

Baryon number violation and neutron oscillations

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## Sakharov's baryogenesis conditions

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- **)** Baryon number violation ( $\Delta B 
  eq 0$ )
- C and CP symmetry violation
- Departure from thermal equilibrium



Sakharov (1967)

assuming "total CPT invariance of the (expanding) Universe"



## Sakharov's baryogenesis conditions

Baryon number violation and neutron oscillations

- Baryon number violation ( $\Delta B 
  eq 0$ )
- C and CP symmetry violation
- Opposition of the second se



Sakharov (1967)

#### assuming "total CPT invariance of the (expanding) Universe"

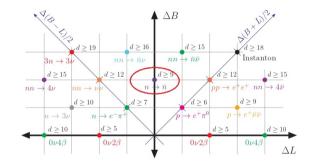
Standard Model:

- B and L are automatically conserved in (renormalizable) couplings
- B L conserved also by non-perturbative effects



## Beyond the Standard Model: $\Delta(B - L) \neq 0$

Baryon number violation and neutron oscillations Landscape of baryon and lepton number violation



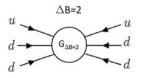
adapted from J. Heeck and V. Takhistov, PRD 101, 015005 (2020), 1910.07647 [hep-ph].

Some representative processes and the minimal mass dimension d of the underlying EFT operator



#### Neutron-antineutron oscillations

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- Six-fermion effective operator  $\frac{1}{M^5}(udd)(udd)$ , M > 1 TeV
- Effective renormalized  $\Delta B = 2$  quadratic Lagrangian for Dirac field n(x):

$$\mathcal{L} = \overline{n}(x)i\gamma^{\mu}\partial_{\mu}n(x) - m\overline{n}(x)n(x) - \epsilon[n^{T}(x)Cn(x) + \overline{n}(x)C\overline{n}^{T}(x)]$$

$$\epsilon = \langle \textit{n}|(\textit{udd})(\textit{udd})|\textit{n}
angle \sim rac{\Lambda_{QCD}^6}{M^5} = rac{1}{ au_{nar{n}}}$$

• nn oscillations distabilize the nuclei

$$(A,Z) 
ightarrow (A-1,ar{n},Z) 
ightarrow (A-2,Z/Z-1) + {
m pions}$$



Baryon number violation and neutron oscillations Present experimental status:

 ILL (Grenoble): free neutron oscillations – the cleanest experimental and theoretical environment to perform the search

 $au_{nar{n}}$  > 2.7 years (8.7 imes 10<sup>7</sup> s)

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Baldo-Ceolin et al., ILL (1994)
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New experiment at ESS promises a sensitivity in oscillation probability up to three orders of magnitude greater than ILL (some proposals go up to 5 orders!).

• Super-Kamiokande: *n* $\bar{n}$  oscillation in <sup>16</sup>O

 $au_{n\bar{n}} >$ 8.6 years

Abe et al., Super-Kamiokande collab. (2011)

• Sudbury Neutrino Observatory: nn oscillation in deuterium

 $\tau_{n\bar{n}} >$  4.1 years

Aharmim et al., SNO collab. (2017)



## $N\bar{N}$ oscillations: short chronology

Baryon number violation and neutron oscillations Anca Tureanu "CP-noninvariance and baryon asymmetry of the Universe"

Kuzmin (1970)

"Neutron Oscillations and the Existence of Massive Neutral Leptons"

Kuo and Love (1980)

"B - L nonconservation and neutron oscillation"

Chang and Chang (1980)

"Phenomenology of neutron oscillation"

Marshak and Mohapatra (1980)



#### Some more recent models:

• Models with large (TeV scale) extra dimensions

Dvali, Gabadadze (2002) Nussinov and Shrock (2002)

• Unified model  $SU(2)_L \times SU(2)_R \times SU(4)_c$  with TeV scale seesaw

Babu, Dev and Mohapatra (2009)

• SO(10) GUT scale seesaw with TeV scalars

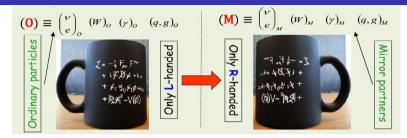
Babu and Mohapatra (2012)

Estimates of oscillation time:  $\tau_{n\bar{n}} = 3 - 30\ 000\ years$ 



# Beyond the Ordinary World: Mirror World

Baryon number violation and neutron oscillations



$$\left( SU(3)_{\textit{c}} imes SU(2)_{\textit{L}} imes U(1)_{\textit{Y}} 
ight) imes \left( SU(3)_{\textit{c}}' imes SU(2)_{\textit{L}}' imes U(1)_{\textit{Y}}' 
ight)$$

- Identical field contents (with opposite chirality) and Lagrangians
- Interactions between the O and M sectors by  $\mathcal{L}_{mix}$  (portals)

$$\mathcal{L}_{\textit{tot}} = \mathcal{L} + \mathcal{L}' + \mathcal{L}_{\textit{mix}}$$

• Mirror worlds can be many...

Dvali, Savicki and Vikman (2009)



# Mirror World as dark matter and LIGO unexpected events

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- $\bullet$  Mirror world reheating temperature lower (constraints from Big Bang Nucleosynthesis): T'/T < 0.64
- Mirror world can explain all dark matter:  $\Omega'_b/\Omega_b \approx 5$
- Stars are composed mainly of He, are more massive and evolve faster
- Number of stars:  $N'(m) \sim 5 \times N(m)$
- Number of BH:  $N_{BH}^\prime \sim 10 \times N_{BH}$

for a review on mirror world, see Berezhiani (2005)



#### Mirror world could provide explanation for LIGO puzzles

Merger objects:	BH-BH	NS-NS	BH-NS	BH-Mass gap
Number of events:	84	2	2	2

- Observed Merger Rates higher than theoretical predictions;
- Only one NS-NS observed event has electromagnetic counterpart
- Mass gap (BH or NS?) events observed



LIGO signals can come from the Mirror World!

Beradze and Gogberashvili (2019, 2021)



### Neutron-mirror neutron oscillations

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#### Berezhiani and Bento (2005)

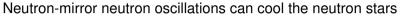
 $n - \bar{n}$  oscillations:  $\Delta B = 2$ 

 $n - \bar{n}'$ , n - n' oscillations:  $\Delta B = 1$ ,  $\Delta B' = 1$ 

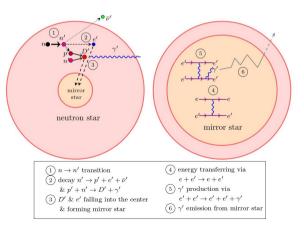
- $\mathcal{L}_{mix}$ : Dirac or Majorana-type mass terms
- The nn' oscillations can speed up in magnetic fields
- Current experimental limits:  $\tau_{nn'}, \tau_{n\bar{n}'} > 50 \ s$



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Goldman, Mohapatra, Nussinov and Zhang (2022)



$$\epsilon_{nn'} < 10^{-17} eV$$

$$\epsilon^2 \gg 10^{-27},$$

 $\epsilon$  is a minute charge of e'



• Most general effective Lagrangian with  $\Delta B = 2$  for neutrons:

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Barvon

number

$$\mathcal{L} = \bar{n}(x)i\gamma^{\mu}\partial_{\mu}n(x) - m\bar{n}(x)n(x) - \frac{i}{2}\epsilon_{1}[e^{i\alpha}\bar{n}^{T}(x)Cn(x) - e^{-i\alpha}\bar{n}(x)C\bar{n}^{T}(x)] - \frac{i}{2}\epsilon_{5}[n^{T}(x)C\gamma_{5}n(x) + \bar{n}(x)\gamma_{5}C\bar{n}^{T}(x)],$$

where *n* is Dirac field and *real* parameters *m*,  $\epsilon_1$ ,  $\epsilon_5$  and  $\alpha$ 

- The Lagrangian violates C, P, CP
- Claim: Neutron-antineutron oscillations violate CP: two of Sakharov's conditions are fulfilled!



• Observable CP violation  $\equiv$  different transition probabilities:

 $P(neutron \rightarrow antineutron) \neq P(antineutron \rightarrow neutron)$ 

• In spite of the intrinsic CP violation in the Lagrangian, the transition probabilities are the same:

 $P(neutron \rightarrow antineutron) = P(antineutron \rightarrow neutron)$ 

- Transfer CP violating chiral U(1) phase to parity violating mass term:  $im'\bar{n}\gamma_5 n$
- Intrinsic CP violation may be observed as contribution to EDM of neutron

Fujikawa and AT (2015, 2019)



#### • In chiral components,

$$-2\mathcal{L}_{mass} = \left( egin{array}{cc} \overline{n_R} & n_L^T(x)C \end{array} 
ight) \left( egin{array}{cc} M_R^\dagger & M_D \ M_D & M_L \end{array} 
ight) \left( egin{array}{cc} C \overline{n_R}^T \ n_L \end{array} 
ight) + h.c.$$

diagonalized using mixing matrix

$$U = \begin{pmatrix} \cos \alpha & \sin \alpha \\ -\sin \alpha & \cos \alpha \end{pmatrix} \begin{pmatrix} 1 & 0 \\ 0 & e^{i\beta} \end{pmatrix}, \quad \beta - \text{Majorana phase}$$

Fujikawa and AT (2021)

- Consider simultaneously nn and nn', nn' oscillations
  - $4 \times 4$  mixing matrix
  - 3 Dirac phases, 3 Majorana phases

CP violation in neutron oscillations can be achieved

Kupiainen and AT (in preparation)