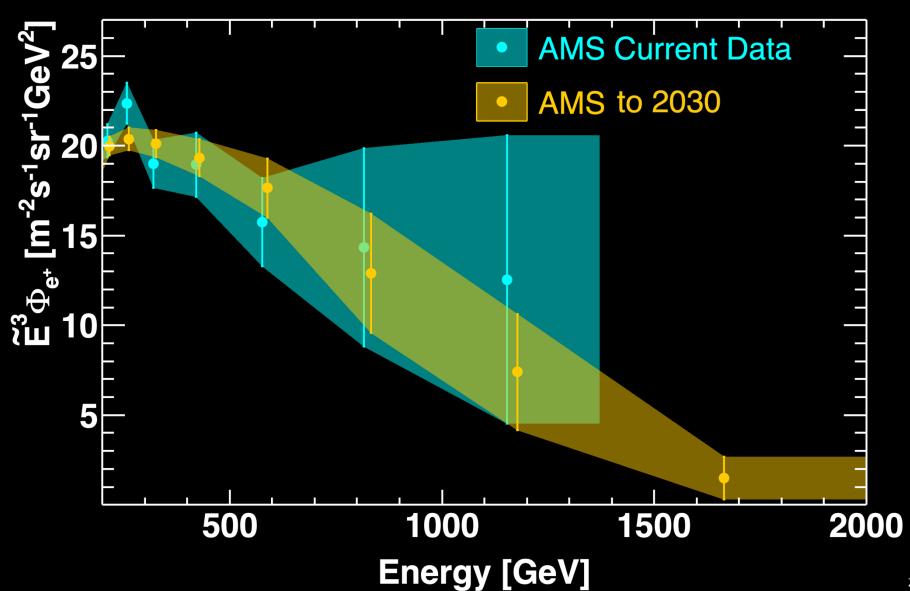


AMS-LO: upgrade status and prospects

G. Ambrosi INFN Perugia

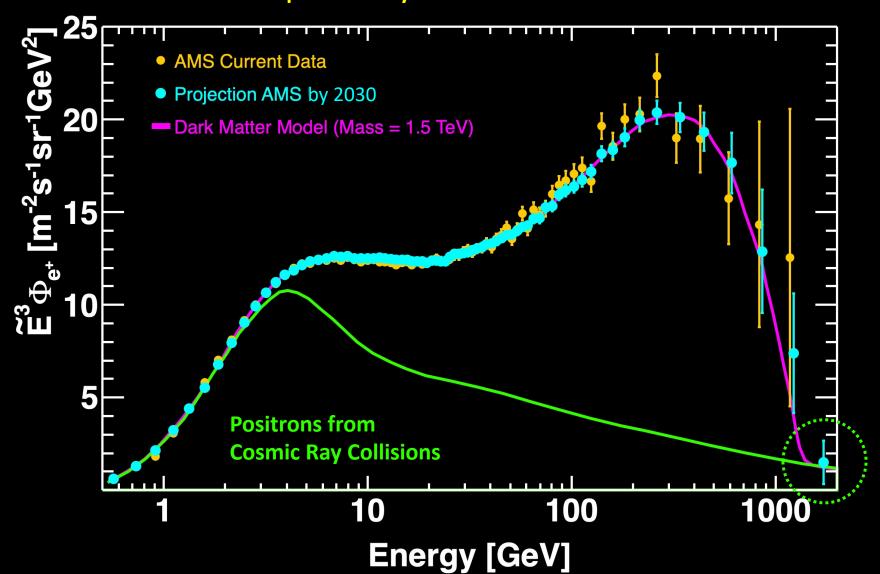


By 2030, AMS will extend the energy range of the positron flux measurement from 1.4 to 2 TeV and reduce the error by a factor of two compared to current data



Determination of the Origin of Cosmic Positrons by 2030

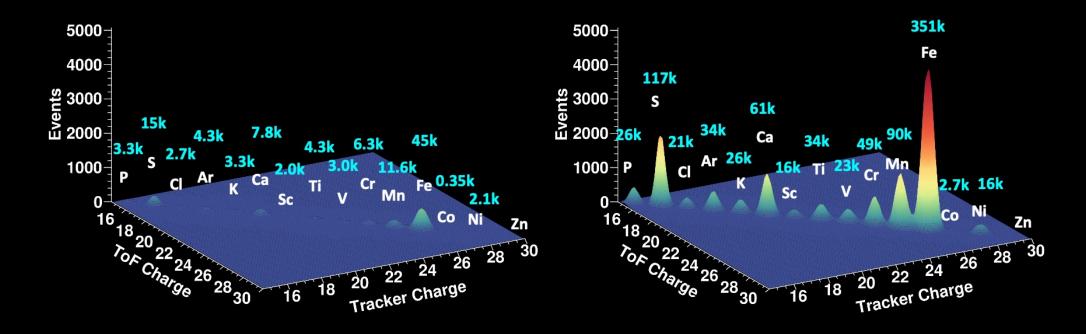
AMS will ensure that the measured high energy positron spectrum indeed drops off quickly and, at the highest energies, the positrons only come from cosmic ray collisions as predicted by dark matter models



Cosmic Ray Nuclei by 2030

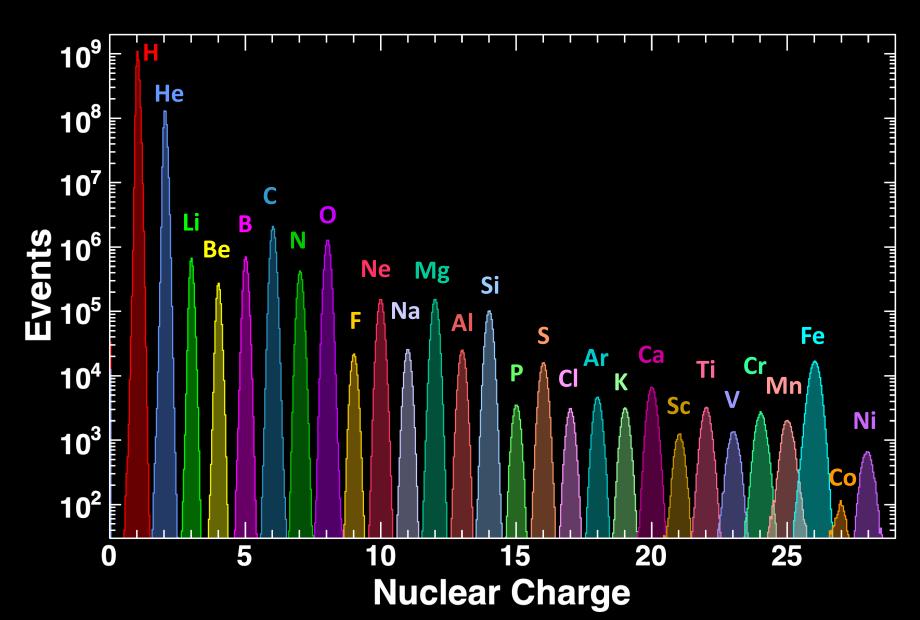
Current Data 2016-

With upgrade



The upgrade of AMS will provide accurate measurements of high-Z (Z≥15) nuclei, particularly with R>35 GV (where no data exists)

Cosmic Ray Nuclei by 2030

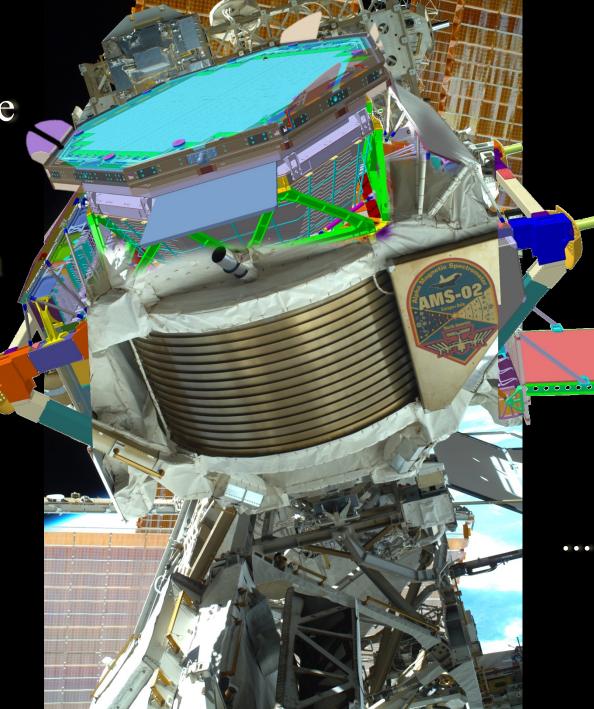


AMS-02 L0 upgrade ... in 1 slide C. Gargiulo

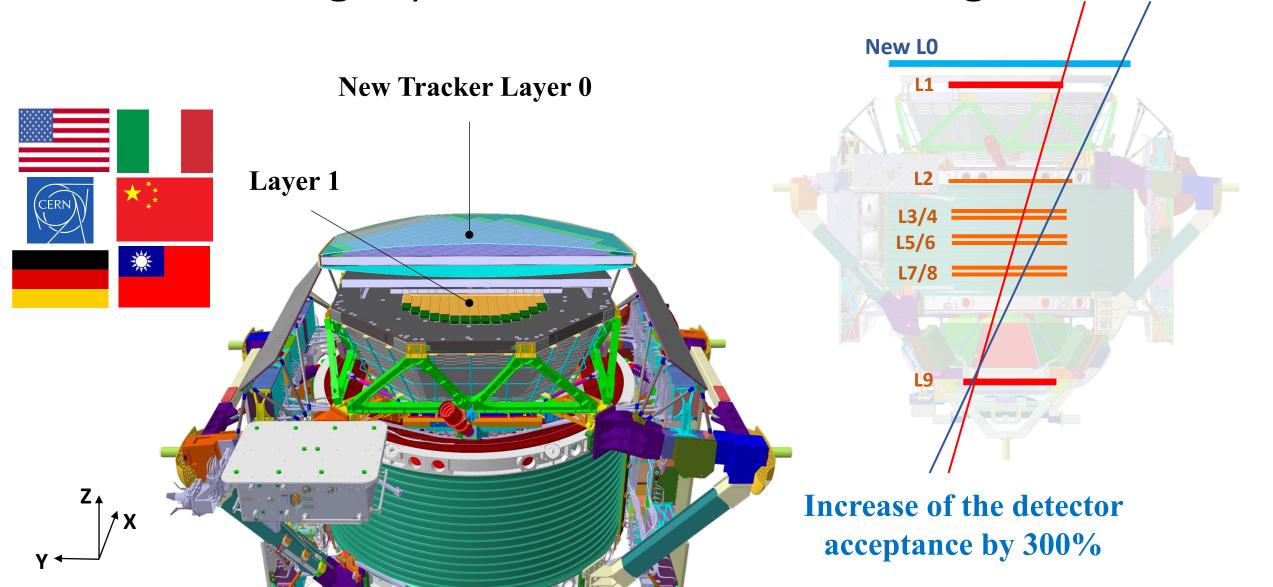
AMS-02 L0

upgrade ... in 1 slide

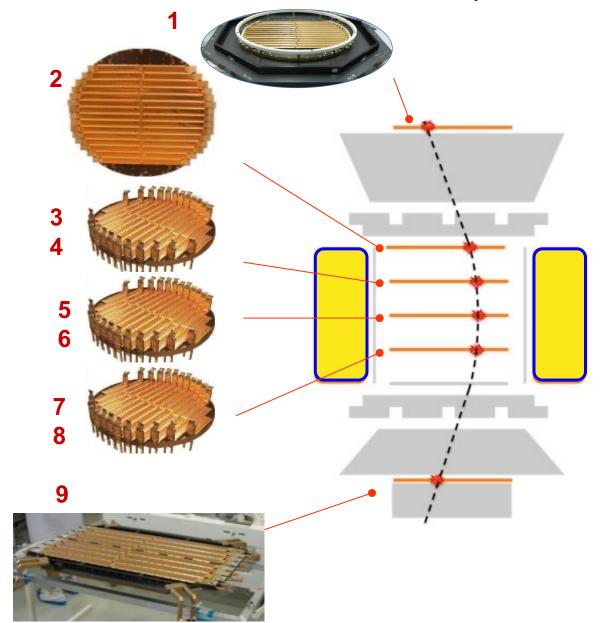
AMS-02 decision in December 2021 ...

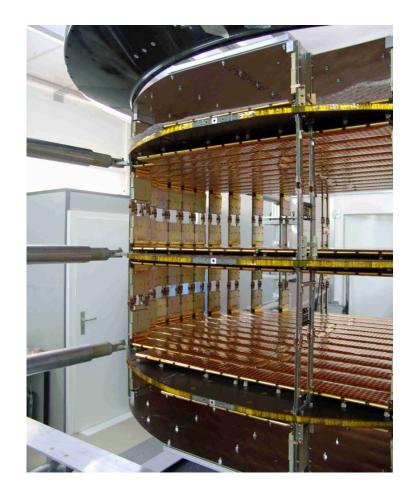


... ready for flight January 2025 LO tracking layer added to the existing ones

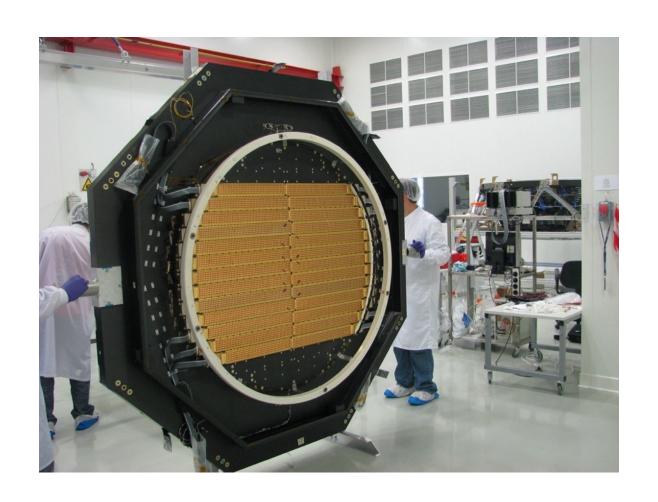


The layout of the AMS-02 Tracker



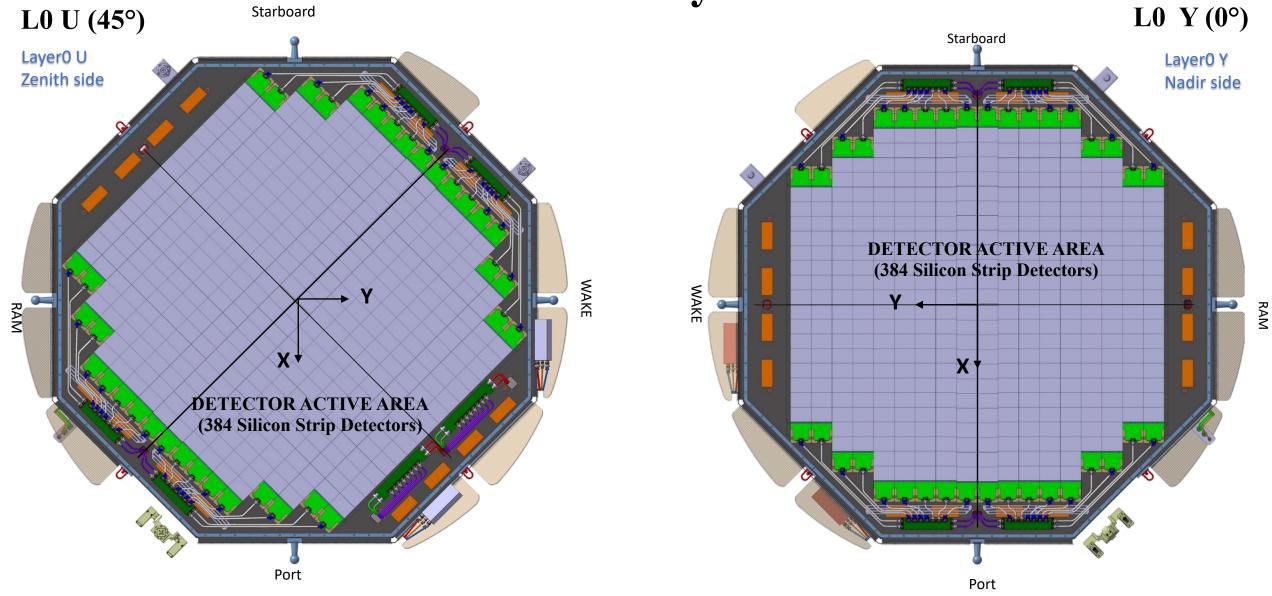


Layer 1





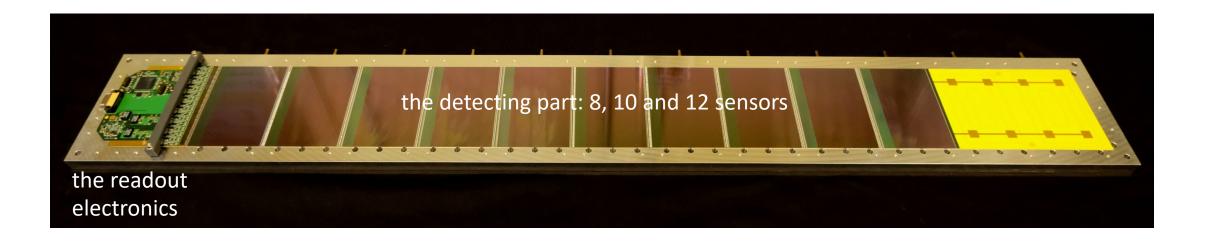
Plane layout

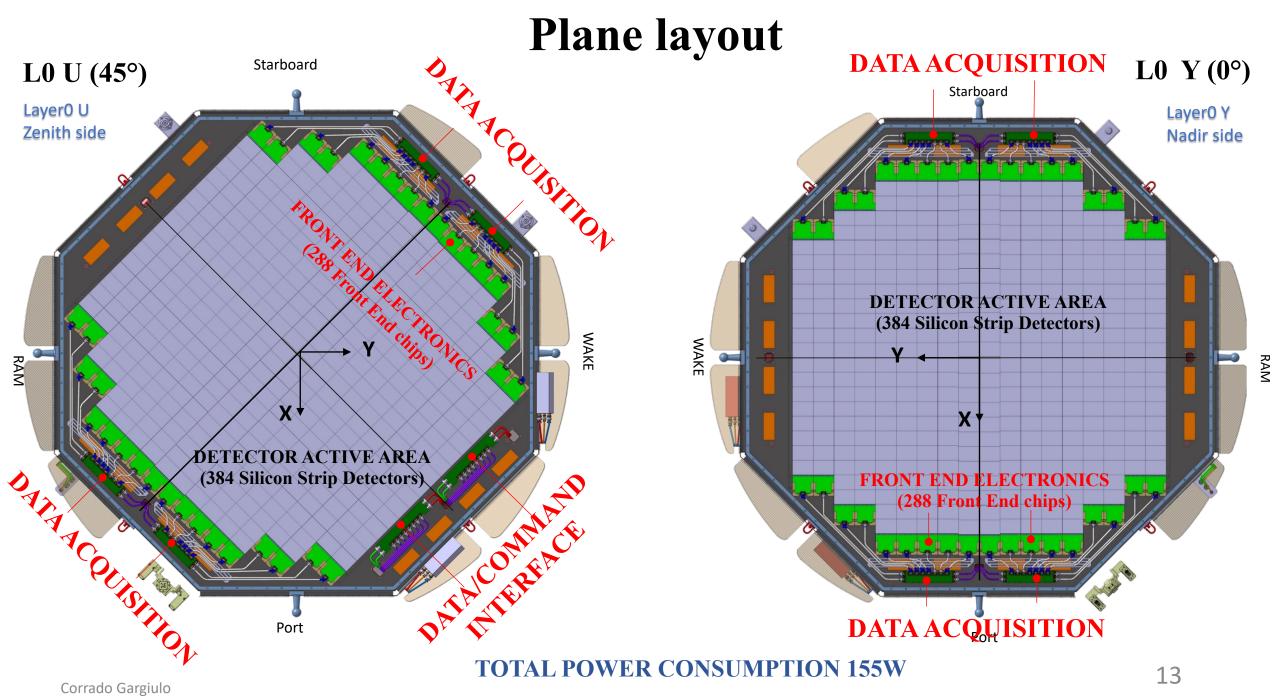


11

The Ladder

• the basic element of the detector

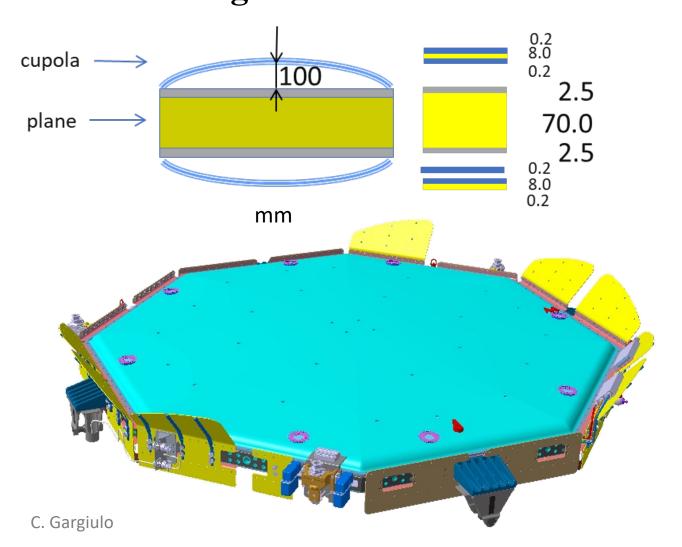




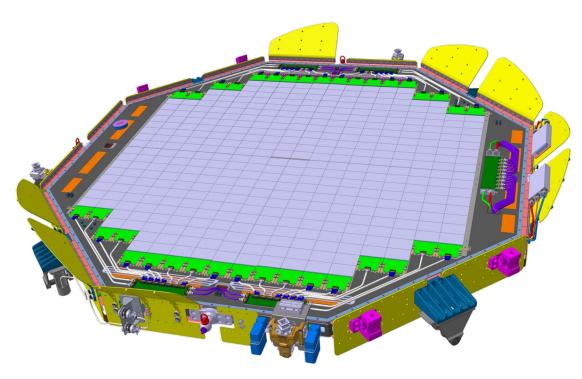
Dimensions

~2.6m diameter

~ 30 cm thick ~ 250 kg



Layer0 overall layout



Cupola Top & Bottom

EMI Shielding

Thermal radiator

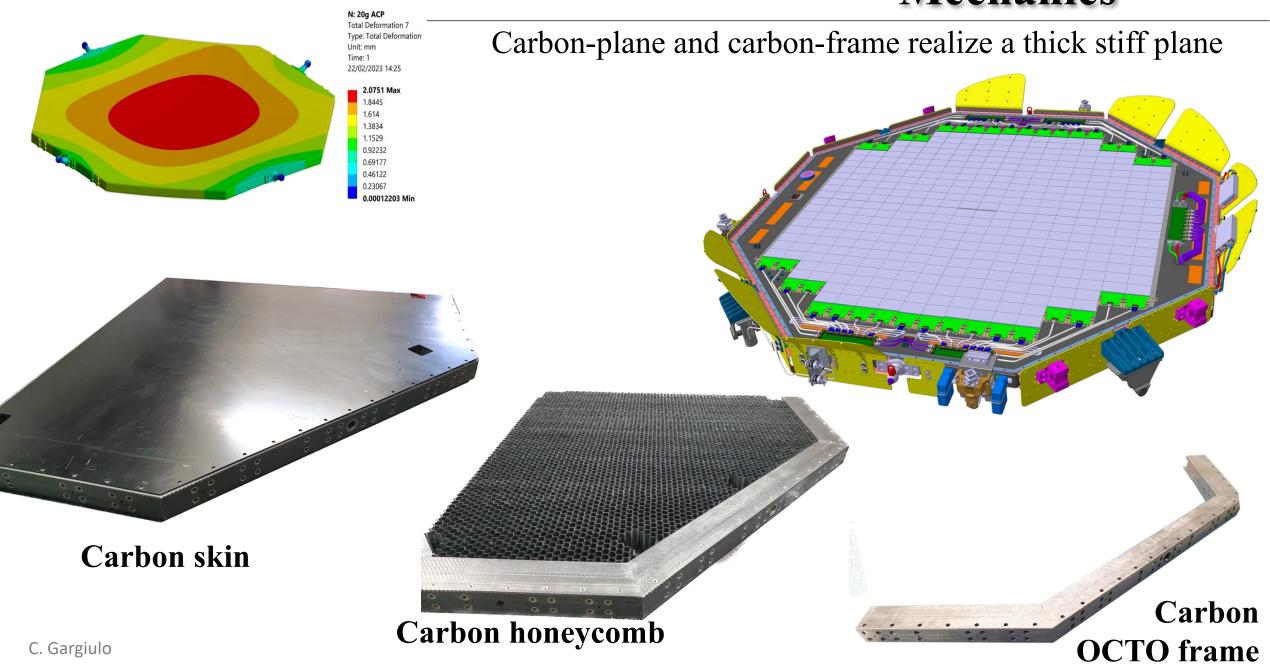
Light shielding

in addition, Cupola Top

Minimum Material Budget

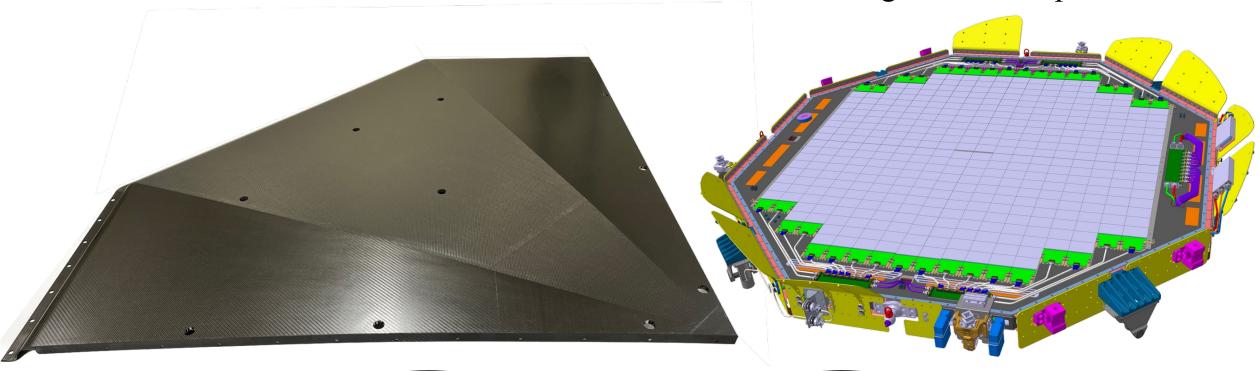
Micrometeorite shielding

Mechanics

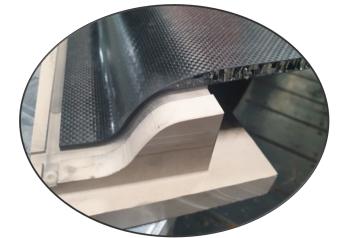


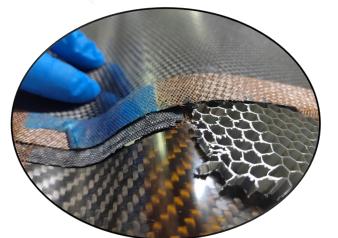
Mechanics

Light carbon cupola



Transition from sandwich to bolted interface





Aluminum honeycomb Carbon skin Copper net for EMI





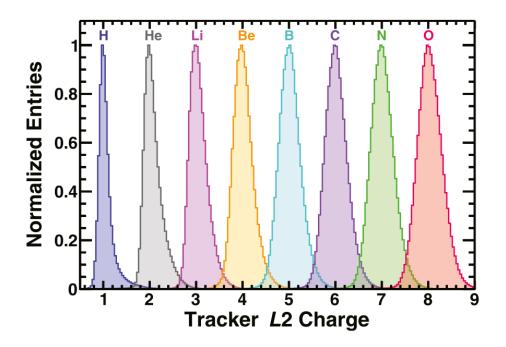


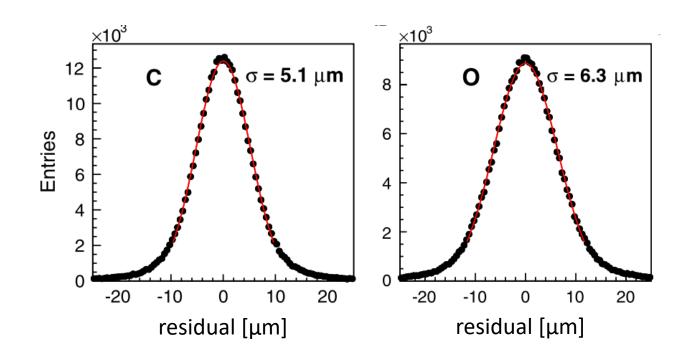
first full-size prototype (IHEP, Beijing)



current Tracker performance

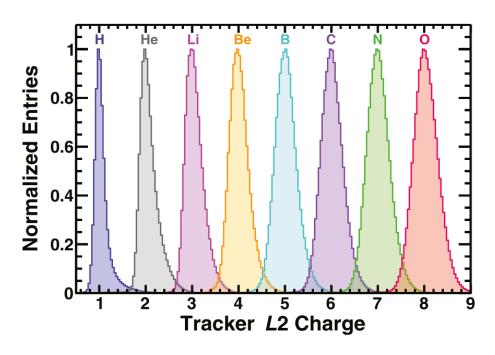
charge measurement



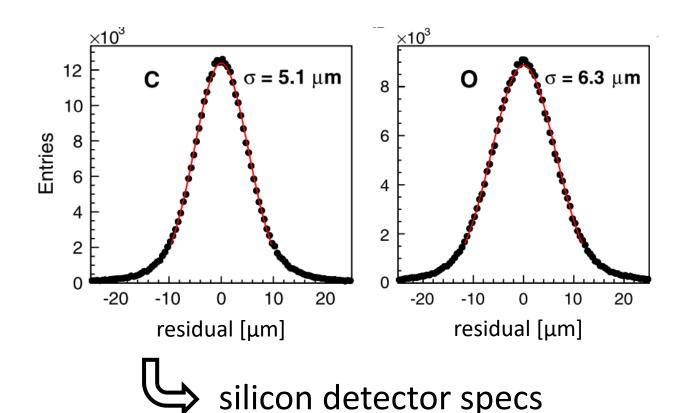


current Tracker performance

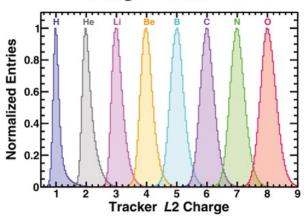
charge measurement

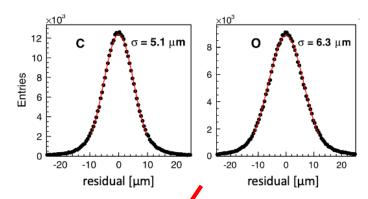


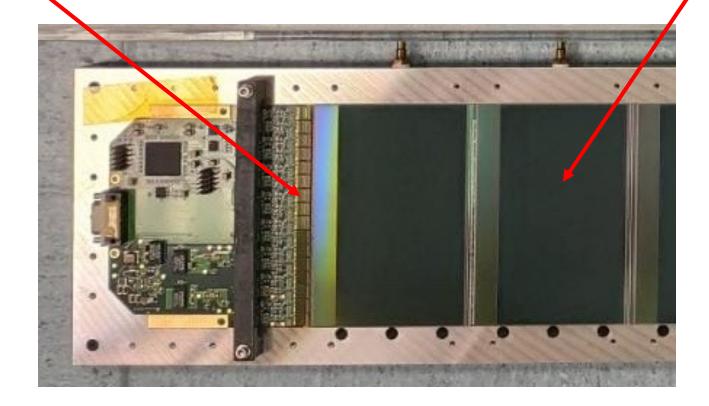
electronics specs



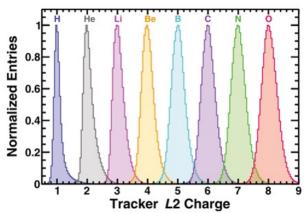
charge measurement

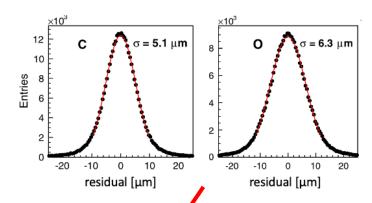


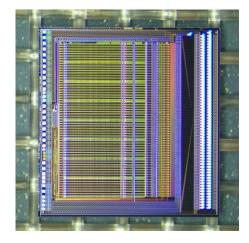




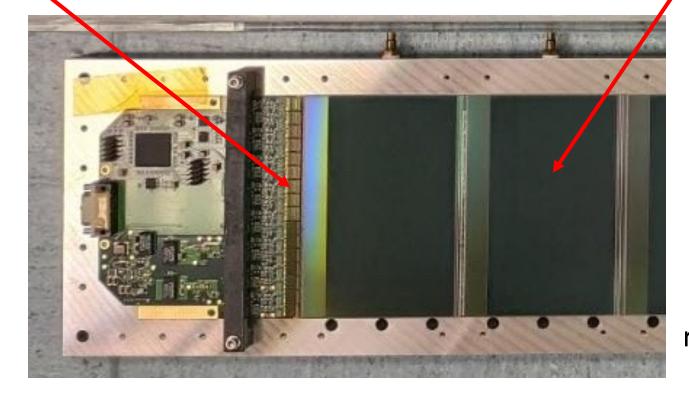
charge measurement

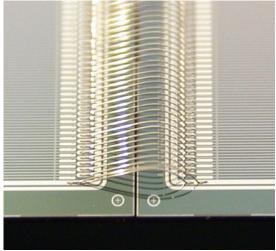






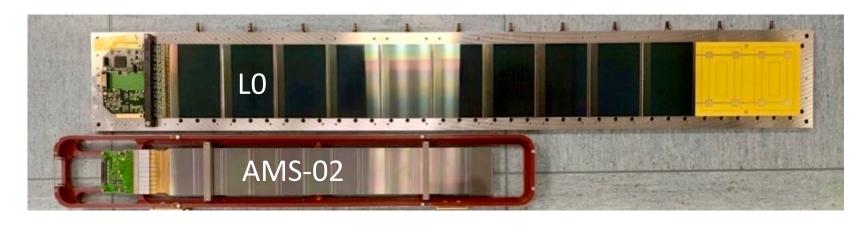
VA1140 high dynamic range

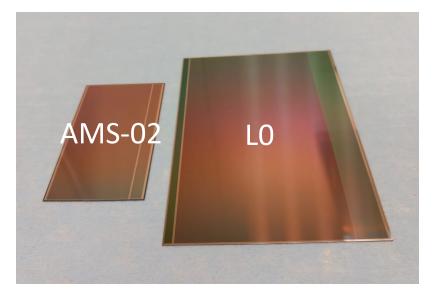




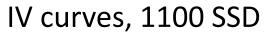
strip pitch 27.25 μm readout pitch 109 μm

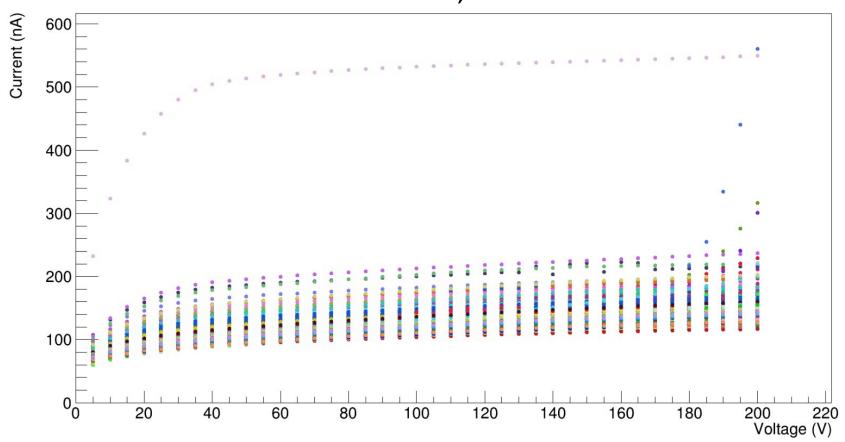
AMS-02 vs L0 ladders



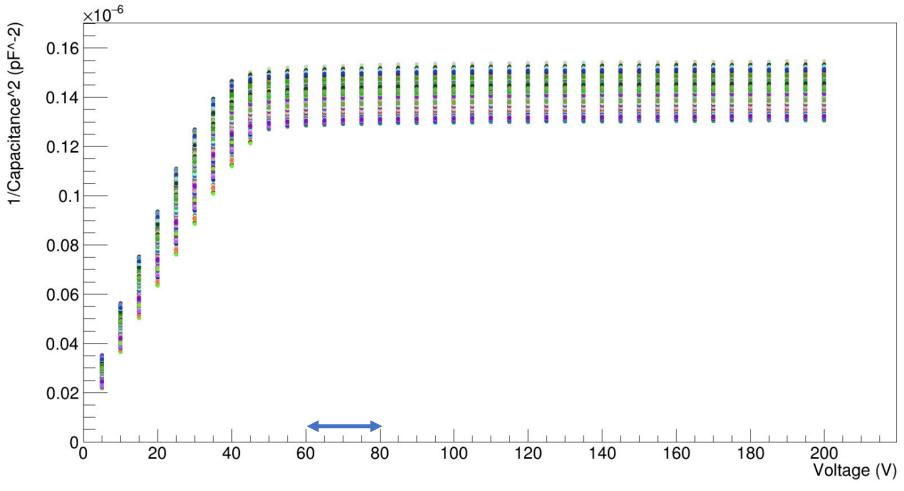


SSD quality ... is excellent



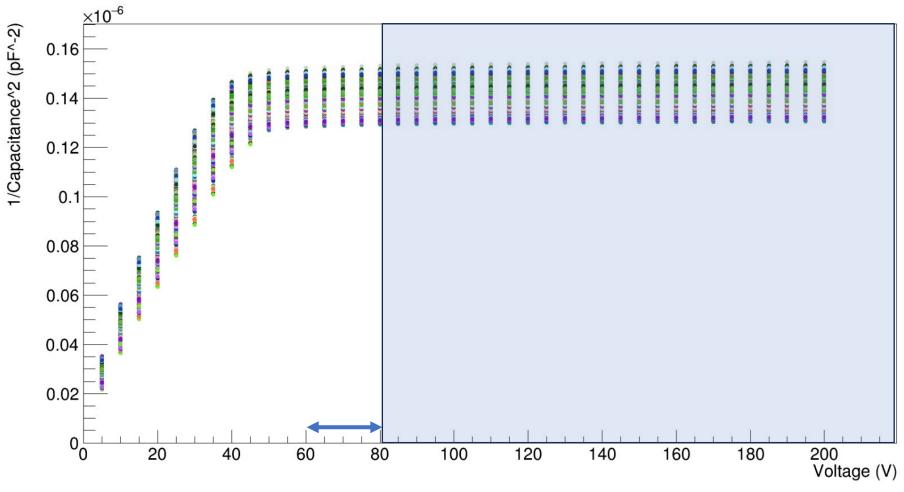


SSD quality ... is excellent



total capacitance define the bias voltage to operate the detector

SSD quality ... is excellent



total capacitance define the bias voltage to operate the detector

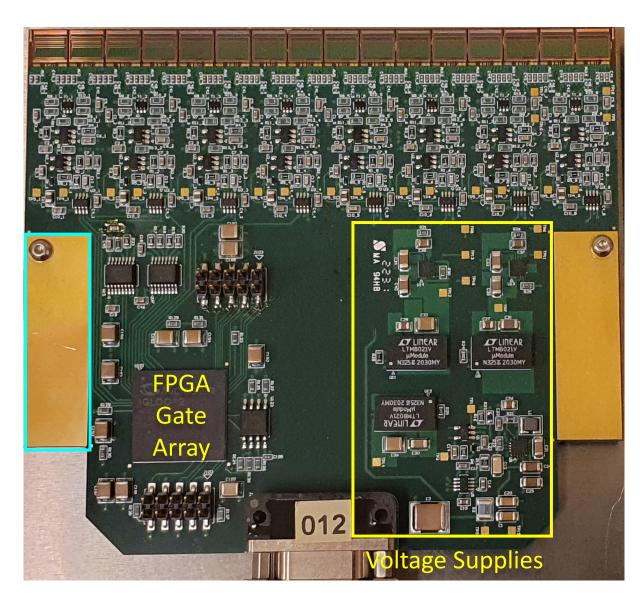
LO Electronic Front End (LEF)

8 Amplifiers

8 14-Bit, 2.5Msps, Serial Sampling ADCs, 4096mV

Thermal Strips

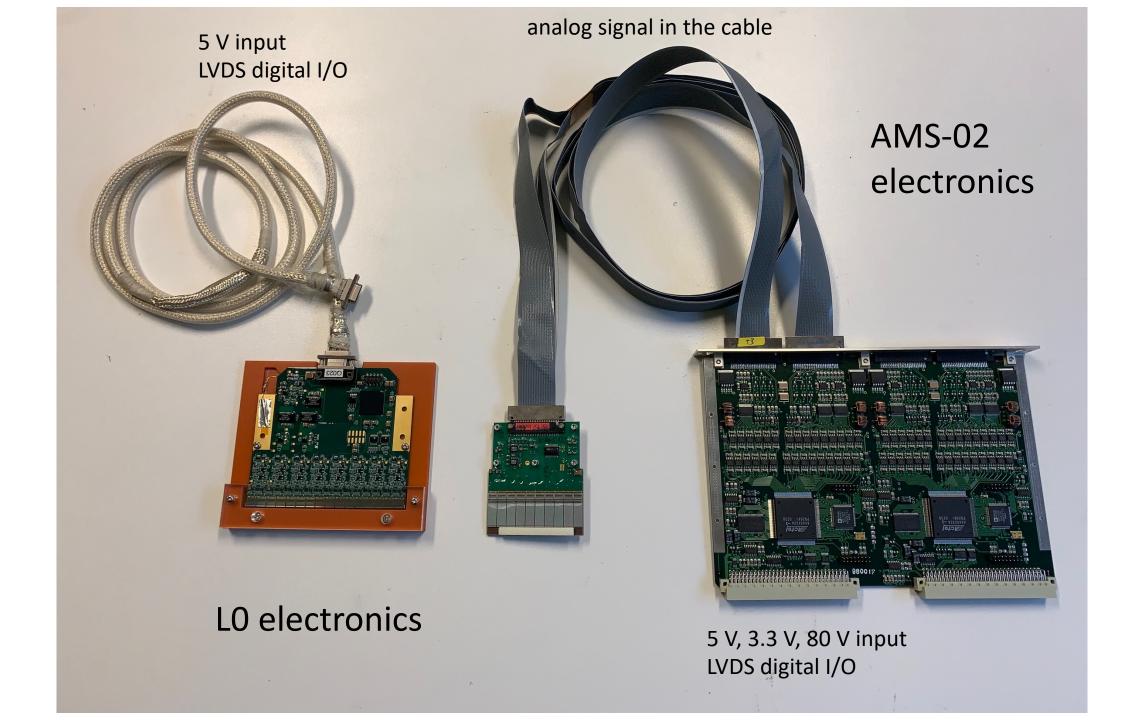
- 1. Heat path to radiators.
- 2. Ground path to chassis

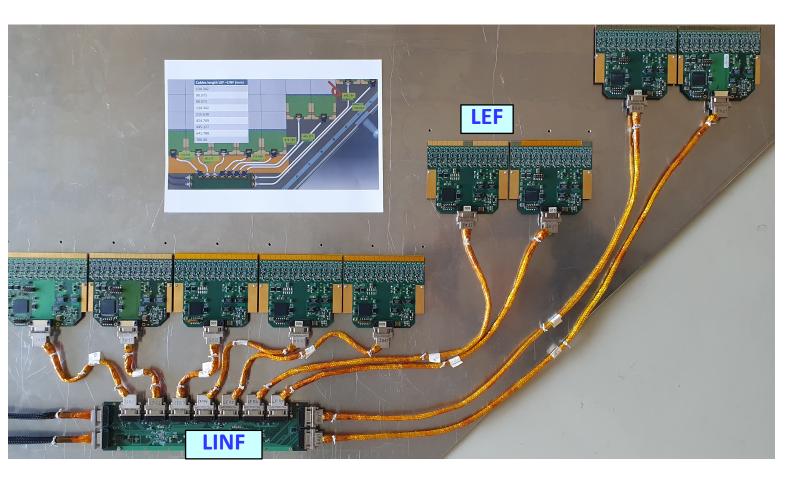


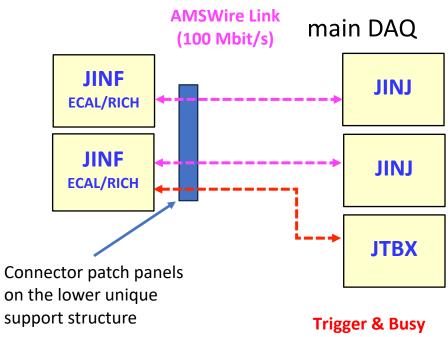
16 IDE1140, 1024 strips

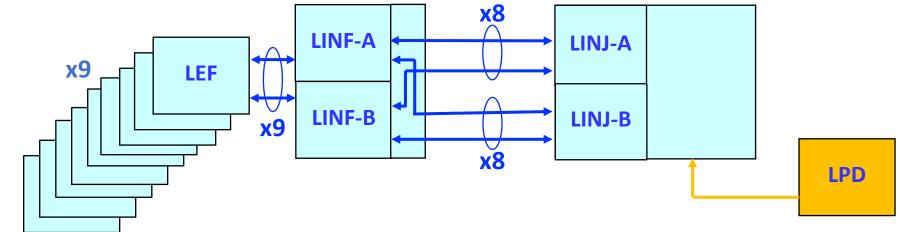
Each IDE1140 ("VA")

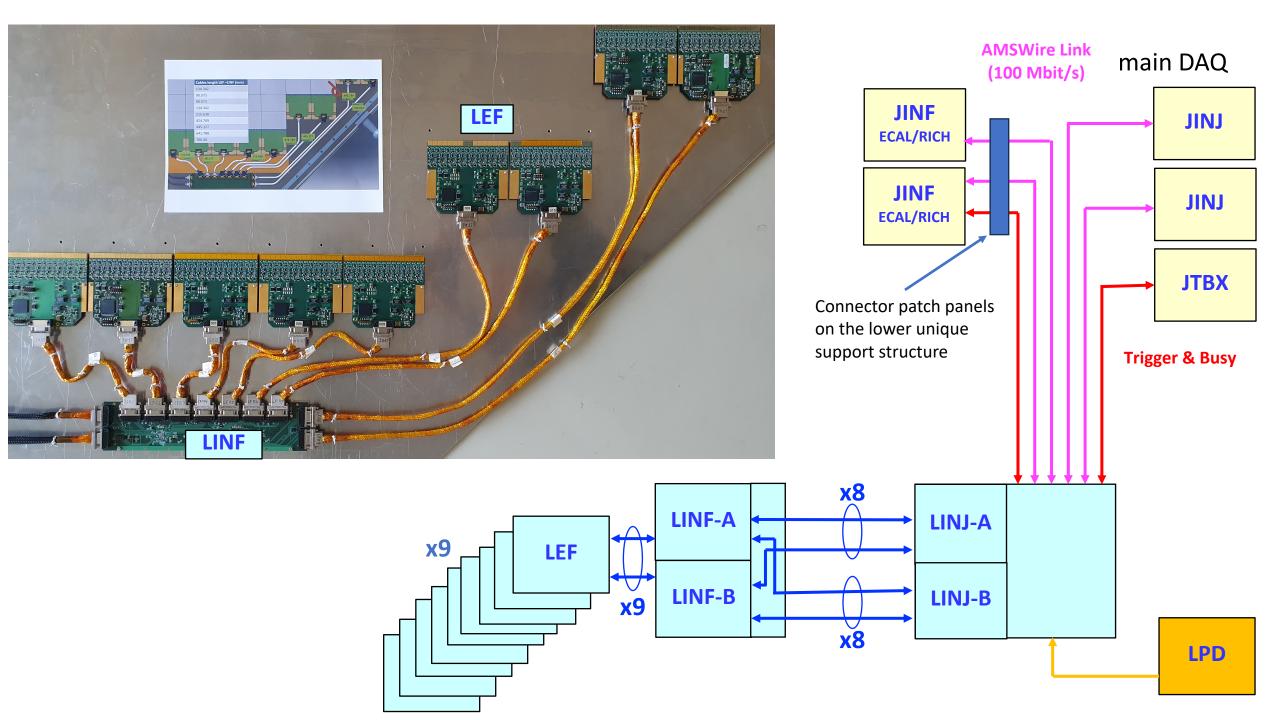
- 64 channels charge amplifier/ shaper.
- Sample and hold.
- 64 channels analog multiplexor.
- 2.6 uA per 1 fC differential current output

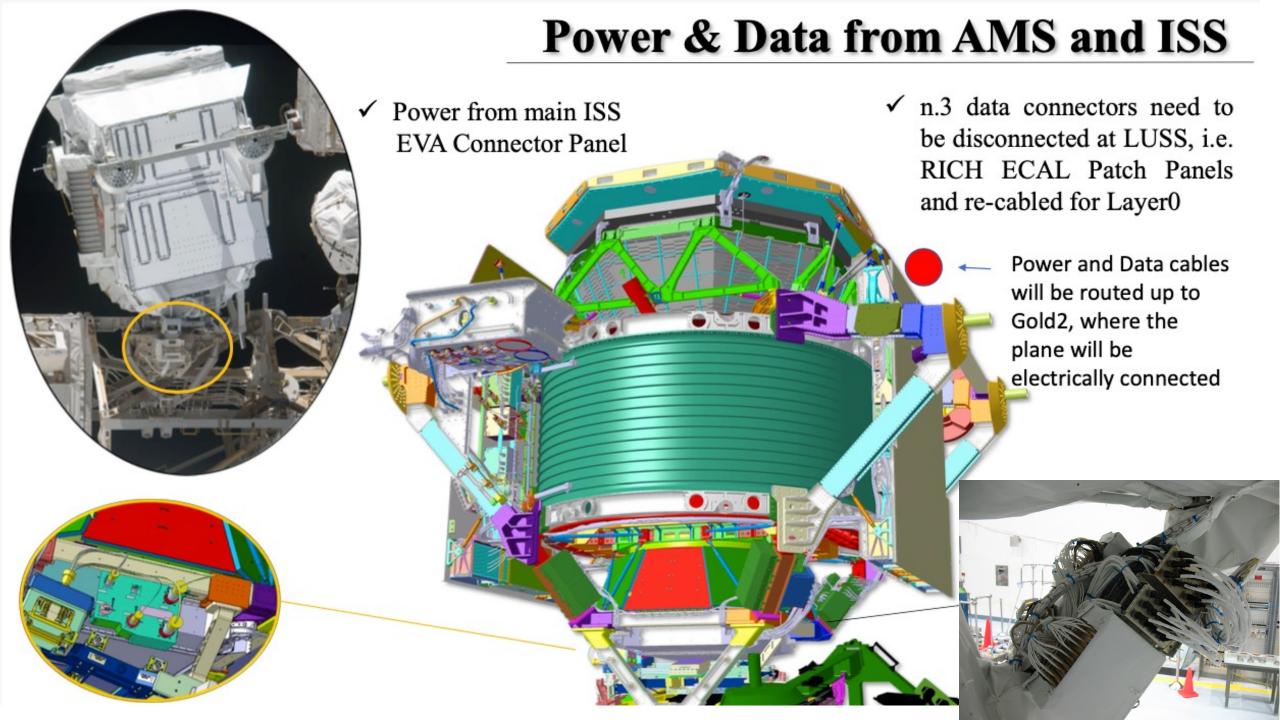


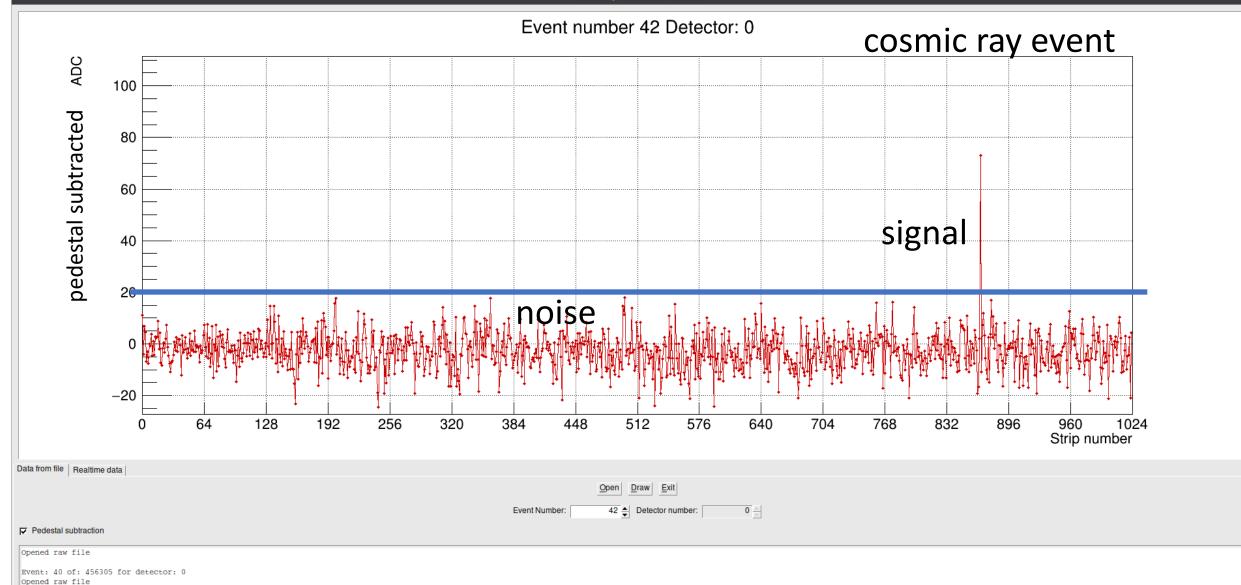












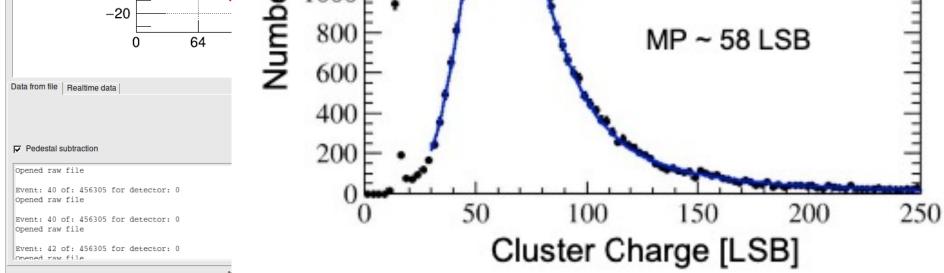
Event: 40 of: 456305 for detector: 0

Event: 42 of: 456305 for detector: 0

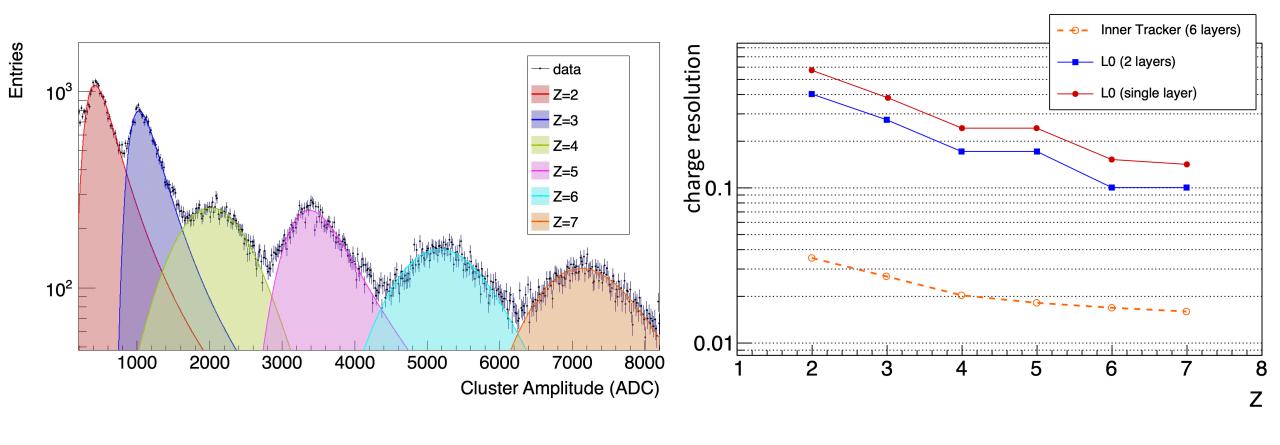
Opened raw file

Onened raw file

G_MUONS/0003/000_662_conv.root



detector performance (ion beam test)



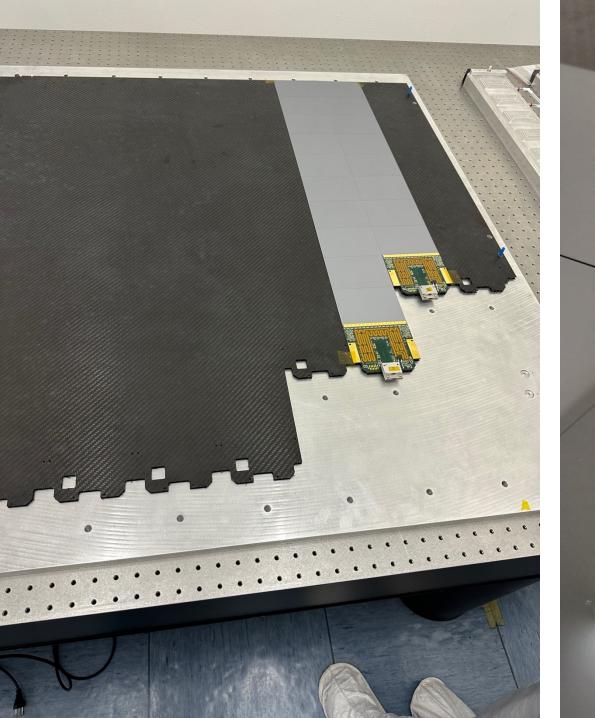
Where are we?

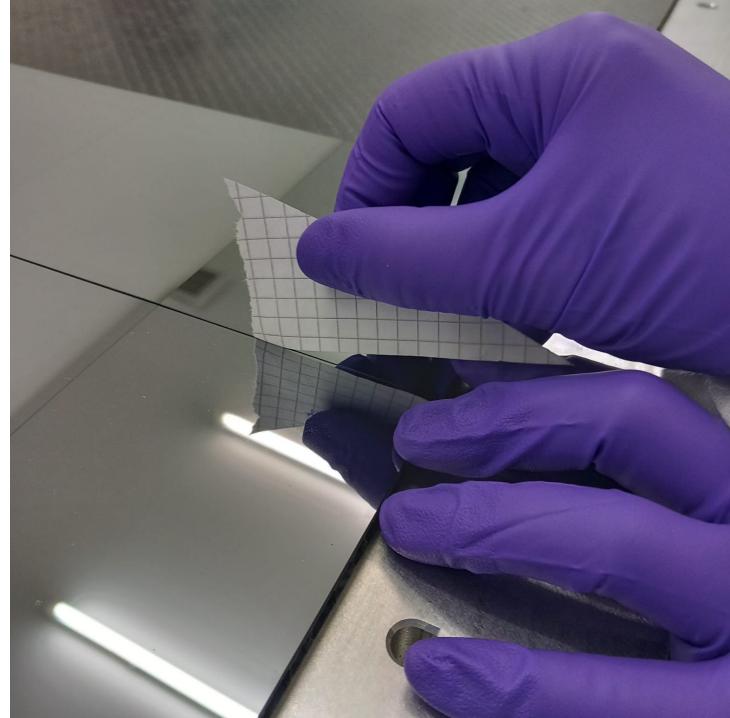
- detector components are all available
- first ladder assembly is done
- performance are being evaluated (preliminary results are according to specifications)

space qualified activity



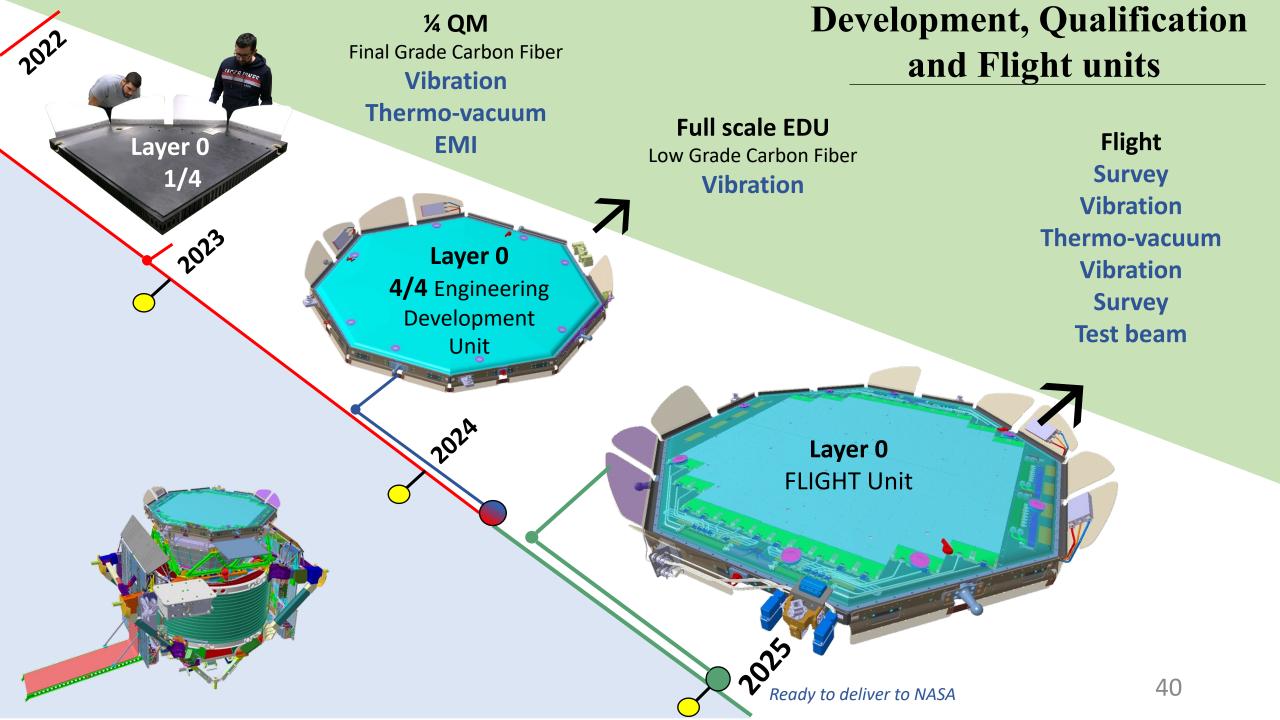






Where are we?

- detector components are all available
- first ladder assembly is done
- performance are being evaluated (preliminary results are according to specifications)
- first integration test is successful
- beam test next August (proton) and October (ions)
- Qualification Model production is started
- Flight Model production is on schedule, start in November 2023



INFN/UniPG laboratory for space qual Terni, Italy

