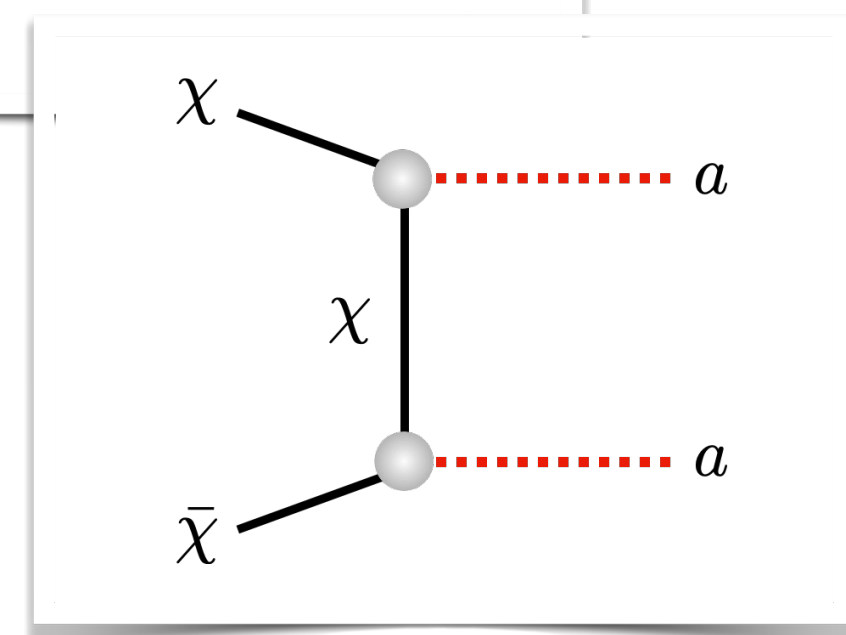
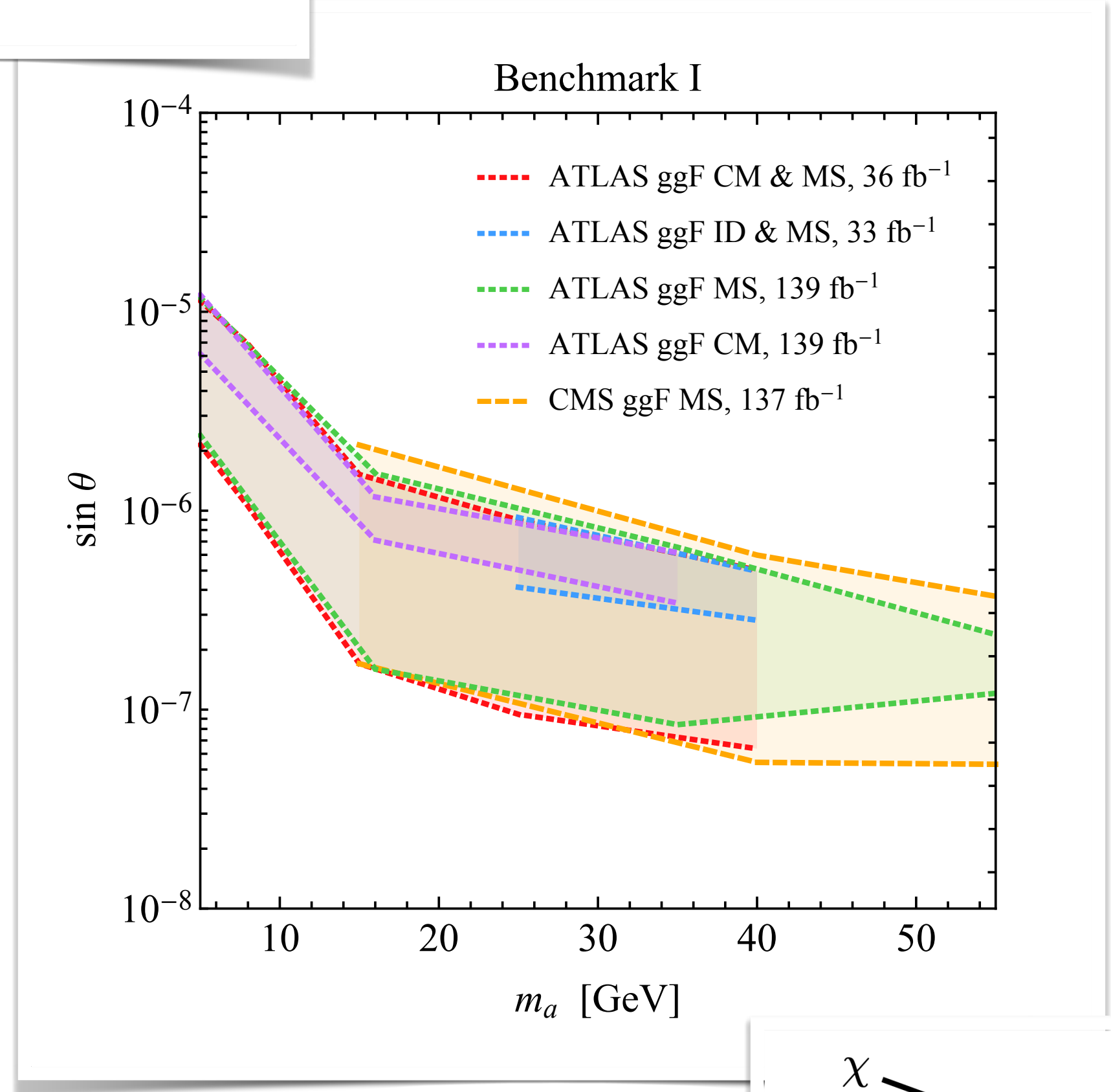
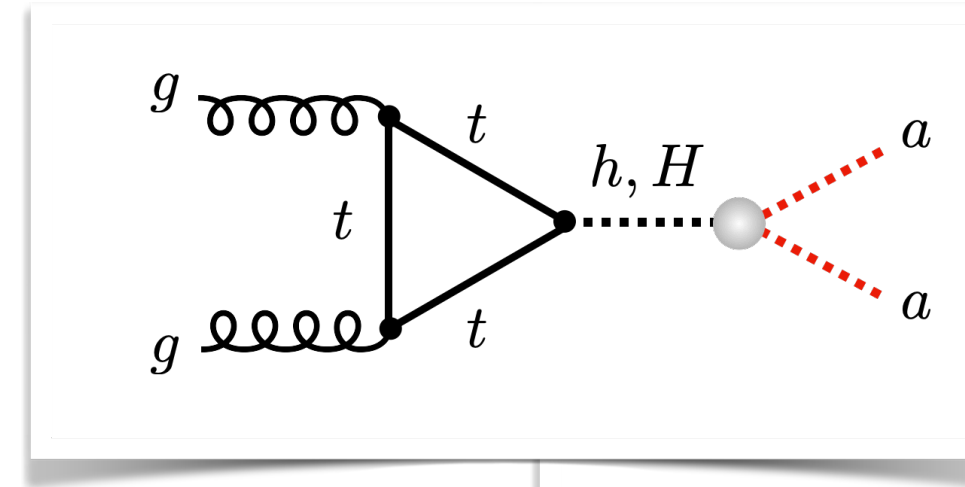


Long-lived particle phenomenology in the 2HDM+ a model

Luc Schnell

13th Workshop of the Long-Lived Particle Community, CERN
June 20, 2023



1. Introduction

1.1 Motivation

1.2 2HDM+ a in a nutshell

1.3 E_T^{miss} signatures

1. Introduction

1.1 Motivation

1. Introduction

1.1 Motivation

UV-complete DM benchmarks

1. Introduction

1.1 Motivation

UV-complete DM benchmarks

- **Dark matter (DM)** has become a prime target of BSM searches at the LHC.

1. Introduction

1.1 Motivation

UV-complete DM benchmarks

- **Dark matter (DM)** has become a prime target of BSM searches at the LHC.
- **Mediator models** are particularly relevant for colliders → how can we explore them systematically?

1. Introduction

1.1 Motivation

UV-complete DM benchmarks

- **Dark matter (DM)** has become a prime target of BSM searches at the LHC.
- **Mediator models** are particularly relevant for colliders → how can we explore them systematically?

Simplified models: e.g.

$$\mathcal{L} \supset -g_q^A Z'^\mu (\bar{q}\gamma_\mu\gamma^5 q) - g_{DM}^A Z'^\mu (\bar{\chi}\gamma_\mu\gamma^5 \chi)$$

1. Introduction

1.1 Motivation

UV-complete DM benchmarks

- **Dark matter (DM)** has become a prime target of BSM searches at the LHC.
- **Mediator models** are particularly relevant for colliders → how can we explore them systematically?

Simplified models: e.g.

$$\mathcal{L} \supset -g_q^A Z'^\mu (\bar{q}\gamma_\mu\gamma^5 q) - g_{DM}^A Z'^\mu (\bar{\chi}\gamma_\mu\gamma^5 \chi) \rightarrow \text{UV-complete benchmarks}$$

1. Introduction

1.1 Motivation

UV-complete DM benchmarks

- **Dark matter (DM)** has become a prime target of BSM searches at the LHC.
- **Mediator models** are particularly relevant for colliders → how can we explore them systematically?

Simplified models: e.g.

$$\mathcal{L} \supset -g_q^A Z'^\mu (\bar{q}\gamma_\mu\gamma^5 q) - g_{DM}^A Z'^\mu (\bar{\chi}\gamma_\mu\gamma^5 \chi) \rightarrow \text{UV-complete benchmarks}$$

Mixing with scalar sector

1. Introduction

1.1 Motivation

UV-complete DM benchmarks

- **Dark matter (DM)** has become a prime target of BSM searches at the LHC.
- **Mediator models** are particularly relevant for colliders → how can we explore them systematically?

Simplified models: e.g. $\mathcal{L} \supset -g_q^A Z'^\mu (\bar{q}\gamma_\mu\gamma^5 q) -g_{DM}^A Z'^\mu (\bar{\chi}\gamma_\mu\gamma^5 \chi) \rightarrow$ **UV-complete benchmarks**

Mixing with scalar sector

- **Mixing of a DM mediator with the (extended) scalar sector** leads to a rich and interesting collider phenomenology.

1. Introduction

1.1 Motivation

UV-complete DM benchmarks

- **Dark matter (DM)** has become a prime target of BSM searches at the LHC.
- **Mediator models** are particularly relevant for colliders → how can we explore them systematically?

Simplified models: e.g.

$$\mathcal{L} \supset -g_q^A Z'^\mu (\bar{q}\gamma_\mu\gamma^5 q) - g_{DM}^A Z'^\mu (\bar{\chi}\gamma_\mu\gamma^5 \chi) \rightarrow \text{UV-complete benchmarks}$$

2HDM+a
model

Mixing with scalar sector

- **Mixing of a DM mediator with the (extended) scalar sector** leads to a rich and interesting collider phenomenology.

1. Introduction

1.2 2HDM+a in a nutshell

1. Introduction

1.2 2HDM+a in a nutshell

**2HDM scalar
potential**

$$V_H = \mu_1 H_1^\dagger H_1 + \mu_2 H_2^\dagger H_2 + \left(\mu_3 H_1^\dagger H_2 + \text{h.c.} \right) + \lambda_1 (H_1^\dagger H_1)^2 + \lambda_2 (H_2^\dagger H_2)^2 \\ + \lambda_3 (H_1^\dagger H_1) (H_2^\dagger H_2) + \lambda_4 (H_1^\dagger H_2) (H_2^\dagger H_1) + \left[\lambda_5 (H_1^\dagger H_2)^2 + \text{h.c.} \right] .$$

1. Introduction

1.2 2HDM+a in a nutshell

2HDM scalar
potential

$$V_H = \mu_1 H_1^\dagger \overbrace{H_1}^{\text{HD 1}} + \mu_2 H_2^\dagger H_2 + \left(\mu_3 H_1^\dagger H_2 + \text{h.c.} \right) + \lambda_1 (H_1^\dagger H_1)^2 + \lambda_2 (H_2^\dagger H_2)^2 \\ + \lambda_3 (H_1^\dagger H_1) (H_2^\dagger H_2) + \lambda_4 (H_1^\dagger H_2) (H_2^\dagger H_1) + \left[\lambda_5 (H_1^\dagger H_2)^2 + \text{h.c.} \right] .$$

1. Introduction

1.2 2HDM+a in a nutshell

2HDM scalar
potential

$$V_H = \mu_1 H_1^\dagger \overbrace{H_1}^{\text{HD 1}} + \mu_2 H_2^\dagger \overbrace{H_2}^{\text{HD 2}} + \left(\mu_3 H_1^\dagger H_2 + \text{h.c.} \right) + \lambda_1 (H_1^\dagger H_1)^2 + \lambda_2 (H_2^\dagger H_2)^2 \\ + \lambda_3 (H_1^\dagger H_1) (H_2^\dagger H_2) + \lambda_4 (H_1^\dagger H_2) (H_2^\dagger H_1) + \left[\lambda_5 (H_1^\dagger H_2)^2 + \text{h.c.} \right] .$$

1. Introduction

1.2 2HDM+a in a nutshell

2HDM scalar
potential

$$V_H = \mu_1 H_1^\dagger \boxed{H_1} + \mu_2 H_2^\dagger \boxed{H_2} + \left(\mu_3 H_1^\dagger H_2 + \text{h.c.} \right) + \lambda_1 (H_1^\dagger H_1)^2 + \lambda_2 (H_2^\dagger H_2)^2 \\ + \lambda_3 (H_1^\dagger H_1)(H_2^\dagger H_2) + \lambda_4 (H_1^\dagger H_2)(H_2^\dagger H_1) + \left[\lambda_5 (H_1^\dagger H_2)^2 + \text{h.c.} \right].$$

Pseudo-
scalar mediator

$$V_P = \frac{1}{2} m_P^2 P^2 + P \left(i b_P H_1^\dagger H_2 + \text{h.c.} \right) + P^2 \left(\lambda_{P1} H_1^\dagger H_1 + \lambda_{P2} H_2^\dagger H_2 \right),$$

Source: [ArXiv:1701.07427](https://arxiv.org/abs/1701.07427) (M. Bauer, U. Haisch and F. Kahlhoefer).

1. Introduction

1.2 2HDM+a in a nutshell

2HDM scalar
potential

$$V_H = \mu_1 H_1^\dagger \boxed{H_1} + \mu_2 H_2^\dagger \boxed{H_2} + \left(\mu_3 H_1^\dagger H_2 + \text{h.c.} \right) + \lambda_1 (H_1^\dagger H_1)^2 + \lambda_2 (H_2^\dagger H_2)^2 \\ + \lambda_3 (H_1^\dagger H_1)(H_2^\dagger H_2) + \lambda_4 (H_1^\dagger H_2)(H_2^\dagger H_1) + \left[\lambda_5 (H_1^\dagger H_2)^2 + \text{h.c.} \right].$$

Pseudo-
scalar mediator

$$V_P = \frac{1}{2} m_P^2 P^2 + \boxed{P} \left(i b_P H_1^\dagger H_2 + \text{h.c.} \right) + P^2 \left(\lambda_{P1} H_1^\dagger H_1 + \lambda_{P2} H_2^\dagger H_2 \right),$$

Source: [ArXiv:1701.07427](https://arxiv.org/abs/1701.07427) (M. Bauer, U. Haisch and F. Kahlhoefer).

1. Introduction

1.2 2HDM+a in a nutshell

2HDM scalar
potential

$$V_H = \mu_1 H_1^\dagger \boxed{H_1} + \mu_2 H_2^\dagger \boxed{H_2} + \left(\mu_3 H_1^\dagger H_2 + \text{h.c.} \right) + \lambda_1 (H_1^\dagger H_1)^2 + \lambda_2 (H_2^\dagger H_2)^2 \\ + \lambda_3 (H_1^\dagger H_1)(H_2^\dagger H_2) + \lambda_4 (H_1^\dagger H_2)(H_2^\dagger H_1) + \left[\lambda_5 (H_1^\dagger H_2)^2 + \text{h.c.} \right].$$

Pseudo-
scalar mediator

$$V_P = \frac{1}{2} m_P^2 P^2 + \boxed{P} \left(i b_P H_1^\dagger H_2 + \text{h.c.} \right) + P^2 \left(\lambda_{P1} H_1^\dagger H_1 + \lambda_{P2} H_2^\dagger H_2 \right),$$

$$\mathcal{L}_\chi = -i y_\chi P \bar{\chi} \gamma_5 \chi,$$

Fermionic
DM

Source: [ArXiv:1701.07427](https://arxiv.org/abs/1701.07427) (M. Bauer, U. Haisch and F. Kahlhoefer).

1. Introduction

1.2 2HDM+a in a nutshell

2HDM scalar
potential

$$V_H = \mu_1 H_1^\dagger \boxed{H_1} + \mu_2 H_2^\dagger \boxed{H_2} + \left(\mu_3 H_1^\dagger H_2 + \text{h.c.} \right) + \lambda_1 (H_1^\dagger H_1)^2 + \lambda_2 (H_2^\dagger H_2)^2 \\ + \lambda_3 (H_1^\dagger H_1)(H_2^\dagger H_2) + \lambda_4 (H_1^\dagger H_2)(H_2^\dagger H_1) + \left[\lambda_5 (H_1^\dagger H_2)^2 + \text{h.c.} \right].$$

Pseudo-
scalar mediator

$$V_P = \frac{1}{2} m_P^2 P^2 + \boxed{P} \left(i b_P H_1^\dagger H_2 + \text{h.c.} \right) + P^2 \left(\lambda_{P1} H_1^\dagger H_1 + \lambda_{P2} H_2^\dagger H_2 \right),$$

$$\mathcal{L}_\chi = -i y_\chi P \bar{\chi} \gamma_5 \boxed{\chi},$$

Fermionic
DM

Source: [ArXiv:1701.07427](https://arxiv.org/abs/1701.07427) (M. Bauer, U. Haisch and F. Kahlhoefer).

1. Introduction

1.2 2HDM+a in a nutshell

2HDM scalar
potential

$$V_H = \mu_1 H_1^\dagger \boxed{H_1} + \mu_2 H_2^\dagger \boxed{H_2} + \left(\mu_3 H_1^\dagger H_2 + \text{h.c.} \right) + \lambda_1 (H_1^\dagger H_1)^2 + \lambda_2 (H_2^\dagger H_2)^2 \\ + \lambda_3 (H_1^\dagger H_1)(H_2^\dagger H_2) + \lambda_4 (H_1^\dagger H_2)(H_2^\dagger H_1) + \left[\lambda_5 (H_1^\dagger H_2)^2 + \text{h.c.} \right].$$

Pseudo-
scalar mediator

$$V_P = \frac{1}{2} m_P^2 P^2 + \boxed{P} \left(i b_P H_1^\dagger H_2 + \text{h.c.} \right) + P^2 \left(\lambda_{P1} H_1^\dagger H_1 + \lambda_{P2} H_2^\dagger H_2 \right),$$

$$\mathcal{L}_\chi = -i y_\chi \boxed{P} \bar{\chi} \gamma_5 \boxed{\chi},$$

Fermionic
DM

Source: [ArXiv:1701.07427](https://arxiv.org/abs/1701.07427) (M. Bauer, U. Haisch and F. Kahlhoefer).

1. Introduction

1.2 2HDM+a in a nutshell

2HDM scalar potential

$$V_H = \mu_1 H_1^\dagger \boxed{H_1} + \mu_2 H_2^\dagger \boxed{H_2} + \left(\mu_3 H_1^\dagger H_2 + \text{h.c.} \right) + \lambda_1 (H_1^\dagger H_1)^2 + \lambda_2 (H_2^\dagger H_2)^2 \\ + \lambda_3 (H_1^\dagger H_1)(H_2^\dagger H_2) + \lambda_4 (H_1^\dagger H_2)(H_2^\dagger H_1) + \left[\lambda_5 (H_1^\dagger H_2)^2 + \text{h.c.} \right].$$

pseudoscalar coupling

Pseudo-scalar mediator

$$V_P = \frac{1}{2} m_P^2 P^2 + \boxed{P} \left(i b_P H_1^\dagger H_2 + \text{h.c.} \right) + P^2 \left(\lambda_{P1} H_1^\dagger H_1 + \lambda_{P2} H_2^\dagger H_2 \right),$$

$$\mathcal{L}_\chi = -i y_\chi \boxed{P} \bar{\chi} \gamma_5 \boxed{\chi},$$

Fermionic DM

Source: [ArXiv:1701.07427](https://arxiv.org/abs/1701.07427) (M. Bauer, U. Haisch and F. Kahlhoefer).

1. Introduction

1.2 2HDM+a in a nutshell

2HDM scalar potential

$$V_H = \mu_1 H_1^\dagger \boxed{H_1} + \mu_2 H_2^\dagger \boxed{H_2} + \left(\mu_3 H_1^\dagger H_2 + \text{h.c.} \right) + \lambda_1 (H_1^\dagger H_1)^2 + \lambda_2 (H_2^\dagger H_2)^2 \\ + \lambda_3 (H_1^\dagger H_1)(H_2^\dagger H_2) + \lambda_4 (H_1^\dagger H_2)(H_2^\dagger H_1) + \left[\lambda_5 (H_1^\dagger H_2)^2 + \text{h.c.} \right].$$

pseudoscalar coupling

Pseudo-scalar mediator

$$V_P = \frac{1}{2} m_P^2 P^2 + \boxed{P} \left(i b_P H_1^\dagger H_2 + \text{h.c.} \right) + P^2 \left(\lambda_{P1} H_1^\dagger H_1 + \lambda_{P2} H_2^\dagger H_2 \right),$$

$$\mathcal{L}_\chi = -i y_\chi \boxed{P} \bar{\chi} \gamma_5 \boxed{\chi},$$

Fermionic DM

Source: [ArXiv:1701.07427](https://arxiv.org/abs/1701.07427) (M. Bauer, U. Haisch and F. Kahlhoefer).



SSB:

$$\langle H_i \rangle = (0, v_i/\sqrt{2})^T \text{ with } v = \sqrt{v_1^2 + v_2^2} \simeq 246 \text{ GeV}$$

1. Introduction

1.2 2HDM+a in a nutshell

2HDM scalar potential

$$V_H = \mu_1 H_1^\dagger \boxed{H_1} + \mu_2 H_2^\dagger \boxed{H_2} + \left(\mu_3 H_1^\dagger H_2 + \text{h.c.} \right) + \lambda_1 (H_1^\dagger H_1)^2 + \lambda_2 (H_2^\dagger H_2)^2 \\ + \lambda_3 (H_1^\dagger H_1)(H_2^\dagger H_2) + \lambda_4 (H_1^\dagger H_2)(H_2^\dagger H_1) + \left[\lambda_5 (H_1^\dagger H_2)^2 + \text{h.c.} \right].$$

pseudoscalar coupling

Pseudo-scalar mediator

$$V_P = \frac{1}{2} m_P^2 P^2 + \boxed{P} \left(i b_P H_1^\dagger H_2 + \text{h.c.} \right) + P^2 \left(\lambda_{P1} H_1^\dagger H_1 + \lambda_{P2} H_2^\dagger H_2 \right),$$

$$\mathcal{L}_\chi = -i y_\chi \boxed{P} \bar{\chi} \gamma_5 \boxed{\chi},$$

Fermionic DM

Source: [ArXiv:1701.07427](https://arxiv.org/abs/1701.07427) (M. Bauer, U. Haisch and F. Kahlhoefer).



SSB:

$$\langle H_i \rangle = (0, v_i/\sqrt{2})^T \text{ with } v = \sqrt{v_1^2 + v_2^2} \simeq 246 \text{ GeV}$$

Physical fields: h, H, a, A, H^\pm (three d.o.f. eaten by Z, W^\pm) and χ .

Physical parameters: $\alpha, \beta, \theta, v, \lambda_3, \lambda_{P1}, \lambda_{P2}, m_h, m_H, m_a, m_A, m_{H^\pm}$.

α : mixing angle for scalars (h, H)

θ : mixing angle for pseudo-scalars (a, A)

1. Introduction

1.2 2HDM+a in a nutshell

2HDM scalar potential

$$V_H = \mu_1 H_1^\dagger \boxed{H_1} + \mu_2 H_2^\dagger \boxed{H_2} + (\mu_3 H_1^\dagger H_2 + \text{h.c.}) + \lambda_1 (H_1^\dagger H_1)^2 + \lambda_2 (H_2^\dagger H_2)^2 + \lambda_3 (H_1^\dagger H_1)(H_2^\dagger H_2) + \lambda_4 (H_1^\dagger H_2)(H_2^\dagger H_1) + [\lambda_5 (H_1^\dagger H_2)^2 + \text{h.c.}] .$$

pseudoscalar coupling

Pseudo-scalar mediator

$$V_P = \frac{1}{2} m_P^2 P^2 + \boxed{P} (ib_P H_1^\dagger H_2 + \text{h.c.}) + P^2 (\lambda_{P1} H_1^\dagger H_1 + \lambda_{P2} H_2^\dagger H_2) ,$$

$$\mathcal{L}_\chi = -iy_\chi \boxed{P} \bar{\chi} \gamma_5 \chi ,$$

Fermionic DM

Source: [ArXiv:1701.07427](https://arxiv.org/abs/1701.07427) (M. Bauer, U. Haisch and F. Kahlhoefer).



SSB:

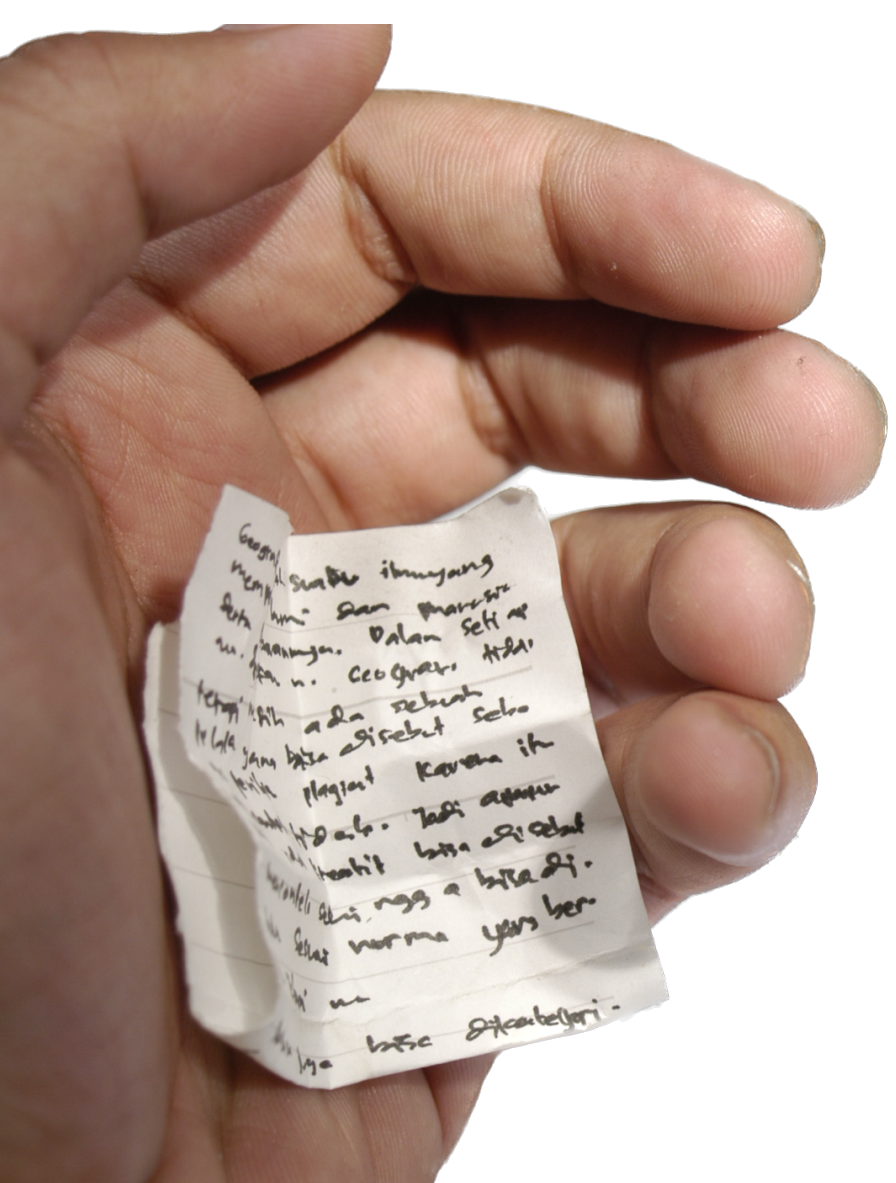
$$\langle H_i \rangle = (0, v_i/\sqrt{2})^T \text{ with } v = \sqrt{v_1^2 + v_2^2} \simeq 246 \text{ GeV}$$

Physical fields: h, H, a, A, H^\pm (three d.o.f. eaten by Z, W^\pm) and χ .

Physical parameters: $\alpha, \beta, \theta, v, \lambda_3, \lambda_{P1}, \lambda_{P2}, m_h, m_H, m_a, m_A, m_{H^\pm}$.

α : mixing angle for scalars (h, H)

θ : mixing angle for pseudo-scalars (a, A)



1. Introduction

1.3 E_T^{miss} signatures

Physical fields: h, H, a, A, H^\pm and χ .

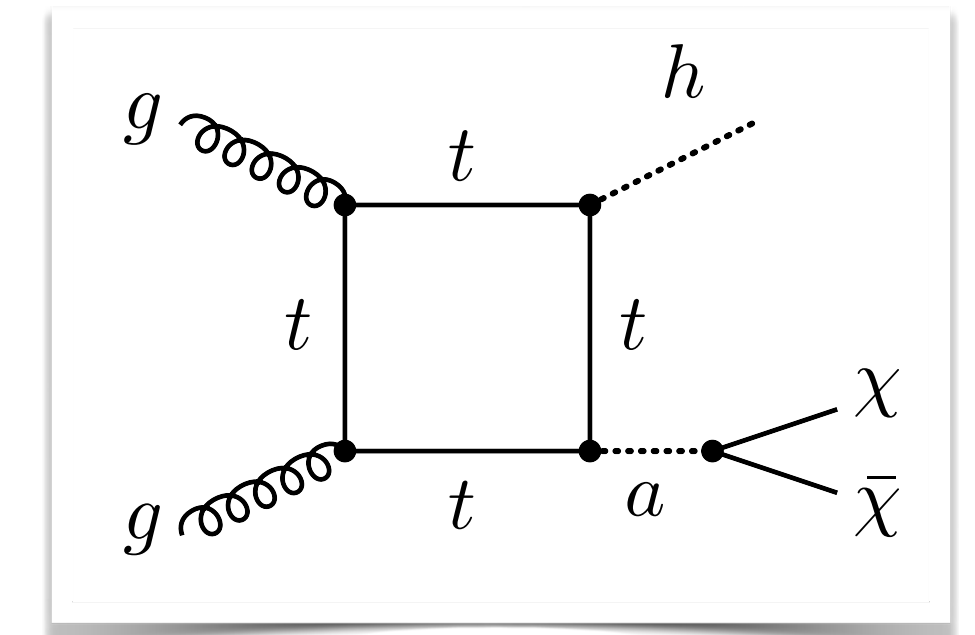
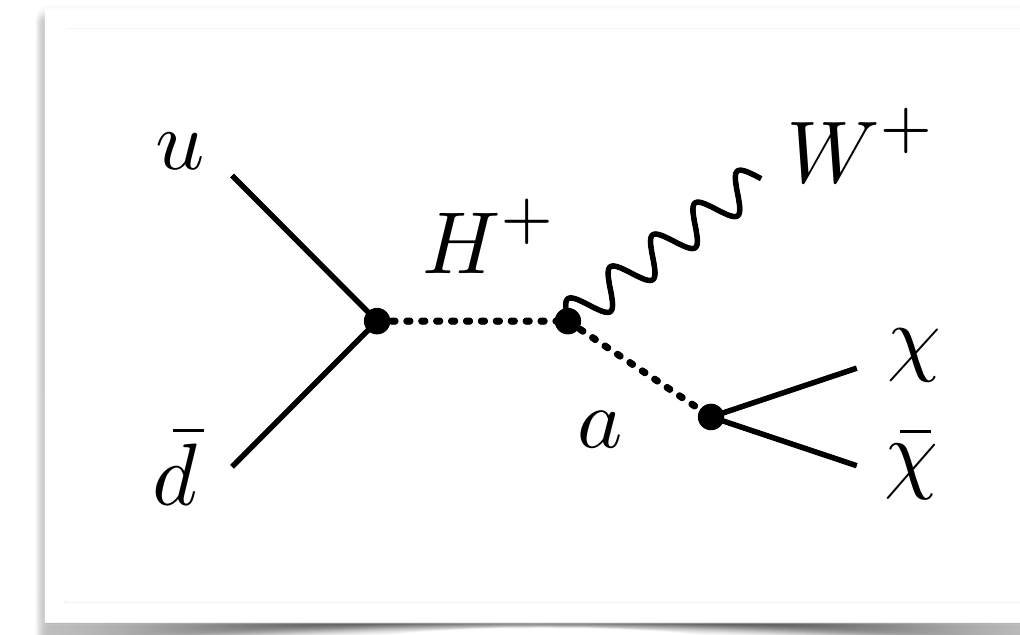
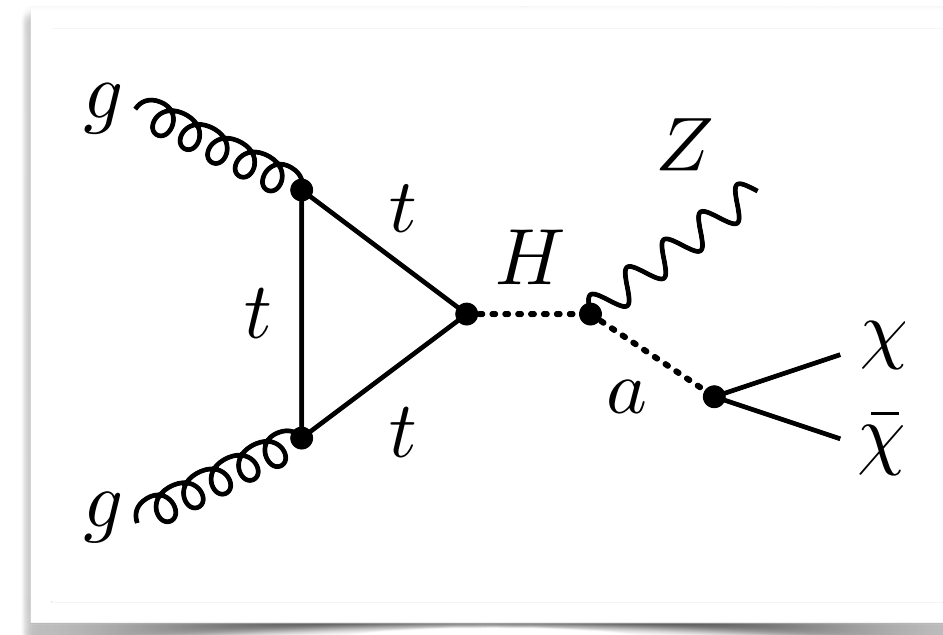
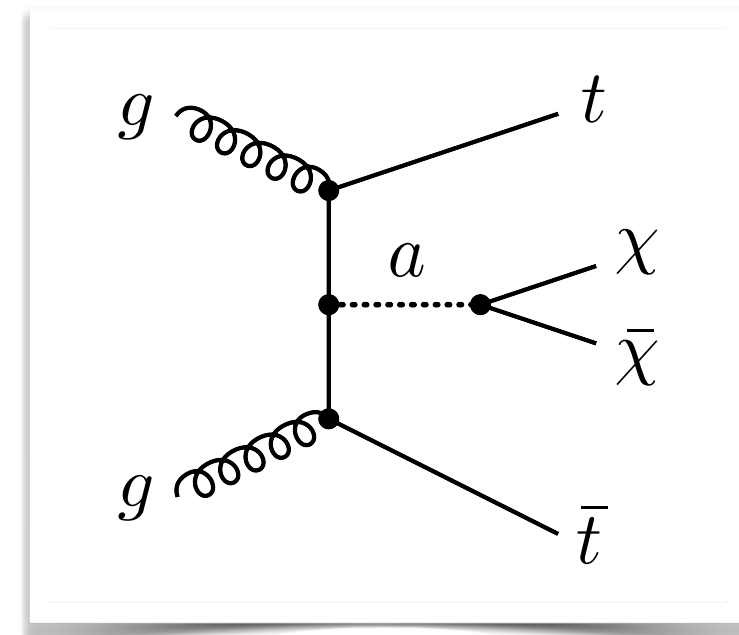
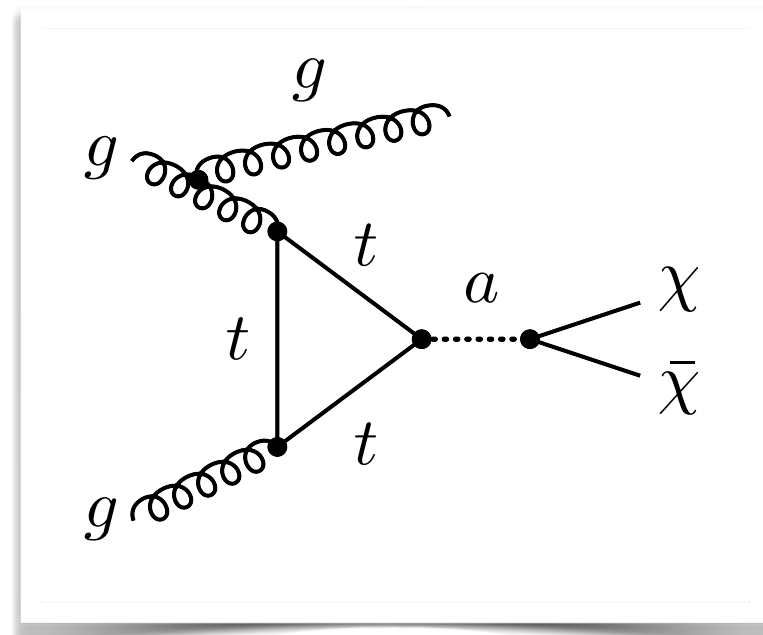
Physical parameters: $\alpha, \beta, \theta, \nu, \lambda_3, \lambda_{P1}, \lambda_{P2}, m_h, m_H, m_a, m_A, m_{H^\pm}$.

α : mixing angle for scalars (h, H)

θ : mixing angle for pseudo-scalars (a, A)

1. Introduction

1.3 E_T^{miss} signatures



Source: [ArXiv:1701.07427](https://arxiv.org/abs/1701.07427) (M. Bauer, U. Haisch and F. Kahlhoefer).

$$\mathbf{j} + E_T^{\text{miss}}$$

$$\mathbf{t\bar{t}} + E_T^{\text{miss}}$$

$$\mathbf{Z} + E_T^{\text{miss}}$$

$$\mathbf{W} + E_T^{\text{miss}}$$

$$\mathbf{h} + E_T^{\text{miss}}$$

Physical fields: h, H, a, A, H^\pm and χ .

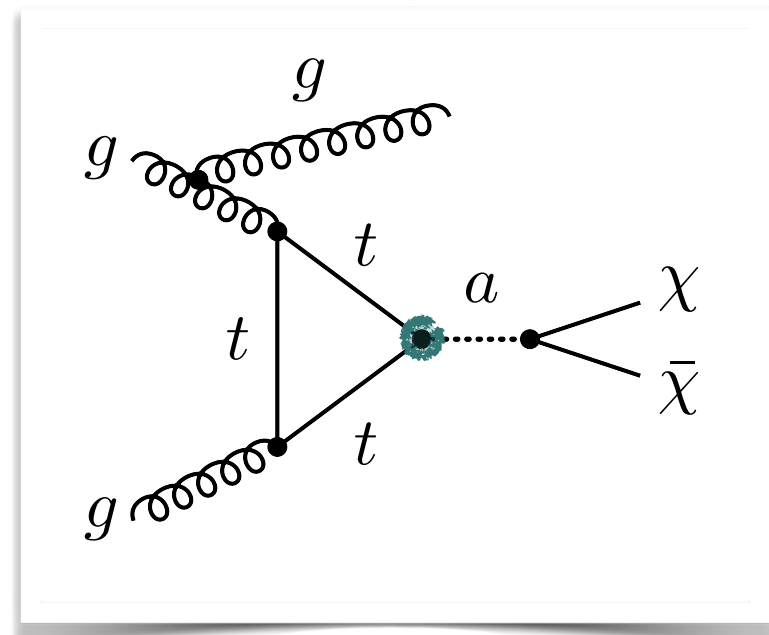
Physical parameters: $\alpha, \beta, \theta, v, \lambda_3, \lambda_{P1}, \lambda_{P2}, m_h, m_H, m_a, m_A, m_{H^\pm}$.

α : mixing angle for scalars (h, H)

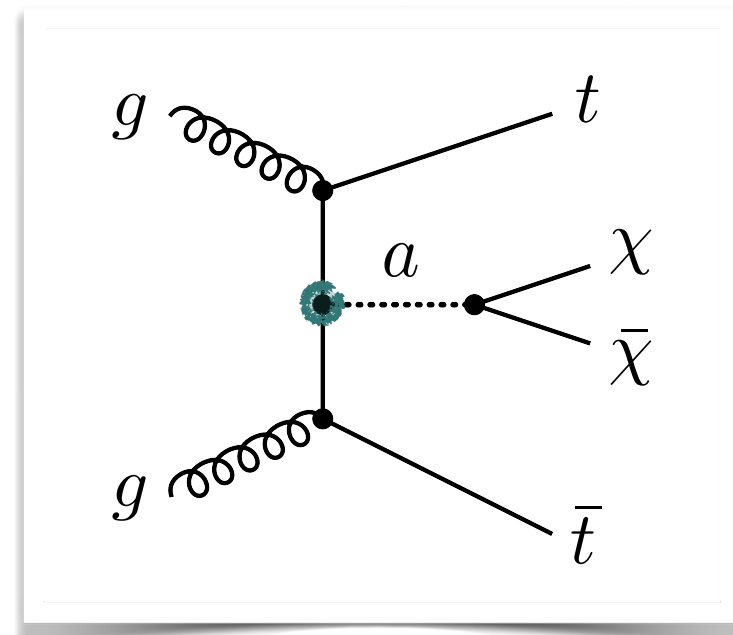
θ : mixing angle for pseudo-scalars (a, A)

1. Introduction

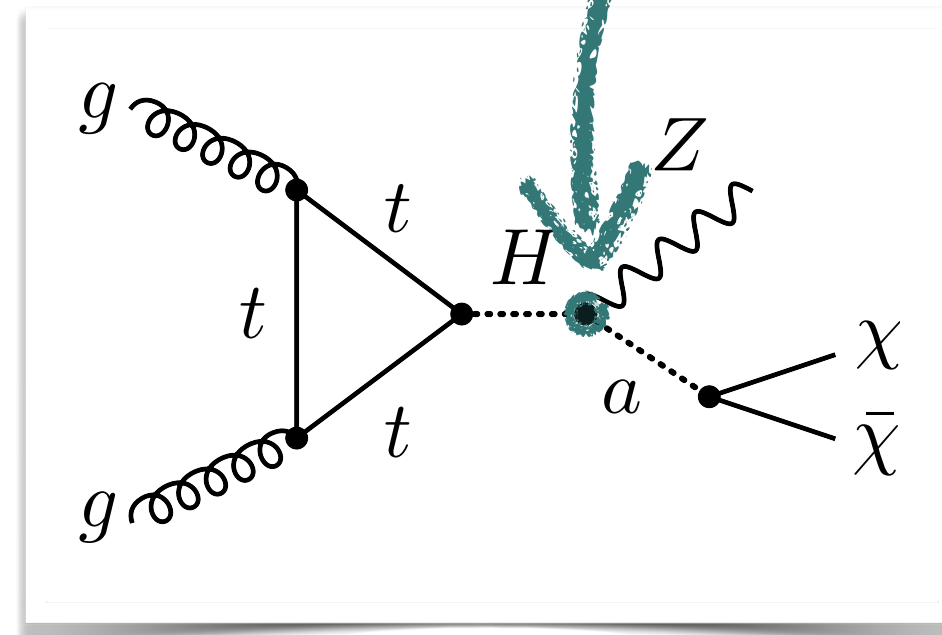
1.3 E_T^{miss} signatures



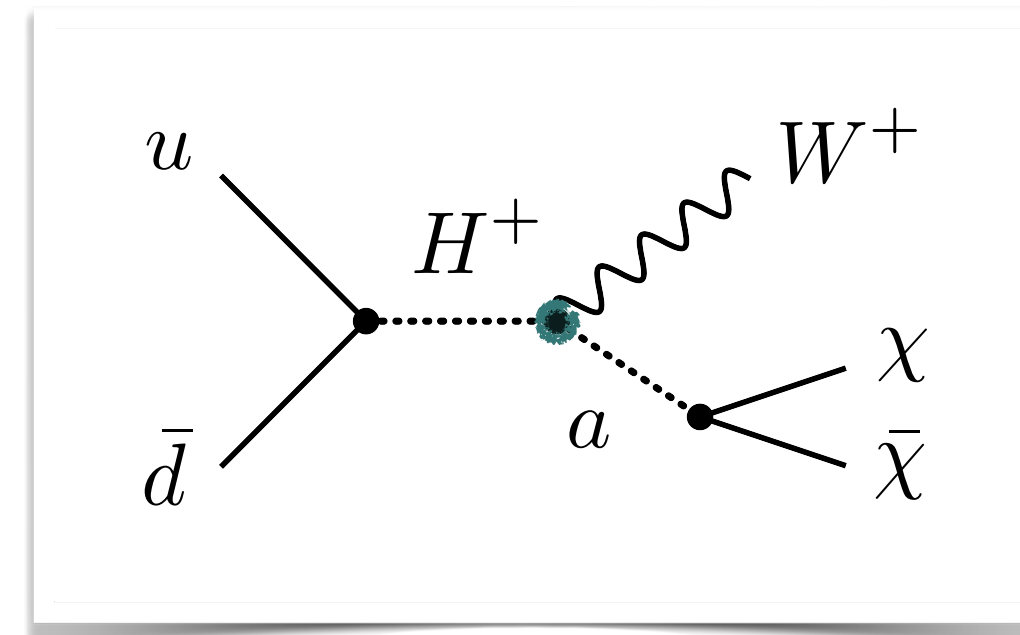
$$\mathbf{j} + E_T^{\text{miss}}$$



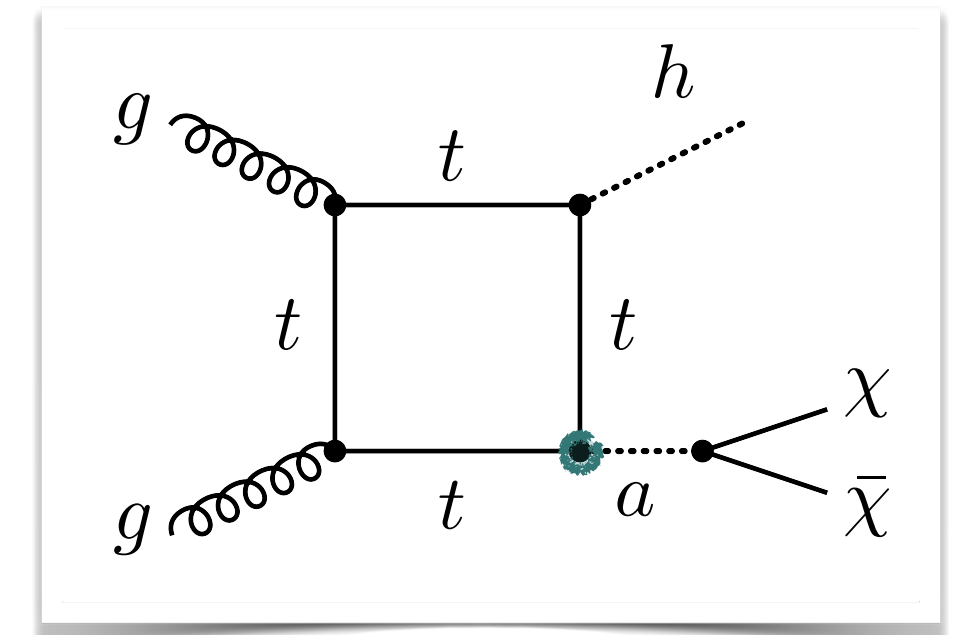
$$\mathbf{t\bar{t}} + E_T^{\text{miss}}$$



$$\mathbf{Z} + E_T^{\text{miss}}$$



$$\mathbf{W} + E_T^{\text{miss}}$$



$$\mathbf{h} + E_T^{\text{miss}}$$

Source: [ArXiv:1701.07427](https://arxiv.org/abs/1701.07427) (M. Bauer, U. Haisch and F. Kahlhoefer).

Physical fields: h, H, a, A, H^\pm and χ .

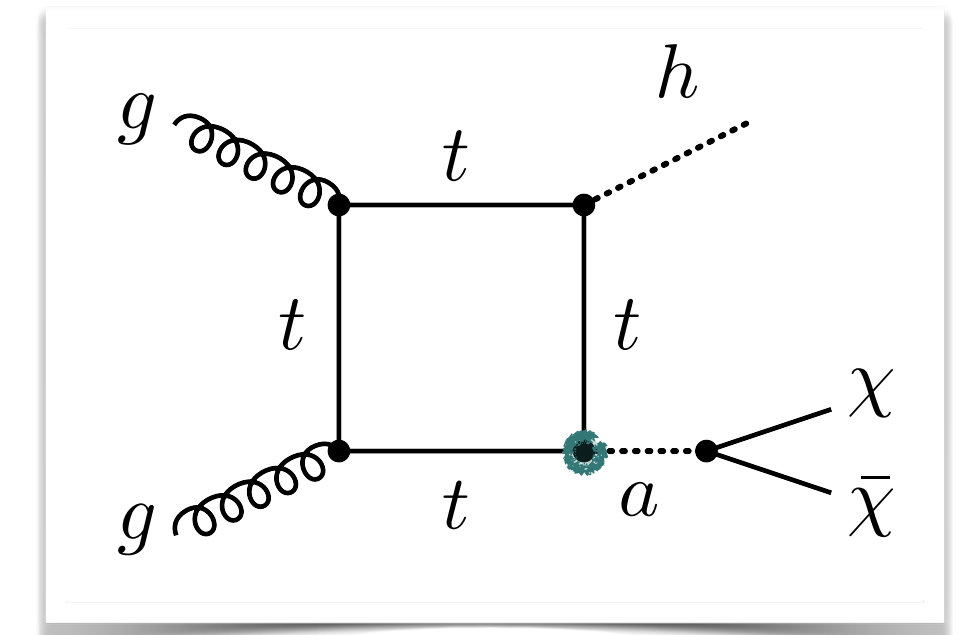
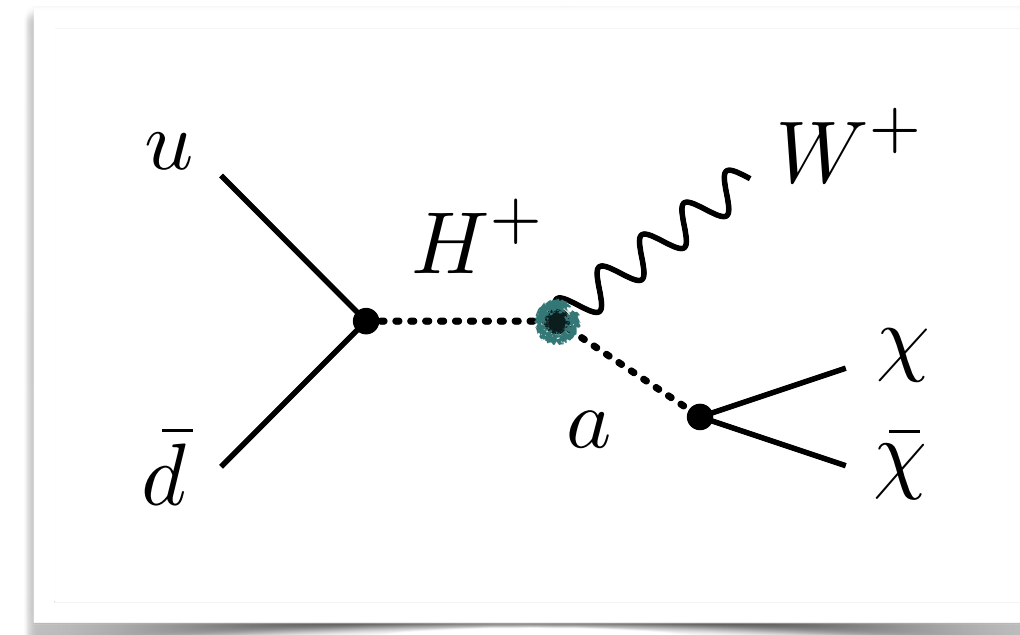
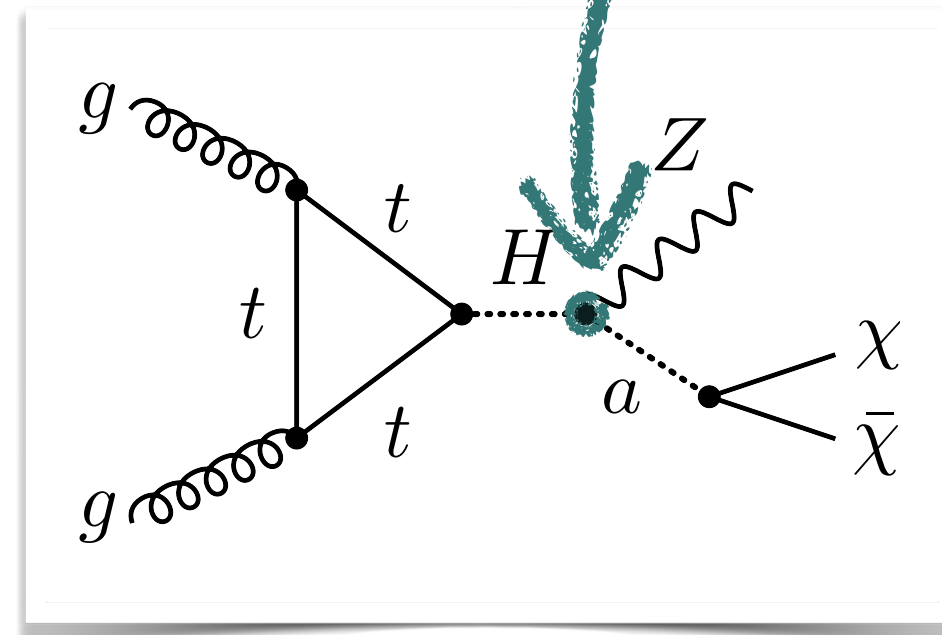
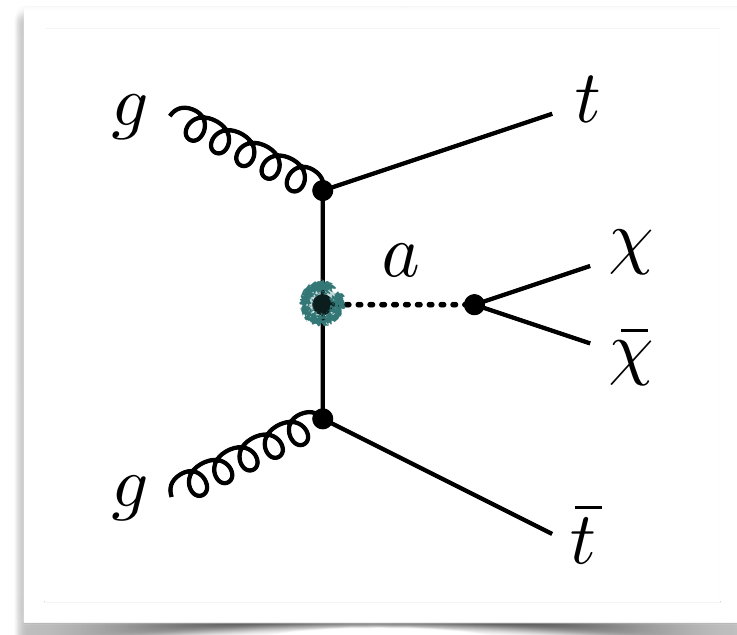
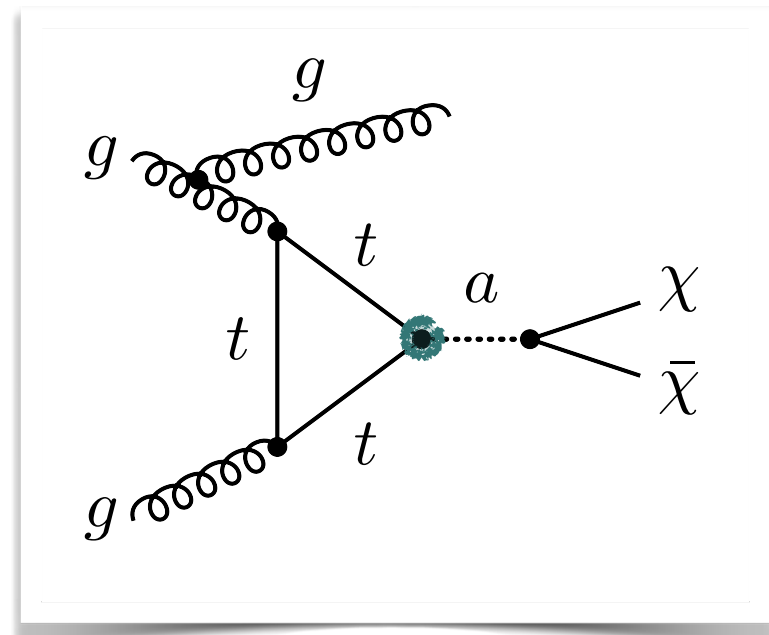
Physical parameters: $\alpha, \beta, \theta, v, \lambda_3, \lambda_{P1}, \lambda_{P2}, m_h, m_H, m_a, m_A, m_{H^\pm}$.

α : mixing angle for scalars (h, H)

θ : mixing angle for pseudo-scalars (a, A)

1. Introduction

1.3 E_T^{miss} signatures



Source: [ArXiv:1701.07427](https://arxiv.org/abs/1701.07427) (M. Bauer, U. Haisch and F. Kahlhoefer).

$$\mathbf{j} + E_T^{\text{miss}}$$

$$\mathbf{t\bar{t}} + E_T^{\text{miss}}$$

$$\mathbf{Z} + E_T^{\text{miss}}$$

$$\mathbf{W} + E_T^{\text{miss}}$$

$$\mathbf{h} + E_T^{\text{miss}}$$

- These E_T^{miss} signatures disappear for small mixing angles $\theta \simeq 0$ ($\rightarrow a \simeq P$).

Physical fields: h, H, a, A, H^\pm and χ .

Physical parameters: $\alpha, \beta, \theta, v, \lambda_3, \lambda_{P1}, \lambda_{P2}, m_h, m_H, m_a, m_A, m_{H^\pm}$.

α : mixing angle for scalars (h, H)

θ : mixing angle for pseudo-scalars (a, A)

2. LLP Phenomenology

2.1 Model parameters

2.2 LLP constraints

2.3 Relic density

2. LLP Phenomenology

2.1 Model parameters

Physical fields: h, H, a, A, H^\pm and χ .

Physical parameters: $\alpha, \beta, \theta, \nu, \lambda_3, \lambda_{P1}, \lambda_{P2}, m_h, m_H, m_a, m_A, m_{H^\pm}$.

α : mixing angle for scalars (h, H)

θ : mixing angle for pseudo-scalars (a, A)

2. LLP Phenomenology

2.1 Model parameters

- For small mixing angles θ , the pseudoscalar $a \simeq P$ can become **long-lived**.

Physical fields: h, H, a, A, H^\pm and χ .

Physical parameters: $\alpha, \beta, \theta, \nu, \lambda_3, \lambda_{P1}, \lambda_{P2}, m_h, m_H, m_a, m_A, m_{H^\pm}$.

α : mixing angle for scalars (h, H)

θ : mixing angle for pseudo-scalars (a, A)

2. LLP Phenomenology

2.1 Model parameters

- For small mixing angles θ , the pseudoscalar $a \simeq P$ can become **long-lived**.

$$\Gamma(a \rightarrow f\bar{f}) = \frac{N_c^f \eta_f^2 y_f^2}{16\pi} m_a \sqrt{1 - \frac{4m_f^2}{m_a^2}} \sin^2 \theta.$$

$$\Gamma(a \rightarrow gg) = \frac{\alpha_s^2}{32\pi^3 v^2} m_a^3 \left| \sum_{q=t,b,c} \eta_q f \left(\frac{4m_q^2}{m_a^2} \right) \right|^2 \sin^2 \theta,$$

$$\Gamma(a \rightarrow \chi\bar{\chi}) = \frac{y_\chi^2}{8\pi} m_a \sqrt{1 - \frac{4m_\chi^2}{m_a^2}} \cos^2 \theta,$$

Physical fields: h, H, a, A, H^\pm and χ .

Physical parameters: $\alpha, \beta, \theta, v, \lambda_3, \lambda_{P1}, \lambda_{P2}, m_h, m_H, m_a, m_A, m_{H^\pm}$.

α : mixing angle for scalars (h, H)

θ : mixing angle for pseudo-scalars (a, A)

2. LLP Phenomenology

2.1 Model parameters

- For small mixing angles θ , the pseudoscalar $a \simeq P$ can become **long-lived**.

$$\Gamma(a \rightarrow f\bar{f}) = \frac{N_c^f \eta_f^2 y_f^2}{16\pi} m_a \sqrt{1 - \frac{4m_f^2}{m_a^2}} \sin^2 \theta$$

$$\Gamma(a \rightarrow gg) = \frac{\alpha_s^2}{32\pi^3 v^2} m_a^3 \left| \sum_{q=t,b,c} \eta_q f \left(\frac{4m_q^2}{m_a^2} \right) \right|^2 \sin^2 \theta,$$

$$\Gamma(a \rightarrow \chi\bar{\chi}) = \frac{y_\chi^2}{8\pi} m_a \sqrt{1 - \frac{4m_\chi^2}{m_a^2}} \cos^2 \theta,$$

Physical fields: h, H, a, A, H^\pm and χ .

Physical parameters: $\alpha, \beta, \theta, v, \lambda_3, \lambda_{P1}, \lambda_{P2}, m_h, m_H, m_a, m_A, m_{H^\pm}$.

α : mixing angle for scalars (h, H)

θ : mixing angle for pseudo-scalars (a, A)

2. LLP Phenomenology

2.1 Model parameters

- For small mixing angles θ , the pseudoscalar $a \simeq P$ can become **long-lived**.

$$\Gamma(a \rightarrow f\bar{f}) = \frac{N_c^f \eta_f^2 y_f^2}{16\pi} m_a \sqrt{1 - \frac{4m_f^2}{m_a^2}} \sin^2 \theta$$

$$\Gamma(a \rightarrow gg) = \frac{\alpha_s^2}{32\pi^3 v^2} m_a^3 \left| \sum_{q=t,b,c} \eta_q f \left(\frac{4m_q^2}{m_a^2} \right) \right|^2 \sin^2 \theta$$

$$\Gamma(a \rightarrow \chi\bar{\chi}) = \frac{y_\chi^2}{8\pi} m_a \sqrt{1 - \frac{4m_\chi^2}{m_a^2}} \cos^2 \theta,$$

Physical fields: h, H, a, A, H^\pm and χ .

Physical parameters: $\alpha, \beta, \theta, v, \lambda_3, \lambda_{P1}, \lambda_{P2}, m_h, m_H, m_a, m_A, m_{H^\pm}$.

α : mixing angle for scalars (h, H)

θ : mixing angle for pseudo-scalars (a, A)

2. LLP Phenomenology

2.1 Model parameters

- For small mixing angles θ , the pseudoscalar $a \simeq P$ can become **long-lived**.

$$\Gamma(a \rightarrow f\bar{f}) = \frac{N_c^f \eta_f^2 y_f^2}{16\pi} m_a \sqrt{1 - \frac{4m_f^2}{m_a^2} \sin^2 \theta}$$

$$\Gamma(a \rightarrow gg) = \frac{\alpha_s^2}{32\pi^3 v^2} m_a^3 \left| \sum_{q=t,b,c} \eta_q f \left(\frac{4m_q^2}{m_a^2} \right) \right|^2 \sin^2 \theta$$

$$\Gamma(a \rightarrow \chi\bar{\chi}) = \frac{y_\chi^2}{8\pi} m_a \sqrt{1 - \frac{4m_\chi^2}{m_a^2} \cos^2 \theta},$$

Physical fields: h, H, a, A, H^\pm and χ .

Physical parameters: $\alpha, \beta, \theta, v, \lambda_3, \lambda_{P1}, \lambda_{P2}, m_h, m_H, m_a, m_A, m_{H^\pm}$.

α : mixing angle for scalars (h, H)

θ : mixing angle for pseudo-scalars (a, A)

2. LLP Phenomenology

2.1 Model parameters

- For small mixing angles θ , the pseudoscalar $a \simeq P$ can become **long-lived**.

$$\Gamma(a \rightarrow f\bar{f}) = \frac{N_c^f \eta_f^2 y_f^2}{16\pi} m_a \sqrt{1 - \frac{4m_f^2}{m_a^2}} \sin^2 \theta$$

$$\Gamma(a \rightarrow gg) = \frac{\alpha_s^2}{32\pi^3 v^2} m_a^3 \left| \sum_{q=t,b,c} \eta_q f \left(\frac{4m_q^2}{m_a^2} \right) \right|^2 \sin^2 \theta$$

$$\Gamma(a \rightarrow \chi\bar{\chi}) = \frac{y_\chi^2}{8\pi} m_a \sqrt{1 - \frac{4m_\chi^2}{m_a^2}} \cos^2 \theta,$$

Physical fields: h, H, a, A, H^\pm and χ .

Physical parameters: $\alpha, \beta, \theta, v, \lambda_3, \lambda_{P1}, \lambda_{P2}, m_h, m_H, m_a, m_A, m_{H^\pm}$.

α : mixing angle for scalars (h, H)

θ : mixing angle for pseudo-scalars (a, A)

2. LLP Phenomenology

2.1 Model parameters

- For small mixing angles θ , the pseudoscalar $a \simeq P$ can become **long-lived**.

Source: [ArXiv:1802.02156](https://arxiv.org/abs/1802.02156)
(U. Haisch, J.F. Kamenik,
A. Malinauskas, M. Spira).

$$\Gamma(a \rightarrow f\bar{f}) = \frac{N_c^f \eta_f^2 y_f^2}{16\pi} m_a \sqrt{1 - \frac{4m_f^2}{m_a^2}} \sin^2 \theta$$

$$\Gamma(a \rightarrow gg) = \frac{\alpha_s^2}{32\pi^3 v^2} m_a^3 \left| \sum_{q=t,b,c} \eta_q f \left(\frac{4m_q^2}{m_a^2} \right) \right|^2 \sin^2 \theta$$

$$\Gamma(a \rightarrow \chi\bar{\chi}) = \frac{y_\chi^2}{8\pi} m_a \sqrt{1 - \frac{4m_\chi^2}{m_a^2}} \cos^2 \theta,$$

**NLO
+ QCD
corr.**

Physical fields: h, H, a, A, H^\pm and χ .

Physical parameters: $\alpha, \beta, \theta, v, \lambda_3, \lambda_{P1}, \lambda_{P2}, m_h, m_H, m_a, m_A, m_{H^\pm}$.

α : mixing angle for scalars (h, H)

θ : mixing angle for pseudo-scalars (a, A)

2. LLP Phenomenology

2.1 Model parameters

- For small mixing angles θ , the pseudoscalar $a \simeq P$ can become **long-lived**.

Source: [ArXiv:1802.02156](https://arxiv.org/abs/1802.02156)
(U. Haisch, J.F. Kamenik,
A. Malinauskas, M. Spira).

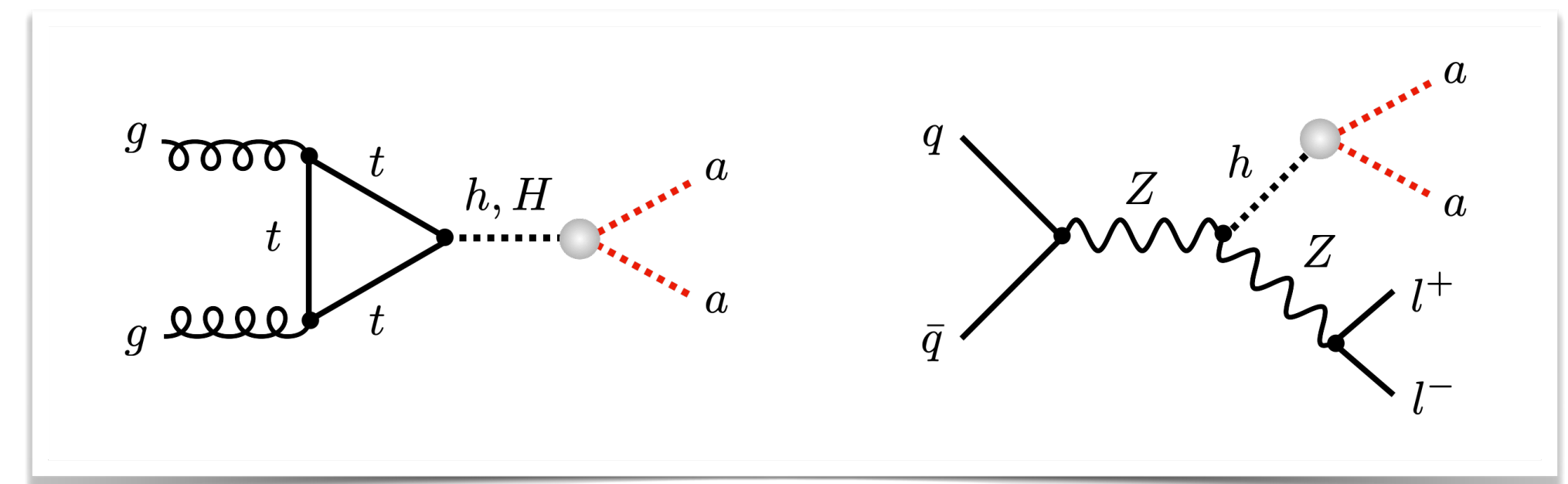
$$\Gamma(a \rightarrow f\bar{f}) = \frac{N_c^f \eta_f^2 y_f^2}{16\pi} m_a \sqrt{1 - \frac{4m_f^2}{m_a^2}} \sin^2 \theta$$

$$\Gamma(a \rightarrow gg) = \frac{\alpha_s^2}{32\pi^3 v^2} m_a^3 \left| \sum_{q=t,b,c} \eta_q f \left(\frac{4m_q^2}{m_a^2} \right) \right|^2 \sin^2 \theta$$

$$\Gamma(a \rightarrow \chi\bar{\chi}) = \frac{y_\chi^2}{8\pi} m_a \sqrt{1 - \frac{4m_\chi^2}{m_a^2}} \cos^2 \theta,$$

**NLO
+ QCD
corr.**

- Production via the **decay of a heavier spin-0 state**:



Source: [ArXiv:2302.02735](https://arxiv.org/abs/2302.02735) (U. Haisch, LS).

Physical fields: h, H, a, A, H^\pm and χ .

Physical parameters: $\alpha, \beta, \theta, v, \lambda_3, \lambda_{P1}, \lambda_{P2}, m_h, m_H, m_a, m_A, m_{H^\pm}$.

α : mixing angle for scalars (h, H)

θ : mixing angle for pseudo-scalars (a, A)

2. LLP Phenomenology

2.1 Model parameters

- For small mixing angles θ , the pseudoscalar $a \simeq P$ can become **long-lived**.

Source: [ArXiv:1802.02156](https://arxiv.org/abs/1802.02156)
(U. Haisch, J.F. Kamenik, A. Malinauskas, M. Spira).

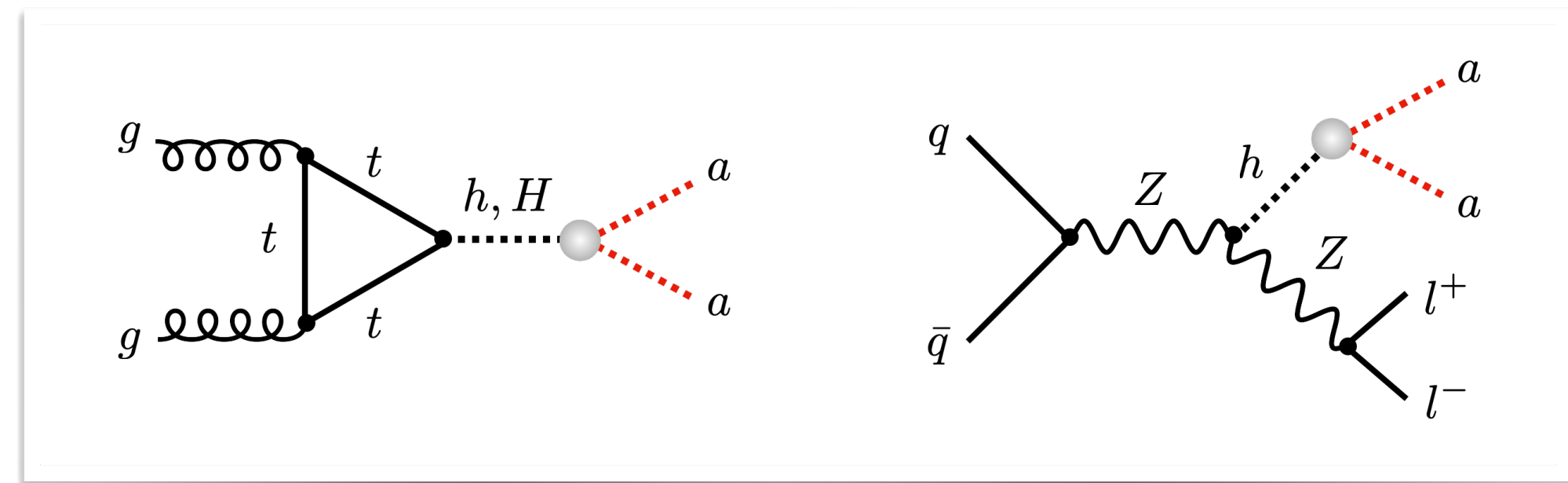
$$\Gamma(a \rightarrow f\bar{f}) = \frac{N_c^f \eta_f^2 y_f^2}{16\pi} m_a \sqrt{1 - \frac{4m_f^2}{m_a^2}} \sin^2 \theta$$

$$\Gamma(a \rightarrow gg) = \frac{\alpha_s^2}{32\pi^3 v^2} m_a^3 \left| \sum_{q=t,b,c} \eta_q f \left(\frac{4m_q^2}{m_a^2} \right) \right|^2 \sin^2 \theta$$

$$\Gamma(a \rightarrow \chi\bar{\chi}) = \frac{y_\chi^2}{8\pi} m_a \sqrt{1 - \frac{4m_\chi^2}{m_a^2}} \cos^2 \theta,$$

**NLO
+ QCD
corr.**

- Production via the **decay of a heavier spin-0 state**:
 - Benchmark I:** $m_a < m_h/2$



Source: [ArXiv:2302.02735](https://arxiv.org/abs/2302.02735) (U. Haisch, LS).

Physical fields: h, H, a, A, H^\pm and χ .
 Physical parameters: $\alpha, \beta, \theta, v, \lambda_3, \lambda_{P1}, \lambda_{P2}, m_h, m_H, m_a, m_A, m_{H^\pm}$.
 α : mixing angle for scalars (h, H) θ : mixing angle for pseudo-scalars (a, A)

2. LLP Phenomenology

2.1 Model parameters

Source: [ArXiv:1802.02156](https://arxiv.org/abs/1802.02156)
(U. Haisch, J.F. Kamenik,
A. Malinauskas, M. Spira).

- For small mixing angles θ , the pseudoscalar $a \simeq P$ can become **long-lived**.

$$\Gamma(a \rightarrow f\bar{f}) = \frac{N_c^f \eta_f^2 y_f^2}{16\pi} m_a \sqrt{1 - \frac{4m_f^2}{m_a^2} \sin^2 \theta}$$

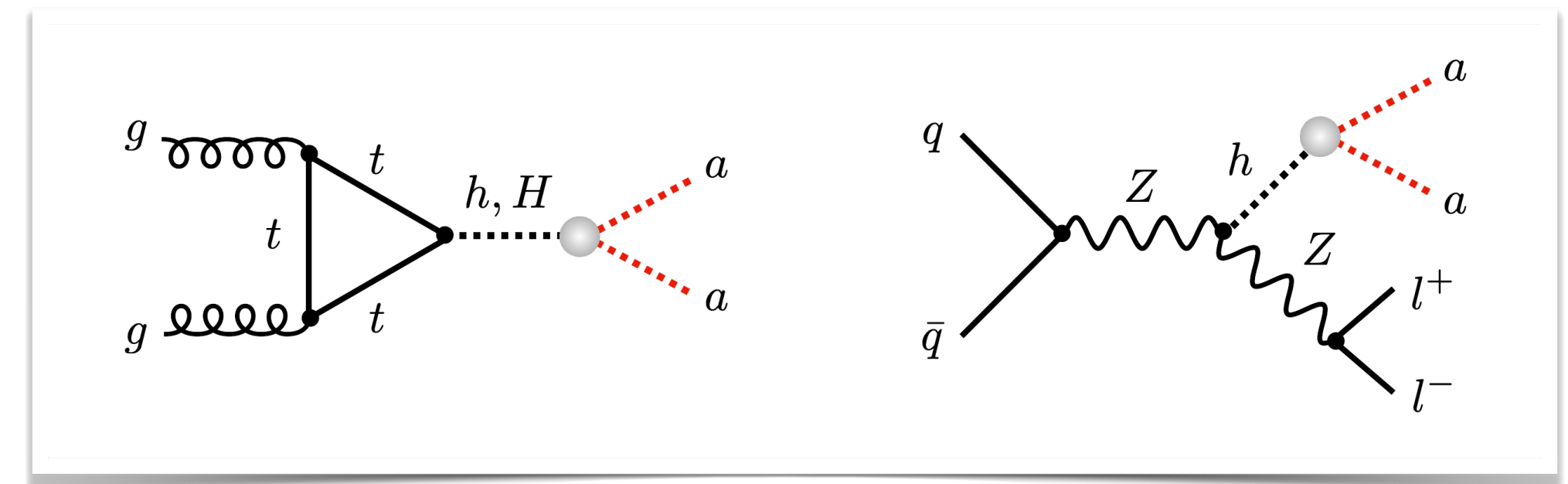
$$\Gamma(a \rightarrow gg) = \frac{\alpha_s^2}{32\pi^3 v^2} m_a^3 \left| \sum_{q=t,b,c} \eta_q f \left(\frac{4m_q^2}{m_a^2} \right) \right|^2 \sin^2 \theta$$

$$\Gamma(a \rightarrow \chi\bar{\chi}) = \frac{y_\chi^2}{8\pi} m_a \sqrt{1 - \frac{4m_\chi^2}{m_a^2}} \cos^2 \theta,$$

**NLO
+ QCD
corr.**

- Production via the **decay of a heavier spin-0 state**:

- **Benchmark I:** $m_a < m_h/2$
- **Benchmark II:** $m_h/2 < m_a < m_H/2$



Source: [ArXiv:2302.02735](https://arxiv.org/abs/2302.02735) (U. Haisch, LS).

Physical fields: h, H, a, A, H^\pm and χ .

Physical parameters: $\alpha, \beta, \theta, v, \lambda_3, \lambda_{P1}, \lambda_{P2}, m_h, m_H, m_a, m_A, m_{H^\pm}$.

α : mixing angle for scalars (h, H)

θ : mixing angle for pseudo-scalars (a, A)

2. LLP Phenomenology

2.1 Model parameters

Source: [ArXiv:1802.02156](https://arxiv.org/abs/1802.02156)
(U. Haisch, J.F. Kamenik, A. Malinauskas, M. Spira).

- For small mixing angles θ , the pseudoscalar $a \simeq P$ can become **long-lived**.

$$\Gamma(a \rightarrow f\bar{f}) = \frac{N_c^f \eta_f^2 y_f^2}{16\pi} m_a \sqrt{1 - \frac{4m_f^2}{m_a^2} \sin^2 \theta}$$

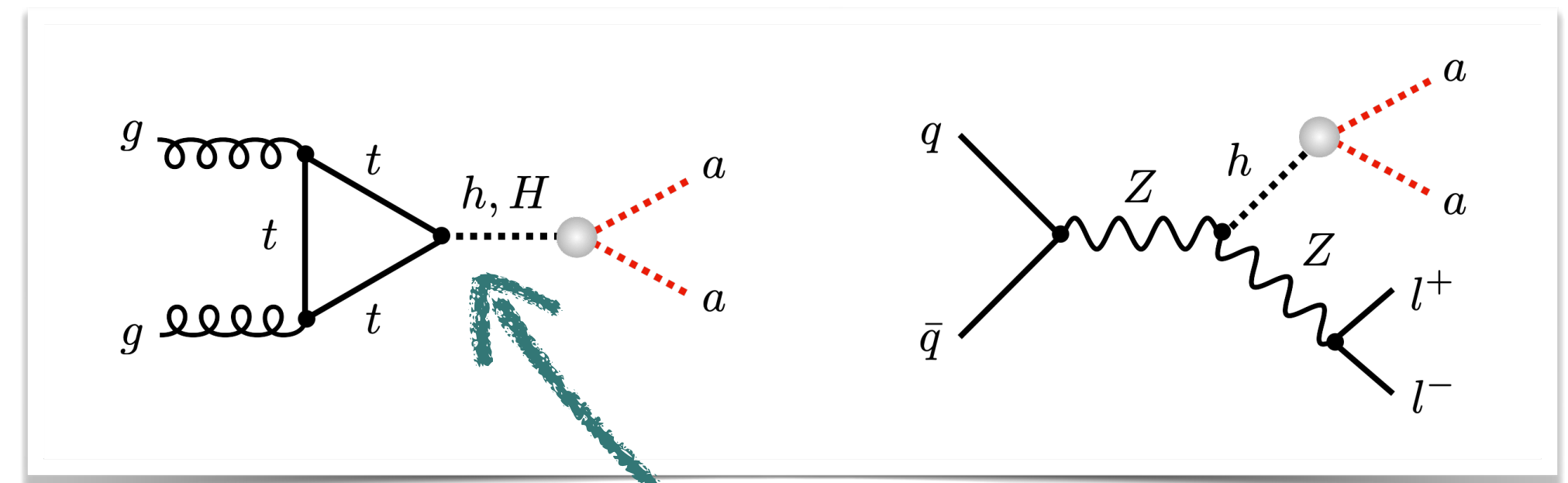
$$\Gamma(a \rightarrow gg) = \frac{\alpha_s^2}{32\pi^3 v^2} m_a^3 \left| \sum_{q=t,b,c} \eta_q f \left(\frac{4m_q^2}{m_a^2} \right) \right|^2 \sin^2 \theta$$

$$\Gamma(a \rightarrow \chi\bar{\chi}) = \frac{y_\chi^2}{8\pi} m_a \sqrt{1 - \frac{4m_\chi^2}{m_a^2}} \cos^2 \theta,$$

**NLO
+ QCD
corr.**

- Production via the **decay of a heavier spin-0 state**:

- **Benchmark I:** $m_a < m_h/2$
- **Benchmark II:** $m_h/2 < m_a < m_H/2$



Source: [ArXiv:2302.02735](https://arxiv.org/abs/2302.02735) (U. Haisch, LS).

BR($h \rightarrow aa$) $\simeq 2\%$
(currently BR($h \rightarrow \text{inv.}$) $\lesssim 9\%$)

Physical fields: h, H, a, A, H^\pm and χ .
Physical parameters: $\alpha, \beta, \theta, v, \lambda_3, \lambda_{P1}, \lambda_{P2}, m_h, m_H, m_a, m_A, m_{H^\pm}$.
 α : mixing angle for scalars (h, H) θ : mixing angle for pseudo-scalars (a, A)

2. LLP Phenomenology

2.2 LLP constraints

Physical fields: h, H, a, A, H^\pm and χ .

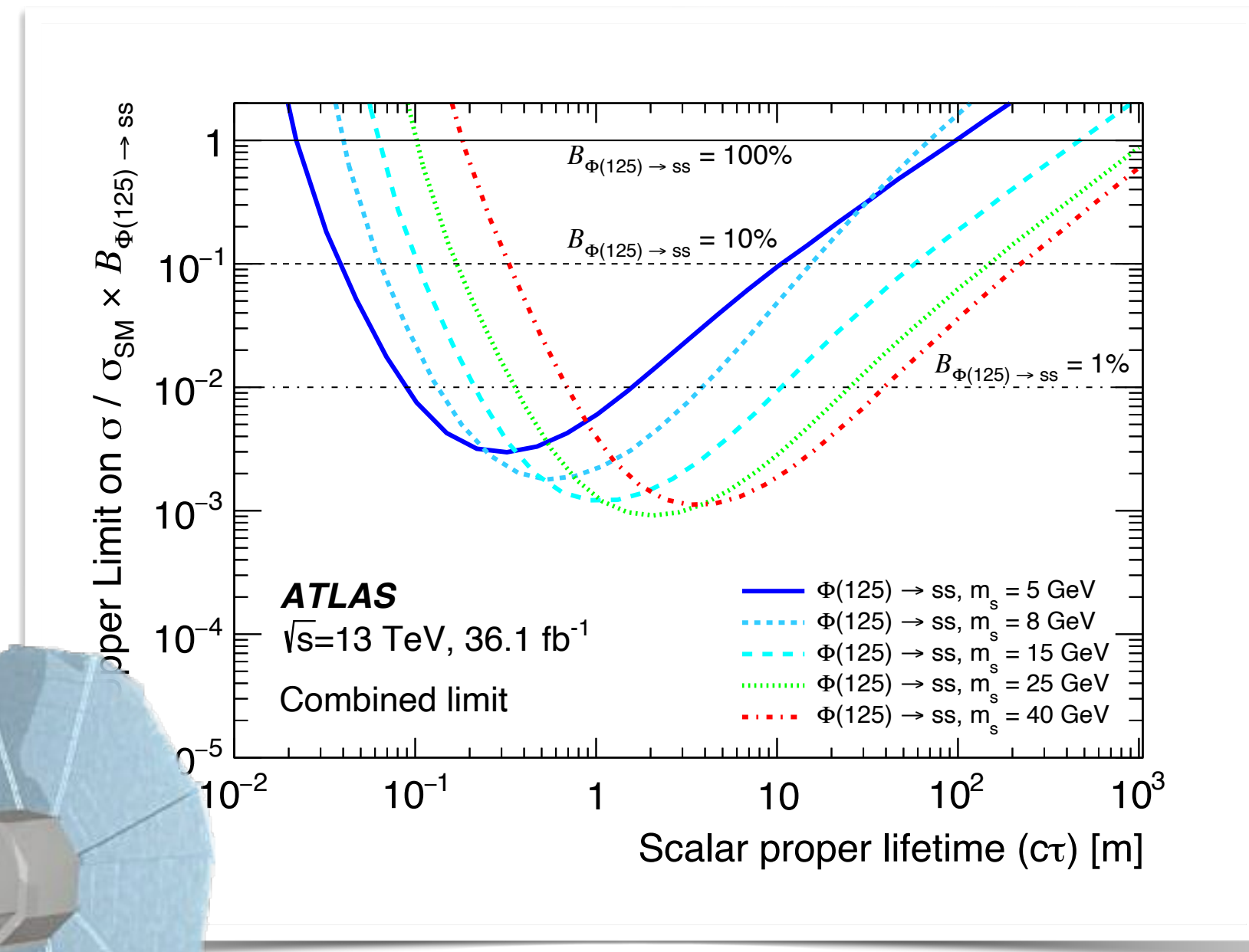
Physical parameters: $\alpha, \beta, \theta, \nu, \lambda_3, \lambda_{P1}, \lambda_{P2}, m_h, m_H, m_a, m_A, m_{H^\pm}$.

α : mixing angle for scalars (h, H)

θ : mixing angle for pseudo-scalars (a, A)

2. LLP Phenomenology

2.2 LLP constraints



ArXiv:1811.07370 (ATLAS).

Physical fields: h, H, a, A, H^\pm and χ .

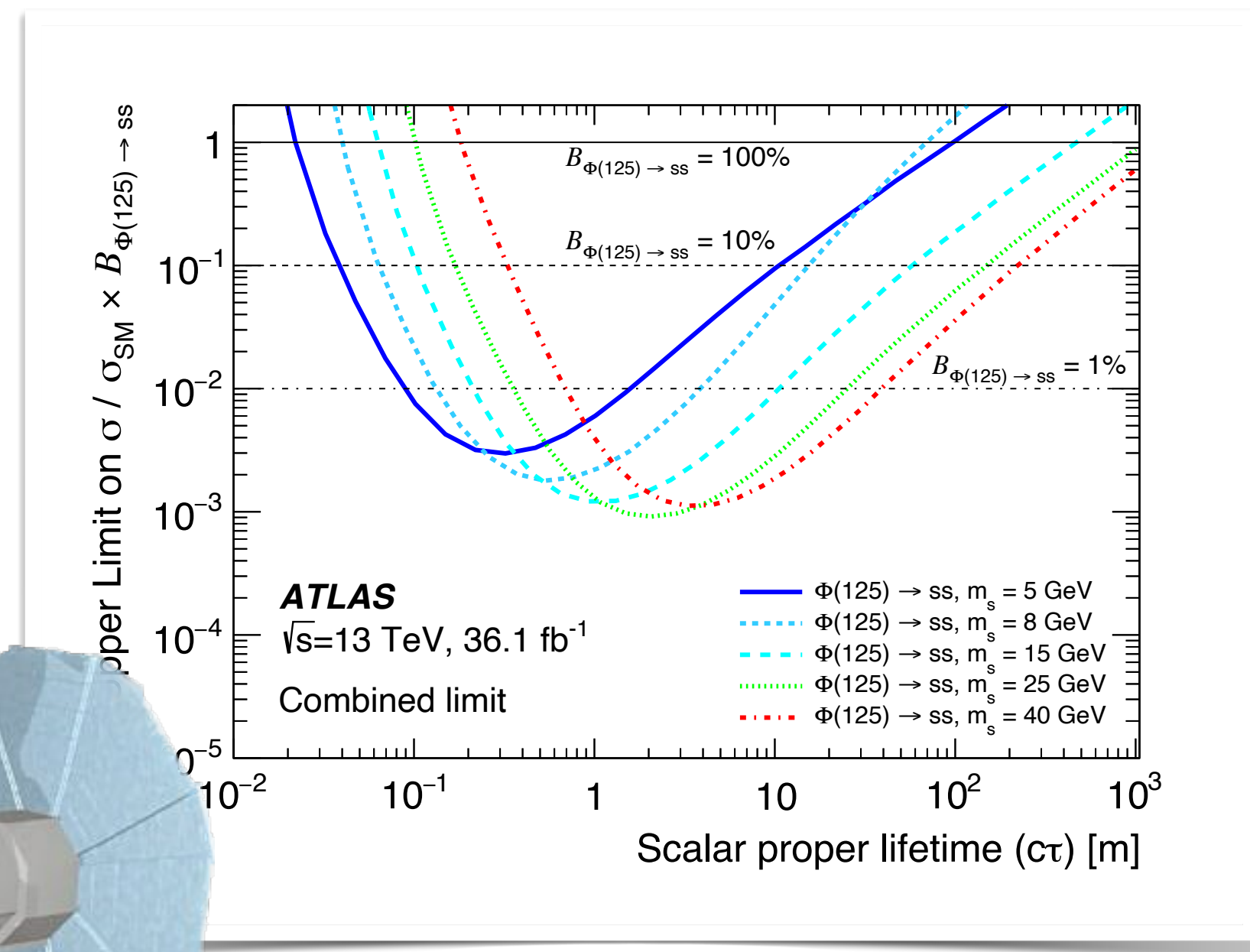
Physical parameters: $\alpha, \beta, \theta, v, \lambda_3, \lambda_{P1}, \lambda_{P2}, m_h, m_H, m_a, m_A, m_{H^\pm}$.

α : mixing angle for scalars (h, H)

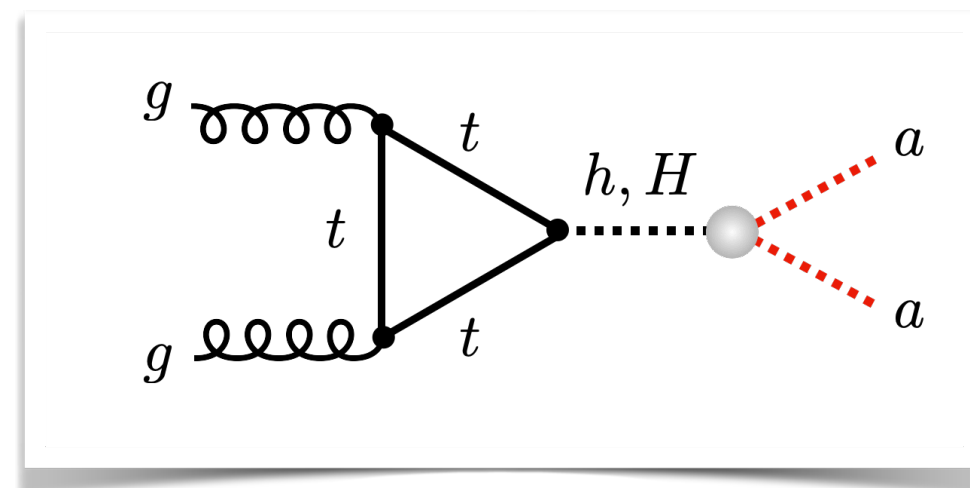
θ : mixing angle for pseudo-scalars (a, A)

2. LLP Phenomenology

2.2 LLP constraints



ArXiv:1811.07370 (ATLAS).



$$\frac{c\tau_a}{m} \simeq 4.8 \cdot 10^{-12} \left(\frac{\text{GeV}}{m_a} \right)^{0.9} \frac{1}{\sin^2 \theta},$$



Physical fields: h, H, a, A, H^\pm and χ .

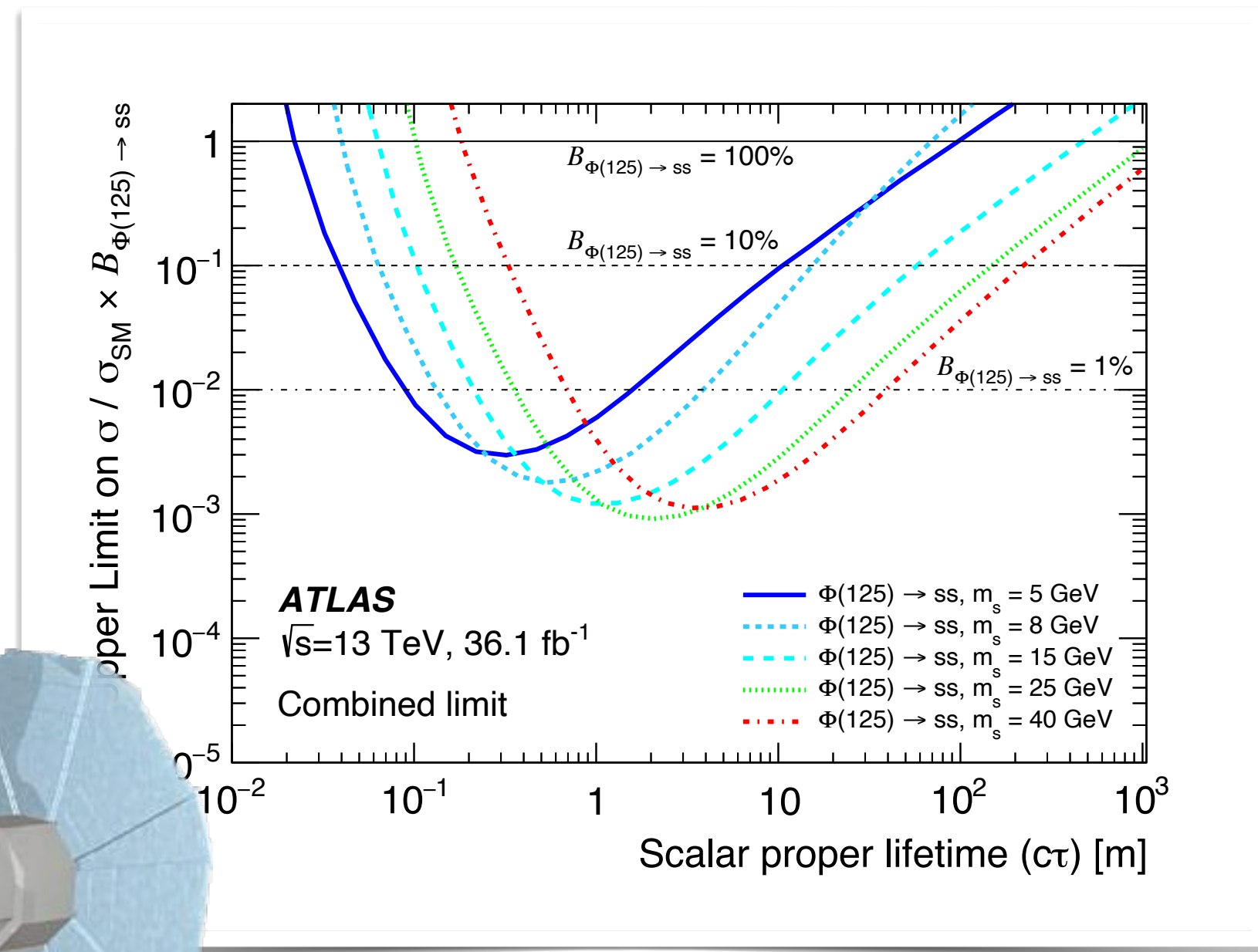
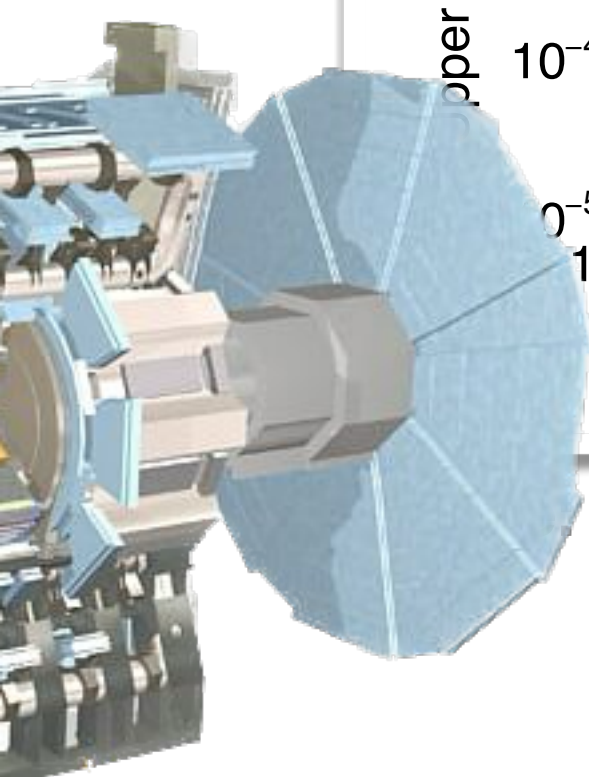
Physical parameters: $\alpha, \beta, \theta, v, \lambda_3, \lambda_{P1}, \lambda_{P2}, m_h, m_H, m_a, m_A, m_{H^\pm}$.

α : mixing angle for scalars (h, H)

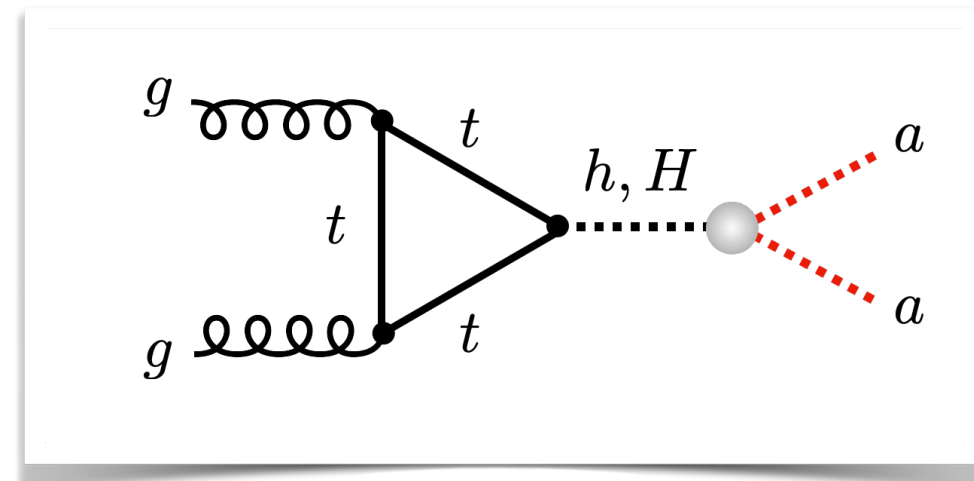
θ : mixing angle for pseudo-scalars (a, A)

2. LLP Phenomenology

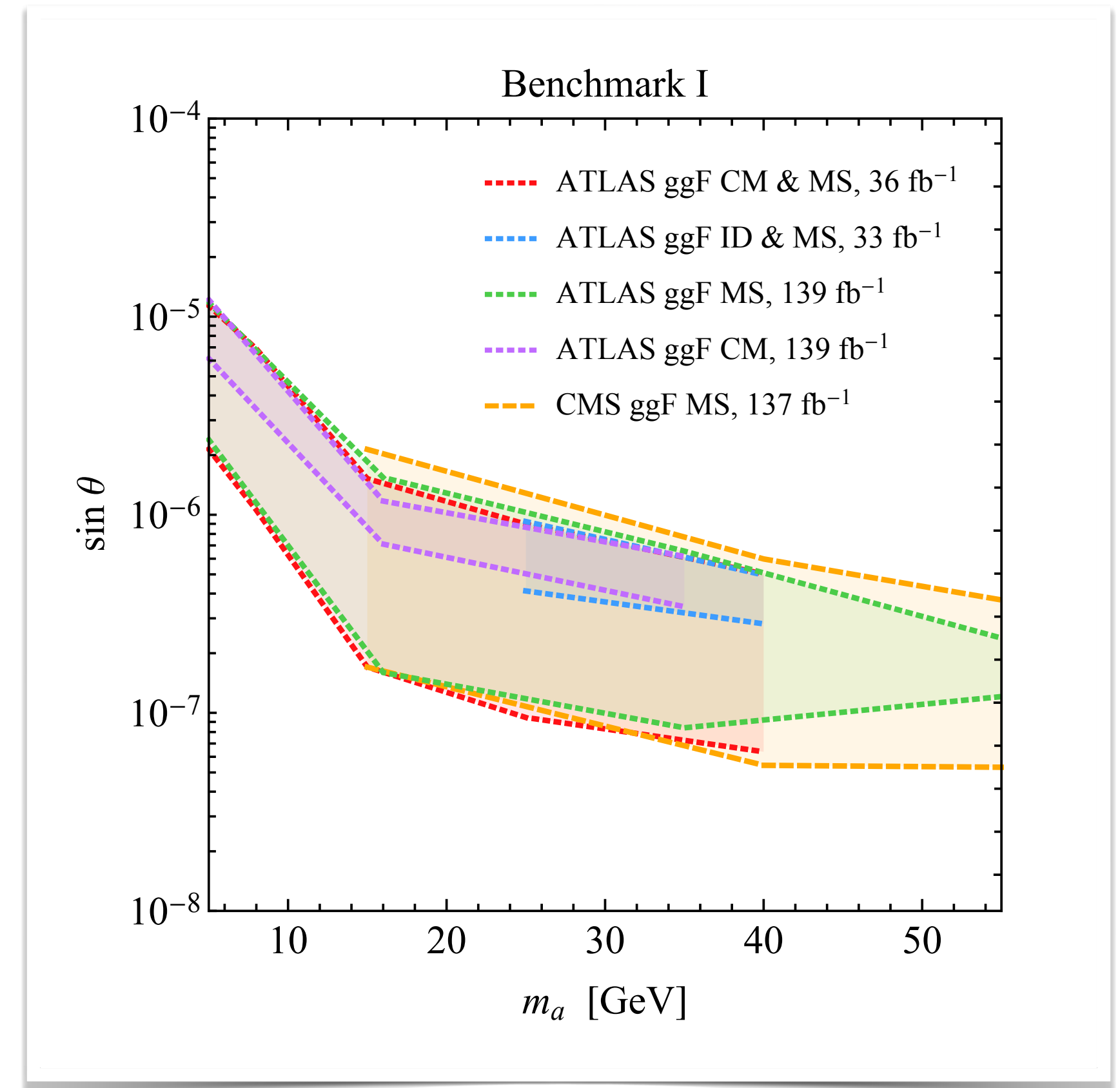
2.2 LLP constraints



ArXiv:1811.07370 (ATLAS).



$$\frac{c\tau_a}{m} \simeq 4.8 \cdot 10^{-12} \left(\frac{\text{GeV}}{m_a} \right)^{0.9} \frac{1}{\sin^2 \theta}$$



Source: ArXiv:2302.02735 (U. Haisch, LS).

Physical fields: h, H, a, A, H^\pm and χ .

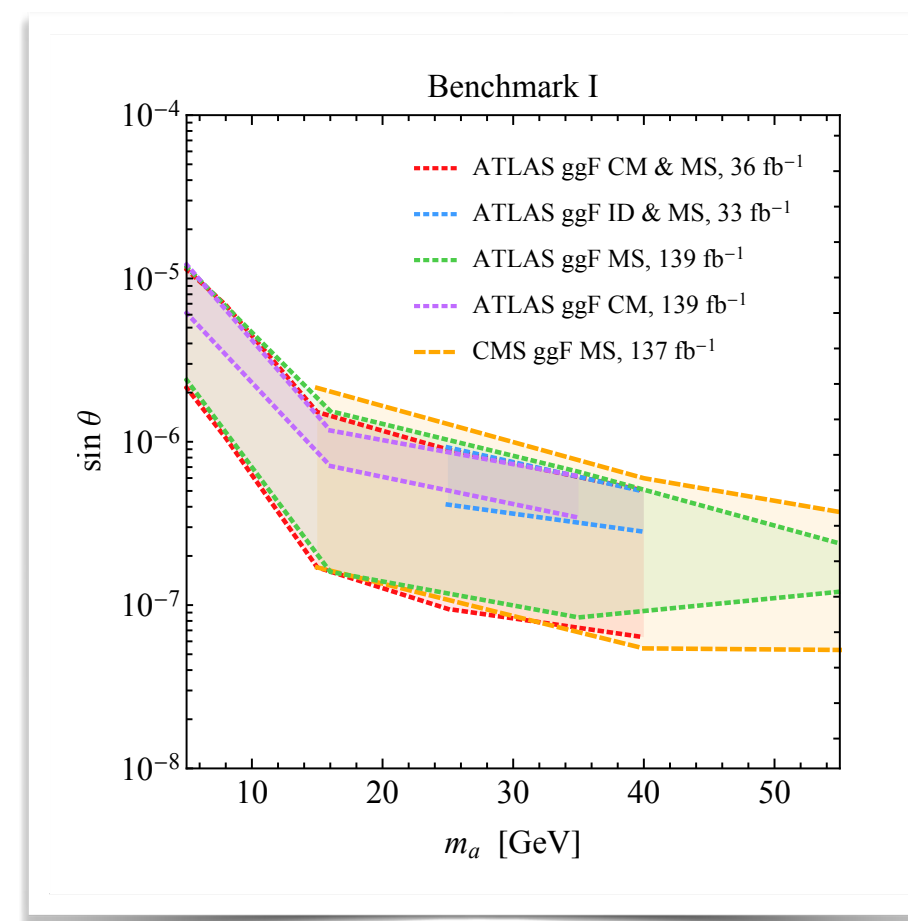
Physical parameters: $\alpha, \beta, \theta, v, \lambda_3, \lambda_{P1}, \lambda_{P2}, m_h, m_H, m_a, m_A, m_{H^\pm}$.

α : mixing angle for scalars (h, H)

θ : mixing angle for pseudo-scalars (a, A)

2. LLP Phenomenology

2.2 LLP constraints



Source: [ArXiv:2302.02735](https://arxiv.org/abs/2302.02735)
(U. Haisch, LS).

2. LLP Phenomenology

2.2 LLP constraints

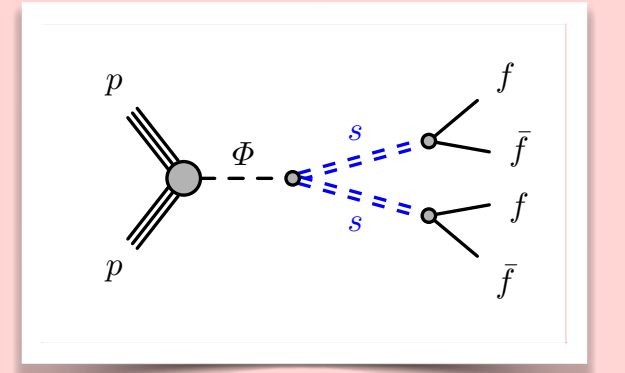
ArXiv:1811.07370 (ATLAS)

36.1 fb⁻¹ of data.

Three benchmark models (one is scalar portal).

Narrow jets in **muon spectrometer (MS)**.

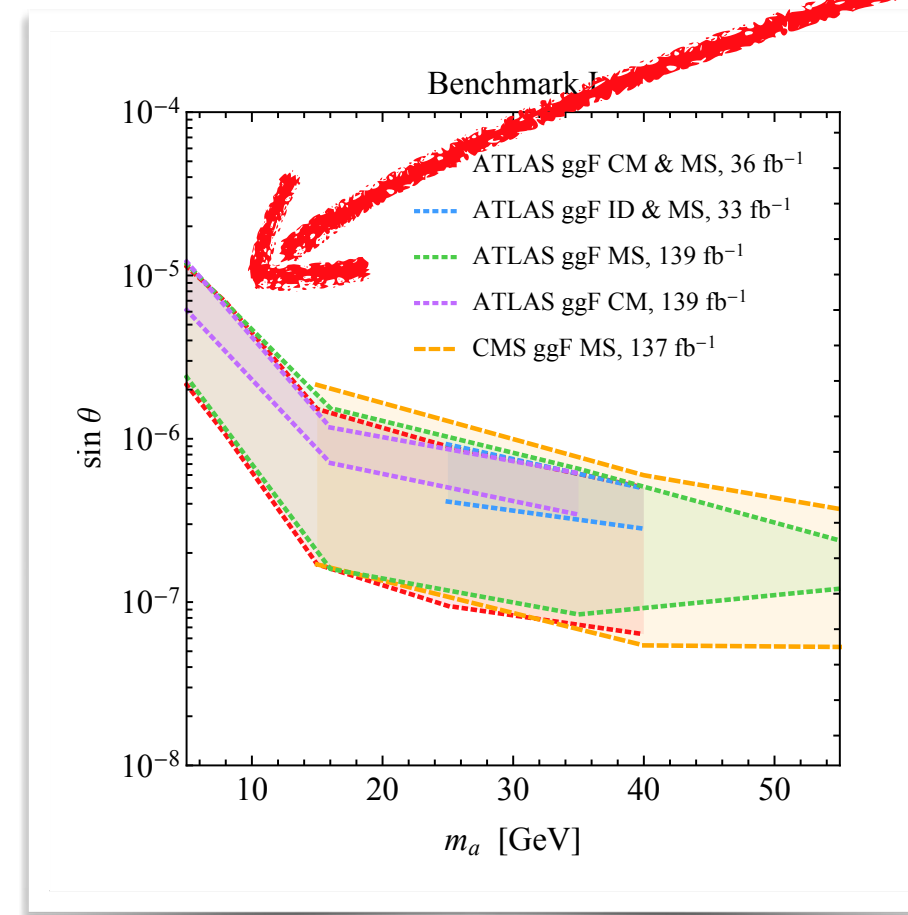
2 MS vertices, 1 MS vertex and $E_T^{miss} > 30$ GeV.



ArXiv:1902.03094 (ATLAS)

10.8 fb⁻¹ of data.

Two narrow jets in **HCal** with no associated activity in tracker, high E_H/E_{EM} („CalRatio“).



Source: [ArXiv:2302.02735](https://arxiv.org/abs/2302.02735)
(U. Haisch, LS).

2. LLP Phenomenology

2.2 LLP constraints

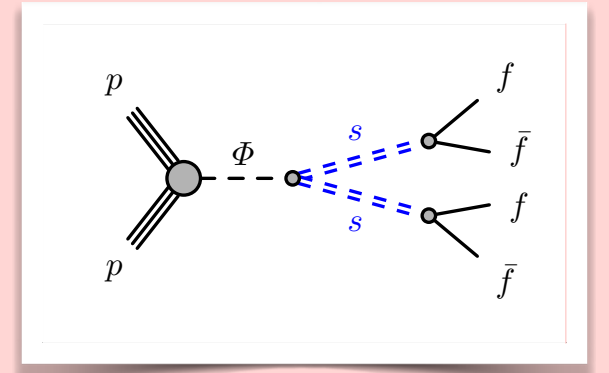
ArXiv:1811.07370 (ATLAS)

36.1 fb⁻¹ of data.

Three benchmark models (one is scalar portal).

Narrow jets in **muon spectrometer (MS)**.

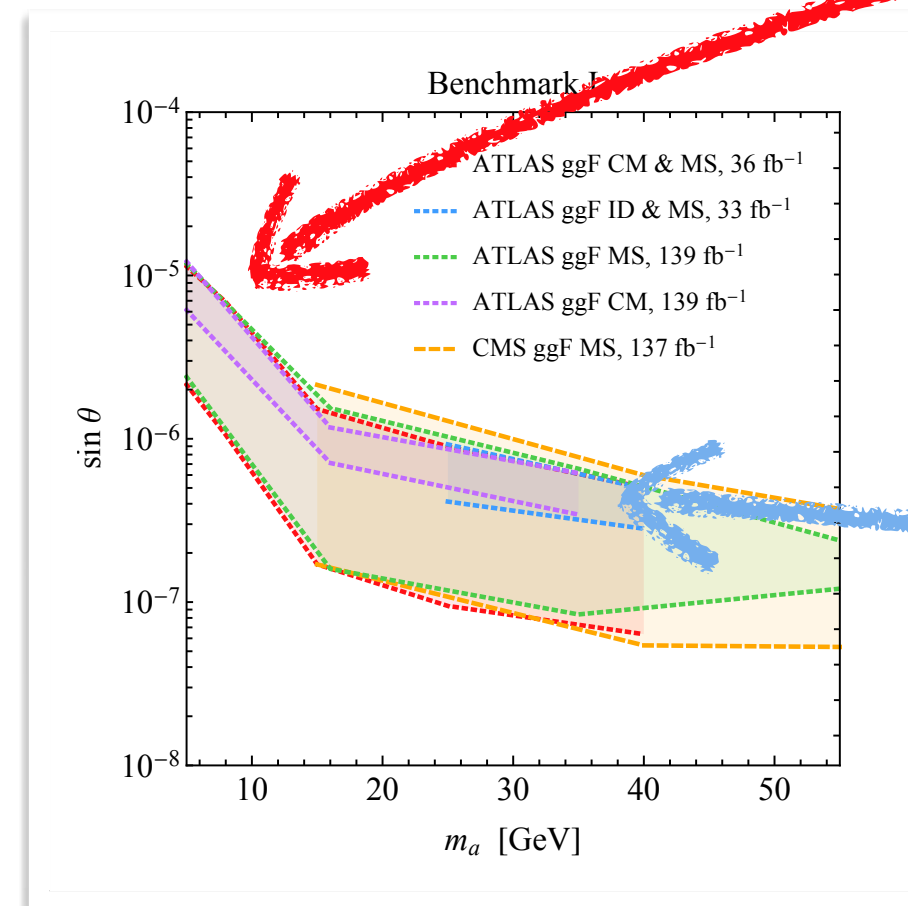
2 MS vertices, 1 MS vertex and $E_T^{miss} > 30$ GeV.



ArXiv:1902.03094 (ATLAS)

10.8 fb⁻¹ of data.

Two narrow jets in **HCal** with no associated activity in tracker, high E_H/E_{EM} („CalRatio“).



Source: [ArXiv:2302.02735](https://arxiv.org/abs/2302.02735)
(U. Haisch, LS).

ArXiv:1911.12575 (ATLAS)

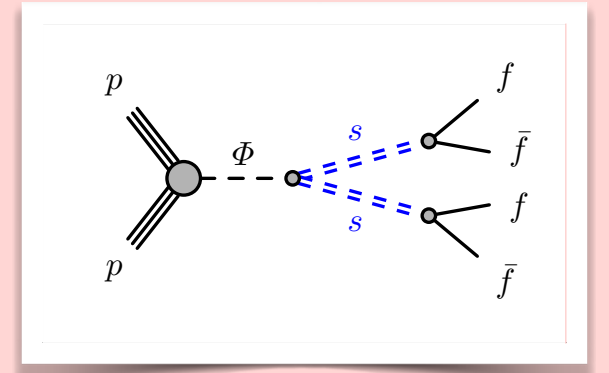
33 fb⁻¹ of data.

Narrow jet in **MS** and displaced track in **inner detector (ID)**.

Branching ratios $a \rightarrow b\bar{b} : c\bar{c} : \tau^+\tau^-$ assumed to be 85:5:8.

2. LLP Phenomenology

2.2 LLP constraints



ArXiv:1811.07370 (ATLAS)

36.1 fb⁻¹ of data.

Three benchmark models (one is scalar portal).

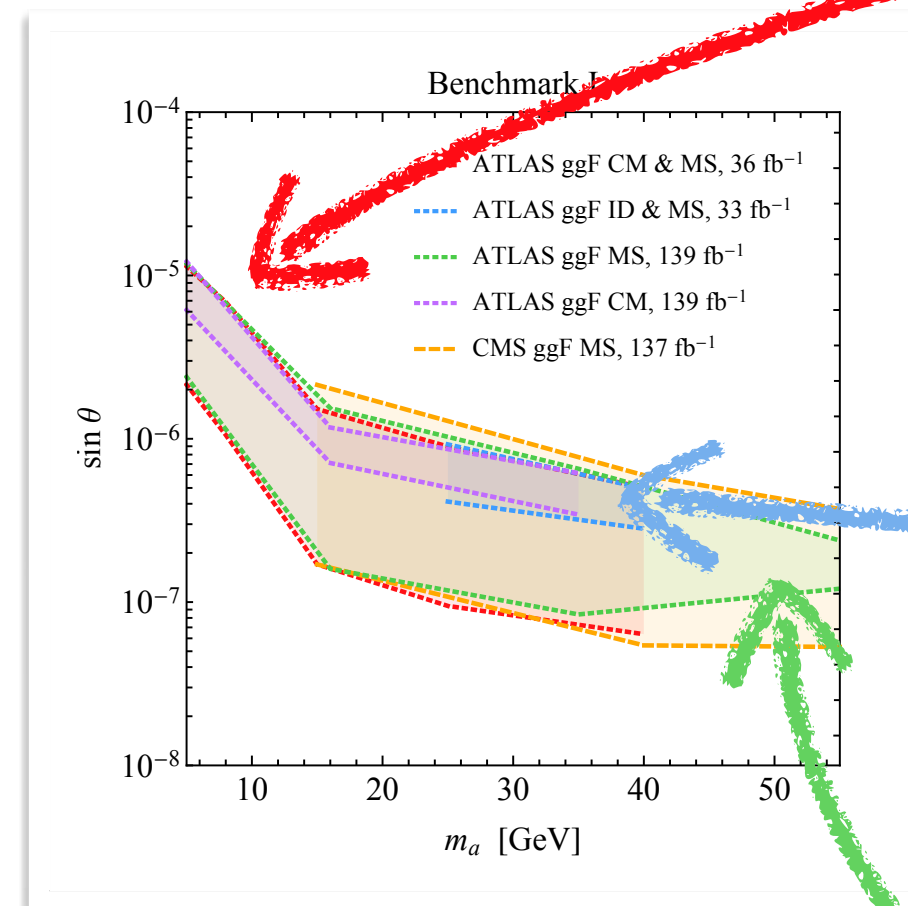
Narrow jets in **muon spectrometer (MS)**.

2 MS vertices, 1 MS vertex and $E_T^{miss} > 30$ GeV.

ArXiv:1902.03094 (ATLAS)

10.8 fb⁻¹ of data.

Two narrow jets in **HCal** with no associated activity in tracker, high E_H/E_{EM} („CalRatio“).



Source: [ArXiv:2302.02735](https://arxiv.org/abs/2302.02735)
(U. Haisch, LS).

ArXiv:1911.12575 (ATLAS)

33 fb⁻¹ of data.

Narrow jet in **MS** and displaced track in **inner detector (ID)**.

Branching ratios $a \rightarrow b\bar{b} : c\bar{c} : \tau^+\tau^-$ assumed to be 85:5:8.

ArXiv:2203.00587 (ATLAS)

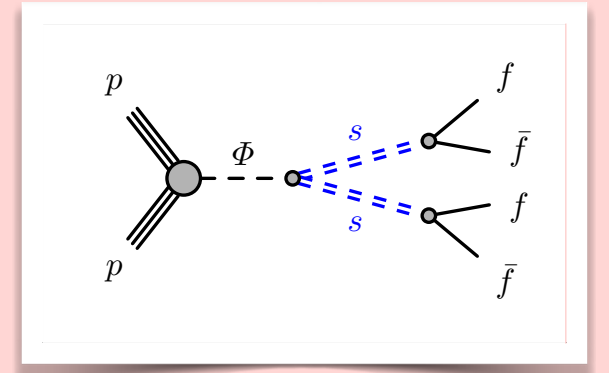
139 fb⁻¹ of data.

Two narrow, high-multiplicity jets in **MS**.

Background (punch-through jets, noncollision background) reduced by requiring two displaced vertices.

2. LLP Phenomenology

2.2 LLP constraints



ArXiv:1811.07370 (ATLAS)

36.1 fb^{-1} of data.

Three benchmark models (one is scalar portal).

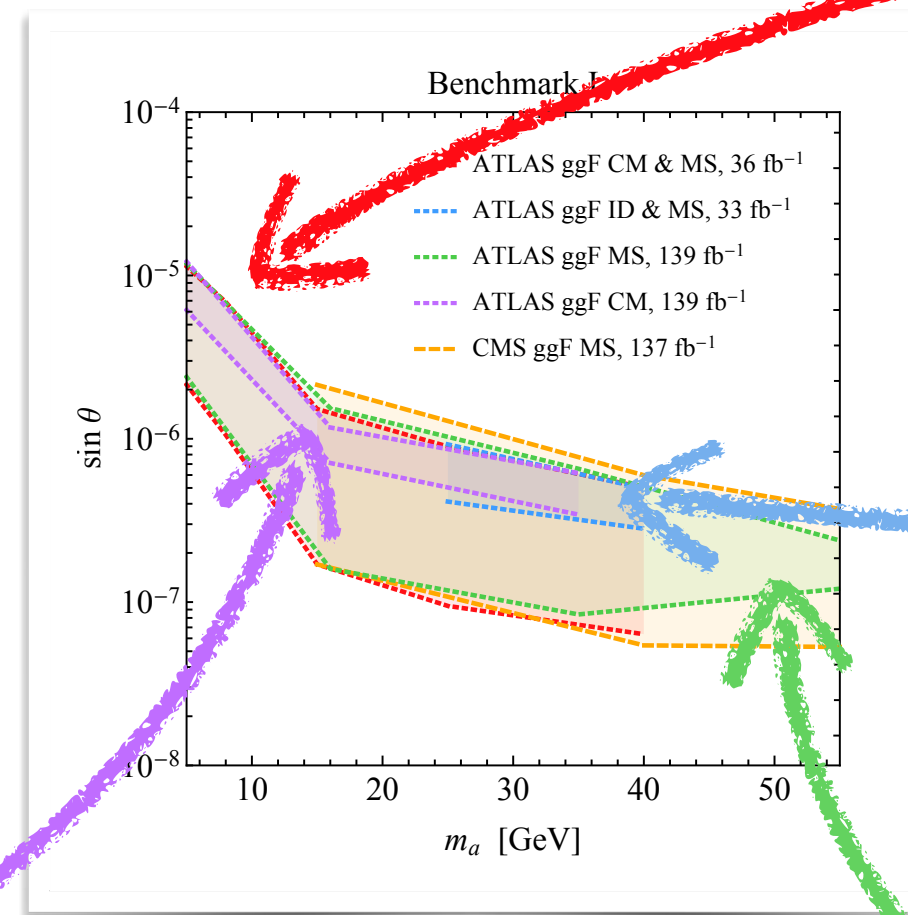
Narrow jets in **muon spectrometer (MS)**.

2 MS vertices, 1 MS vertex and $E_T^{miss} > 30$ GeV.

ArXiv:1902.03094 (ATLAS)

10.8 fb^{-1} of data.

Two narrow jets in **HCal** with no associated activity in tracker, high E_H/E_{EM} („*CalRatio*“).



Source: [ArXiv:2302.02735](https://arxiv.org/abs/2302.02735)
(U. Haisch, LS).

ArXiv:1911.12575 (ATLAS)

33 fb^{-1} of data.

Narrow jet in **MS** and displaced track in **inner detector (ID)**.

Branching ratios $a \rightarrow b\bar{b} : c\bar{c} : \tau^+\tau^-$ assumed to be 85:5:8.

ArXiv:2203.01009 (ATLAS)

139 fb^{-1} of data.

Two narrow jets in **HCal** with no associated activity in tracker, high E_H/E_{EM} („*CalRatio*“).

Improved displaced-jet identification (NN).

ArXiv:2203.00587 (ATLAS)

139 fb^{-1} of data.

Two narrow, high-multiplicity jets in **MS**.

Background (punch-through jets, noncollision background) reduced by requiring two displaced vertices.

2. LLP Phenomenology

2.2 LLP constraints

ArXiv:2107.04838 (CMS)

137 fb⁻¹ of data.

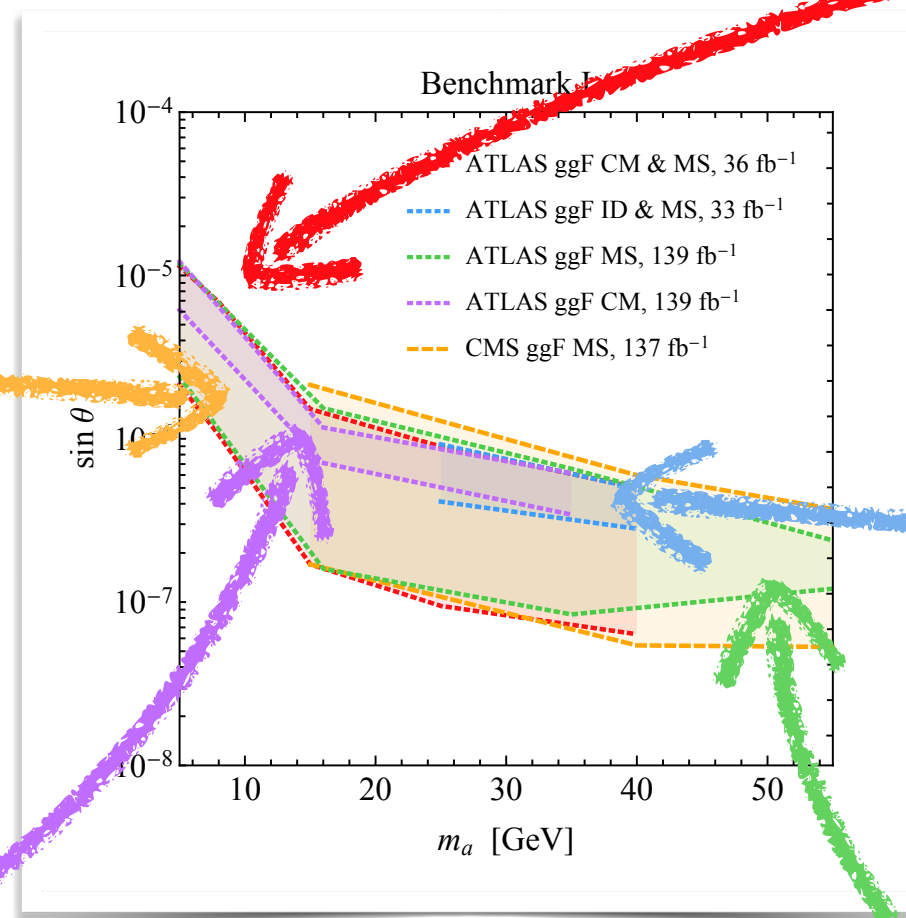
One particle shower in **endcap muon detectors (EMD)** and $p_T^{miss} > 200$ GeV.

Sufficient level of shielding in front of the EMD makes background low enough to only search for one shower.

ArXiv:2203.01009 (ATLAS)

139 fb⁻¹ of data.

Two narrow jets in **HCal** with no associated activity in tracker, high E_H/E_{EM} („CalRatio“). Improved displaced-jet identification (NN).



Source: ArXiv:2302.02735 (U. Haisch, LS).

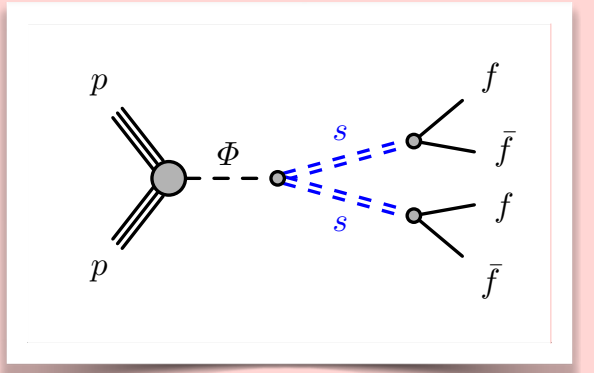
ArXiv:1811.07370 (ATLAS)

36.1 fb⁻¹ of data.

Three benchmark models (one is scalar portal).

Narrow jets in **muon spectrometer (MS)**.

2 MS vertices, 1 MS vertex and $E_T^{miss} > 30$ GeV.



ArXiv:1902.03094 (ATLAS)

10.8 fb⁻¹ of data.

Two narrow jets in **HCal** with no associated activity in tracker, high E_H/E_{EM} („CalRatio“).

ArXiv:1911.12575 (ATLAS)

33 fb⁻¹ of data.

Narrow jet in **MS** and displaced track in **inner detector (ID)**. Branching ratios $a \rightarrow b\bar{b} : c\bar{c} : \tau^+\tau^-$ assumed to be 85:5:8.

ArXiv:2203.00587 (ATLAS)

139 fb⁻¹ of data.

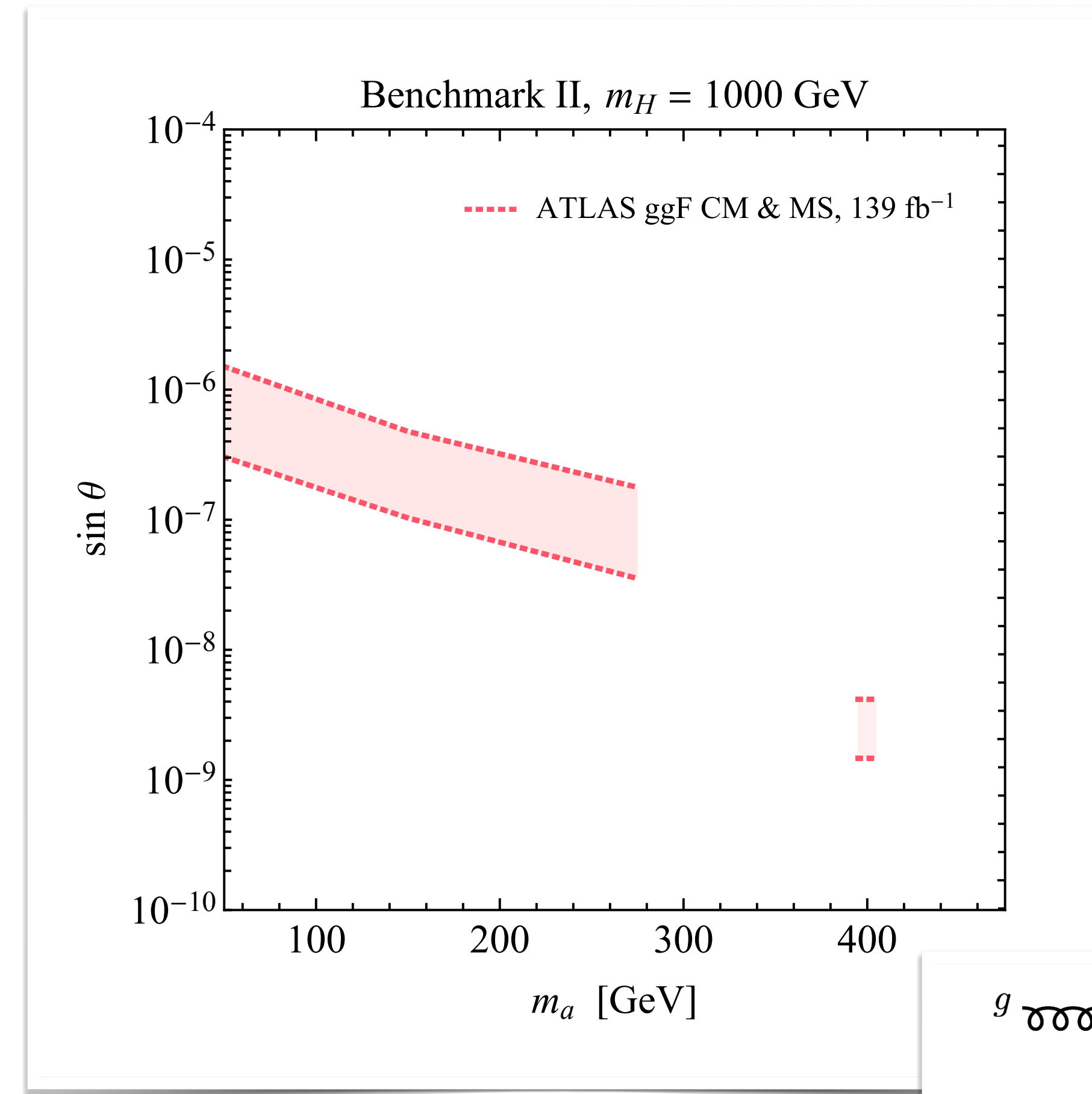
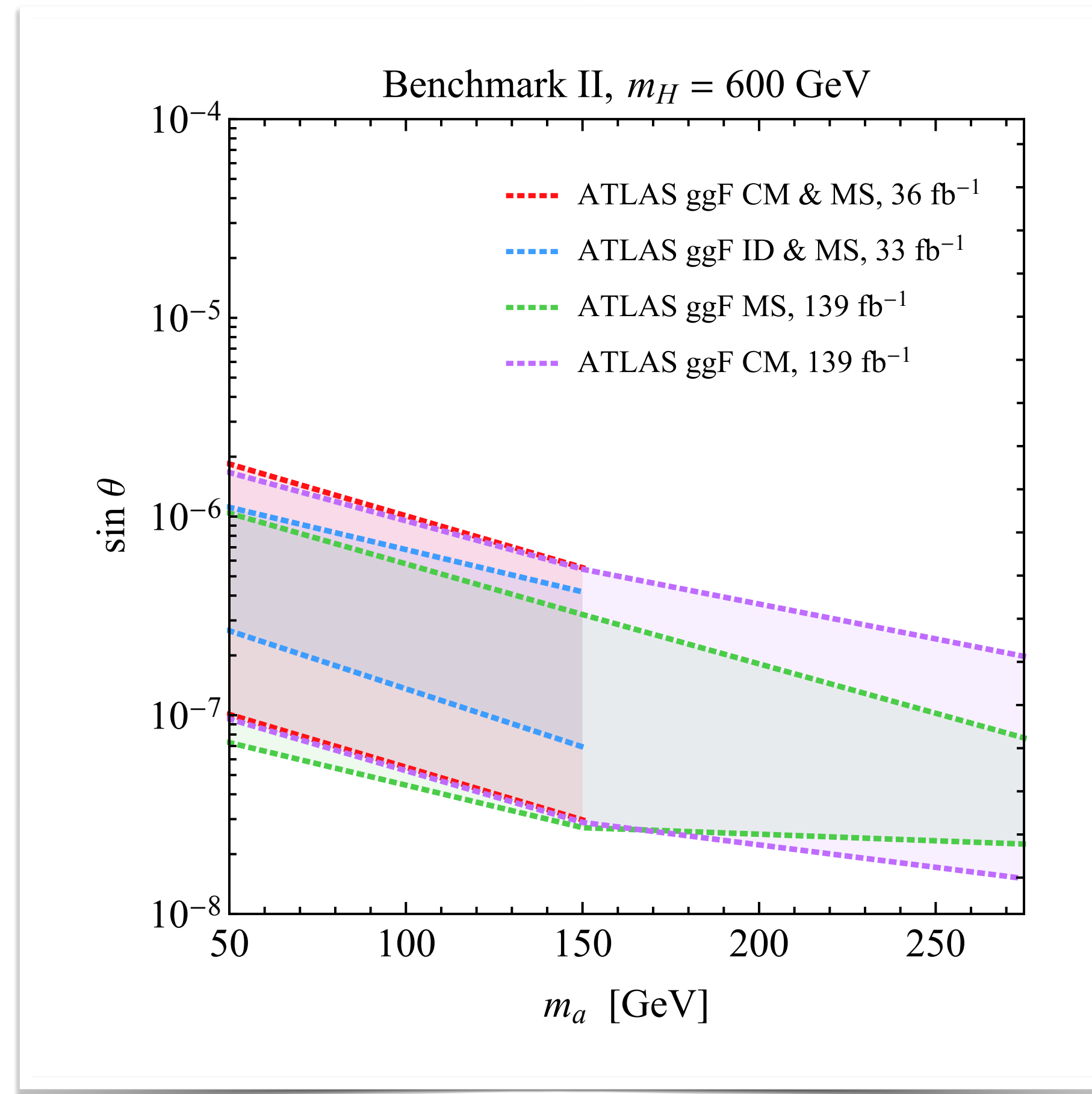
Two narrow, high-multiplicity jets in **MS**. Background (punch-through jets, noncollision background) reduced by requiring two displaced vertices.

2. LLP Phenomenology

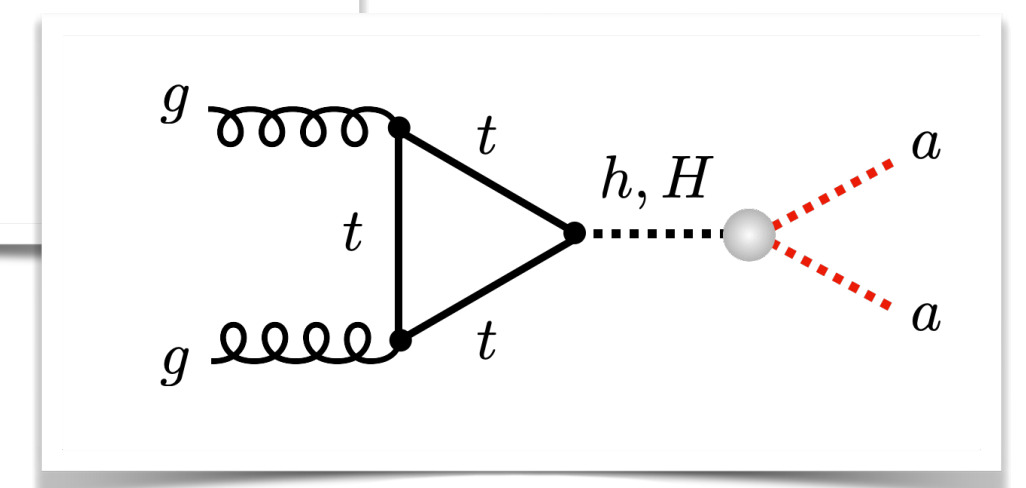
2.2 LLP constraints

2. LLP Phenomenology

2.2 LLP constraints



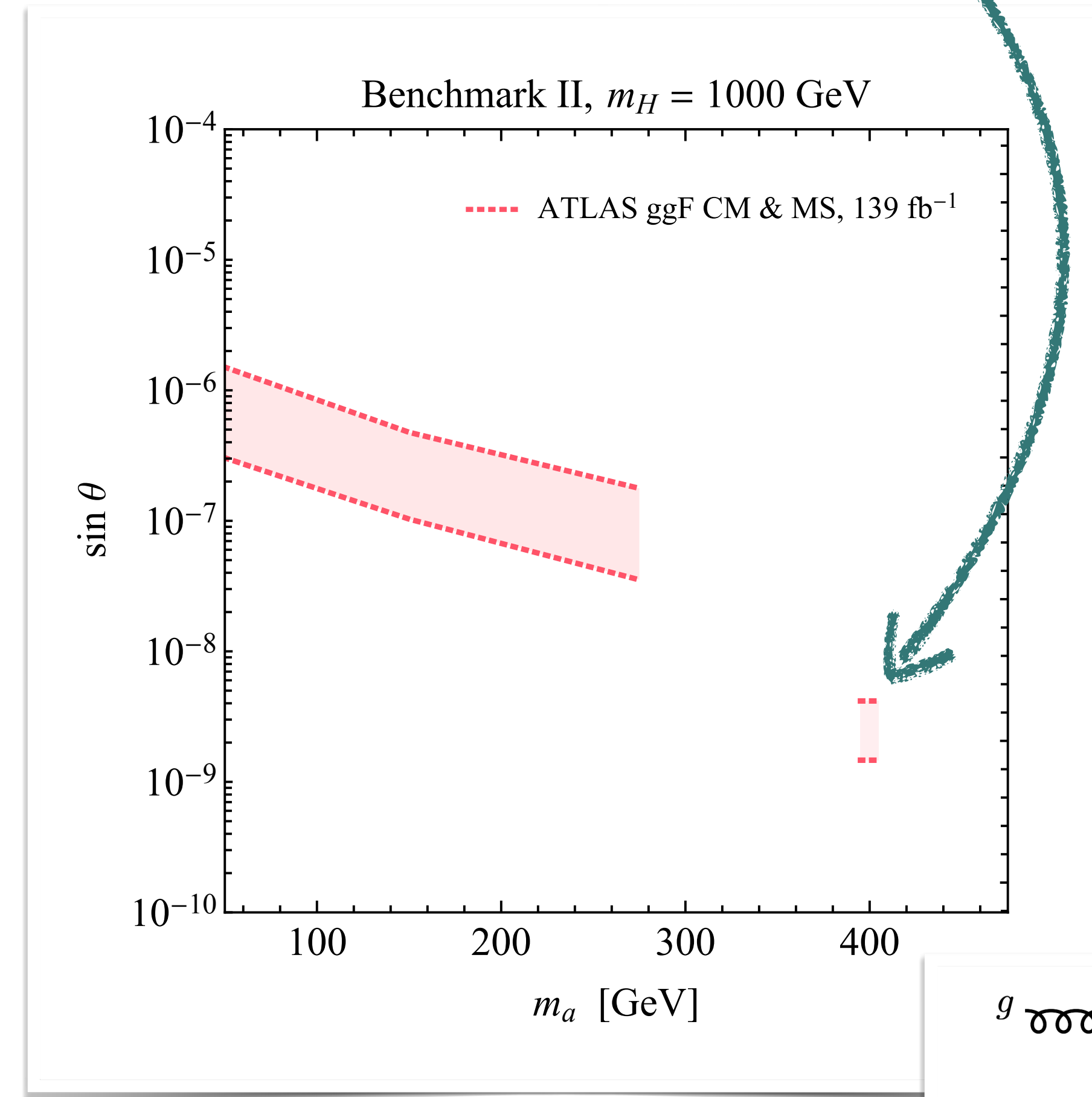
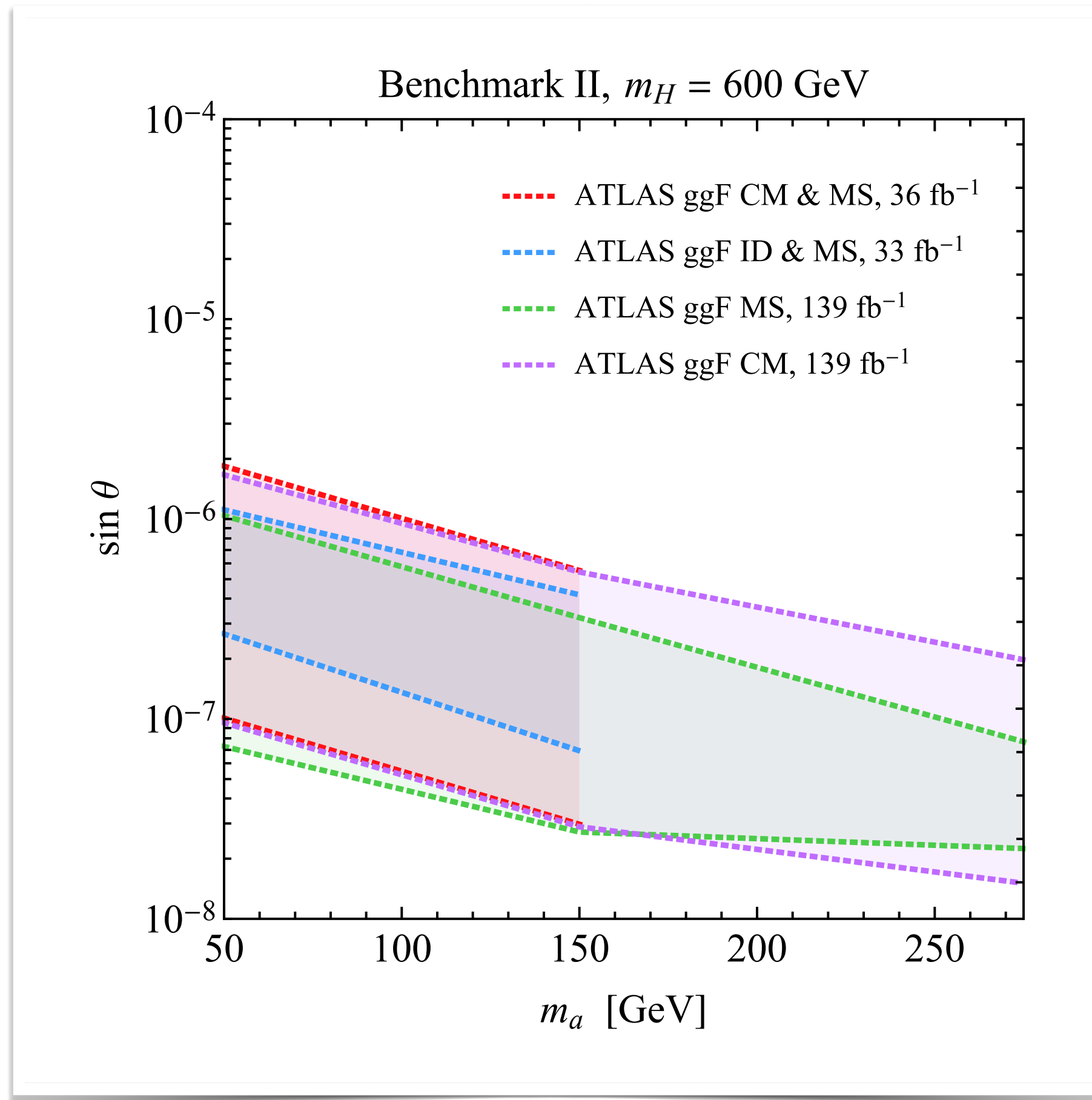
Source: [ArXiv:2302.02735](https://arxiv.org/abs/2302.02735) (U. Haisch, LS).



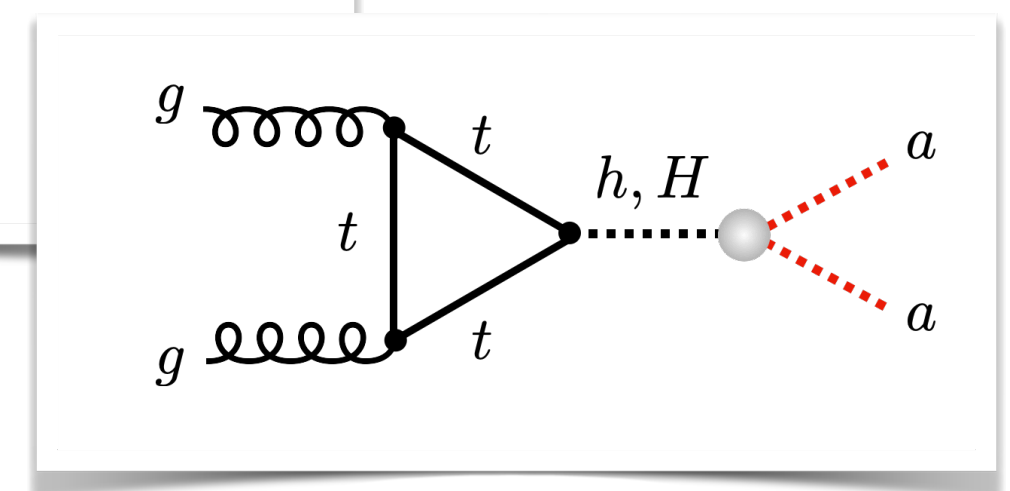
2. LLP Phenomenology

2.2 LLP constraints

$$\Gamma(a \rightarrow f\bar{f}) = \frac{N_c^f \eta_f^2 y_f^2}{16\pi} m_a \sqrt{1 - \frac{4m_f^2}{m_a^2}} \sin^2 \theta.$$



Decay
 $a \rightarrow t\bar{t}$
opens up



Source: [ArXiv:2302.02735](https://arxiv.org/abs/2302.02735) (U. Haisch, LS).

2. LLP Phenomenology

2.3 Relic density

Physical fields: h, H, a, A, H^\pm and χ .

Physical parameters: $\alpha, \beta, \theta, \nu, \lambda_3, \lambda_{P1}, \lambda_{P2}, m_h, m_H, m_a, m_A, m_{H^\pm}$.

α : mixing angle for scalars (h, H)

θ : mixing angle for pseudo-scalars (a, A)

2. LLP Phenomenology

2.3 Relic density

- Can we get the **DM relic density** $\Omega h^2 = 0.120(1)$ right?

Physical fields: h, H, a, A, H^\pm and χ .

Physical parameters: $\alpha, \beta, \theta, \nu, \lambda_3, \lambda_{P1}, \lambda_{P2}, m_h, m_H, m_a, m_A, m_{H^\pm}$.

α : mixing angle for scalars (h, H)

θ : mixing angle for pseudo-scalars (a, A)

2. LLP Phenomenology

2.3 Relic density

- Can we get the **DM relic density** $\Omega h^2 = 0.120(1)$ right?
- DM density evolution („freeze-out“):

$$\frac{dn_\chi}{dt} + 3Hn_\chi = - \langle \sigma v_{rel} \rangle \left(n_\chi^2 - n_\chi^{(eq)2} \right)$$

Physical fields: h, H, a, A, H^\pm and χ .

Physical parameters: $\alpha, \beta, \theta, v, \lambda_3, \lambda_{P1}, \lambda_{P2}, m_h, m_H, m_a, m_A, m_{H^\pm}$.

α : mixing angle for scalars (h, H)

θ : mixing angle for pseudo-scalars (a, A)

2. LLP Phenomenology

2.3 Relic density

- Can we get the **DM relic density** $\Omega h^2 = 0.120(1)$ right?
- DM density evolution („freeze-out“):

$$\frac{dn_\chi}{dt} + 3Hn_\chi = - \langle \sigma v_{rel} \rangle \left(n_\chi^2 - n_\chi^{(eq)2} \right)$$
$$n_\chi/T^3 \sim \exp(-m_\chi/T)$$

Physical fields: h, H, a, A, H^\pm and χ .

Physical parameters: $\alpha, \beta, \theta, v, \lambda_3, \lambda_{P1}, \lambda_{P2}, m_h, m_H, m_a, m_A, m_{H^\pm}$.

α : mixing angle for scalars (h, H)

θ : mixing angle for pseudo-scalars (a, A)

2. LLP Phenomenology

2.3 Relic density

- Can we get the **DM relic density** $\Omega h^2 = 0.120(1)$ right?
- DM density evolution („freeze-out“):

$$\frac{dn_\chi}{dt} + 3Hn_\chi = - \langle \sigma v_{rel} \rangle \left(n_\chi^2 - n_\chi^{(eq)2} \right) \quad n_\chi/T^3 \sim \exp(-m_\chi/T)$$

\rightarrow freeze-out: $n_\chi \langle \sigma v_{rel} \rangle \sim H$

Physical fields: h, H, a, A, H^\pm and χ .

Physical parameters: $\alpha, \beta, \theta, v, \lambda_3, \lambda_{P1}, \lambda_{P2}, m_h, m_H, m_a, m_A, m_{H^\pm}$.

α : mixing angle for scalars (h, H)

θ : mixing angle for pseudo-scalars (a, A)

2. LLP Phenomenology

2.3 Relic density

- Can we get the **DM relic density** $\Omega h^2 = 0.120(1)$ right?
- DM density evolution („freeze-out“):

$$\frac{dn_\chi}{dt} + 3Hn_\chi = - \langle \sigma v_{rel} \rangle \left(n_\chi^2 - n_\chi^{(eq)2} \right)$$

$n_\chi/T^3 \sim \exp(-m_\chi/T)$
 \rightarrow freeze-out: $n_\chi \langle \sigma v_{rel} \rangle \sim H$
 $n_\chi/T^3 \equiv \text{const.}$

Physical fields: h, H, a, A, H^\pm and χ .

Physical parameters: $\alpha, \beta, \theta, v, \lambda_3, \lambda_{P1}, \lambda_{P2}, m_h, m_H, m_a, m_A, m_{H^\pm}$.

α : mixing angle for scalars (h, H)

θ : mixing angle for pseudo-scalars (a, A)

2. LLP Phenomenology

2.3 Relic density

- Can we get the **DM relic density** $\Omega h^2 = 0.120(1)$ right?
- DM density evolution („freeze-out“):

$$\frac{dn_\chi}{dt} + 3Hn_\chi = - \langle \sigma v_{rel} \rangle \left(n_\chi^2 - n_\chi^{(eq)2} \right)$$

$$n_\chi/T^3 \sim \exp(-m_\chi/T)$$

$$\rightarrow \text{freeze-out: } n_\chi \langle \sigma v_{rel} \rangle \sim H$$

$$n_\chi/T^3 \equiv \text{const.}$$

$$\frac{\Omega h^2}{0.12} \simeq \frac{1.6 \cdot 10^{-10} \text{ GeV}^{-2} x_f}{\langle \sigma v_{rel} \rangle_f},$$

$$\langle \sigma v_{rel} \rangle_f \simeq \frac{y_\chi^2}{128\pi m_\chi^2} \left[\frac{(g_{haa}^2 + g_{Haa}^2) v^2}{4m_\chi^2} + \frac{y_\chi^2}{x_f} \right].$$

(hold for $m_\chi \gg m_a, m_h, m_H$)

Physical fields: h, H, a, A, H^\pm and χ .

Physical parameters: $\alpha, \beta, \theta, v, \lambda_3, \lambda_{P1}, \lambda_{P2}, m_h, m_H, m_a, m_A, m_{H^\pm}$.

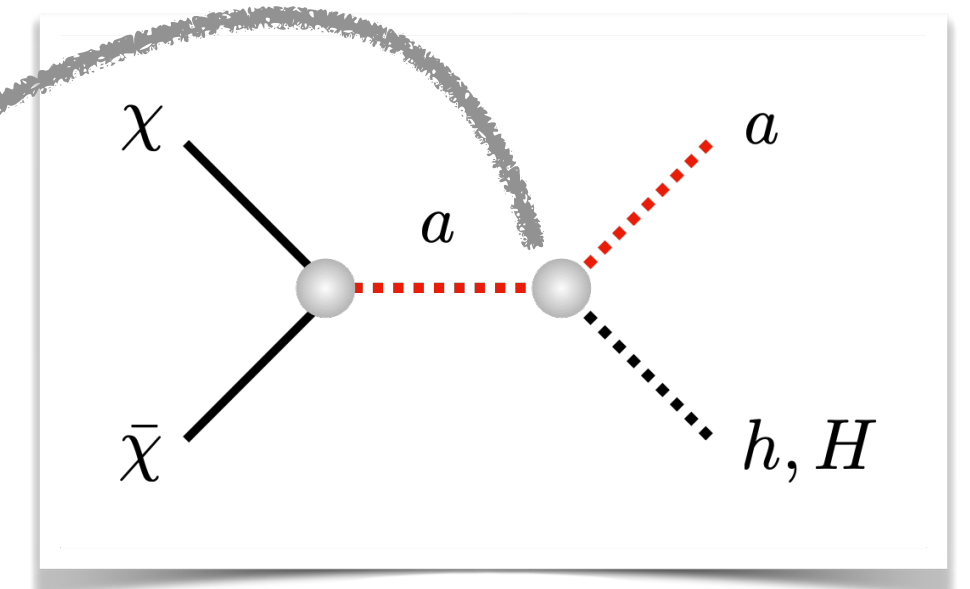
α : mixing angle for scalars (h, H)

θ : mixing angle for pseudo-scalars (a, A)

2. LLP Phenomenology

2.3 Relic density

- Can we get the **DM relic density** $\Omega h^2 = 0.120(1)$ right?
- DM density evolution („freeze-out“):



$$\frac{dn_\chi}{dt} + 3Hn_\chi = - \langle \sigma v_{rel} \rangle \left(n_\chi^2 - n_\chi^{(eq)2} \right)$$

$$n_\chi/T^3 \sim \exp(-m_\chi/T)$$

$$\rightarrow \text{freeze-out: } n_\chi \langle \sigma v_{rel} \rangle \sim H$$

$$n_\chi/T^3 \equiv \text{const.}$$

$$\frac{\Omega h^2}{0.12} \simeq \frac{1.6 \cdot 10^{-10} \text{ GeV}^{-2} x_f}{\langle \sigma v_{rel} \rangle_f},$$

$$\langle \sigma v_{rel} \rangle_f \simeq \frac{y_\chi^2}{128\pi m_\chi^2} \left[\frac{(g_{haa}^2 + g_{Haa}^2) v^2}{4m_\chi^2} + \frac{y_\chi^2}{x_f} \right].$$

(hold for $m_\chi \gg m_a, m_h, m_H$)

Physical fields: h, H, a, A, H^\pm and χ .

Physical parameters: $\alpha, \beta, \theta, v, \lambda_3, \lambda_{P1}, \lambda_{P2}, m_h, m_H, m_a, m_A, m_{H^\pm}$.

α : mixing angle for scalars (h, H)

θ : mixing angle for pseudo-scalars (a, A)

2. LLP Phenomenology

2.3 Relic density

- Can we get the **DM relic density** $\Omega h^2 = 0.120(1)$ right?
- DM density evolution („freeze-out“):

$$\frac{dn_\chi}{dt} + 3Hn_\chi = - \langle \sigma v_{rel} \rangle \left(n_\chi^2 - n_\chi^{(eq)2} \right)$$

$$n_\chi/T^3 \sim \exp(-m_\chi/T)$$

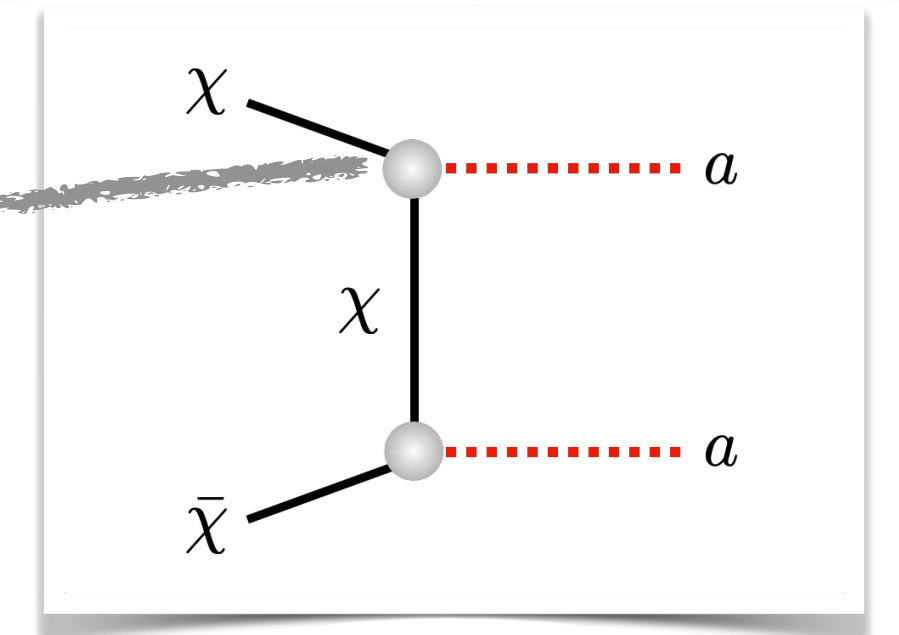
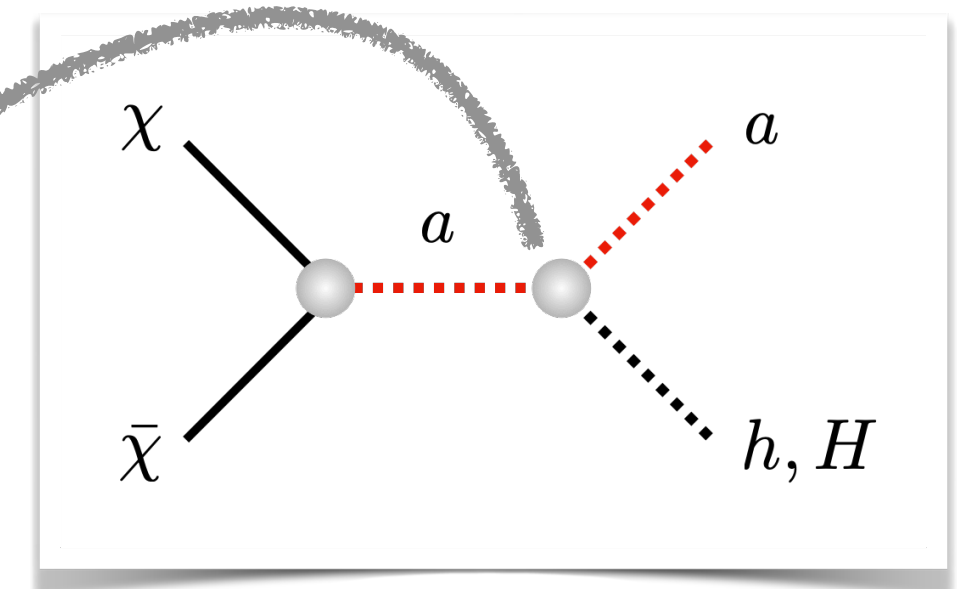
$$\rightarrow \text{freeze-out: } n_\chi \langle \sigma v_{rel} \rangle \sim H$$

$$n_\chi/T^3 \equiv \text{const.}$$

$$\frac{\Omega h^2}{0.12} \simeq \frac{1.6 \cdot 10^{-10} \text{ GeV}^{-2} x_f}{\langle \sigma v_{rel} \rangle_f},$$

$$\langle \sigma v_{rel} \rangle_f \simeq \frac{y_\chi^2}{128\pi m_\chi^2} \left[\frac{(g_{haa}^2 + g_{Haa}^2) v^2}{4m_\chi^2} + \frac{y_\chi^2}{x_f} \right].$$

(hold for $m_\chi \gg m_a, m_h, m_H$)



Source: [ArXiv:2302.02735](https://arxiv.org/abs/2302.02735) (U. Haisch, LS).

Physical fields: h, H, a, A, H^\pm and χ .

Physical parameters: $\alpha, \beta, \theta, v, \lambda_3, \lambda_{P1}, \lambda_{P2}, m_h, m_H, m_a, m_A, m_{H^\pm}$.

α : mixing angle for scalars (h, H)

θ : mixing angle for pseudo-scalars (a, A)

2. LLP Phenomenology

2.3 Relic density

2. LLP Phenomenology

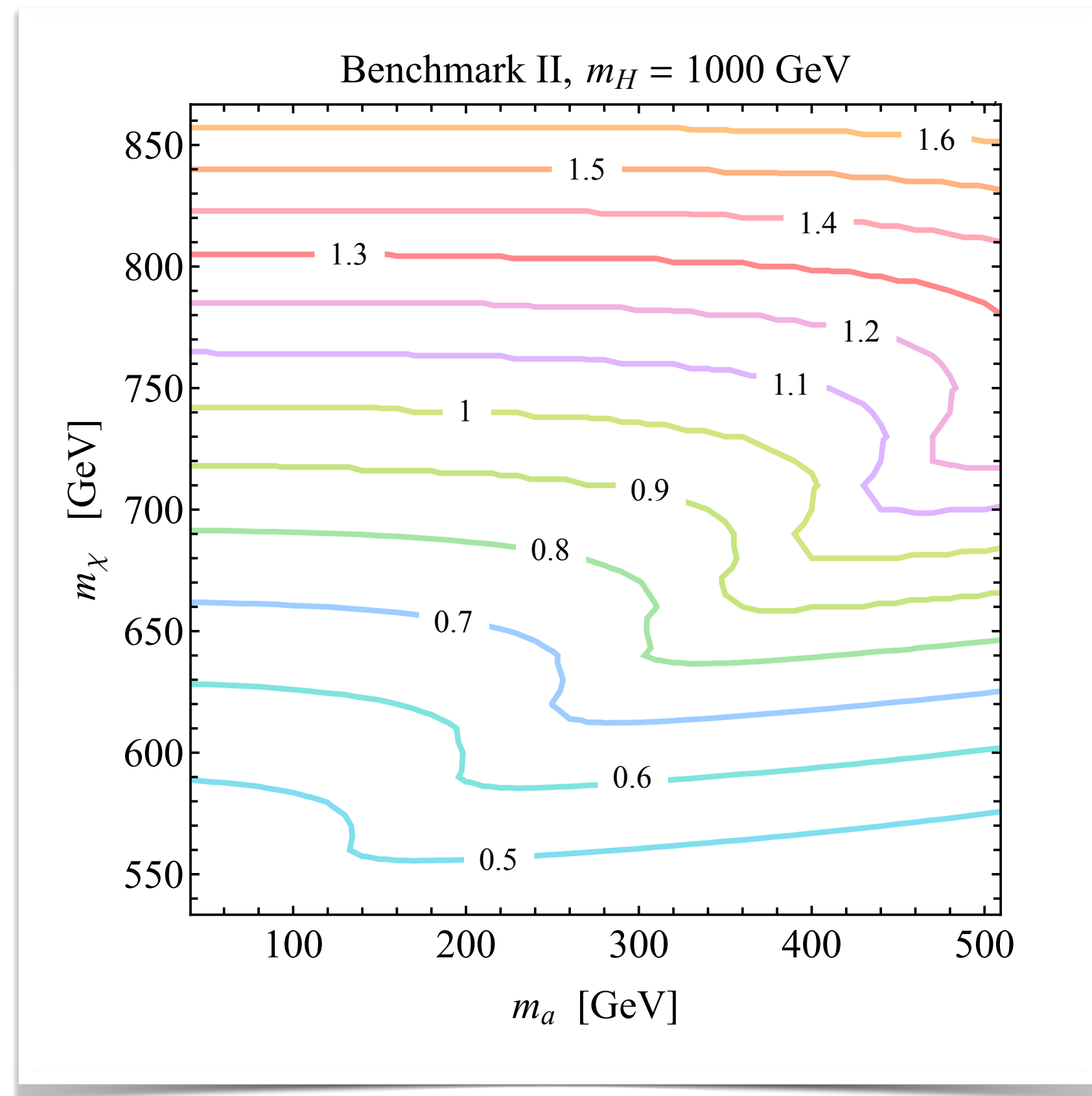
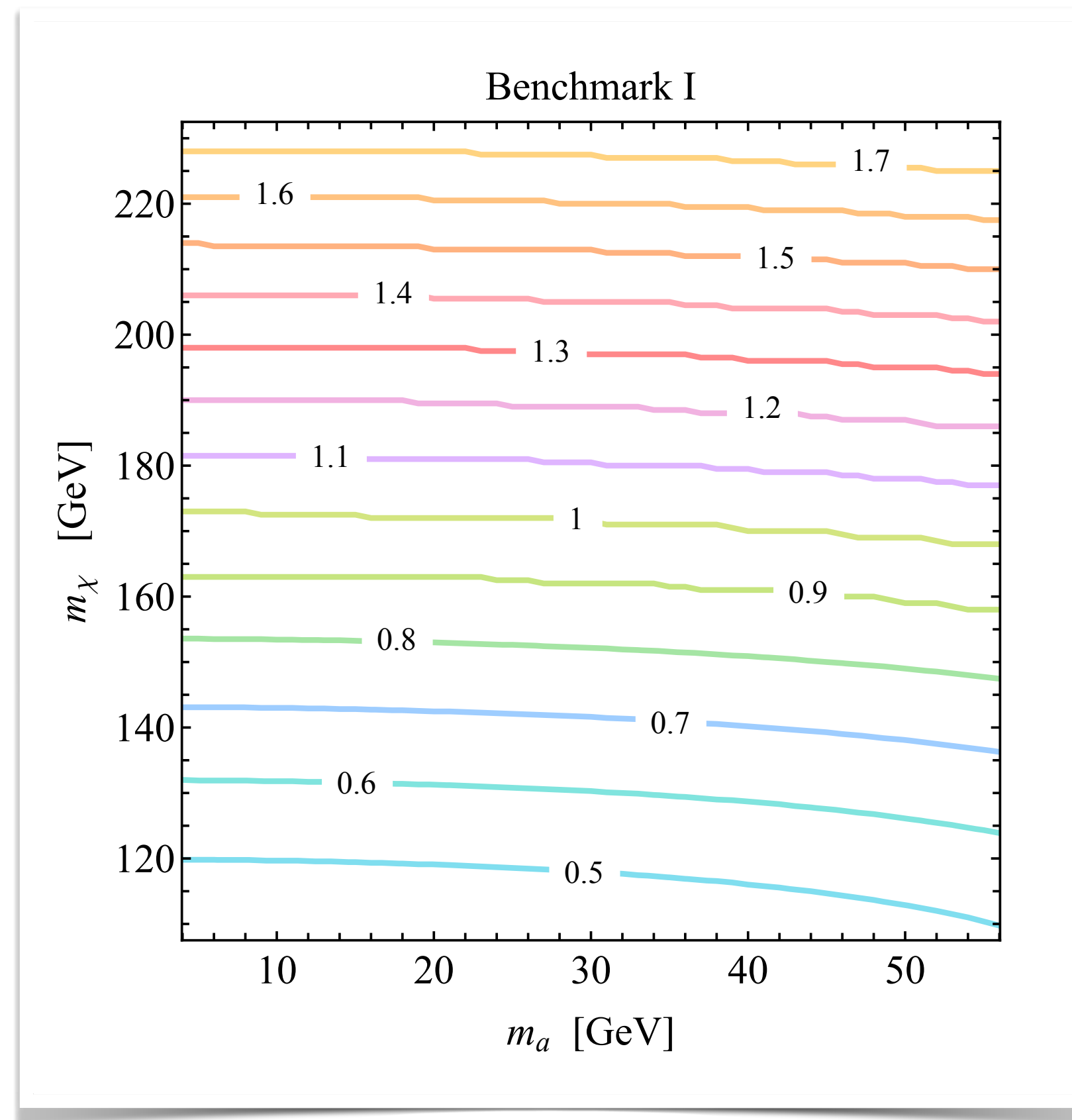
2.3 Relic density

- Can we get the **DM relic density** $\Omega h^2 = 0.120(1)$ right?

2. LLP Phenomenology

2.3 Relic density

- Can we get the **DM relic density** $\Omega h^2 = 0.120(1)$ right?

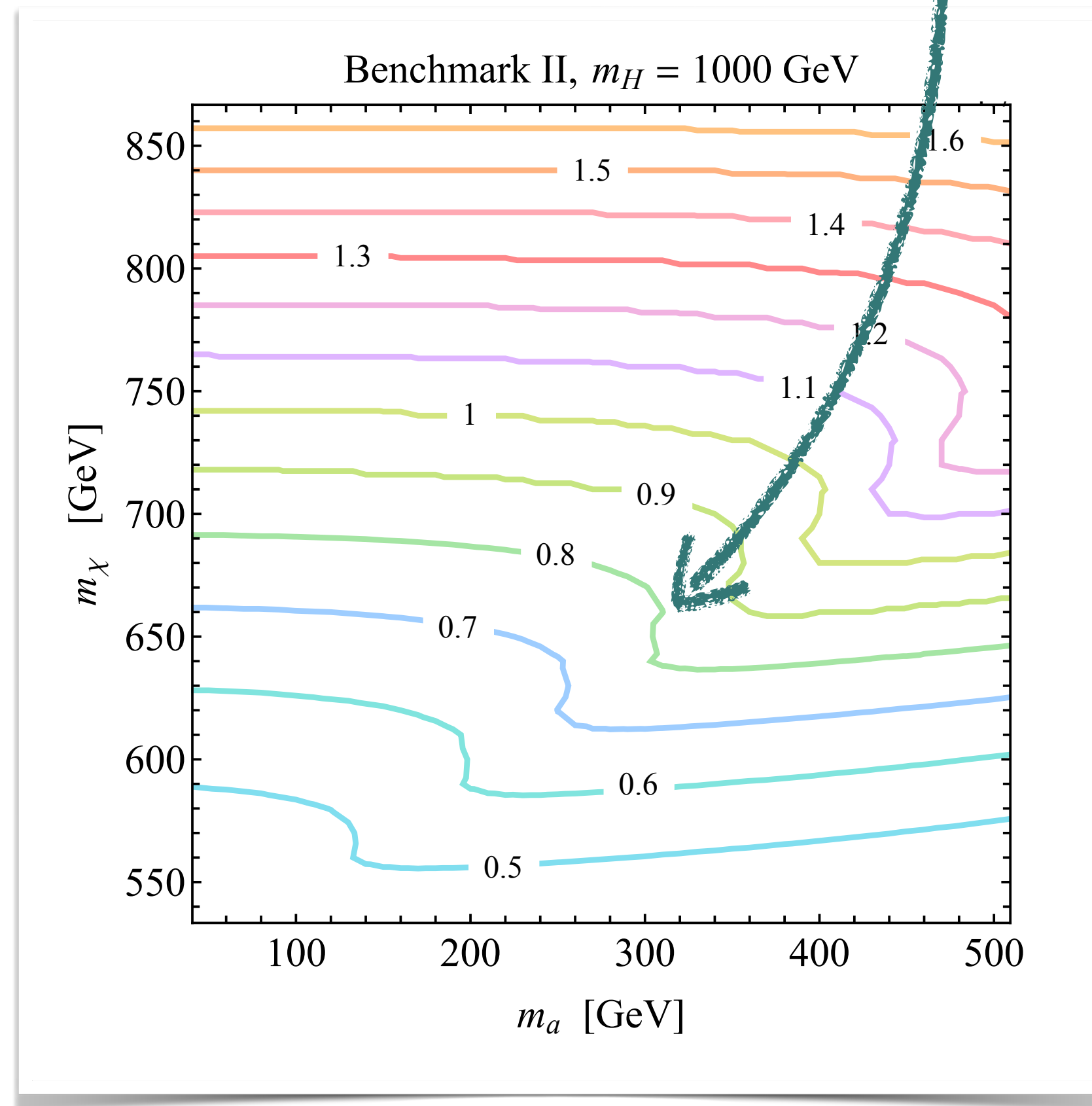
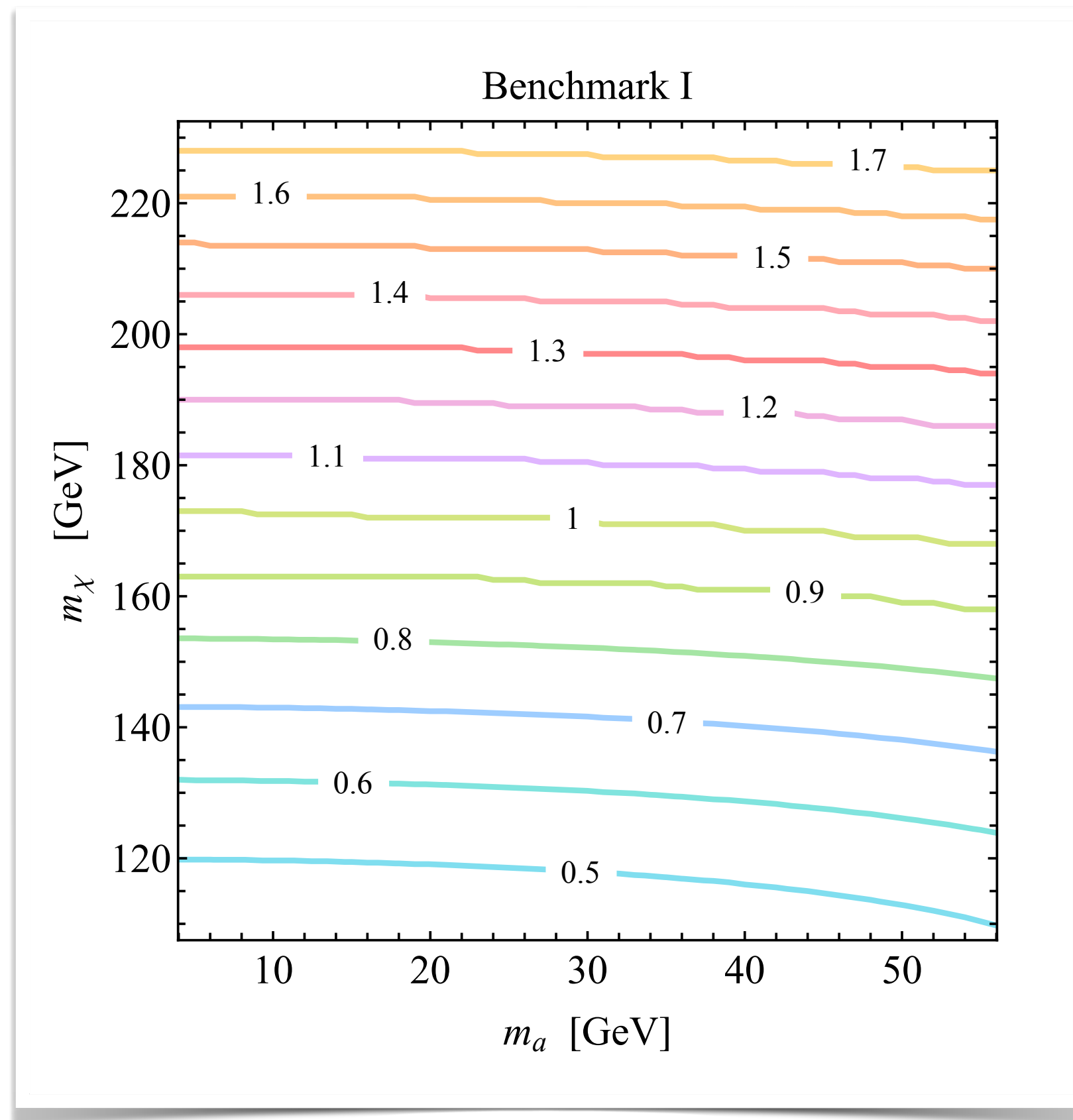


2. LLP Phenomenology

2.3 Relic density

- Can we get the **DM relic density** $\Omega h^2 = 0.120(1)$ right?

$m_\chi \gg m_H$ no longer true
→ we get more intricate behaviour modelled with MadDM.



3. Conclusions

3. Conclusions

3. Conclusions

- The 2HDM+ a model combines an **extended scalar sector** (2HDM) with a UV-complete **pseudoscalar DM mediator scenario**.

This leads to an **interesting collider phenomenology** → **important benchmark**.

3. Conclusions

- The 2HDM+ a model combines an **extended scalar sector** (2HDM) with a UV-complete **pseudoscalar DM mediator scenario**.

This leads to an **interesting collider phenomenology** → **important benchmark**.

- The additional pseudo-scalar a can become **long-lived** for small mixing angles θ .
 - **Interesting LLP signatures** that can be probed for at colliders.
 - This scenario is **compatible with current relic density measurements**.

Thank you for your attention!