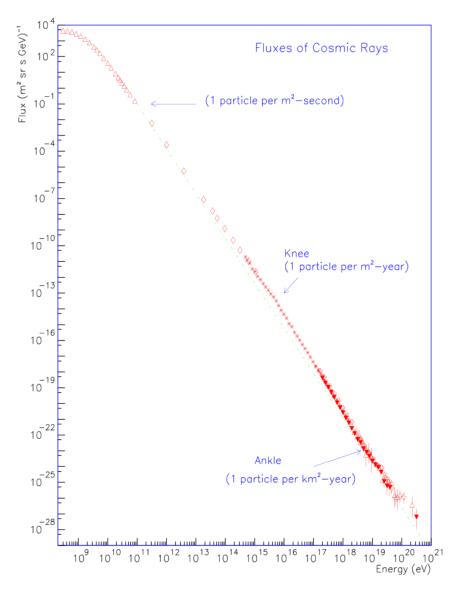


# The Pierre Auger Observatory: Cosmic Rays at the Highest Energies

Jeff Allen (New York University)

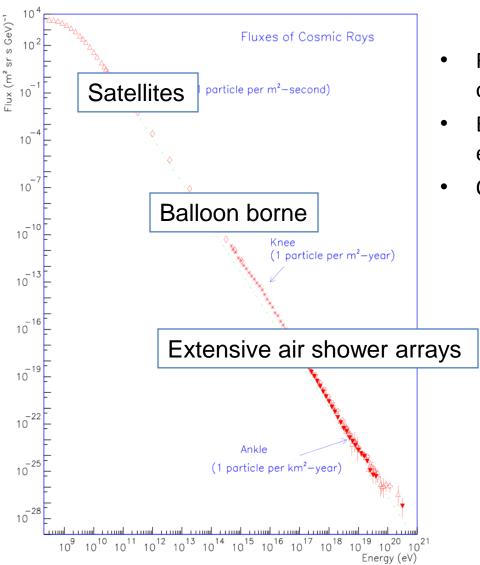
For the Pierre Auger Collaboration

### **Cosmic Rays**



- Fairly constant power law spectrum over 10 decades in energy
- Extremely low flux at the highest energies, ~1 per km2 per century
- Open questions
  - Galactic to extra-galactic transition
  - Sources
  - Composition

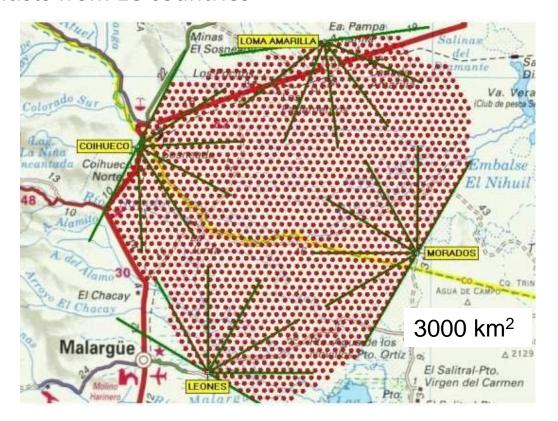
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### The Pierre Auger Observatory

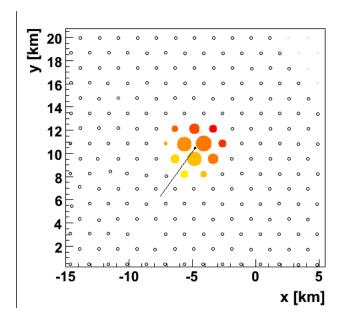
- Largest detector in operation
- 490 scientists from 18 countries



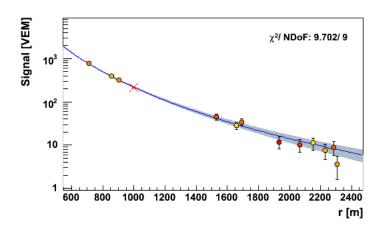
Argentina, Australia, Bolivia, Brazil, Croatia, the Czech Republic, France, Germany, Italy, Mexico, Netherlands, Poland, Portugal, Slovenia, Spain, the United Kingdom, the United States, and Vietnam.

## Surface Array (SD)

- High energy air showers leave large footprint on the ground
- 1600 water Cherenkov tanks with 1.5 km spacing
  - Each detector self calibrates and triggers
- Fully efficient around 3 EeV
- Angular reconstruction from GPS timing
  - < 1° at high energies</p>

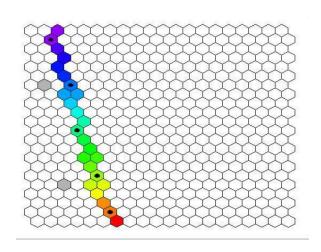


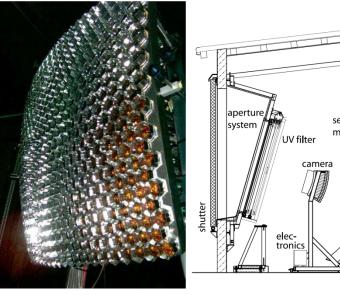


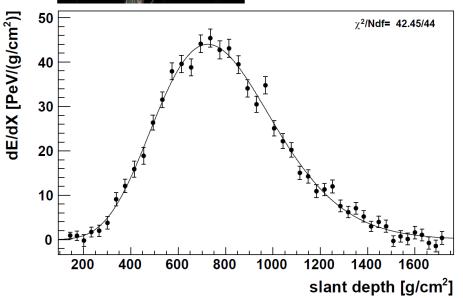


### Fluorescence Telescopes (FD)

- Charged particles ionize the atmosphere
- Resulting fluorescence light viewed by 24 telescopes at four sites
- Provides nearly calorimetric energy reconstruction

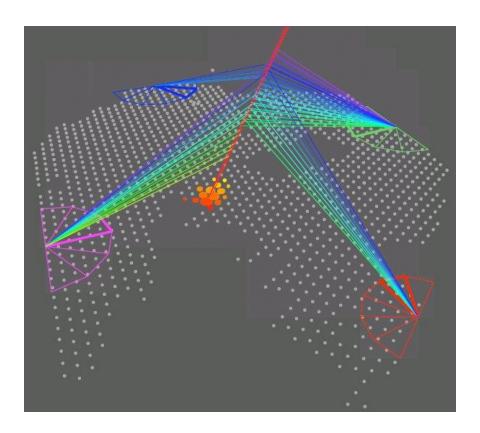






# **Hybrid Detector**

- Hybrid technique makes Auger truly unique
- Air showers detected with both techniques are some of the best constrained showers detected

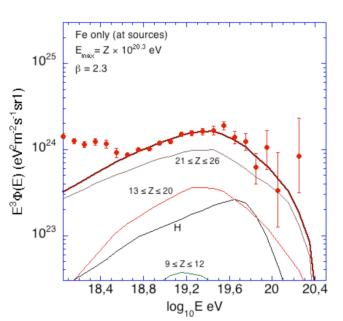


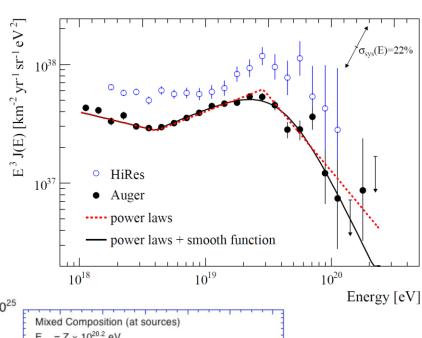
#### Science Goals

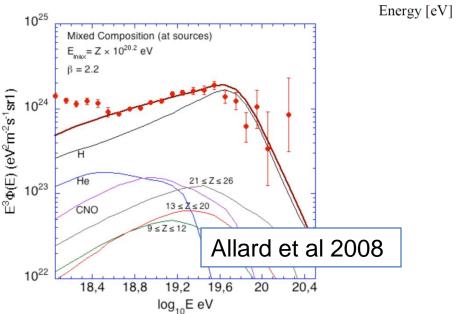
- Measure the flux of UHECR (E > 1EeV)
- Determine sources and nature of acceleration
  - Correlation seen in the arrival directions with AGN within 70 Mpc
- Study composition of UHECR
  - Proton, iron, intermediate?
- Look for air showers generated by all types of particles
  - Photons
  - Neutrinos
  - Exotics?

### **Energy Spectrum**

- Ankle seen at 4 EeV
  - Possible location of galactic to extra-galactic transition
- High energy cutoff seen above 30 EeV
  - Spectral index drops to -4.3
  - GZK? Photodisintegration?
- Current energy spectrum can be fit by iron or proton

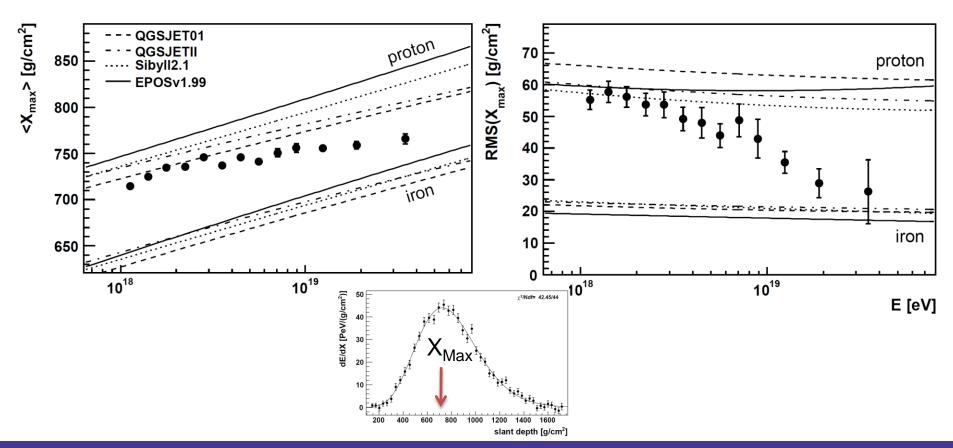






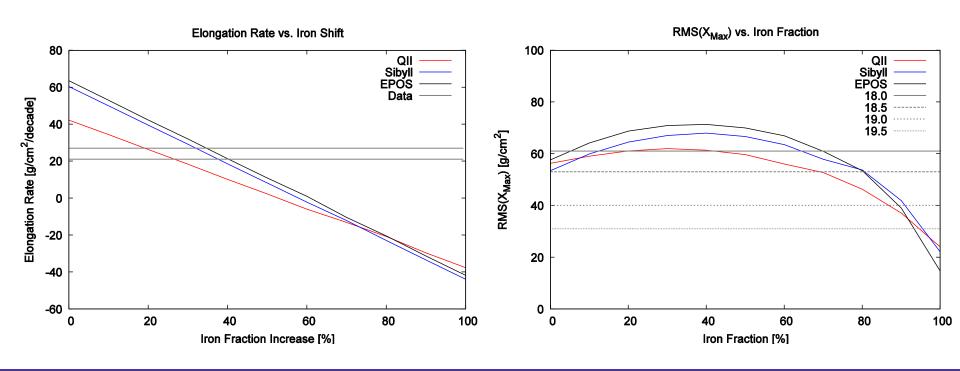
# **X**<sub>Max</sub>: Composition

- Depth of shower maximum  $(X_{Max})$  sensitive to the composition
  - Both mean and variance are useful tools in determining the mass and evolution
- Current results point to increasingly heavy mass, from MC simulations



# **X**<sub>Max</sub>: Comparison to Models

- Quantitative interpretation is possible with the use of MC simulations and models of the high-energy interactions
  - Rely on extrapolations from accelerator data (Take with grain of salt)
- Taking MC predictions at face value, results consistent within measurement and theoretical uncertainty

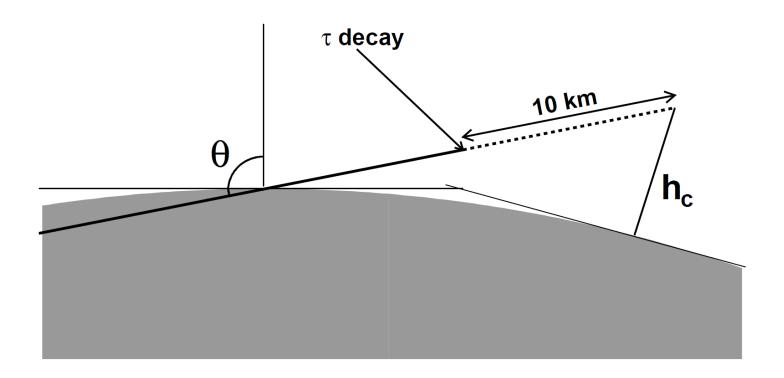


### **Beyond Cosmic Rays**

- Auger detects air showers.
  - Can be initiated by photons, neutrinos, or even exotic particles!

#### Neutrinos

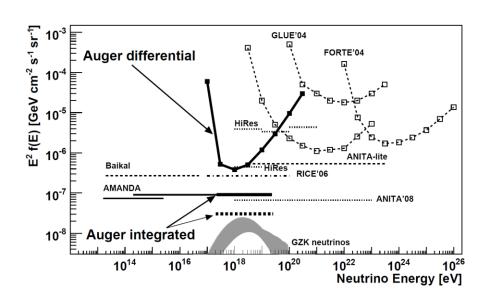
- Large target mass provided by the earth
  - "Earth Skimming" tau-neutrinos can interact in near the surface of the earth
  - Decaying tau initiates an EM rich shower at nearly horizontal showers

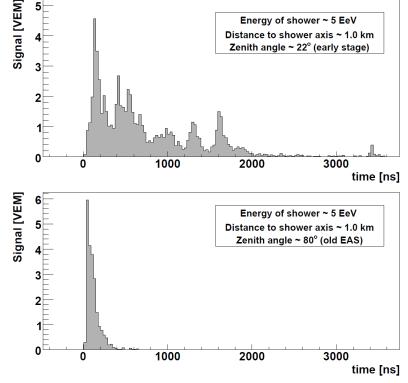


#### **Neutrinos**

- Decaying tau initiates an EM rich shower at nearly horizontal angles
  - EM component can be detected through the FADC trace of surface detectors
- Horizontal showers will produce elongated footprints
  - Time difference between neighboring tanks ~light travel time

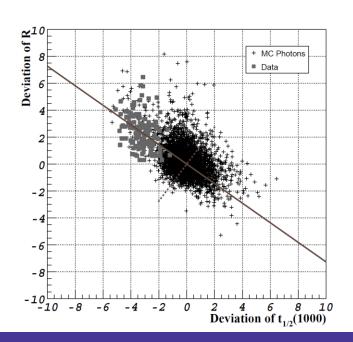
No candidates observed, but limit is competitive in the field

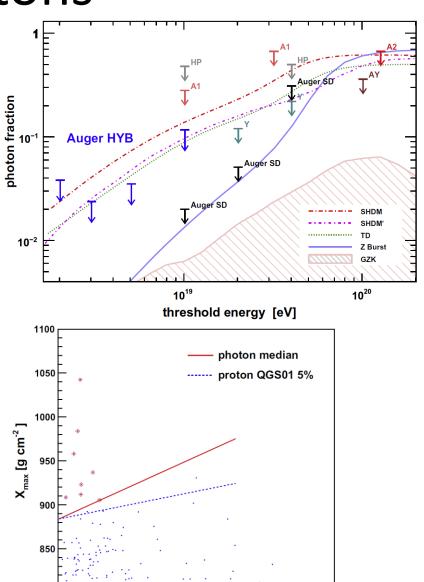




#### **Photons**

- Photons can be seen in SD as muon poor showers
- Can be seen in FD as showers with deep X<sub>Max</sub>
- Separation not perfect in either case
  - Candidates exist
  - Limits the photon fraction
- Current limit rules out some exotic models of UHECR creation





Ig(E/eV)

19.6

#### Conclusion

- Construction of Auger completed in 2008, but work not done
  - AMIGA and HEAT are low energy extensions, take Auger down to the second knee
  - Radio R&D could provide a FD type detector with ~100% duty cycle
- Composition of cosmic rays
  - Taking models at face value, suggests mixed composition around 1 EeV transitioning to heavier elements
- Large air shower observatories are useful beyond cosmic rays
  - Neutrino limits competitive with dedicated neutrino experiments
  - Photon limits can rule out several "exotic" models of UHECR creation