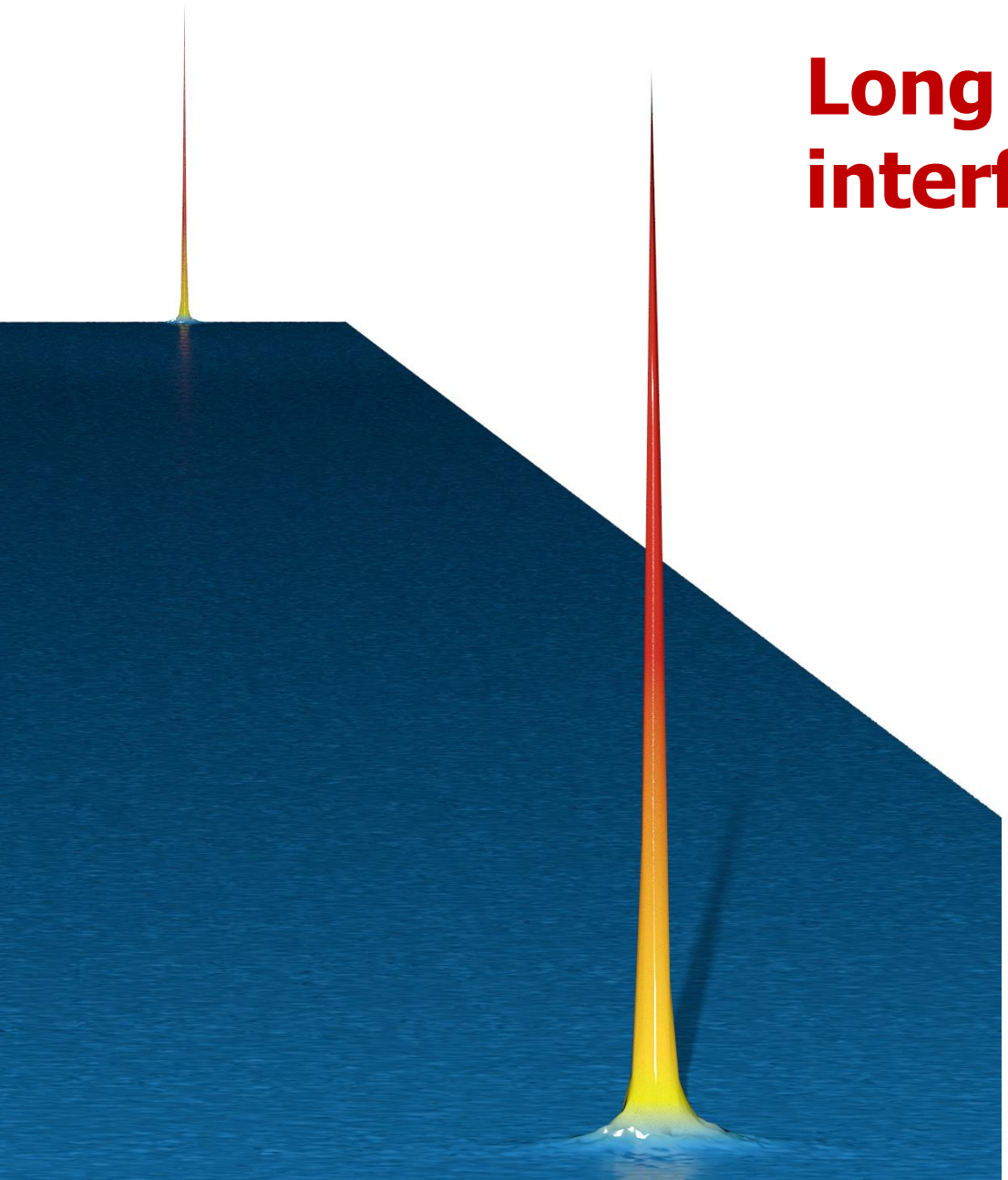
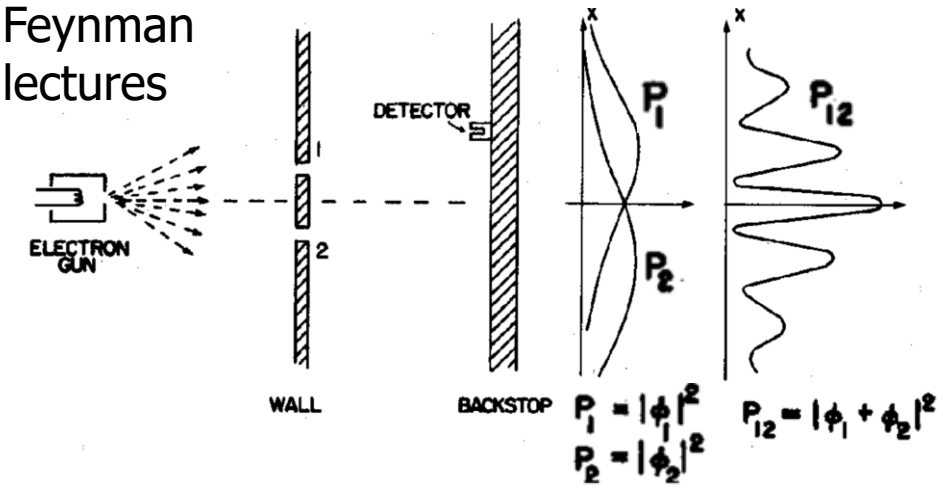


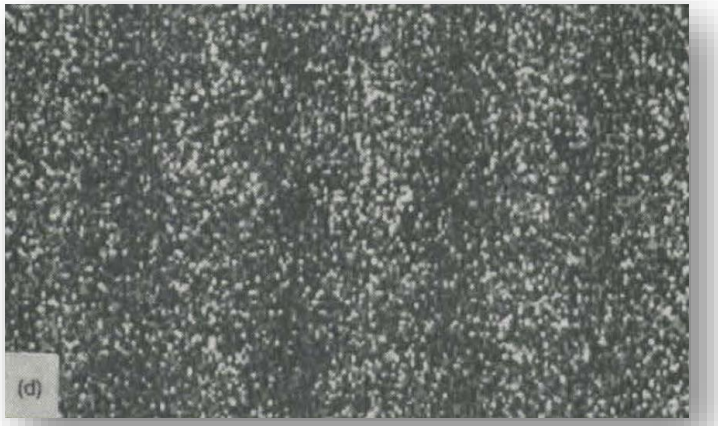
Long baseline atom interferometry

Mark Kasevich
Stanford University
kasevich@stanford.edu

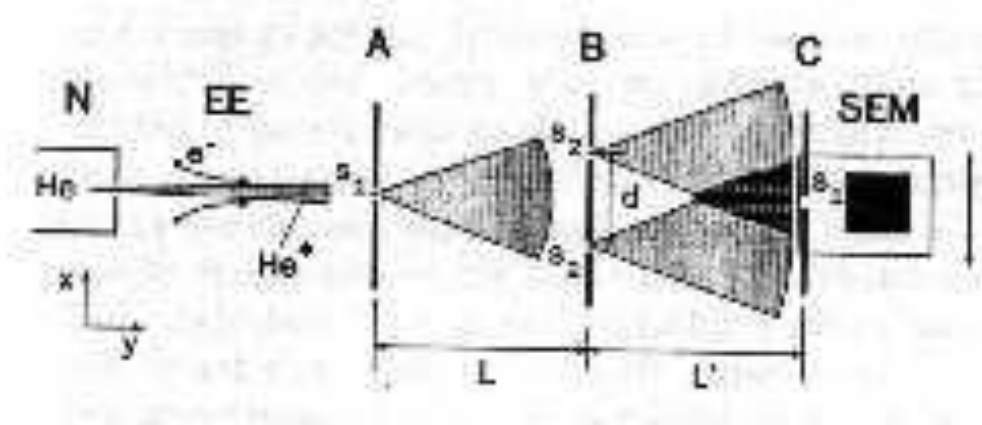
Young's double slit exp't with particles

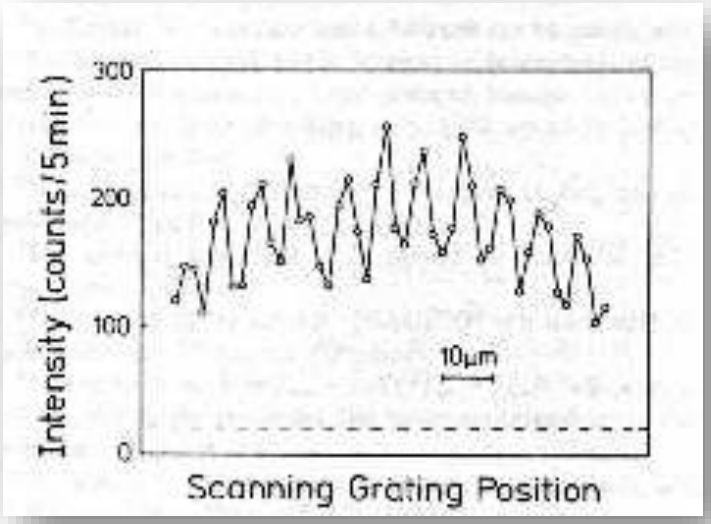
Feynman lectures





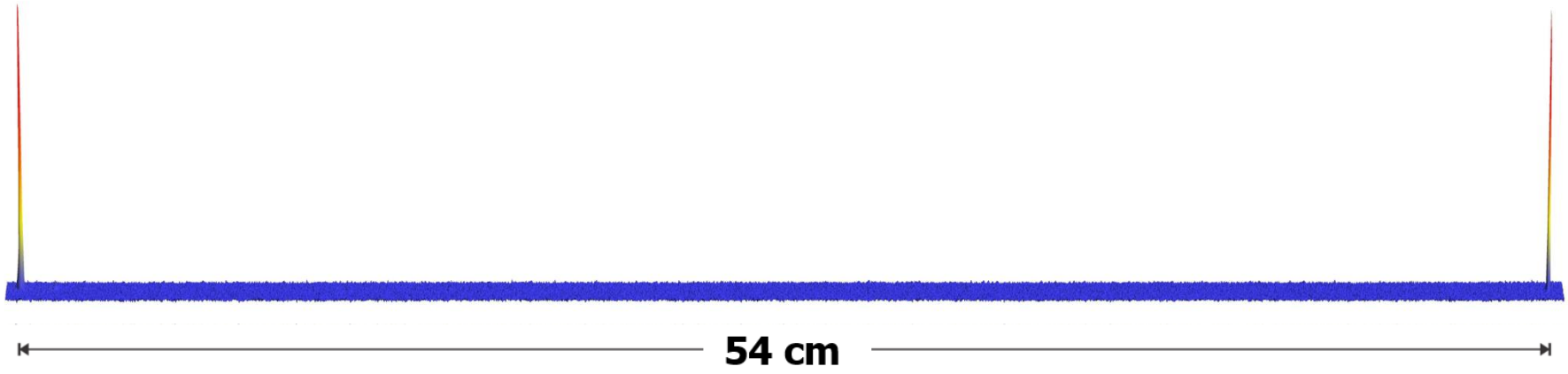
Tonomura, PRL, 1989





Mlynek, PRL, 1991

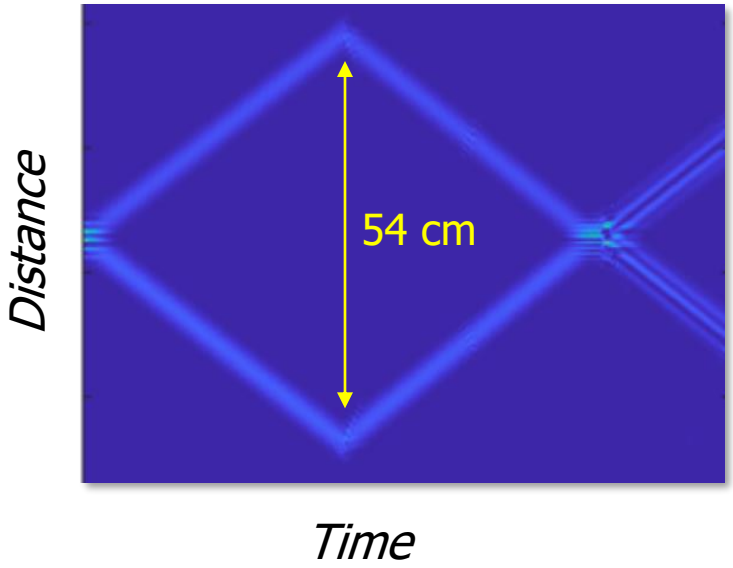
Atomic wavepacket superposition

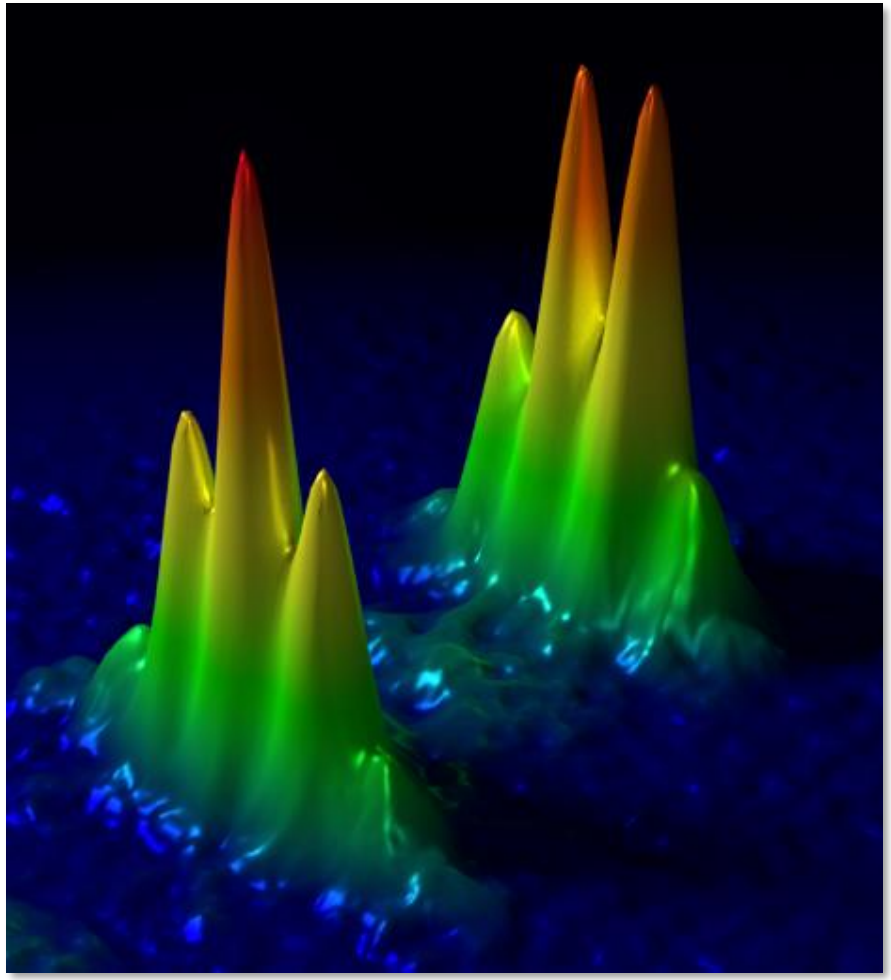


Kovachy, et al., Nature (2015).

Atomic wavepacket interference

Interferometer atomic
center-of-mass wavefunction

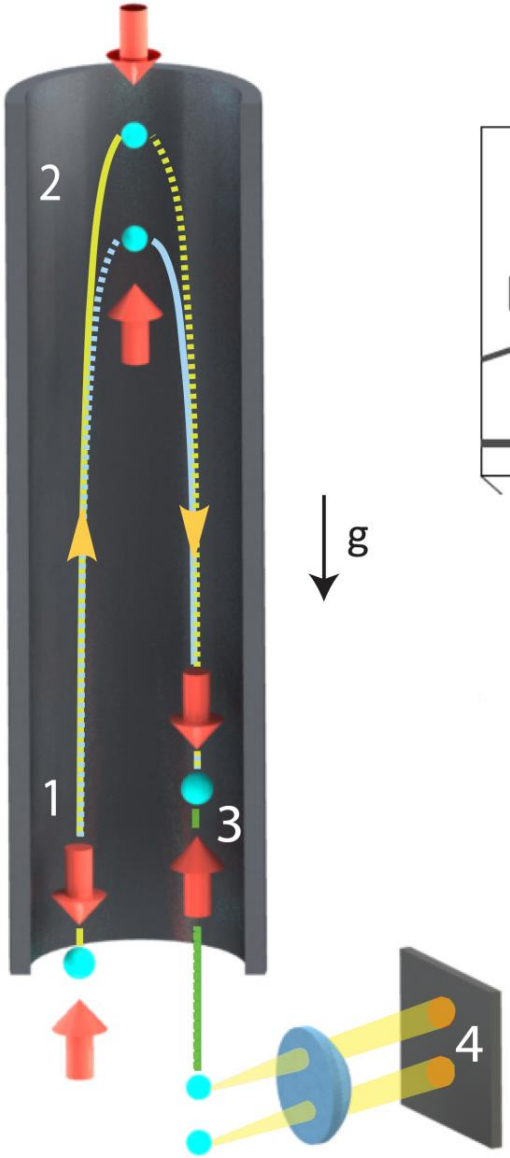


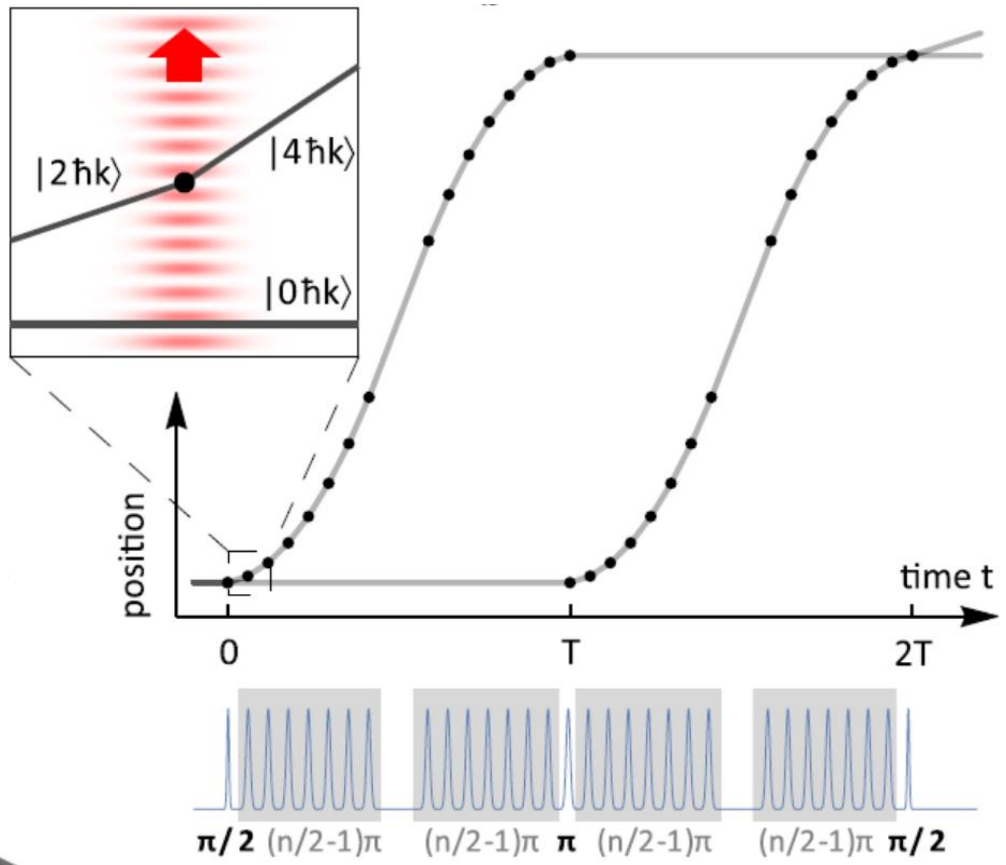


Superposed atomic wavepackets interfere

Asenbaum, PRL (2017).

Light pulse atom interferometry

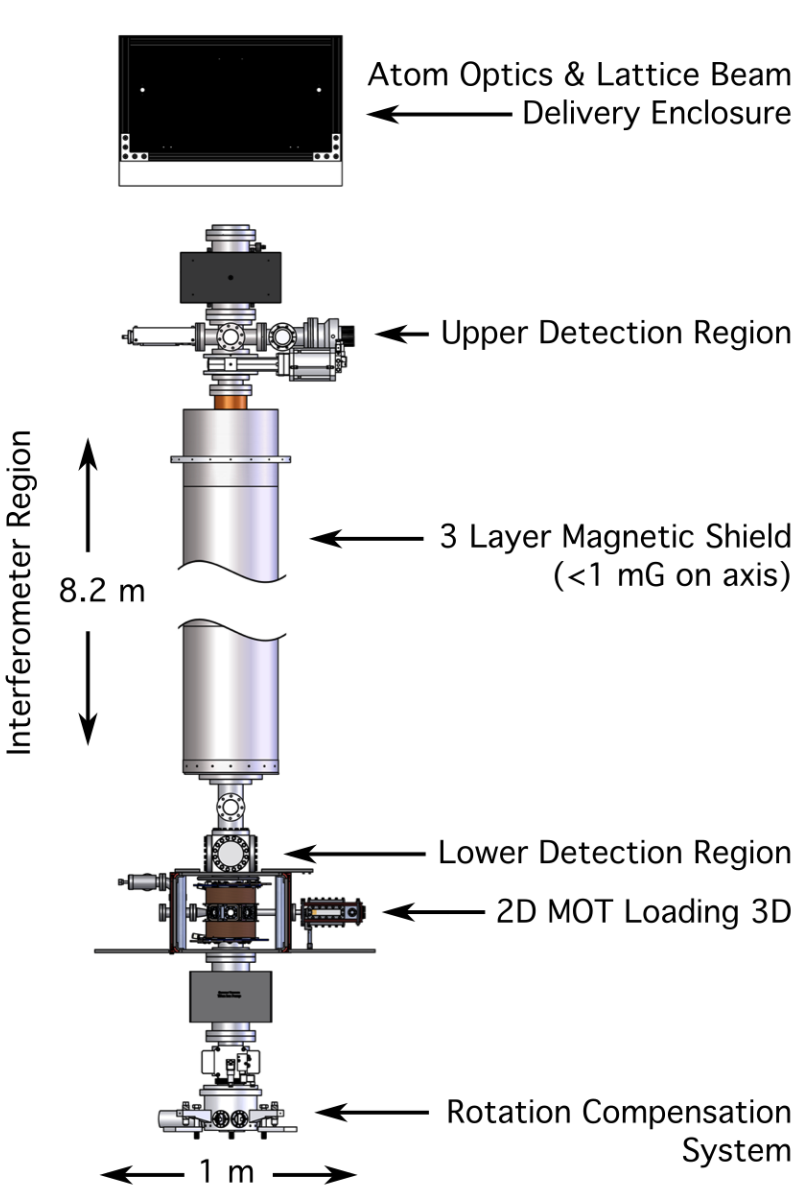




Pulse sequence duration: ~2 s

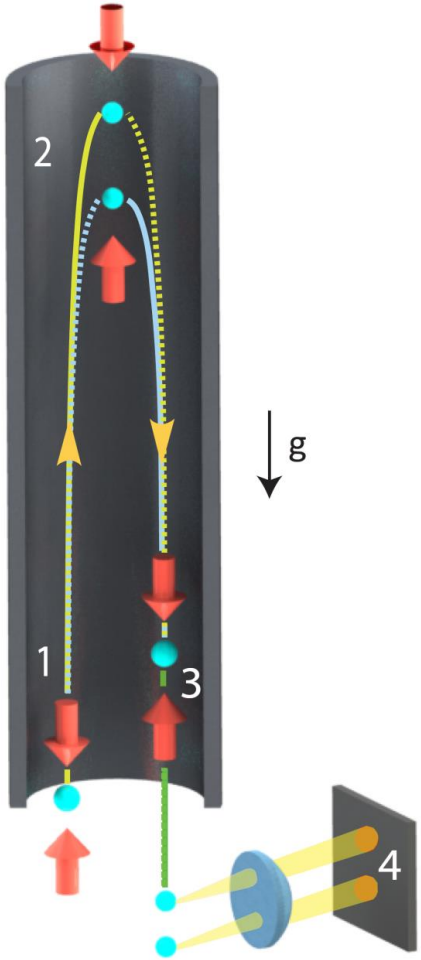
Kovachy, Nature, 2015

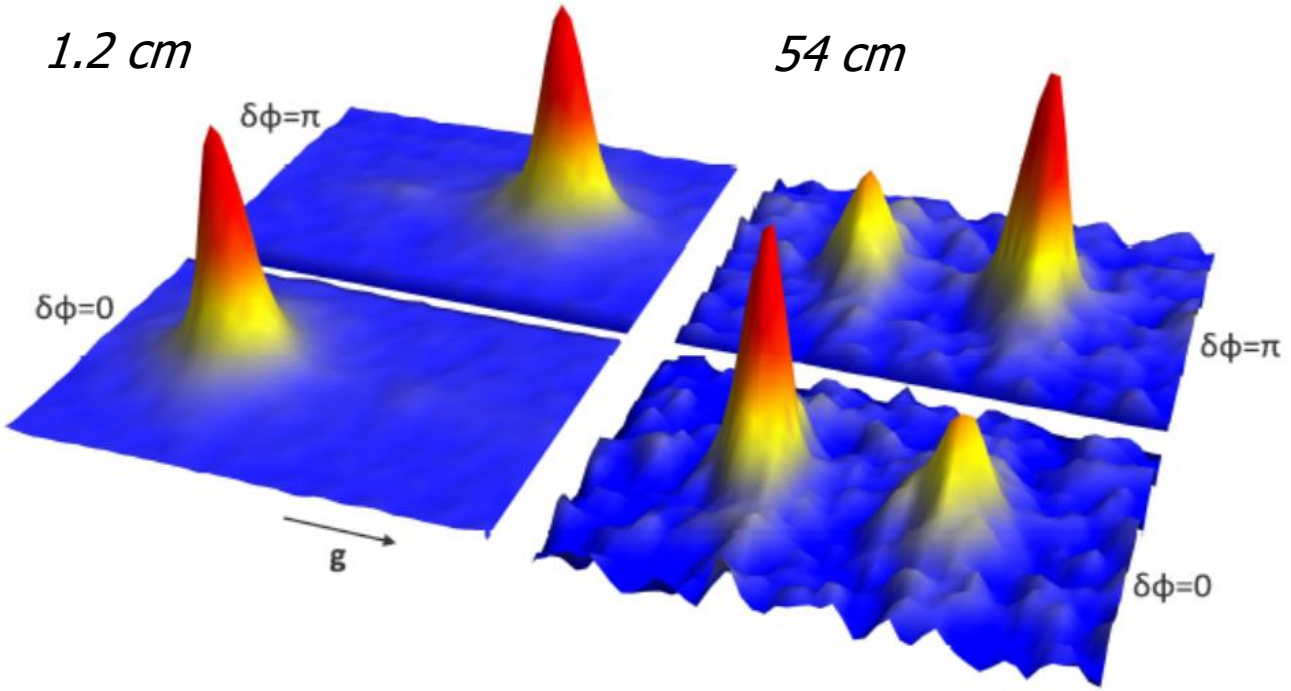
Apparatus



~ 100 pK
1e5 atoms/shot

Interference at output ports





Interference causes population modulation between the output ports

T. Kovachy, et al., Nature (2015)

Bounds for gravitational decoherence

IOP Publishing
Classical and Quantum Gravity
Class. Quantum Grav. 35 (2018) 145005 (18pp)
https://doi.org/10.1088/1361-6382/aac72f

Gravity is not a pairwise local classical channel

Natacha Altamirano[1,2], Paulina Corona-Ugalde[2,3], Robert B Mann[1,2,3] and Magdalena Zych[4,5]

$$\tilde{\Gamma}^{C}_{\mathrm{KTM}} = C\frac{GMm}{\hbar R^3}\Delta x^2$$

Experiment	m (Kg)	M (Kg)	d (m)	Δx (m)	$1/\Gamma_{\mathrm{DP}}$ (s)	$1/\Gamma^{\mathrm{min}}_{\mathrm{KTM}}$ (s)
10 m atomic fountain with ^{87}Rb [38]	1.4×10^{-25}	$M_\oplus$	$R_\oplus$	0.54	3×10^{10}	2×10^{-3}
Two atomic fountains with ^{87}Rb [33]	1.4×10^{-25}	$M_\oplus$	$R_\oplus$	1.86×10^{-3}	3×10^{10}	2×10^{1}
(Operating as gravity-gradiometer)		4×129	0.11, 0.18, 0.28, 0.31			
Large-molecule interferometry [43]	1.6×10^{-23}	$M_\oplus$	$R_\oplus$	2.7×10^{-7}	3×10^{6}	6×10^{7}
PcH_2 diffraction on alga skeleton [44]	8.2×10^{-25}	$M_\oplus$	$R_\oplus$	2×10^{-7}	1×10^{9}	2×10^{9}

Large wavepacket separation AI constrains the KTM model

Phase shifts (non-relativistic)

8e9 rad

Term	Phase Shift	
1	$k_{\text{eff}} g T^2$	Gravity
2	$2\mathbf{k}_{\text{eff}} \cdot (\boldsymbol{\Omega} \times \mathbf{v}) T^2$	Coriolis
3	$k_{\text{eff}} v_z \delta T$	Timing asymmetry
4	$\frac{\hbar k_{\text{eff}}^2}{2m} T_{zz} T^3$	Curvature, quantum (tidal)
5	$k_{\text{eff}} T_{zi} (x_i + v_i T) T^2$	Gravity gradient
6	$\frac{1}{2} k_{\text{eff}} \alpha (v_x^2 + v_y^2) T^2$	Wavefront

T_{ij}, gravity gradient

v_i, velocity; x_i, initial position

g, acceleration due to gravity

T, interrogation time

k_{eff}, effective propagation vector

Equivalence Principle

Acceleration of co-falling ^{85}Rb and ^{87}Rb ensembles is measured using light-pulse atom interferometry

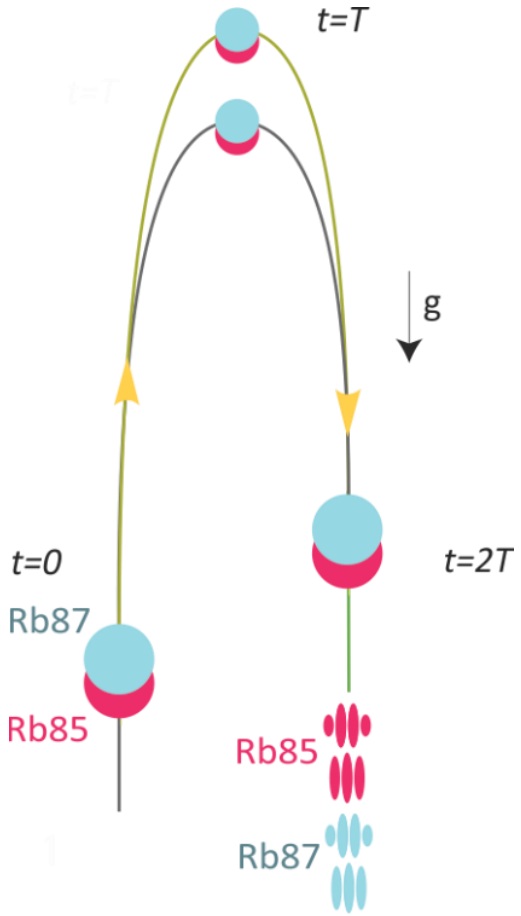
Statistical sensitivity

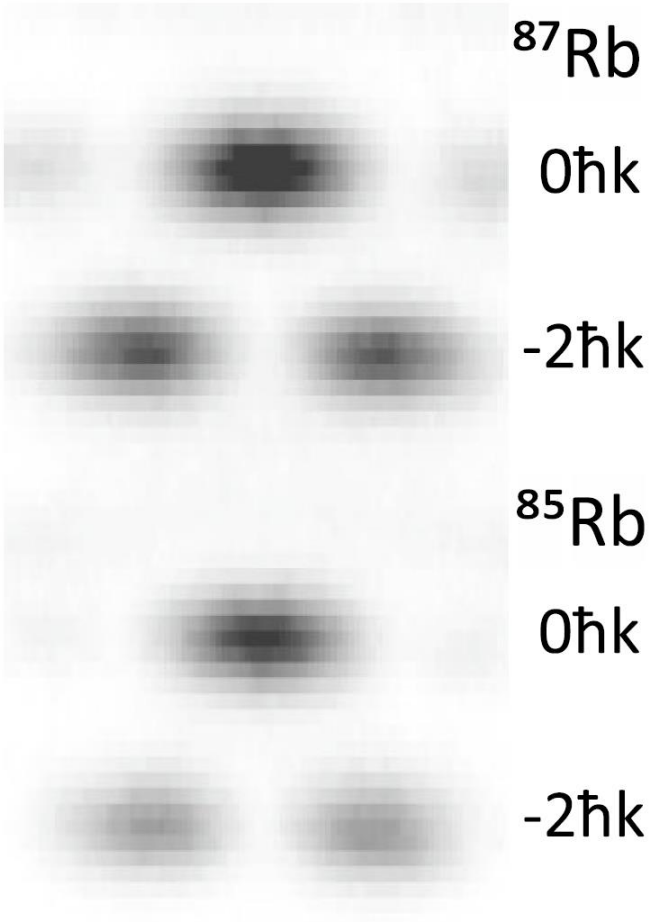
$\delta g \sim 10^{-15}$ g with 1 month data collection.

Limited by atom number.

Systematic uncertainties

$\delta g/g \sim 10^{-14}$ limited by magnetic field inhomogeneities, gravity anomalies, and light-shifts.



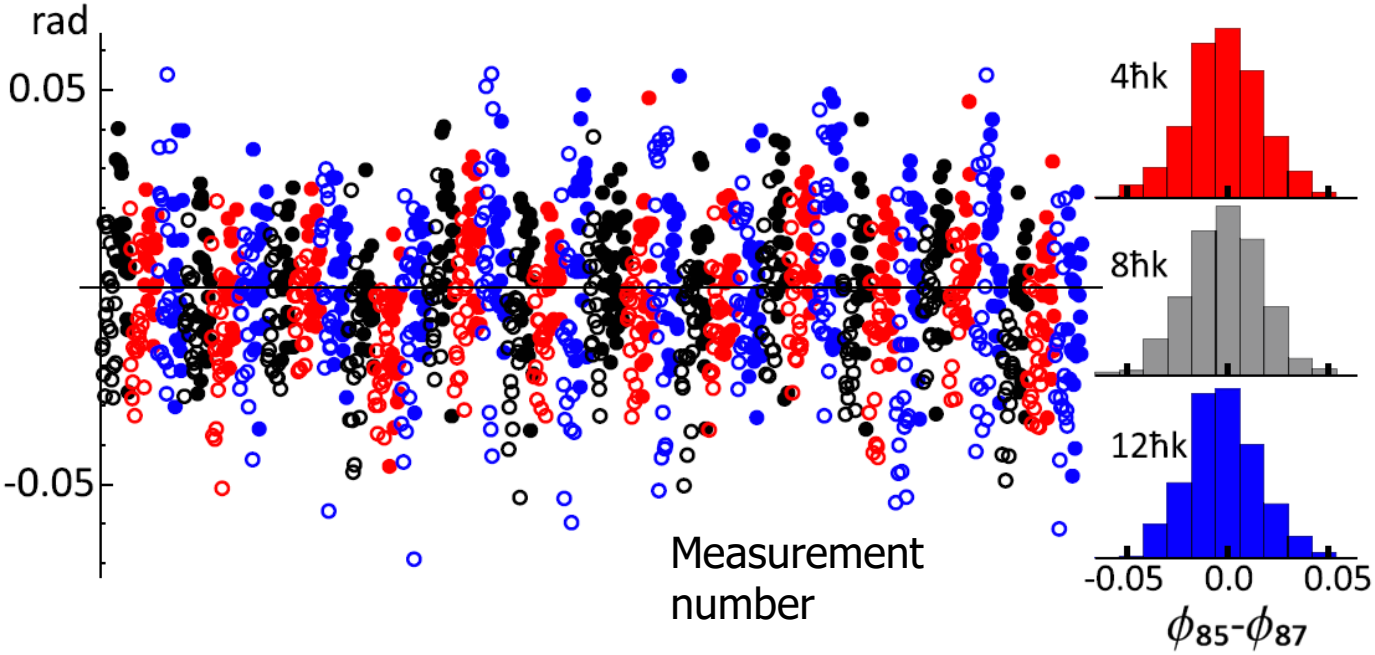


Overstreet, et al., PRL 2020

Data

The differential accelerations of 85Rb and 87Rb are inferred by comparing phase shifts for atom interferometers with differing wavepacket separation (recoil momenta).

This suppresses systematic phase shifts associated with the initial and final beam-splitting interactions.



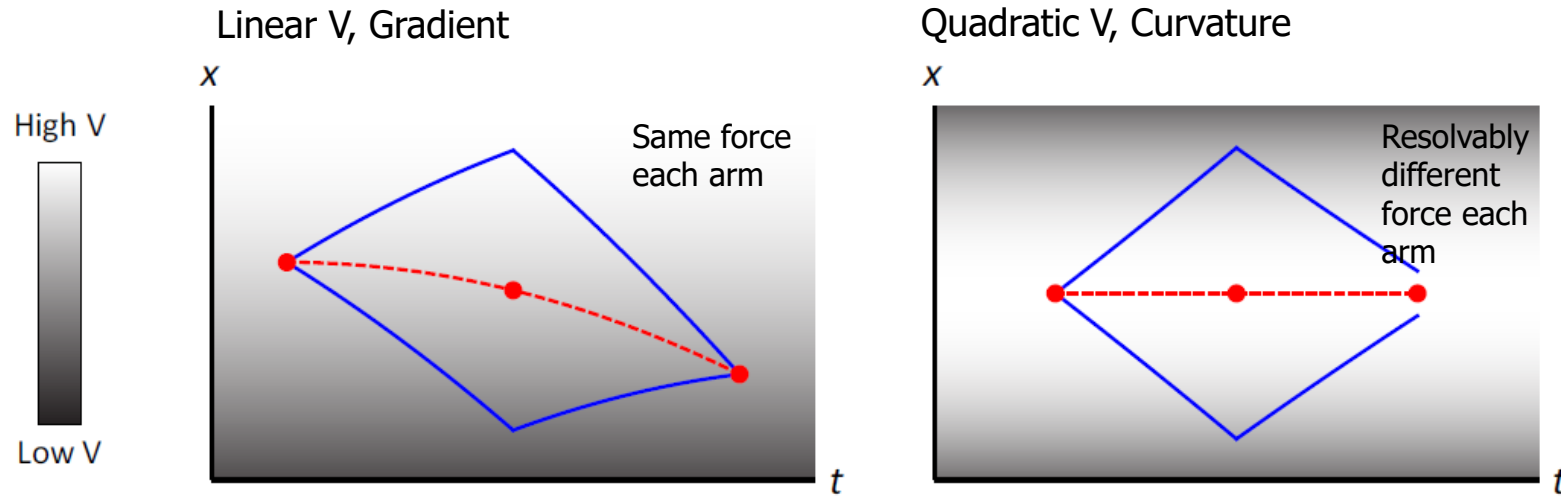
Equivalence Principle Test Results

$$\eta = [1.6 \pm 1.8(\text{stat}) \pm 3.4(\text{syst})] \times 10^{-12}$$

Parameter	Shift	Uncertainty
Total kinematic	1.5	2.0
Δz		1.0
Δv_z	1.5	0.7
Δx		0.04
Δv_x		0.04
Δy		0.2
Δv_y		0.2
Width		1.6
ac-Stark shift		2.7
Magnetic gradient	−5.9	0.5
Pulse timing		0.04
Blackbody radiation		0.01
Total systematic	−4.4	3.4
Statistical		1.8

Overstreet, et al., PRL 2020

Atom interferometer vs. classical measurements



In both cases, interferometer phase shift is well described by the classical trajectories associated with the interferometer arms:

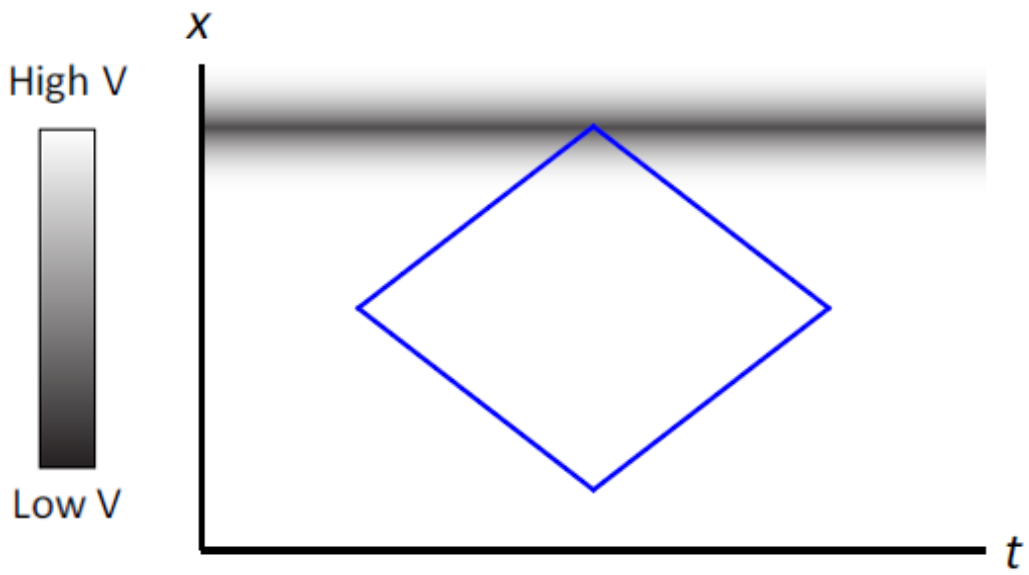
$$\phi_{\mathrm{MP}} \equiv \sum_{i=1}^{N} \left[\left(k_{1,i} - k_{2,i}\right) \bar{x}_i - \left(\omega_{1,i} - \omega_{2,i}\right) t_i + \left(\phi_{1,i} - \phi_{2,i}\right) \right].$$

(k_i and x_i are propagation vectors and wavepacket positions at the i^{th} pulse.)

These atom interferometric measurements are conceptually similar to classical measurements. Phase shift is given by the force acting on atomic wavepackets.

Antoine and Borde, JOSA B, 2013.
Overstreet, et al., AJP, 2021.

Mass dependent phase shifts



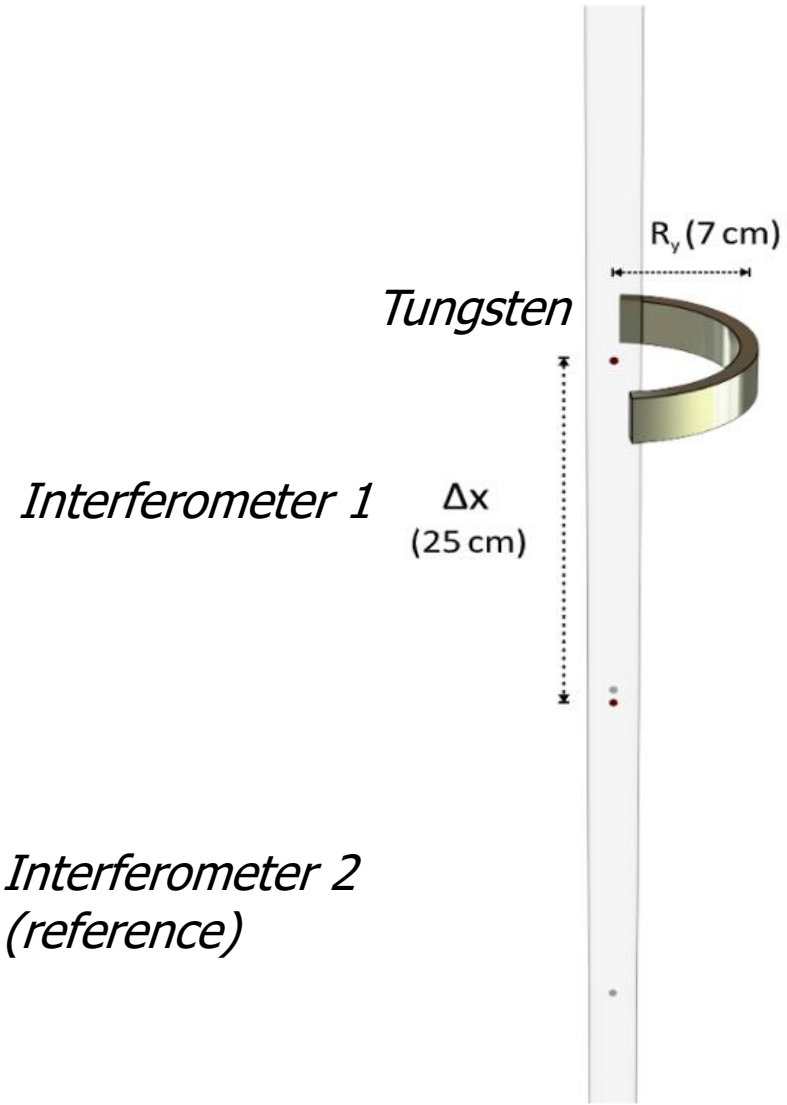
For higher order curvature, the phase shift is mass dependent.

Can be interpreted as a gravitational Aharonov-Bohm effect.

Possible systematic for future EP measurements based on atom interferometry.

Overstreet, et al., AJP, 2021.

Gravitational Aharonov-Bohm Experiment

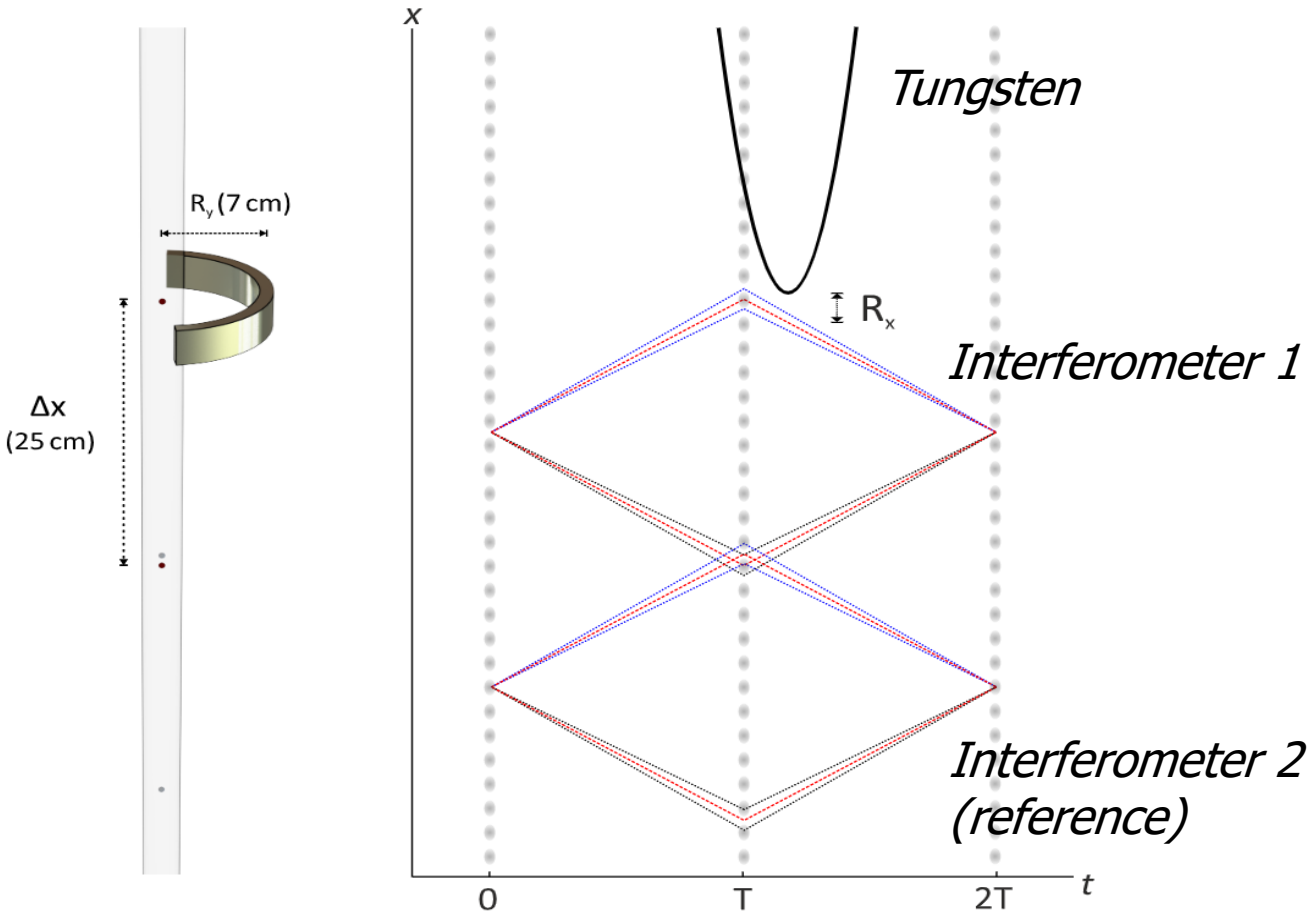


Wavepacket separation greater than distance of nearest wavepacket to source mass

Overstreet, et al., Science, 2022

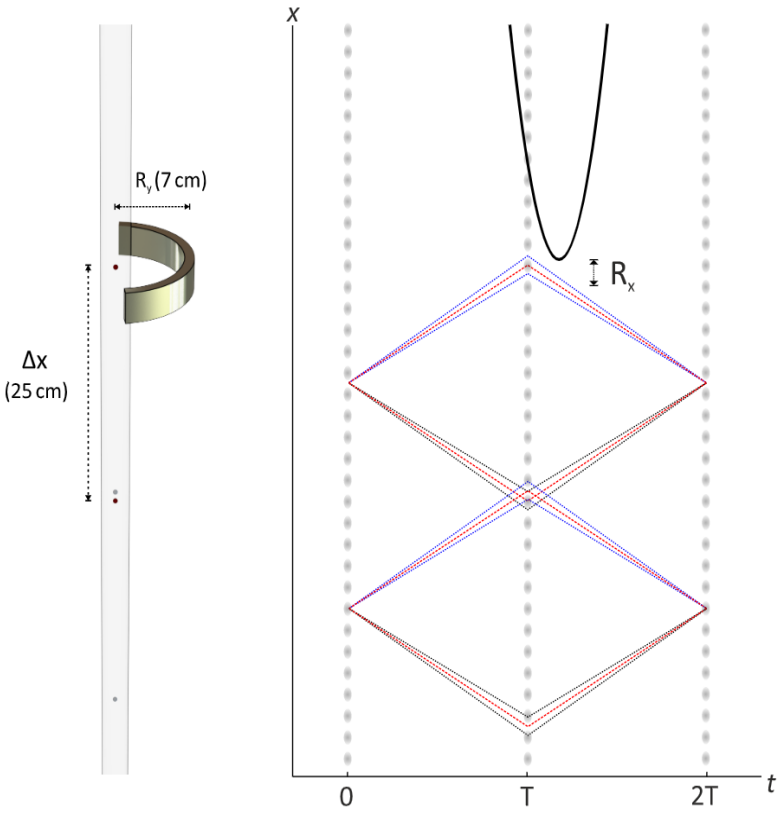
Prior proposals:
Audretsch and Lammerzhal, 1983
Hohensee, et al., 2012

Interferometer trajectories in freely falling frame



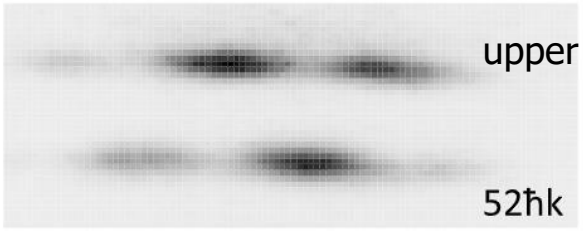
Phase shift due to gravitational action

Interferometer geometry

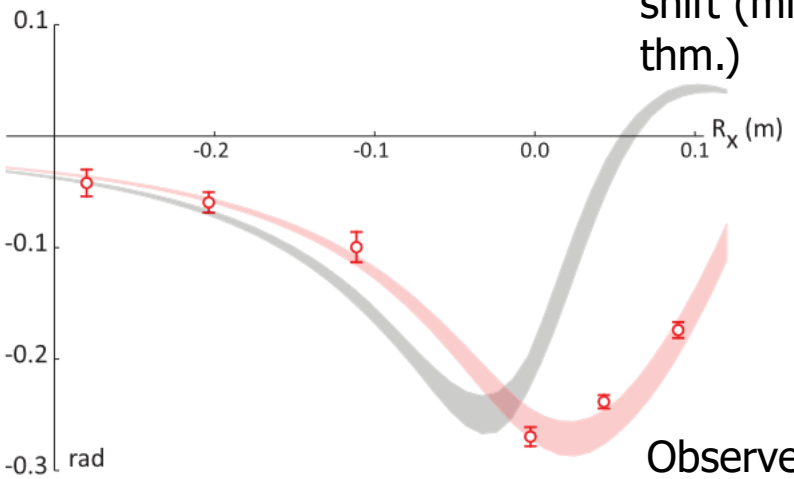


(~7e-12 g/shot resolution for each accelerometer)

Raw data



Phase shift

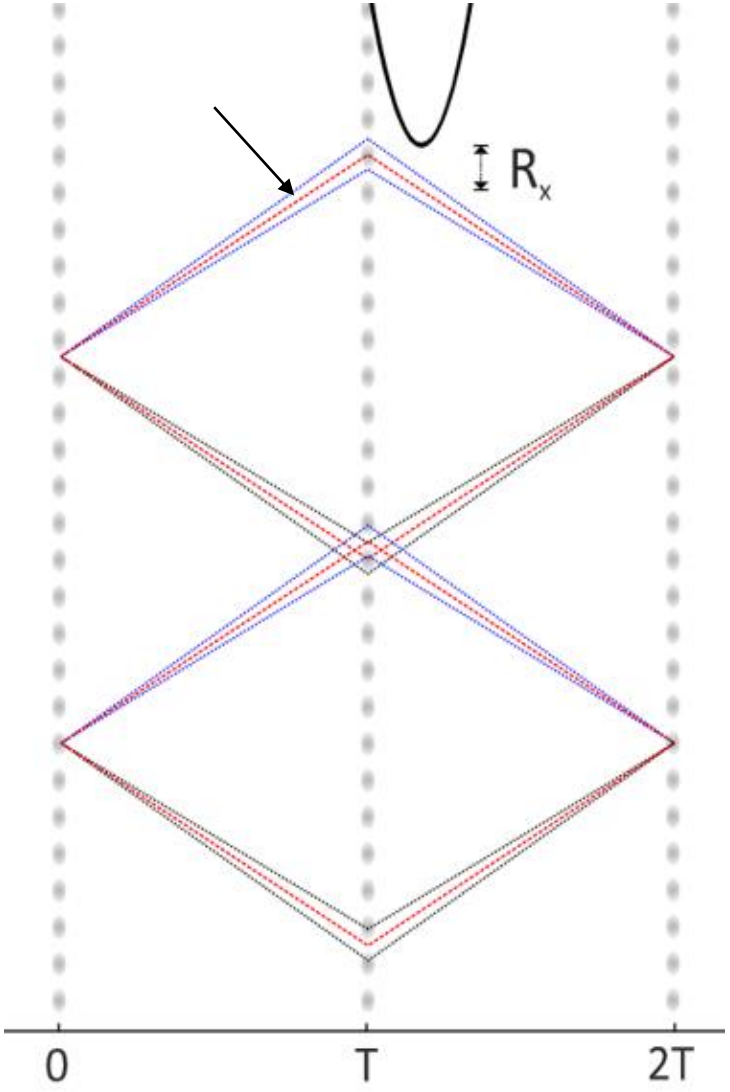


Deflection-induced phase shift (midpoint thm.)

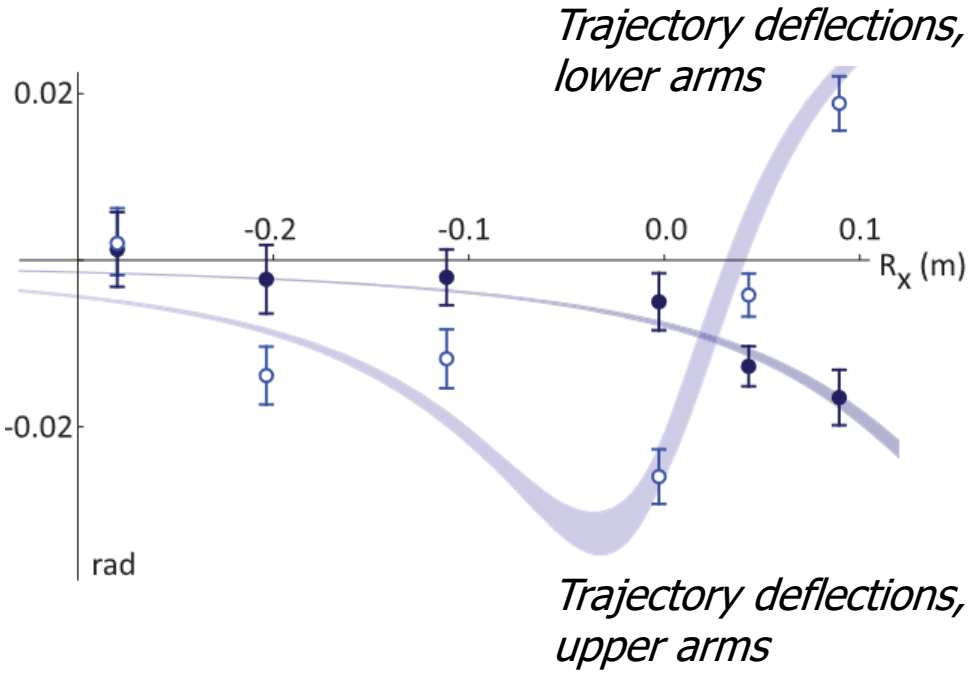
Observed phase shift

Overstreet, at al., Science, 2022

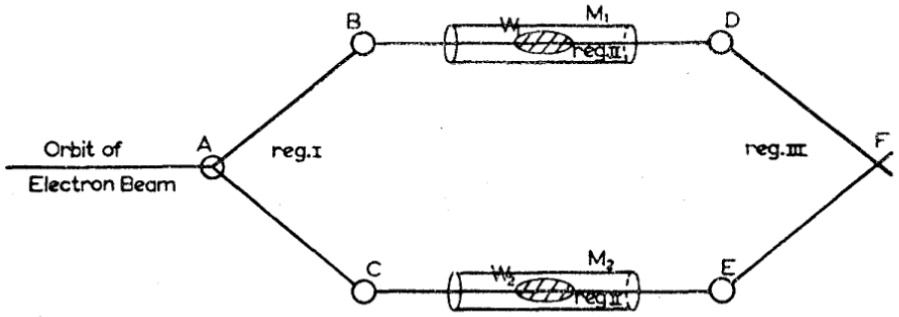
Deflection-induced phase shifts



Auxiliary (small wavepacket separation) interferometers allow independent characterization of deflection-induced phase shifts



Scalar Aharonov-Bohm Effect



$$\psi = \psi_1{}^0 e^{-iS_1/\hbar} + \psi_2{}^0 e^{-iS_2/\hbar}$$

$$S_1 = e\int \varphi_1 dt, \quad S_2 = e\int \varphi_2 dt.$$

Aharonov and Bohm, Phys. Rev. 1959

Negligible contribution to phase shift from forces on wavepackets (!)

One interpretation: physical original of the phase shift is the energy required to establish potential in the presence of the electron's electric field. Implies electron electric field is in superposition.

By analogy, observation of the gravitational Aharonov-Bohm shift implies the atom's gravitational field is in superposition.

Newtonian gravitational field energy

Field energy (weak field limit):

$$E_{\mathrm{G}} = -\frac{1}{8\pi G}\int |\mathbf{g}|^2 dV$$

$$\mathbf{g} = \mathbf{g}_{\mathrm{atom}} + \mathbf{g}_{\mathrm{tungsten}}$$

Phase shift:

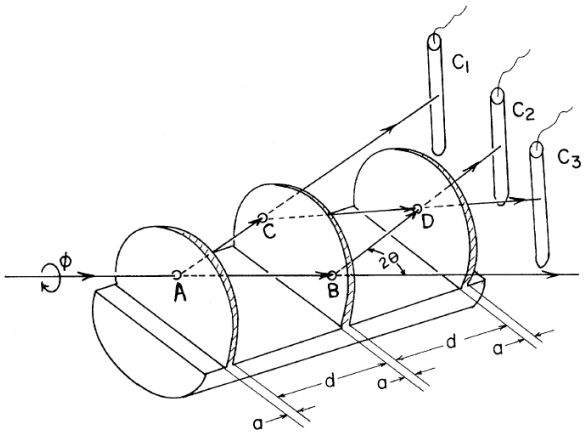
$$\phi = \frac{1}{\hbar}\int (E_1 - E_2)\, dt$$

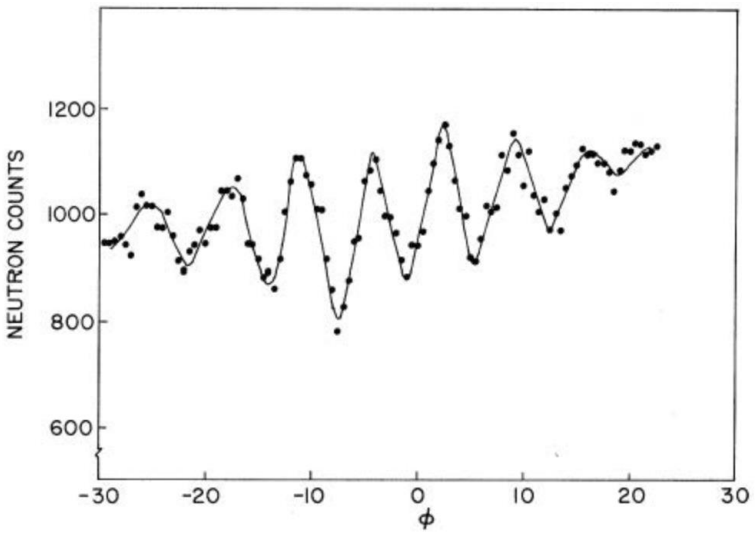
E_1, E_2 are gravitational energies for each arm.

Phase shift is interpreted as resulting from energy stored in superposed gravitational fields.

Overstreet, et al. arXiv:2209.02214

Collela, Overhauser and Werner (1975)





Uniform gravitational field implies gravitational action phase shift is zero -- uniform gravitational fields are not observable.

Physical original of phase shift: relative (kinematic) displacement of Si crystal with respect to de Broglie waves due to non-gravitational forces.*

*textbook treatments use perturbation theory, which masks the physical origin of the phase shift.

(Very) long baseline sensors

Exploit sensitivity for fundamental physics by comparing outputs of distant atom interferometers

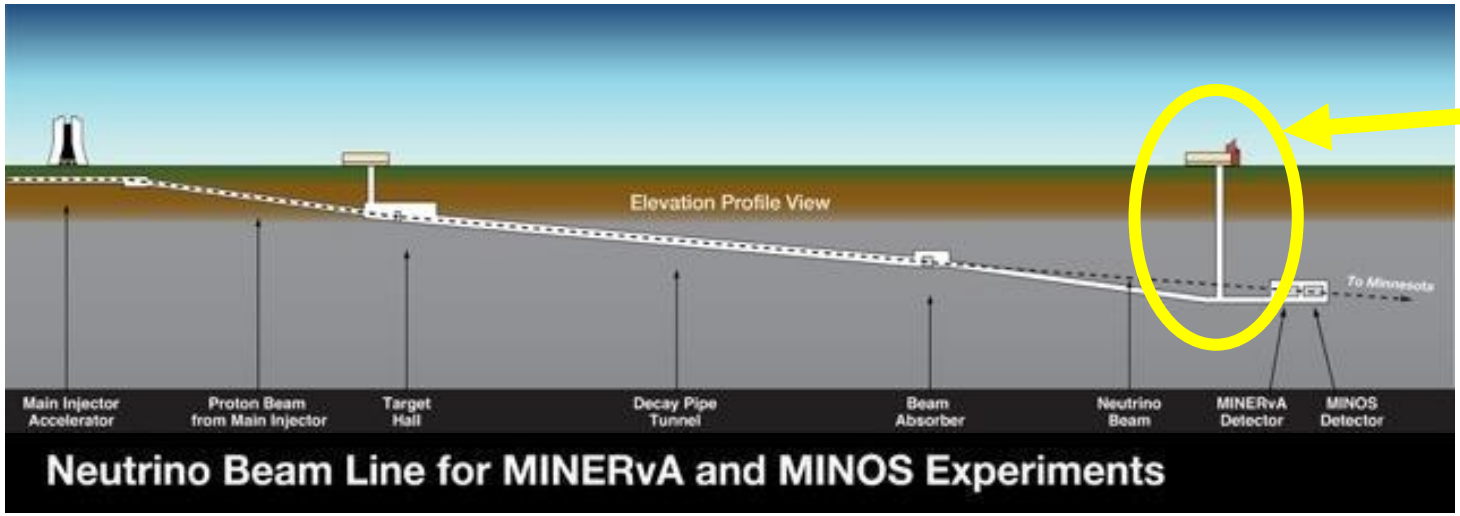
- gravitational wave detection
- ultra-light dark matter detection

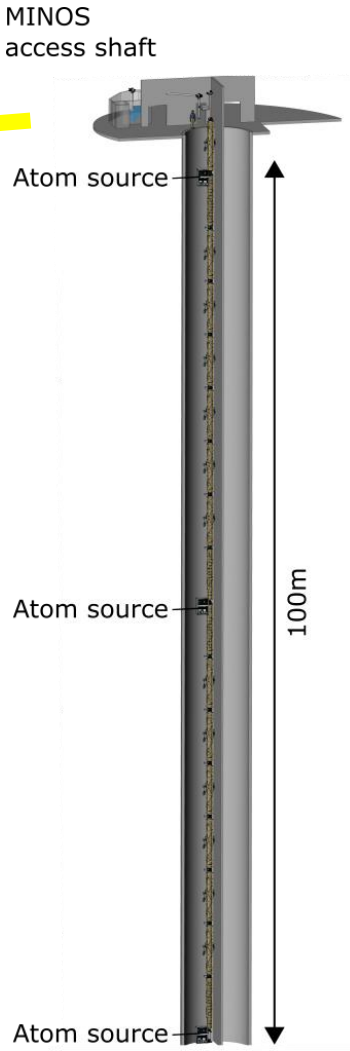
Physics reach scales with separation of interferometers

- L ~ 100 – 1000+ m for envisioned ground-based detectors
- Precision atom interferometers at each end of the baseline
- Separation limited by Earth rotation and seismic noise

MAGIS-100: Detector prototype at Fermilab

MAGIS-100



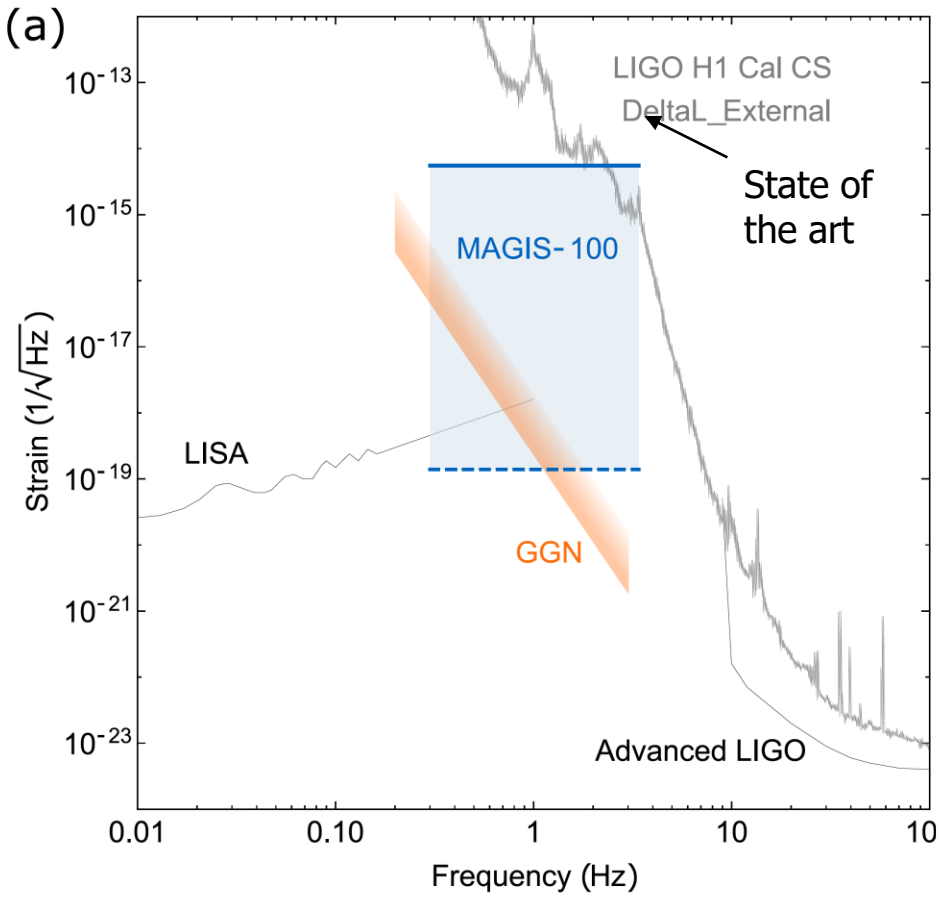


- 100-meter baseline atom interferometry in existing shaft at Fermilab
- Intermediate step to full-scale (km) detector for gravitational waves
- Atom sources (Sr) at three positions to realize a gradiometer
- Probes for ultralight scalar dark matter beyond current limits (Hz range)
- Extreme quantum superposition states: >meter wavepacket separation

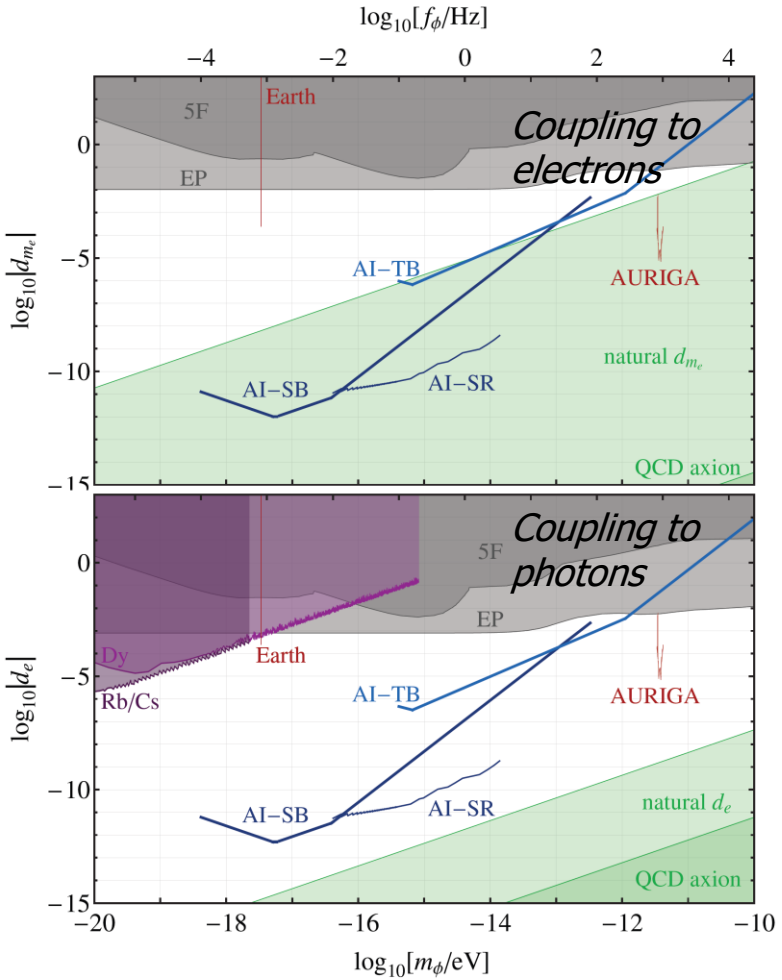
STANFORD · JOHNS HOPKINS UNIVERSITY · Northwestern University · Fermilab · SLAC · UNIVERSITY OF CAMBRIDGE · OLD DOMINION UNIVERSITY · UNIVERSITY OF OXFORD · UNIVERSITY OF LIVERPOOL

MAGIS-100 GW and DM sensitivity

Gravitational wave (GW) sensitivity

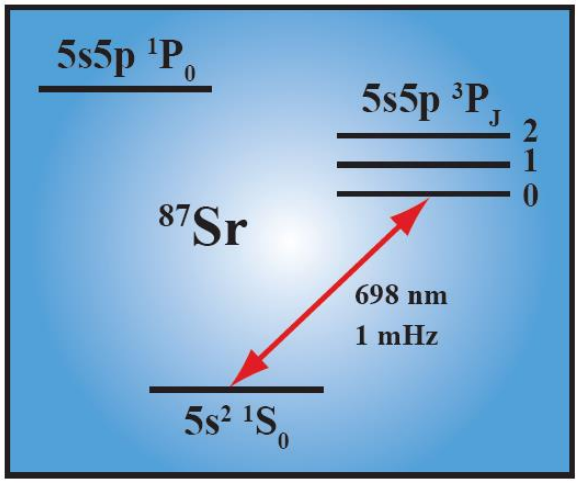


Dark matter (DM) sensitivity



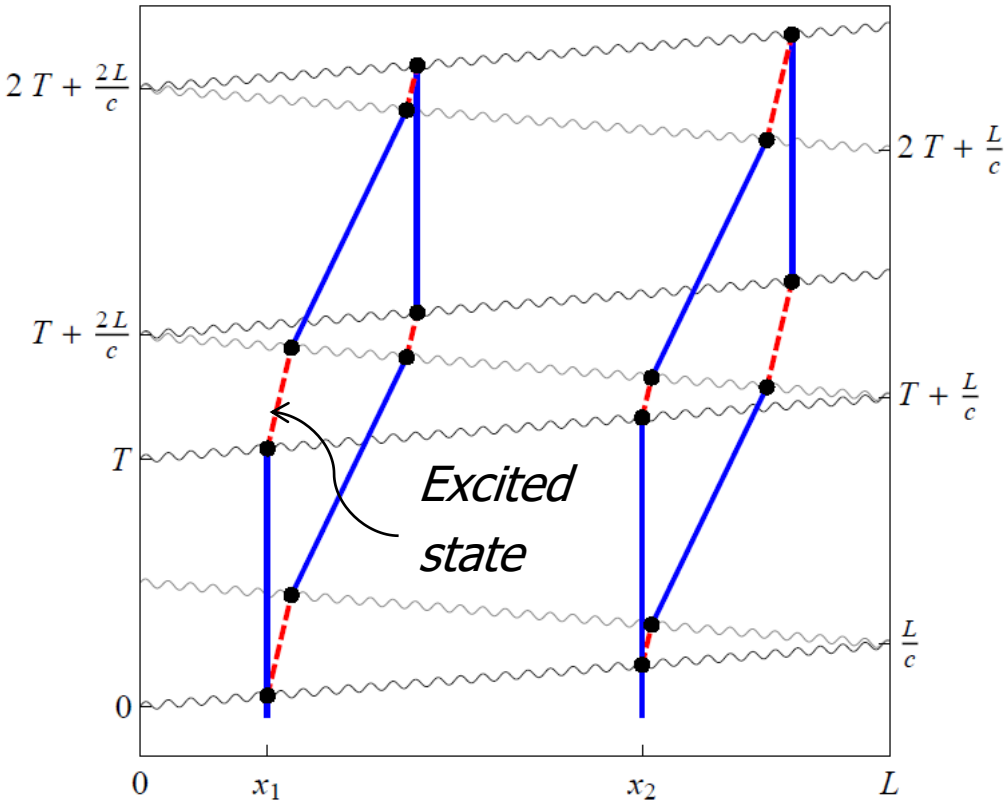
Laser frequency noise insensitive detector

Atomic excitation scheme



Laser phase noise common to both interferometers

Atom interferometer and laser space-time trajectories.



Graham, *et al.*, arXiv:1206.0818, PRL (2013)

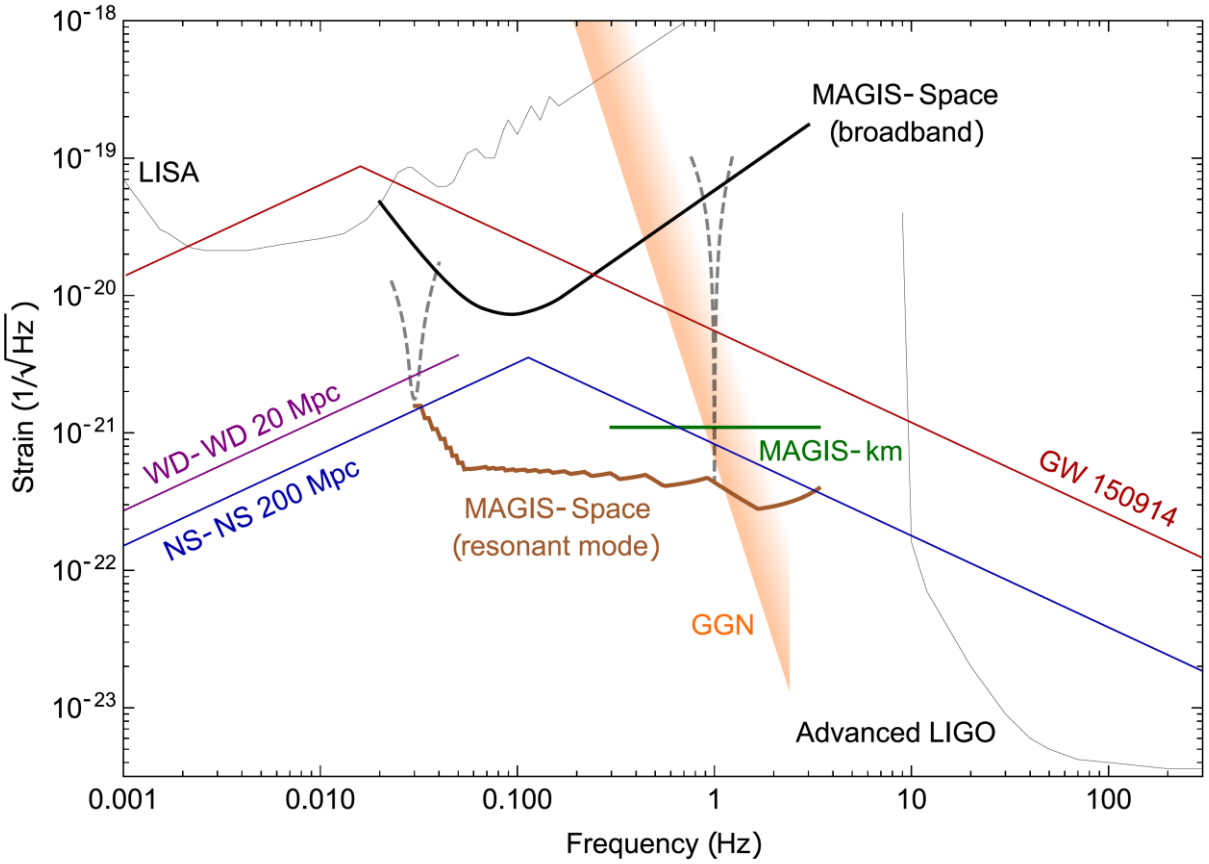
Gravitational wave (GW) response

$$\Delta\phi = \frac{4N\omega_a h}{c}(x_1 - x_2)\sin^2\left(\frac{\omega T}{2}\right)\sin(\phi_0 + \omega T)$$

Atomic transition freq.

GW phase

GW phase shift for interferometer pair



Literature: B. Lamine, et al., Eur. Phys. J. D **20**, (2002); R. Chiao, et al., J. Mod. Opt. **51**, (2004); S. Foffa, et al., Phys. Rev. D **73**, (2006); A. Roura, et al., Phys. Rev. D **73**, (2006); P. Delva, Phys. Lett. A **357** (2006); G. Tino, et al., Class. Quant. Grav. **24** (2007), Dimopoulos, et al., PRD (2008), Graham, et al., PRL (2013).

Ultra-light scalar field dark matter detection

PRL 117, 261301 (2016) PHYSICAL REVIEW LETTERS week ending 23 DECEMBER 2016

Sensitivity of Atom Interferometry to Ultralight Scalar Field Dark Matter

Andrew A. Geraci and Andrei Derevianko
Department of Physics, University of Nevada, Reno, Nevada 89557, USA
(Received 24 September 2016; published 20 December 2016)

We discuss the use of atom interferometry as a tool to search for dark matter (DM) composed of virialized ultralight fields (VULFs). Previous work on VULF DM detection using accelerometers has considered the possibility of equivalence-principle-violating effects whereby gradients in the dark matter field can directly produce relative accelerations between media of differing composition. In atom interferometers, we find that time-varying phase signals induced by coherent oscillations of DM fields can also arise due to changes in the atom rest mass that can occur between light pulses throughout the interferometer sequence as well as changes in Earth's gravitational field. We estimate that several orders of magnitude of unexplored phase space for VULF DM couplings can be probed due to these new effects.

DOI: 10.1103/PhysRevLett.117.261301

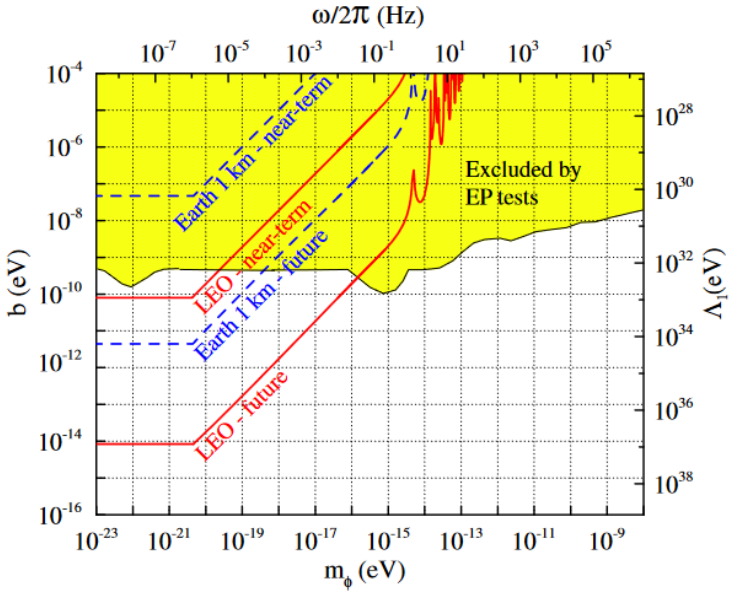
Physical origin of the sensitivity:

$$g(t) = g_0[1 + \delta_g \cos(\omega t + \theta_0)],$$

$$m(t) = m_0[1 + \delta_m \cos(\omega t + \theta_0)],$$

Corresponding atom interferometer response:

$$\Delta\phi \approx -k_{\rm eff} g_0 (1+\delta_g) T^2 - \delta_m \frac{k_{\rm eff}}{\omega}(v_L + v_R/2)(\omega T)^3.$$



Also, clock coupling:
(Arvanitaki et al. 2018)

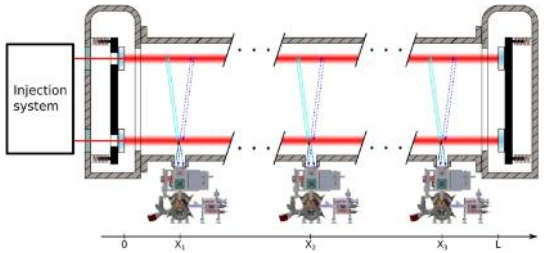
$$\omega_{\rm A}(t) \simeq \omega_{\rm A} + \Delta\omega_{\rm A} \cos(m_\phi t);$$

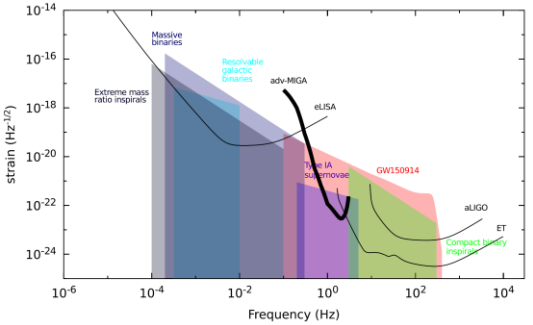
$$\Delta\omega_{\rm A} \equiv \omega_{\rm A} \sqrt{4\pi G_N}\phi_0 \left(d_{m_e} + \xi d_e\right).$$

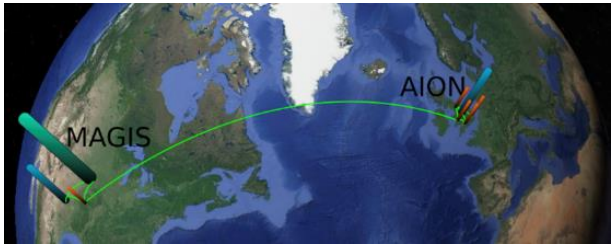
International efforts in long baseline atomic sensors

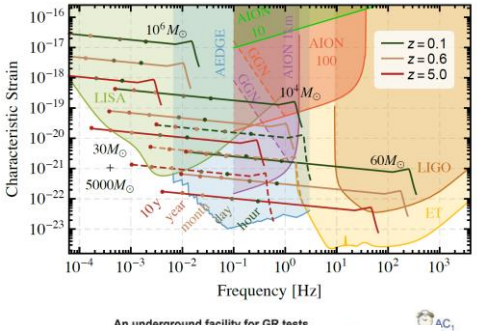
Project	Baseline Length	Number of Baselines	Orientation	Atom	Atom Optics	Location
MAGIS [ ]	100 m	1	Vertical	Sr	Clock AI, Bragg	USA
AION [ ]	100 m	1	Vertical	Sr	Clock AI	UK
MIGA [ɔ]	200 m	2	Horizontal	Rb	Bragg	France
ZAIGA [δ]	300 m	3	Vertical	Rb, Sr	Raman, Bragg, OLC	China

MIGA: Matter Wave laser Interferometric Gravitation Antenna (France)



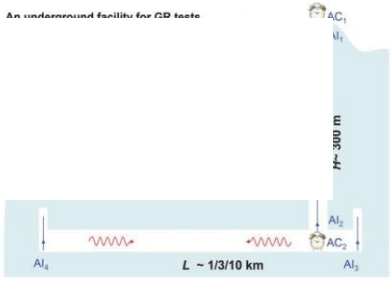


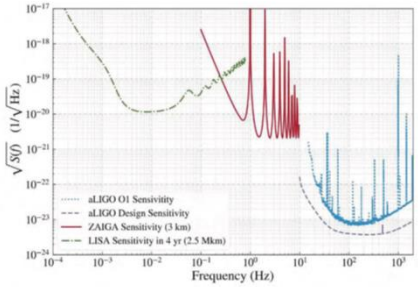




AION: Atom Interferometer Observatory and Network (UK)

ZAIGA: Zhaoshan Long-baseline Atom Interferometer Gravitation Antenna (China)





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