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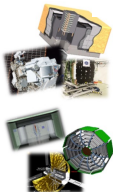
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## Silicon Microstrip detectors in space

Most of space detectors for charged cosmic ray and  $\gamma$ -ray measurements require **solid state tracking systems based on Si-microstrip (SIMS) sensors**.

SIMS detectors are the only solution to instrument **large area detectors** with larger number of electronics channels coping with the **limitations on power consumption in space**



Operating Missions						
	Mission Start	Si-sensor area	Strip length	Readout channels	Readout pitch	Spatial resolution
Fermi-LAT	2008	$\sim 74 \text{ m}^2$	38 cm	$\sim 880 \cdot 10^3$	228 $\mu\text{m}$	$\sim 66 \mu\text{m}$
AMS-02	2011	$\sim 7 \text{ m}^2$	29-42 cm	$\sim 200 \cdot 10^3$	110 $\mu\text{m}$	$\sim 7 \mu\text{m}$
DAMPE	2015	$\sim 7 \text{ m}^2$	38 cm	$\sim 70 \cdot 10^3$	242 $\mu\text{m}$	$\sim 40 \mu\text{m}$

Future Missions						
	Planned operations	Si-sensor area	Strip length	Readout channels	Readout pitch	Spatial resolution
HERD	2030	$\sim 35 \text{ m}^2$	48-67 cm	$\sim 350 \cdot 10^3$	$\sim 242 \mu\text{m}$	$\sim 40 \mu\text{m}$
ALADiO	2050	$\sim 80 \cdot 100 \text{ m}^2$	$\sim 2.5 \cdot 10^6$	$\sim 100 \mu\text{m}$	$\sim 5 \mu\text{m}$	$\sim 5 \mu\text{m}$
AMS-100	2050	$\sim 180 \cdot 200 \text{ m}^2$	$\sim 100 \text{ cm}$	$\sim 8 \cdot 10^6$	$\sim 100 \mu\text{m}$	$\sim 5 \mu\text{m}$

## "5D tracking"

high accuracy position, charge and timing measurement for each hit in a high rate environment

Possible experimental opportunities enabled

### IMPROVED TRACK FINDING

Hit timing information improves track reconstruction on high rate environments and identifies backscattering hits from downstream calorimeters

### REMOVE "GHOST" HITS

Separating tracks in time can mitigate the ambiguity of "ghost" hits in SIMS with strips running in perpendicular directions

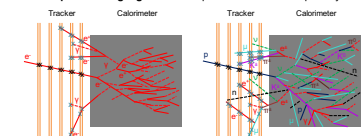
### TIME OF FLIGHT

Hit timing resolutions of  $\sim 100$  ps enable ToF measurements with SIMS complementary to scintillators with fast light readout

### PARTICLE ID

Track timing identifies slow low-energy particles backscattering from downstream calorimeters for primary hadronic particle crossings

Different backplash timing signatures in upstream tracker from primary  $e^\pm$  or hadrons



A detailed discussion on possible advantages from 5D tracking in space astroparticle experiments is available at M. Duranti, V. Vagelli et al., Instruments **2021**, 5(2), 20



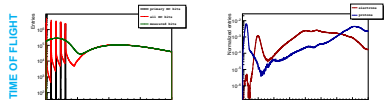
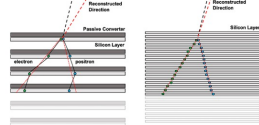
## Sub-GeV $\gamma$ -ray detectors

Opportunity for improved PSF below 1 GeV

**Thin high signal Si sensors:** the LGAD intrinsic gain improves the SNR for thin sensors and allows for reduced active material budget tracking planes

The PSF of  $\gamma$ -ray experiments (Fermi-LAT, DAMPE, ...) is degraded at low energies by Coulomb MS in materials (passive converter and Si-sensors)

- Remove passive materials
- Use thin active detectors
- Increase number of active layers to boost the GR conversion probability (approach first proposed in [5])



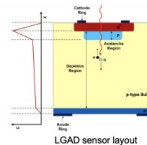
100 ps resolution allows to separate hits from primary particles and from secondary backplash

Different timing signatures from p and  $e^\pm$  showers

PARTICLE ID

Prospects to use in MVA classifiers for  $e/p$  separation

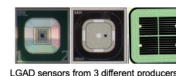
## Tracking with Low-Gain Avalanche Diodes



**LGAD: standard Si sensors with intrinsic gain layer typical of APD devices**

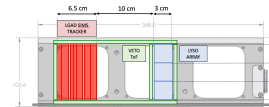
- High E-field region allows carrier multiplication and provides gain
- Moderate gain (10-70) increases the signal-to-noise ratio (SNR)
- Provides accurate timing resolution down to tens of ps

Developed for pixel sensors for HL-LHC detectors, **LGAD SIMS detectors and readout FEE require dedicated R&D for space qualification and power consumption**



LGAD is the most suitable, mature enough, candidate technology to **enable 5D tracking** and have **very thin but efficient SIMS detectors in Space**.

## A CubeSat demonstrator for the Space qualification of LGAD microstrip detectors in space



A conceptual design of the demonstrator compatible with the constraints in weight, volume and power budget of a CubeSat platform hosted in 2 units of a 3U CubeSat, with one additional units dedicated to the FEE and DAQ of the demonstrator.

### LGAD SIMS Tracker

40 layers of 150  $\mu\text{m}$  thick SIMS LGAD sensors  
readout pitch: 150  $\mu\text{m}$   $\rightarrow$  expected  $\Delta x \sim 15 \mu\text{m}$   
Target timing resolution  $\sim 100$  ps

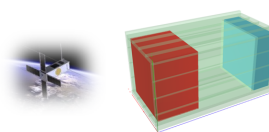
### Veto / Time of Flight system

0.5 cm thick Sci-paddles  
SiPM readout using commercial FEE.  $\Delta t \sim 30$  ps

### Electromagnetic Calorimeter

3x3x3  $\text{cm}^3$  array of LYSO crystals  
SiPM readout using commercial FEE  
Feasibility to add another stack of LYSO array under study

**Weight < 3 kg Power < 20 W**



Simulation of the detector performances is ongoing  
FEE power mitigation techniques under investigation

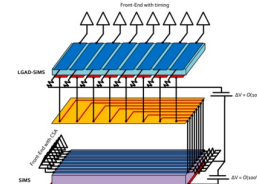
**R&D activity to increase LGAD SIMS TRL for space from TRL=2 to TRL=5 funded by Italian MUR (PRIN 2022)**

### PTSD - Penta Tracking Space Detector

Institutes: INFN, ASI and FBK:

1. Proto-instrument requirement identification
2. Design and production of LGAD SIMS sensors with large active area optimized for 5D-tracking measurements in space
3. Development of SIMS sensor assembly procedures in a large area detector
4. assembly of a breadboard laboratory model for verification of requirements, functionalities and space qualification

**Funding and activities planned to start in Fall 2023**



Possible connections for a 5D detector made of a thin LGAD SIMS sensor for timing and moderate resolution coordinate measurement coupled with a thicker standard SIMS sensor for charge and high resolution coordinate measurement. This layout drives the main investigation in the context of the project.