

Vanishing soft
masses under
 $E_7/SO(10)$

S. K. Mandal

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Vanishing SUSY soft masses under $E_7/SO(10)$ coset models

Work in preparation w/ M. Nojiri, T. T. Yanagida

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PHENO Symposium, UW-Madison
May 9, 2011

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Hierarchy of Yukawa couplings?

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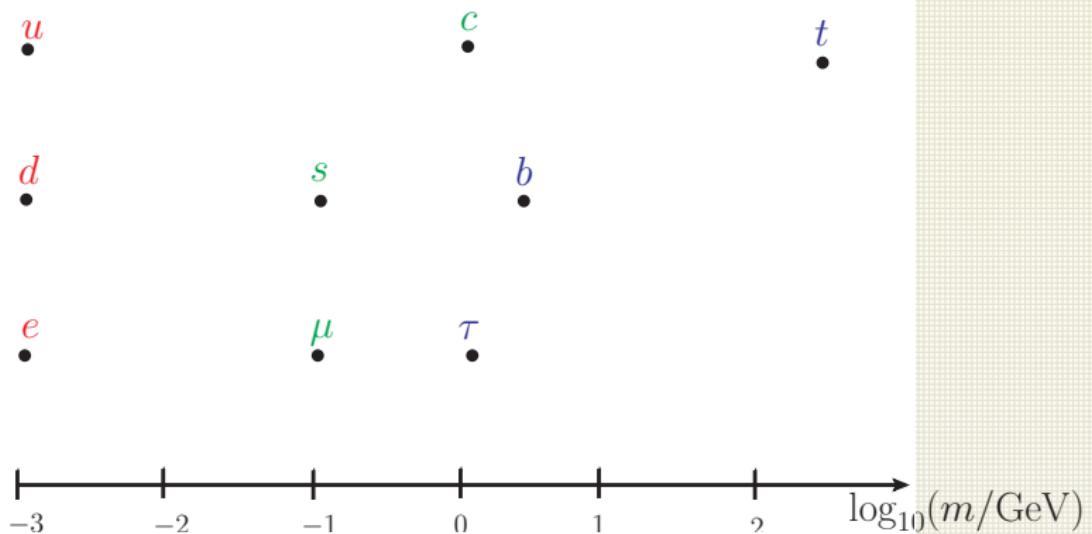
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Ad-hoc \Rightarrow dynamical origin of mass?

Field content of standard model?

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Gauged adjoint and chiral fundamental

- Gauge some group $H \subset G$ in real representation
- Where does chiral field in fundamental of H come from?

Hint from pions

- Take $SU(2)/U(1)$ toy model
- Broken X and $Y \Rightarrow Z = X + iY, Z^* = X - iY$
- $SU(2) \sim S^3 \implies S^3/U(1) \sim \mathbb{CP}^1$
- Gauge $U(1)$, toy pions in fundamental complex representation
- Need fermionic analogue

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Other problems ...

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Problems of the Standard Model

- Higgs hierarchy problem
- Dark matter
- Gauge unification
- ⋮

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Problems of supersymmetry

- μ problem $\Rightarrow m_H \rightarrow 0$?
- Flavor/CP problem $\Rightarrow m_{0(3)} \gg m_{0(1,2)}$?
- ⋮

Natural way to accomplish this? Follow hint ...

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Strategy

Coset spaces

- $M_P, M_{GUT}, M_\Lambda \gg M_{SUSY}$
- Higgs bosons = Pseudo-Nambu-Goldstone (NG) bosons at M_Λ

Pseudo-Nambu-Goldstone fermions

- 1st and 2nd generation Standard Model fermions = SUSY partners of NG bosons
- Leave 3rd generation as matter fields $\sim M_{SUSY}$

⇒ Need SUSY nonlinear σ model!

How to construct one?

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Formulation of SUSY nonlinear σ model

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Nonlinear σ model

Under global symmetry breaking $G \rightarrow H$,

$$\mathcal{L}_{NGB}^{NL} = -(\partial^\mu \pi_i) g^{ij} (\pi/f_\pi) (\partial_\mu \pi_j)$$

$$\mathcal{L}_{NGF}^{NL} = \frac{c_{ijkl}}{f_\pi^2} (\bar{\psi}_i \gamma^\mu \psi_j) (\bar{\psi}_k \gamma_\mu \psi_l)$$

The Kähler potential

Supersymmetric generalization of nonlinear σ model:

$$\mathcal{L}_{SUSY}^{NL} = K(\Phi, \Phi^\dagger) \Big|_{\theta\bar{\theta}\bar{\theta}\bar{\theta}}$$

Broken d.o.f.'s parameterize Kähler manifold

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Technical points

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Invariance of K

- Kähler transformation:

$$K(\Phi, \Phi^\dagger) \rightarrow K(\Phi, \Phi^\dagger) + F(\Phi) + F^\dagger(\Phi^\dagger)$$

- Under supergravity, W can break G -invariance on Kähler manifold

Solution criteria

- G/H must be non-compact [Bagger and Witten, Phys. Lett. B **118**, 103 (1982)]
- H must have no $U(1)$ factors [Kugo and Yanagida, Prog. Theor. Phys. **124**, 555 (2010)]

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Technical points (cont'd)

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- General form of Kähler potential:

$$K = v^2 F(\det[\Psi_\alpha^\dagger \Psi^\alpha])$$

- v is scale of G breaking, $F(x)$ freedom from broken $U(1)$'s
- G nonlinearily realized by Ψ^α
- For a simple example manifold, see:

SKM, Nojiri, Sudano, Yanagida, JHEP **01**, 131 (2011)

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Supergravity and SUSY breaking

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- Couple to supergravity:

$$\Delta\mathcal{L} = \frac{3}{4} \int d^2\theta d^2\bar{\theta} \mathcal{E} (\bar{D}\bar{D} - 8R) \exp\left(\frac{1}{3}K(\Psi, \Psi^\dagger)\right)$$

- Introduce hidden SUSY breaking sector:

$$K = K(\Psi, \Psi^\dagger) + Z^\dagger Z + \dots$$

- NG bosons remain massless, but QNGB d.o.f.'s from broken $U(1)$'s inside Ψ get mass \Rightarrow “novino” superfields

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Choosing $G = E_7$

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Formal motivations for E_7

- Appears in $N = 8$ supergravity
- Nonlinear realization gives the “simplest field theory”
[Arkani-Hamed, Cachazo, Kaplan arXiv:0808.1446]

GUT representations

- $SU(5) \supset (5 + 5^*)_H \oplus 10 \oplus 5^*$
- $SO(10) \supset 10_H \oplus 16 \rightarrow (5 + 5^*)_H \oplus 10 \oplus 5^* \oplus 1$
- E_7 gives right flavor cosets:
 - $E_7/[SU(5) \times U(1)^3] \rightarrow (10 \oplus 5^*) \times 3 + 5_H$
 - $E_7/[SO(10) \times U(1)^2] \rightarrow (10 \oplus 5^*) \times 2 + (5 + 5^*)_H$
- E_6 too small, E_8 gives mirror families

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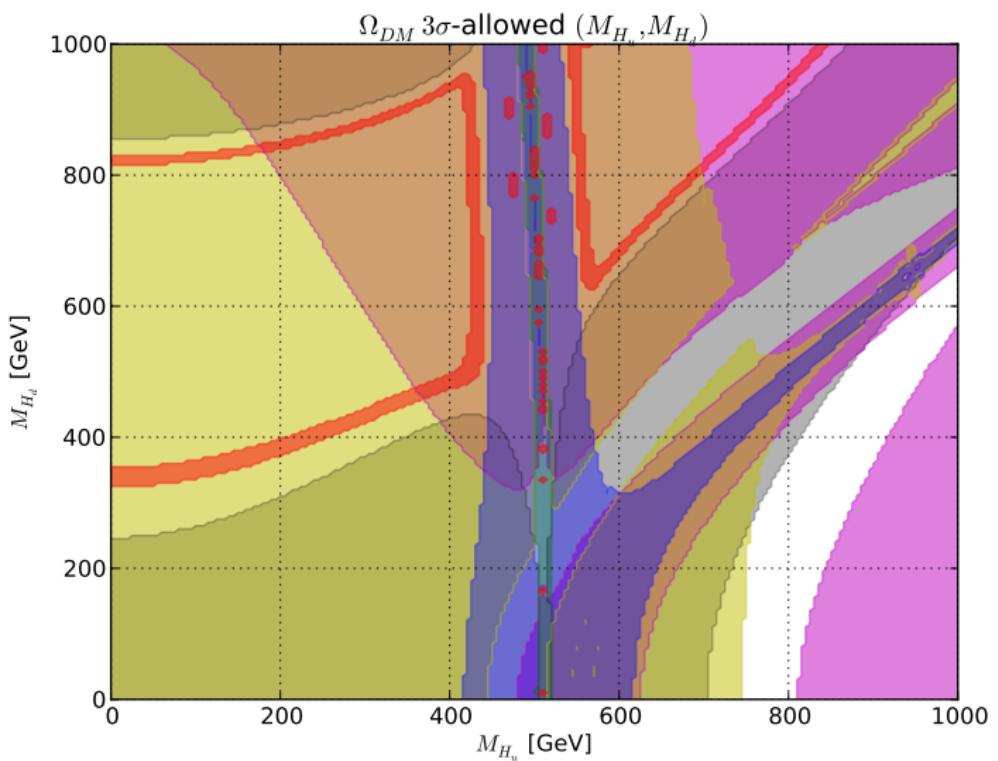
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Light 5*, Light 10 (3rd generation)

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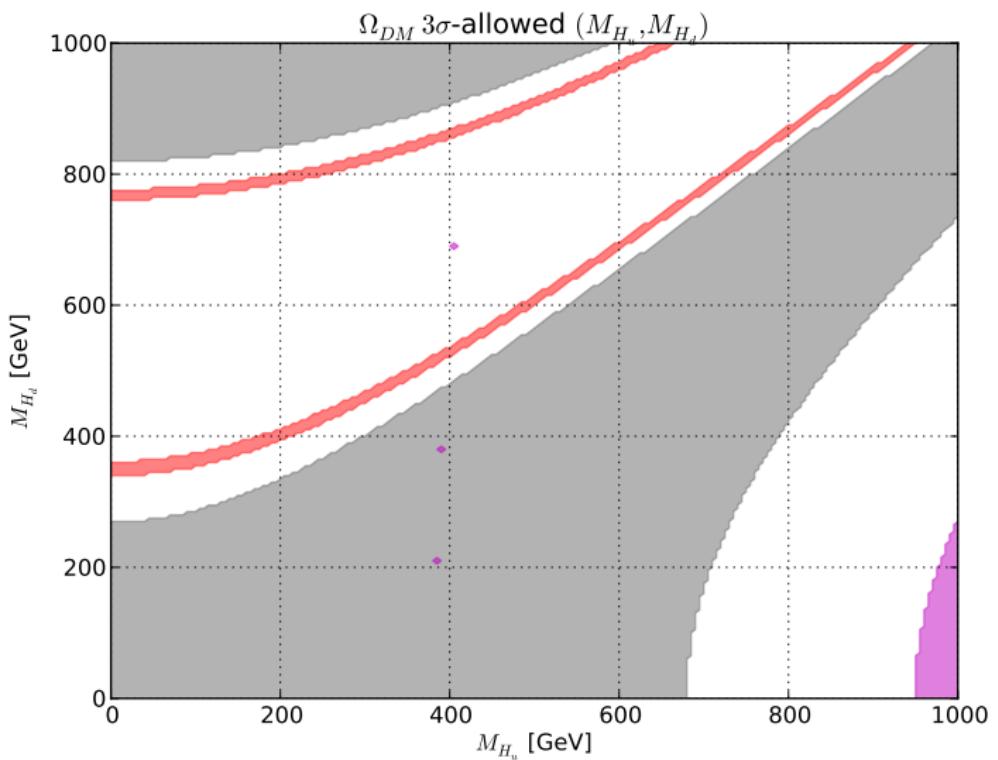
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Heavy 5^* , Heavy 10

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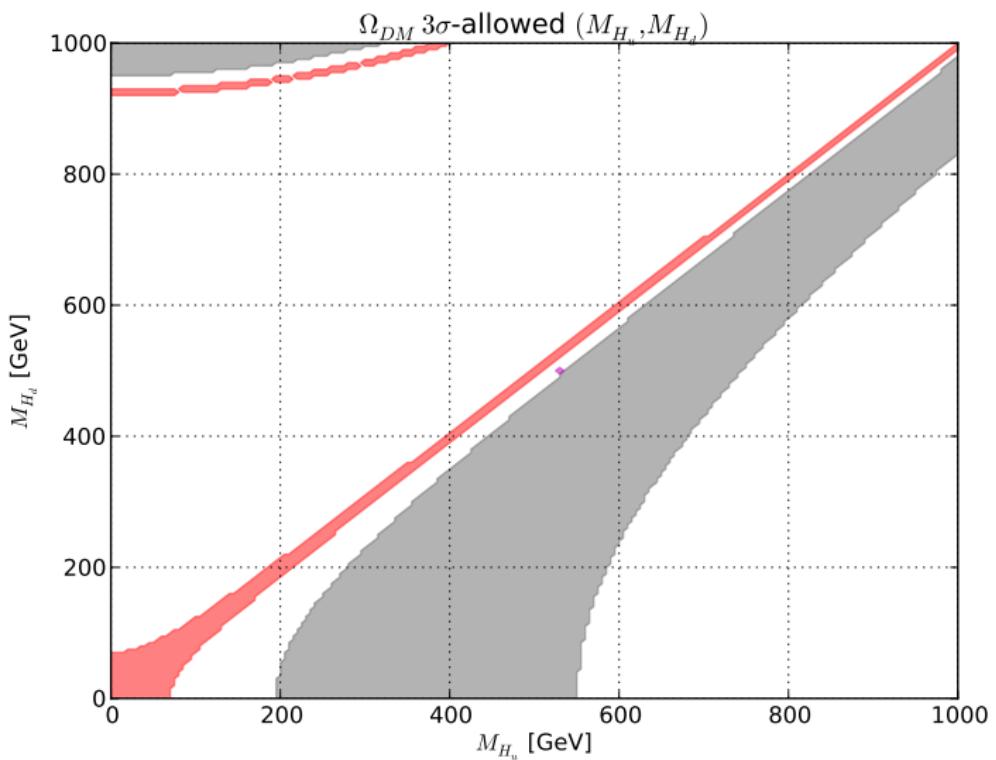
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Heavy 5^* , Heavy 10 ($M_{\text{input}} = M_P$)

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$SU(5)$ of $E_7/SO(10)$

- Global symmetry must be gauged anyway \Rightarrow gauge $SU(5) \subset SO(10)$
- Explicitly break $U(1)^2$ for coupling to supergravity, gives two novinos $N_1, N_2 \sim M_P$
- Superpotential:

$$W = W_Y + W_S + W_\Sigma + W_H + W_{SSB}$$

where

$$W_Y = Y_u \cdot 10 \cdot 10 \cdot 5_H + Y_d \cdot 10 \cdot 5^* \cdot 5_H^*$$

$$W_S = M_\nu \cdot 1 \cdot 1 + M_N \cdot N \cdot N$$

$$W_\Sigma = M_\Sigma \text{Tr } \Sigma^2 + \lambda \text{Tr } \Sigma^3$$

$$W_H = \lambda_1 \cdot 5_H \cdot \Sigma \cdot 5_H^* + \mu \cdot 5_H \cdot 5_H$$

$$W_T = \lambda_T \cdot 5_H \cdot 75 \cdot 50 + \text{c.c.}$$

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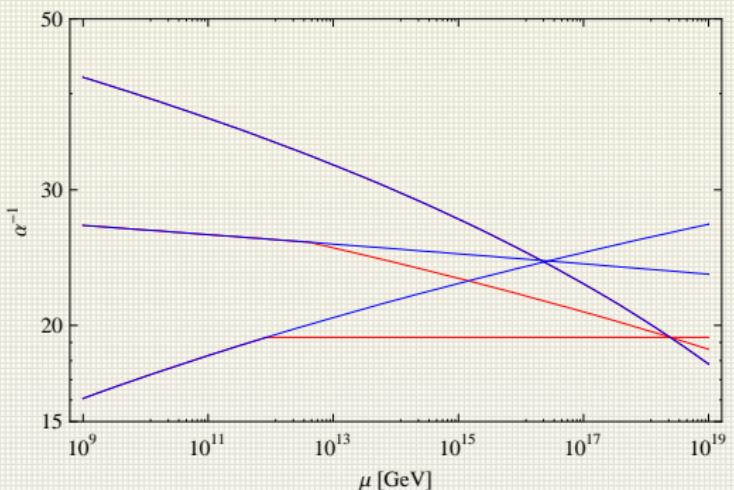
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Making $M_{GUT} \approx M_P$

$$\Sigma \sim 24 \supset (8, 1) \oplus (1, 3) \oplus (1, 1) \oplus (3, 2) \oplus (3^*, 2)$$

- Since gauging anyway, gauge un-Higgsed part of 24
- R -charge of $\text{Tr } \Sigma^3 \neq 2 \implies \lambda \ll 1$ natural $\implies m_{3,8} \ll M_{GUT}$

Choose: $\sim 10^{12}$ GeV



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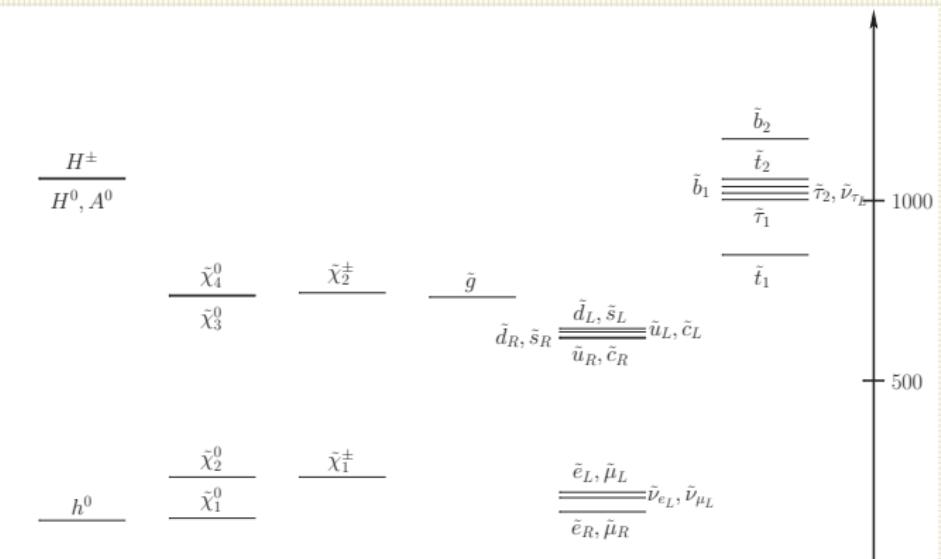
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Spectrum of model point

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$$\tan \beta = 10, m_{1/2} = 300 \text{ GeV}, m_{0(3)} = 1000 \text{ GeV}$$

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Reconstruction of model point

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- 7 fb⁻¹ @ 14 TeV in the LHC
- For this point:

$$\%n_b \geq 1 = 0(+5)\%$$

$$\%n_\tau \geq 1 = 2(+7)\%$$

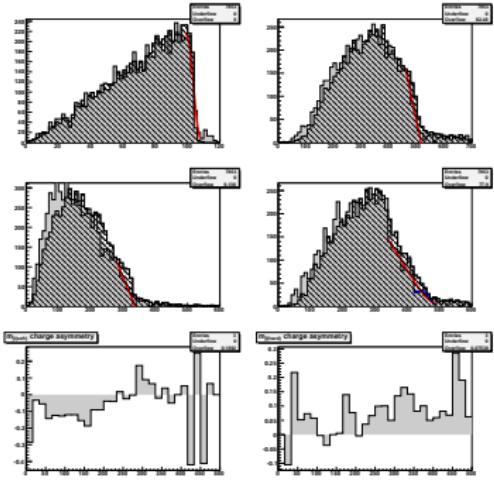
$$\%n_l \geq 2 = 14\%$$

- Typical mSUGRA point:

$$\%n_b \geq 1 = 30(+5)\%$$

$$\%n_\tau \geq 1 = 17(+7)\%$$

$$\%n_l \geq 2 = 4\%$$



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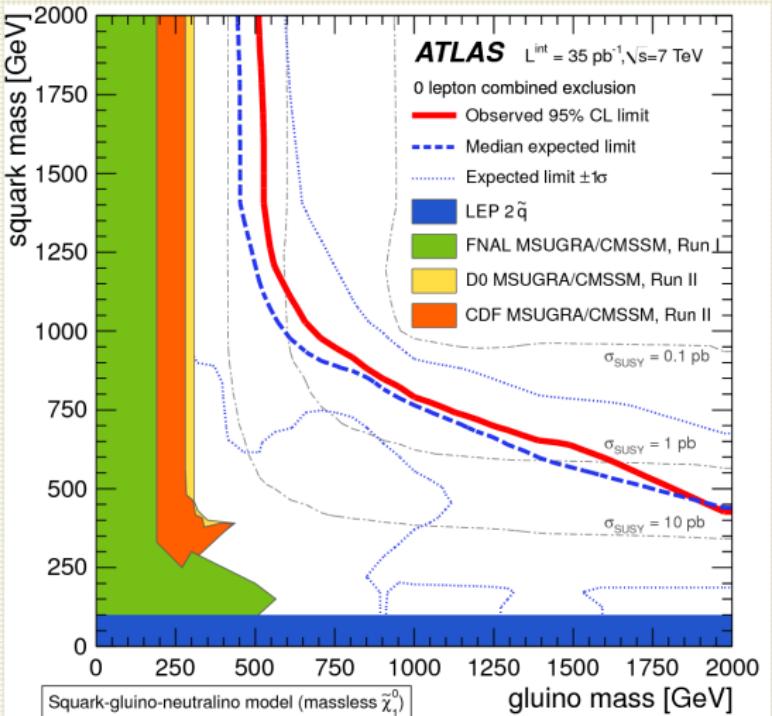
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Newest model-independent constraints

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Simulation events

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Comparing with ATLAS for 35pb^{-1}

- 0-lepton search limited by systematics
- Tension will increase slightly with \mathcal{L}
- Can raise M_3 independently of $M_{1,2}$ for Ω_{DM}

Region	SM	Δ_{JES}	Model
A	118	+32/-23	5
B	10	+4.0/-1.9	5
C	88	+26/-18	4
D	2.5	+1.0/-0.4	0.44

Similar $\sim 1\sigma$ tension v. systematics in jets+1-lepton search

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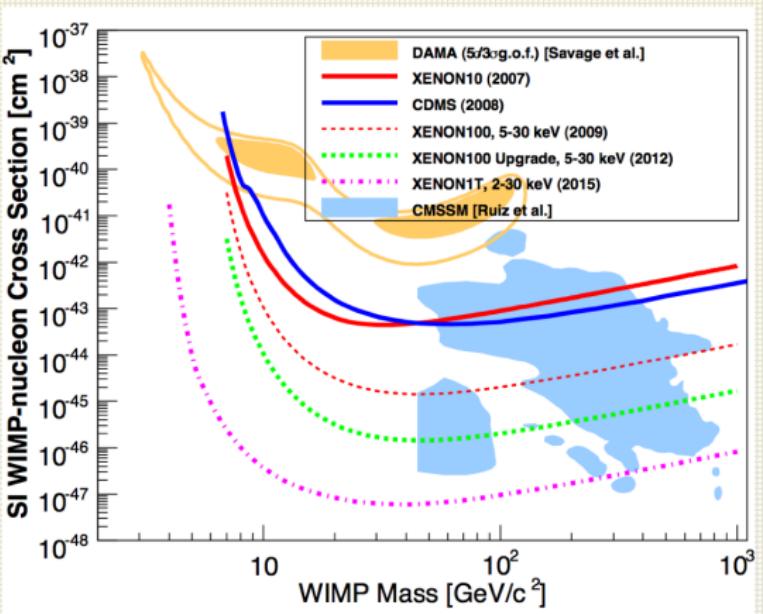
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Dark matter phenomenology

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- $\sigma_{\chi N} \simeq 10^{-46}$ \rightarrow excludable by XENON1T
- χ_1^0 mostly bino, $\langle \sigma v \rangle_{\chi\chi}$ is p -wave suppressed



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Created consistent $E_7/SO(10)$ GUT with distinctive
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ATLAS constraints must be followed carefully; if exclude
the model, must work out explicitly how to lift M_3