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# High Precision Timing: Detector Simulation Challenges - an ATLAS HGTD perspective

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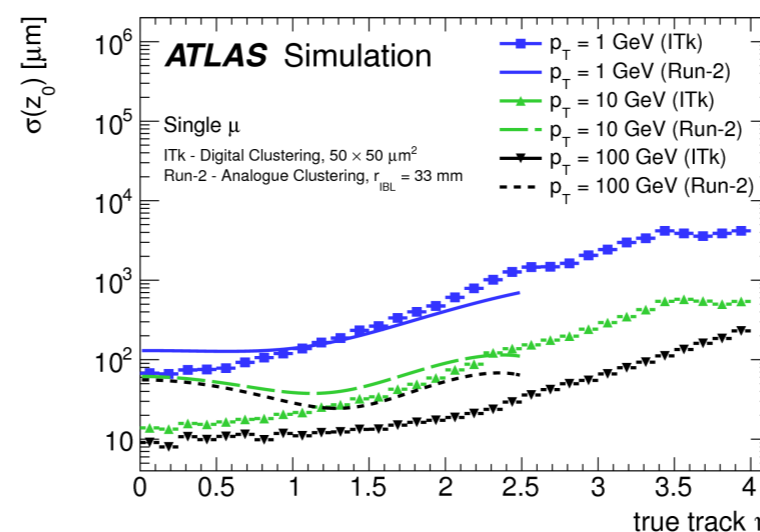
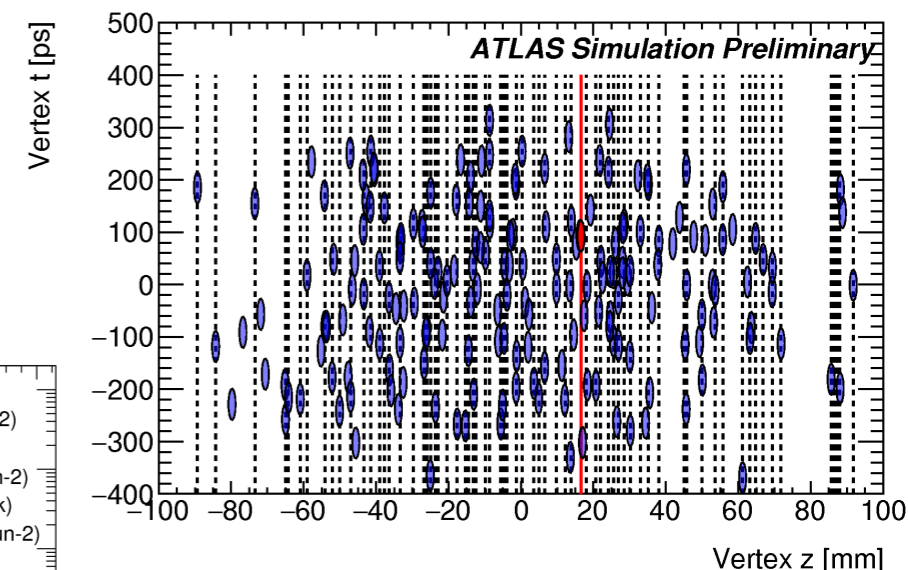
*On behalf of the ATLAS HGTD group*

# Introduction - HGTD in a nutshell

The **High Luminosity** phase of the LHC represents great **challenges** and **opportunities** to test new technologies for future detectors.

In the forward regions, the **pile-up density** will be comparable to the **longitudinal ( $z_0$ ) track resolution**.

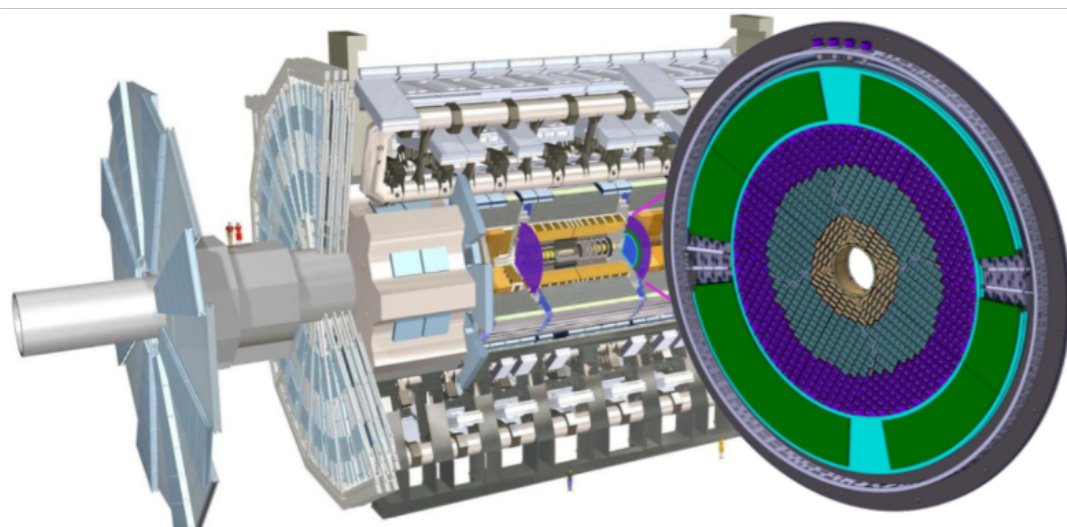
- Object performance reduced;
- Track-to-vertex ambiguity;
- **Solution**: add timing information with a small enough resolution, especially at high  $\eta$ .



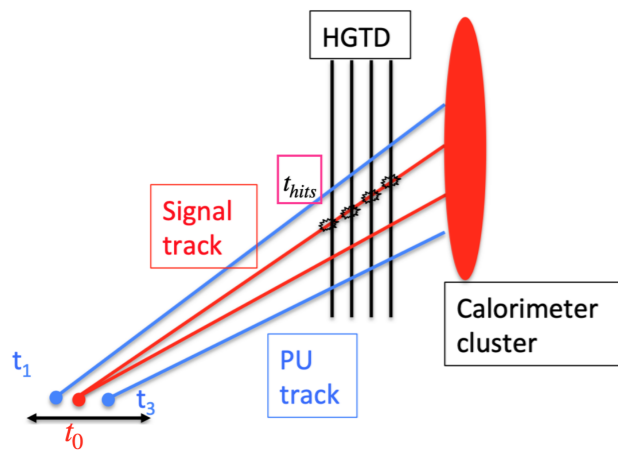
Coverage on **both endcap regions** outside of upgraded Inner Tracker:  $2.4 < |\eta| < 4.0$

Per track (2/3 hits per track) timing resolution of  $30 - 50 \text{ ps}$  up to a fluence of  $2.5e15 \text{ n}_{eq}/\text{cm}^2$  (replacement foreseen) :

- **LGAD** (Low Gain Avalanche Detectors) of  $50 \text{ micrometers}$  thickness;
- $< 50 \text{ ps}$  resolution per hit.

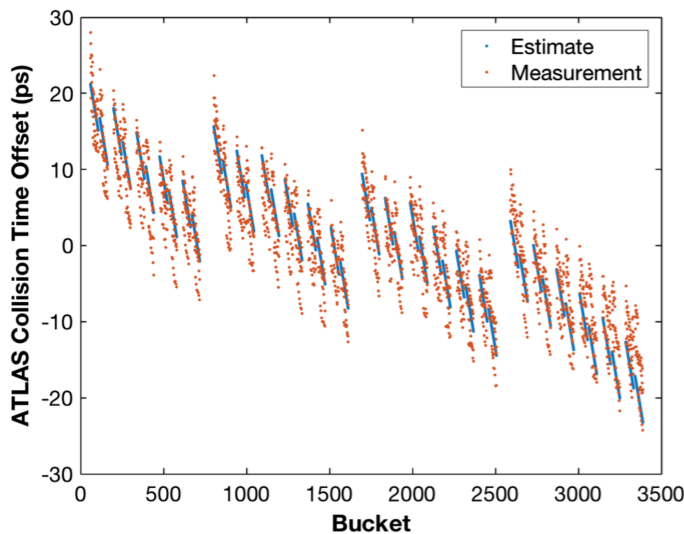


# How time is distributed in HGTD



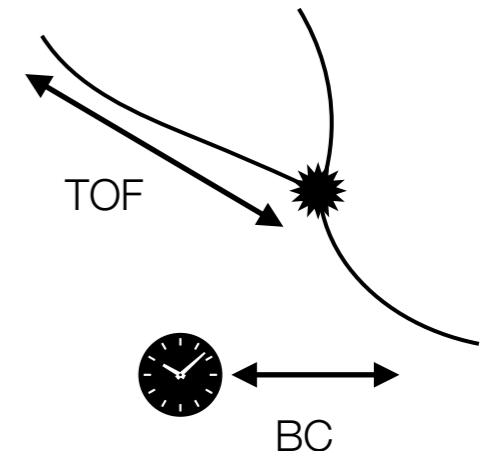
**TOF:** The first effect we need to account is the time for the particle to reach the detector. For a same z-position, the time will differ depending on the hit R and the layer.

- One of the biggest effect we have to account for ~11 ns, but not related to clock.
- This is parameterised for each hit depending on its coordinates.



**BC:** with time the bunch crossing collision time changes slightly.

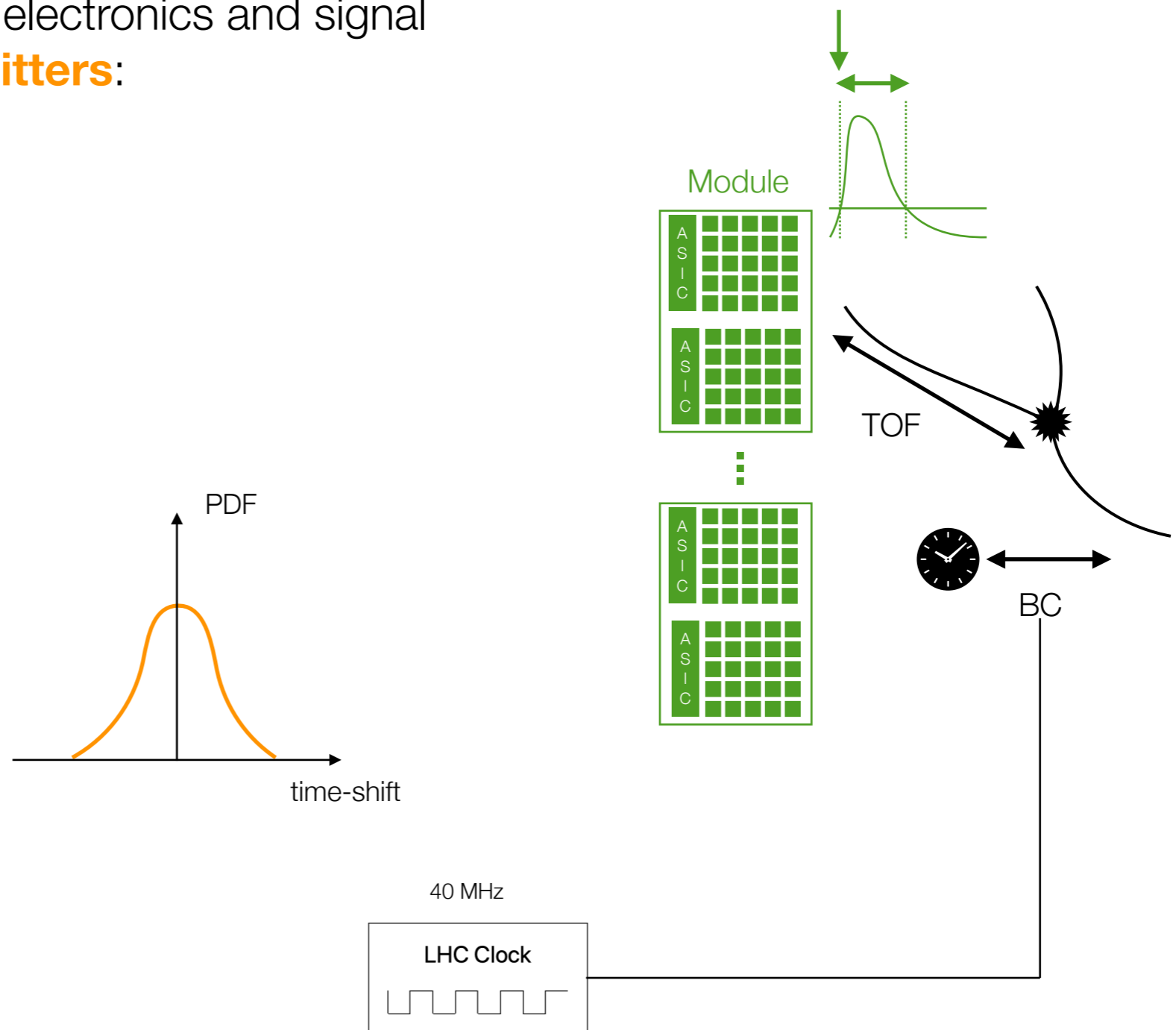
- This is measured and can be parametrised easily with respect to the time (or event number for simulation).



# How time is distributed in HGTD

The first effects concerning the electronics and signal propagation are accounted as **jitters**:

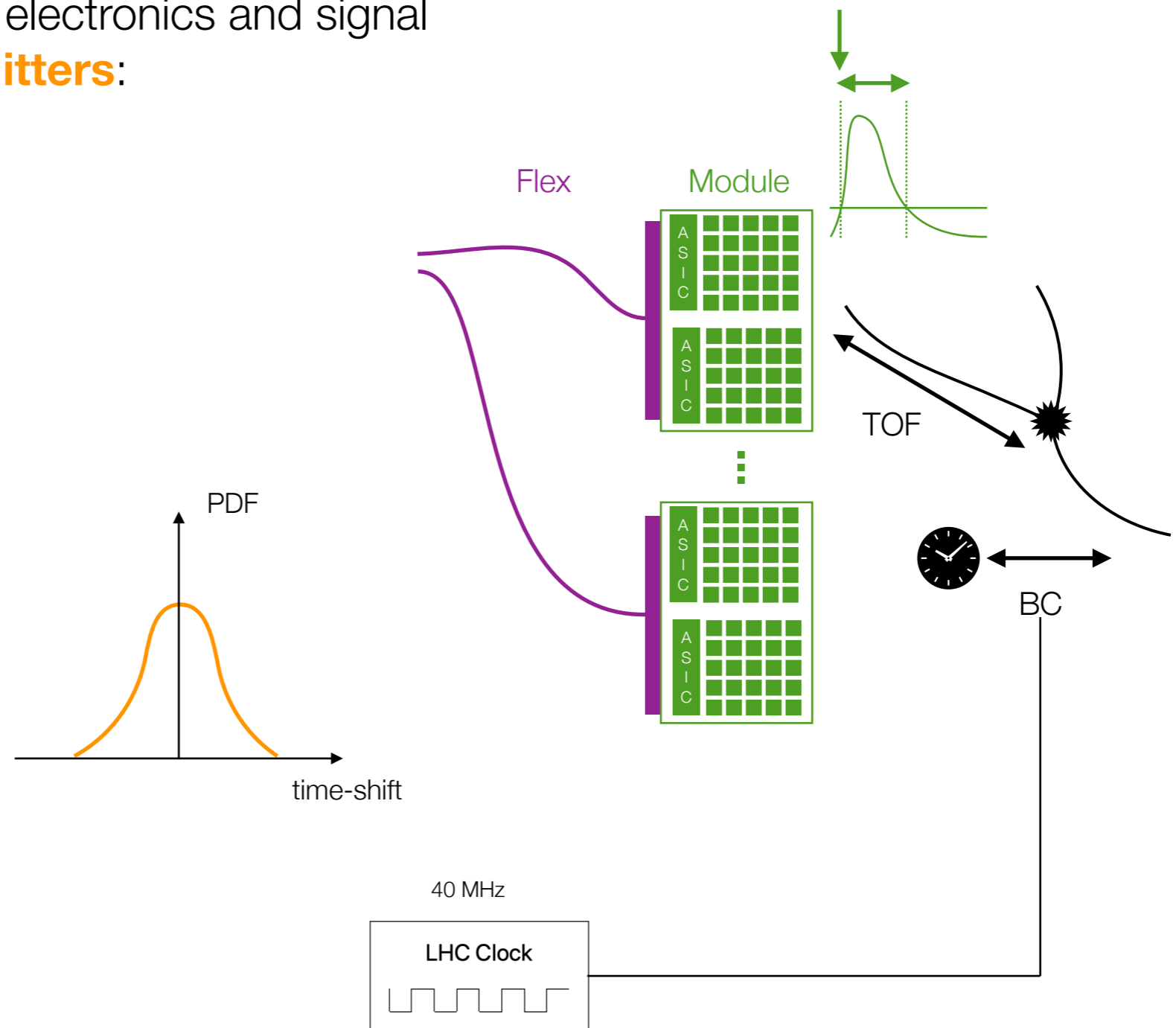
- For the **ASIC** a conservative **35 ps jitter**: path-length differences and internal jitter.



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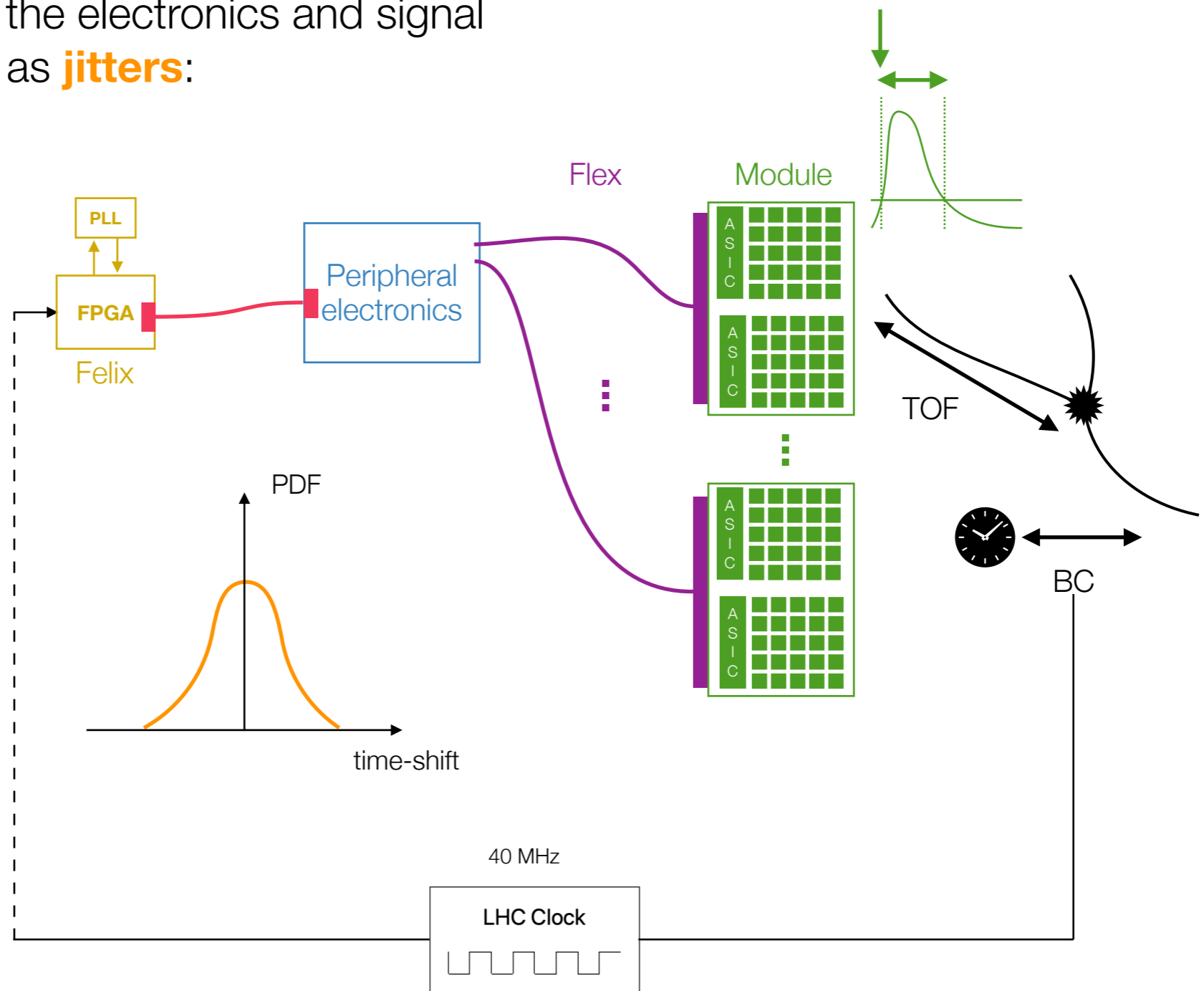
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- The **Flex** effects include environment noise, inherent time jitters performance: measured by the HPTD group: **5 ps**.



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- For the **ASIC** a conservative **35 ps jitter**: path-length differences and internal jitter.
- The **Flex** effects include environment noise, inherent time jitters performance: measured by the HPTD group: **5 ps**.
- For **Felix** the internal clock jitter has been estimated to be **5.2 ps**.

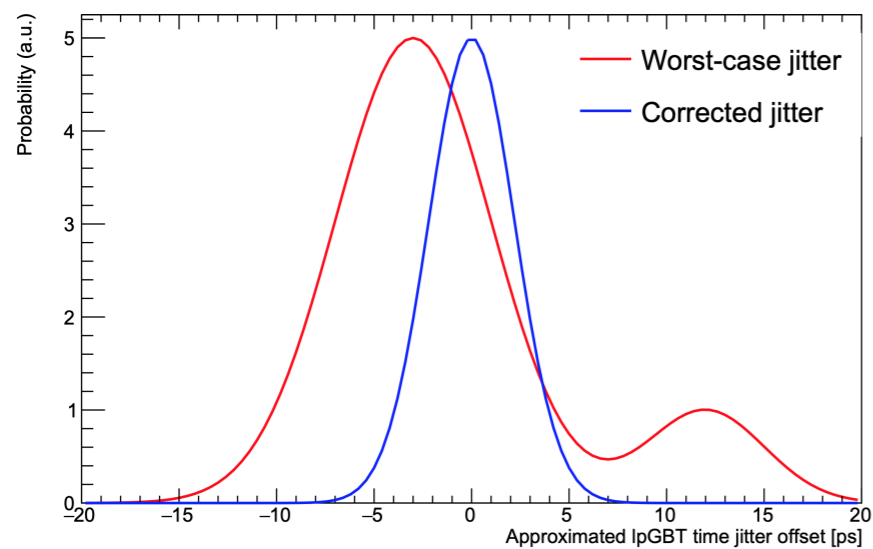


All these effects are **additive** in this simple analysis!

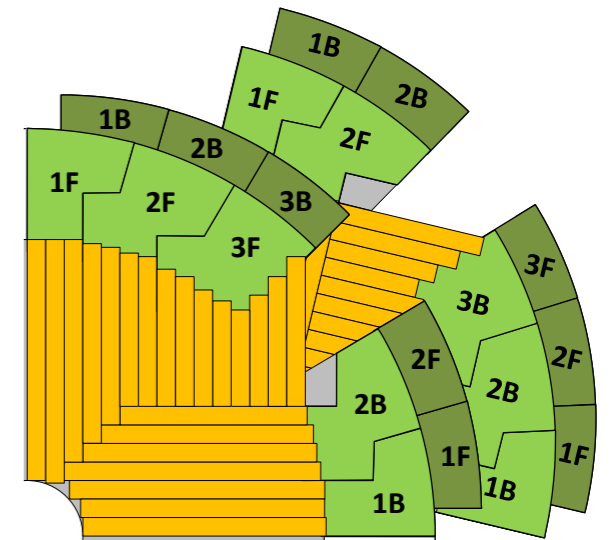
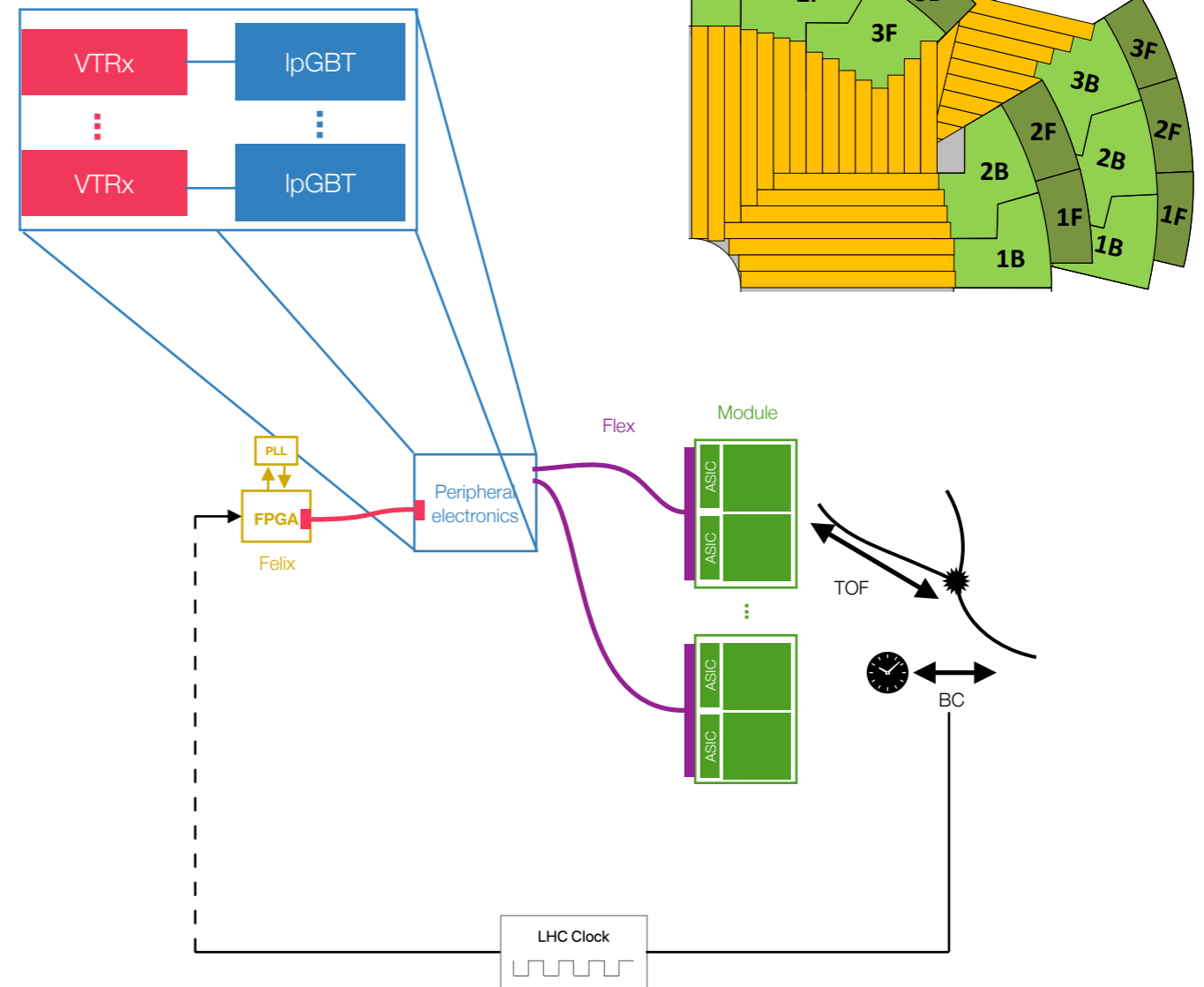
# How time is distributed in HGTD

In between the ASIC and the off-detector electronics, the Peripheral Electronic Boards (PEBs) contain all the ingredients to deal with: Trigger, Data, DCS and timing distribution.

It also contains a specific tool to help us clean part of the jitters: the **Low Power Giga Bit Transceiver** (lpGBT) chip:

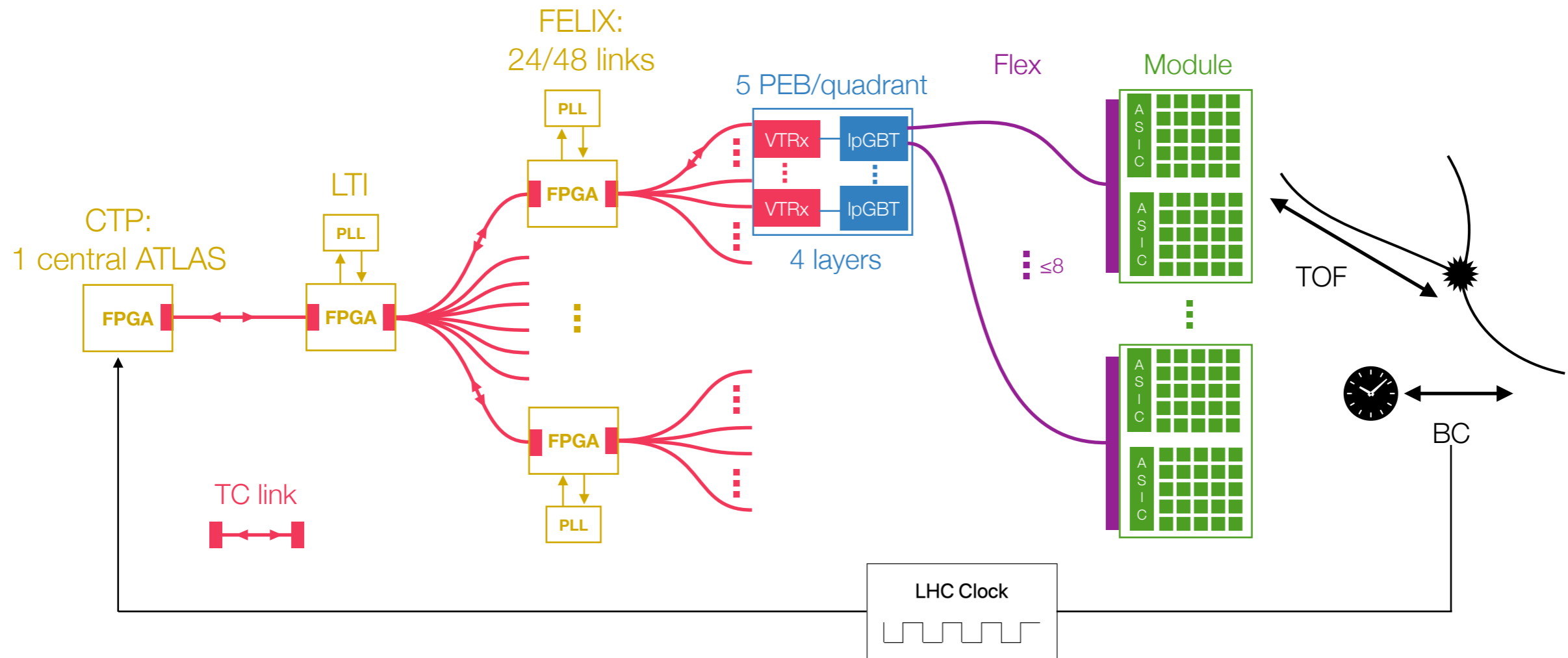


Recent results are showing that we could *remove the FELIX jitter* and only consider a **global 2.2 ps effect**.



# How time is distributed in HGTD

The global sketch of the time distribution should more look like this:



One aspect that hasn't been mentioned here is the **phase determinism** and its potential impact: we have many links that can reset at some point.

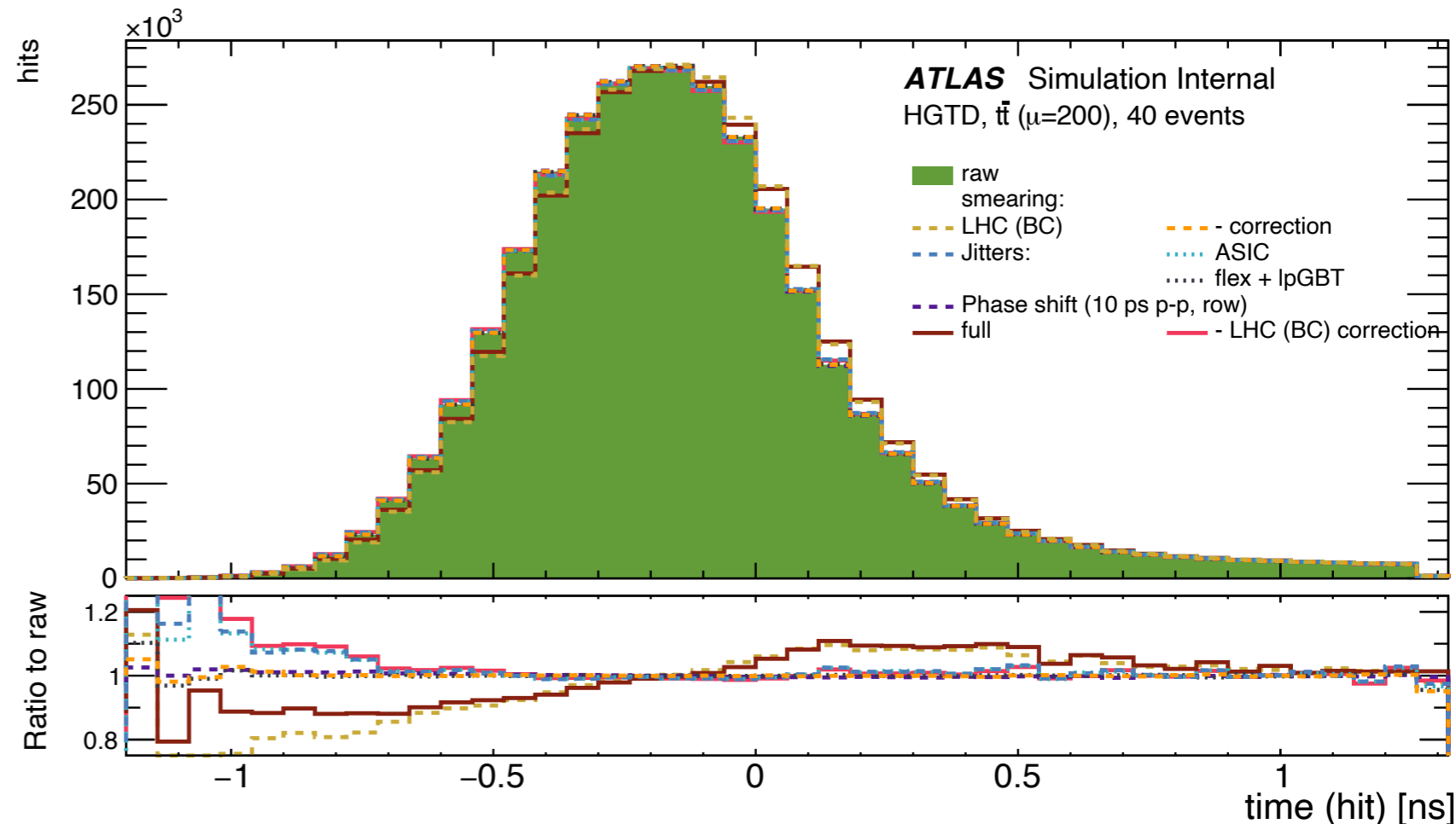
It may be **rare (TBD)**, but will desynchronise some **part of the detector** (which fraction ?) and by a **"big" amount** ( $O(10 \text{ ps})$ ).



# How to look at these effects

The first type of plots one can show is the effect of the various effect on a simulated **Time of Arrival (ToA)** inside the ASIC time window selected:

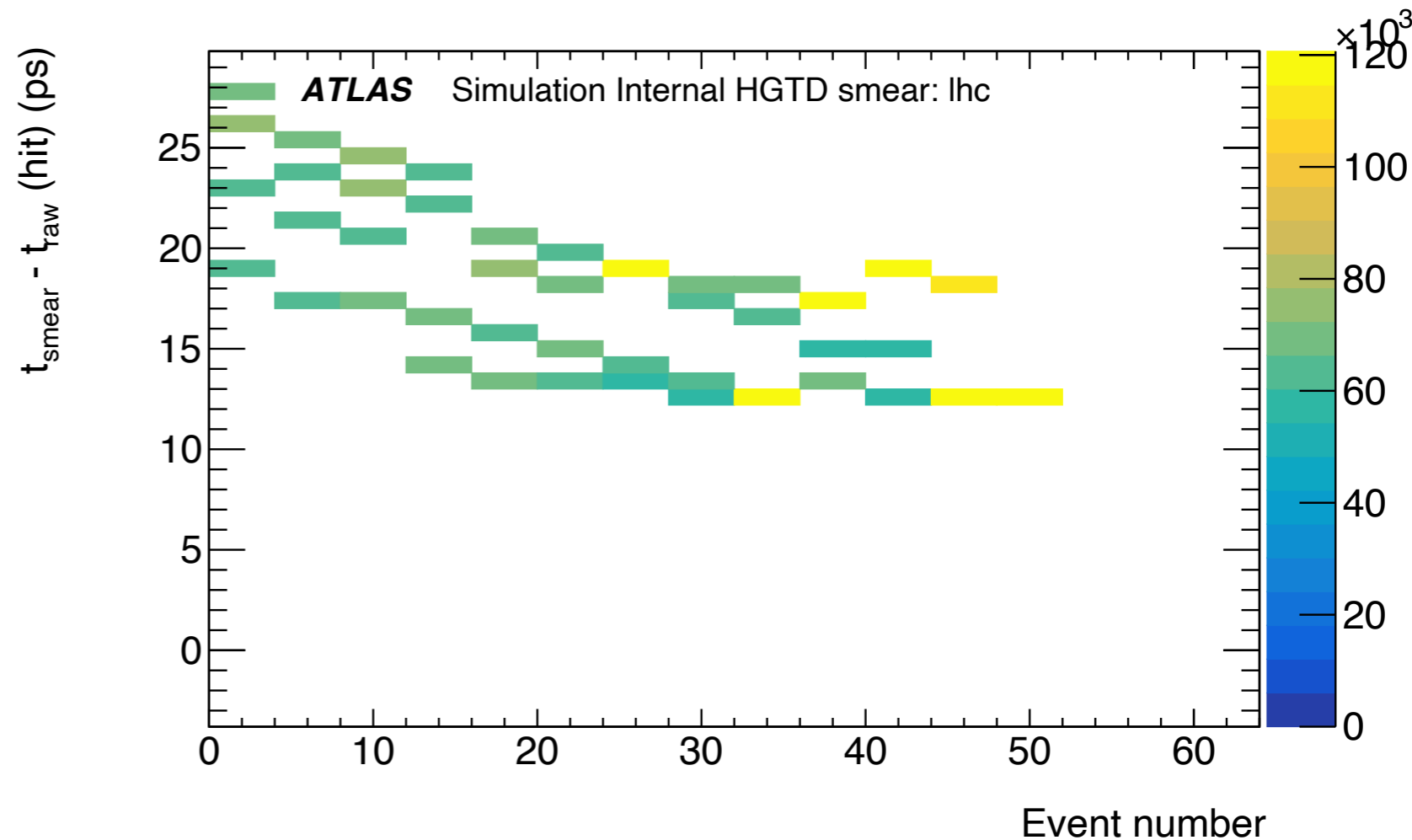
- A 2.5 ns window with a 100 ps step is selected by a "Phase Shifter";
- This selection is made once per ASIC and correct for most of Time of Flight effects;
- The other effects are shown as deviation from this "raw time".



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- A 2.5 ns window with a 100 ps step is selected by a "Phase Shifter";
- This selection is made once per ASIC and correct for most of Time of Flight effects;
- The other effects are shown as deviation from this "raw time".
- These effects can be also change over time.
- The remaining question is whether we can try to correct them ?



## Residual time calibration

$$t[\text{calib}, \text{channel } i] = t[\text{raw}, \text{channel } i] - \langle t_{\text{hits}} \rangle$$

$$C_N = \langle t_{\text{hits}} \rangle$$

Average of the hits on some time period

# Opening discussion

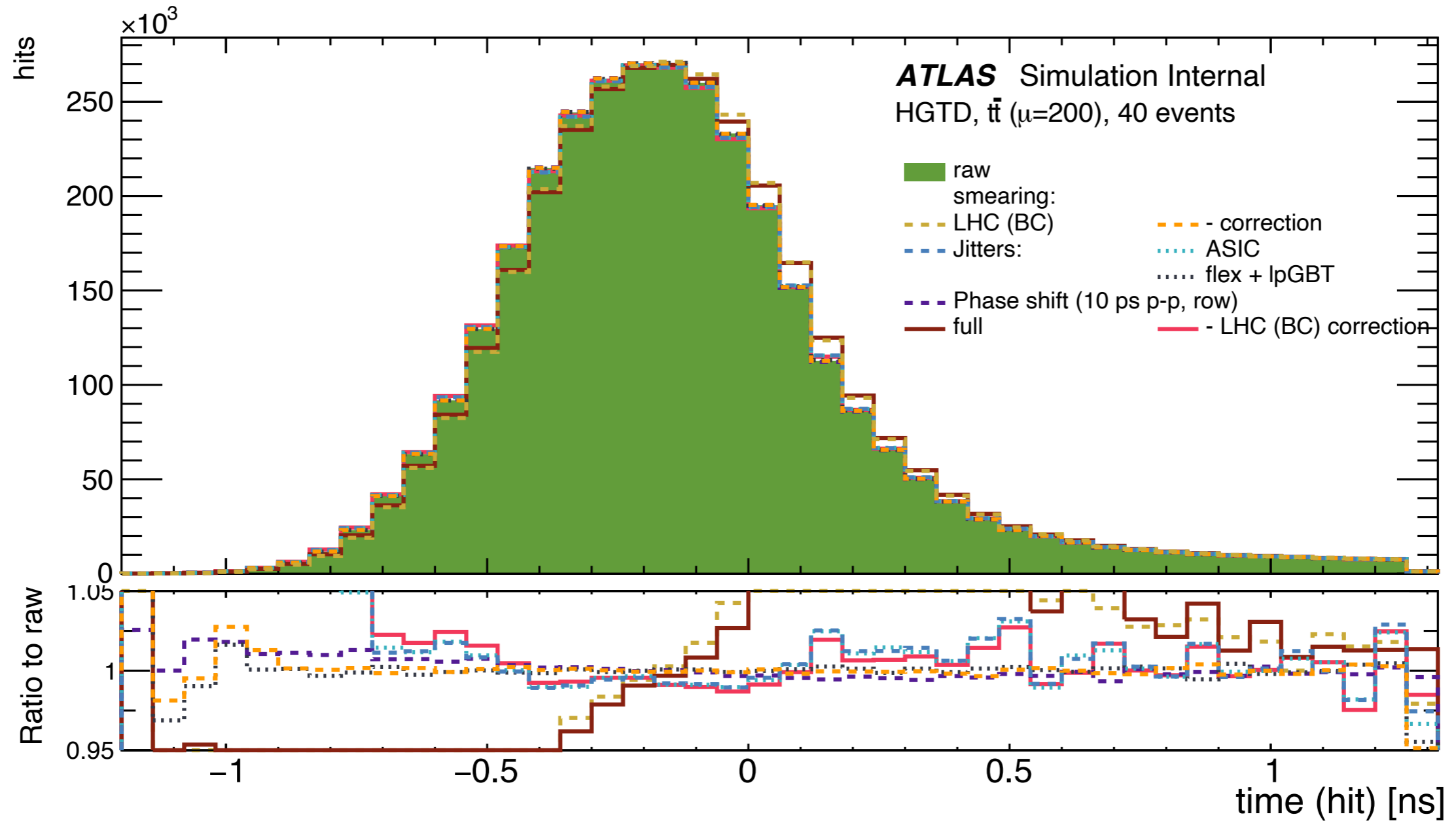
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In the studies conducted we have opened maybe **more questions** than solved issues, allowing to connect **detector level** and **global LHC requirements**:

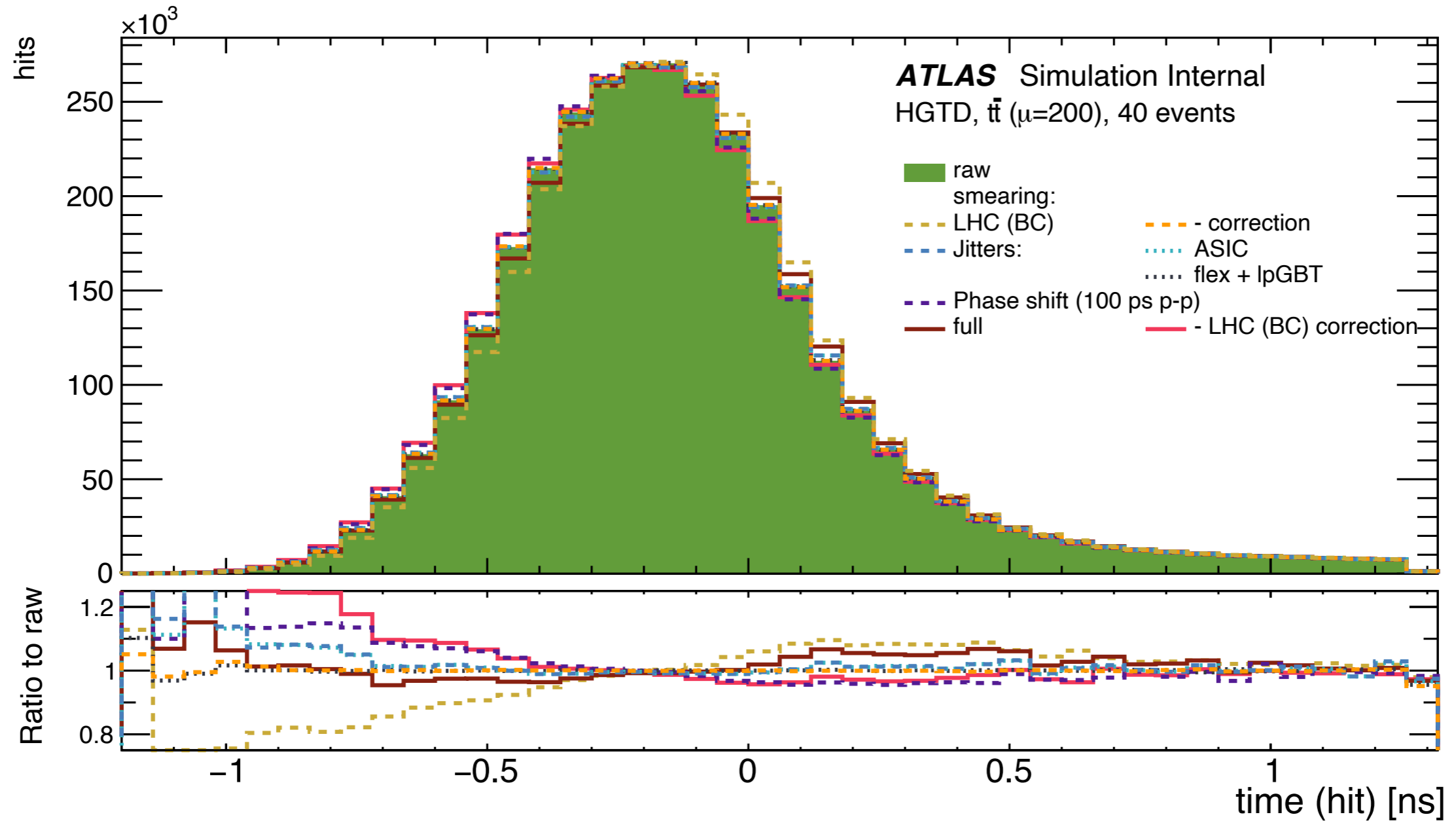
- One of the key question is the **mapping** of the effects:
  - it's not easy to get it clear when the project is still in the R&D phase.
- The **frequency** is also challenging:
  - Maybe we can rely on current detector experience ? But will LHC behave the same in High-lumi ?
- Is the **correlation** between the effects something we are sensitive to ?
- Which **figure of merit** to use:
  - The hit level is easy to extract but is it the most meaningful one ?
  - What about per track effects ?
- The **calibration** procedure also triggers a lot of decisions:
  - Offline vs Online ?
  - How often and what granularity ?
  - How well this mitigates the effects described earlier ?

**BACK-UP**

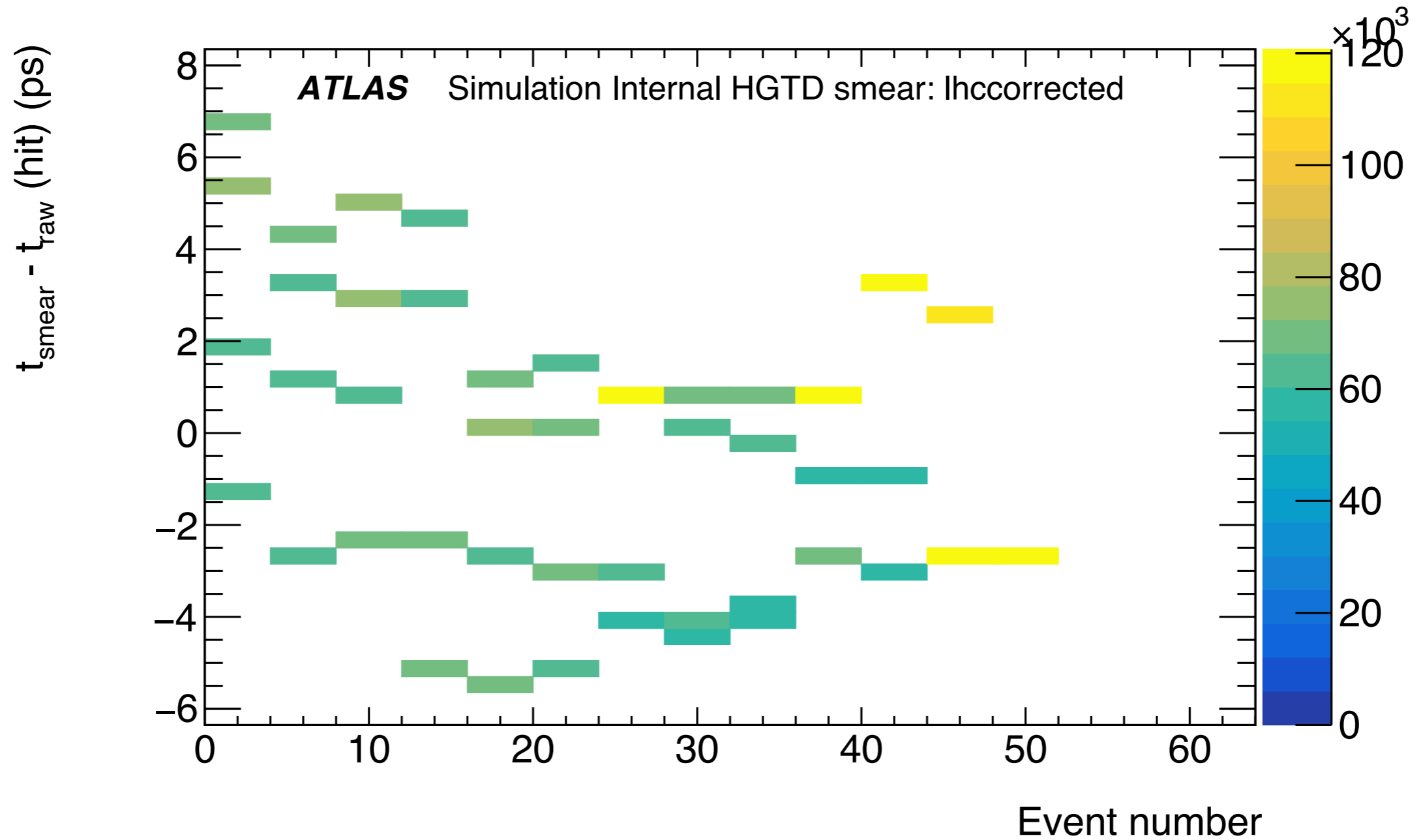
# Zoomed effects



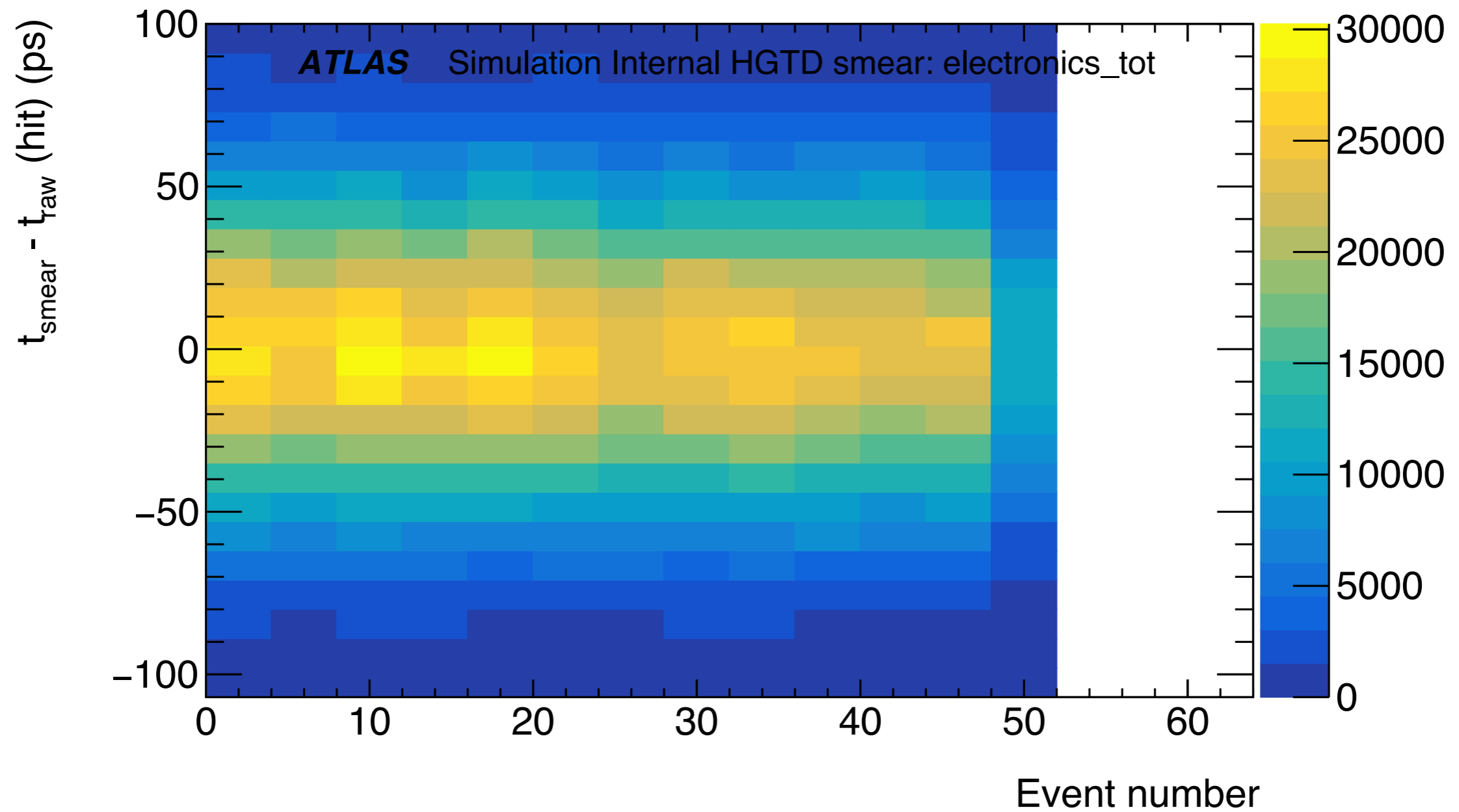
# 100ps Phase Shift



# Other plots

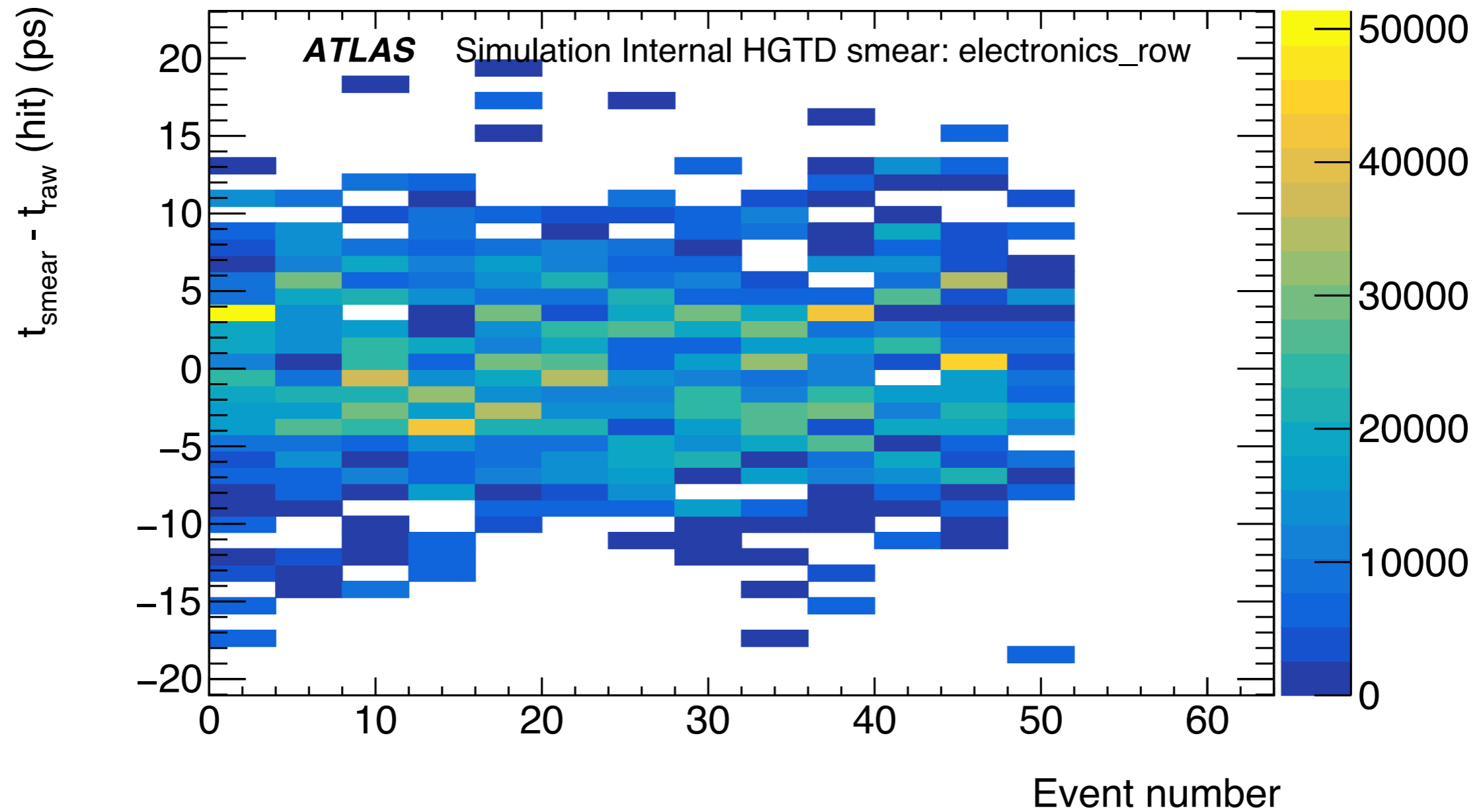


# Other plots

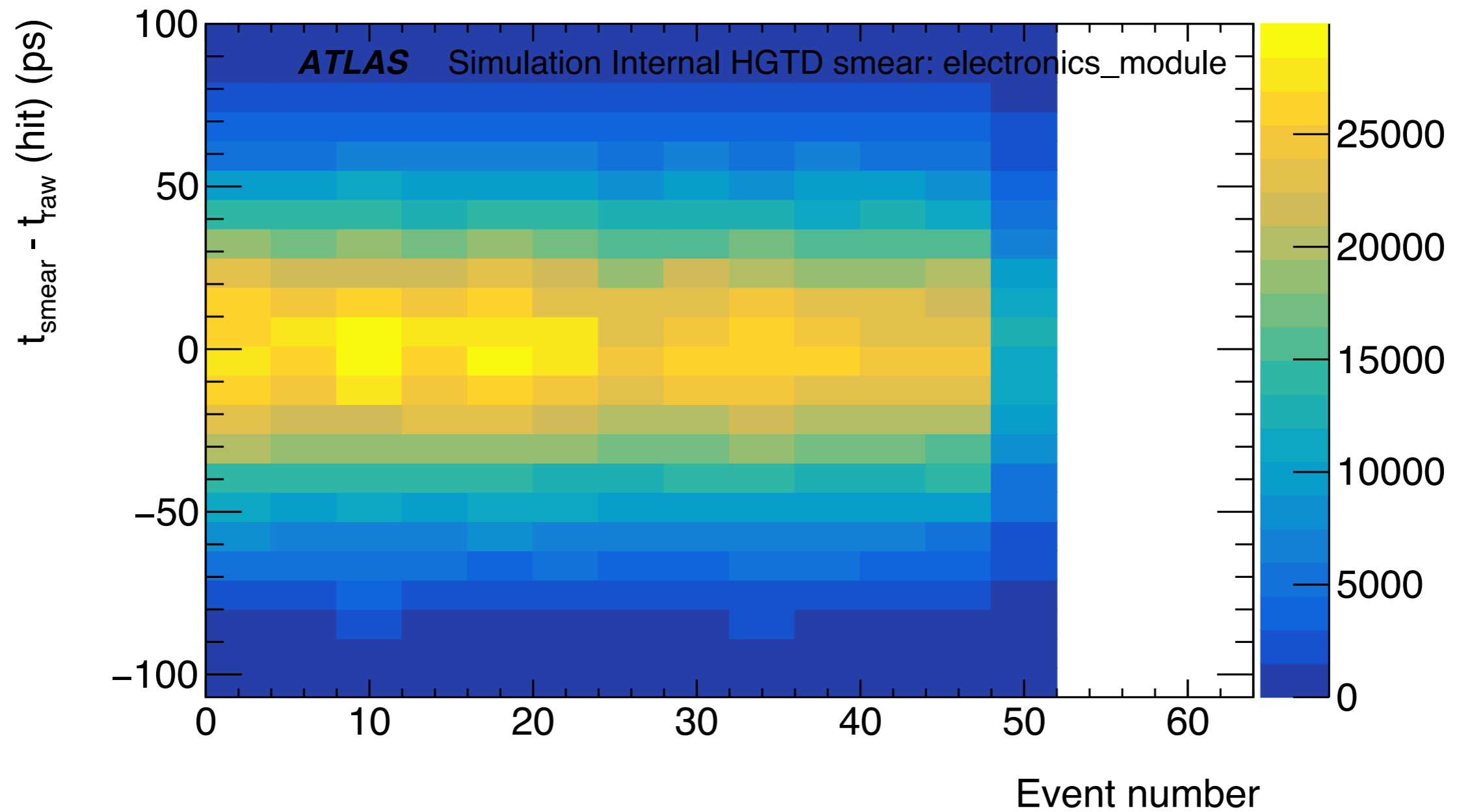




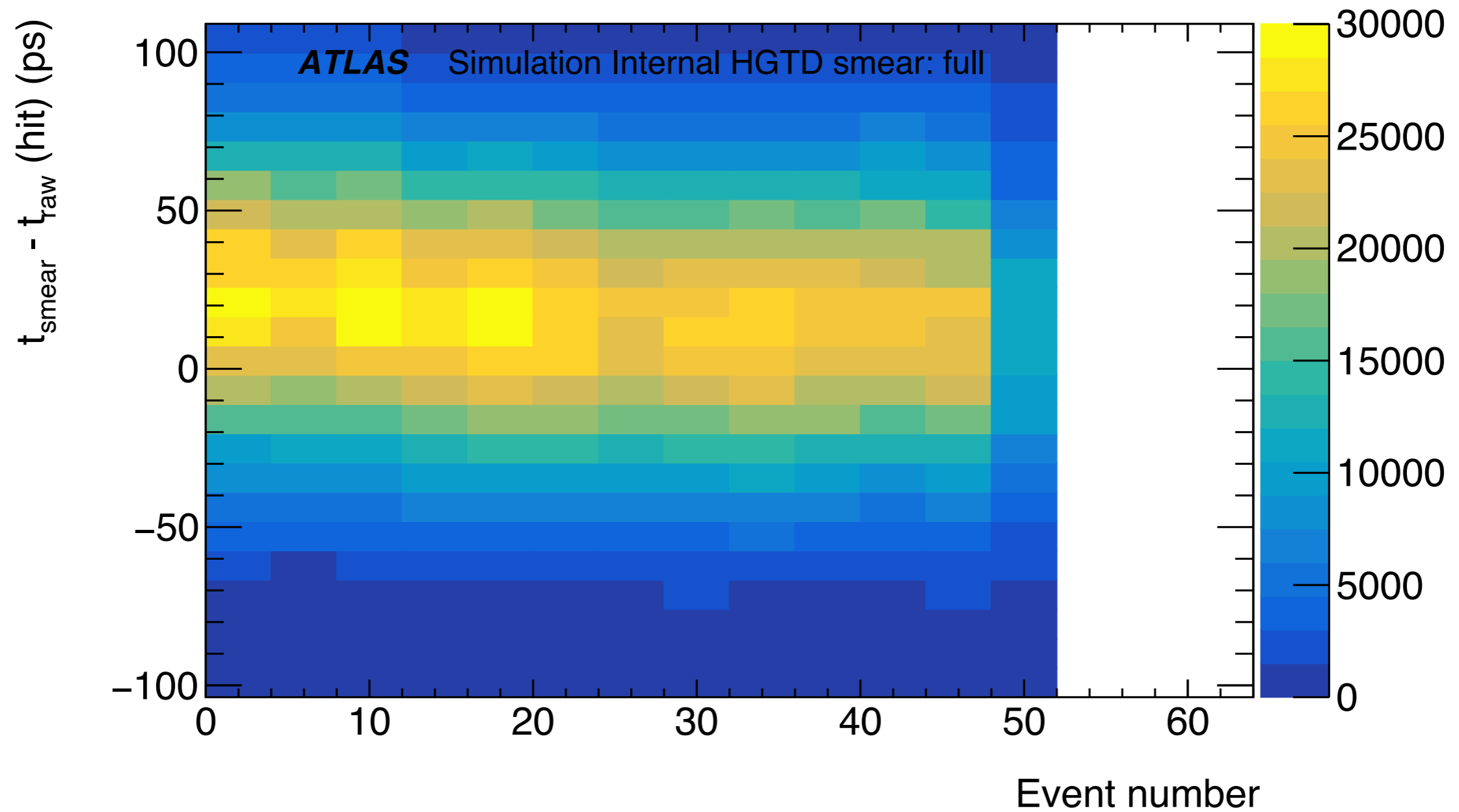
# Other plots



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# Other plots

