SpaceQ – Direct Detection of Ultralight Dark Matter with Space Quantum Sensor

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- <u>https://arxiv.org/abs/2112.07674</u>, Nature Astronomy (2022)
- <u>https://arxiv.org/abs/2107.04038</u>, under review by NA
- <u>https://arxiv.org/abs/2210.03749</u>, under review by NA
- INSPIRE: <u>https://inspirehep.net/authors/1274923</u>

Image: Parker Solar Probe Credit: NASA/Johns Hopkins APL/Steve Gribben

Collaborations on Various Projects



• Quantum technologies in Space: <u>Q-SEnSE</u> + SpaceQ informal meeting

Big Questions

Can planetary data set meaningful constraints on
 Dark matter?
 General Relativity?
 5th forces?

• Can we use current or future **Space Quantum Technologies** to study fundamental physics?

Answers

- Can planetary data set meaningful dark matter constraints?
 General Relativity?
 5th forces?
 Yes! Many opportunities
- Can we use current or future space Quantum Technology to study fundamental physics? Yes! I will show you an example today.

Outline

- New Technologies & Ultralight Dark Matter
- Space Quantum Clocks & Sensitivity
- Model-Independent Constraints on Dark Matter & Cosmic Neutrinos: Rubin + Einstein!

Bridging Planetary Science, Space/Quantum Technologies, and Fundamental Physics

Many real-life applications & consequences!



Sun Devils / Anteaters - Starship

Why Space Quantum Clocks? Auto-Navigating Spacecraft & Space Travel





Artist's concept for Mars-ready habitat. Image Credit: SpaceWorks Torpor/NASA collaboration

Exploring the deep space: auto-driving Spacecraft; needs precision timing!!!

NASA Deep Space Atomic Clocks (current technology!) & Deep space and global navigation satellite system (GNSS)

Can we use the technology to study fundamental physics?

NASA DSAC & Parker Solar Probe



- Deep Space Atomic Clock loses one second every 10 million years, as proven in controlled tests on Earth.
- The clock has operated for more than 12 months in space; demonstrated long-term fractional frequency stability of 3 × 10⁻¹⁵

Burt, Prestage, Tjoelker, Enzer, Kuang, Murphy et al., Nature 595 (2021) 43.

• Exceeds previous space clock performance by up to an order of magnitude



(1.0 m × 3.0 m × 2.3 m)

Parker Solar Probe

Kasper, Klein, Lichko, Huang, Chen, Badman et al., Parker solar probe enters the magnetically dominated solar corona, Phys. Rev. Lett. (2021)

Why don't we put a quantum clock on a solar probe?
 What can we do with that?

Atomic Clock & Caesium Standard

- Atomic clocks: used to measure the distance between objects by timing how long it takes a signal to travel from A to B.
- For space exploration, clocks must be extremely precise:
- An error of even one second can mean the difference between landing on Mars or missing it by hundreds of thousands of miles.



http://hyperphysics.phy-astr.gsu.edu/hbase/acloc.html Reference: U.S. Naval Observatory, Cesium Clocks

Wave-Like (Fuzzy) Particles as Dark Matter

$$\lambda_{
m dB}\equiv rac{2\pi}{mv}$$

$$N_{
m dB} \sim \left(rac{34\,{
m eV}}{m}
ight)^4 \left(rac{250\,{
m km/s}}{v}
ight)^3 {
m in} \ \lambda_{
m dB}^3$$



Schive, Chiueh, Broadhurst, Nature Physics '14 arXiv:1406.6586,

- For ultralight m << 30 eV, the occupancy NdB is so large that the particles are best described by classical waves
- like electromagnetism, a state with a large number of photons is described by the classical EM fields.
- It would consist of extremely light scalar particles with masses go as low as 10⁻²² eV (rough lower bound):

de Broglie wavelength $\lambda \sim 1$ kpc: affect **structure formation.**

Oscillation of Wave-like Scalars

$$V(\phi)=rac{1}{2}m_{\phi}^2\phi^2+rac{1}{3}a_{\phi}\phi^3+rac{1}{4}\lambda_{\phi}\phi^4.$$
Dark matter potential

$$\phi(t, \vec{x}) = \phi_0 \cos(m_\phi t - \vec{k}_\phi \cdot \vec{x} + \dots).$$

(Non-relativistic solutions)

 $\omega \simeq m_{\phi}.$

Oscillation frequency ~ dark matter mass

Dark Matter Coupling



where e is the electron field, $F^{\mu\nu}$ ($G^{A\mu\nu}$) is the electromagnetic (QCD) field strength, g_s and β_3 are the strong interaction coupling constant and beta function (respectively), and $\kappa = \sqrt{4\pi}/M_P$ with $M_P = 1.2 \times 10^{19}$ GeV.

Atomic Physics Probe

$$\mathcal{L} \supset \kappa \phi \left(d_{m_e} m_e \bar{e} e + \frac{d_\alpha}{4} F_{\mu\nu} F^{\mu\nu} + \frac{d_g \beta_3}{2g_s} G^A_{\mu\nu} G^{A\mu\nu} \right),$$
(1)

$$\mu(\phi) \simeq \mu_0 \left(1 + d_{m_e} \kappa \phi \right), \quad \alpha(\phi) \simeq \alpha_0 \left(1 - d_\alpha \kappa \phi \right)$$
$$\alpha_s(\phi) \simeq \alpha_{s,0} \left(1 - \frac{2d_g \beta_3}{g_s} \kappa \phi \right), \quad (2)$$

where $\mu = m_e/m_p$ is the electron-proton mass ratio, and the subscript $_0$ denotes the central (time-independent) value of μ , α , and α_s .

Atomic Probe Basics

$$\mathcal{L} \supset \kappa \phi \left(d_{m_e} m_e \bar{e} e + \frac{d_\alpha}{4} F_{\mu\nu} F^{\mu\nu} + \frac{d_g \beta_3}{2g_s} G^A_{\mu\nu} G^{A\mu\nu} \right),$$
(1)

Turning off d_{m_e} and d_g for demonstrations,

 $f_A \propto \alpha^{\xi_A+2}$, f is the frequency of a (clock) transition.

$$\alpha = \alpha_0 (1 + d_\alpha \kappa \phi(t)).$$

$$\frac{\delta(f_A/f_B)}{f_A/f_B} \simeq (\xi_A - \xi_B) d_{\alpha} \kappa \phi(t).$$

- Experimental observable! See arXiv:1405.2925, Arvanitaki, Huang, Tilburg, PRD 15
- For example, if A is a hyperfine microwave transition and B is an electronic optical transition, ζA = 1 and ζB = 0.
- Clock (~ 10⁻¹⁵ for DSAC) stability translate to how well we can measure $\frac{\delta(f_A/f_B)}{f_A/f_B}$

Solar Bound-State Halo An example of DM overdensity

Yu-Dai Tsai, UC Irvine, '22 yudait1@uci.edu

Scalar DM Halo

Stable solution can be supported by external potential

$$V_{\text{ext}} = \begin{cases} -\frac{G \, m_{\phi} \, M_{\text{ext}}}{r} & \text{for } R_{\star} > R_{\text{ext}} ,\\ -\frac{3 \, G \, m_{\phi} \, M_{\text{ext}}}{2 \, R_{\text{ext}}} \left[1 - \frac{1}{3} \left(\frac{r}{R_{\text{ext}}} \right)^2 \right] & \text{for } R_{\star} \le R_{\text{ext}} , \end{cases}$$

$$ho(r)\simeq
ho_\star\exp\left(-2r/R_\star
ight)$$
, for $R_\star>R_{
m ext}$

 $R_{\star} \simeq \frac{M_P^2}{M_{\text{ext}} m_{\phi}^2},$ where $M_{\text{ext}} = M_{\odot}$ is the mass of the external host body; note that R_{\star} is independent of the total mass in the halo $v_{\star} = (m_{\phi} R_{\star})^{-1},$

Banerjee, Budker, Eby, Flambaum, Kim, Matsedonskyi, and Perez, 1912.04295

Dark matter in solar system? Planetary constraint!



• Black data points are model-independent constraints!

• Dark matter induce precessions to the planets Mercury, Venus, Earth, Mars, Jupiter, Saturn Pitjev, Pitjeva, 1306.5534, Astronomy Letters '13 Tsai, Eby, Safronova, 2112.07674

Enhancement of the DM Density



PSP: Parker Solar Probe Tsai, Eby, Safronova, arXiv:2112.07674

Results





- Motivate Specific Frequency Region!
- Motivate Nuclear Clocks!
- **Tsai**, Eby, Safronova, arXv:2112.07674, Nature Astronomy (2022)

$$\mathcal{L} \supset \kappa \phi \left(d_{m_e} m_e \bar{e} e + \frac{u_\alpha}{4} F_{\mu\nu} F^{\mu\nu} + \frac{u_g \rho_3}{2g_s} G^A_{\mu\nu} G^{A\mu\nu} \right),$$
(1)

$$rac{g_e^2\Lambda^2}{(4\pi)^2} \lesssim m_\phi^2, \ \Lambda = 4\pi v_{EW} \simeq 3 \ {
m TeV}.$$

Naturalness condition

Spatial Variation of Fundamental Constants

$$k_X \equiv c^2 \frac{\delta X}{X \, \delta U}$$
. $X = \alpha, \mu, \text{ or } m_q / \Lambda_{QCD}$.

 δU : change in gravitational potential .

$$\delta U/c^2\simeq 3.3 imes 10^{-10},~$$
 Earth variation.

 $\delta U/c^2 \sim 9 imes 10^{-8},~$ from Earth to Solar probe at 0.1 AU.

• Achieve constraints on k_X that are a factor of ~ 300 stronger!

New Directions at Precision Frontier

 Planetary Defense & Dark Matter: Model-Independent Probes of Any Dark Matter Candidates (especially purely gravitational dark matter)



Vera Rubin Carnegie Institution for Science





Albert Einstein Mount Wilson Observatory, California

Asteroids hitting the Earth

Fireballs Reported by US Government Sensors (1988-Apr-15 to 2021-Jul-30)





~ 65 million years ago

Engin Akyurt – Unsplash - free usage for commercial & non-commercial purposes

Tracking asteroids is extremely important e.g., unexpected 2013 Chelyabinsk meteor injured >1500 people Also, near-Earth asteroid search accidentally found 'Oumuamua

New Project: New Model Independent Constraints!



Tsai, Farnocchia, Eby, Arakawa, Safronova, <u>arXiv:2210.03749</u> Obtained results from NASA JPL Sentry-II code & OSIRIS-Rex+ data Can be applied to ideas like Solar-Basin & Axion Mini-cluster (more refs in paper)

Astrometric Probe of the DM Gravity (and Cosmic Neutrinos!)



Highly exaggerated eccentricity

Collaborate with NASA expert, Davide Farnocchia, to derive the bounds



 New bounds on local DM profile from gravity (applicable to all models, but most meaningful for solar bound halo, axion mini-cluster, solar basin, etc.)

To reach usual galactic DM density $\overline{\rho}_{DM} = 0.3 \text{ GeV}/cm^3$

- 1. Quantum sensor (e.g., NASA DSAC) can improve the result
- 2. Can consider **DM-SM long-range interaction**, only 4-order stronger than gravity



Probing Fundamental Physics with Solar-System Study & Quantum Sensors



 Space quantum sensor to study solar bound dark matter, Tsai, Eby, Marianna, <u>arXiv:2112.07674</u> to appear on Nature Astronomy



Millions of objects to study LSST can increase the observed objects by a factor of 5-10.

- GR/Modified Gravity (inc. Dark Energy models)
- Gravitational Waves
- Pure Gravitational Constraints on Dark Matter & Cosmic Neutrino Profiles Tsai et al., <u>arXiv:2210.03749</u>
- My outreach interview on this topic gathered
 > 77 K views worldwide! Click link below: https://www.youtube.com/watch?v=xDX9XwLHBuM
- Fifth Forces,
 Tsai, Wu, Vagnozzi, Visinelli,
 arXiv:2107.04038
- Primordial Black Holes
- Topological Defects

Let's protect the Earth & find dark matter; happy to discuss more

Thank you!

Thank Josh, Marianna, Luca, Sunny, Youjia for comments Outreach / interview:

<u>https://www.youtube.com/watch?v=xDX9XwLHBuM</u> Yu-Dai Tsai, UC Irvine, <u>yt444@cornell.edu</u> / <u>yudait1@uci.edu</u>

I also work on ...

- Fixed-target searches for dark matter & long-lived particles (FerMINI & LongQuest) with Pospelov et al.
- LHC Forward Experiments: Forward Physics Facility, FORMOSA (a millicharge experiment I proposed), with Feng et al.
- Dark matter model building (dark sector QCD, Strongly Self-Interacting Dark Matter, SIMP/ELDER), with Murayama, Slatyer, Perelstein et al.
- Dark matter searches using neutron star / compact merger / multimessenger astronomy, with Profumo, Sathyaprakash et al.
- Neutrino physics (cosmic neutrino background) & neutrino BSM, with Shoemaker et al.
- My works on <u>Inspire HEP</u> & <u>my outreach interview</u> (> 76K views!)

More References

- Seto, Cooray, arXiv:0405216, PRD 04
- LLR Experiments: Williams, Turyshev, Boggs, PRL 04 Murphy, Rept. Prog. Phys 13
- Atomic / nuclear clocks for fundamental physics: Peik, Schumm, Safronova, Pálffy, Weitenberg, Thirolf, 2012.09304
- GW background, Fedderke, Graham, Rajendran, PRD21
- Quantum Technologies in Space, Kaltenbaek, Exp Astron 21