Neutrinos from the Sun as a discovery tool for dark matter-electron scattering

Based on: TNM, A K Saha, S Mondal, R Laha; in preparation



DM



Tarak Nath Maity

Dark Interactions

New Perspectives from Theory and Experiment





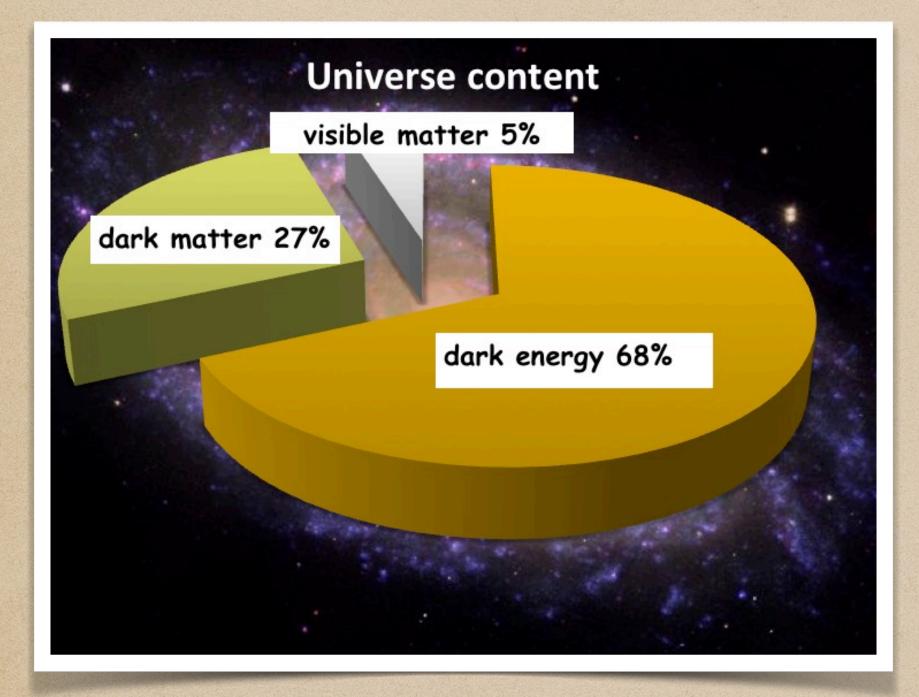


Image credit: QUANTUM DIARIES



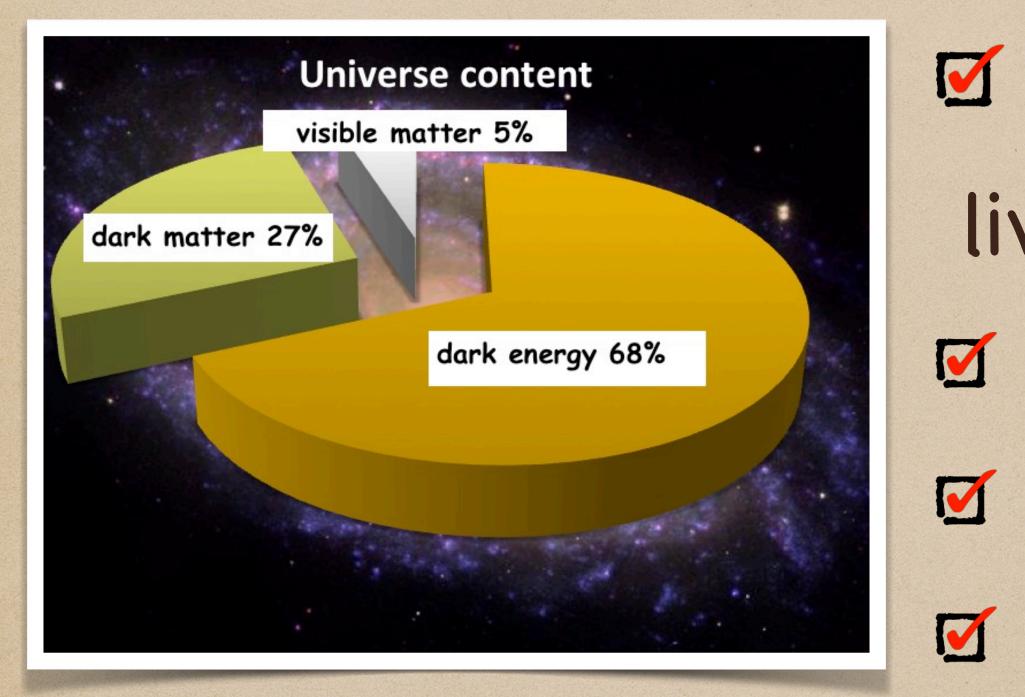


Image credit: QUANTUM DIARIES

Stable: No decay, very longlived

Cold: Non-relativistic
Massive: Wide range



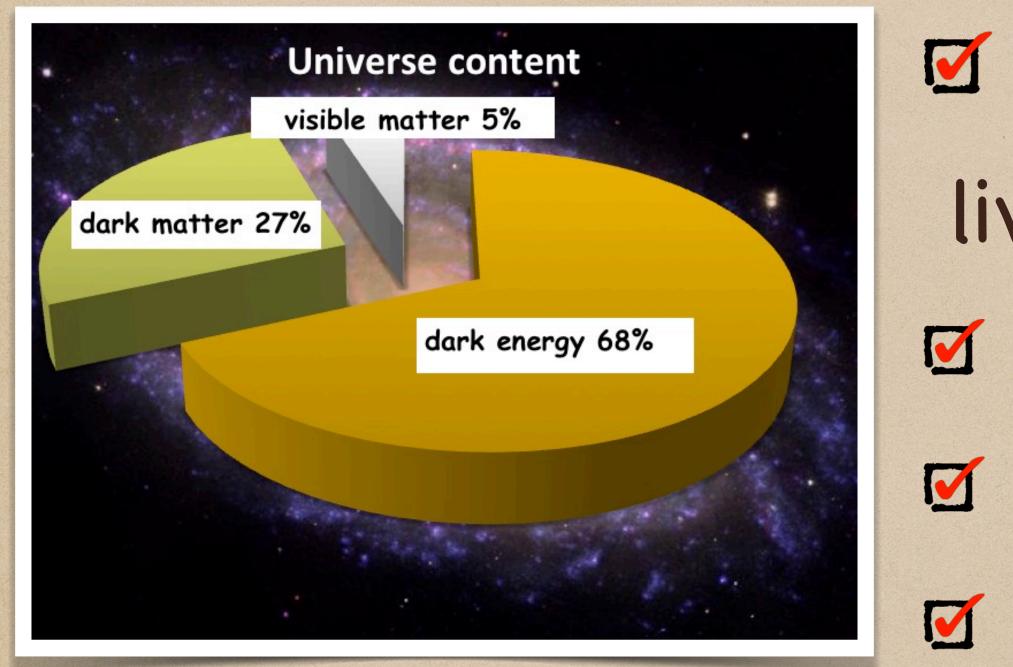


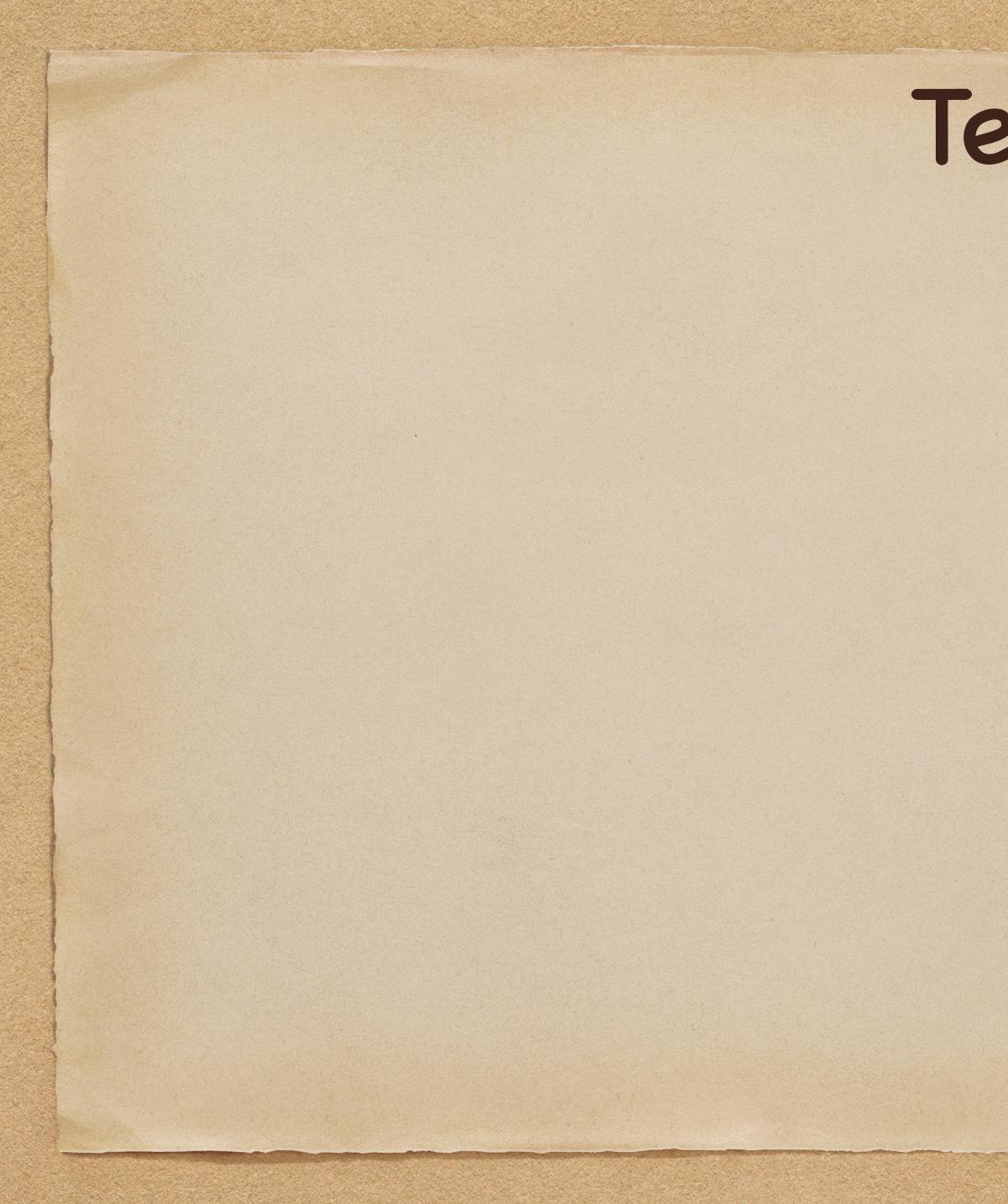
Image credit: QUANTUM DIARIES

Assume dark matter standard model coupling

Stable: No decay, very longlived

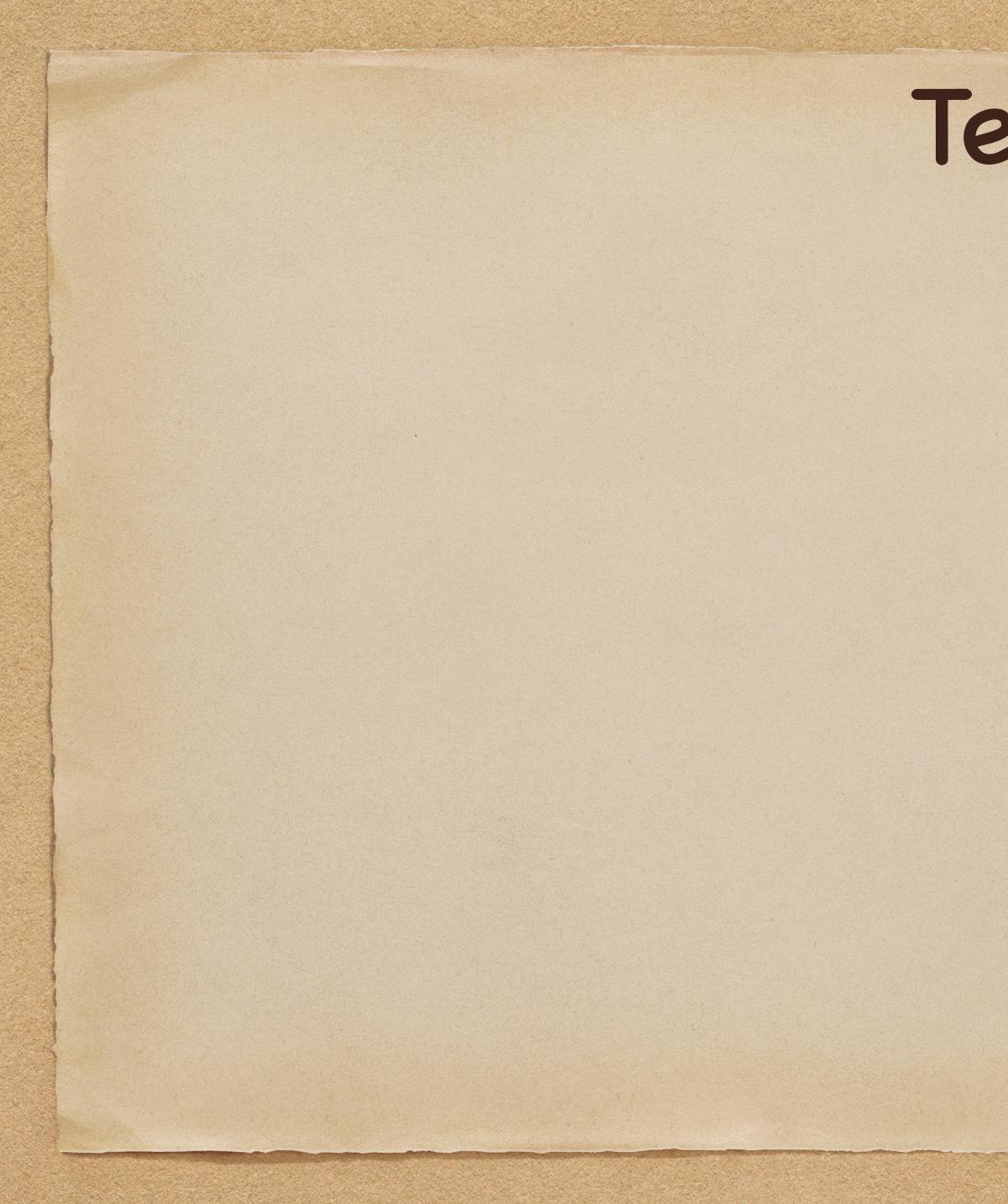
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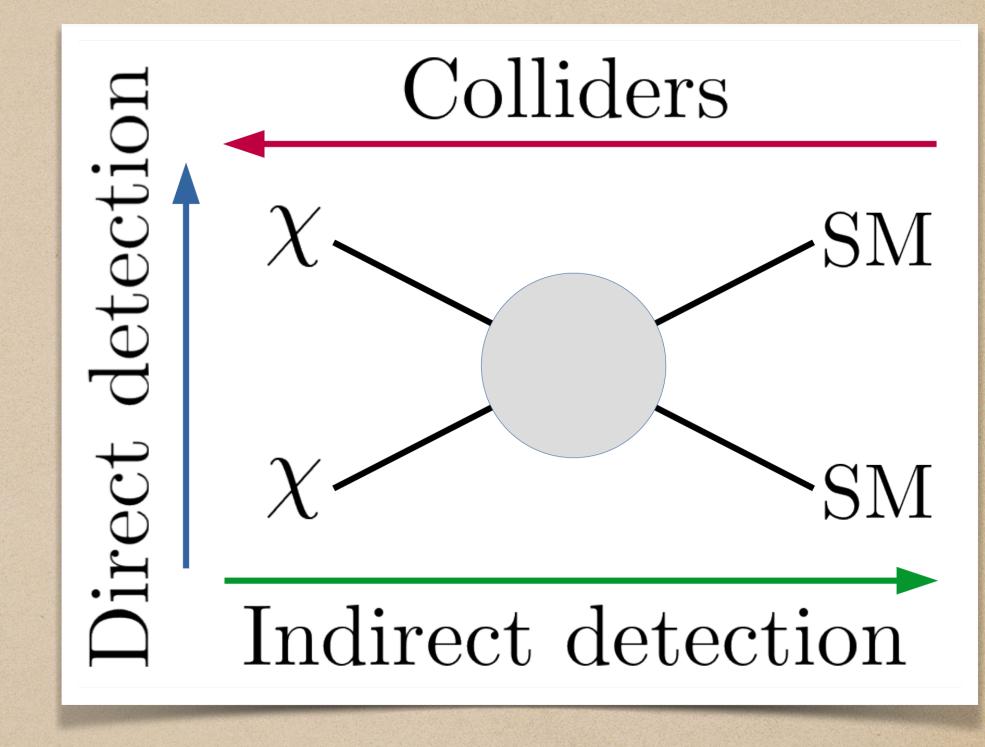


Test it





Test it



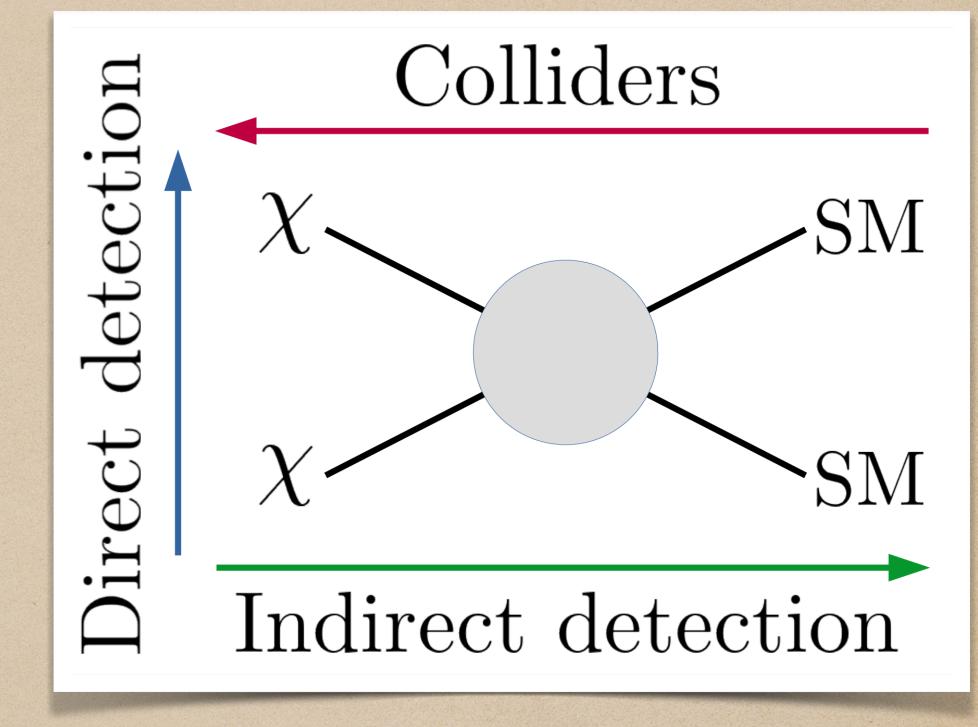
2104.11488





Direct detection Mostly focuses on DMnuclear scattering Made extraordinary progress

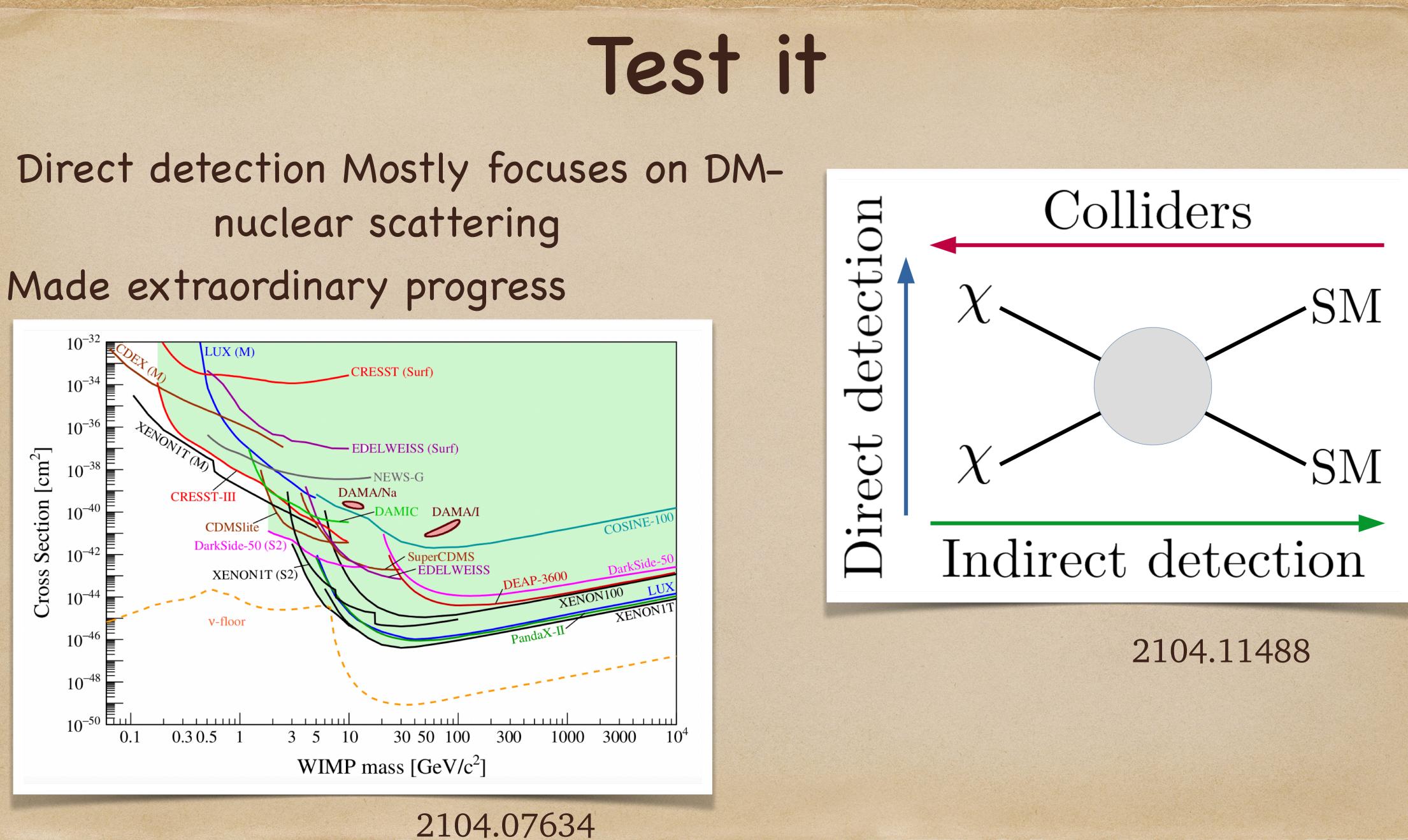
Test it



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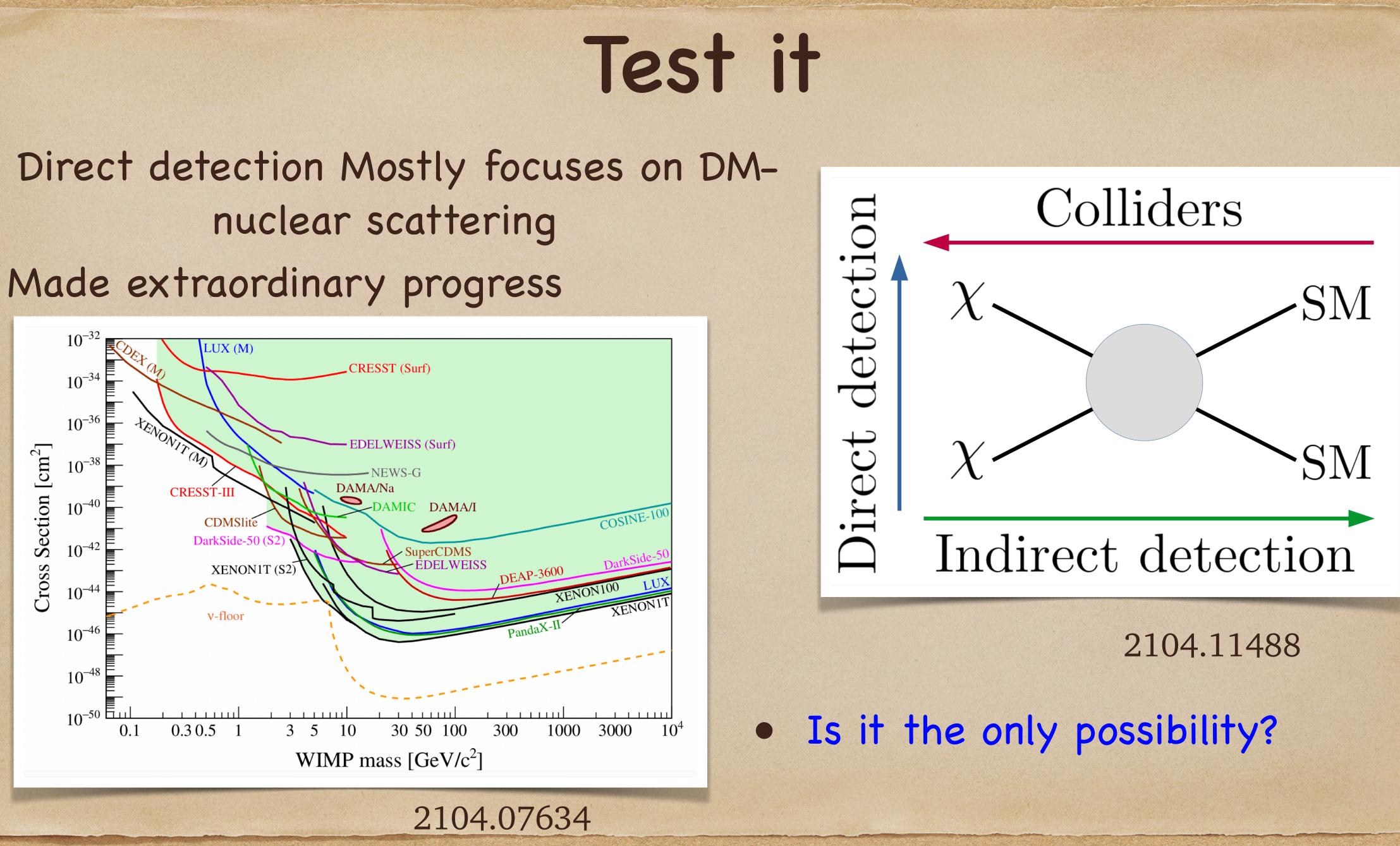


nuclear scattering





nuclear scattering



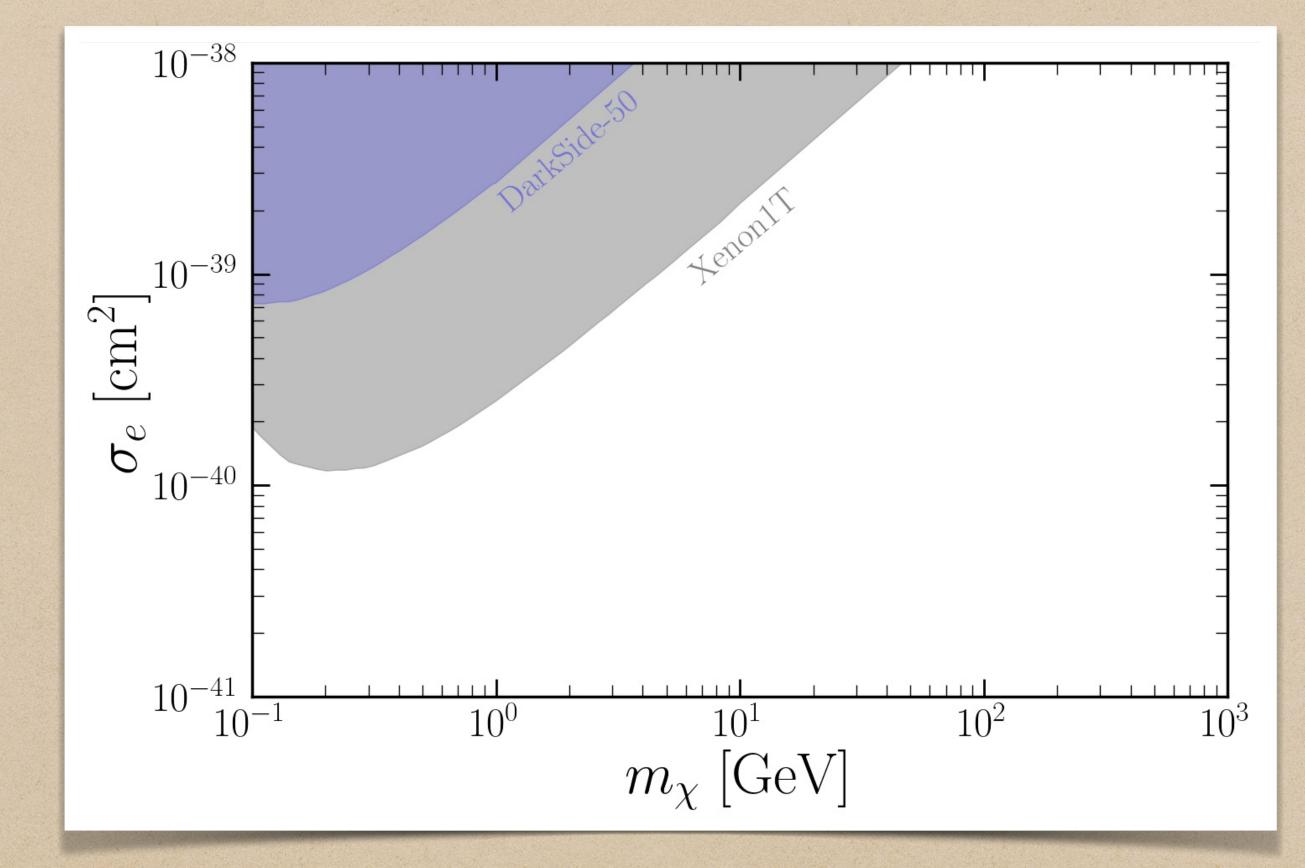
3



DM-electron scattering

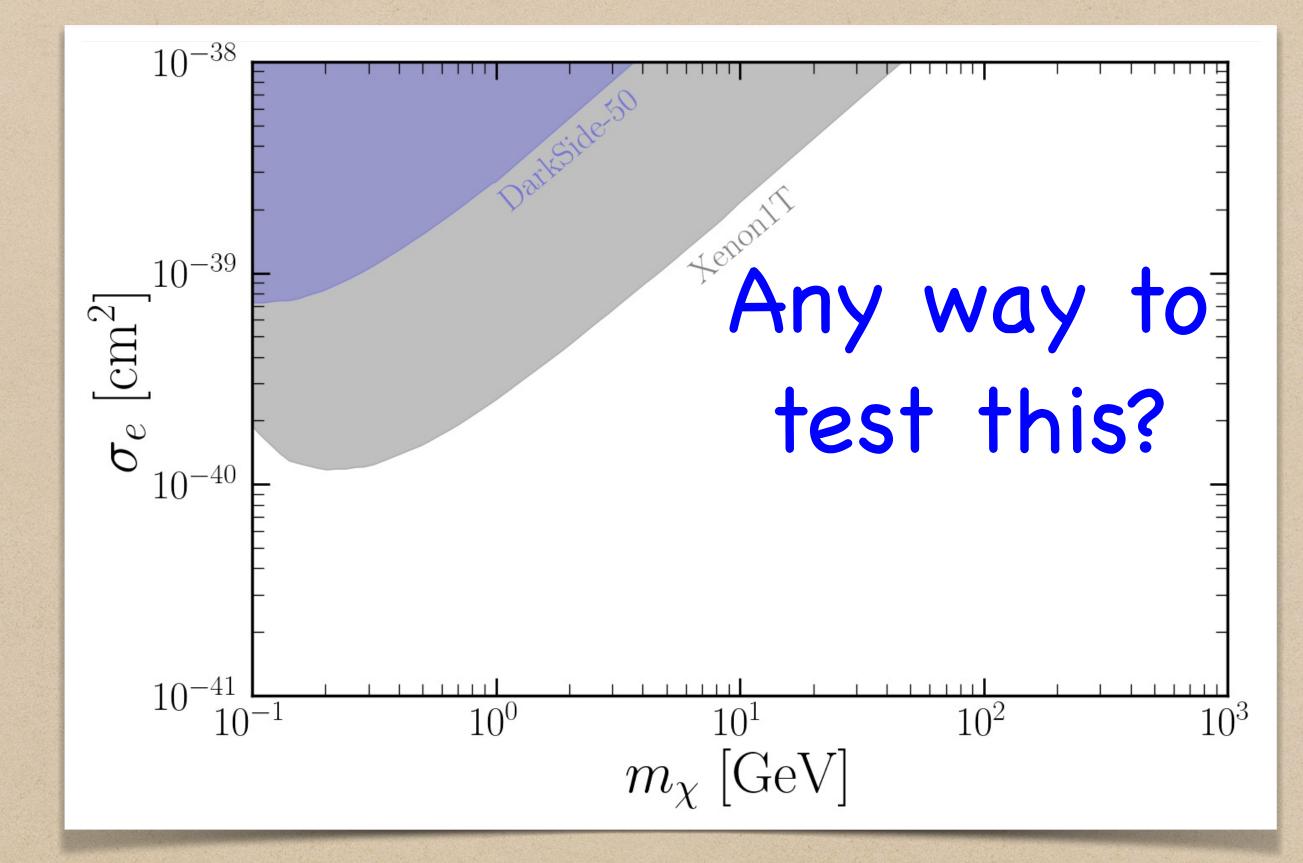


DM-electron scattering



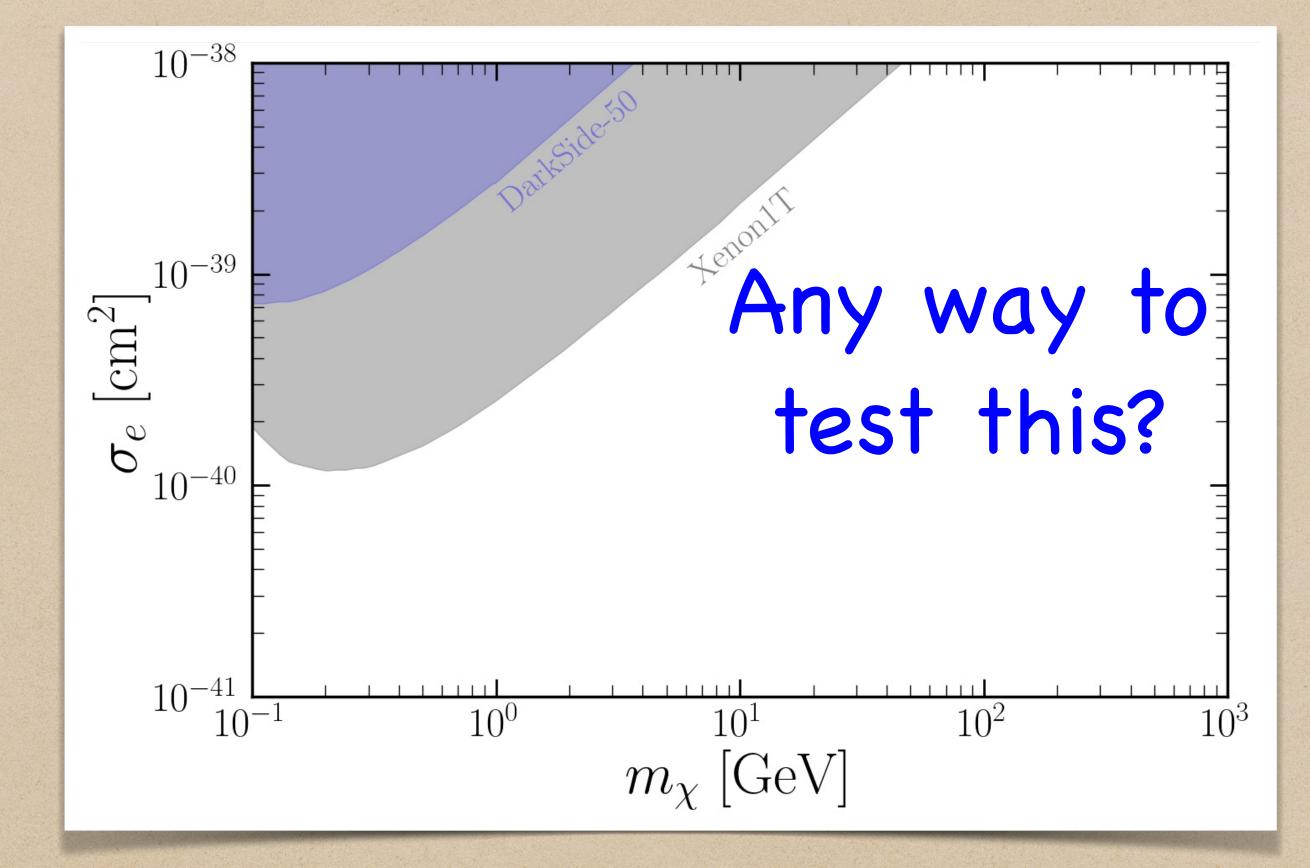


DM-electron scattering



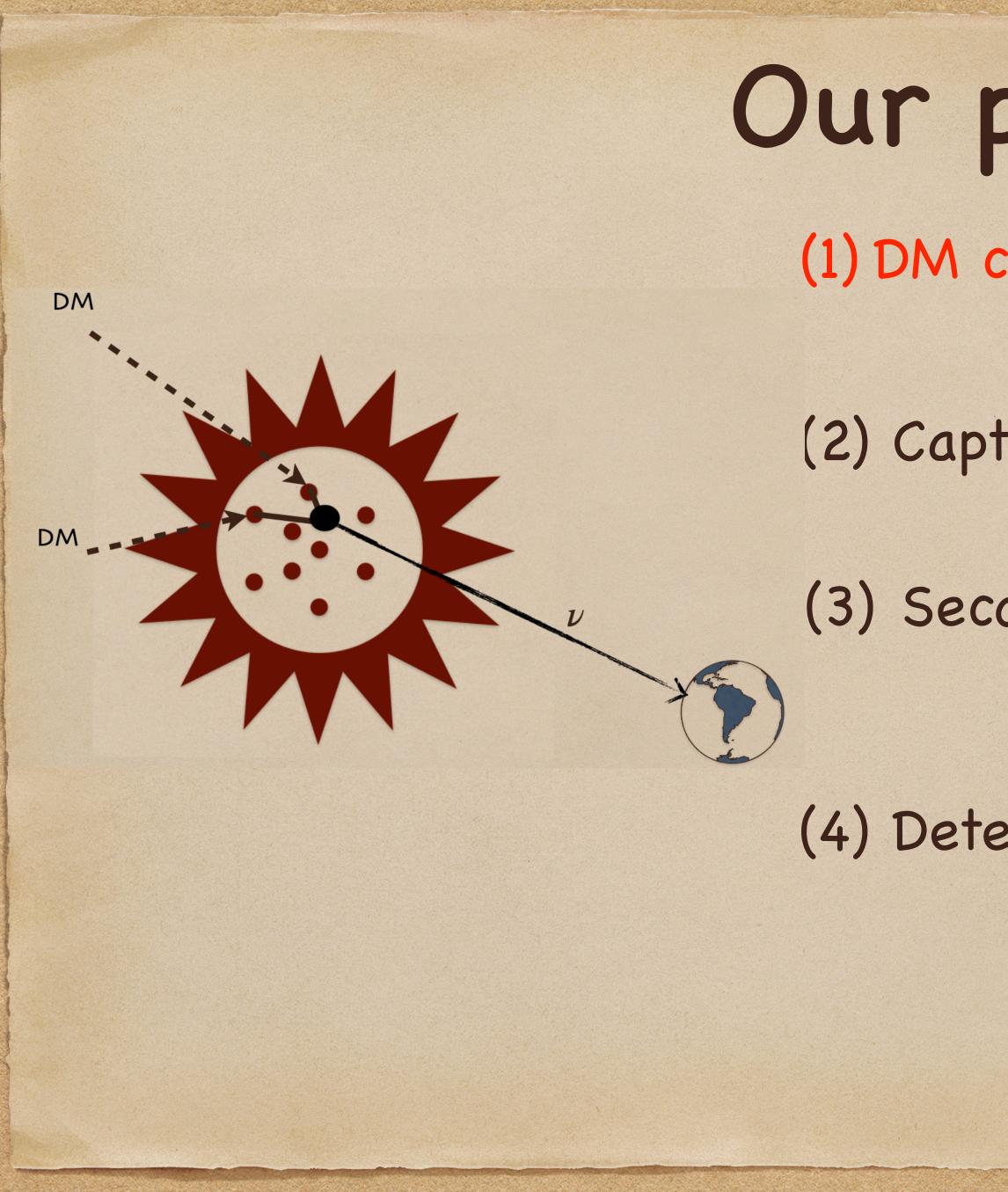


DM-electron scattering



Caveats: Model dependently loop level DM-nucleon interactions Kopp et al, In progress with D Bose, R Pramanik





Our proposal (1) DM capture in Sun through electron scattering (2) Captured DM annihilation in the Sun (3) Secondary Neutrinos produce from annihilation will reach earth (4) Detect these neutrinos in neutrino telescopes IceCube / DeepCore Super Kamiokande

5



DM particle velocity (u_{χ})

velocity at the surface

 $\sqrt{u_{\chi}^2 + v_{\rm esc}^2}$

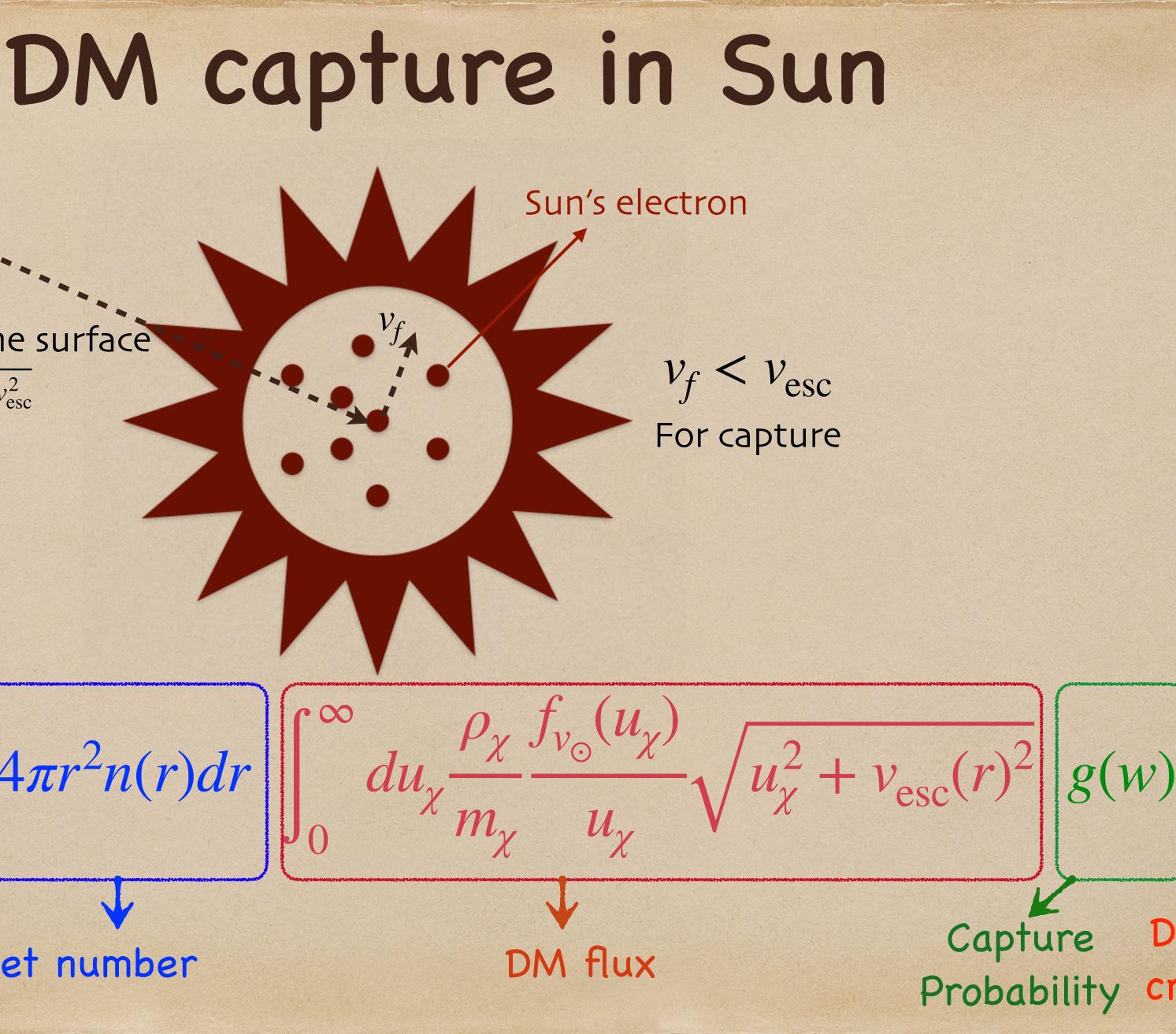
C =

Gould, ApJ, 1987 Kopp et al 0907.3159 Garani et al 1702.02768

Target number

 $4\pi r^2 n(r) dr$

V



DM-electron Probability cross section



Neutrinos from captured DM Propagation (nuSQuIDS)



DM

DM

within Sun Vacuum

Atmosphere



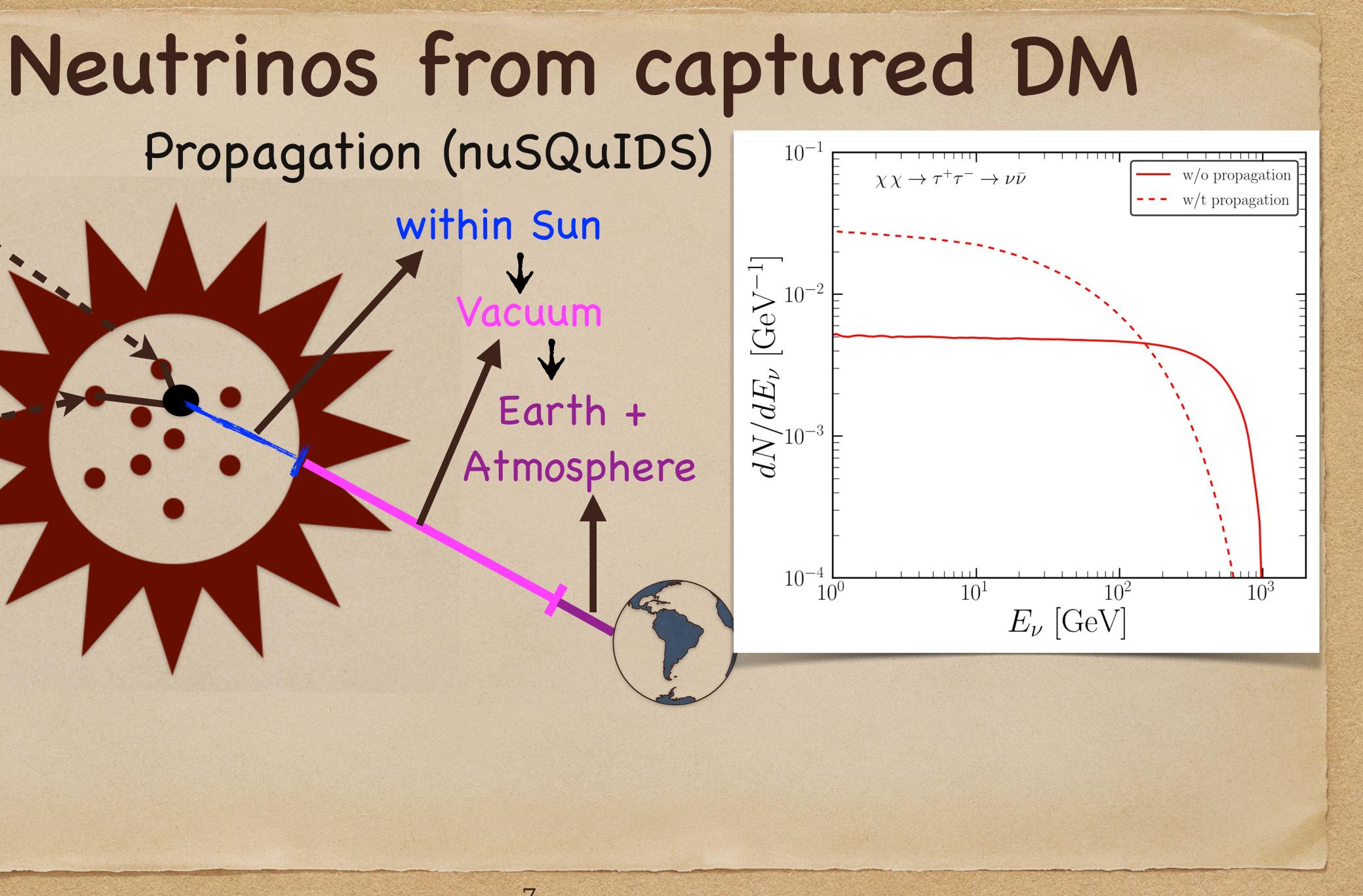


Propagation (nuSQuIDS)



DM

DM



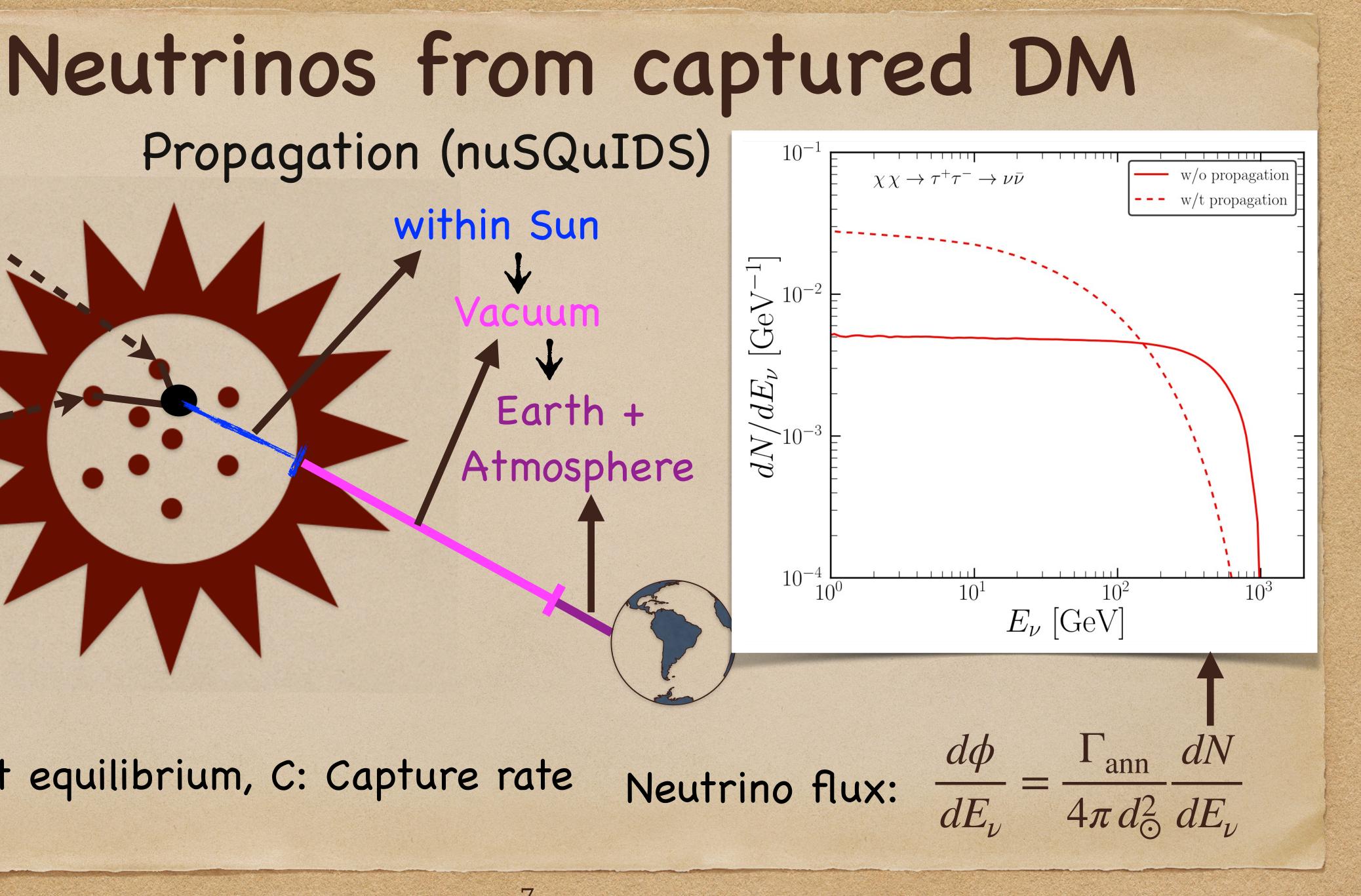
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DM

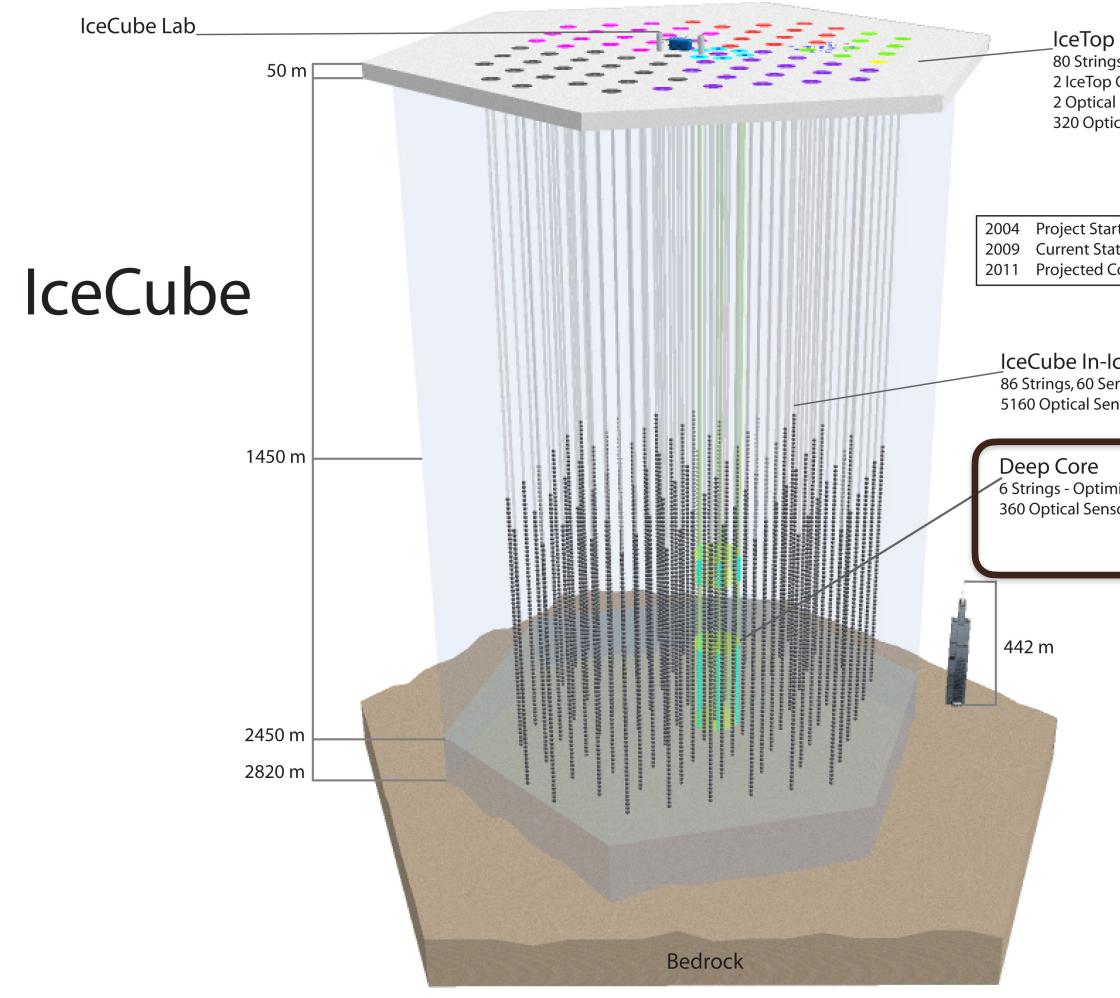
DM

 $\Gamma_{ann} = \frac{1}{2}$ At equilibrium, C: Capture rate



7

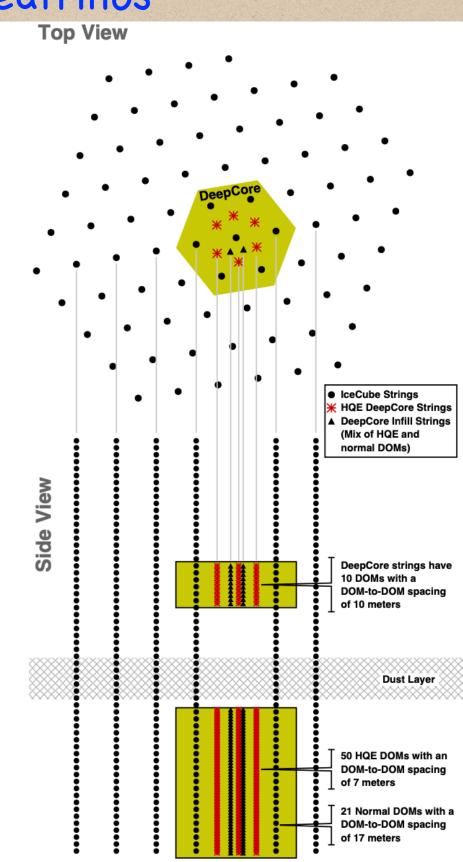
Detect neutrinos Requires BIG detector and patience; IceCube (DeepCore) is one such



At South Pole, 1 km³

Cherenkov radiation form by product

of neutrinos



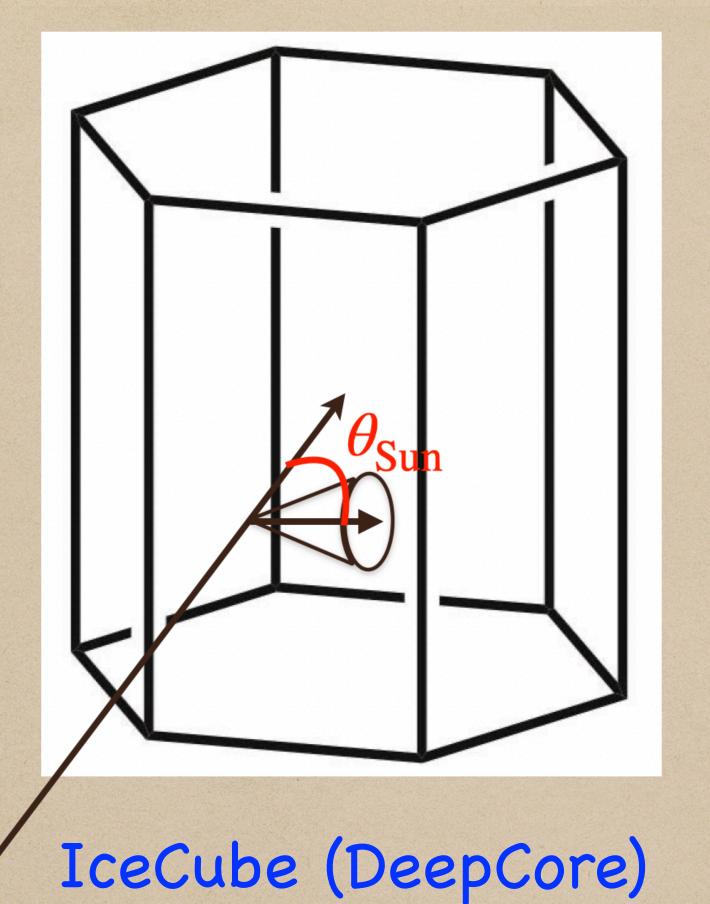
80 Strings each with 2 IceTop Cherenkov Detector Tanks 2 Optical Sensors per tank 320 Optical Sensors

roject Start	1 Hole
urrent Status	59 Holes
ojected Completion	86 Holes

IceCube In-Ice Arrav 86 Strings, 60 Sensors 5160 Optical Sensors

6 Strings - Optimized for low energies 360 Optical Sensors

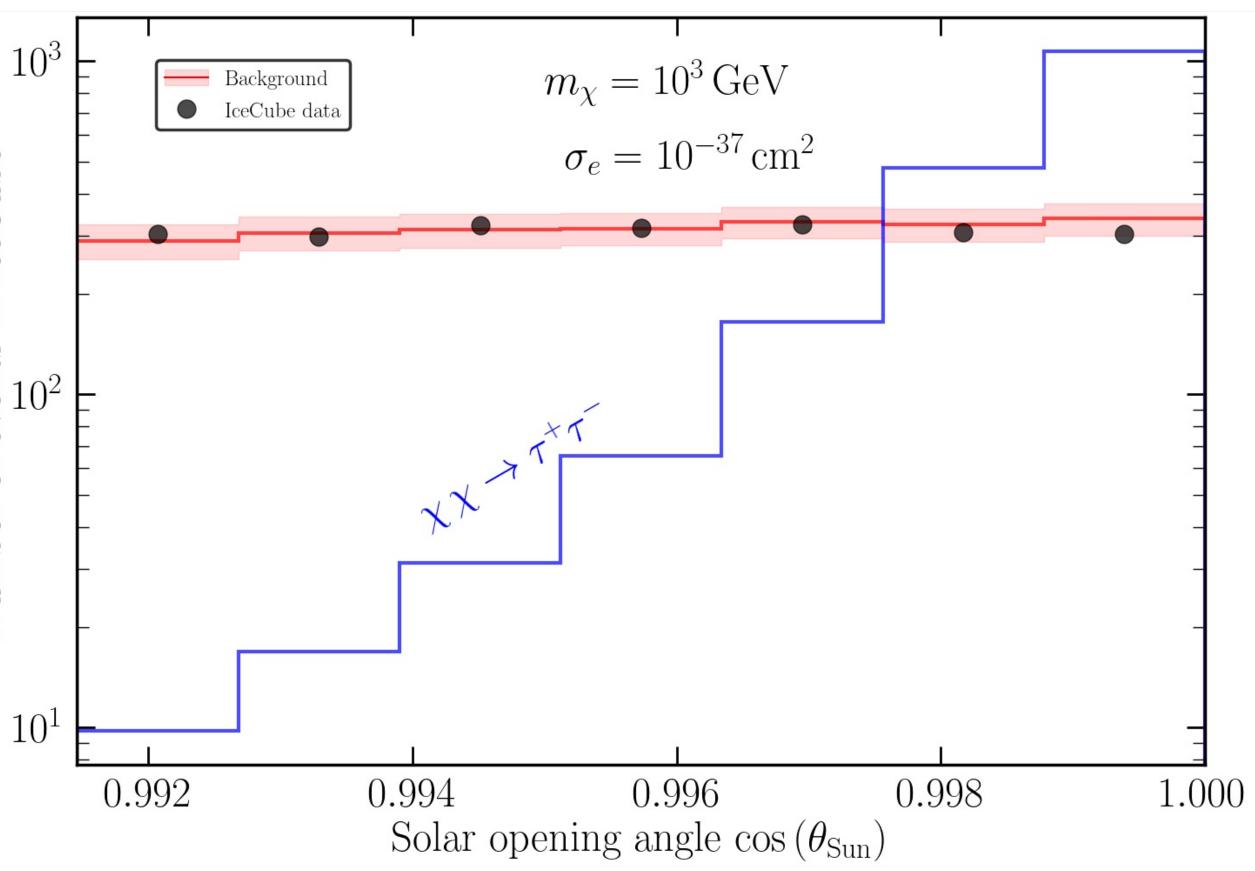






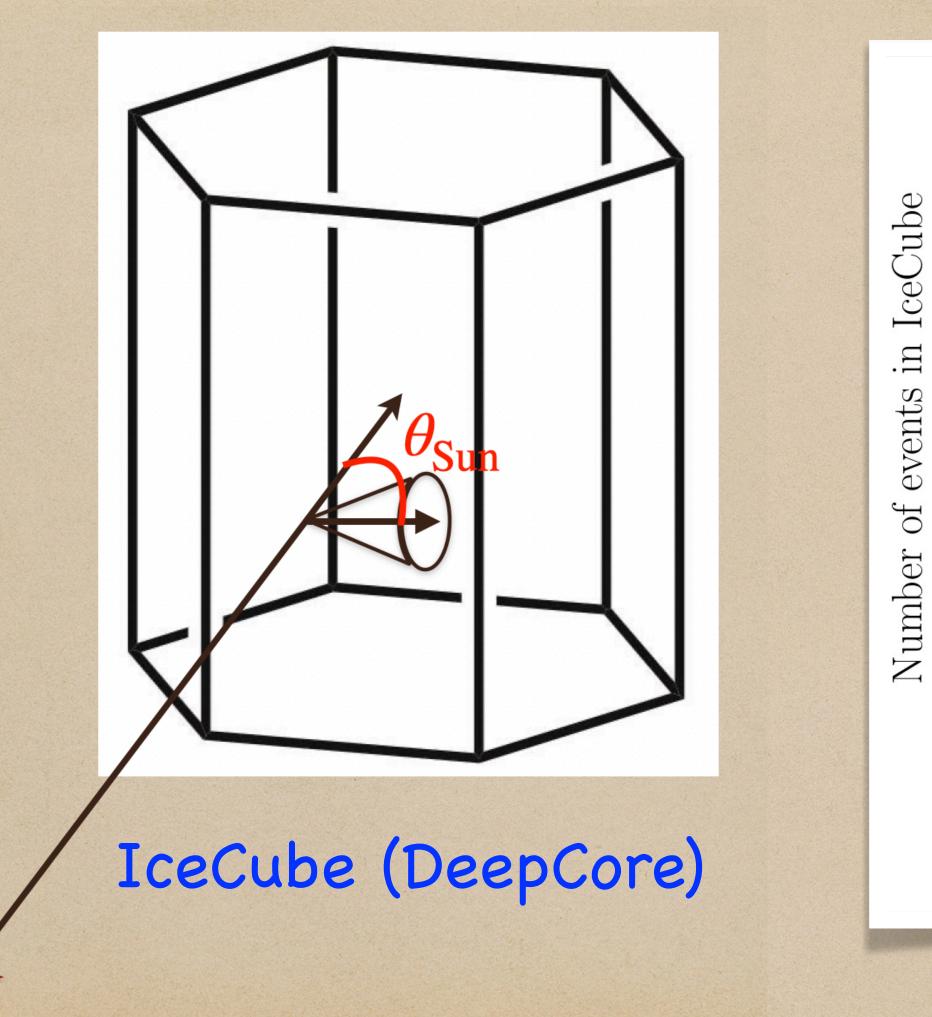
Number of events in IceCube

Look towards the Sun



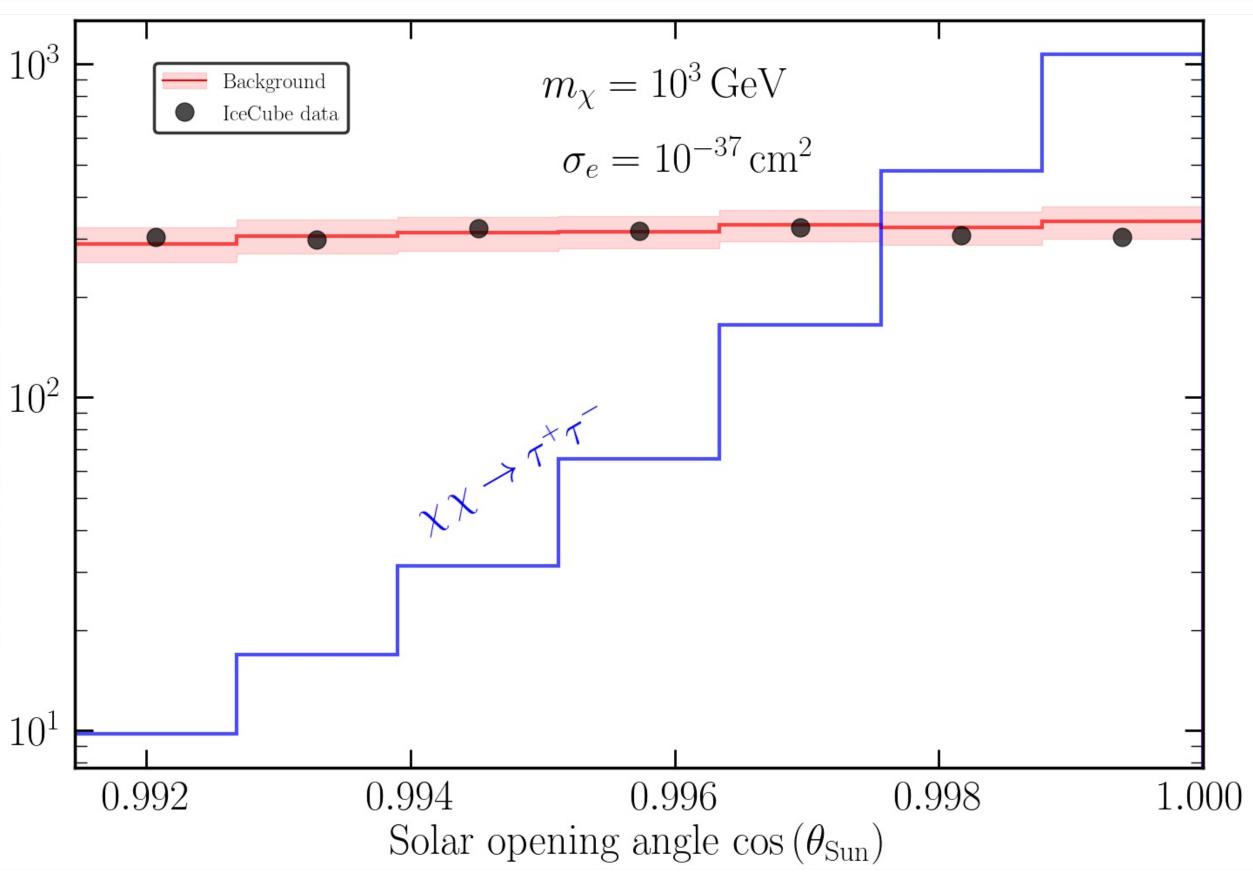
9





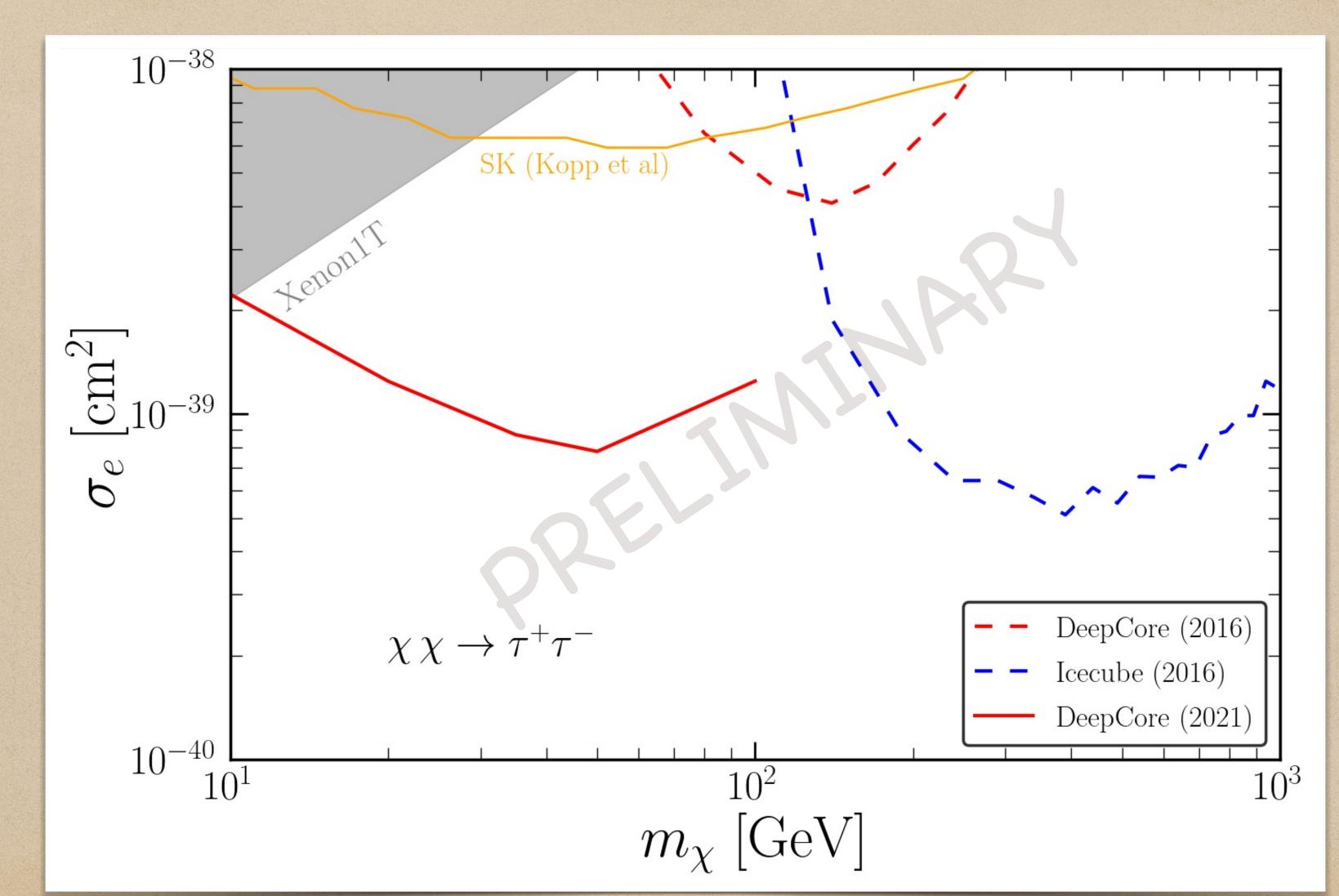
m

Look towards the Sun



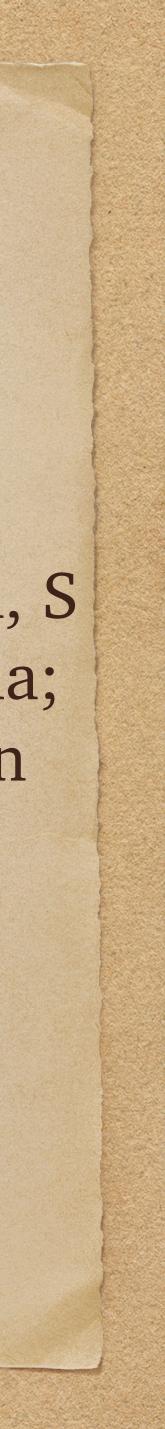
No excess found! --- Constraint on DM-e scattering cross section





Look towards the Sun: result

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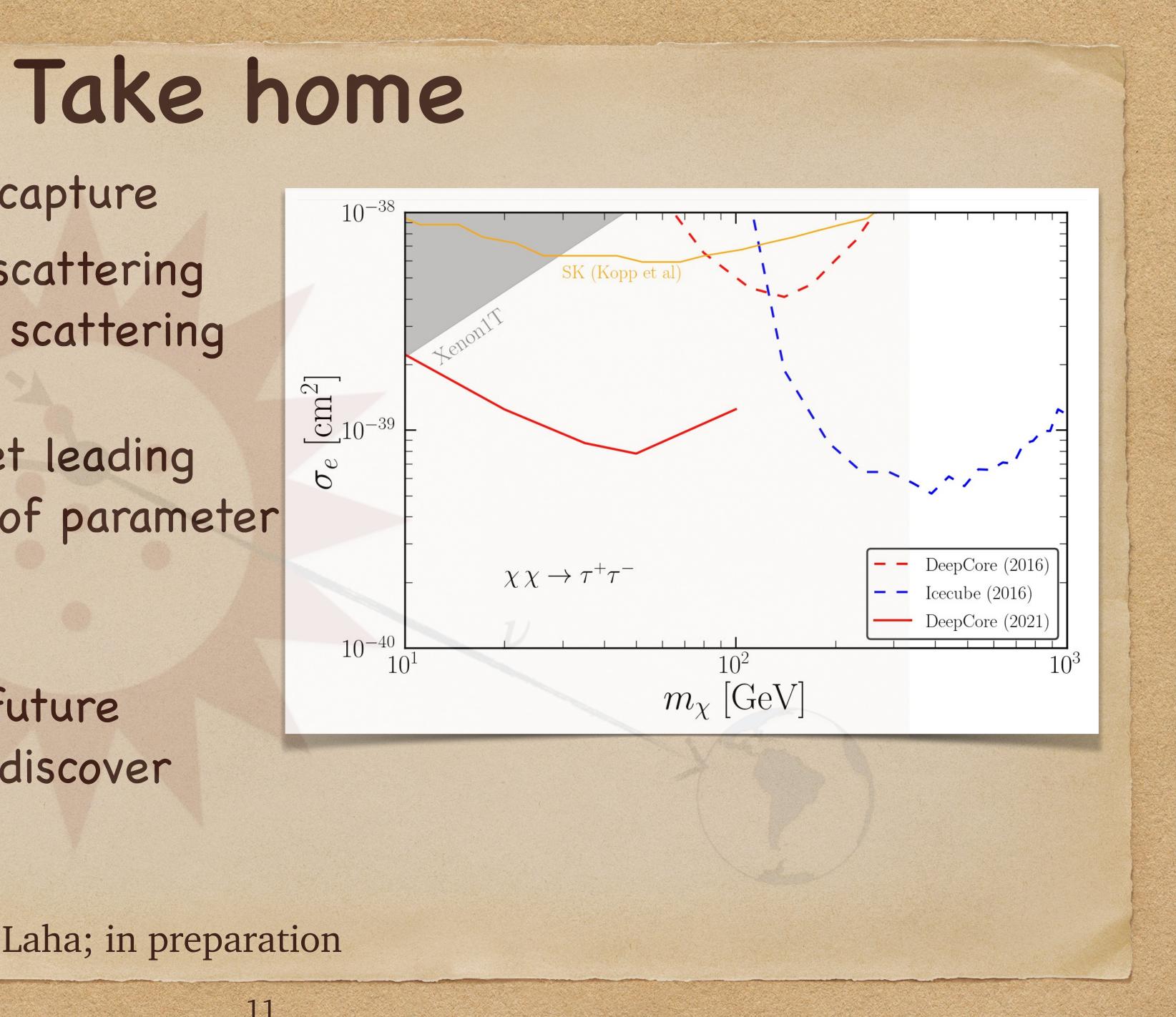


Neutrinos produce DM capture in Sun through electron scattering useful probe DM-electron scattering

IceCube/ DeepCore set leading 0 constraint in large region of parameter Space

Using this framework future O neutrinos telescopes may discover DM-electron scattering

TNM, A K Saha, S Mondal, R Laha; in preparation



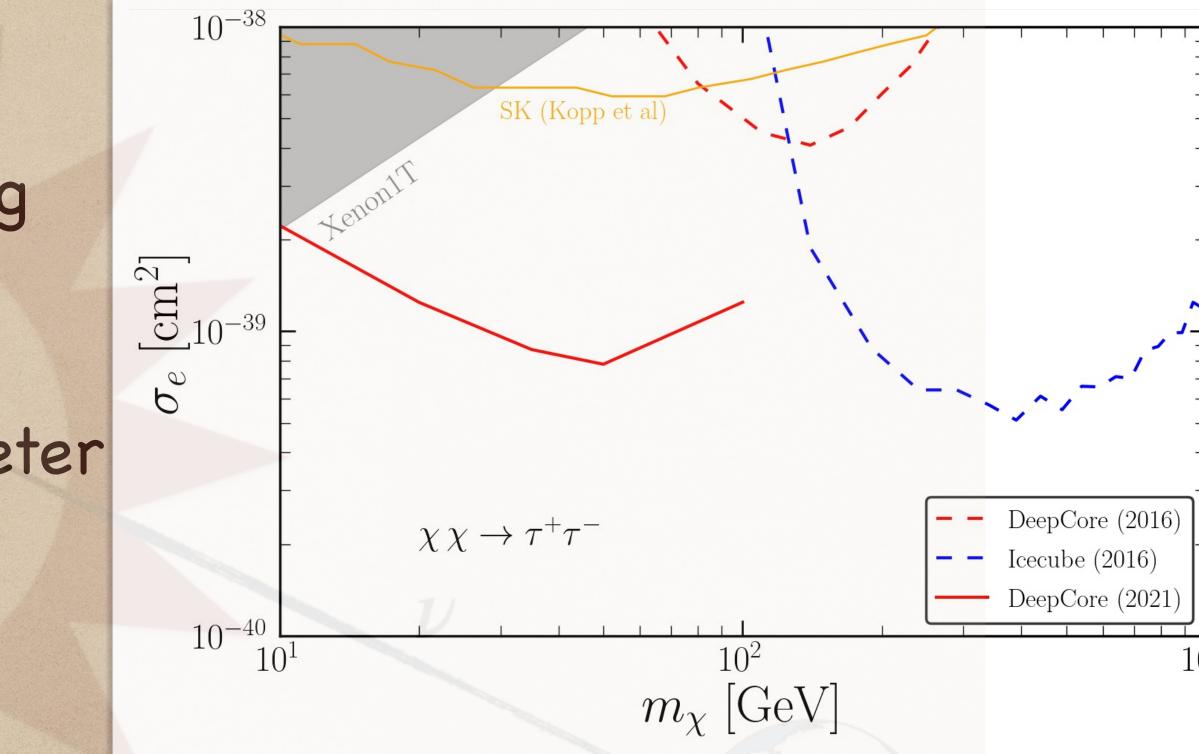
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Take home



email: tarak.maity.physics@gmail.com

Thank you

