

# HIP SEMINAR

## PROFILE LIKELIHOOD METHODS IN TOP QUARK MASS MEASUREMENTS AT THE CMS

BACKGROUND

PRECISION  
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IN-SITU  
SYSTEMATICS

BINNING

CONSEQUENCES

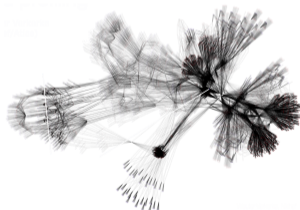
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25th October, 2022



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② PRECISION LIMIT?

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③ IN-SITU SYSTEMATICS

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④ BINNING

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⑤ CONSEQUENCES

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⑦ RESULTS

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⑧ SUMMARY

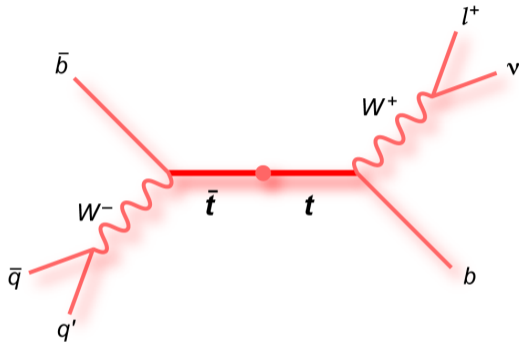
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# TOP QUARK MEASUREMENTS AT THE LHC

- At the LHC, direct measurements of the top quark mass ( $m_t$ ) are the most precise
  - Direct  $\equiv$  reconstruction of decay products
  - Indirect  $\equiv$  anything else, e.g. based on cross-section
- In Run 2 conditions ( $\sqrt{s} = 13 \text{ TeV}$ ):
  - Collision events producing a top quark pair are the most frequent ( $832 \text{ fb}^{-1}$ )
  - Events with a single top quark make a good number 2 ( $264 \text{ fb}^{-1}$ )

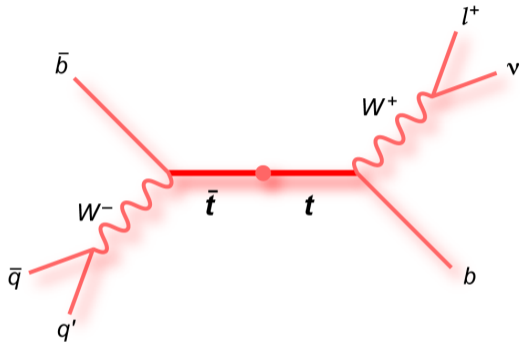


Visualized: the semileptonic top quark pair decay channel



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# DIRECT TOP QUARK MASS MEASUREMENTS

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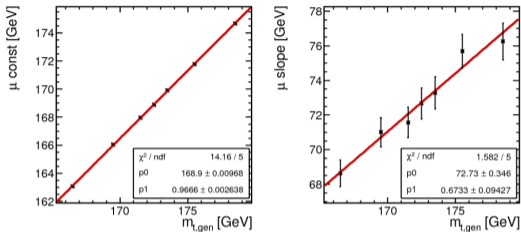
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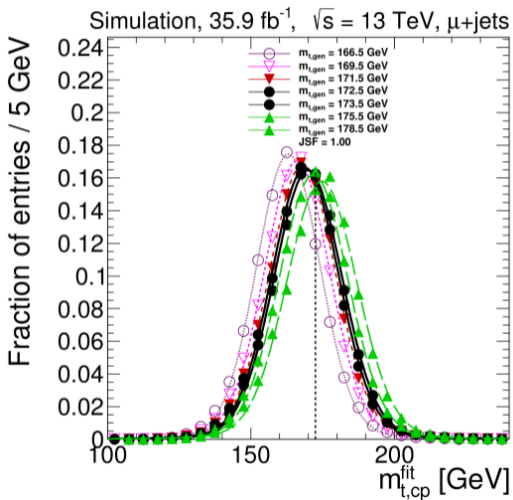
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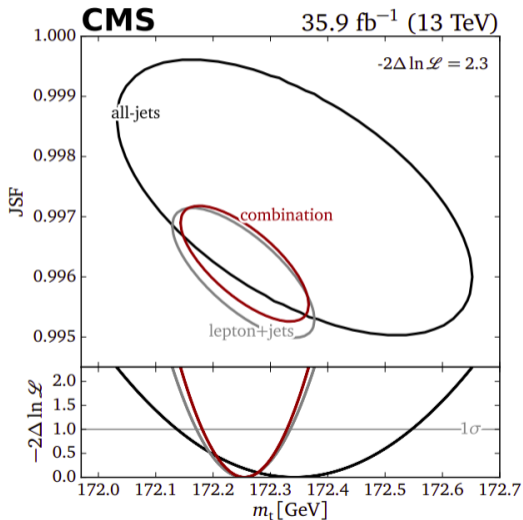


- Classical strategy:
  - Intricate parametrized fits on the observables are made against simulation truth values
  - The parameter dependence on e.g.  $m_t$  is fit against the **simulation truth**



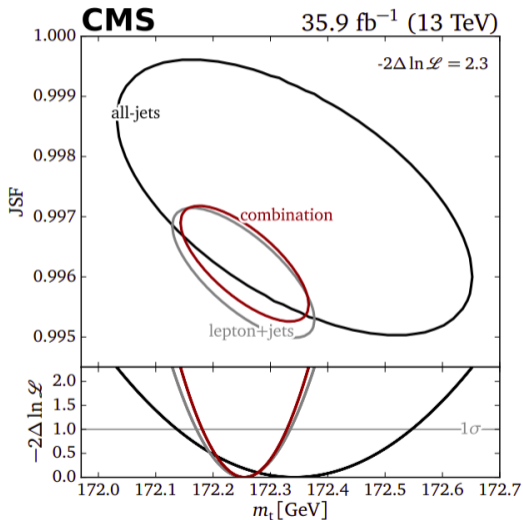
# CONSTRUCTING A LIKELIHOOD

- Using the parametrized fits, a likelihood function for the observables can be constructed
  - This will depend on the chosen parameters
- CMS used to utilize  $m_t$  and Jet Scale Factor (JSF)
  - On the right is the CMS lepton+jets and all-jets likelihood combination on the 2016 data
  - The yielded result:  $m_t = 172.26 \pm 0.07$  (stat+JSF)  $\pm 0.61$  (syst) GeV
- ATLAS used in addition bJSF during Run1



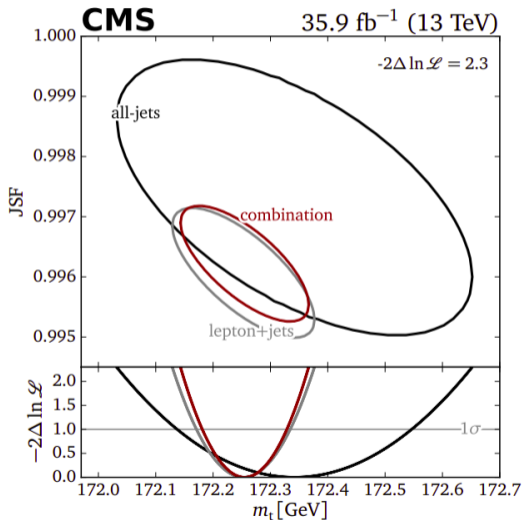
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# HAS THE PRECISION LIMIT BEEN REACHED?

- The stat. (+ JSF) errors in the CMS 2016 measurement are vanishing vs. the systematics: **0.07 GeV vs. 0.61 GeV**
  - **In the classic treatment, the most important systematic error sources cannot typically be reduced by adding statistics**
  - Such error sources include e.g. modelling uncertainties (in the simulations)
  - This indicates that adding limitless statistics to the measurement would at best yield an error of  $\pm 0.00$  (stat+JSF)  $\pm 0.61$  (syst) GeV
- As modelling advances, an inverse trend is observed:
  - With better modelling, the number of potential systematic error sources tends to increase
  - With traditional methods, there is always a risk of double-counting and adding statistical noise to each additional error source

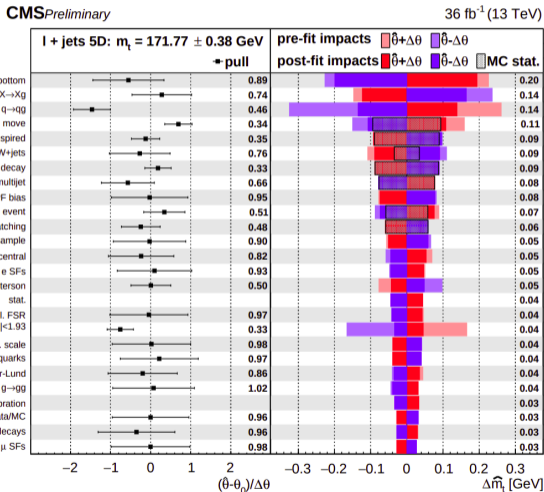


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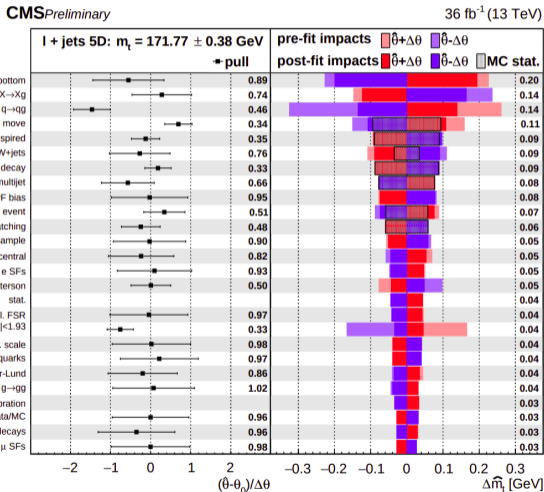


- Earlier this year preliminary results on a new measurement on the 2016 data was released:

- $m_t = 171.77 \pm 0.04$  (stat)  $\pm 0.38$  (syst) GeV
- The analysis uses a split scheme of Final State Radiation (FSR) uncertainties: light quarks and heavy quarks handled separately
- With the old FSR definitions one measures  $m_t = 172.14 \pm 0.04$  (stat)  $\pm 0.31$  (syst) GeV



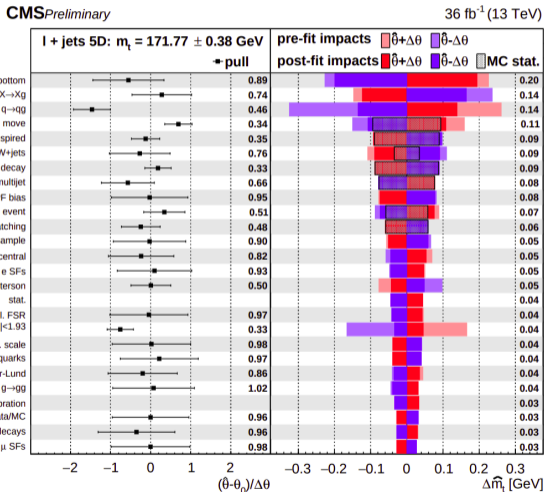
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# WHAT HAS CHANGED?

- Earlier CMS measurements only used two observables, which excelled in a fit on  $m_t^{gen}$  and JSF
  - ① The reconstructed  $m_t$  resonance
  - ② The reconstructed hadronic  $m_W$  resonance
- Three new observables were introduced in the new study
  - These include a ratio between b-jet and W jet  $p_T$  values, earlier utilized by ATLAS
  - Most importantly, the use of the partial resonance between a b-jet and the charged lepton was added in a **phase-space region that was earlier unused**
- **These updates are still not sufficient for explaining the improvement**
  - The missing piece is the introduction of a profile likelihood approach



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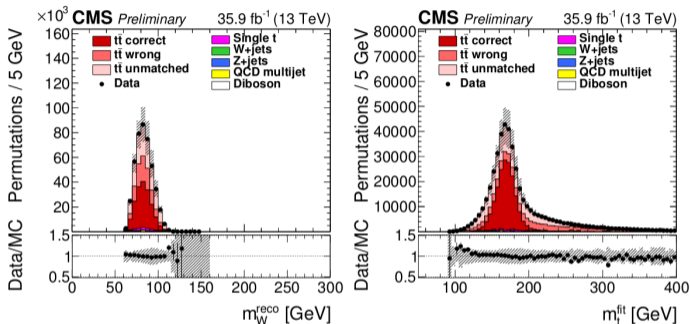
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# A “SIMPLE” IN-SITU MEASUREMENT: JSF



- In the earlier iterations, the JSF **nuisance parameter** is measured in association to  $m_t$  in an *in-situ* manner
  - Measuring the additional dependence is possible, as the  $m_W$  distribution is present in the full likelihood
  - This allows reducing the jet calibration systematics



# MORE NUISANCE PARAMETERS

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- **Could we do the same thing more comprehensively, making nuisance parameters from all systematic error sources?**



# A GENERALIZED IN-SITU MEASUREMENT

- There is no fundamental obstacle for interpreting the systematic variations similarly as JSF variations
  - **This is exactly what is done in the new CMS  $m_t$  analysis on the 2016 Data with a profile likelihood method**
  - In practice this requires a higher level of automation
  - In a final analysis there can easily exist more than 100 systematic uncertainties
- In consequence, the auxiliary JSF parameter becomes unnecessary:
  - This task is already covered by the Jet Energy Uncertainties
- Systematic uncertainties are in general modelled with variations on the central simulated samples
  - This includes two-sided variations, interpreted as the corresponding  $\pm\sigma$  uncertainties
  - ... and one-sided variations, which are interpreted as  $+\sigma$  uncertainties



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# SYSTEMATICS IN THE LIKELIHOOD

- In the old CMS analyses the best results were found with the **hybrid method**:
  - Here, a physically motivated Gaussian prior was imposed on the value of JSF
- In the profile likelihood approach the systematics are interpreted as nuisance parameters  $\theta_k$ 
  - The  $\pm\sigma$  variations are defined to correspond to the values  $\theta_k = \pm 1$
  - The knowledge that the variations are the  $\pm\sigma$  uncertainties is enforced by a Gaussian prior  $\mathcal{G}$  on all of the nuisance parameters  $\theta_k$
  - Similar Gaussian priors are also set on the normalization scales ( $\eta_j$ ) of the simulated samples, according to the cross-section and luminosity uncertainties
- In summary, starting from the likelihood without priors ( $\mathcal{L}_0$ ), the full likelihood stands as:

$$\mathcal{L} = \mathcal{L}_0 \times \prod_{k \in \text{nuisances}} \mathcal{G}(\theta_k) \times \prod_{j \in \text{samples}} \mathcal{G}(\eta_j). \quad (1)$$



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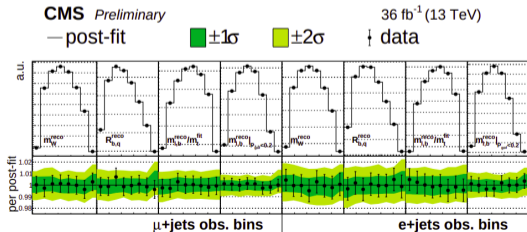
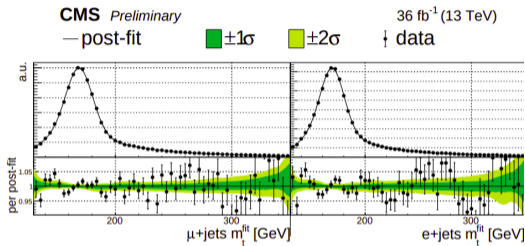
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- The automatization ends up being a major issue in the practical implementation
  - One method for tackling the issue is using a binned likelihood model
  - Here, the only parameters are the bin edges, and the only variables to fit the bin event counts
  - Each bin forms its own Poisson counting experiment
- The 2016 CMS  $m_t$  analysis started with the 5 observables being handled by fit functions, but ended up handling 4/5 observables with binning (left)



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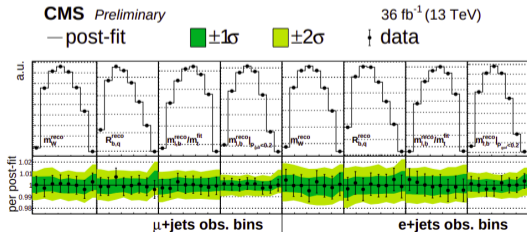
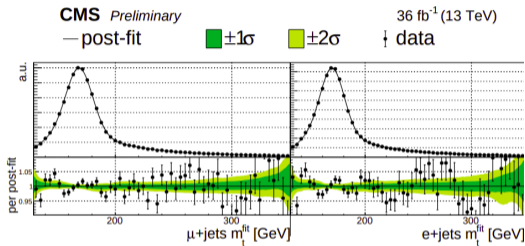
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# CONSIDERATIONS FOR A BINNED ANALYSIS

- Maybe counter-intuitively, the binned approach is the most robust choice
  - **Parametrized fit functions are notoriously volatile, even with a very good function**
  - In contrast, a binned analysis derives its robustness from the **simplicity of Poisson statistics**
- The CMS 2017–2018  $m_t$  analysis uses the same 5 observables as the 2016 analysis, but all with a binned approach
  - This will be further handled in my **thesis defence in Chemicum A129 on 22nd November, 1 pm**



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# A BINNED LIKELIHOOD

- For a fully binned analysis the core likelihood in Eq. (1) can be expressed as:

$$\mathcal{L}_0 = \prod_{i \in \text{bins}} \mathcal{P} \left( n_i \mid \sum_{j \in \text{samples}} (1 + \kappa_j)^{\eta_j} \times \nu_i^j \left( \vec{\theta}, m_t \right) \right) \quad (2)$$

- $\mathcal{P}(n_i | \lambda)$  is the Poisson probability distribution for the (bin) event yield  $n_i$
  - $\nu_i^j$  the expected event yield for the simulated sample  $j$  in the bin  $i$
  - $\vec{\theta}$  collects the nuisance parameters
  - $\kappa_j$  is the fraction of normalization uncertainty for the sample  $j$ , controlled by the nuisance parameter  $\eta_j$ .
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# INTERPOLATION?

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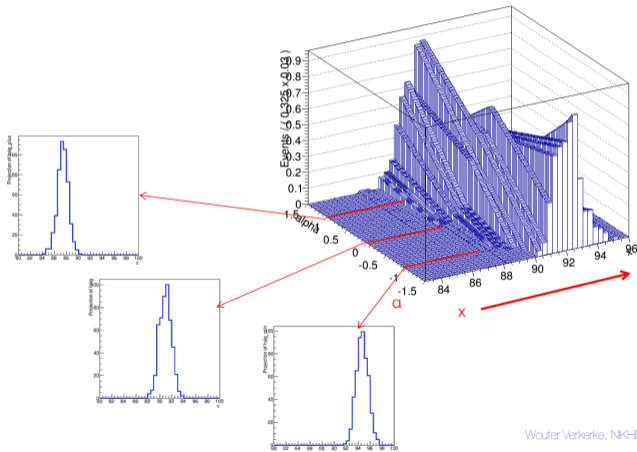
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- For a binned approach, interpolation between the central simulation sample and the variations is achieved through **histogram morphing techniques**

Visualization of bin-by-bin linear interpolation of distribution



Wouter Verkerke, NIKHEF

# MORPHING STRATEGIES

- The two main classes of histogram morphing are **vertical morphing** (see previous slide) and **horizontal morphing**
  - **Vertical morphing** (seen on the previous slide) is relatively simple: each bin is considered separately
  - **Horizontal morphing** can be implemented with various algorithms, and it usually considers both bin migration and vertical effects
- **Vertical morphing** is preferred by its simplicity:
  - No migration between bins, so each bin receives its own parameters
    - The morphing parameters are easy to determine automatically
    - Typically a smooth interpolation function is utilized
  - The parametrization of each nuisance for each bin can be parametrized either ...
    - as a multiplicative factor (more common) or ...
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# VERTICAL VS. HORIZONTAL MORPHING

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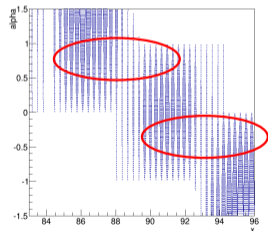
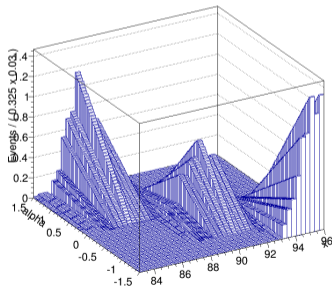
SOURCES

- Even if **vertical morphing** is preferable, **horizontal morphing** becomes necessary when the variations are too big

## Limitations of piece-wise linear interpolation

- Bin-by-bin interpolation looks spectacularly easy and simple, but be aware of its limitations
  - Same example, **but with larger 'mean shift' between templates**

Note double peak structure around  $|a|=0.5$



Wouter Verkerke, NIKHEF

# DETERMINATION OF A GOOD BINNING

- Strict rules follow from the preference of vertical morphing:
  - **Bins must be wide enough**, so that the variations impose relatively small changes in bin contents
  - Simultaneously the bins should be **narrow enough** to provide resolution to measure the problem at hand
- Different problems may prefer different solutions:
  - A “bump hunt” is best done on an evenly spaced binning on the mass axis
  - A precision measurement (such as that of  $m_t$ ) is best performed with bins with even statistics



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# WHAT IS THIS ALL GOOD FOR?

- Nomenclature:
  - There can be one or more **unconstrained Parameters of Interest**
  - For us, there is only one:  $\mathbf{m}_t$
  - This is in contrast to the **nuisance parameters**  $\vec{\theta}$  with Gaussian constraints
- We can define the following **profile likelihood ratio**:

$$\lambda(m_t) = \frac{\mathcal{L}(m_t, \hat{\vec{\theta}}_{m_t})}{\mathcal{L}(\hat{m}_t, \hat{\vec{\theta}})} \quad (3)$$

- Here,  $\hat{\vec{\theta}}$  is the global maximum likelihood solution for the nuisance parameters
- And  $\hat{\vec{\theta}}_{m_t}$  the solution at  $m_t$
- According to **Wilks' theorem**  $-2 \ln \lambda(m_t)$  asymptotically approaches the  $\chi^2$  distribution at good statistics (error  $\mathcal{O}\left(\frac{1}{\sqrt{n}}\right)$  for the 1D profile)



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# NUISANCE PARAMETER CONSTRAINTS

- The *a priori*  $\pm\sigma$  uncertainties correspond to the index  $k$  nuisance parameter values  $\theta_k = \pm 1$
- Considering the full likelihood function, the nuisance parameter can be constrained *a posteriori* to a smaller uncertainty
  - This can be understood through applying Wilks' theorem to the nuisance
  - The nuisance parameter  $\pm\sigma$  (*a posteriori*) **limits are found at**  $-2 \ln \lambda(\theta_k) = 1$
  - Here,  $\lambda(\theta_k)$  is understood in analogy to  $\lambda(m_t)$  – i.e. all parameters but  $\theta_k$  being fixed at the maximum likelihood values
- On the next two slides an example borrowed from Ref. 4 is presented



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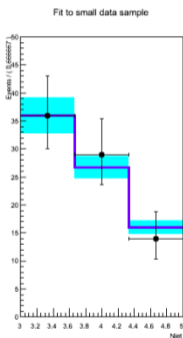
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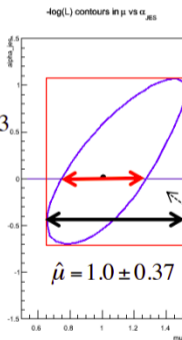


# EXAMPLE: LOOSE CONSTRAINT

Now measure  $(\mu, \alpha)$  from data – 80 events



$$\hat{\alpha} = 0.01 \pm 0.83$$



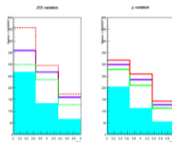
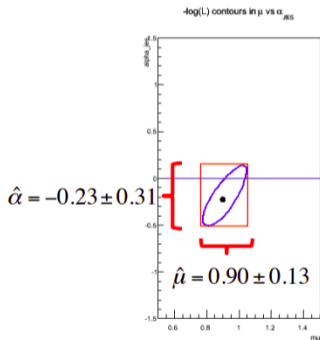
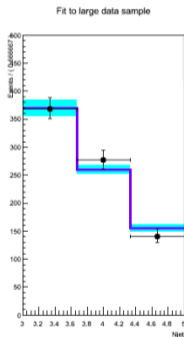
Estimators of  $\mu$ ,  $\alpha$  correlated due to similar response in physics measurement

Uncertainty on  $\mu$  with/without effect of JES

- The Parameter of Interest  $\mu$  is measured aside the Jet Energy Scale (JES) nuisance  $\alpha$ , yielding a slight constraint on  $\alpha$  from the *a priori* values  $\alpha = \pm 1$

# EXAMPLE: TIGHT CONSTRAINT

The next year – 10x more data (800 events)  
repeat measurement with same model



Estimators of  $\mu$ ,  $\alpha$  correlated due to similar response in physics measurement

- Here, statistics start to take over, constraining  $\alpha$  to  $\pm 0.31$ , bringing down also the uncertainty on  $\mu$

# WHICH NUISANCES CAN BE CONSTRAINED

- **It is important to emphasize that the systematics that can be constrained usually have an experimental connection**
  - This includes e.g. the energy scales of jets and charged leptons
  - ...and jet flavor tagging (mainly b-jets)
- But also the choices in the simulation are uncertain and (maybe less obviously) connected to the experiment:
  - Parton Showers (Initial and Final State Radiation)
  - Parton Density Functions
  - Hadronization
  - Underlying event





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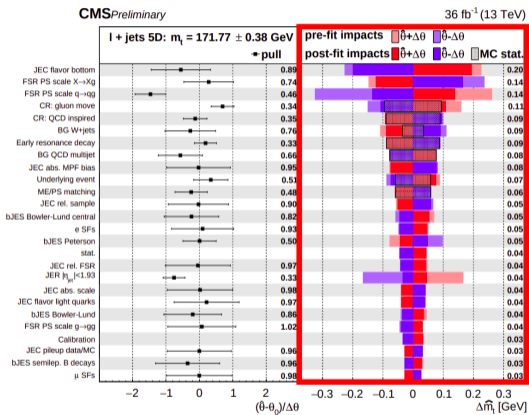
# NUISANCE PARAMETER IMPACTS

- The uncertainty  $\delta m_t$  imposed by the index  $k$  nuisance  $\theta_k$  on the parameter of interest ( $m_t$ ) is called the **impact** of  $\theta_k$  on  $m_t$

- The impacts are found in a few steps

- First fitting all the nuisance parameters and the parameter of interest
- Then, fixing  $\theta_k$  to  $+1$  and  $-1$ , and then fitting on all the other parameters
- The difference in the value of the parameter of interest is the impact

- The same approach works both on the *a priori* nuisance limits  $\theta_k = \pm 1$  and the possibly constrained *a posteriori* limits  $\theta_k = \pm \sigma_{\text{posteriori}}$ , where  $|\sigma_{\text{posteriori}}| < 1$



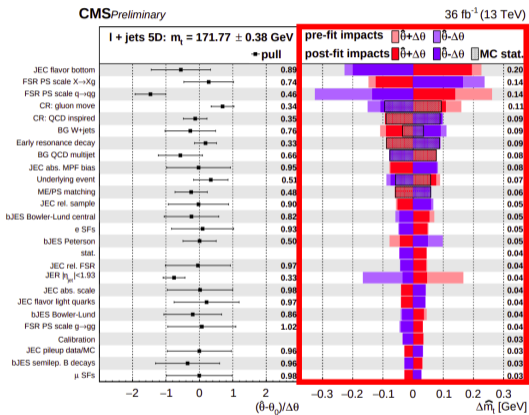
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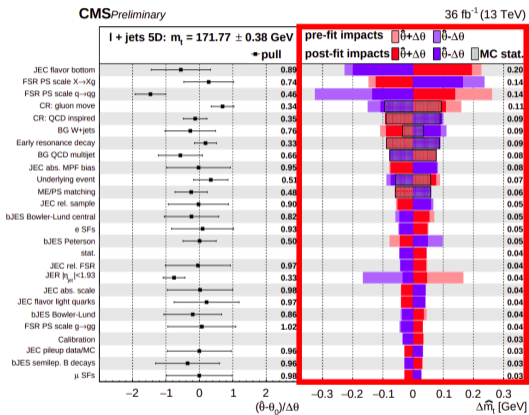
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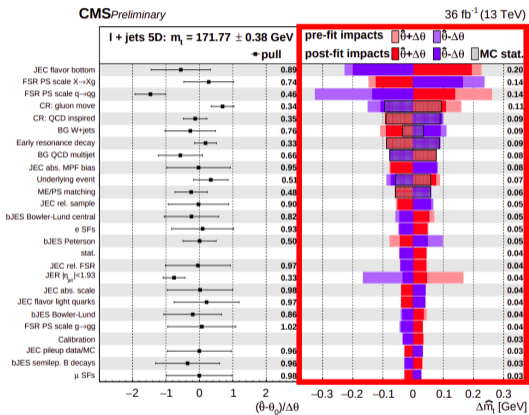
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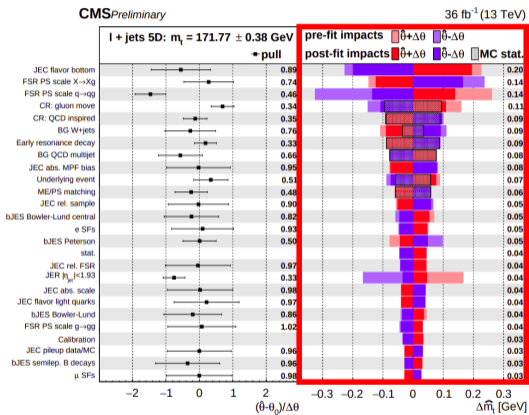
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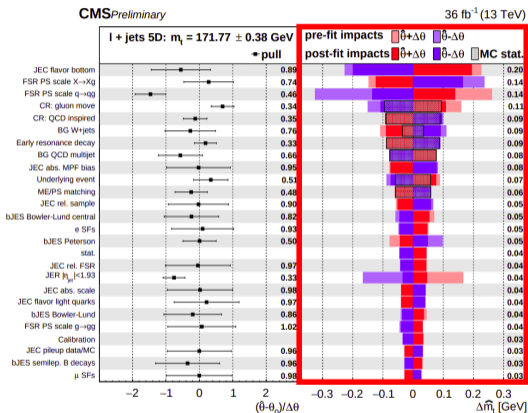
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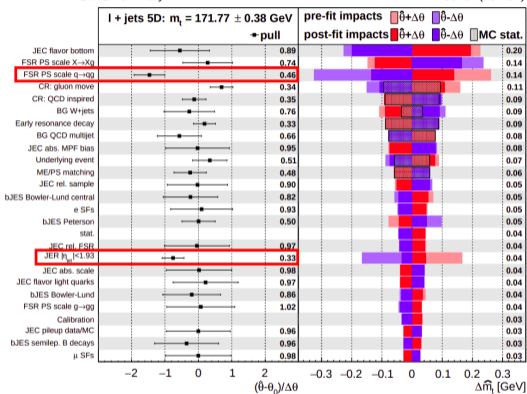
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- A **pull** is defined as the nuisance parameter offset from zero (expressed in the units of the *a priori*  $\pm\sigma$  uncertainty)
- Impacts an pulls can be **blinded or unblinded**:
  - In the blinded mode simulation is used instead of real data
  - The simulation agrees with the central values in the model, so all blinded pulls are zero
  - However, the impacts are meaningful estimators for the true errors that will be measured in data

CMS Preliminary

36 fb<sup>-1</sup> (13 TeV)

- In the unblinded case notable pulls can appear but they are expected to be less than a unity (one sigma)





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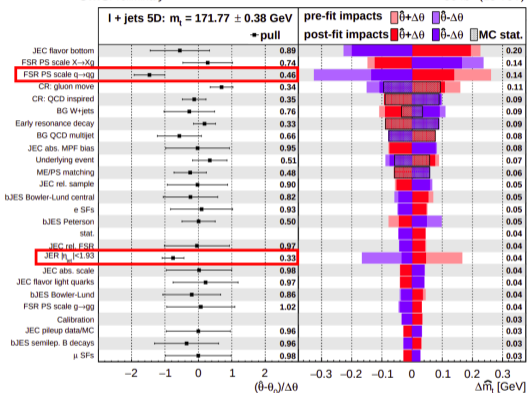
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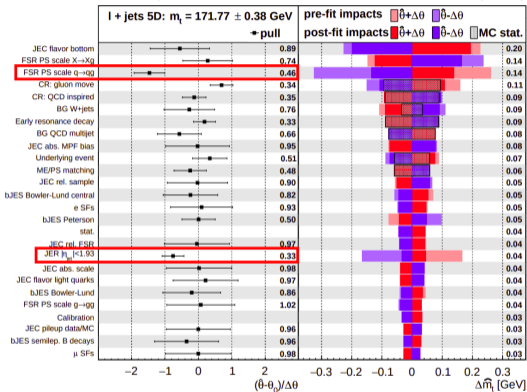
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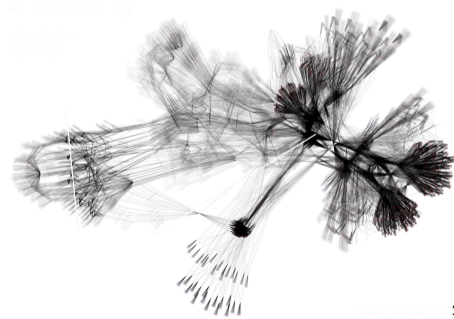
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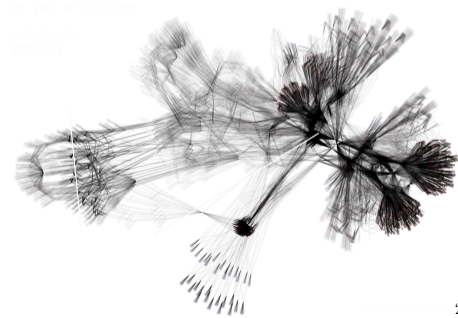
- When estimating the expected number of events in a bin, the **statistical uncertainties are also a kind of modelling uncertainty**:
  - This means that every simulated sample in each bin should receive their statistical nuisance parameter
  - At low statistics they follow the Poisson distribution, which converges to a Gaussian at large statistics
  - If around 10 simulated samples are combined, the number of parameters quickly explodes

- Visualized from Ref. 4: *Atlas Higgs combination model (23,000 functions, 1600 parameters)*
- Partial cure: the *Barlow-Beeston* approach allows the combination of statistical errors from separate samples



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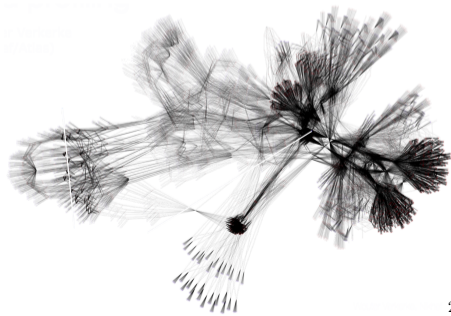
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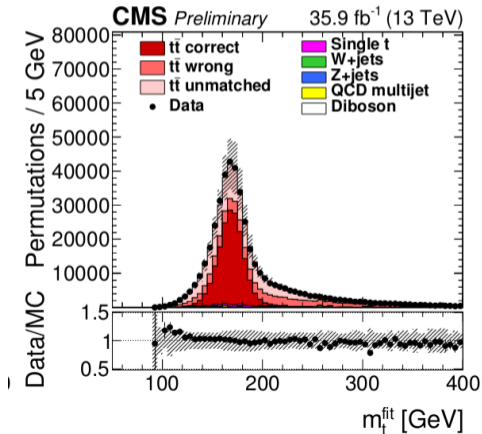


# COMBINATION OF SIMULATIONS

- Integrated luminosity  $\mathcal{L}$  is the general measure for the amount of events in a sample
- When simulation and data are compared, each sample should be weighted by:

$$w^{\text{Sim}} = \frac{\mathcal{L}^{\text{Data}}}{\mathcal{L}_{\text{Eff}}^{\text{Sim}}} = \frac{\sigma^{\text{Sim}} \mathcal{L}^{\text{Data}}}{N_{\text{Eff}}^{\text{Sim}}} \quad (4)$$

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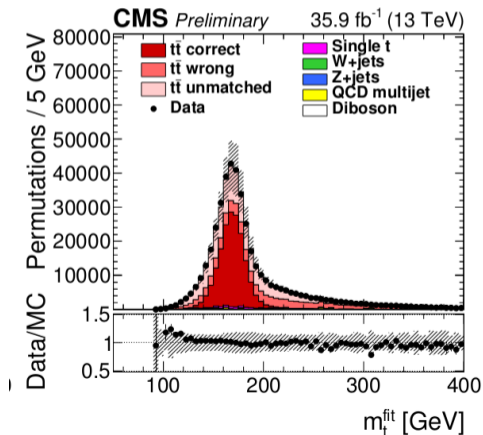


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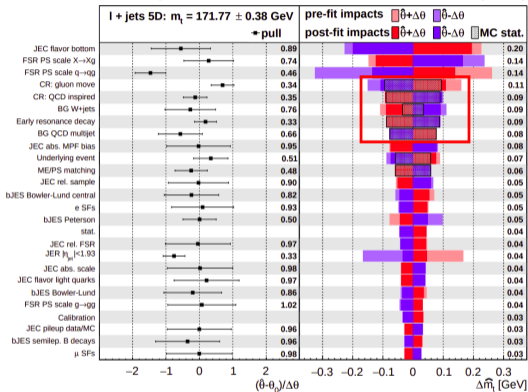
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- There are three common methods for quantifying systematics

- 1 Reweighting the simulated sample (almost 100% statistical correlation)
- 2 Rescaling object (e.g. jet) energies (quite close to 100% statistical correlation)
- 3 Separate simulated samples (no statistical correlation)

- In the third case one should provide separate bin-wise nuisance parameters for the systematic variation sample
  - This is often a weak point and poorly implemented in the common tools

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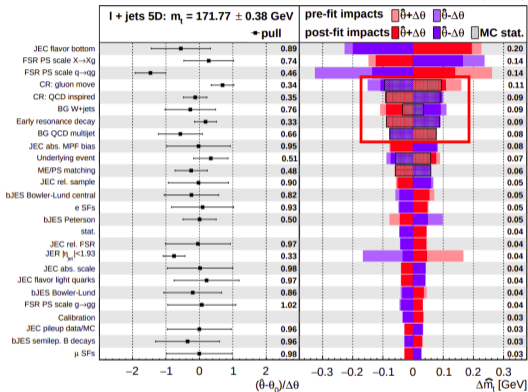
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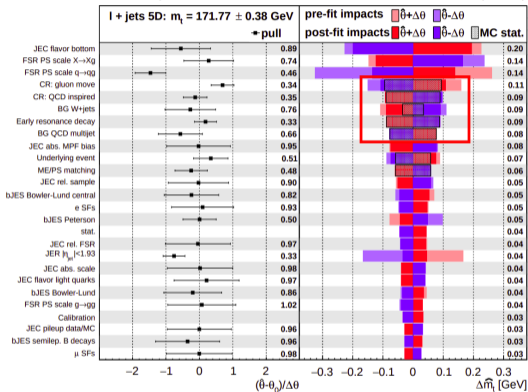
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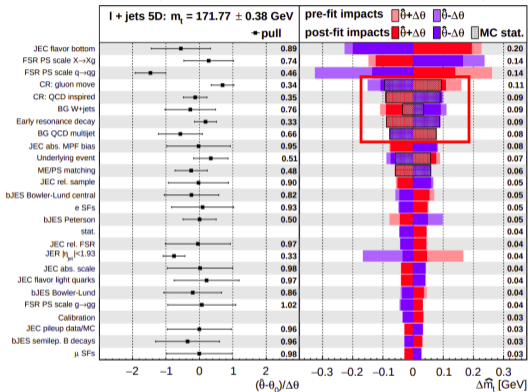
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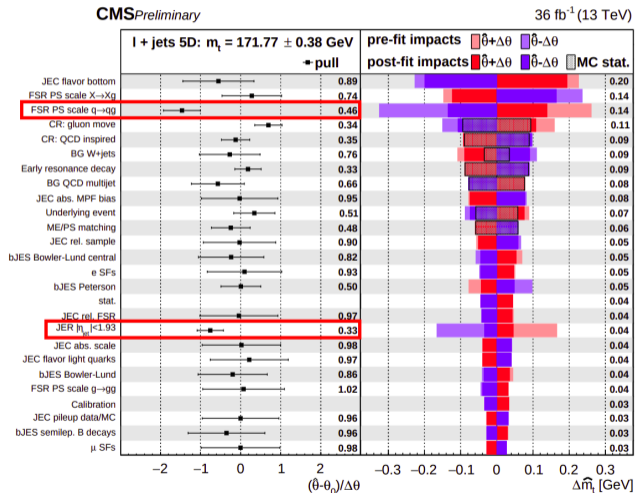
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- In the third case one should provide separate bin-wise nuisance parameters for the systematic variation sample
  - This is often a weak point and poorly implemented in the common tools

CMS Preliminary

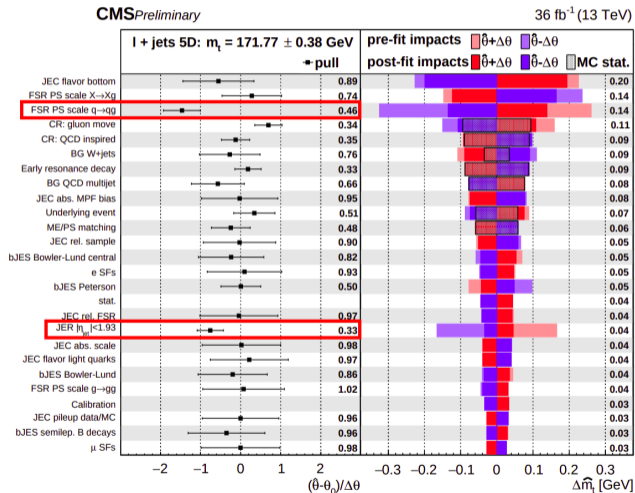
36 fb<sup>-1</sup> (13 TeV)

# A FINAL LOOK AT THE NEW CMS RESULTS



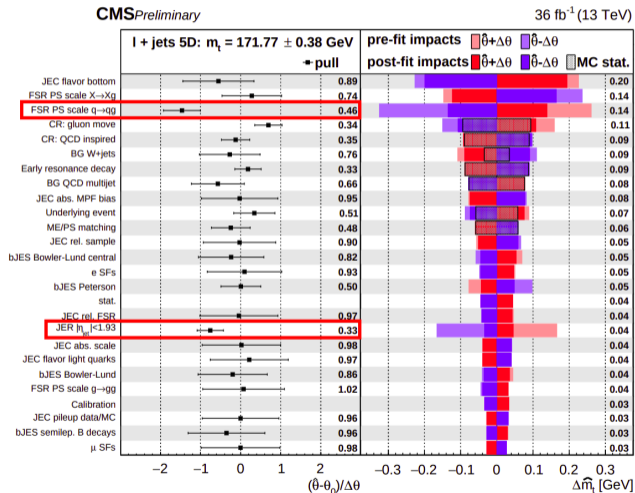
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- **The number of recorded events is around 3 times larger in 2017–2018 w.r.t. 2016**
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  - For systematics that are well constrained by the measured data, one can optimistically expect similar scaling as for statistical errors
  - In consequence, the error scaling factor in 2017–2018 w.r.t. 2016 can go to  $1/\sqrt{3} \approx 0.58$
- Other improvements:
  - The method is more stable with a binned with also for  $m_t$
  - The 2016 analysis scales away the even yield both in the central simulation and systematic variations
  - With Poisson statistics the (absolute) event yields are an important part of the whole picture, so we see including these as an important improvement



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# HISTORICAL REVIEW

- **One could ask *why now?***
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- Slow diffusion from one field of study to another:
  - The early adopters at the LHC were the Higgs hunters, as can be reviewed from [this profile likelihood article](#) and [this article on CMS/ATLAS methods on Higgs boson searches](#)
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- **The most precise  $m_t$  measures at the LHC are currently performed on the top quark pair topology, based directly on the top quark decay products**
- In early Run 2 it seemed that the precision limit for such  $m_t$  measurements had been reached
  - This is ruled by irreducible (systematic) error sources
- Profile likelihood methods ruled this as a false assumption
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## SOURCES

- These slides can be reached at [tinyurl.com/hiphannu](https://tinyurl.com/hiphannu)
- ① <https://cds.cern.ch/record/2674989>
- ② <https://arxiv.org/pdf/1812.10534.pdf>
- ③ <https://cds.cern.ch/record/2806509/files/T0P-20-008-pas.pdf>
- ④ <https://www.precision.hep.phy.cam.ac.uk/wp-content/people/mitov/lectures/GraduateLectures/Advanced-Statistics-Verkerke.pdf>

