

# Mini-EUSO on board the ISS: mission status and prospects

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*ASAPP 2023 Conference – Perugia, June 19-23, 2023*



# *The JEM-EUSO program*

**1. EUSO-TA:** *Ground detector installed in 2013 at Telescope Array site: currently operational*

## 2. EUSO-BALLOONS:

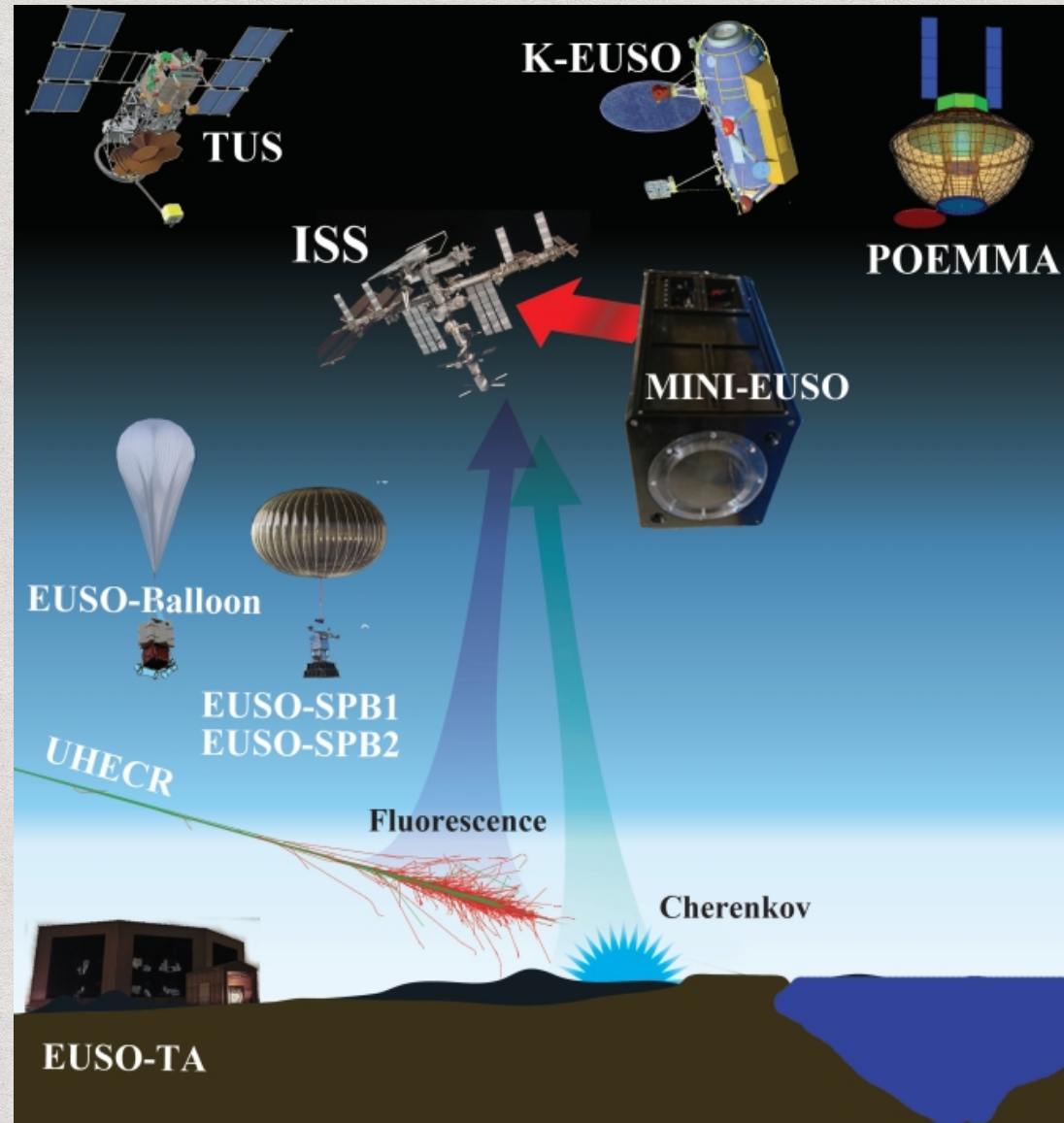
- **2014: EUSO-Balloon Timmins, Canada**
- **2017: EUSO-SPB, NASA, Ultra long duration flight**
- **2023: EUSO-SPB2, NASA, Ultra long duration flight**

### 3. TUS (2016): *free-flyer on Lomonosov Russian Satellite*

#### 4. Mini-EUSO (2019): ISS, Beyond Mission (L. Parmitano)

## 5. K-EUSO (2026+): ISS, Phase A, Russian Space Agency

## 6. POEMMA (2030+): *NASA twin free-flyer*



JEM-EUSO collaboration

16 Countries, 93 Institutes, 351 people





# Mini-EUSO / UV-Atmosphere

*Multiwavelength Imaging New Instrument for the Extreme Universe Space Observatory*

**Installed on a UV-transparent window on board the ISS**

Weight: 35 kg

Power consumption: 60 W

Dimensions: 37x37x62 cm<sup>3</sup>



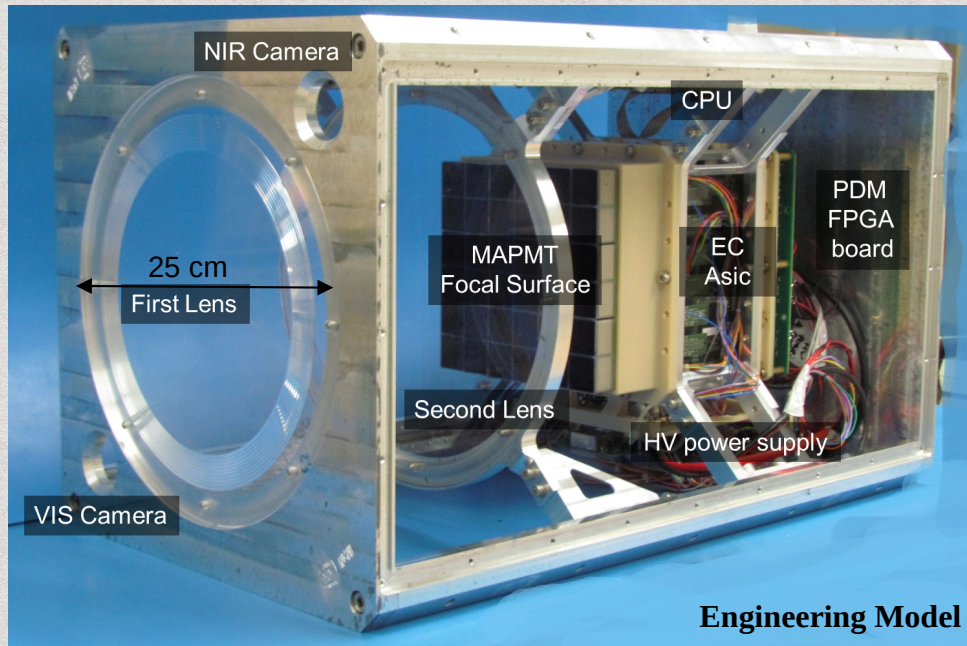


# Mini-EUSO / UV-Atmosphere

*Multiwavelength Imaging New Instrument for the Extreme Universe Space Observatory*

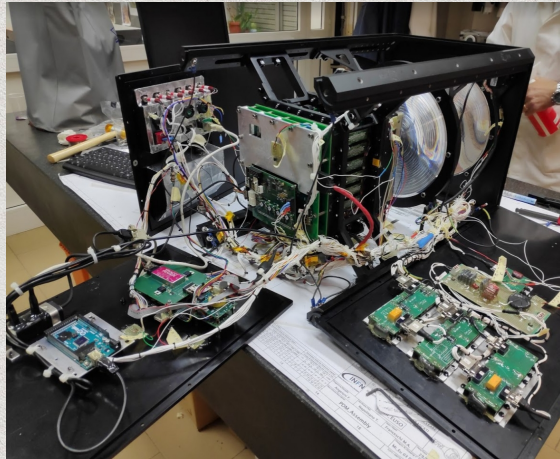
Mini-EUSO main sensors:

- Ultraviolet telescope with Fresnel lenses (48\*48 pixels, FoV= 44 deg,  $\sim 320 \times 320 \text{ km}^2$ , 2.5  $\mu\text{s}$  and above)
- Near Infrared camera (1280\*960 pixels, FoV=33.2\*24.8 deg,  $231 \times 174 \text{ km}^2$ , 1s)
- Visible camera (1280\*960 pixels, FoV=33.2\*24.8 deg,  $231 \times 174 \text{ km}^2$ , 4s)
- SiPM (8\*8 pixels) and UV sensors





# Integration and Tests of EM and FM @ INFN 2017-2019



INFN, Tor Vergata and LNF  
Integration & Mechanics

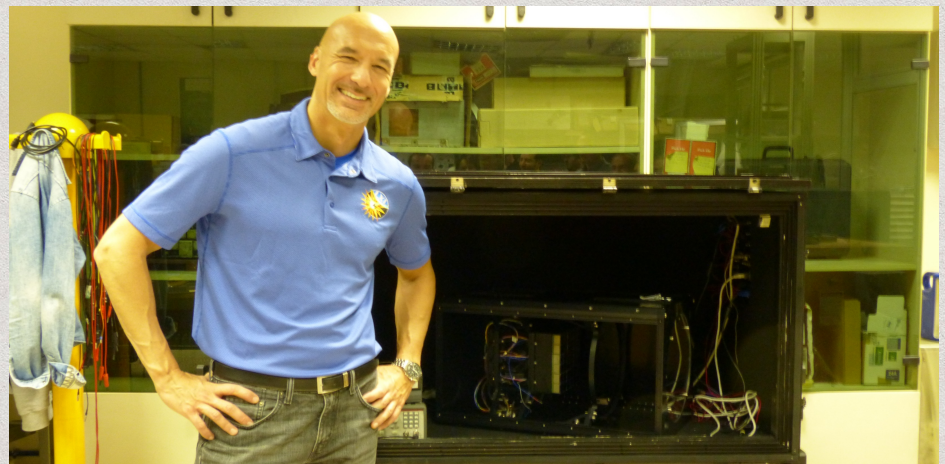
TUR-LAB, Univ and INFN Torino  
Test on EM and emulation of ISS



INFN, Tor Vergata  
Sky tests (FM)



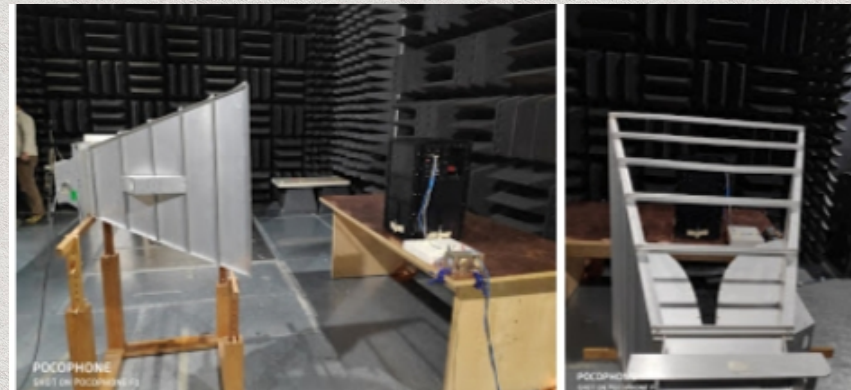
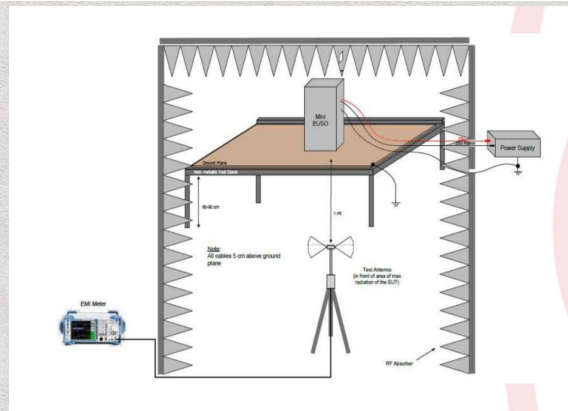
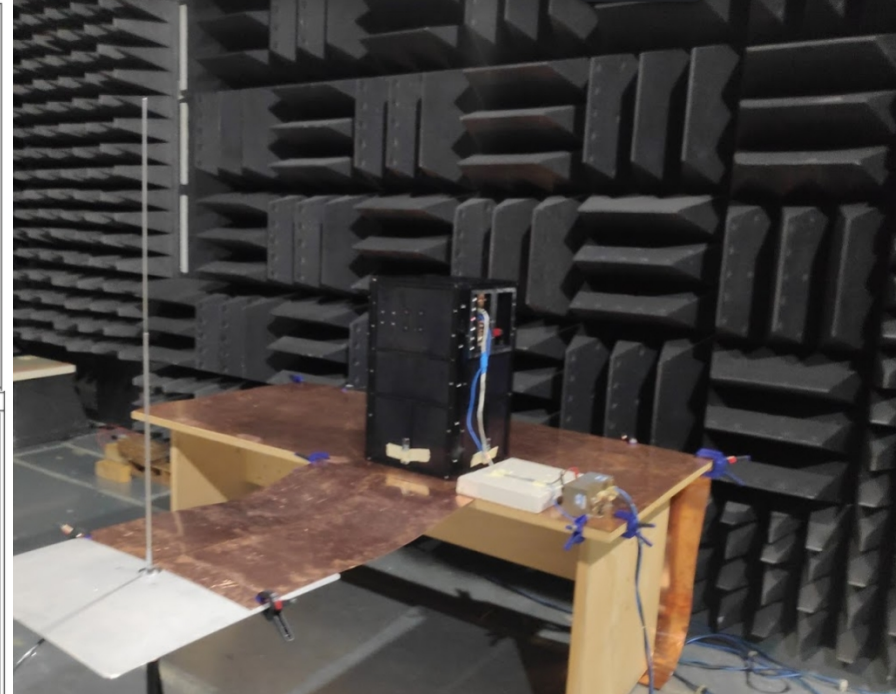
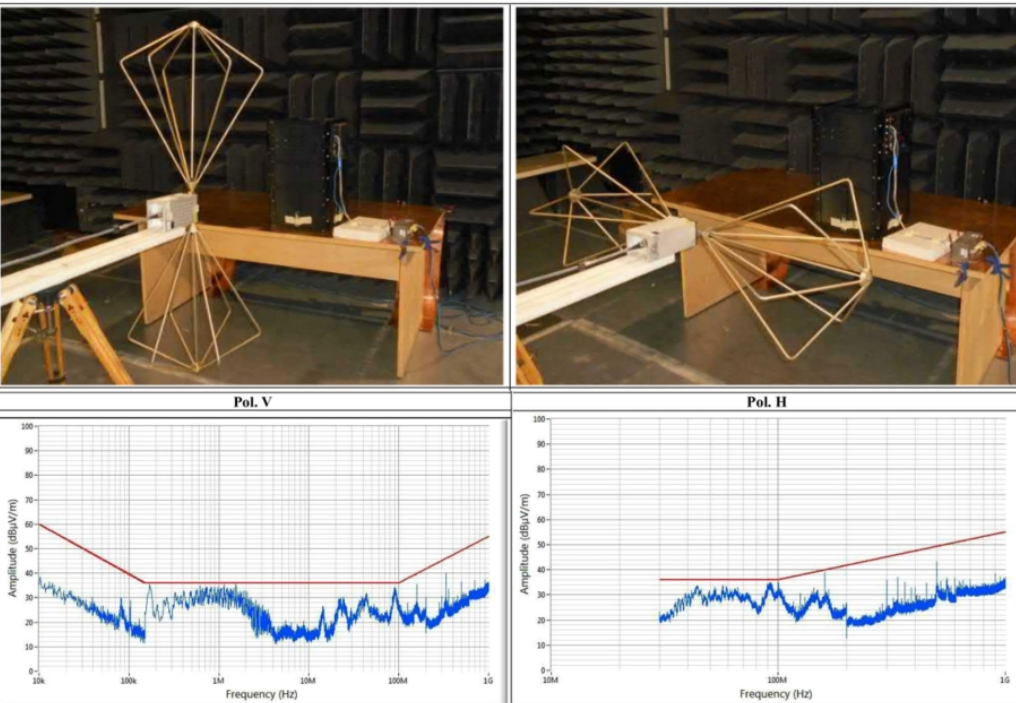
INFN, Tor Vergata  
Qualification and  
Acceptance tests



INFN, Tor Vergata  
L. Parmitano & Mini-EUSO in the black box

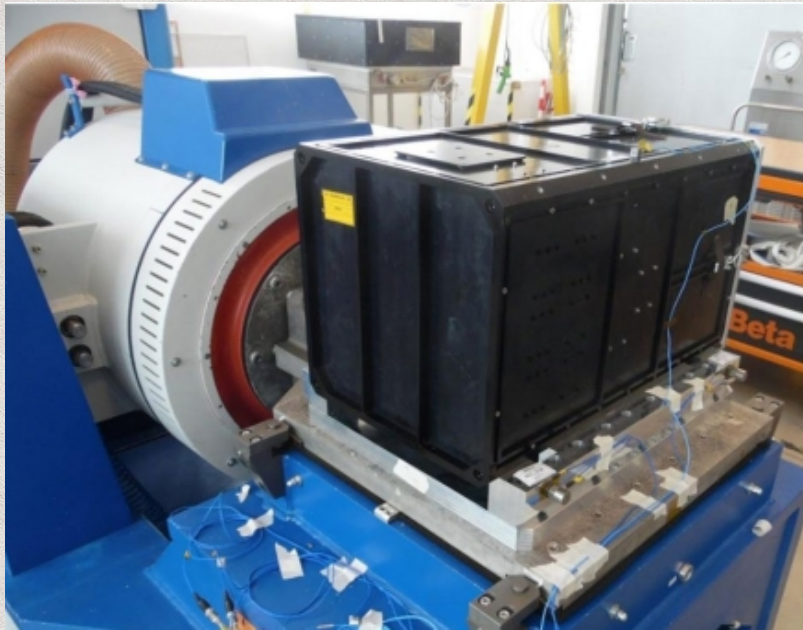


# EMI/EMC tests in Pisa (GSD Laboratories) - 14/1/2019

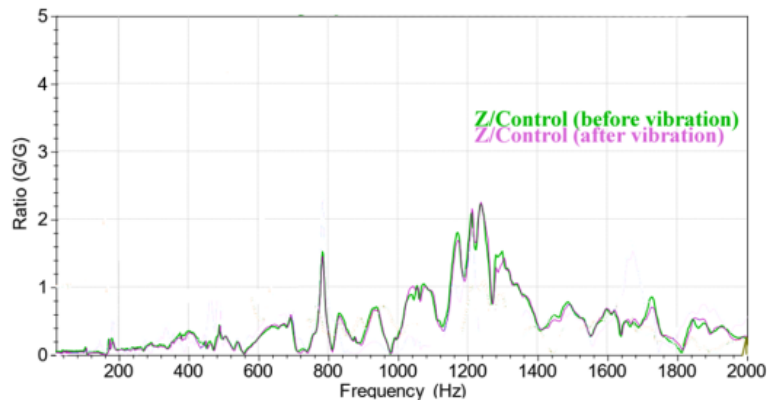




# Vibration tests in Torrita di Siena (MATE Laboratories) - 31/5/2019



QVT-Z-Resonance Survey Comparison



Displacement of natural frequencies lower than 5% → Test successful



*Integration, qualification, and launch of the Mini-EUSO telescope on board the ISS*

*L. Marcelli et al., Rendiconti Lincei (2023)*



# Acceptance tests in Baikonur and integration with Soyuz – August 2019



**Building 254, assembly of Soyuz/Progress**



Roll-out of Soyuz MS-14, 19/8/2019



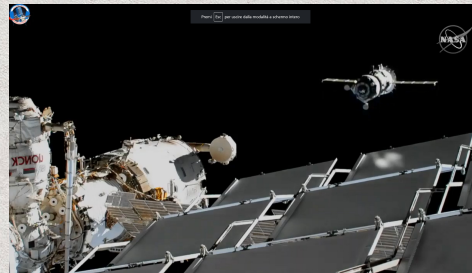
Launch, 22/8/2019



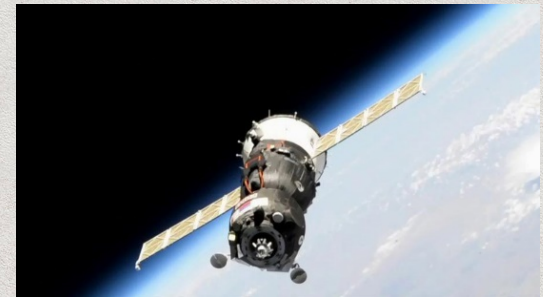
First docking, 24/8/2019 unsuccessful



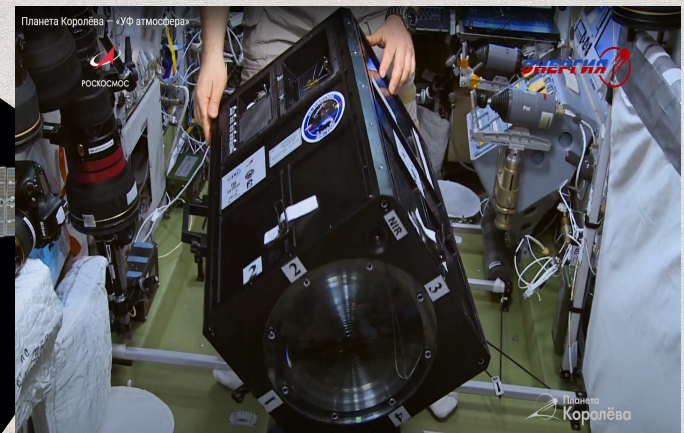
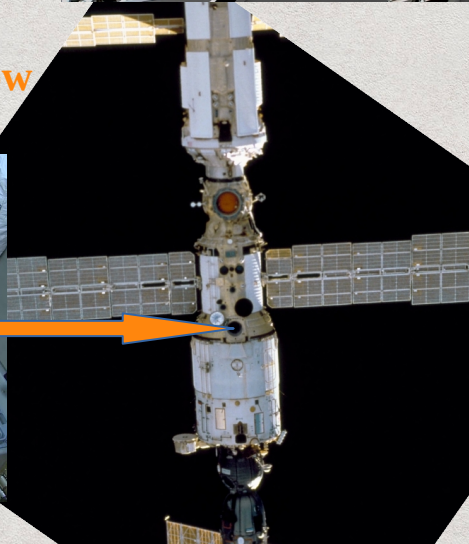
Relocation of MS-13 from Zvezda to Poisk



Second docking, 27/8/2019 successful

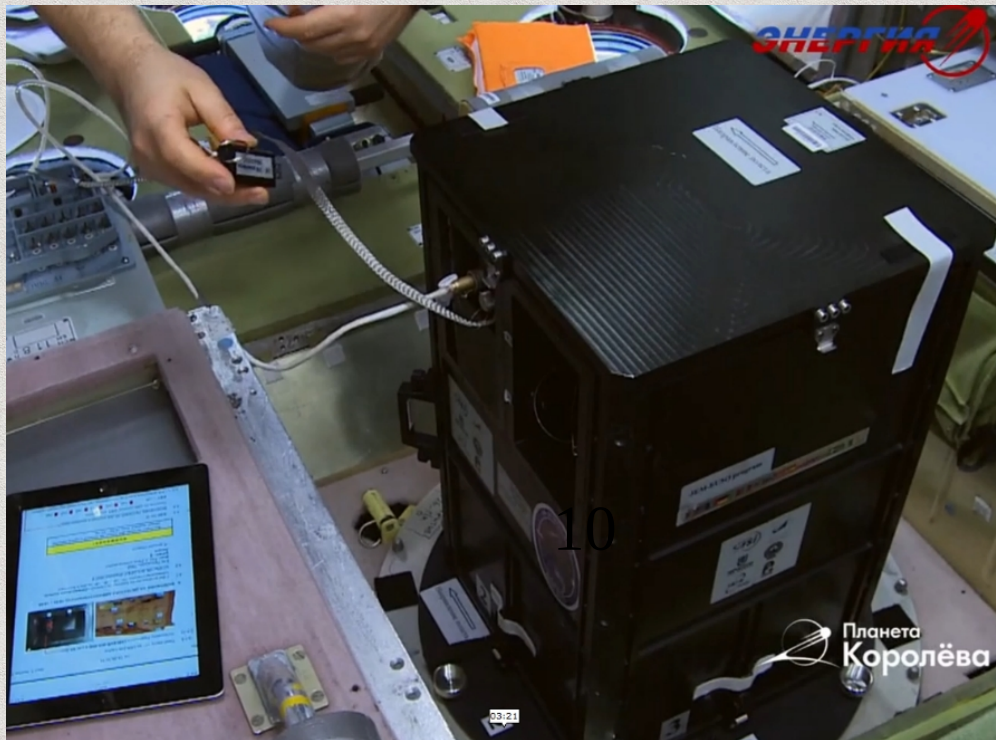


Installation - UV transparent window  
Zvezda module, 07/10/2019





# Mini-EUSO in-flight operations



10

## START OF SESSION

- Latch on the UV-transparent window
- Connect 27V power supply cable
- Connect grounding cable
- Insert and latch the USB stick
- Turn on the switch



## END OF SESSION

- Turn off the switch
- Remove and store USB stick
- Periodically copy of selectet files on station computer for later downlink
- Unlatch and store the instrument



# Mini-EUSO in-flight operations



Launch: August 2019

Pouch003: sessions 1-14  
Returned in April 2020

Pouch004: sessions 15-44  
Returned in October 2021

Pouch003\_v2: sessions 45-(?)  
Currently in use

Pouch004\_v2:  
Ready to be launched!

~3 sessions/month  
~12 hours/session

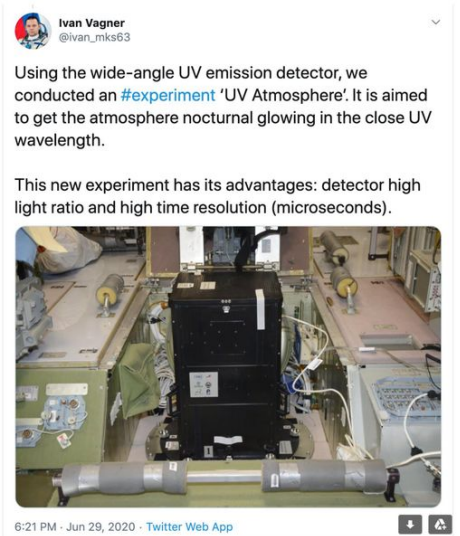
~20 GB stored each session  
(data subset downlinked  
via telemetry channel)



# Mini-EUSO on the ISS



Sergei Kud-Sverchkov



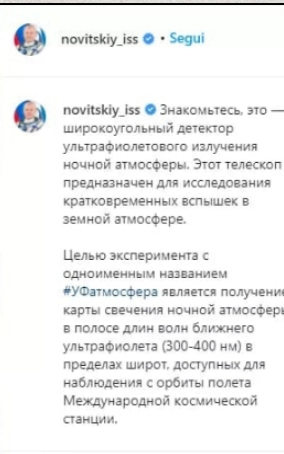
Twitter from I. Vagner



Oleg Artemyev



Twitter from O. Novitskiy



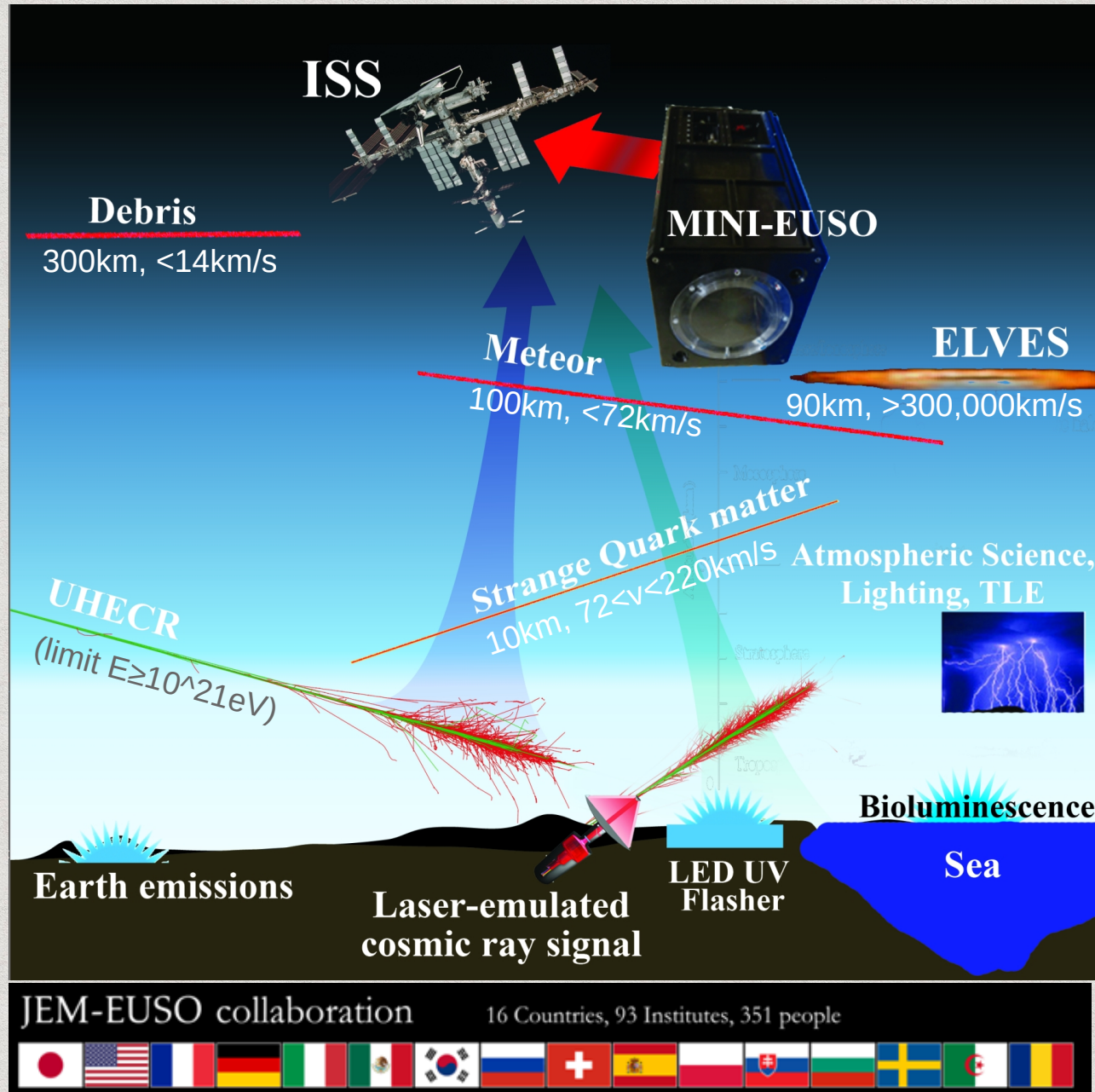
O. Skripochka: outreach video from ISS  
<https://www.youtube.com/watch?v=IXedBGVHc4o&t=62s>

85 sessions performed

>150 crew hours



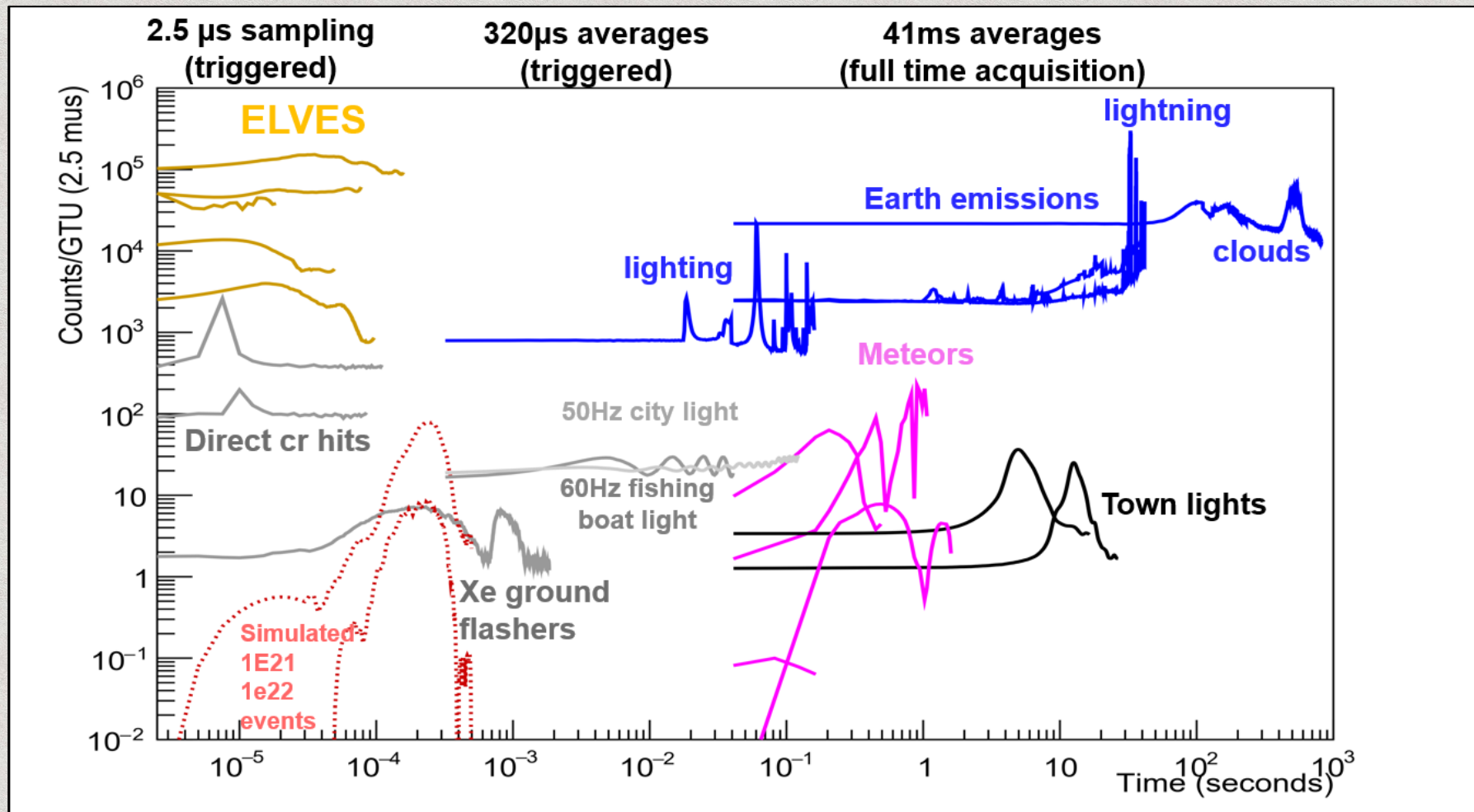
# Scientific Objectives





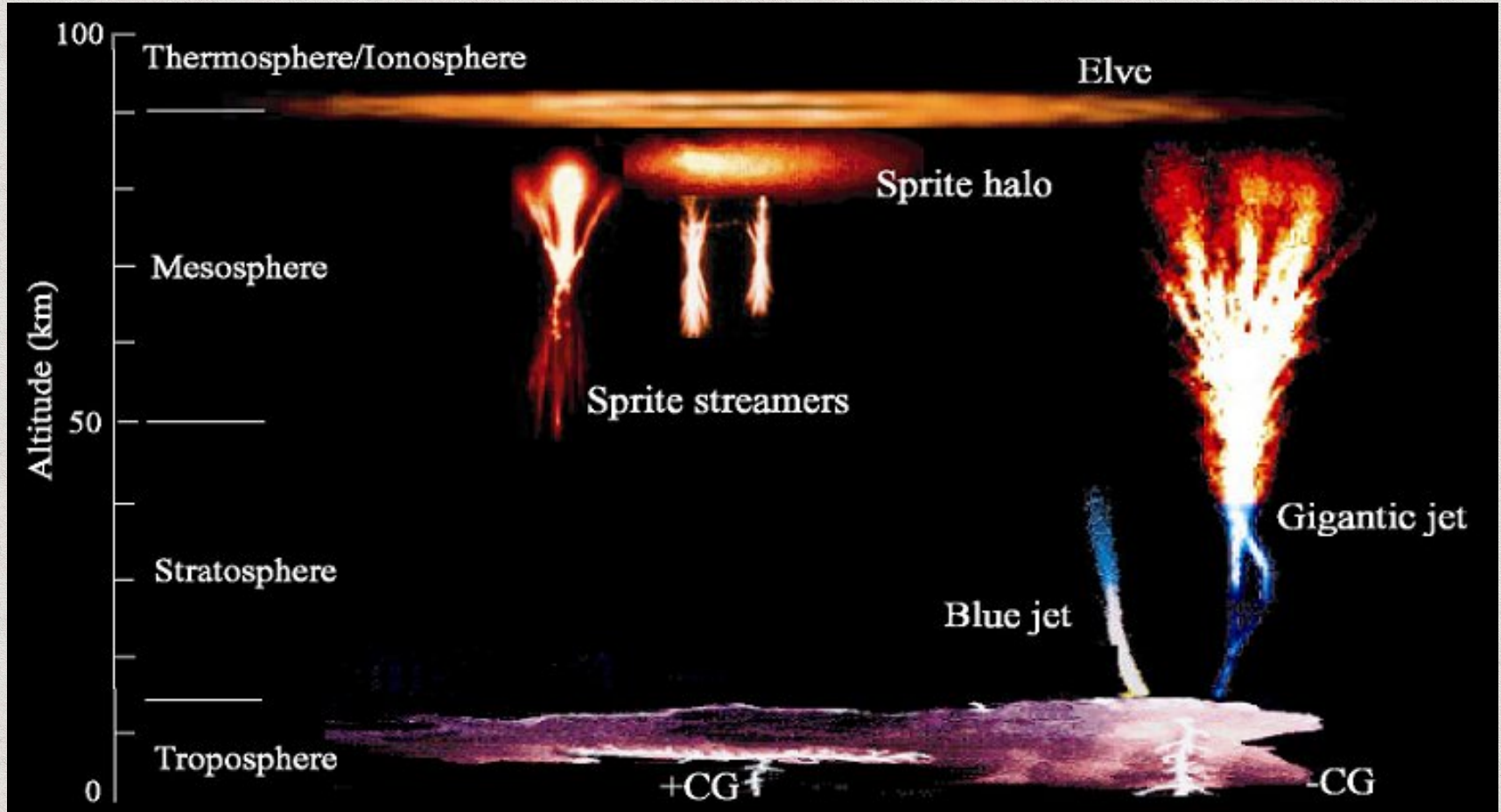
# Time profile of various events

2.5  $\mu$ s X 128      320  $\mu$ s X 128  
D1 acq. mode    $\longrightarrow$    D2 acq. mode    $\longrightarrow$    D3 acq. mode





# ELVES



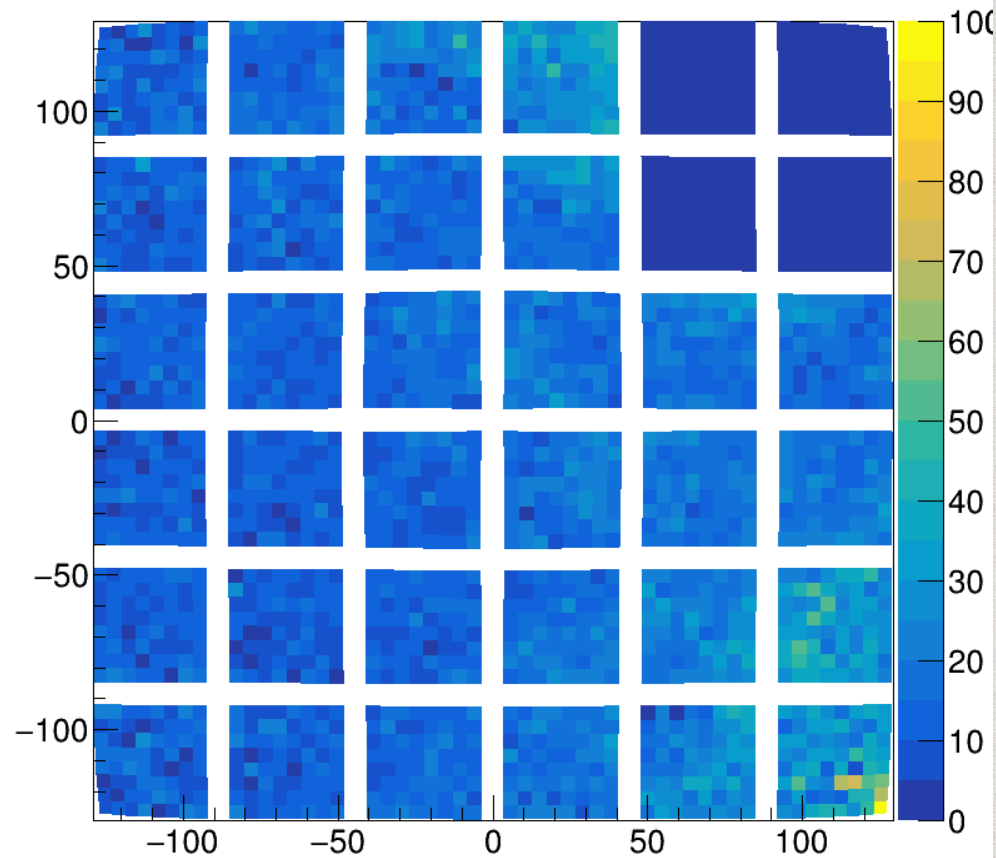
Apparently superluminal rings

Upper atmospheric lighting releases e.m.  
wave which heats the ionosphere



# ELVES

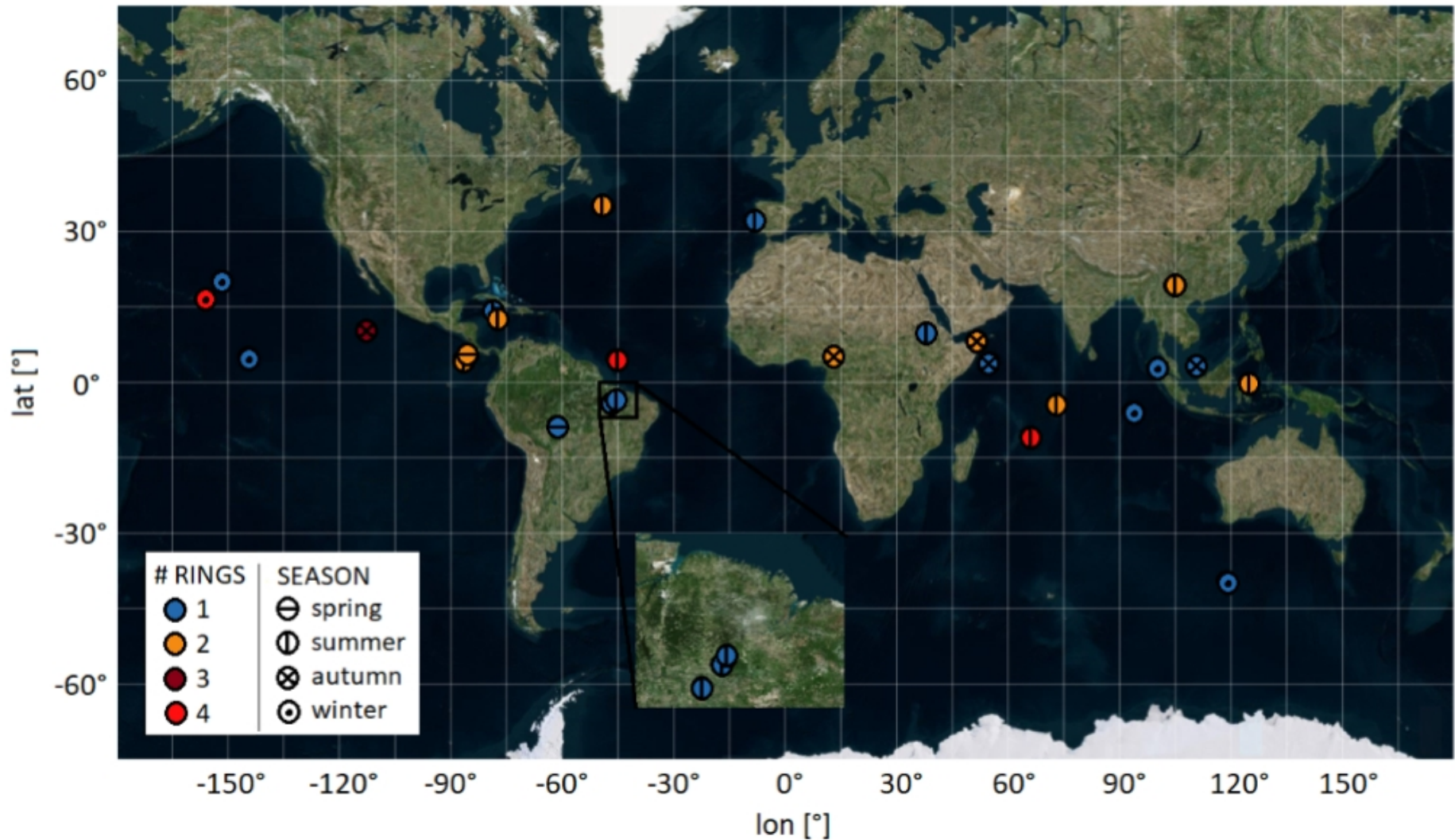
Time sampling:  $2.5\ \mu\text{s}$   
Pixel size:  $5 \times 5\ \text{km}^2$  (@ 90 km)  
Elve lifetime: about  $400\ \mu\text{s}$





# ELVES

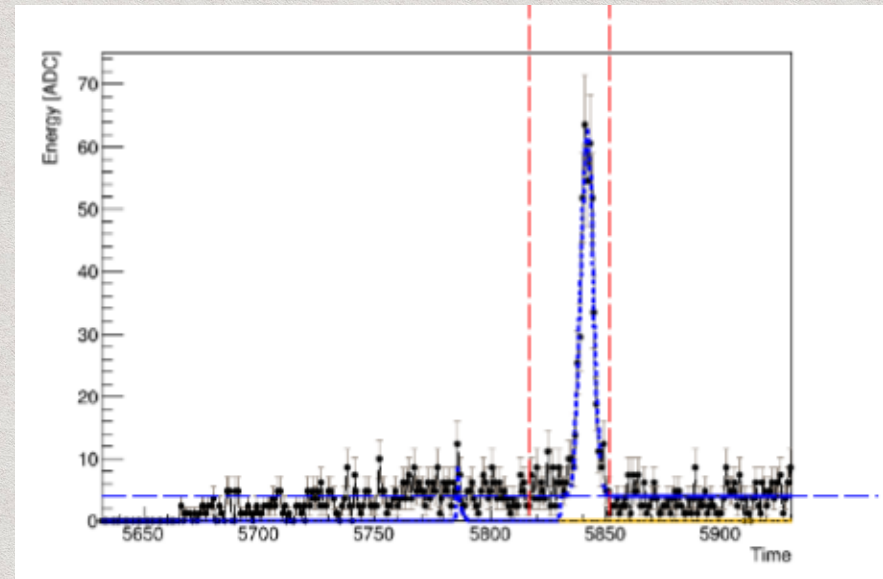
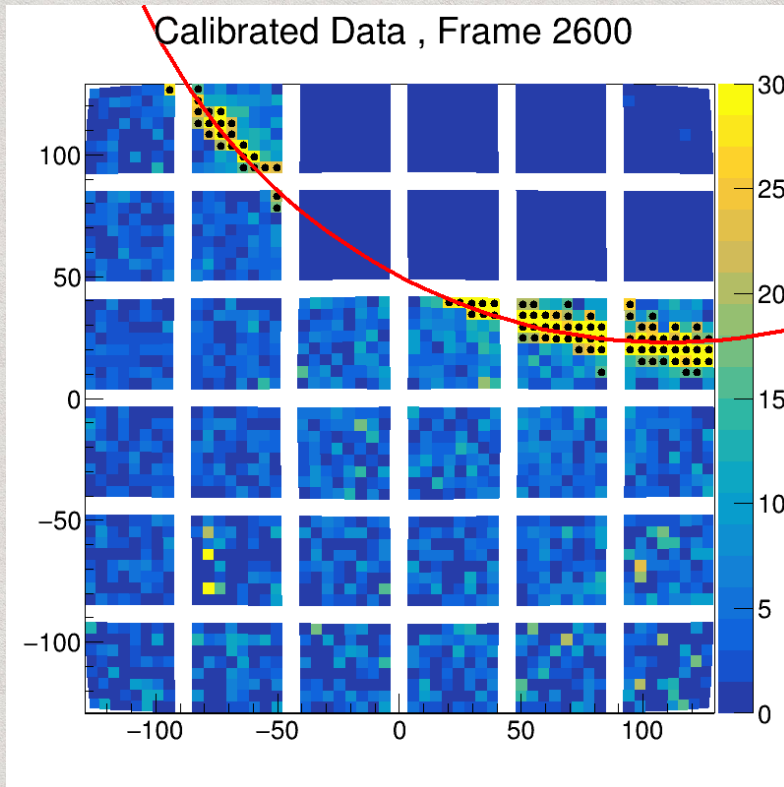
**30 ELVES** detected, mostly in the equatorial region



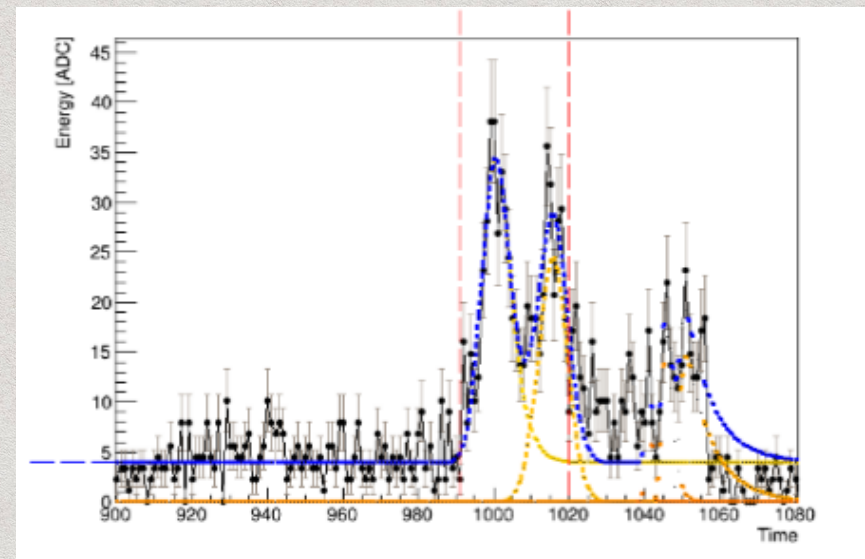


# ELVES

From G. Romoli



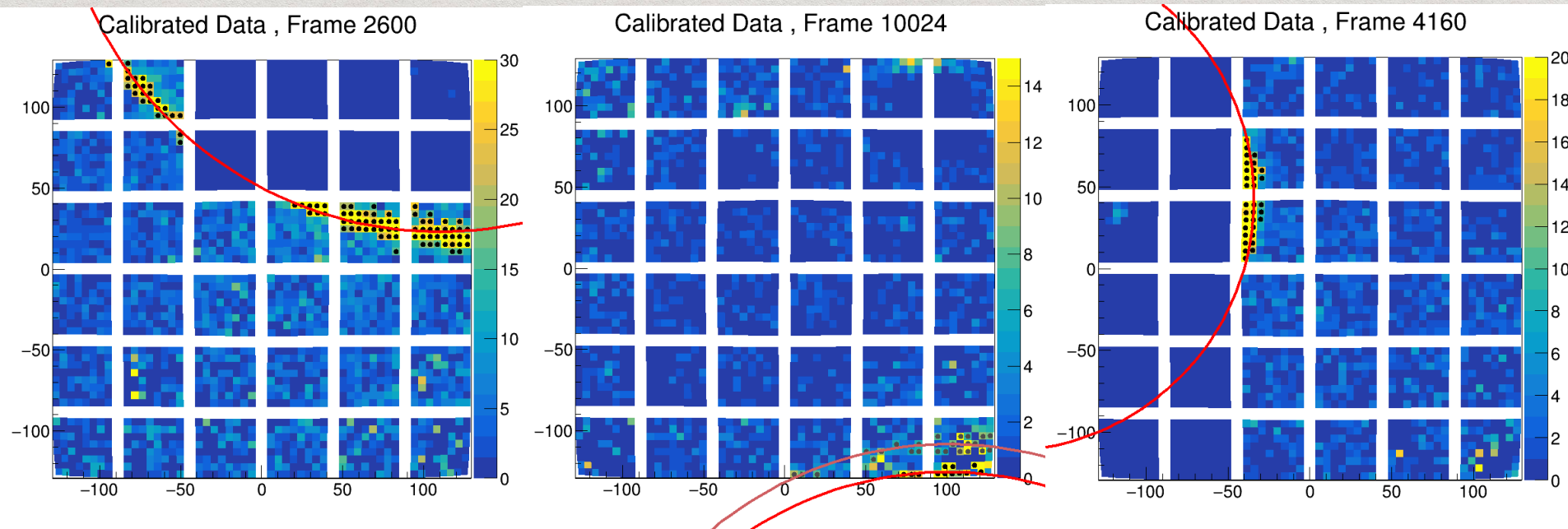
Pixel signal (pixel lightcurve) -> ELVES classification





# ELVES

From G. Romoli

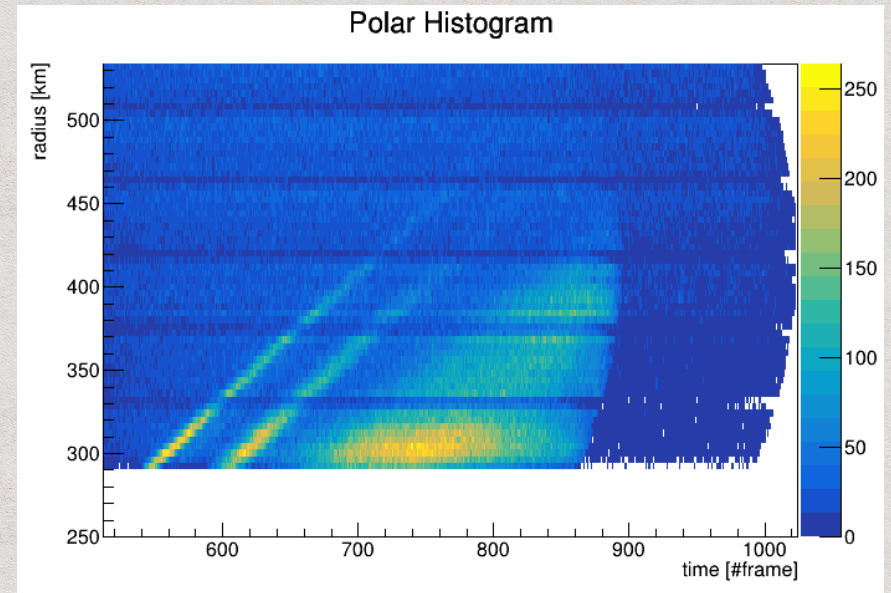
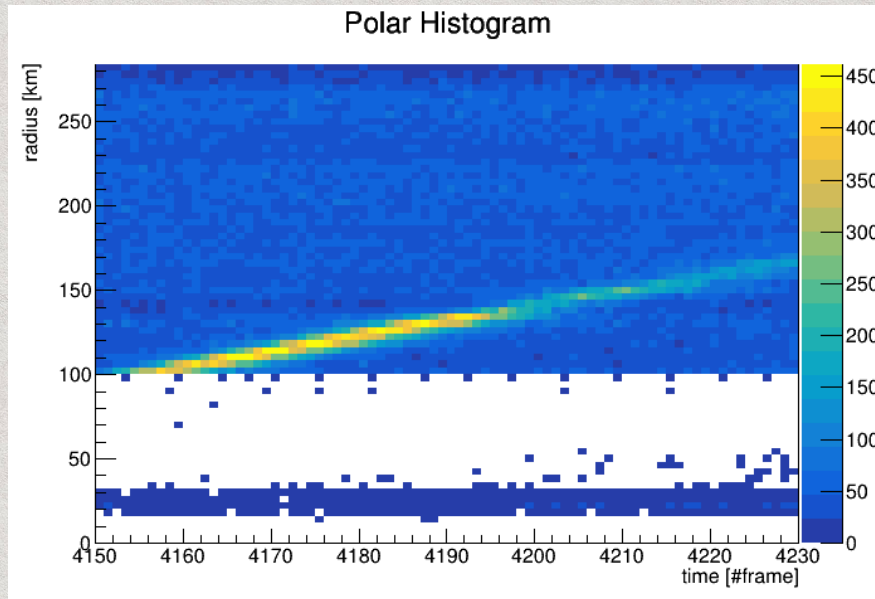


Tracking fitting algorithm -> ELVES centre



# ELVES

From G. Romoli / Z. Plebaniak

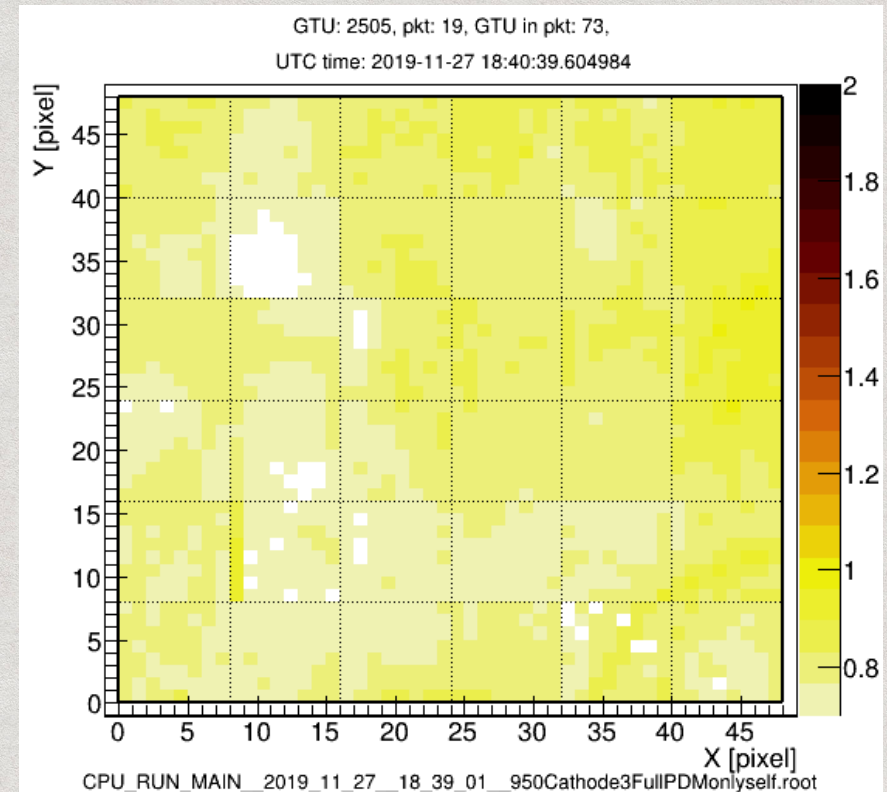
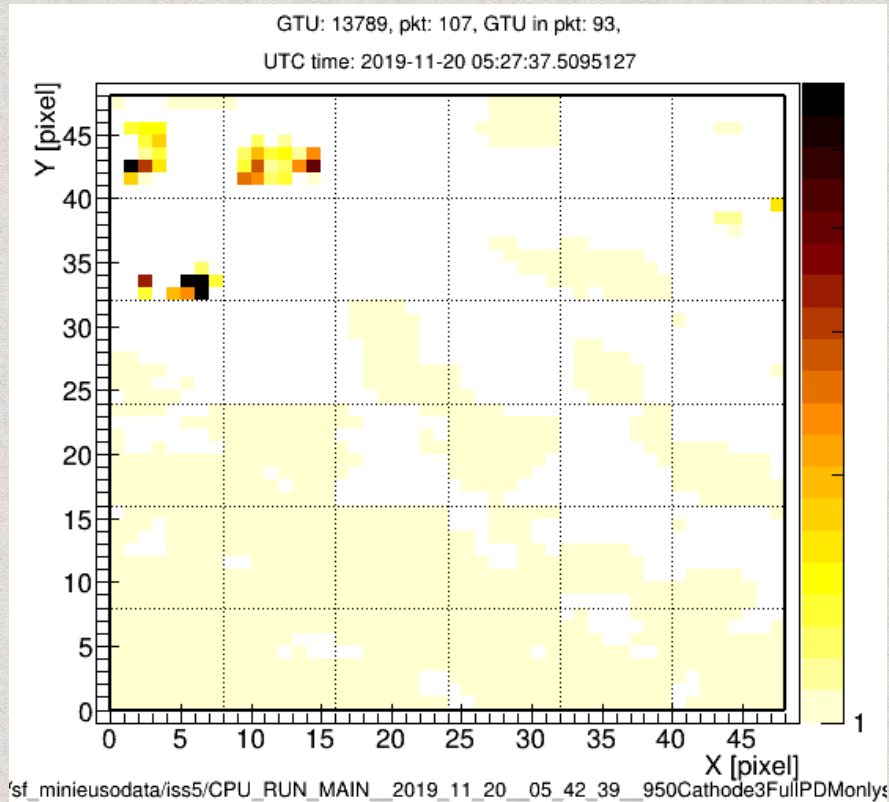


Polar histogram -> ELVES speed and classification



# Meteors

Time sampling: 40.96 ms

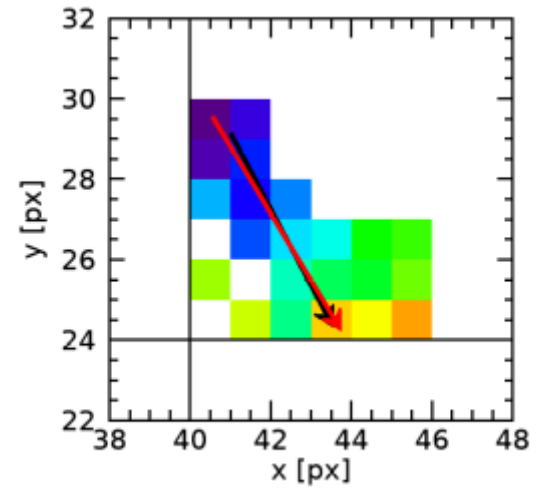
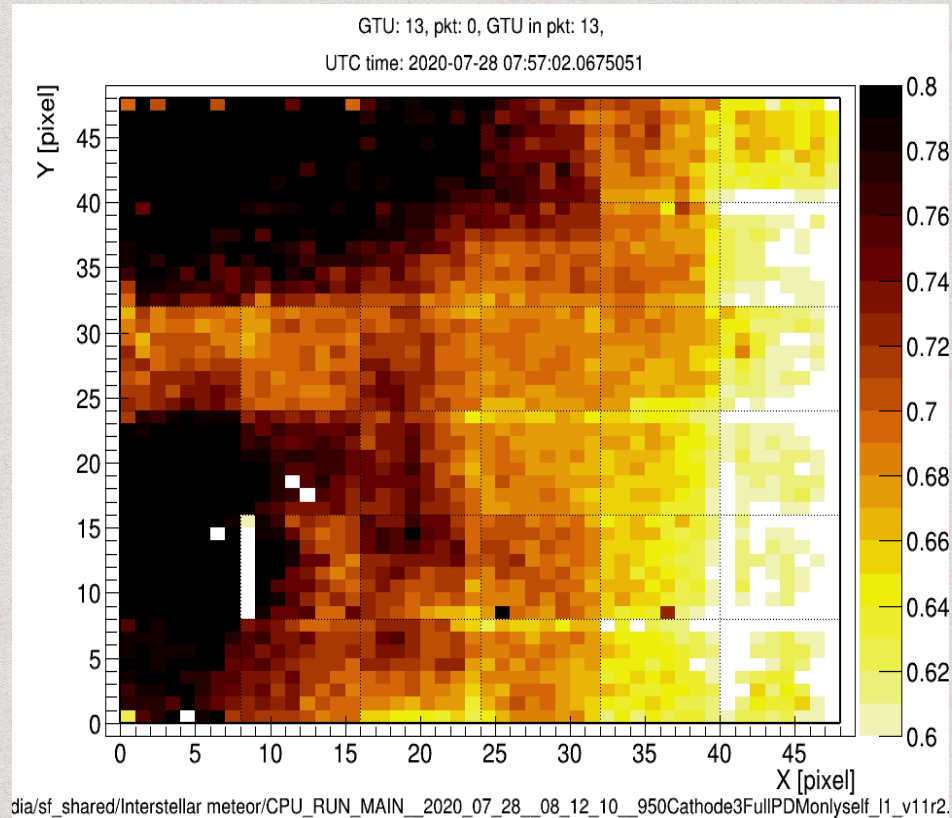


Mini-EUSO meteor dataset: 24k



# Meteor tracking algorithm

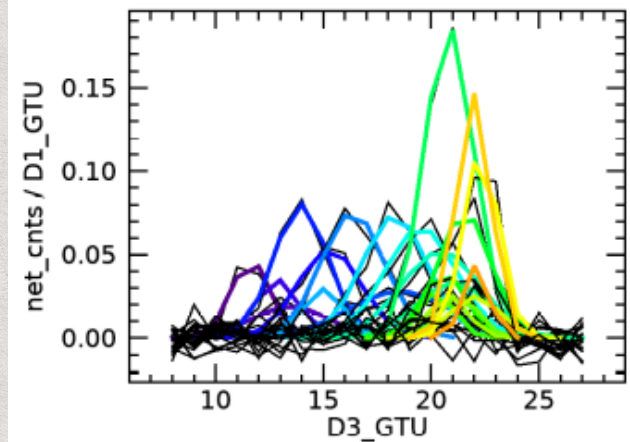
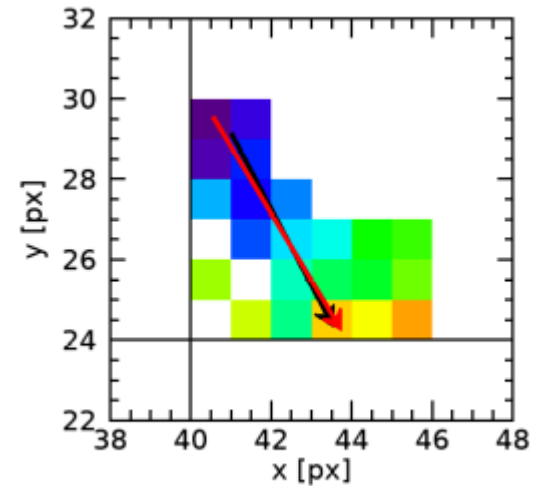
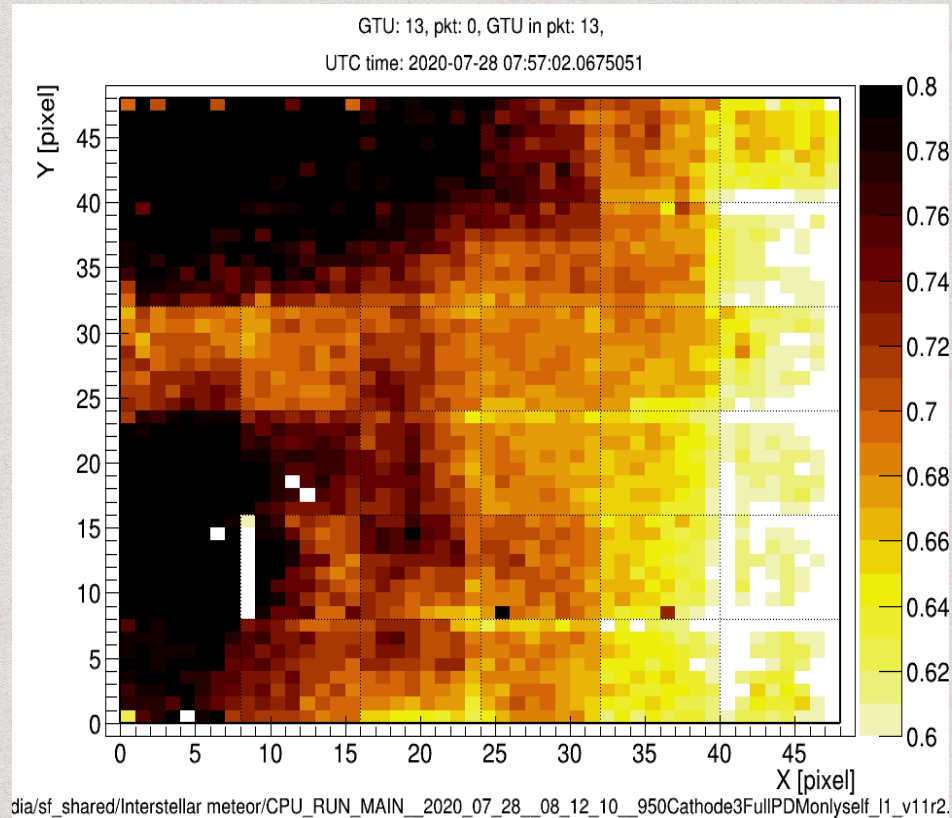
From D. Barghini





# Meteor tracking algorithm

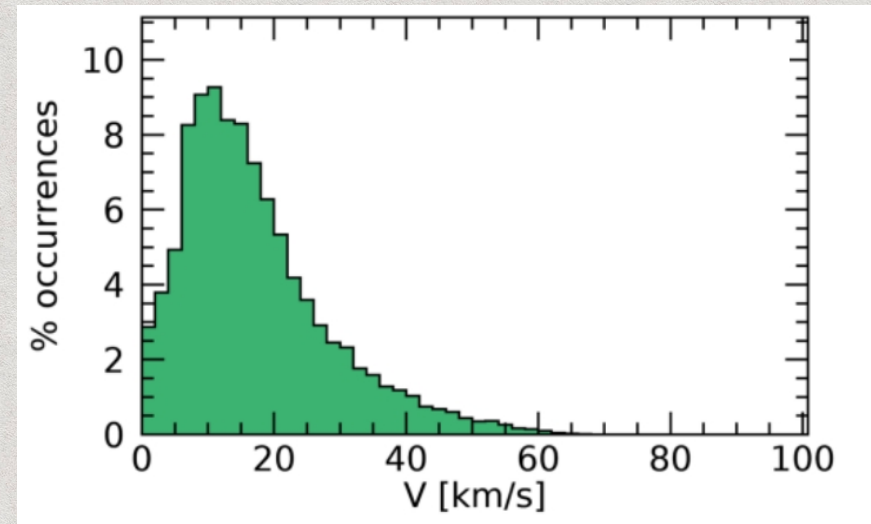
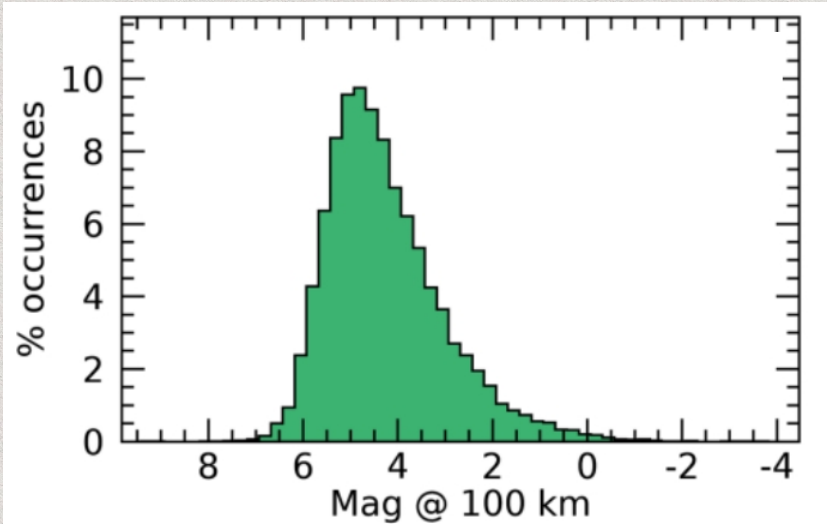
From D. Barghini





# Meteors

From D. Barghini

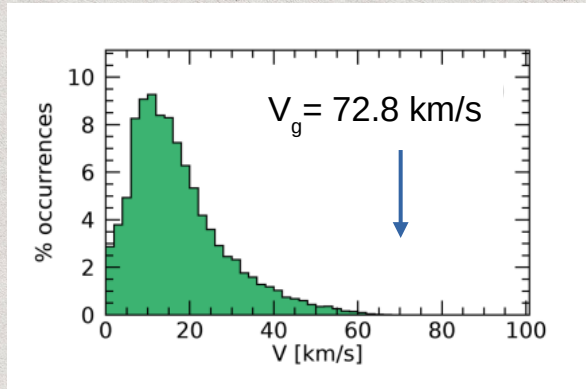


Mag 7 → ~ 2 mg  
Mag 5 → ~ 10 mg  
Mag 0 → ~ 0.1 g

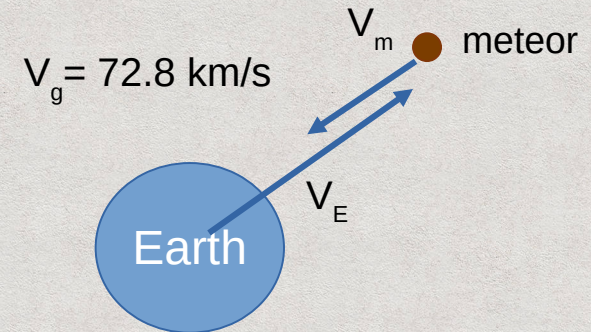
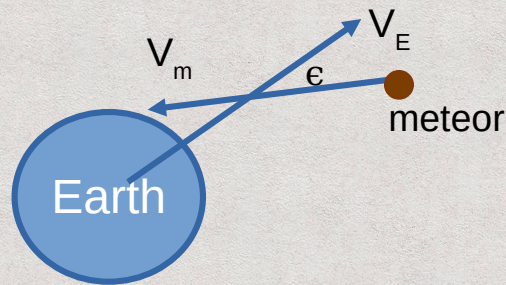
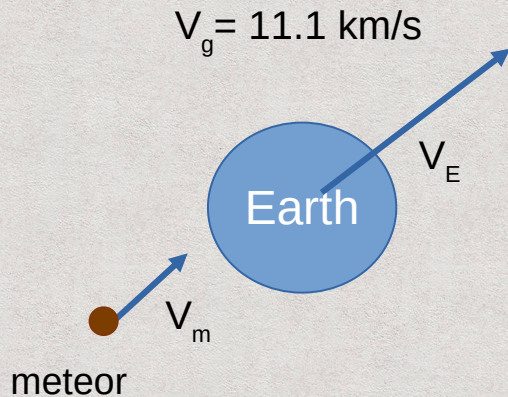
Mini-EUSO meteor dataset: 24k



# Interstellar meteors



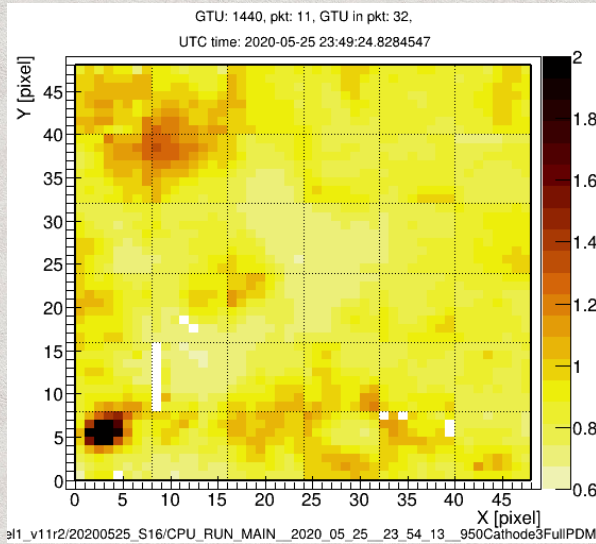
Solar system meteors:  $11.1 \text{ km/s} < v_g < 72.8 \text{ km/s}$



**Interstellar meteors:** meteor speed above solar system allowed speed



# Interstellar meteors: three candidates

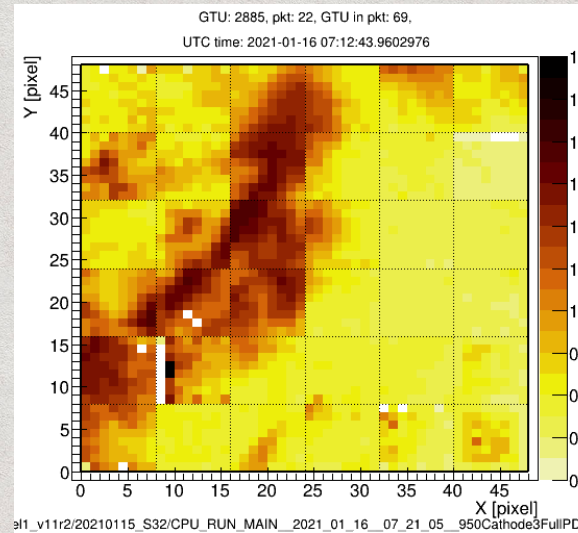


$$V_{\text{hor}} = 33 \pm 2 \text{ km/s}$$

$$M = 2.7 \pm 0.1$$

$$N_{\text{pix}} = 9$$

$$\Delta T = 0.70 \text{ s}$$

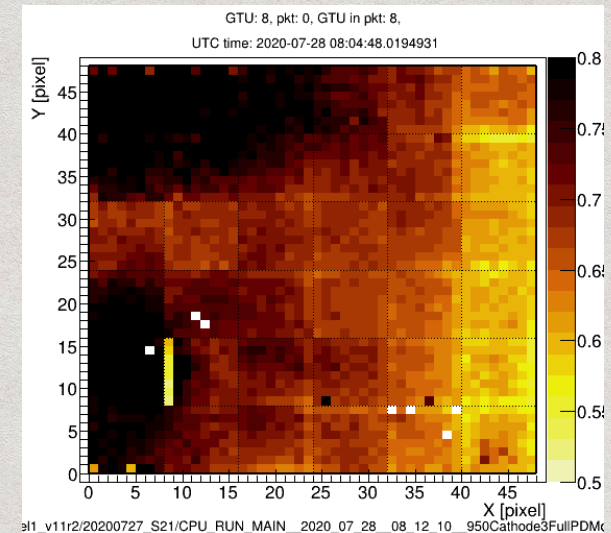


$$V_{\text{hor}} = 42 \pm 4 \text{ km/s}$$

$$M = 3.3 \pm 0.1$$

$$N_{\text{pix}} = 22$$

$$\Delta T = 0.82 \text{ s}$$



$$V_{\text{hor}} = 37 \pm 3 \text{ km/s}$$

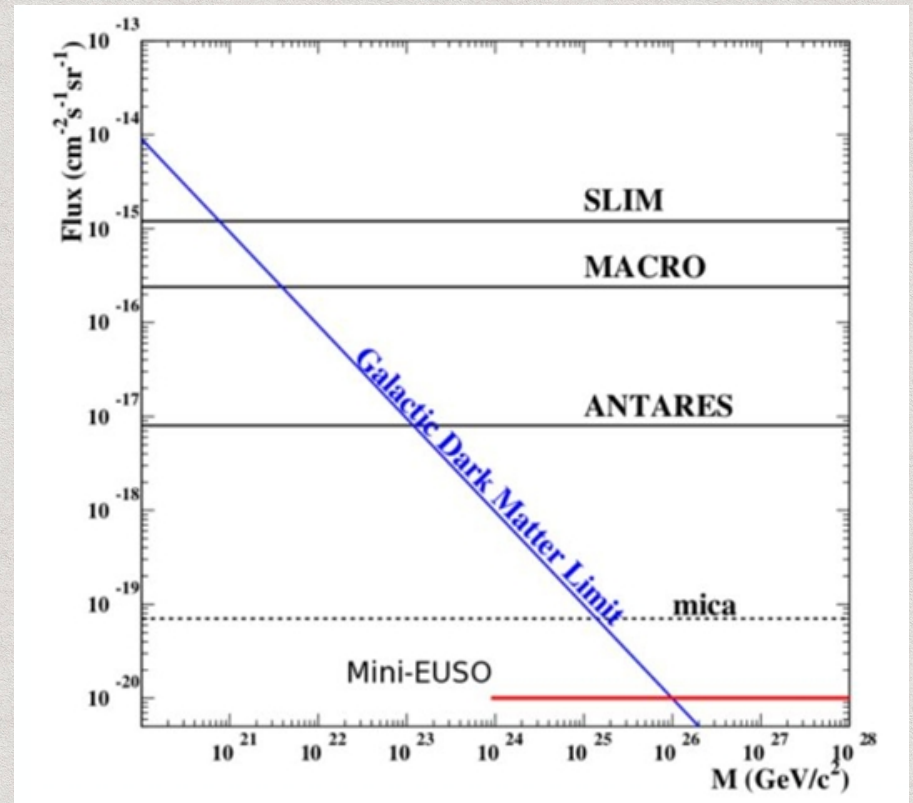
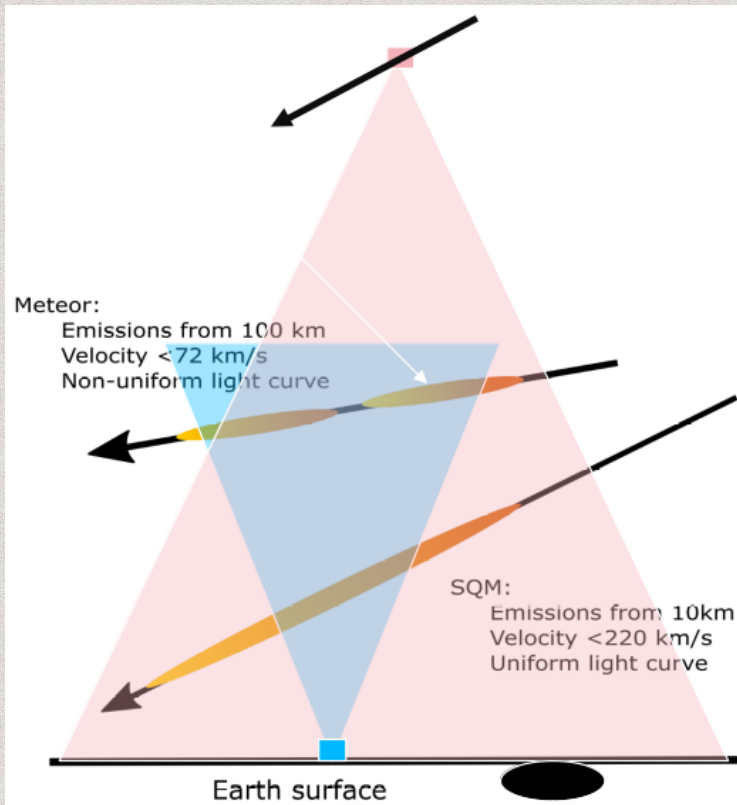
$$M = 2.9 \pm 0.1$$

$$N_{\text{pix}} = 10$$

$$\Delta T = 0.78 \text{ s}$$

Selection: robust track reconstruction (n. pixel, magnitude, ...)  
correct estimation of uncertainty on velocity measurement  
(trajectory inclination missing)

# Search for Strange Quark Matter



*Meteor studies in the framework of the JEM-EUSO program. PLANETARY AND SPACE SCIENCE, 143(SI):245-255, 2017.*

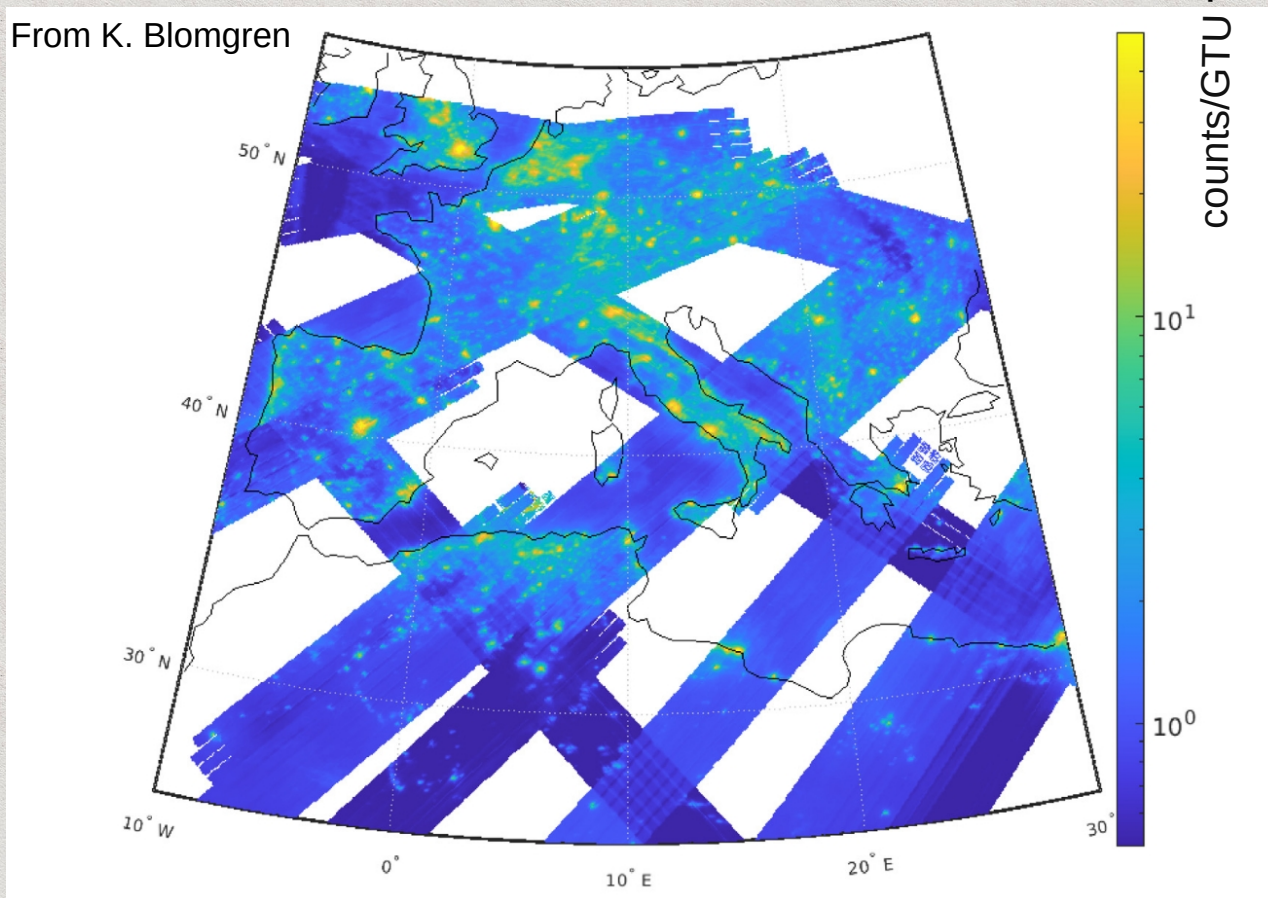
*JEM-EUSO: Meteor and nuclearite observations. Experimental Astronomy, 40:253- 279, 2015.*



# Night-time Earth Emissions

Time sampling: 40.96 ms

From K. Blomgren



Observation of night-time emissions of the Earth in the near UV range from the International Space Station with the Mini-EUSO detector  
M. Casolino et al., *Remote Sensing of Environment* 284 (2023) 113336

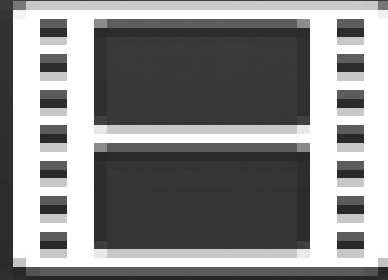
Dataset of night-time emissions of the Earth in the near UV range (290-430 nm), with 6.3 km resolution in the latitude range - 51.6<L<+51.6 degrees, acquired on board the International Space Station with the Mini-EUSO detector  
L. Marcelli et al., *Data in Brief*, Volume 48, June 2023, 109105

Mendeley database: <https://data.mendeley.com/datasets/57fmo7rh4n/4>

Youtube video: [https://youtu.be/X\\_OATf38Cg](https://youtu.be/X_OATf38Cg)

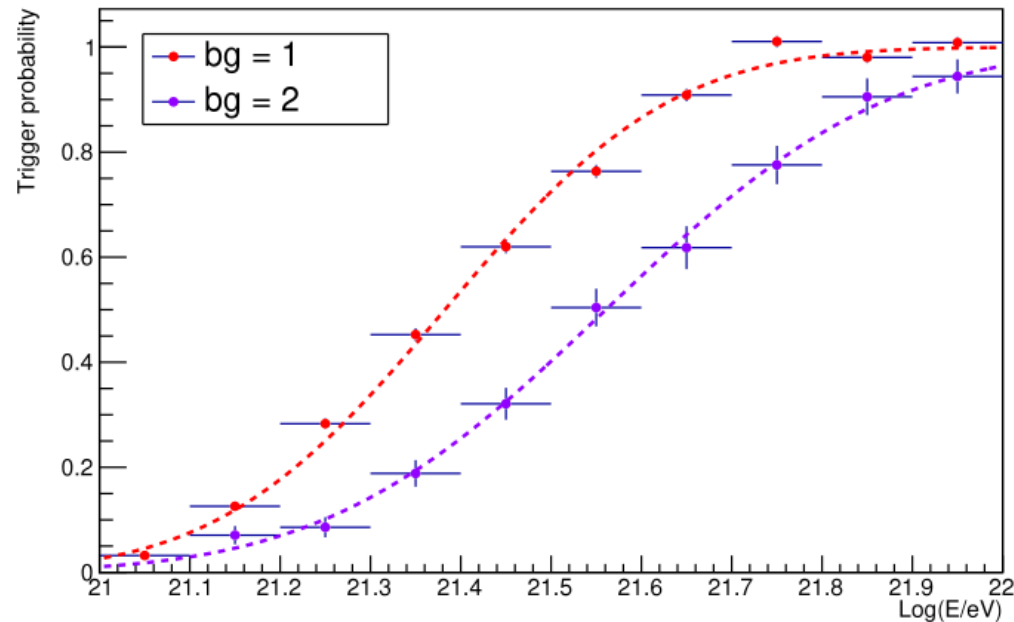


# Night-time Earth Emissions





# UHECR detection efficiency



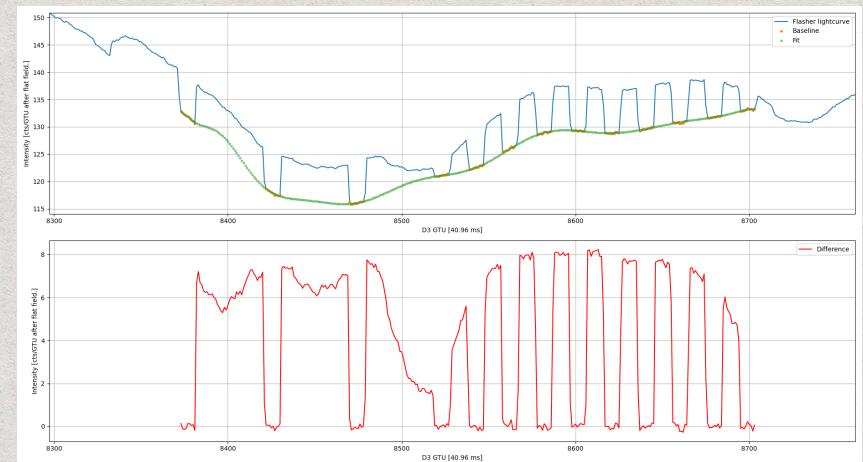
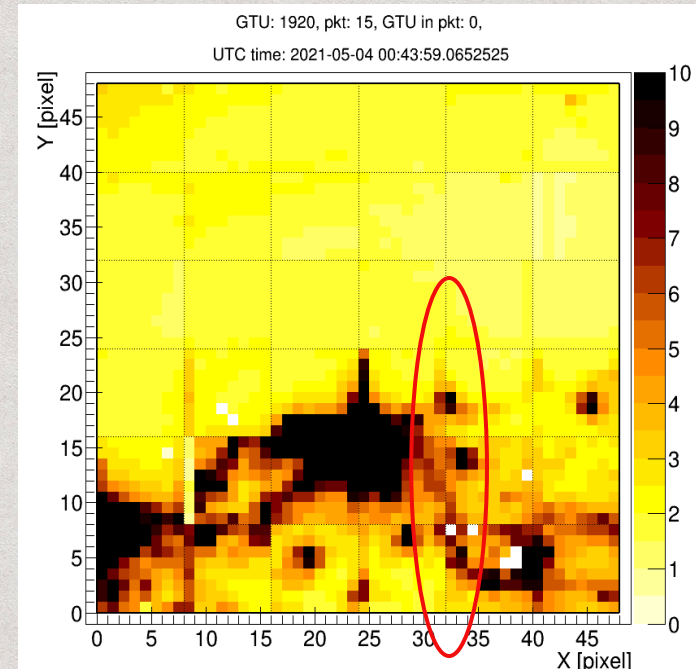
**Fig. 27.** UHECR efficiency curve as a function of energy for the Mini-EUSO detector in two different background conditions (bg = 1 and bg = 2 counts/GTU) (Bianciotto, 2020). Trigger threshold (the energy where trigger efficiency is 50%) is  $10^{21.38}$  eV and  $10^{21.56}$  eV for bg = 1 count/GTU and bg = 2 counts/GTU, respectively.

Detection thresholds (trigger efficiency = 50%):  $10^{21.38}$  eV for bg = 1 count/pixel/GTU  
 $10^{21.56}$  eV for bg = 2 counts/pixel/GTU

# End-to-end Calibration with UV flashers

2-3 kW pulsed UV LED portable arrays

- Calibration from ground
- Shoot when in field of view
- Pulsed and coded shots
- Many flasher campaigns performed
- Partial results published, waiting for new data



*An end-to-end in-flight calibration of Mini-EUSO detector  
H. Miyamoto, EPJ Web of Conference 283, 06017 (2023)*



# Selected publications

**Dataset of night-time emissions of the earth in the near uv range (290-430 nm), with 6.3 km resolution in the latitude range  $-51.6 < L < +51.6$  degrees, acquired on board the international space station with the mini-euso detector.**

**Data in Brief, 48, 2023.**

**Observation of night-time emissions of the earth in the near uv range from the international space station with the mini-euso detector.**  
**Remote Sensing Of Environment, 284, 2023.**

**Onboard performance of the level 1 trigger of the mini-euso telescope.**  
**Advances in Space Research, 70(9):2750–2766, 2022.**

**Pre-flight qualification tests of the mini-euso telescope engineering model.**  
**Experimental Astronomy, 53(1):133–158, 2022.**

**Mini-EUSO Mission to Study Earth UV Emissions on board the ISS**  
**The Astrophysical Journal Supplement Series, 253, 2 , 36, 2021.**

**Mini-EUSO: A high resolution detector for the study of terrestrial and cosmic UV emission from the International Space Station.**  
**Advances in Space Research, 62(10):2954-2965, 2018.**

**Secondary cameras onboard the Mini-EUSO experiment: Control software and calibration.**  
**Advances in Space Research, 64(5):1188-1198, 2019.**

**Mini-EUSO data acquisition and control software.**  
**Journal of Astronomical Telescopes Instruments and Systems, 5(4), 2019.**

**The integration and testing of the Mini-EUSO multi-level trigger system,**  
**Advances in Space Research, 62(10), 2966-2976 , 2018**



# Conclusions

After 85 sessions Mini-EUSO works nominally

Mini-EUSO is a multidisciplinary experiment (ELVES, UV Earth maps, SQM...)

UHECR detection efficiency estimation validated for future missions

**SPB2 + Mini-EUSO results pave the way for UHECR detection from space**

