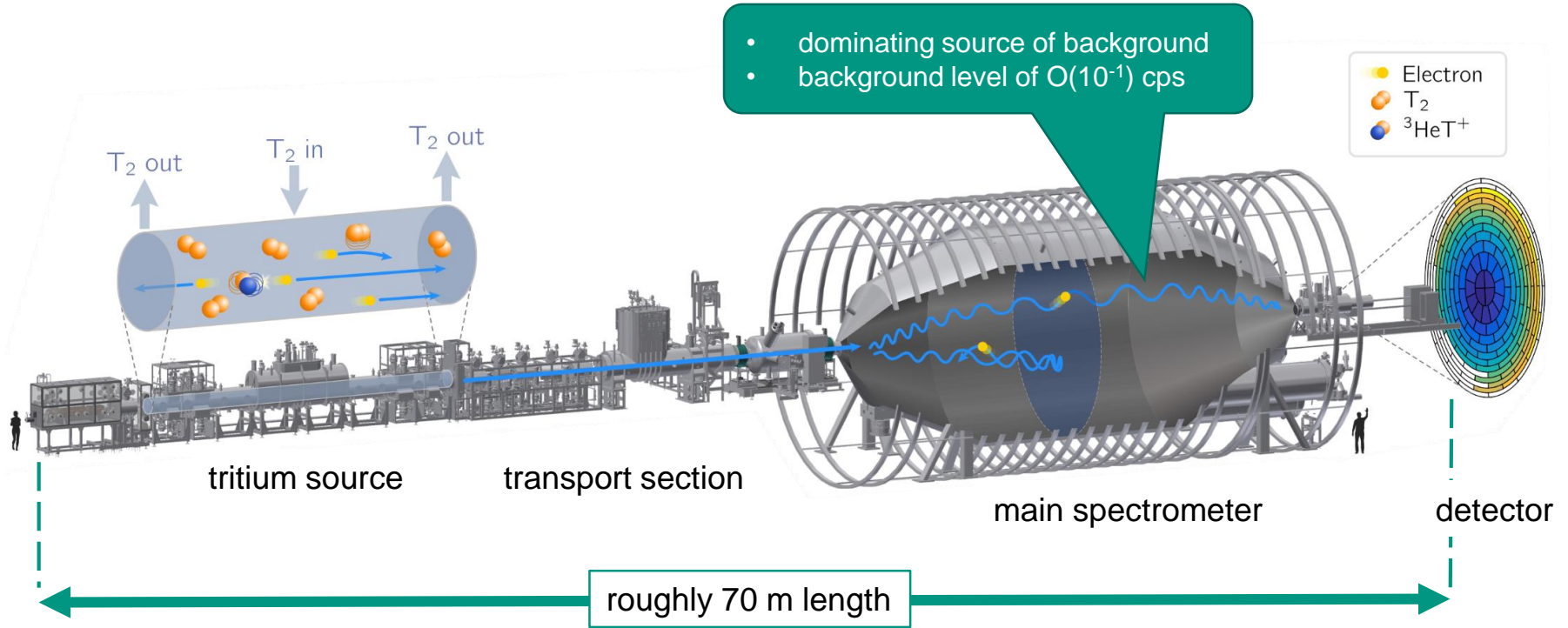
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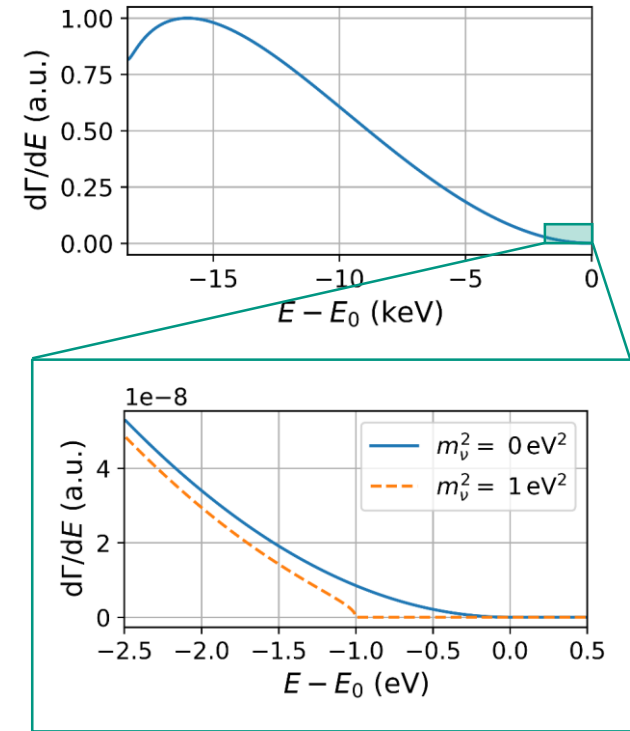


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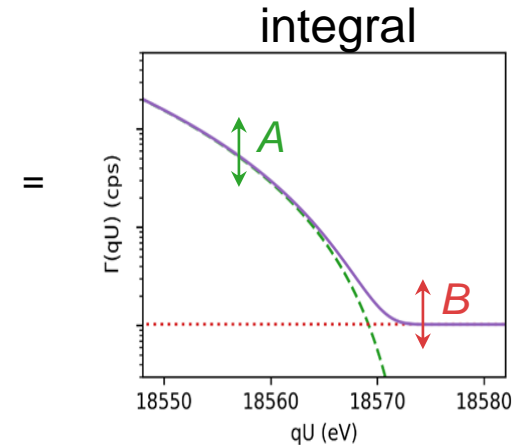
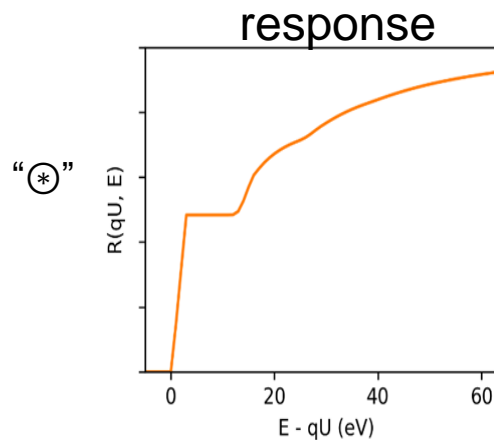
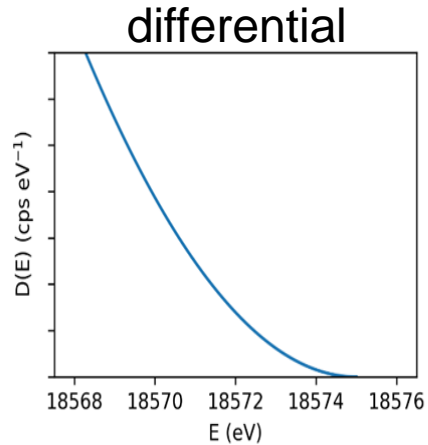
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# Exact understanding of spectrum shape

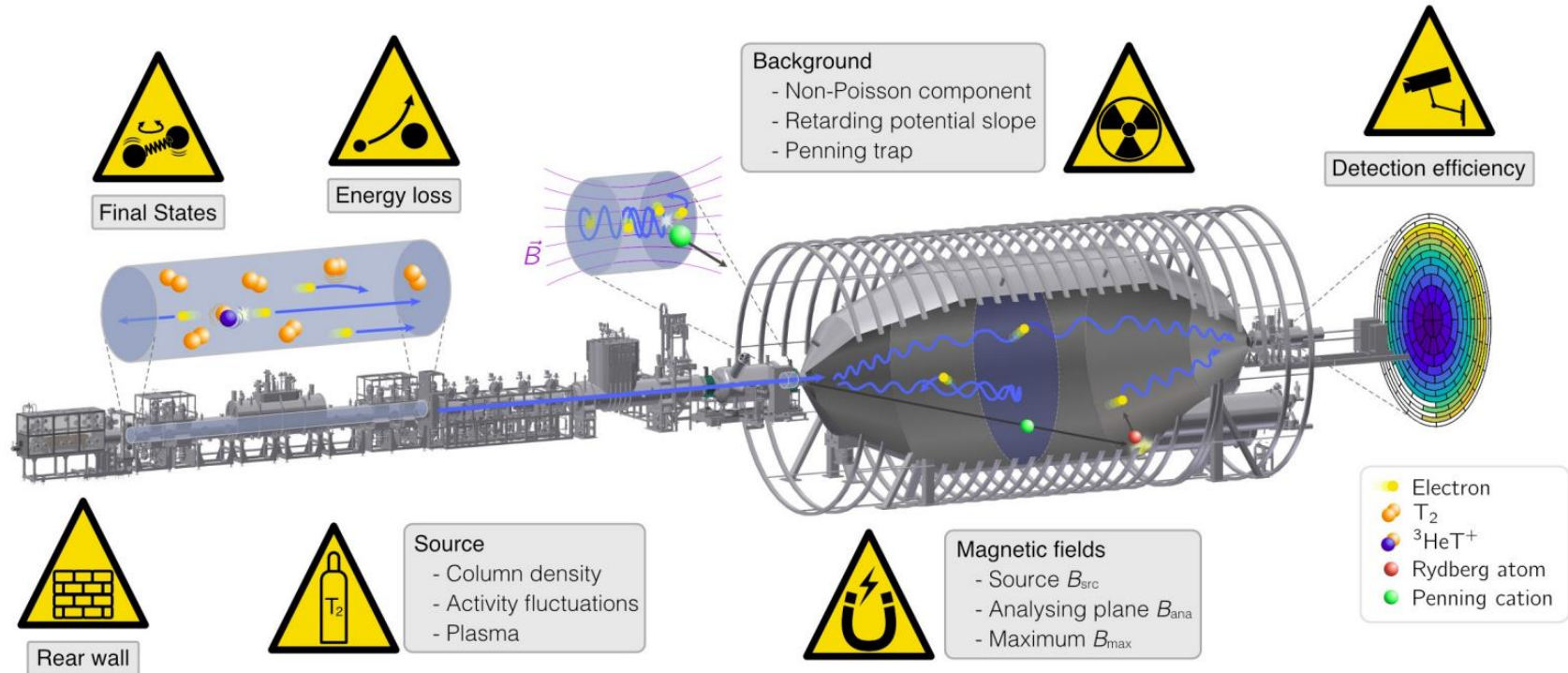
- KATRIN measures the integral  $\beta$ -decay spectrum

$$\Gamma(qU) \propto A \int_{qU}^{E_0} D(E; m_\beta^2, E_0) R(qU, E) dE + B$$







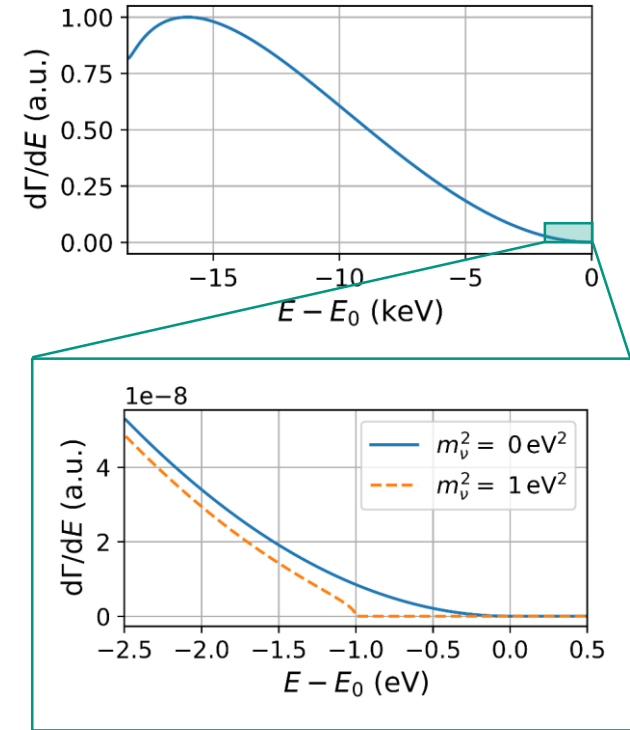


# Exact understanding of spectrum shape



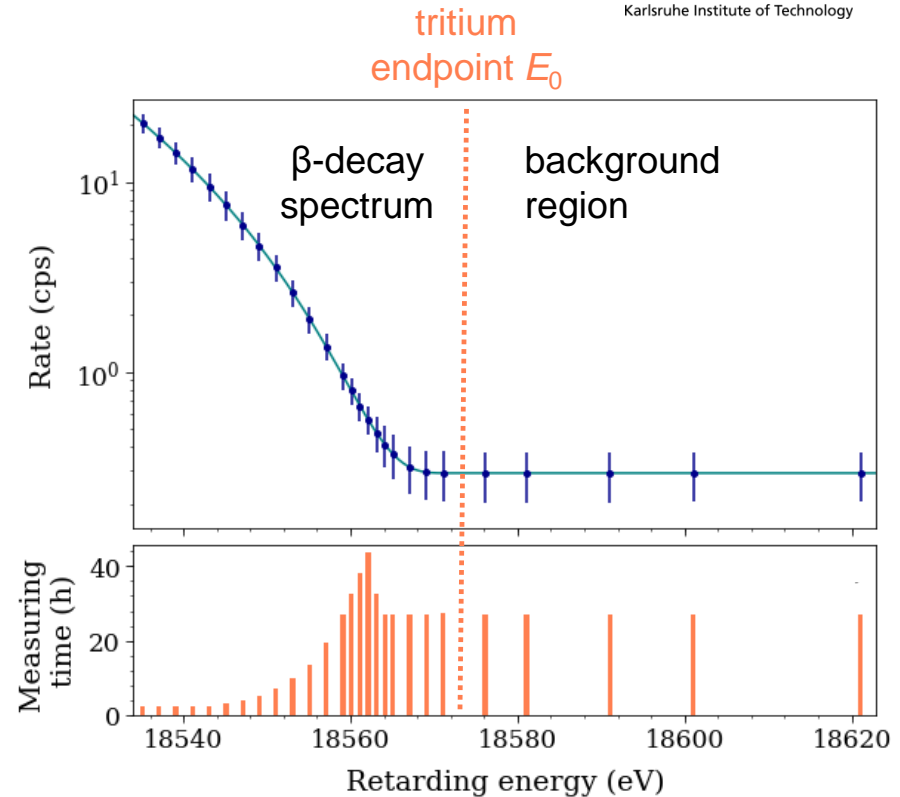
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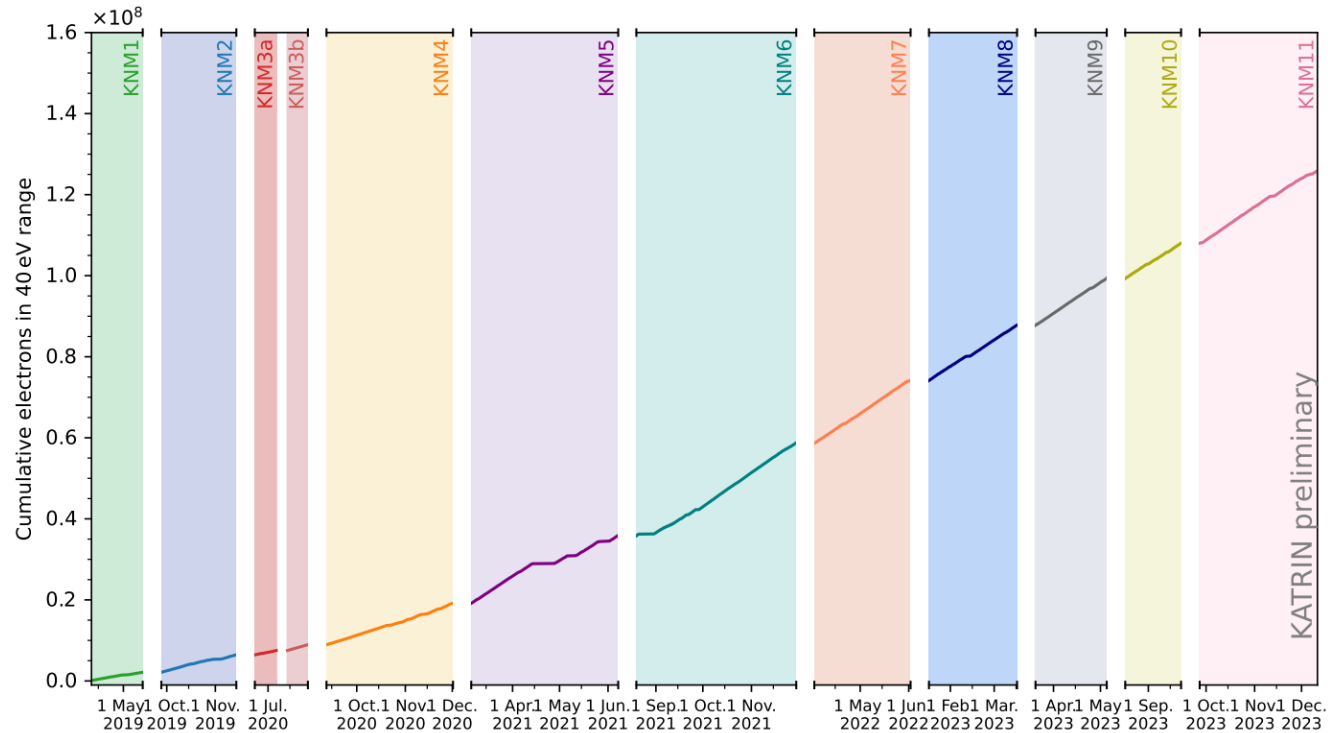


# KATRIN tritium scan

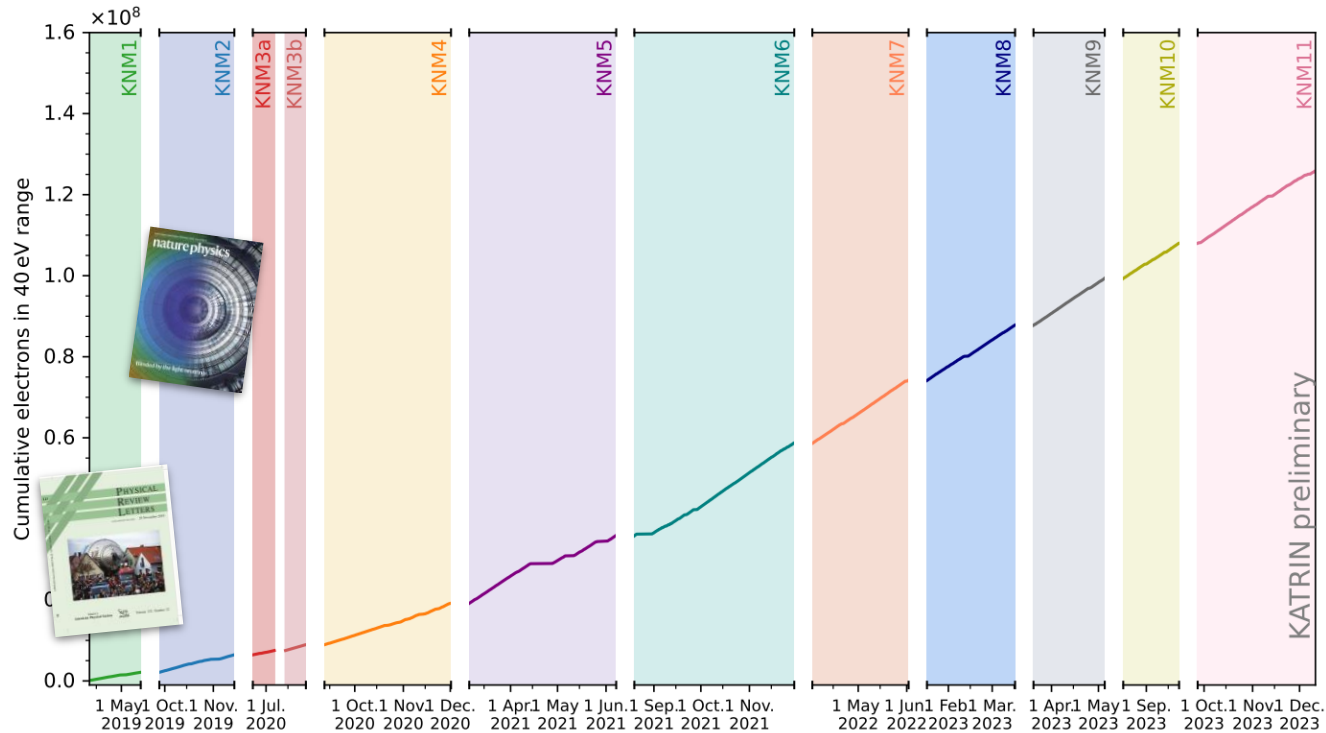
- ~30 set points of the retarding potential with varying duration
- scan interval ranges from  $E_0 - 40$  eV up to  $E_0 + 130$  eV
- ~ 3 hours measurement time per scan
- O(500) scans per year



# Overview of data taking



# Overview of data taking



# Latest neutrino mass result

## ■ first measurement campaign

■ best fit:  $m_\beta^2 = (-1.0_{-1.1}^{+0.9}) \text{ eV}^2$

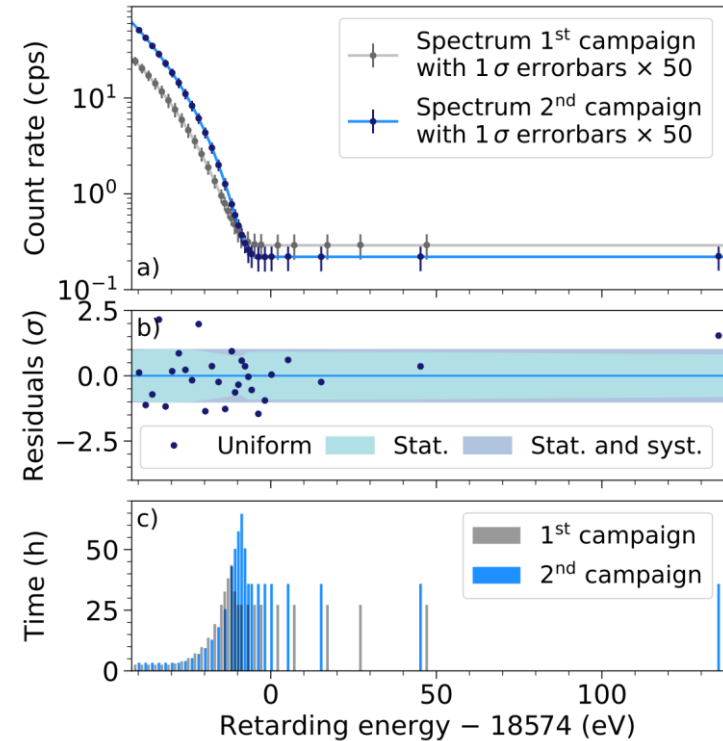
■ limit:  $m_\beta < 1.1 \text{ eV}$  (90% CL)

M. Aker et al., Phys. Rev. Lett. **123**, 221802

## ■ second measurement campaign

■ best fit:  $m_\beta^2 = (0.26_{-0.34}^{+0.34}) \text{ eV}^2$

■ limit:  $m_\beta < 0.9 \text{ eV}$  (90% CL)

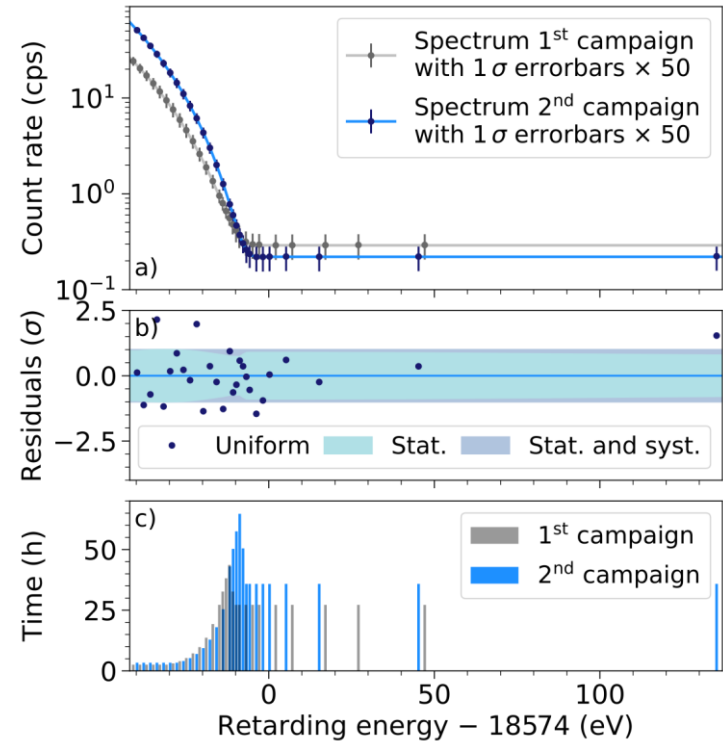


# Latest neutrino mass result

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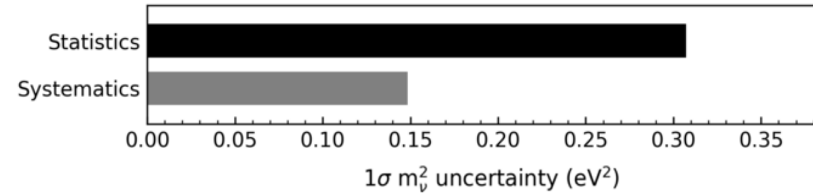
M. Aker et al., Phys. Rev. Lett. **128**, 101801 (2022)
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    - best fit:  $m_\beta^2 = (0.26^{+0.34}_{-0.34}) \text{ eV}^2$
    - limit:  $m_\beta < 0.9 \text{ eV}$  (90% CL)
  - combined result:  $m_\beta < 0.8 \text{ eV}$  (90% CL)
- M. Aker et al., Nat. Phys. **18**, 160–166 (2022)

first direct neutrino-mass measurement  
with sub-eV sensitivity



# Uncertainty breakdown of KNM2

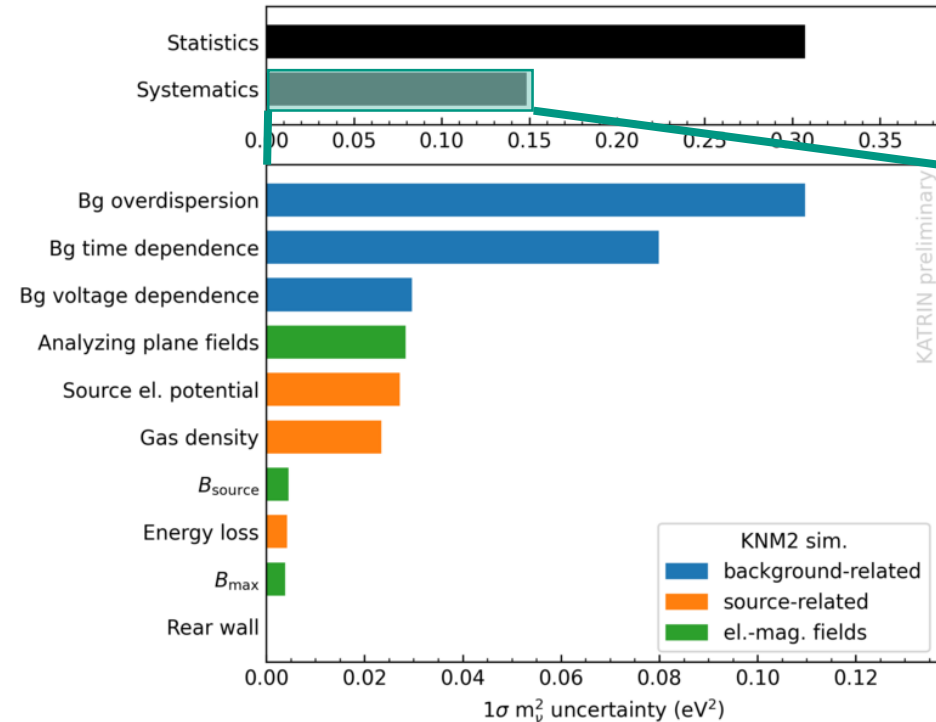
- statistical uncertainty dominates
  - systematics non-negligible





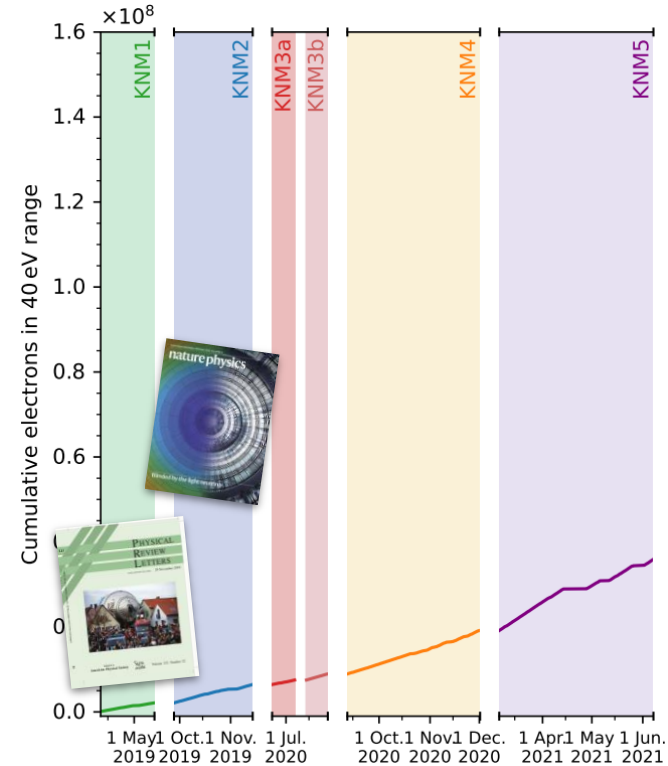
# Uncertainty breakdown of KNM2

- statistical uncertainty dominates
  - systematics non-negligible
- background-related uncertainties dominate systematics budget



# Next analysis release

- combined analysis of first five campaigns
- currently in unblinding process
- next data release in summer



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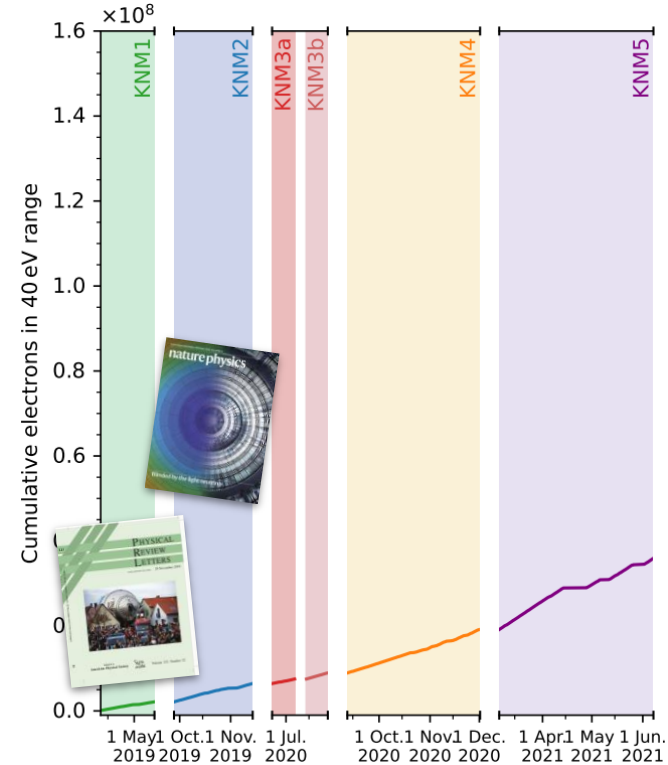
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- sensitivity projection:

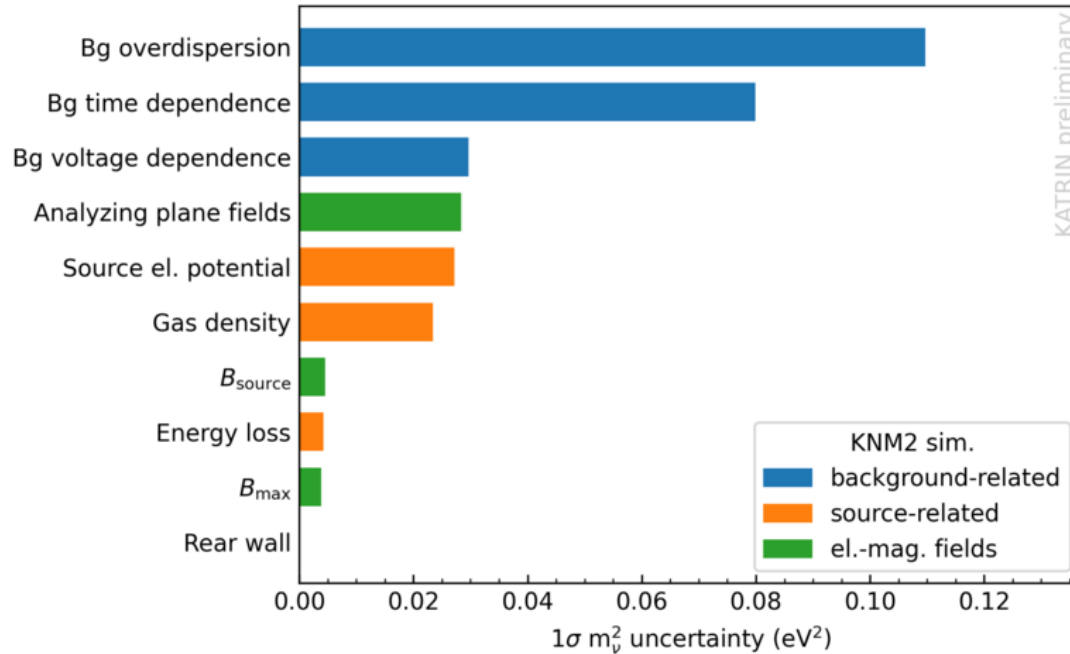
$$m_{\beta} < 0.5 \text{ eV (90\% CL)}$$

- improvements

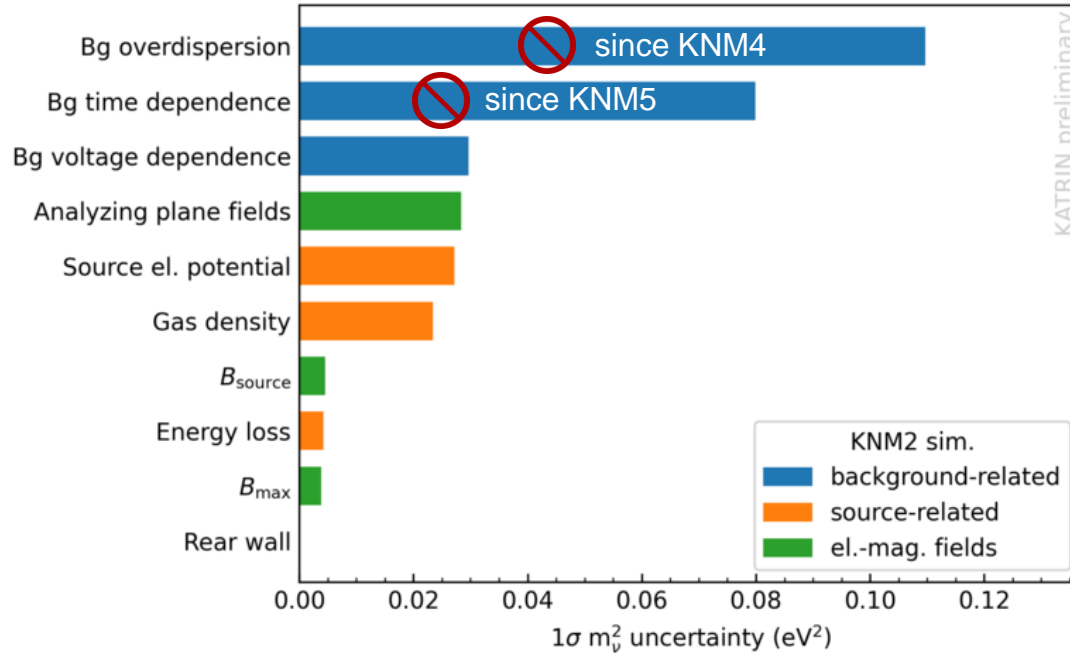
- factor six in statistics
- background reduction
- more accurate knowledge of systematics




# Next analysis release - uncertainty breakdown

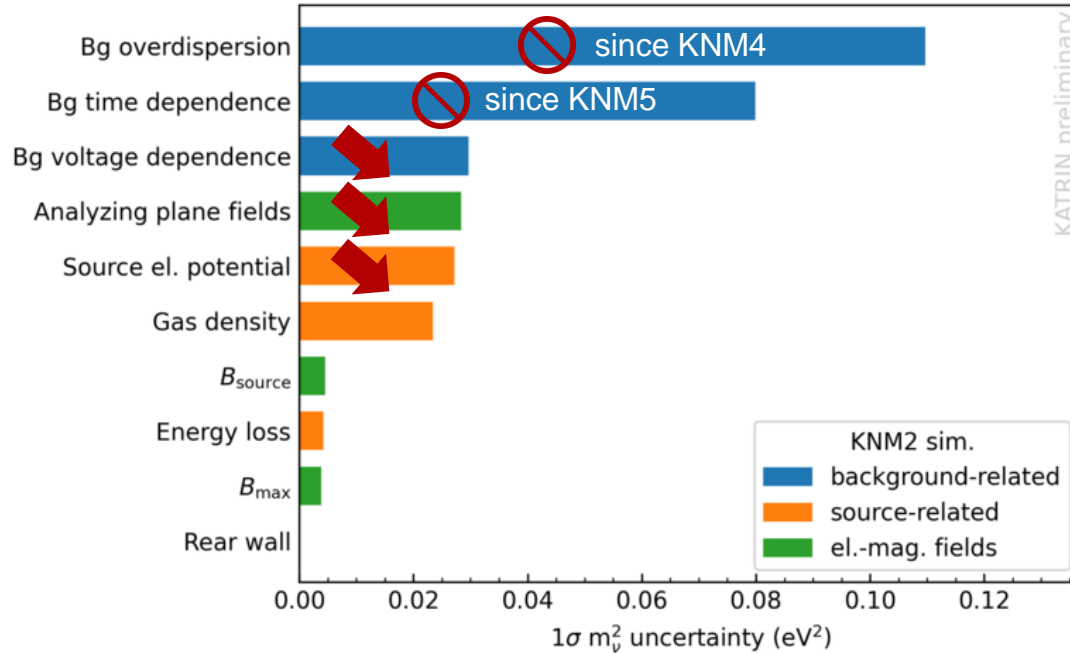



# Next analysis release - uncertainty breakdown




 counter measures implemented for data taking

# Next analysis release - uncertainty breakdown



 counter measures implemented for data taking

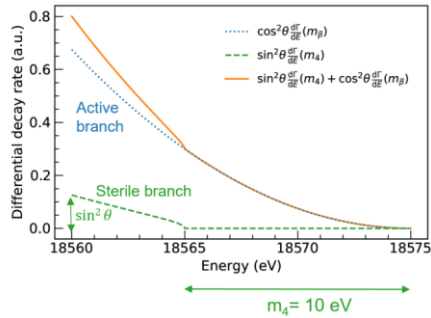
 decrease due to new calibration measurements

# “Beyond neutrino mass” in KATRIN

high statistics and  
precise  $\beta$  spectrum

# “Beyond neutrino mass” in KATRIN

search for sterile neutrinos



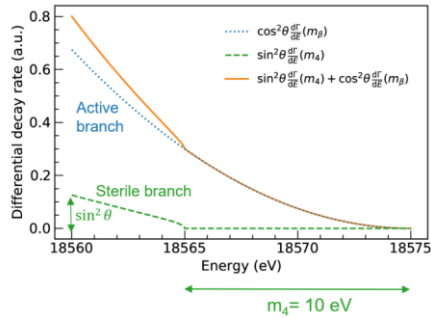
high statistics and precise  $\beta$  spectrum

Aker et al., 2022 *Phys. Rev. D* **105** 072004  
 Aker et al., 2021 *Phys. Rev. Lett.* **126** 091803  
 Mertens et al., 2015 *J. Cosmol. Astropart. Phys.*  
**JCAP02(2015)020**  
 etc.



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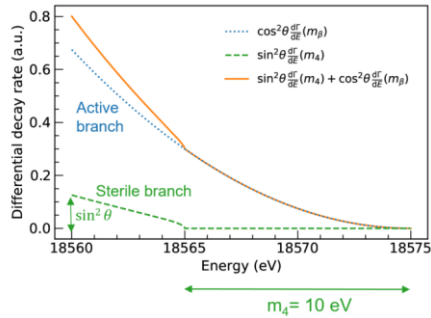
search for exotic interactions

Arcadi et al., 2019 *J. High Energy Phys.* JHEP01(2019)206  
 Fengler et al., 2024 *PoS, DISCRETE2022* 011  
 etc.

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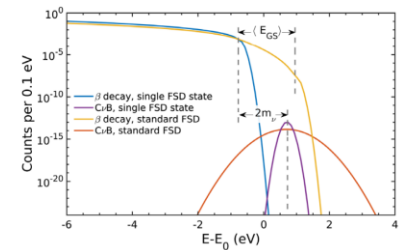
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etc.

set constraints on local density of cosmic relic neutrinos

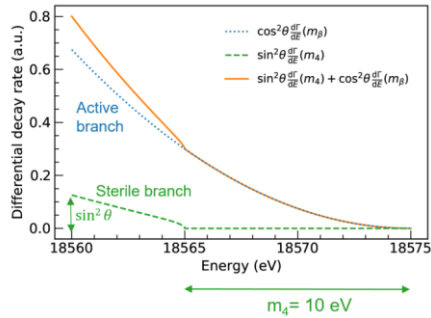
Aker et al., 2022 *Phys. Rev. Lett.* **129** 011806

Aker et al., 2022 *Phys. Rev. D* **105** 072004  
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# “Beyond neutrino mass” in KATRIN

search for sterile neutrinos



Aker et al., 2022 *Phys. Rev. D* **105** 072004  
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 Mertens et al., 2015 *J. Cosmol. Astropart. Phys.*  
**JCAP02(2015)020**  
 etc.

high statistics and precise  $\beta$  spectrum

search for Lorentz invariance violation

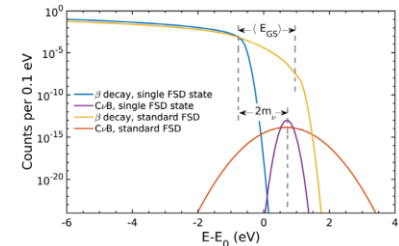
Aker et al., 2023 *Phys. Rev. D* **107** 082005

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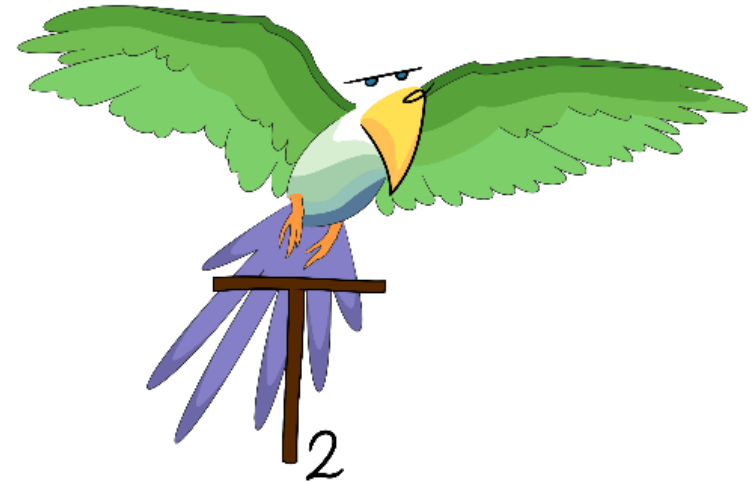


# Summary and outlook

- first direct neutrino-mass measurement with sub-eV sensitivity

$$m_{\beta} < 0.8 \text{ eV (90\% CL)}$$

- combination of first five campaigns
  - significant improvement of systematics and statistics
  - sensitivity:  $m_{\beta} < 0.5 \text{ eV (90\% CL)}$
- multiple physics searches beyond the neutrino mass



# Thank you for your attention!



# Back-up

# Tests of new physics – light sterile neutrinos

- eV-scale sterile neutrino motivated by anomalies in short-baseline neutrino oscillation experiments

Aguilar et al., Phys. Rev. D 64, 112007 (2001)

Aguilar-Arevalo et al., Phys. Rev. Lett. 121, 221801 (2018)

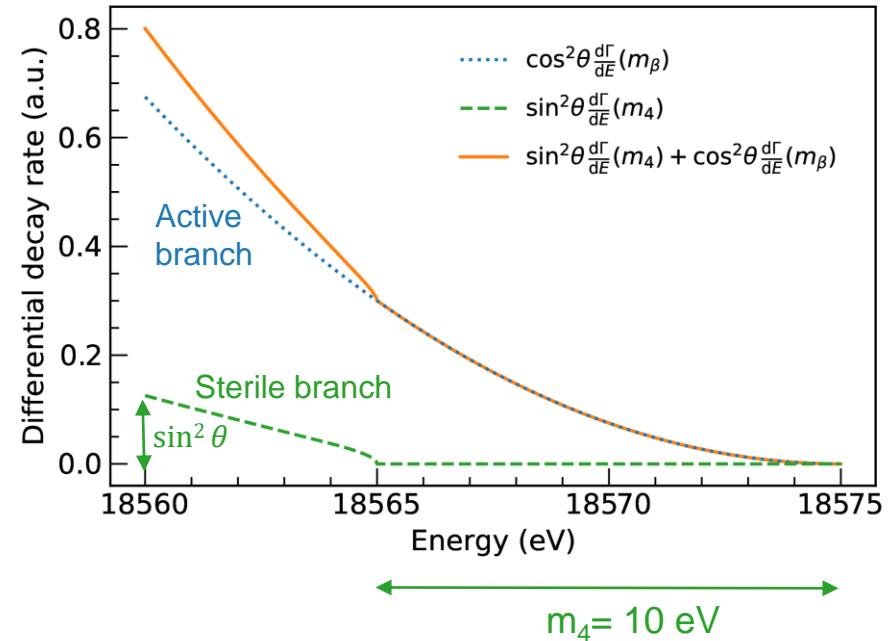
Hampel et al., Phys. Lett. B 420, 114 (1998)

Abdurashitov et al., Phys. Rev. C 80, 015807 (2009)

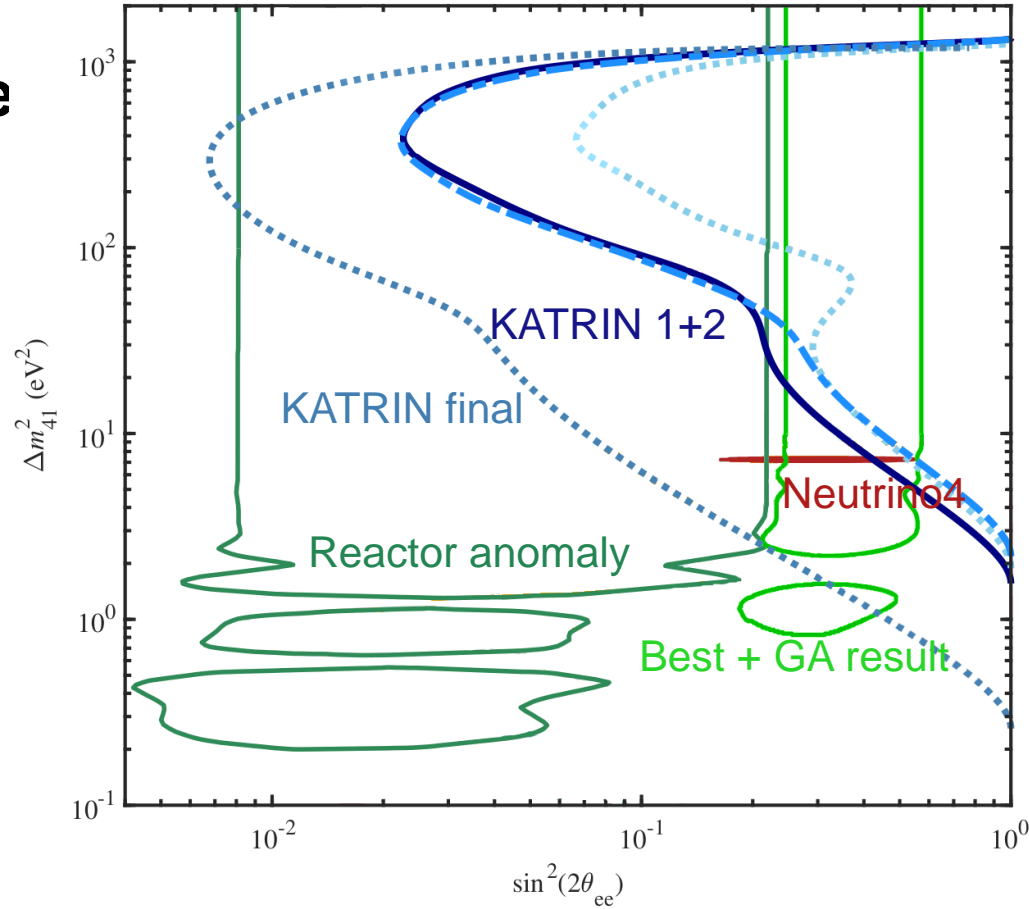
Barinov et al., Phys. Rev. Lett. 128, 232501

Mention et al., Phys. Rev. D 83, 073006 (2011)

Serebrov et al., Pis'ma Zh. Eksp. Teor. Fiz. 109, 209 (2019)



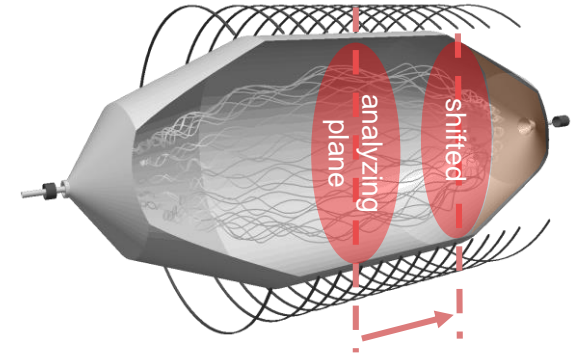
# Tests of $\nu_e$



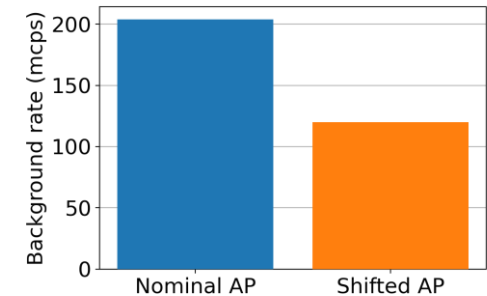


# Background reduction

- dominating background contribution is produced in the main spectrometer
- volume-dependent background rate
  - sensitive to volume between analyzing plane and detector
  - adjusted electromagnetic field layout in main spectrometer
  - significant reduction of background rate



Lokhov et al., *Eur. Phys. J. C* 82, 258 (2022).

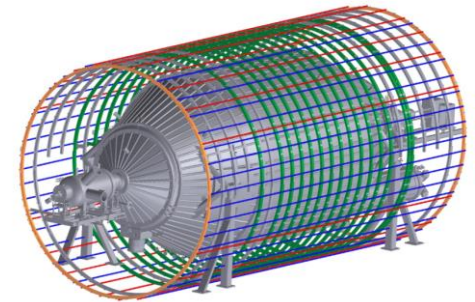


# Main spectrometer as MAC-E filter

- fine-tuned drop of magnetic field in combination with electrostatic retarding potential (MAC-E filter principle) allows main spectrometer energy resolution of

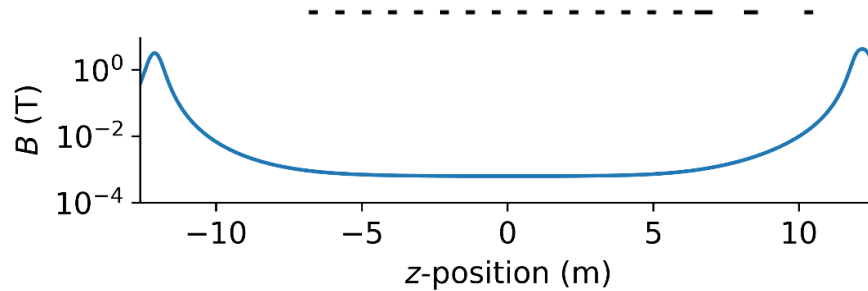
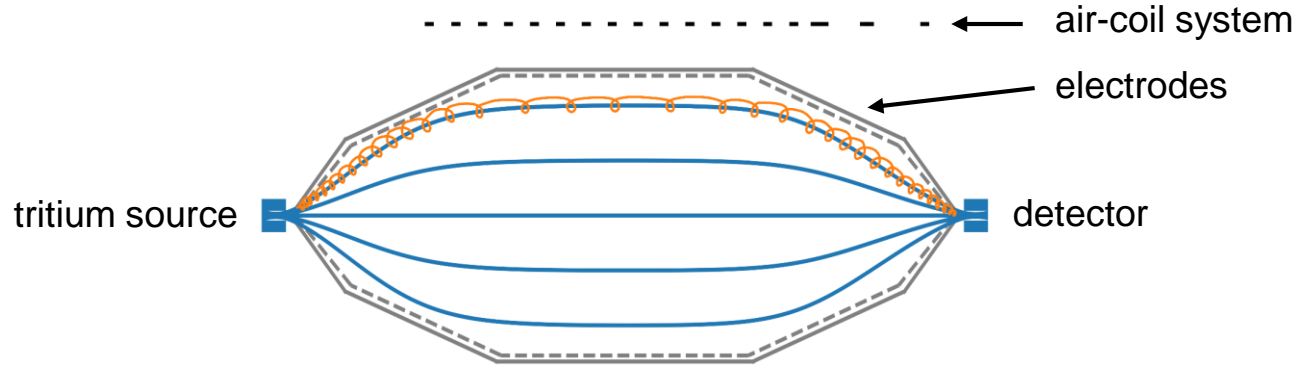
$$\Delta E = 1 \text{ eV (at 18.6 keV)}$$

- air-coil system and inner electrode system allow fine-tuning of electromagnetic field configuration
- high-voltage system can provide up to -40 kV

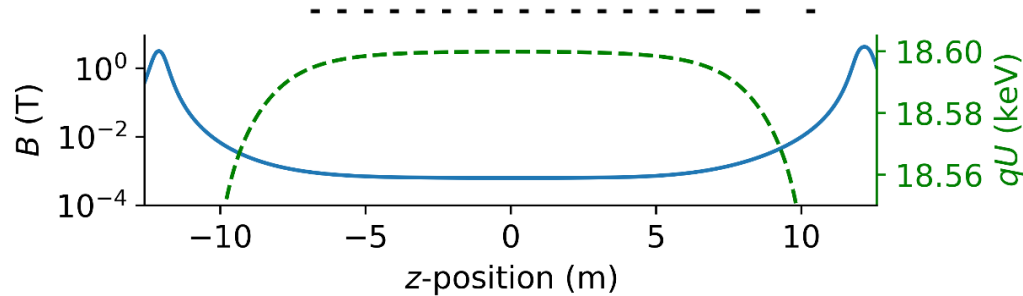
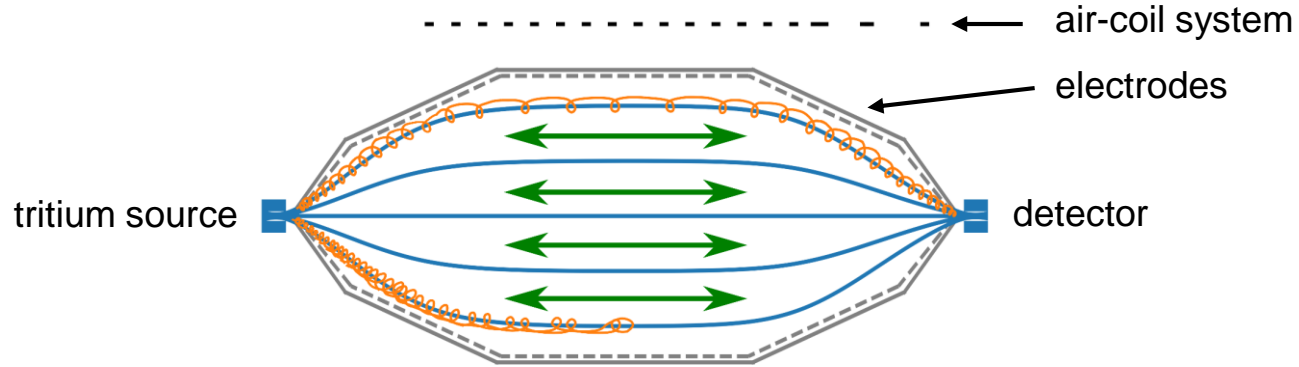


The KATRIN collaboration, 2021 *JINST* 16 T08015

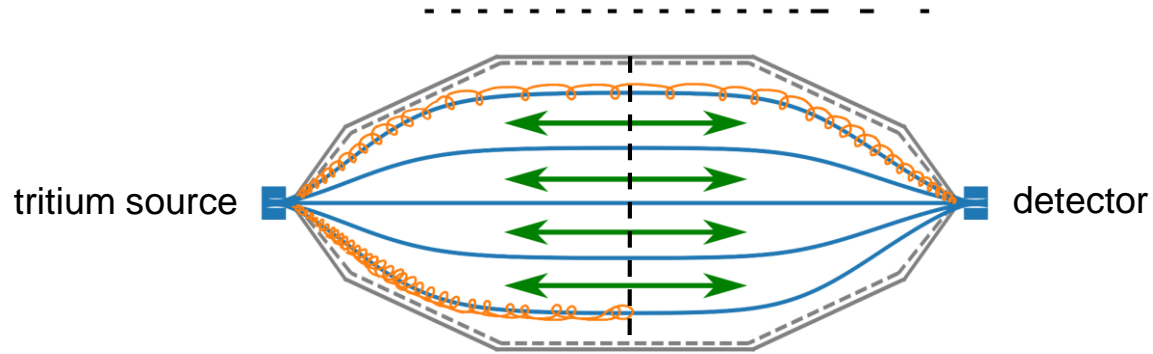
# Main spectrometer as MAC-E filter



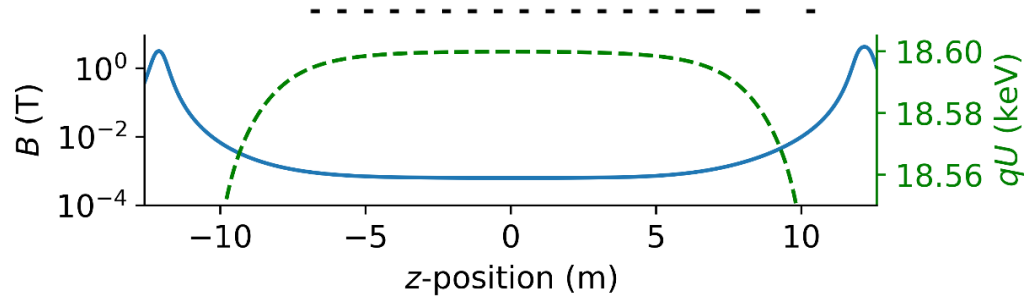
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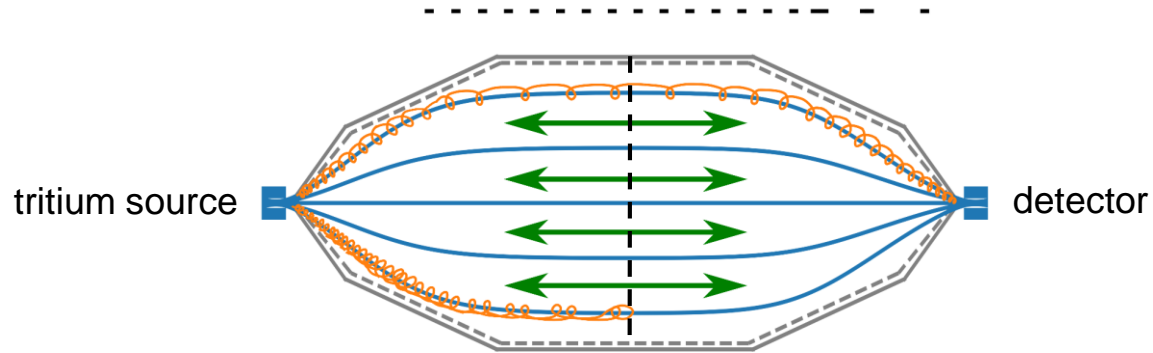
# Main spectrometer as MAC-E filter



analyzing plane:  
 - minimal  $B$   
 - maximal  $qU$



# Main spectrometer as MAC-E filter



analyzing plane:  
 - minimal  $B$   
 - maximal  $qU$



described in  
 transmission  
 function  $T(E, U)$

