Newtonian Noise Simulations

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14 March 2023 Terrestrial Very-long-baseline atom interferometry workshop







Overview

Goal: Briefly summarize state of Newtonian Noise simulations

- Background What is Newtonian Noise?
- Current studies Sources and models
- Simulations Time domain, Frequency domain
- What's next?



- Density perturbations affecting acceleration of test masses
- Modeled through continuity equation

$$\delta\phi(\boldsymbol{r}_0) = G \int dV \frac{\delta\rho(\boldsymbol{r})}{|\boldsymbol{r} - \boldsymbol{r}_0|}$$



Sources

Anthropogenic

Human foot-fall

Motors/pumps

Vehicles Trains/Planes/Automobiles

Earth generated

Seismic: Long-duration standing waves Transient impacts (earthquakes)

Atmospheric: Pressure/Infrasound waves

Atmospheric: Temperature gradients



Anthropogenic



* Or Rat foot-fall (See poster by John Carlton, KCL)



Earth generated



Note: Gravity gradient noise (GGN) usually in reference to indirect coupling



Earth Generated



Workshop dedicated to investigating the interplay between Earth Sciences and atom interferometry

Atom Interferometric Sensing of Earth's Spheres





Natural Environment Research Council

КК

 Research into Newtonian Noise strongly initiated by laser interferometry gravitational wave community

Saulson. Phys. Rev. D **30,4**, (August 1984)

Hughes, Thorne. Phys. Rev. D 58, 122002 (1998)

Teviet Creighton. 2008 Class. Quantum Grav. 25 125011

Harms. Living Reviews in Relativity Volume 22, Article number: 6 (2019)

 First investigations for atom interferometry advanced by MIGA also being investigated by MAGIS and VLBAI

MIGA. Proc. SPIE 9900, Quantum Optics, 990008 (29 April 2016) Mitchell et al. 2022 *JINST* **17** P01007



Simulations

- 1. Choose a source
- 2. Derive gravitational potential perturbation
- 3. Simulate AI phase shift in response
- 4. Study output

 $\Delta \phi = k_{eff} g T^2 \rightarrow k_{eff} (g + \delta g) T^2$

Models give us intuition on experimental parameter scaling

Simulation outputs informs necessary passive levels and site requirements

Highlights needs for extra monitoring systems

Used to understand feasible methods of mitigation



Simulations



$$\Phi_{\mathrm{GGN},m}^{(i)} = \sum_{a} \xi_a \left(\widetilde{A}_a \exp\left(-q\frac{\omega z_i}{c_H}\right) + \widetilde{B}_a \exp\left(-\frac{\omega z_i}{c_H}\right) \right) \cos\widetilde{\phi}_{a,m} \, .$$

$$\widetilde{A}_a = -\frac{16 G \pi \rho_0}{\omega_a^2 (s^2 - 1)} n k_A q (1 + s^2) \sin\left(\frac{\omega_a T}{2}\right)^2$$
$$\widetilde{B}_a = \frac{8 G \pi \rho_0}{\omega_a^2 q (s^2 - 1)} n k_A (1 + (1 + 2q)s^2) \sin\left(\frac{\omega_a T}{2}\right)^2$$
$$\widetilde{\phi}_{a,m} = \omega_a (T + m \Delta t) + \widetilde{\theta}_a ,$$

Badurina, et al. Phys. Rev. D **107**, 055002



g

100 m - 1 km

Simulations guide us

- Methods predominantly used:
 - Fourier analysis and power spectral densities
 - Filtering via Weiner filters
 - Kalman filtering for active feedback and feedforward
- Serve to derive strategies for measuring and controlling





- Generate general formalisms for use with input spectra
- Find transfer functions for various atom interferometry configurations
- Understand the atmospheric effects and turbulence more robustly





- Newtonian noise and specifically GGN is a crucial aspect of very-longbaseline AI detectors
- Understanding impact theoretically and experimentally helps us to prepare for data analysis
- Simulations provide framework to guide designs and instrumentation

Thanks: Kavli-Newton Trust Tim Kovachy, Jason Hogan, Peter Graham Chris McCabe, Leonardo Badurina Swapan Chattopadhyay Val Gibson









