

MATHUSLA status update

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on behalf of the MATHUSLA Collaboration

13 LHC LLP workshop

22 June 2023



Gen=T

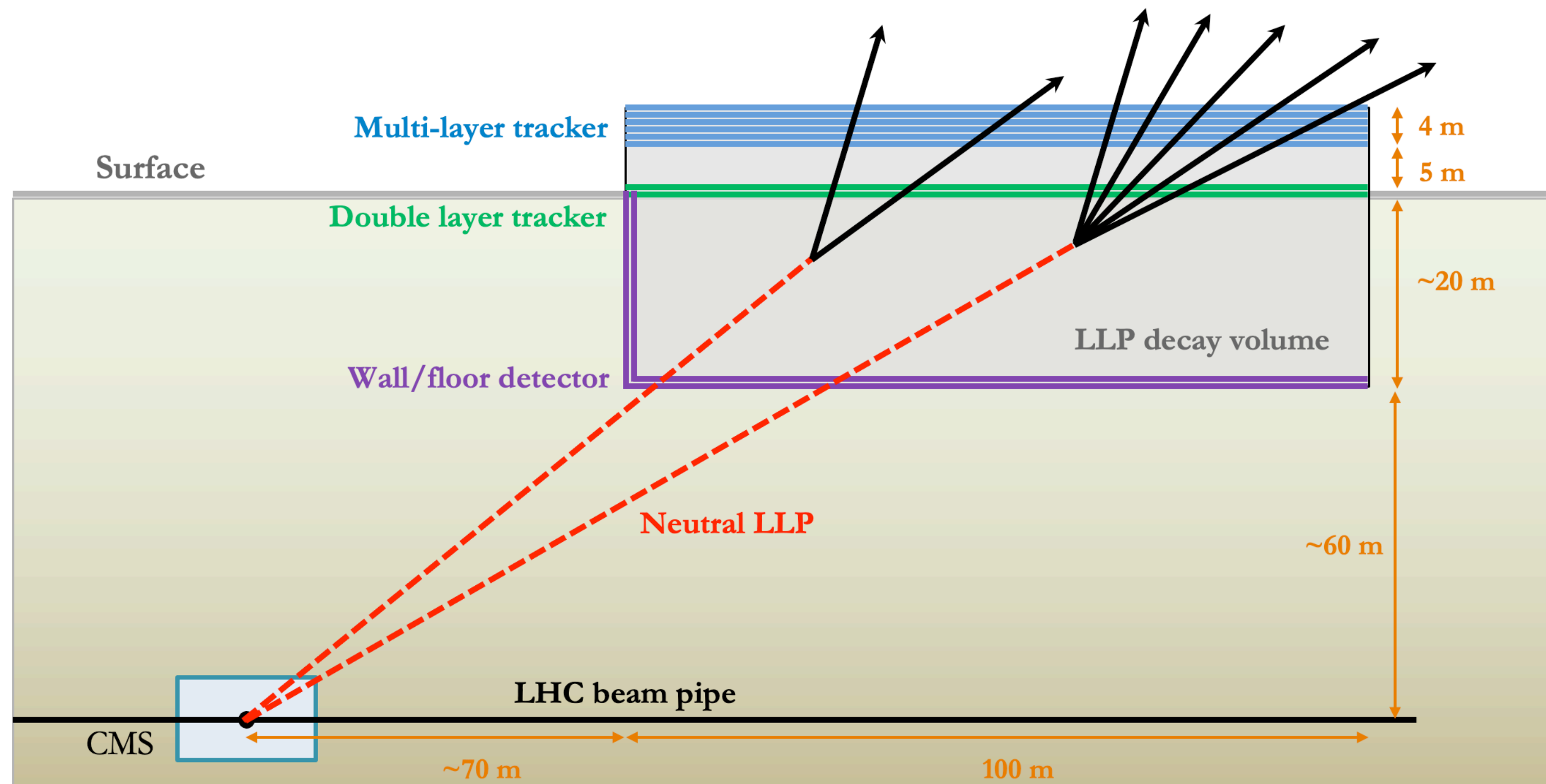


MATHUSLA concept

Updated LHCC Lol: <https://arxiv.org/abs/2009.01693>

MAasive Timing Hodoscope for Ultra Stable neutral pArticles

- Proposed LHC auxiliary transverse LLP Detector
- Target: ultra long-lived particles, to have sensitivity to lifetimes up to 0.1s (BBN limit)
- To be placed on the surface **above CMS**
- To take data **during the entire HL-LHC** operation

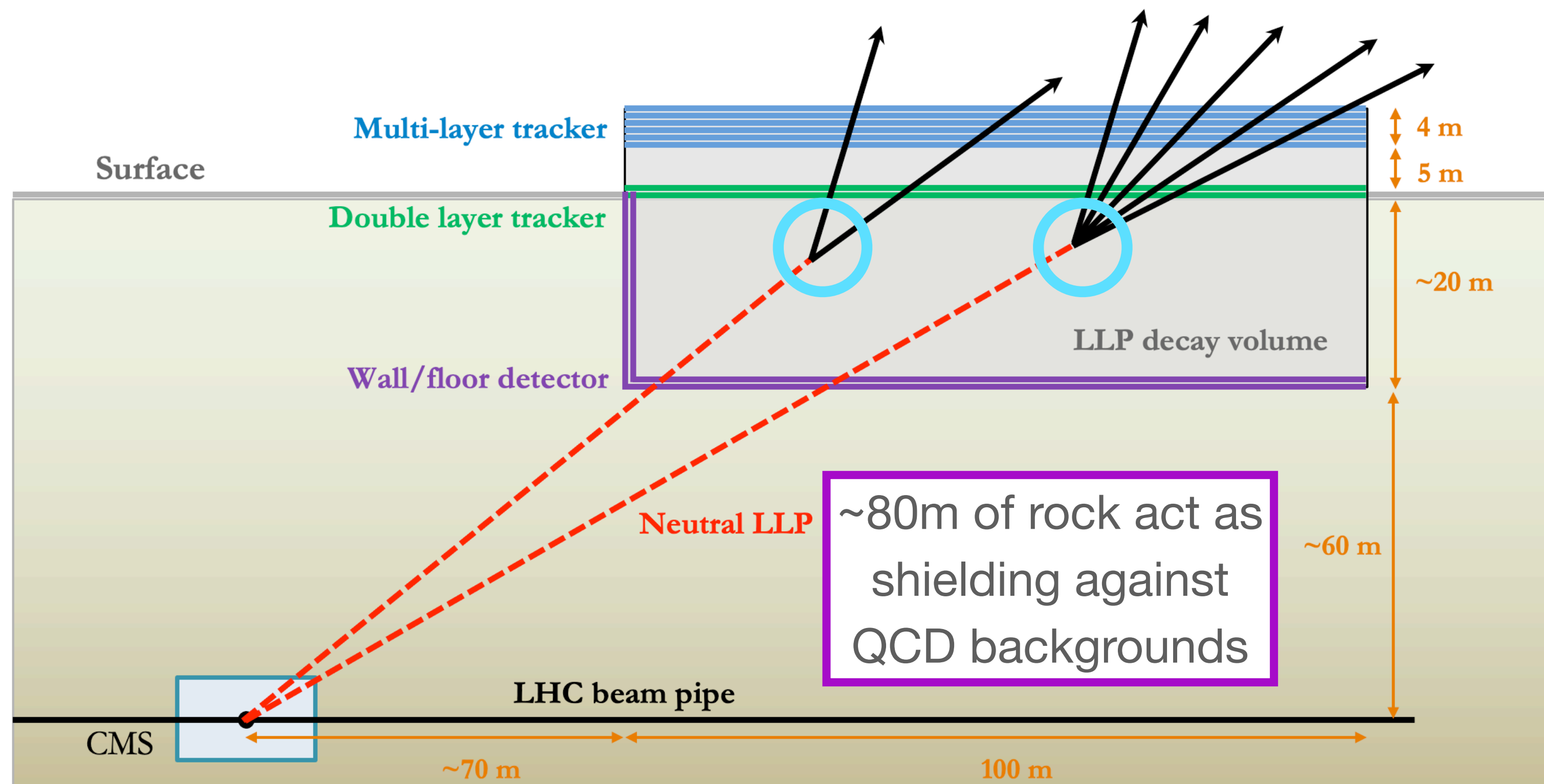


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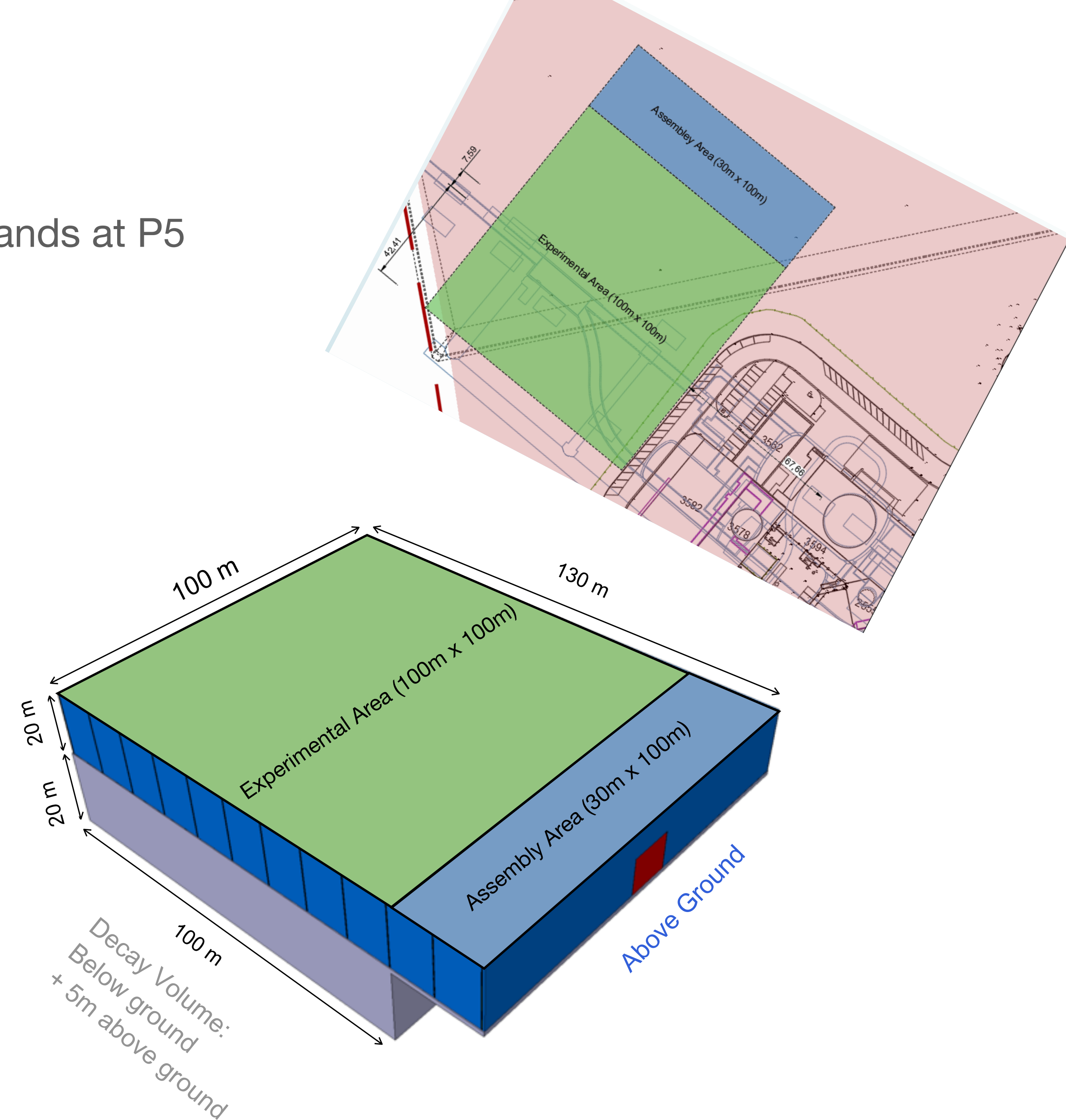
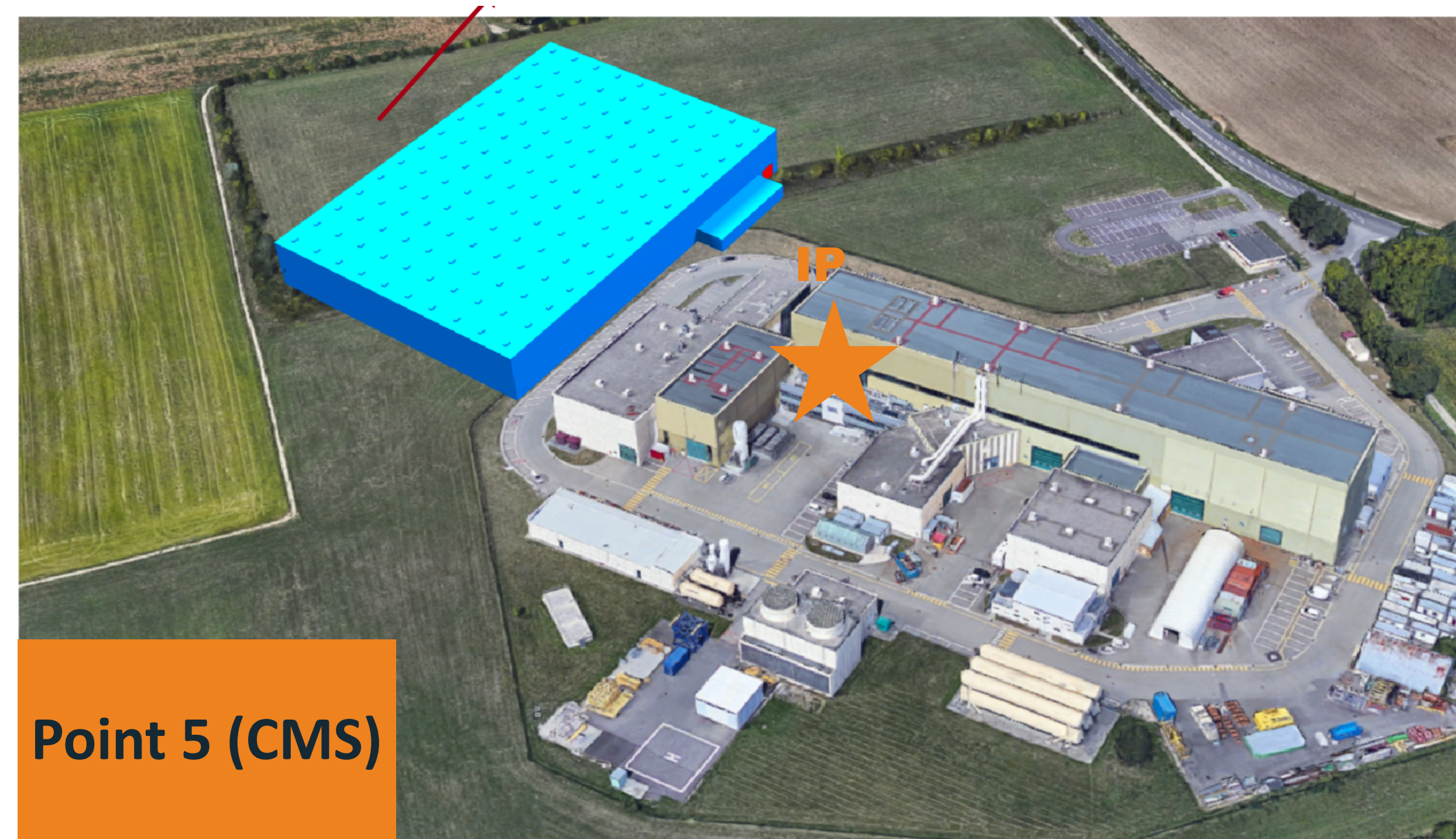
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- Large air decay volume with:
 - walls and floor scintillator layers to flag LHC muons
 - several layers of scintillators for tracking
 - Scintillator timing separates upward tracks from LLP decays from downward cosmic rays
 - LLPs decaying inside MATHUSLA are reconstructed as **displaced vertices** (robust tracking)

MATHUSLA at P5

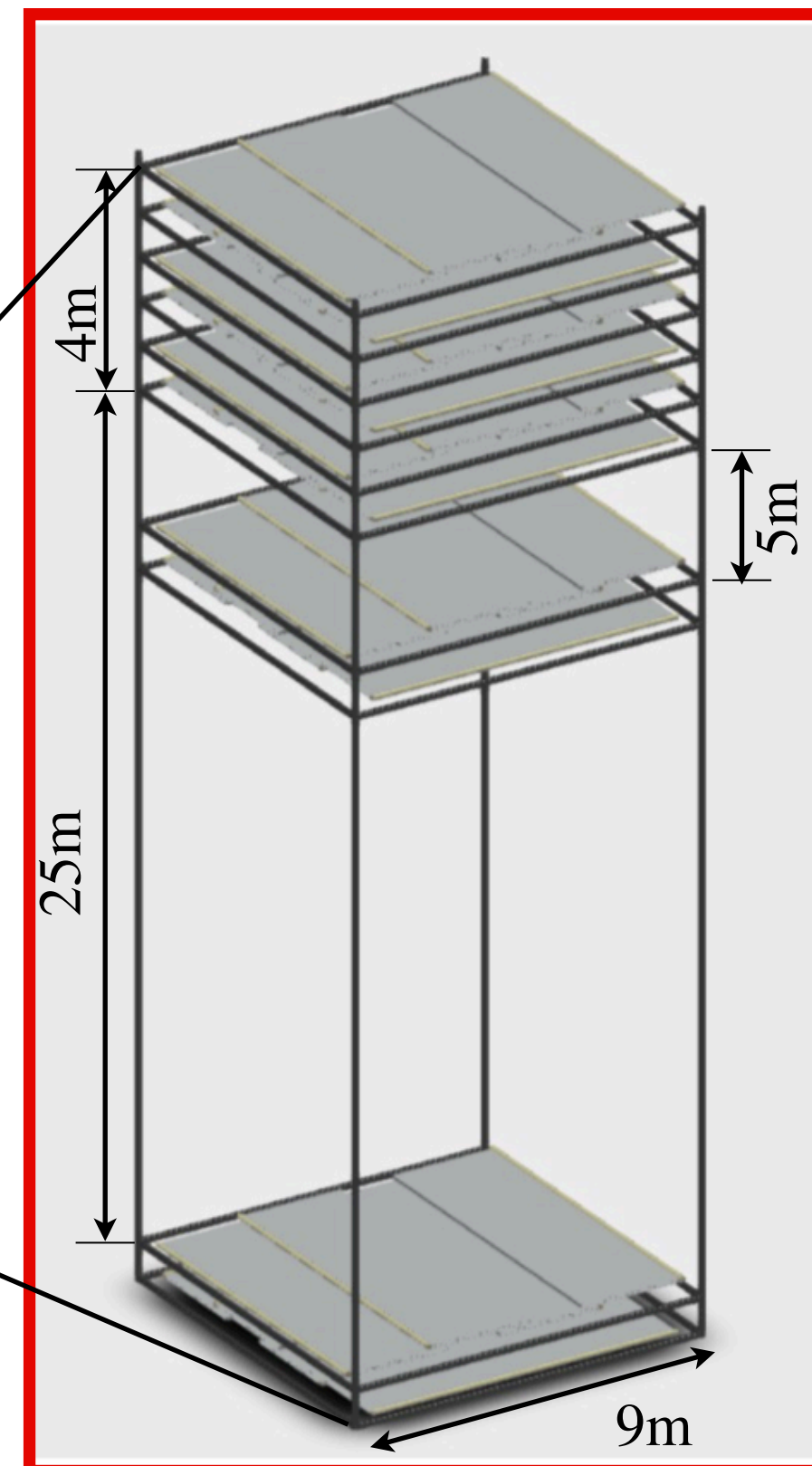
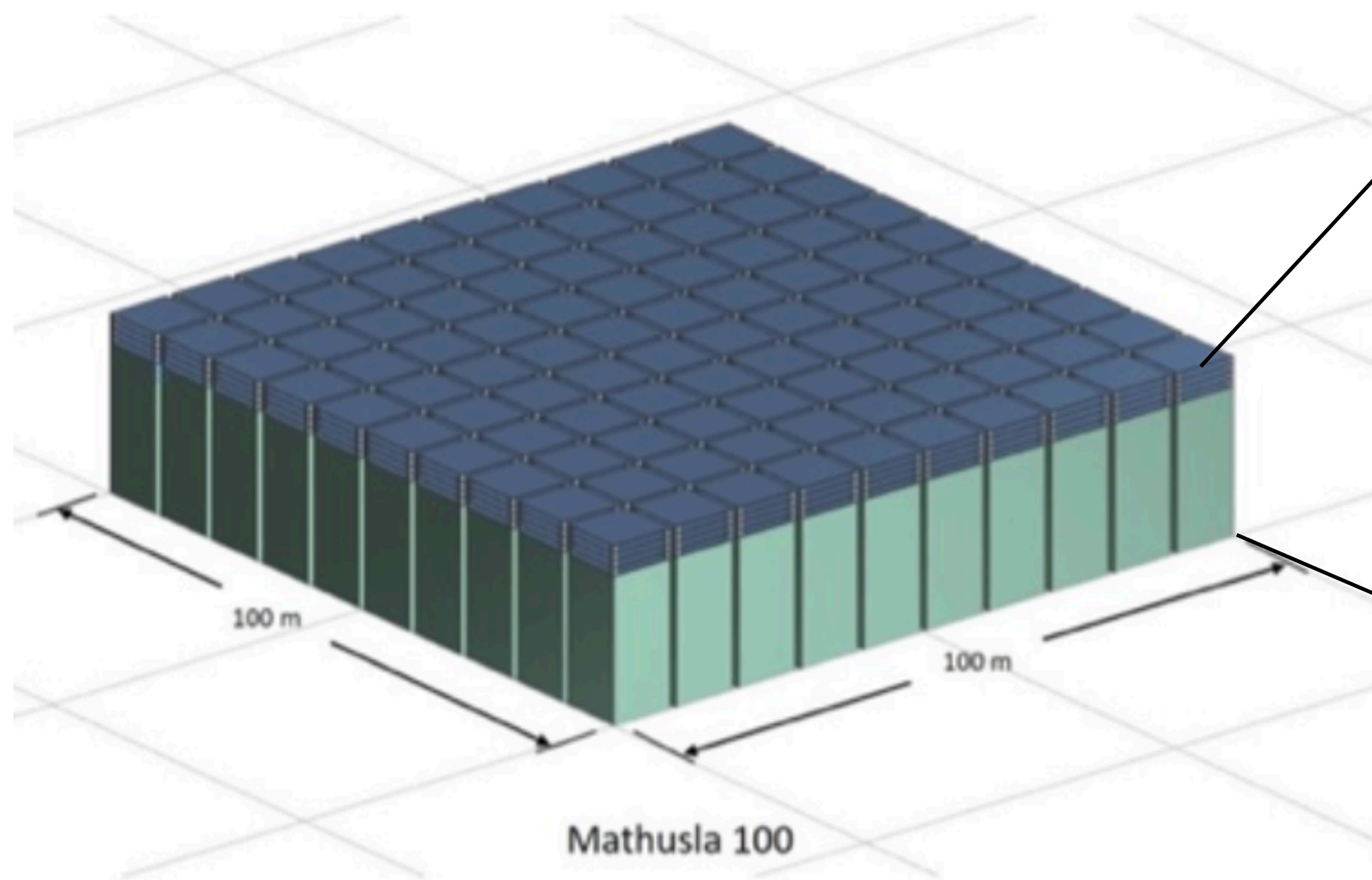
- Proposed building to house MATHUSLA on CERN owned lands at P5



- Layout restricted by existing structures based on current concept and engineering requirements
- Building at the surface extends ~20m below ground

Detector Plane layout

- Decay volume $\sim 100 \times 100 \times 25 \text{ m}^3$
- **Modular** design (100 modules of $9 \times 9 \times 30 \text{ m}^3$)
 - Assembly time line not governed rigidly by HL-LHC beam schedule
 - Data taking can **start after installation of the first module**

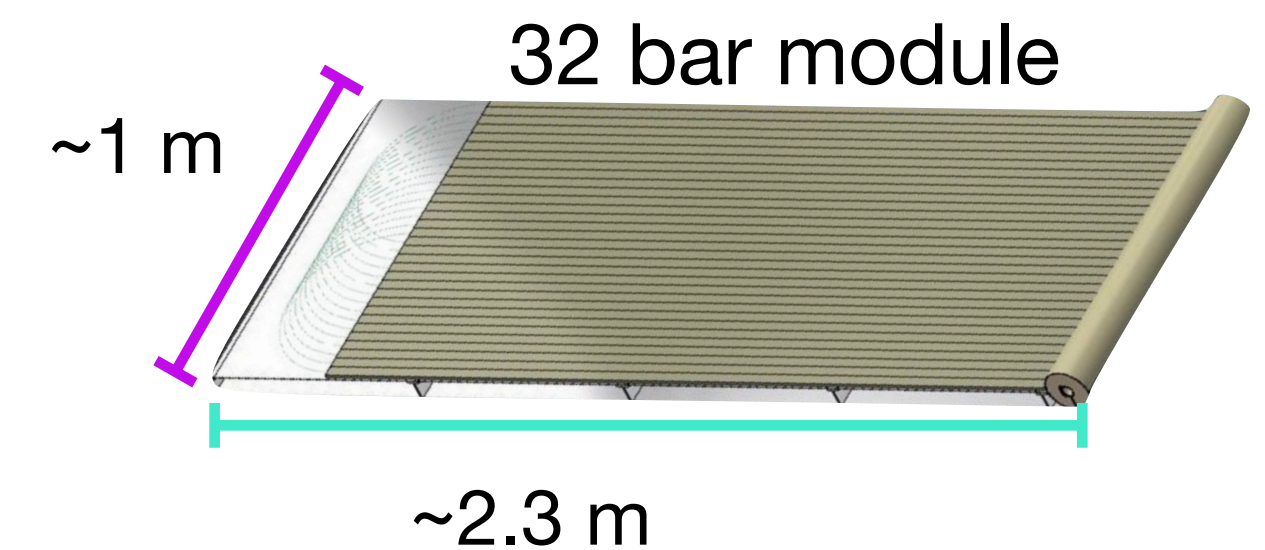
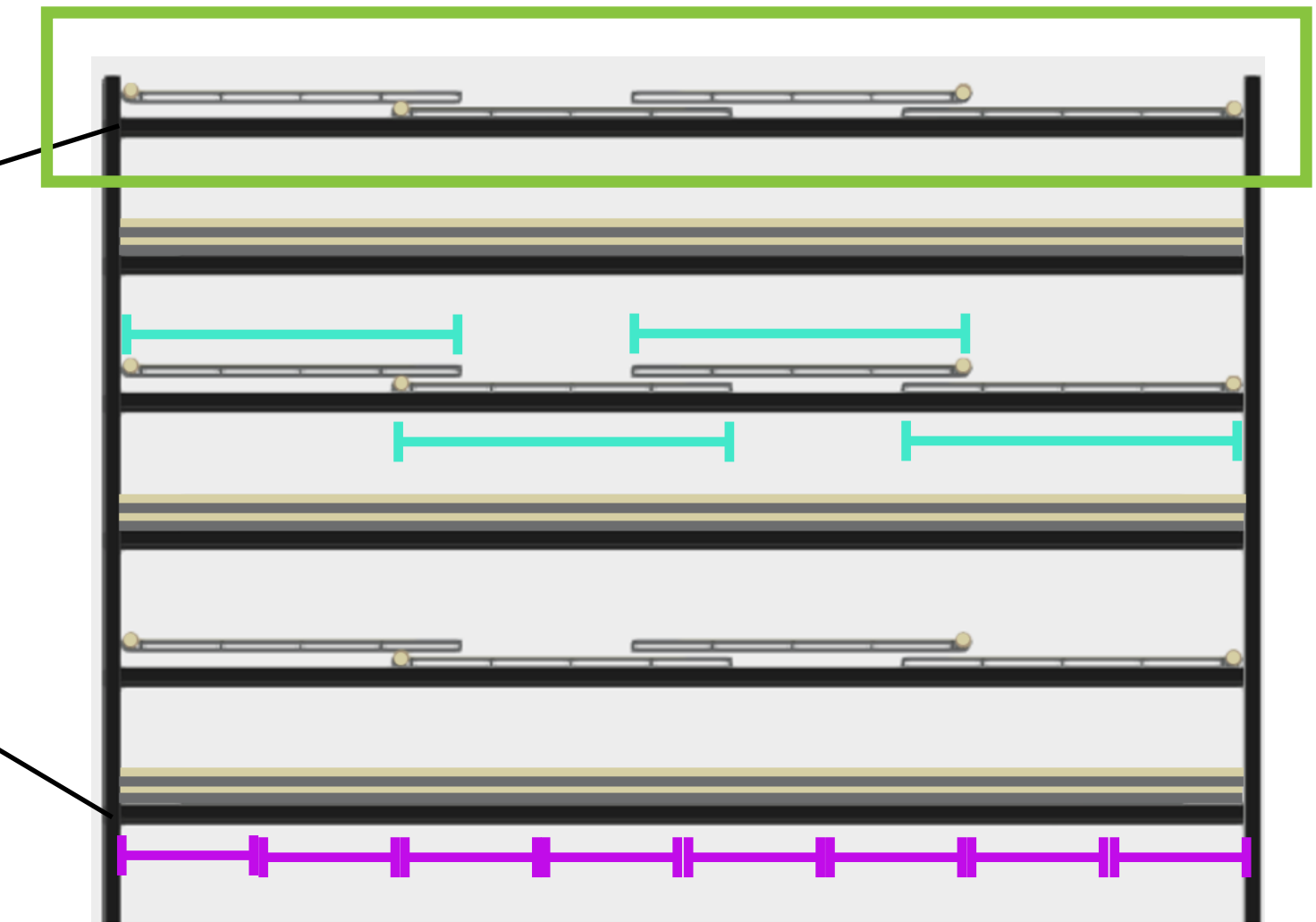
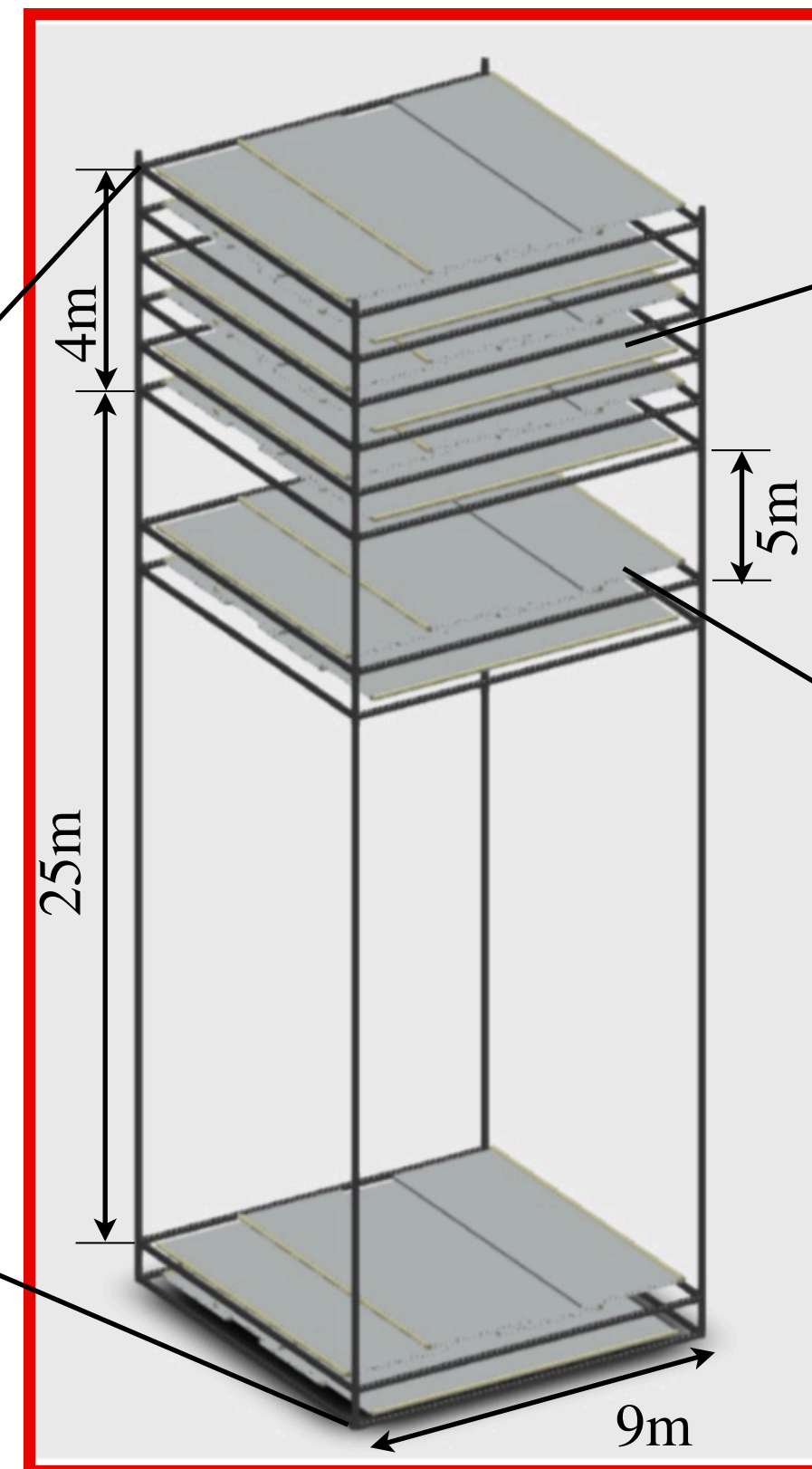
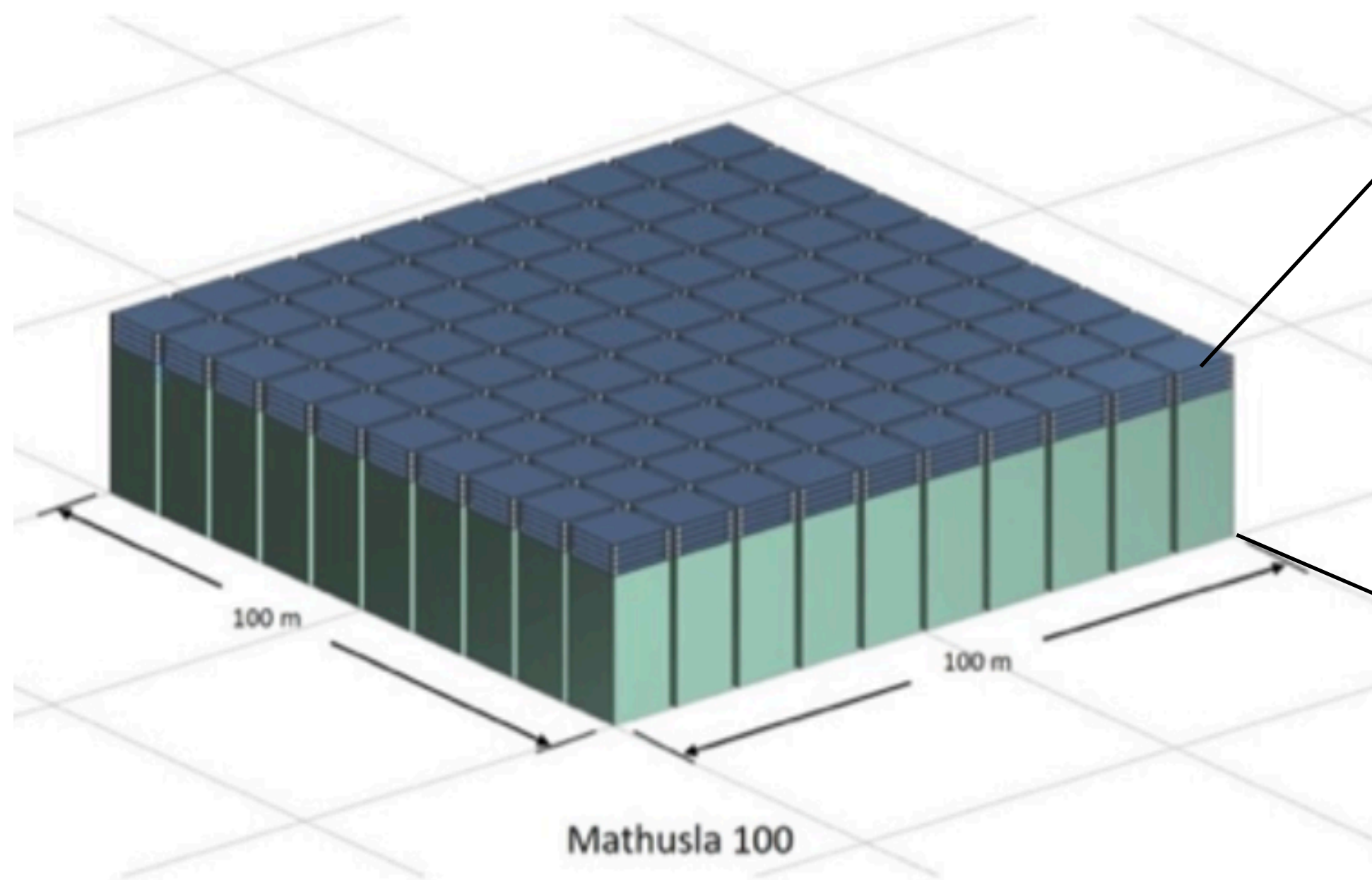


- 6 layers of tracking / timing detector
 - 80 cm between planes
- Additional double layer 5 m below
- Double layer floor detector to veto charged particles from the LHC

Detector Plane layout

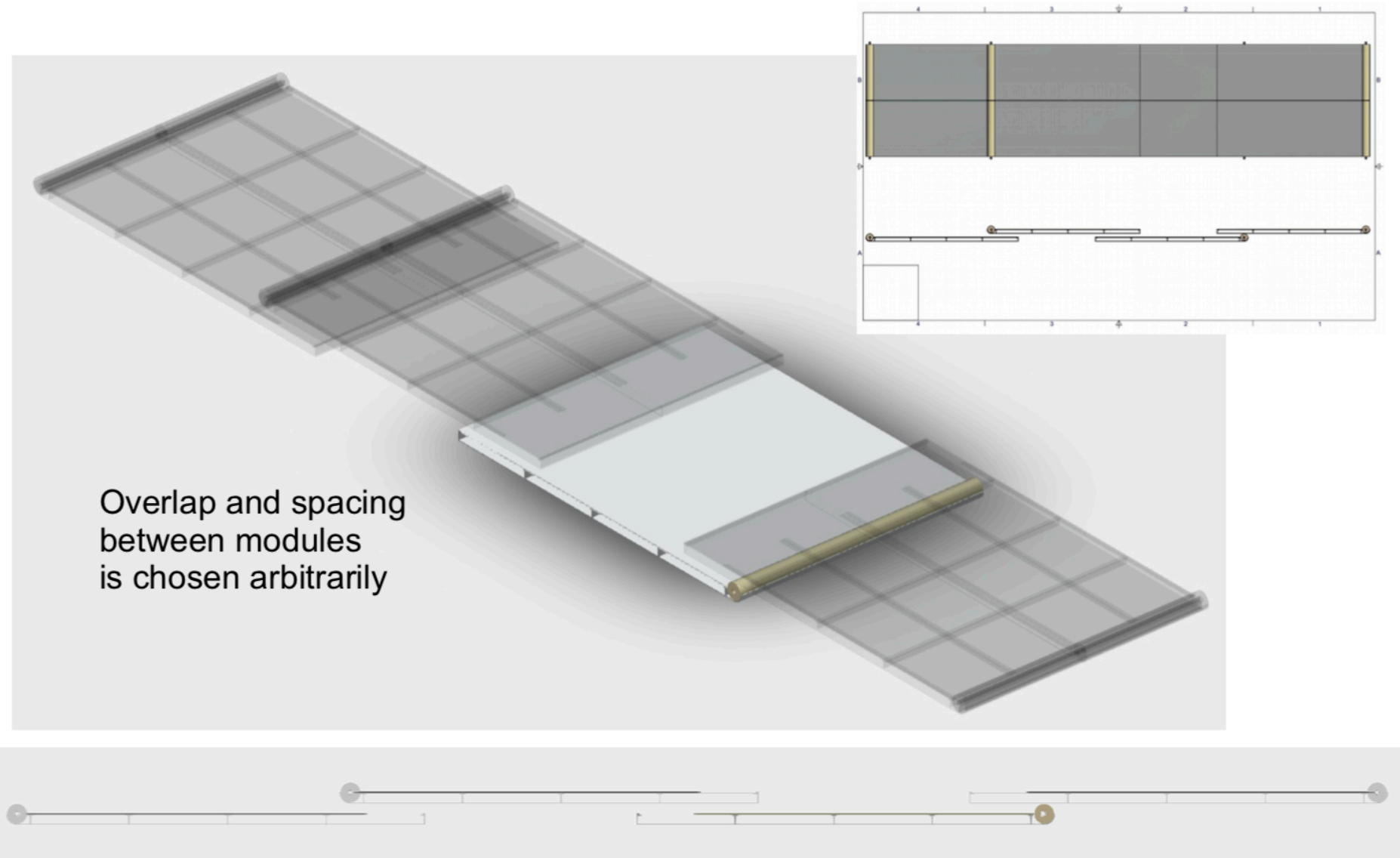
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- Each **tracking layer** is formed by **4 sub-planes** consisting of **8 adjacent modules**
- Each module contains **32 scintillator bars**

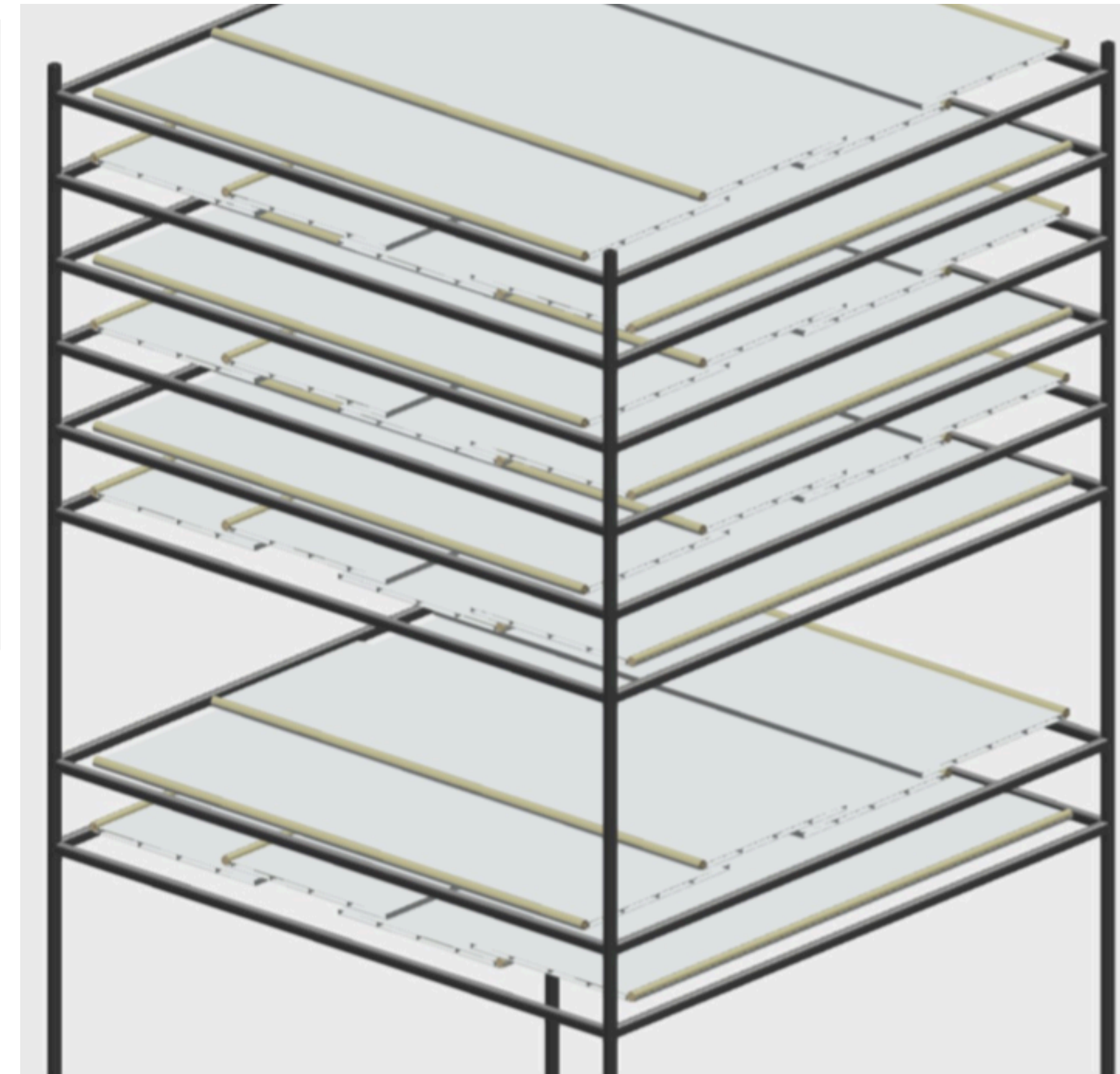


Detector Plane layout

- Each scintillator layer made of 4 sub-planes (2.3 m x 2.25 m) to cover (9 m x 9 m) with overlaps



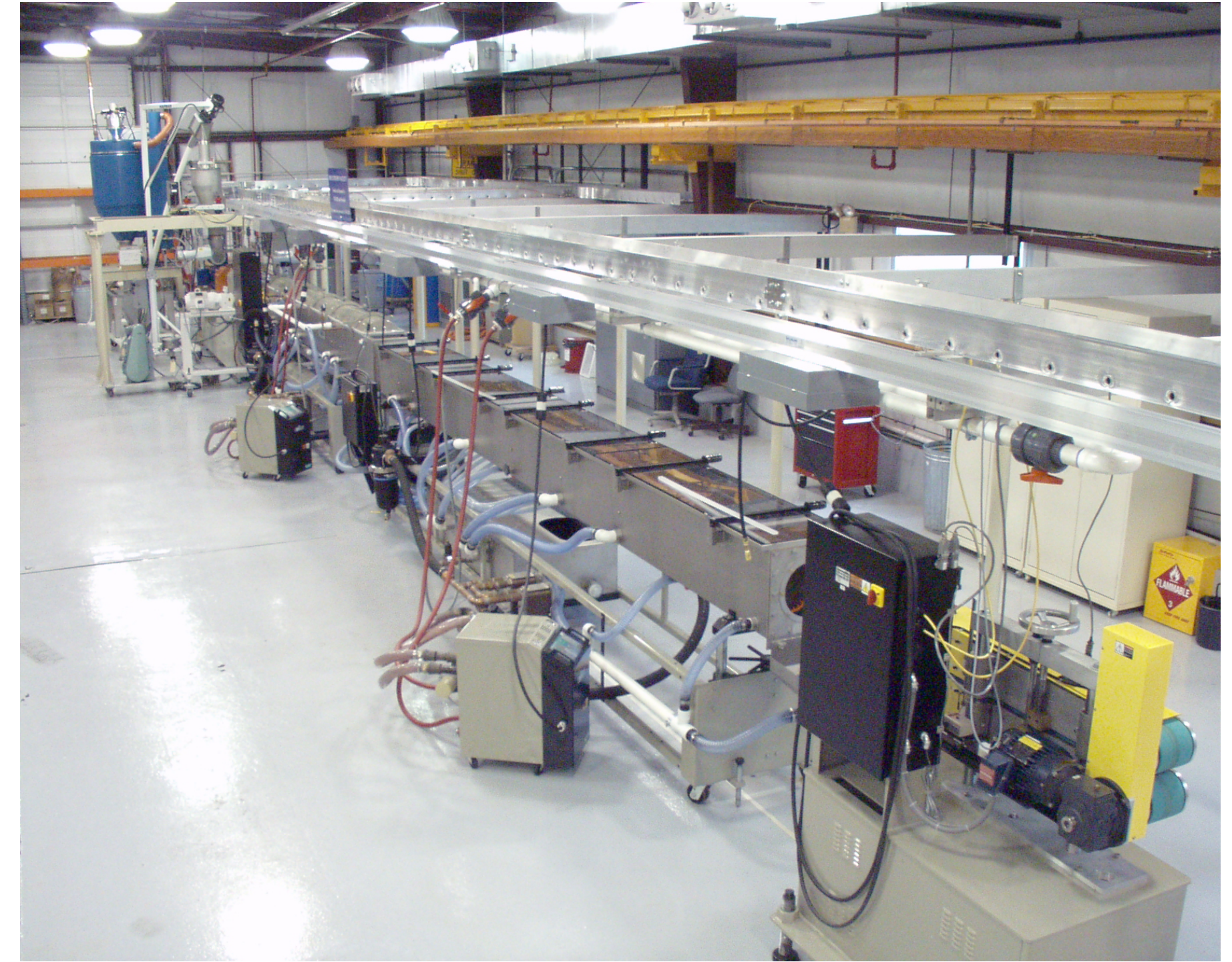
- Extrusions rotated by 90 degrees for alternating scintillating layers that gives X-Y segmentation



Scintillators / SiPMs R&D

- Bar modules are extruded scintillators
- Scintillator extrusions would be fabricated at Fermilab
- Extruded scintillators from Fermilab widely used:
 - Mu2e cosmic ray veto
 - MINERVA
 - Belle-2
 - ...

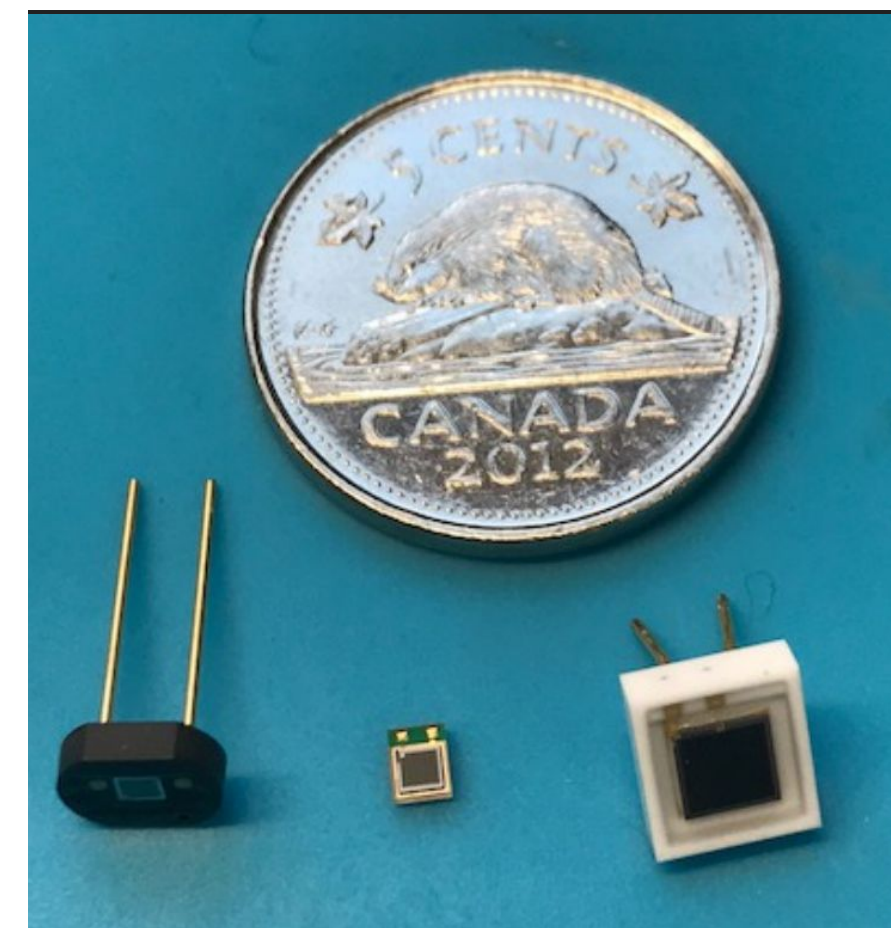
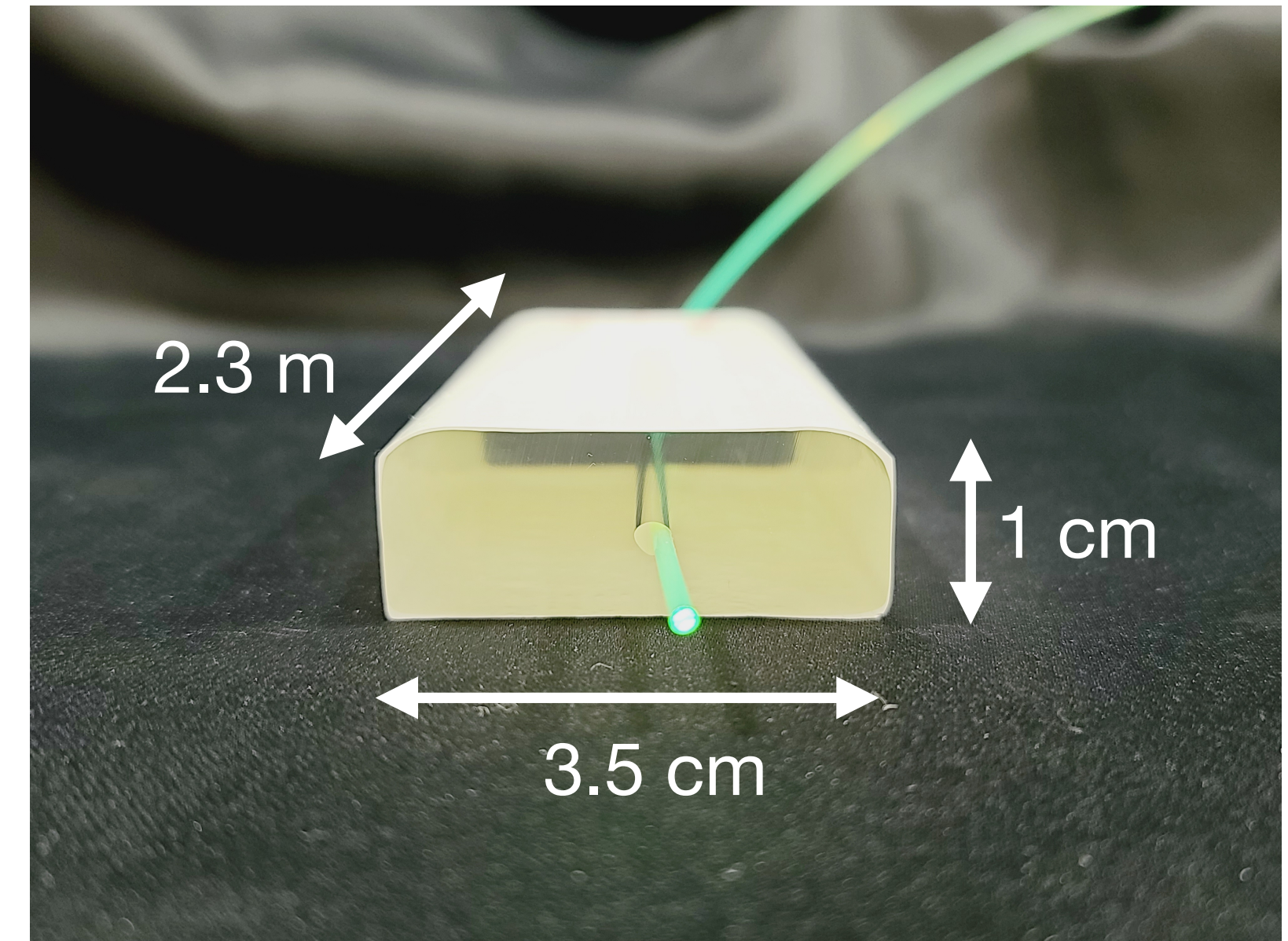
- Critical features for detector design
 - Separate downward from upward going tracks
 - Reject low beta particles from neutrinos
 - **4D tracking and vertexing** to reduce fakes/combinatorics



Scintillators / SiPMs R&D

- Extruded scintillators have a ~ 25 cm attenuation length
 - Light is carried through a wavelength-shifting (WLS) fibre running through the bar
 - Detected by silicon PMs (SiPMs) on both ends of the fibre
- Good resolution both in time and space:
 - Timing **hit resolution is ~ 1 ns** (corresponds to ~ 15 cm along the bar)
 - Transverse **hit resolution is ~ 1 cm**
- Lab tests of multiple WLS fibres show these resolutions are achievable

- Currently testing a number of SiPMs performance

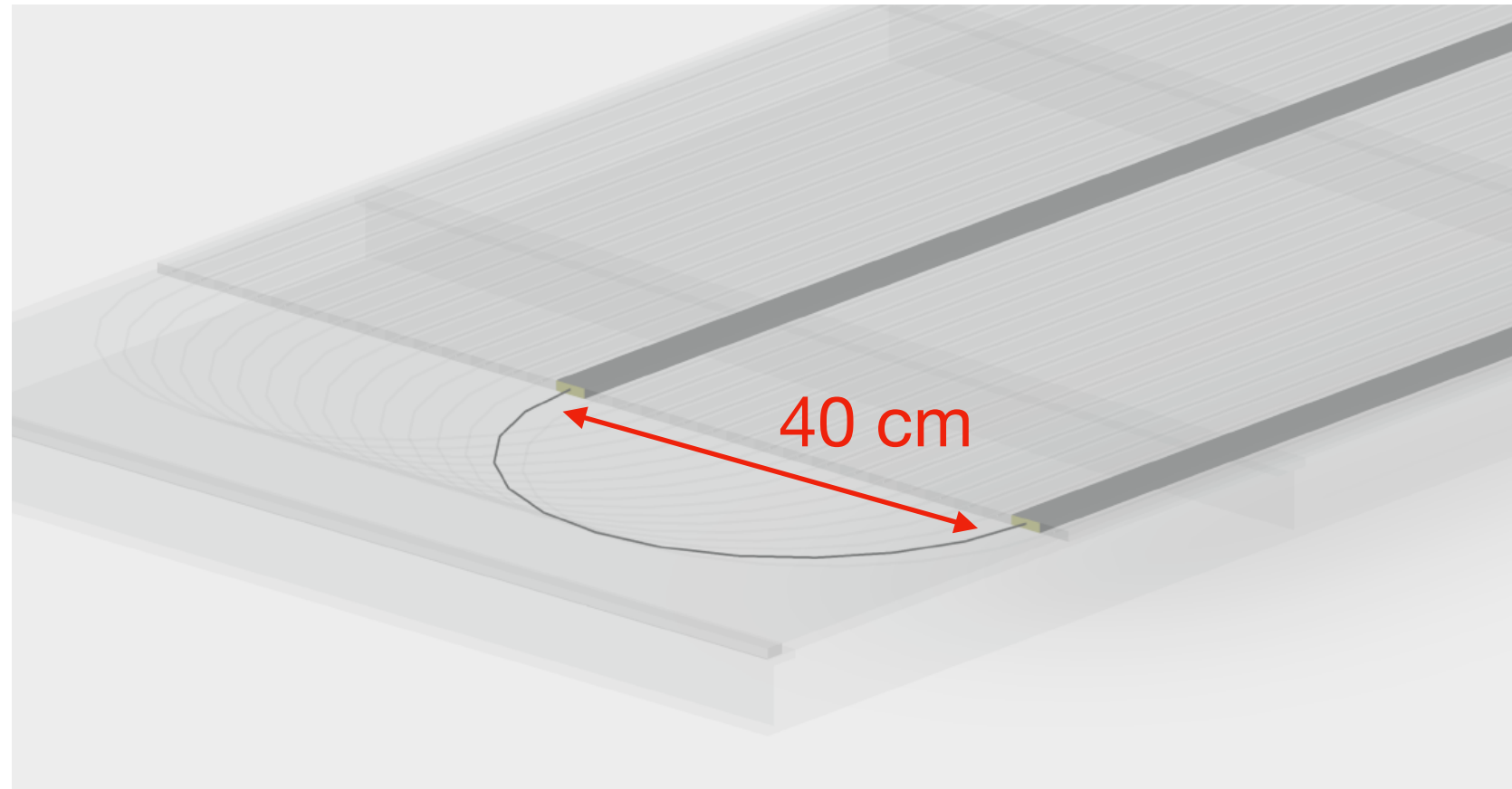


Example SiPMs that have been tested

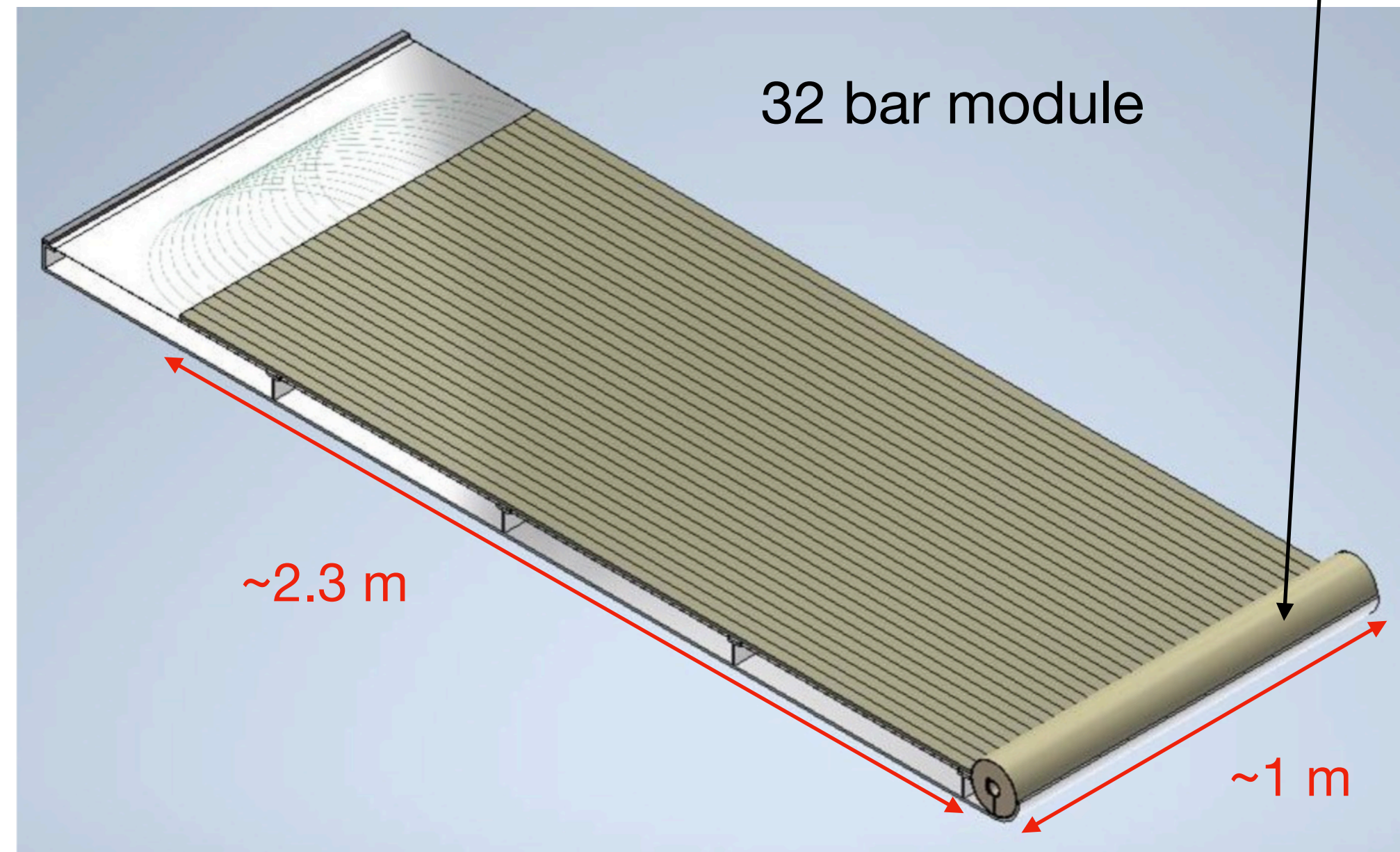
Detector Plane layout

- Extruded scintillator bars with WLSF connected to SiPMs
 - low operating voltage (~ 30 V), low sensitive to temperature and pressure variations
- Extruded bars 2.3m x 3.5 cm x 1 cm
- All SiPM connections on one side of the layer
 - Unit with 2 x (32 bars in ~ 2.3 m x ~ 1 m)
- Cylindrical region: SiPMs, connections to electronics, and cooling for temperature stabilization

- Layout of a detector unit with U-readout



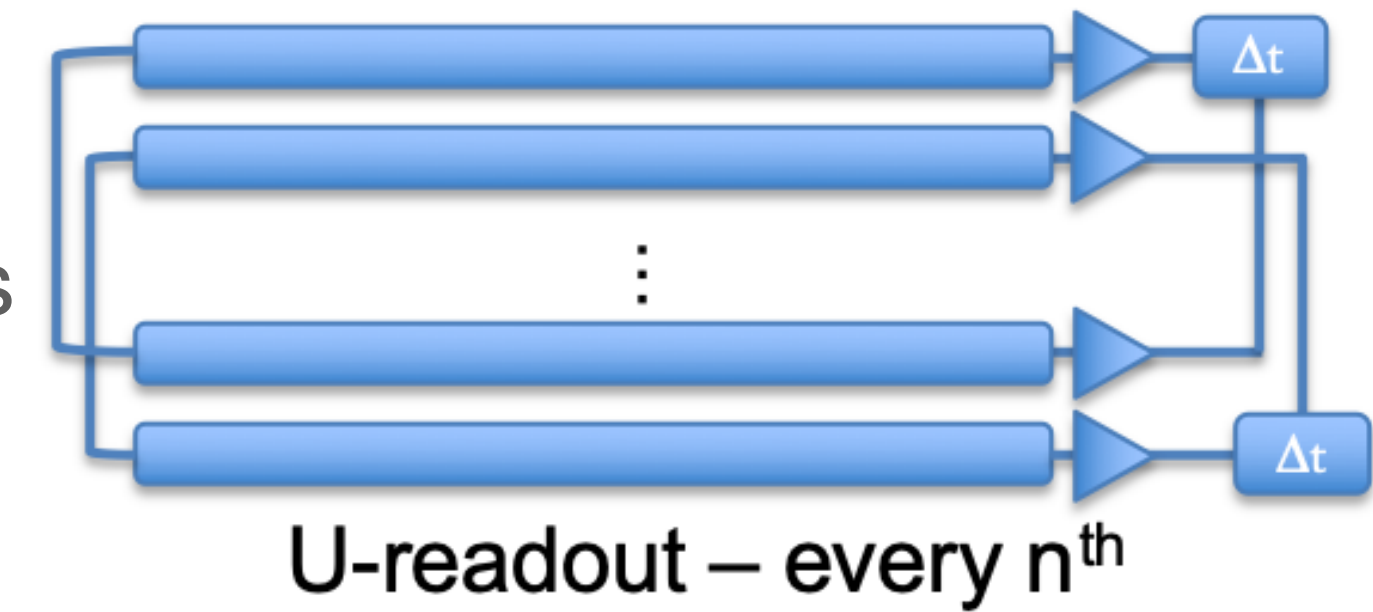
- Separation of extrusions to satisfy minimum bend radius of the fiber



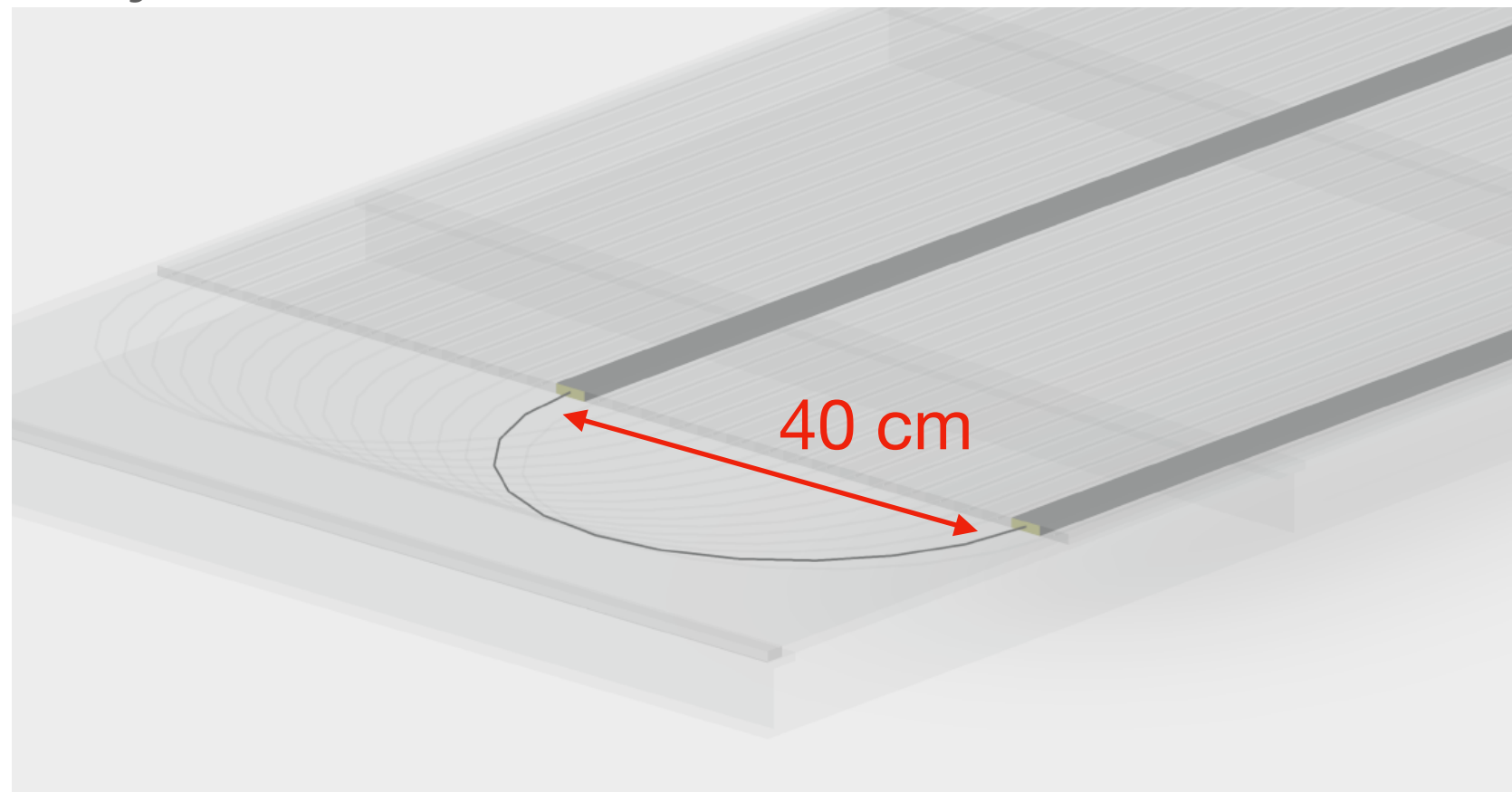
- All SiPMs on the same side **simplifies DAQ** readout
- But **require protective** cover on WLS fibres / more delicate assembly

Detector Plane layout

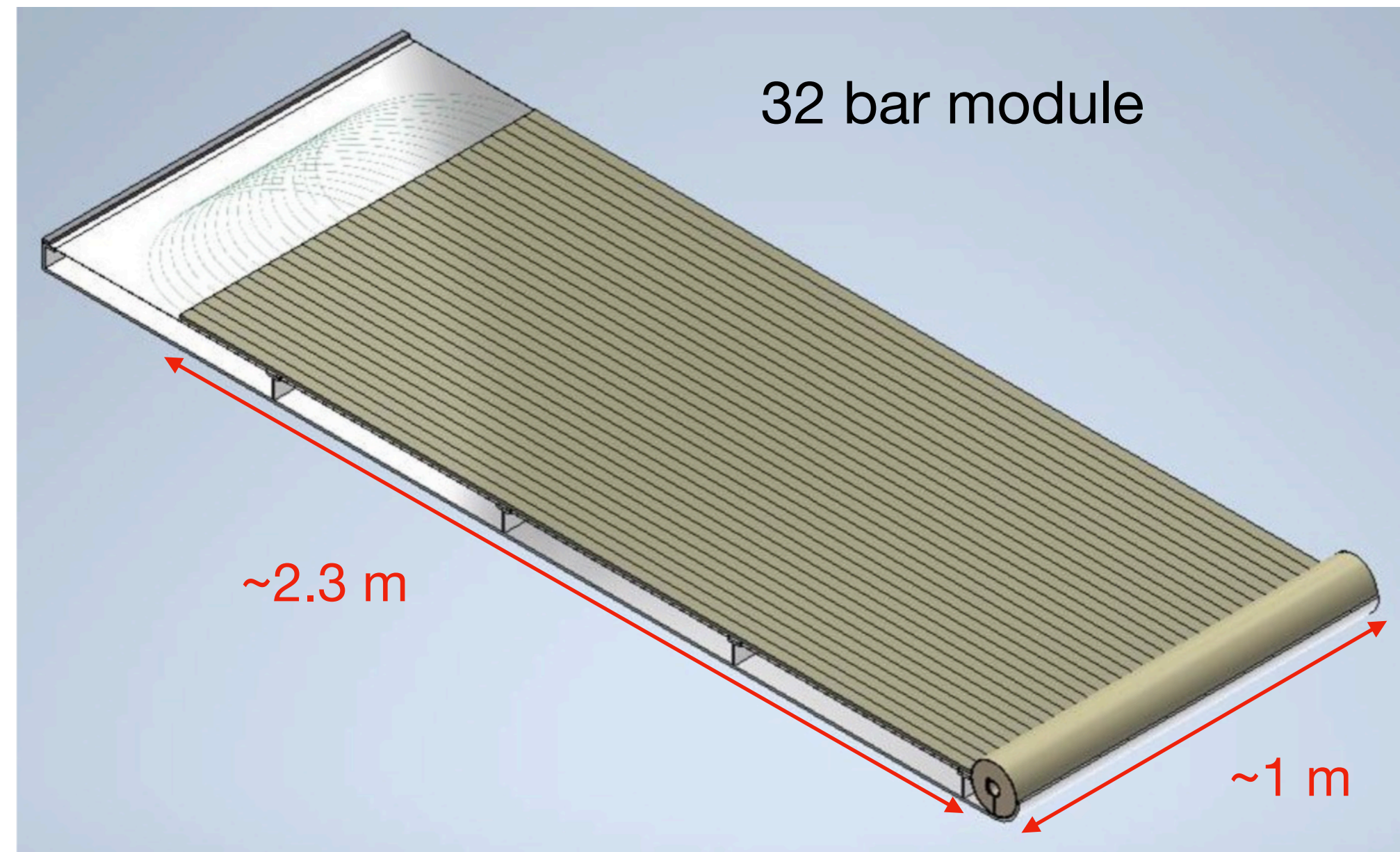
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- Time difference between light pulses at the ends of the WLS fibre: complementary coordinate measurement in each scintillating layer

Trigger and data acquisition

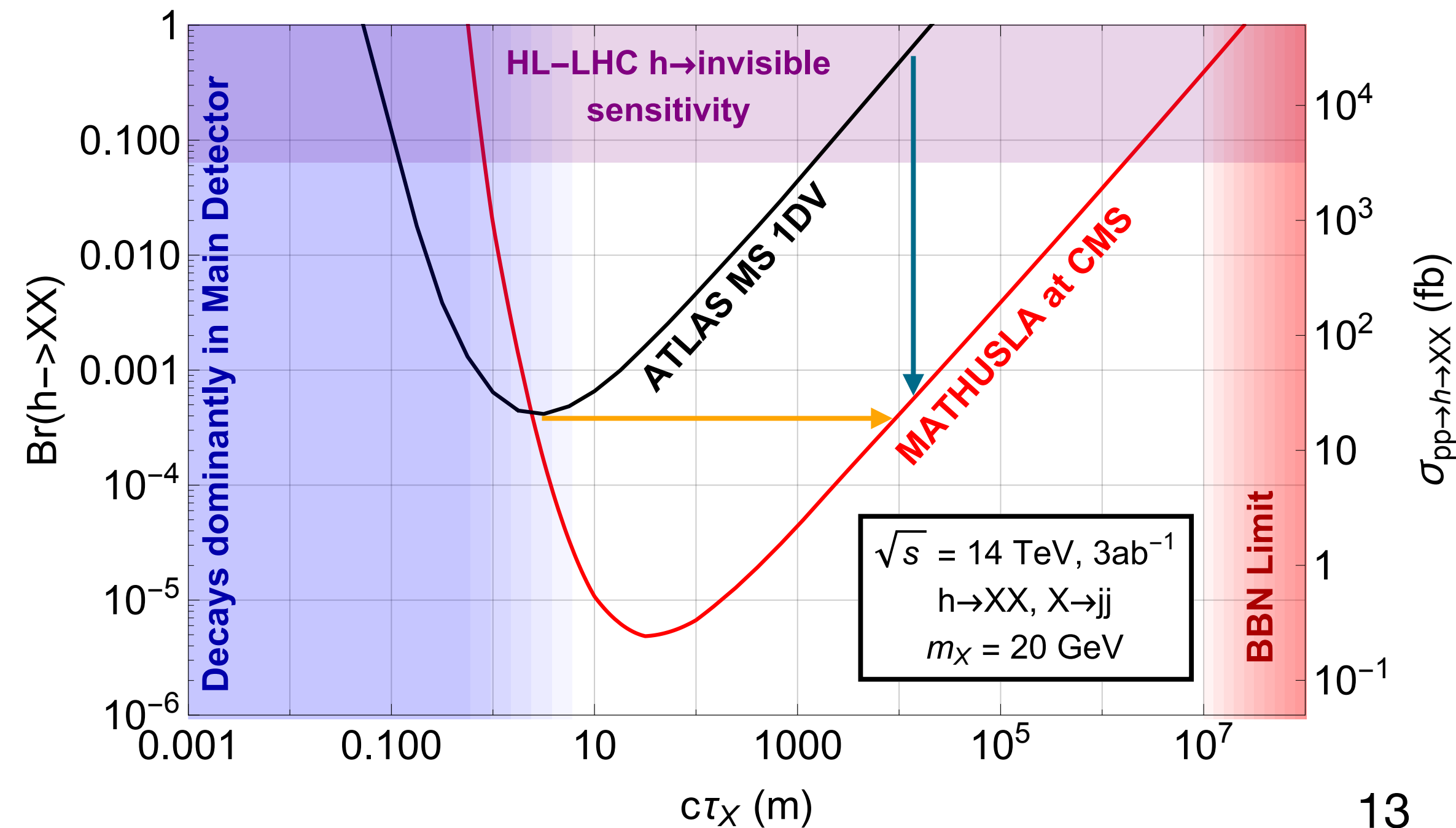
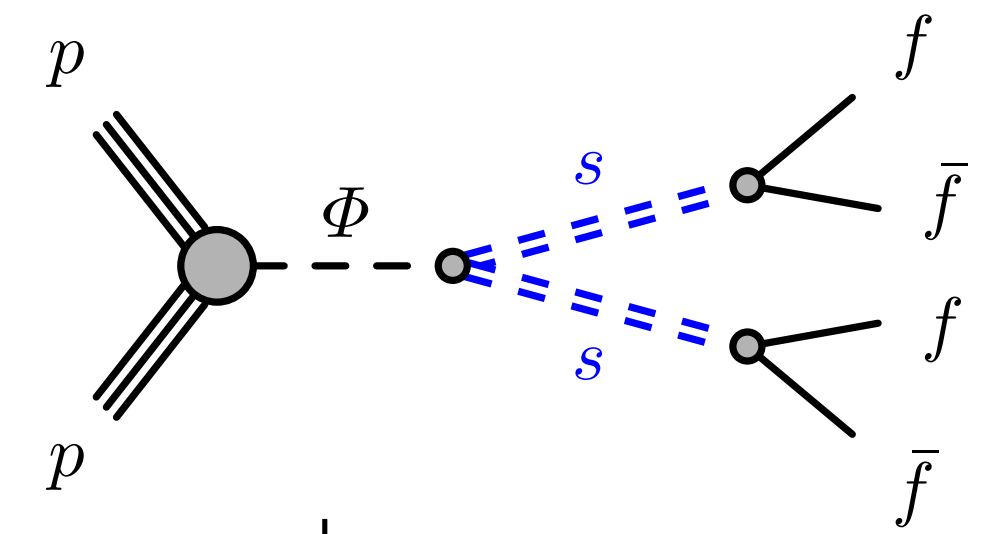
- Naturally shielded from IP, so hit rates are dominated by cosmic rays
- Plan to use commodity hardware for trigger and data acquisition
 - Stream all hits to a buffer storage
 - Relevant hits selected for permanent storage
 - Kind of high-level trigger with hours of trigger latency
- Able to trigger CMS readout!
 - Allows matching to full event information
 - Can check prompt activity and missing transverse momentum in CMS collision
 - Challenging due to tight CMS L1 latency with their Phase-2 trigger upgrade
 - **Feasibility has been confirmed in detailed study**
 - Trigger is defined by identification of upward going decay vertex using 4D tracking
 - Necessary to include particle transit times from IP, signal propagation time in MATHUSLA, trigger algorithm time, and signal transmission time back to CMS
 - Resource usage reasonable with current generation FPGAs

Expected sensitivity

- Great background rejection to \sim zero thanks to:
 - Rock shielding: punch-through jets in ATLAS (CMS) background is the limiting factor for the mu-vertex search
 - In MATHUSLA this is eliminated in the rock
 - \sim zero QCD background
 - Good space-time resolution: rejection of 1.7MHz from cosmic μ
 - Robust tracking: rejection of 10Hz from LHC μ
- **Background hypotheses confirmed** with the Test Stand
- Primary physics case: hadronically decaying $O(10-100 \text{ GeV})$ LLPs
 - Example: exotics Higgs decays in Hidden sector $H \rightarrow ss$

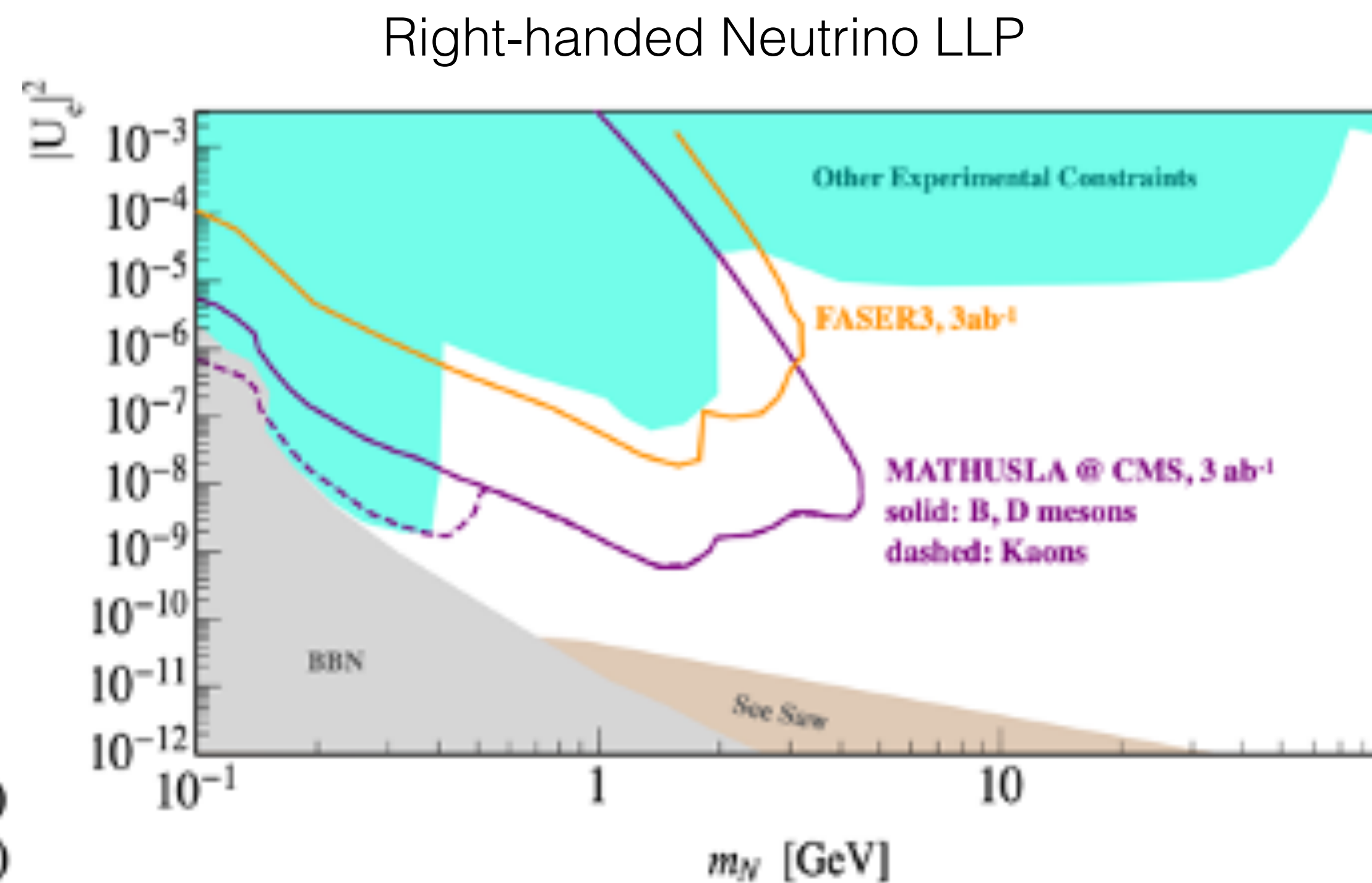
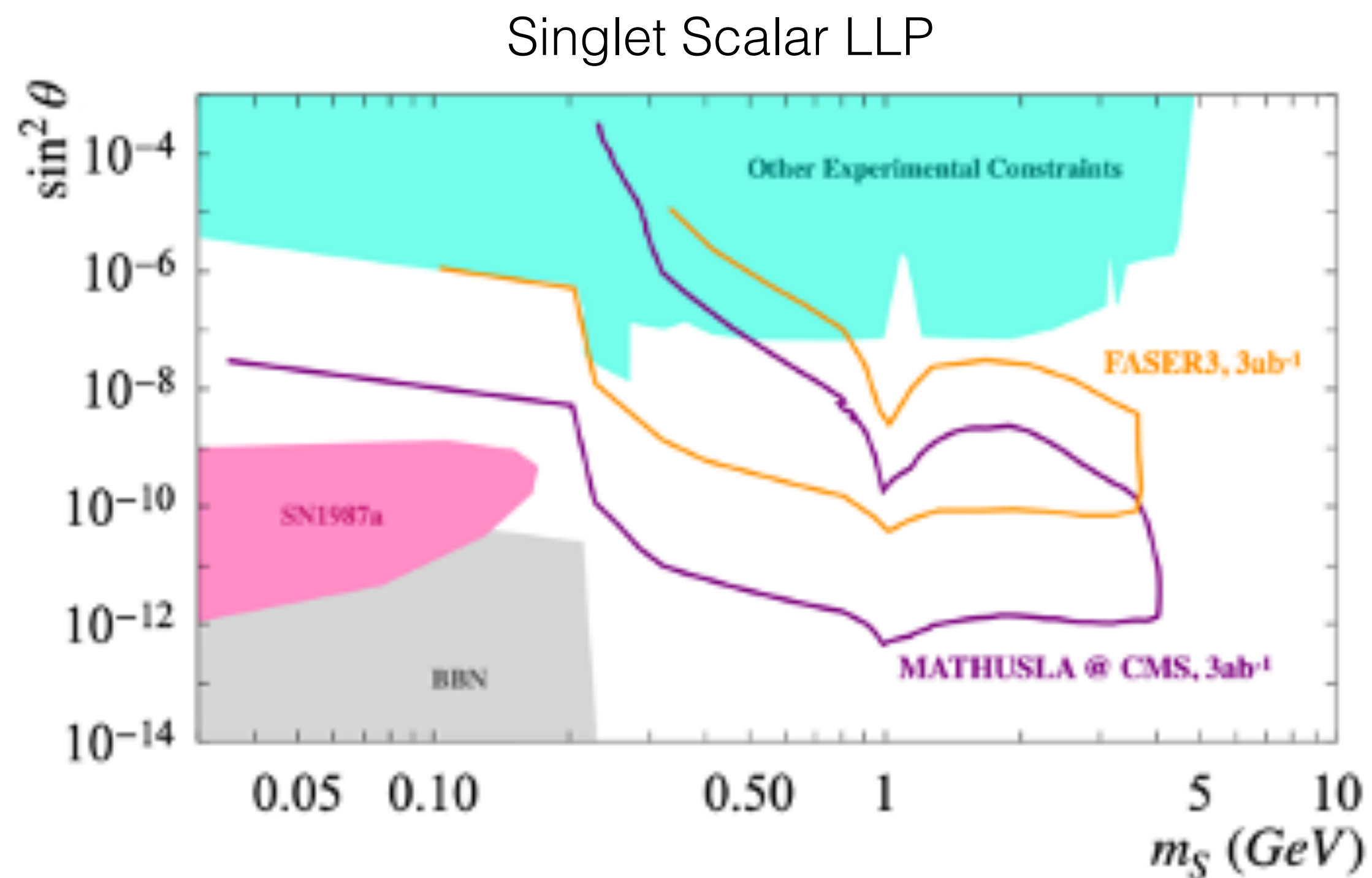
Absence of collision backgrounds allow MATHUSLA for a good gain wrt best ATLAS expected result:

- 3 orders of magnitude in $c\tau$
- 3 orders of magnitude in sensitivity



Expected sensitivity

- Secondary physics case: GeV-scale LLPs from B, D meson decays



- Greatly extends the discovery potential to smaller angles
- Significant complementarity with FASER, which probes similar masses but shorter lifetimes

Status and outlook

- Current status:
 - Detector technology has been studied extensively
 - Small lab-scale prototype units are under construction
 - Conceptual design report (CDR) in preparation
- Looking ahead:
 - Full detector construction can be installed in stages (per detector unit)
 - Installation of one 9m x 9m unit $O(1 \text{ week})$
 - Can begin taking data before all units are installed
 - Goal is to be ready for the start of HL-LHC running

Documentation

Auxiliary LHC detectors are widely supported in Snowmass study and European Strategy Update.

- Web: <https://mathusla-experiment.web.cern.ch>
- Original idea: J.P. Chou, D. Curtin, H. Lubatti [arXiv 1606.06298](https://arxiv.org/abs/1606.06298)
- Mathusla physics case - theory white paper [arXiv 1806.07396](https://arxiv.org/abs/1806.07396)
- LHCC Letter of Intent [arXiv 1811.00927](https://arxiv.org/abs/1811.00927)
- Input to **European Strategy** for Particle Physics [arXiv 1901.04040](https://arxiv.org/abs/1901.04040)
- Updated LHCC Letter of Intent: [arXiv 2009.01693](https://arxiv.org/abs/2009.01693)
- MATHUSLA Test Stand: [NIMA 985 \(2021\) 164661](https://arxiv.org/abs/2101.16466)
- Snowmass White Paper: <https://arxiv.org/abs/2203.08126>

MATHUSLA Collaboration

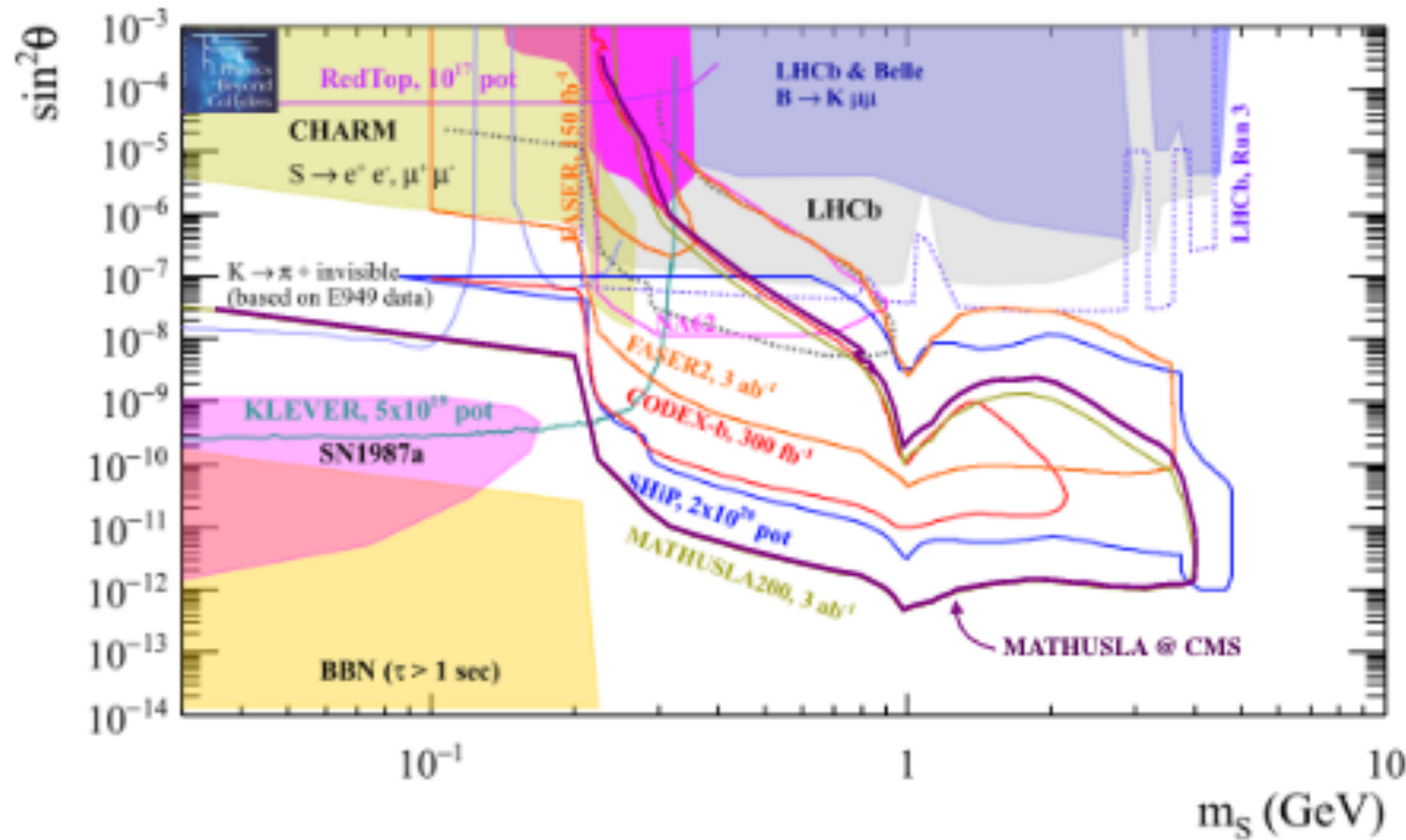


Backup

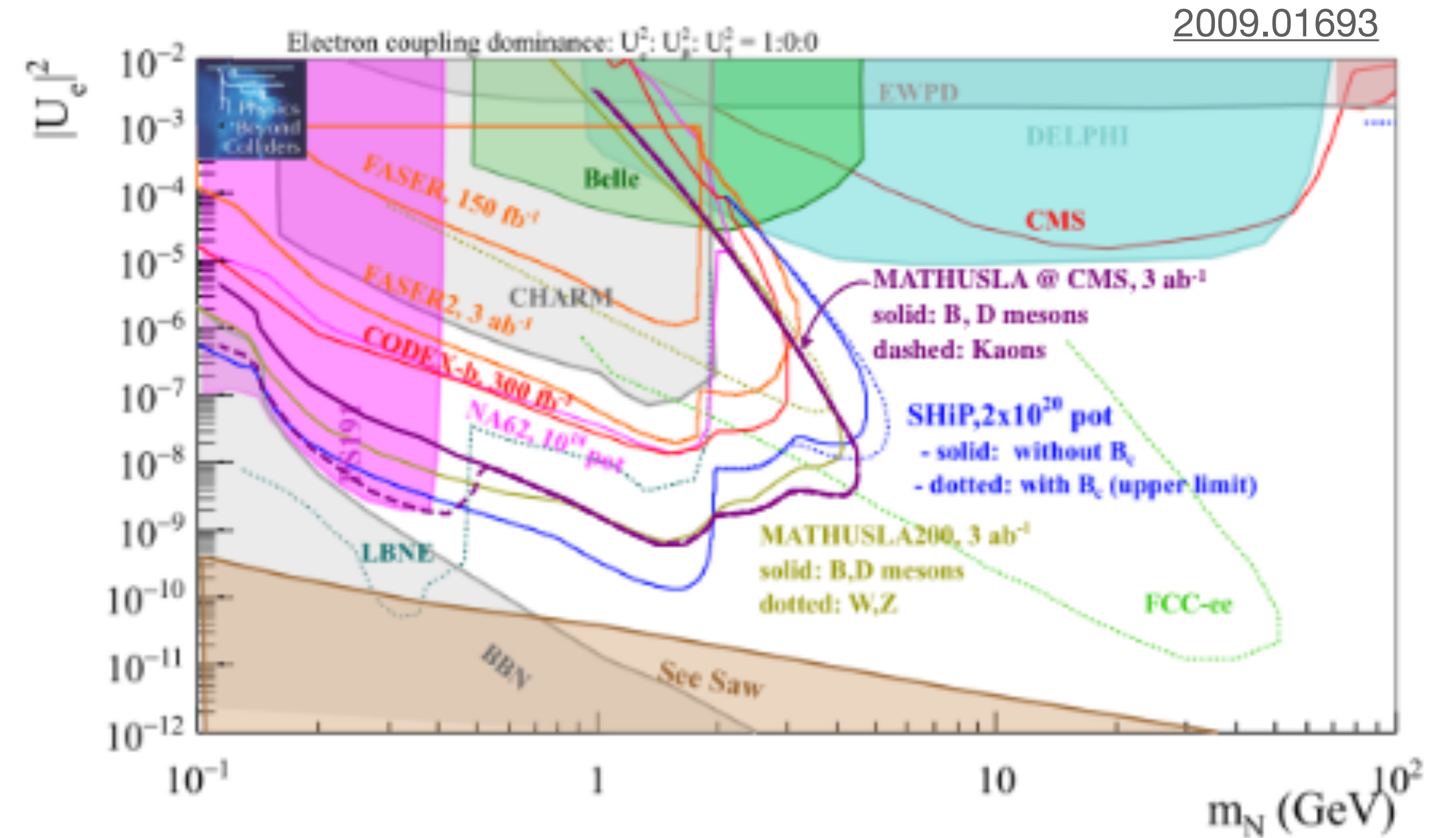
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Singlet Scalar LLP



Right-handed Neutrino LLP



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