

---

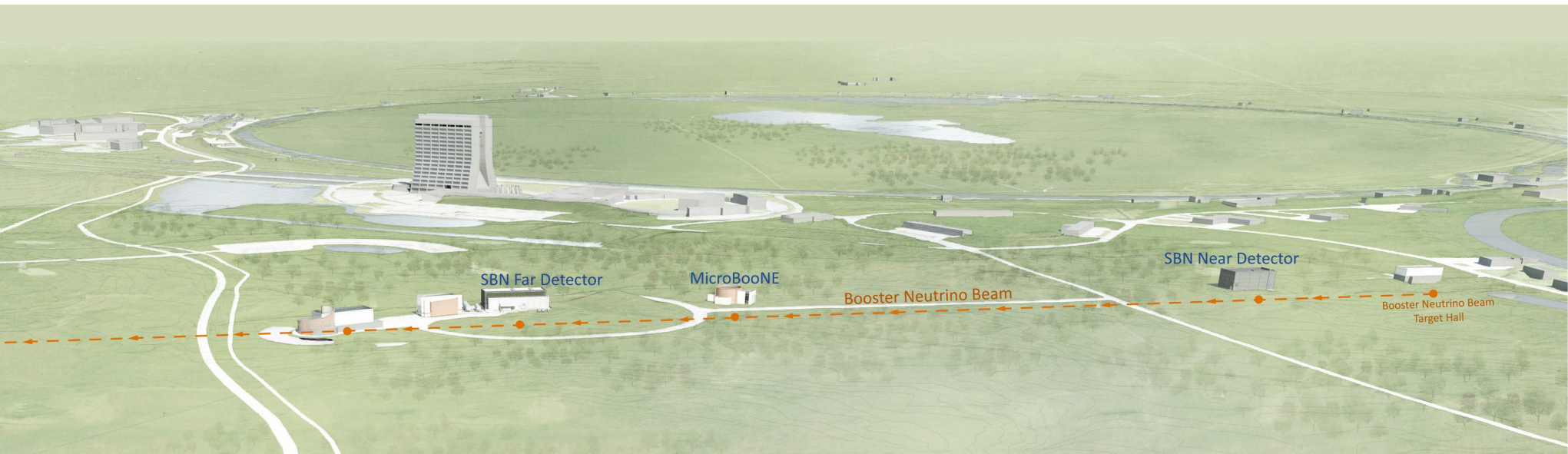
# Neutrino Interaction Capabilities of SBND

Andrew Furmanski  
NuFact 2023, Seoul, South Korea



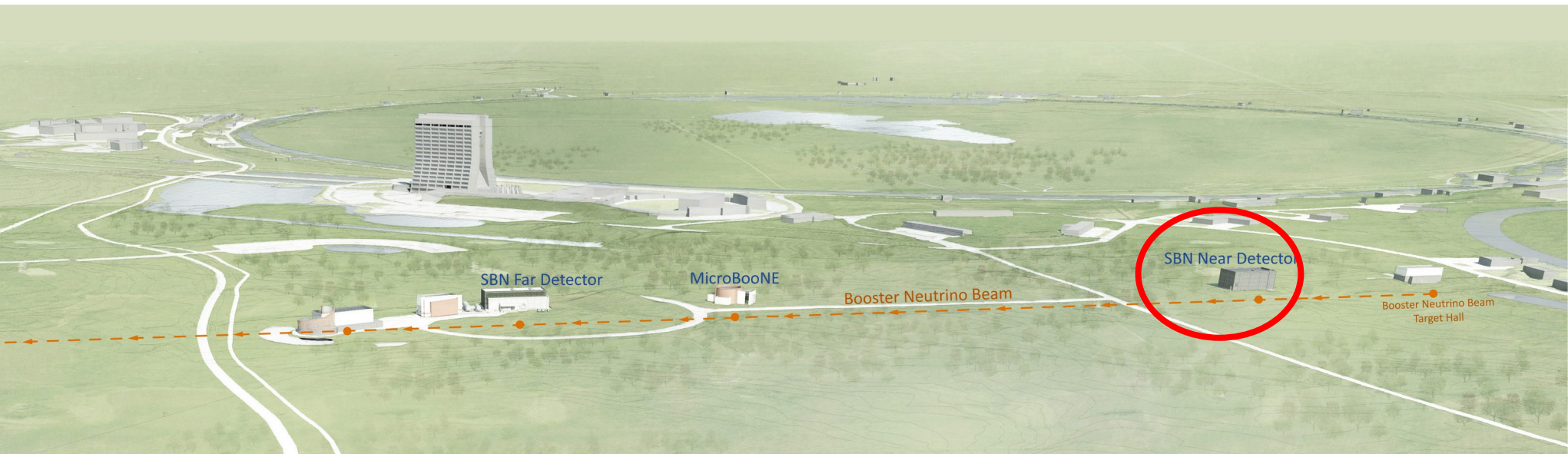
# What is SBND?

- “Short Baseline Near Detector”
- Near Detector for SBN program
  - With world-leading eV-scale sterile neutrino sensitivity
- Single-detector physics program of its own



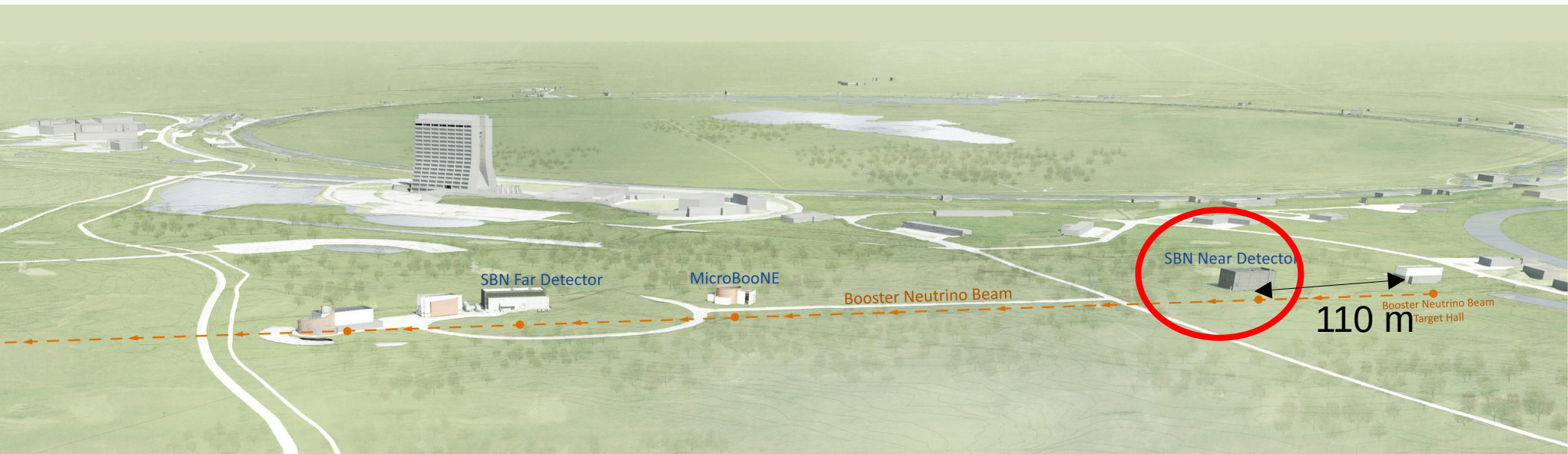
# What is SBND?

- “Short Baseline Near Detector”
- Near Detector for SBN program
  - With world-leading eV-scale sterile neutrino sensitivity
- Single-detector physics program of its own



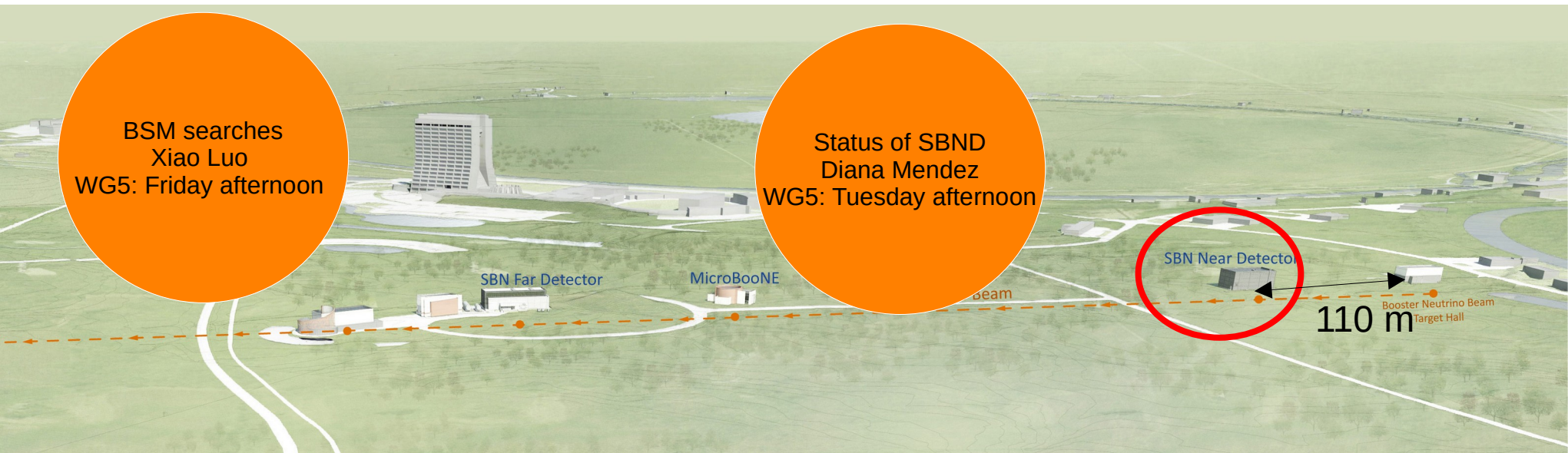
# What is SBND?

- “Short Baseline Near Detector”
- Near Detector for SBN program
  - With world-leading eV-scale sterile neutrino sensitivity
- Single-detector physics program of its own



# What is SBND?

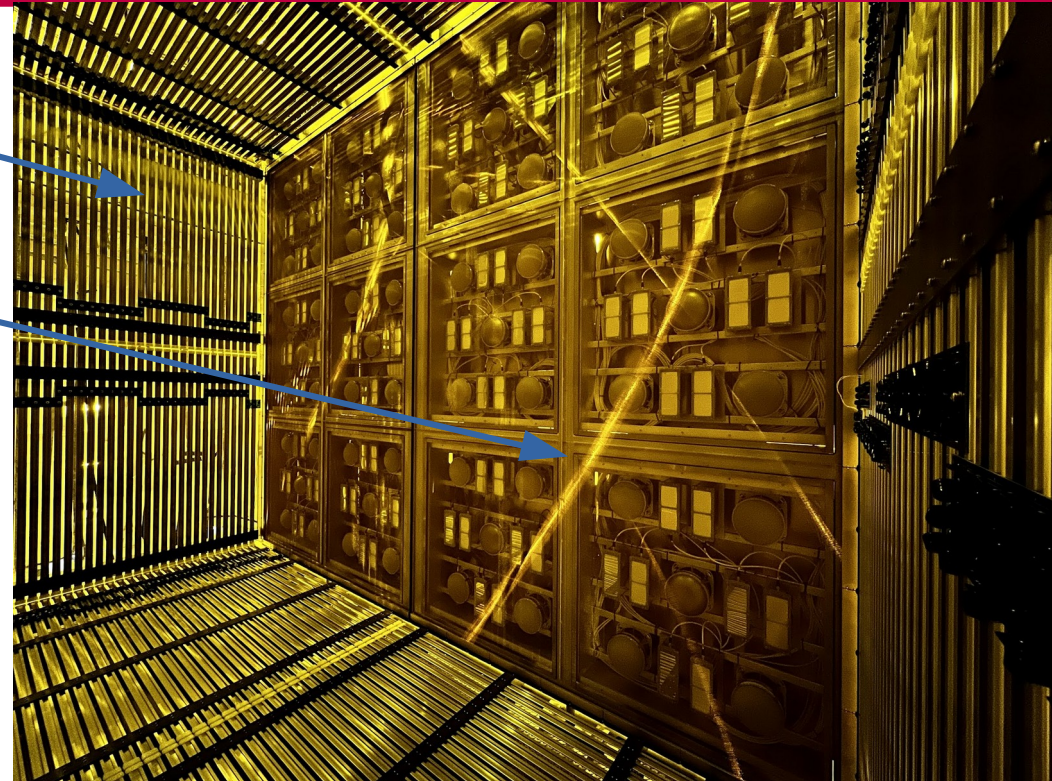
- “Short Baseline Near Detector”
- Near Detector for SBN program
  - With world-leading eV-scale sterile neutrino sensitivity
- Single-detector physics program of its own



# Detector Design

Two TPCs (one shown)

~10,000 readout wires (few-mm resolution)



Few-mm spatial resolution leads to low particle tracking thresholds and excellent momentum/direction resolution



# Detector Design

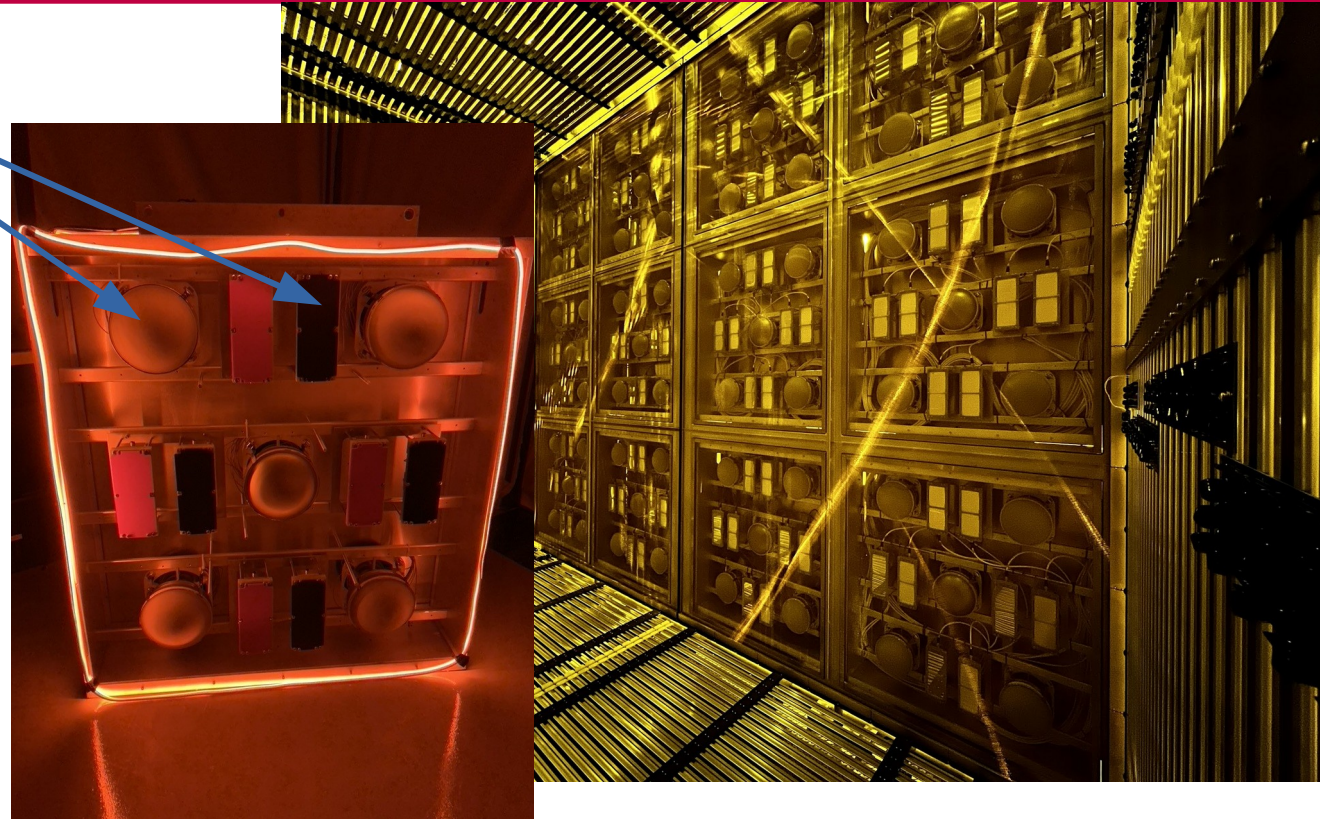
Large photocoverage

PMTs and X-ARAPUCAs



# Detector Design

Some UV-sensitive

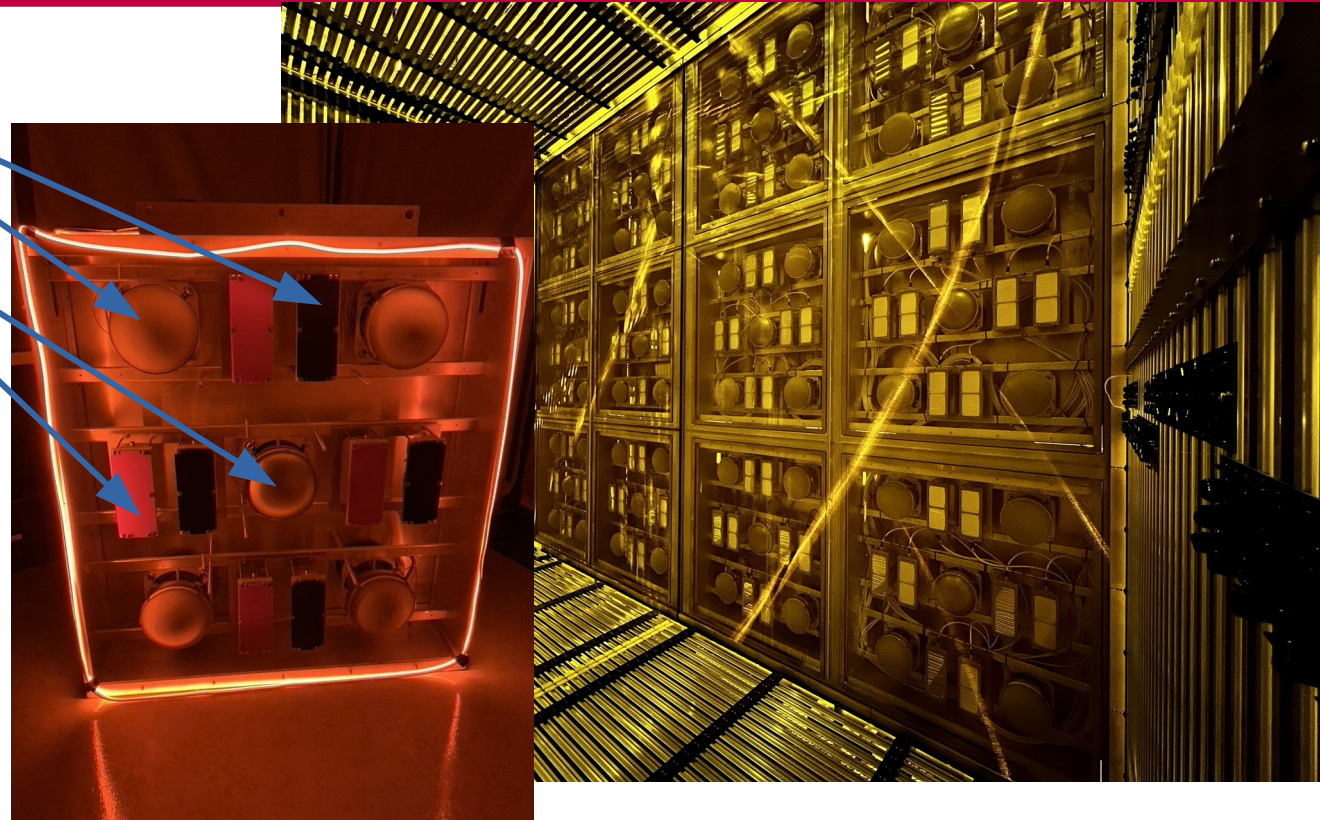




# Detector Design

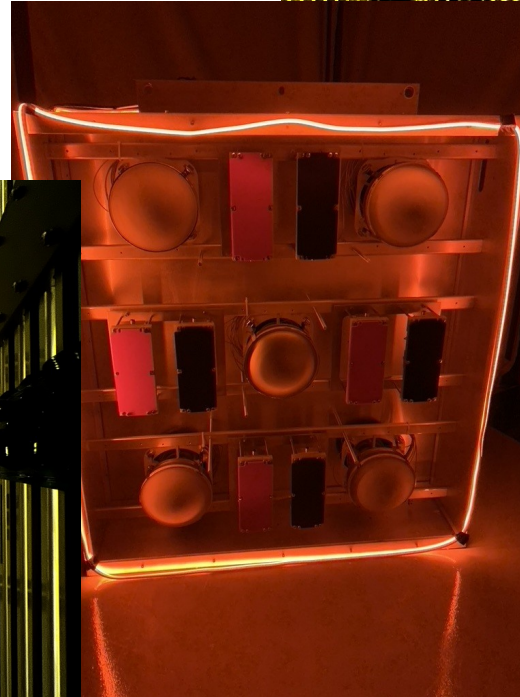
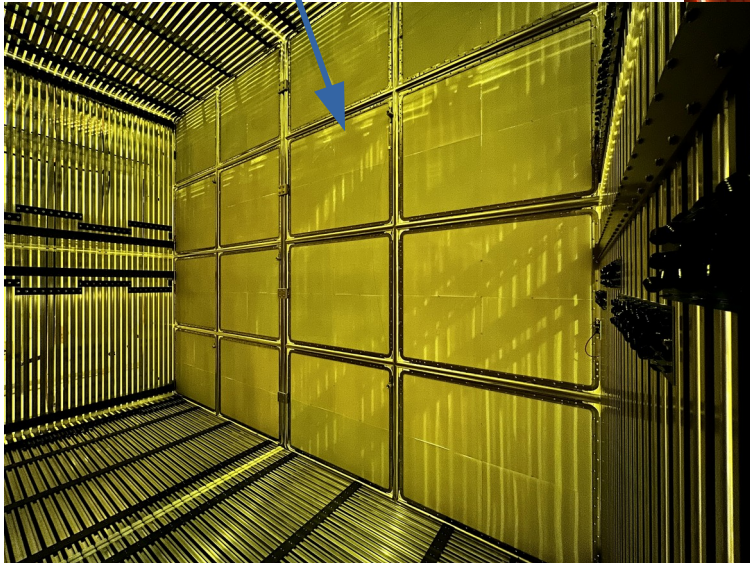
Some UV-sensitive

Some visible-sensitive



# Detector Design

Wavelength-shifting reflective foils

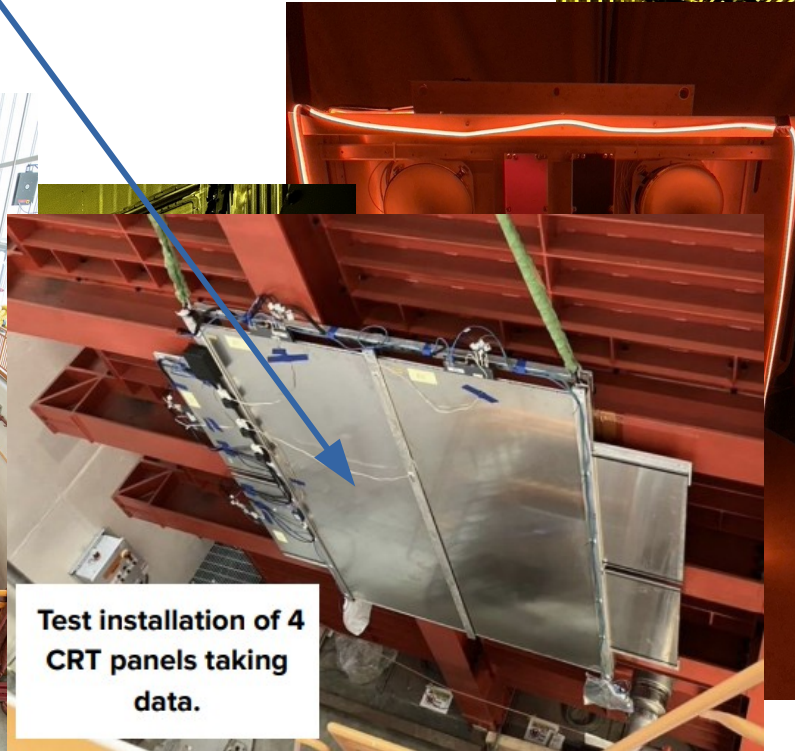
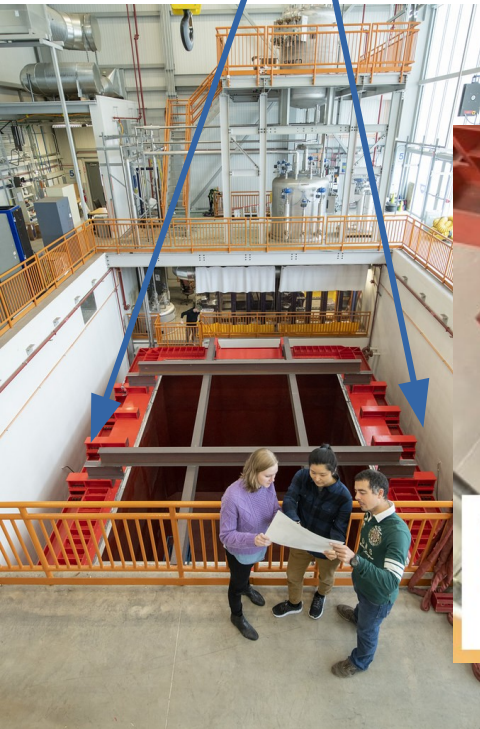


High-coverage photon detection allows for improved calorimetry, lower thresholds, and better background rejection!



# Detector Design

Cosmic Ray Tagger – high coverage

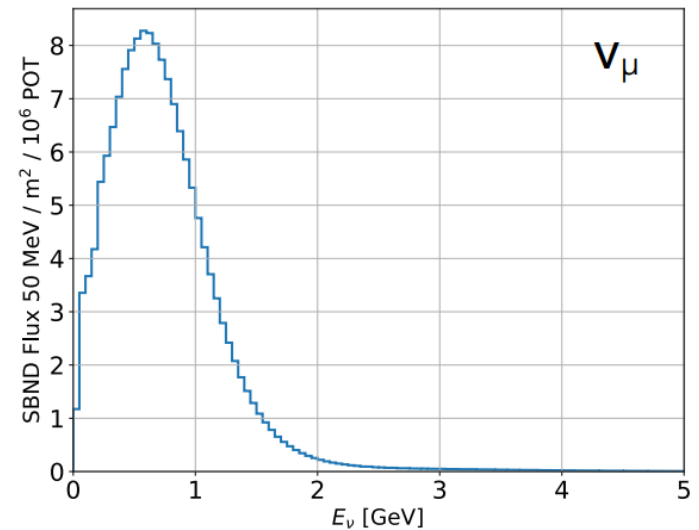


Almost  $4\pi$  coverage CRT suppresses high cosmic rate at surface



# Booster Neutrino Beam

- 8 GeV proton beam from Fermilab Booster
- Beryllium target
- Single focusing horn
- 50m decay pipe
- 600MeV peak, 800MeV mean, tail up to ~4-5 GeV



# SBND Event Rates

| Process            |   | No.<br>Events |
|--------------------|---|---------------|
|                    | <i><math>\nu_\mu</math> Events (By Final State Topology)</i>  |               |
| CC Inclusive       |   | 5,212,690     |
| CC 0 $\pi$         | $\nu_\mu N \rightarrow \mu + Np$                              | 3,551,830     |
|                    | · $\nu_\mu N \rightarrow \mu + 0p$                            | 793,153       |
|                    | · $\nu_\mu N \rightarrow \mu + 1p$                            | 2,027,830     |
|                    | · $\nu_\mu N \rightarrow \mu + 2p$                            | 359,496       |
|                    | · $\nu_\mu N \rightarrow \mu + \geq 3p$                       | 371,347       |
| CC 1 $\pi^\pm$     | $\nu_\mu N \rightarrow \mu + \text{nucleons} + 1\pi^\pm$      | 1,161,610     |
| CC $\geq 2\pi^\pm$ | $\nu_\mu N \rightarrow \mu + \text{nucleons} + \geq 2\pi^\pm$ | 97,929        |
| CC $\geq 1\pi^0$   | $\nu_\mu N \rightarrow \mu + \text{nucleons} + \geq 1\pi^0$   | 497,963       |
| NC Inclusive       |   | 1,988,110     |
| NC 0 $\pi$         | $\nu_\mu N \rightarrow \text{nucleons}$                       | 1,371,070     |
| NC 1 $\pi^\pm$     | $\nu_\mu N \rightarrow \text{nucleons} + 1\pi^\pm$            | 260,924       |
| NC $\geq 2\pi^\pm$ | $\nu_\mu N \rightarrow \text{nucleons} + \geq 2\pi^\pm$       | 31,940        |
| NC $\geq 1\pi^0$   | $\nu_\mu N \rightarrow \text{nucleons} + \geq 1\pi^0$         | 358,443       |

30x MicroBooNE stats

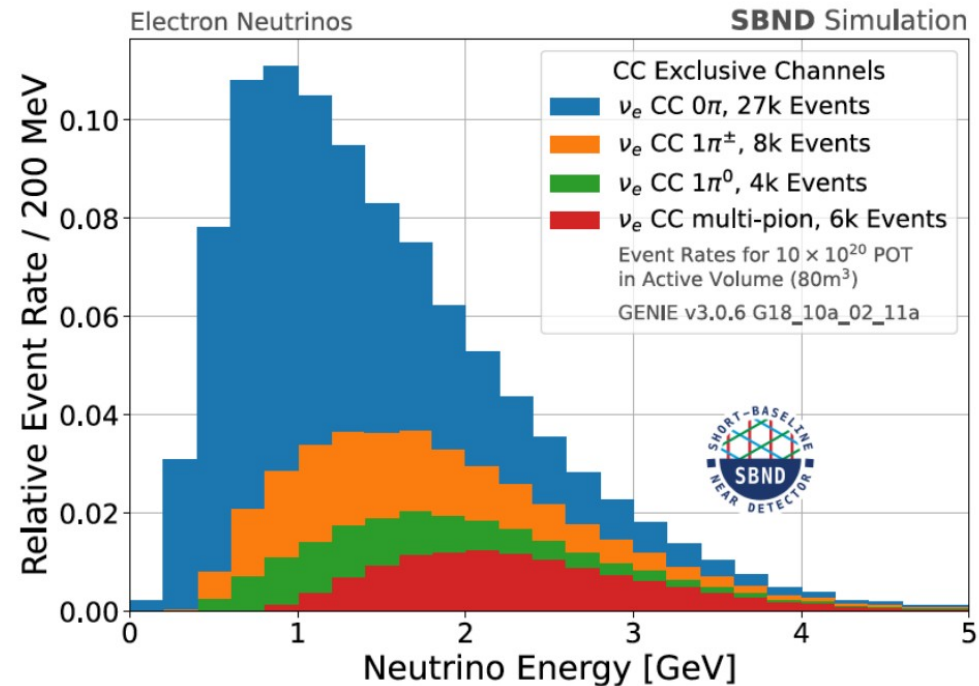
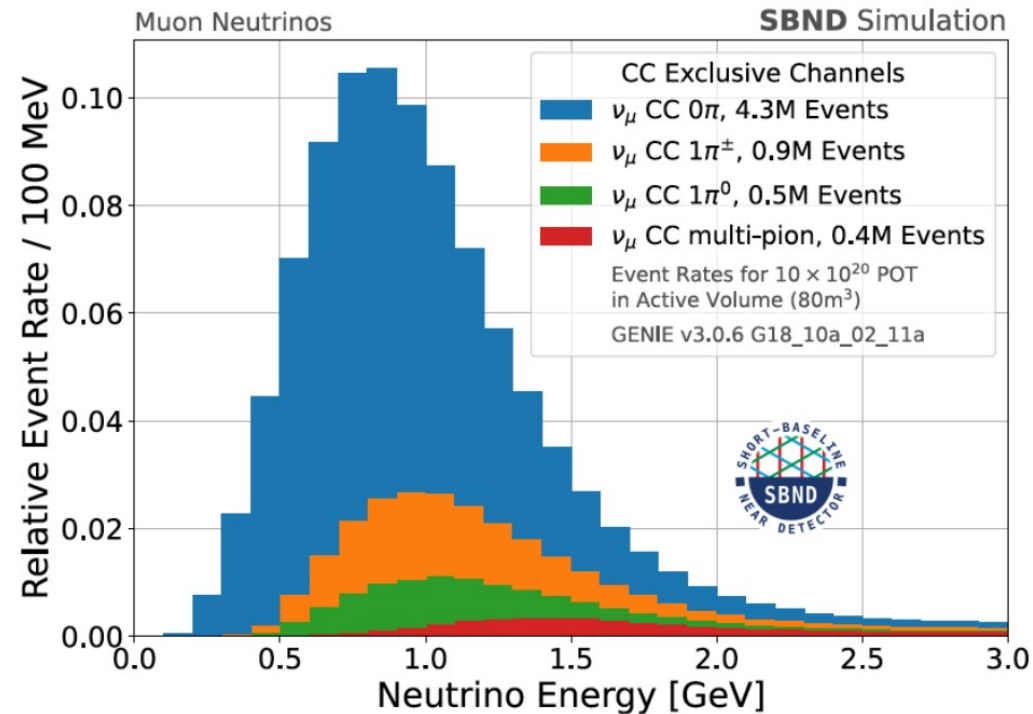
5M  $\nu_\mu$  CC per year  
 2M NC per year  
 12k  $\nu_e$  CC per year



# SBND Event Rates

Sufficient statistics for differential measurements of all topologies

Even electron neutrino multi-pion events (SIS/DIS region)

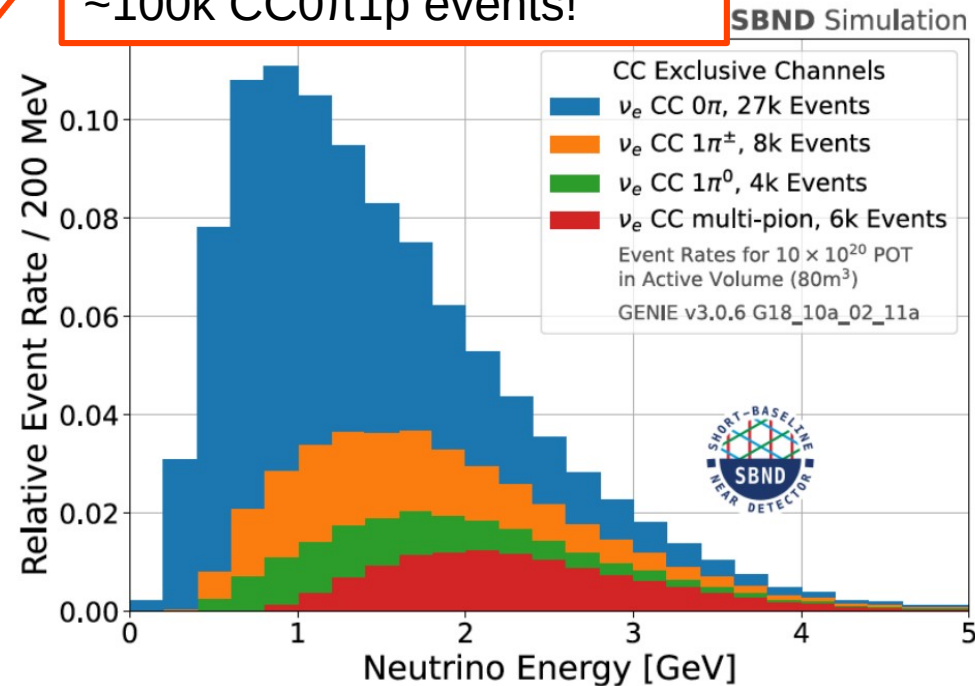
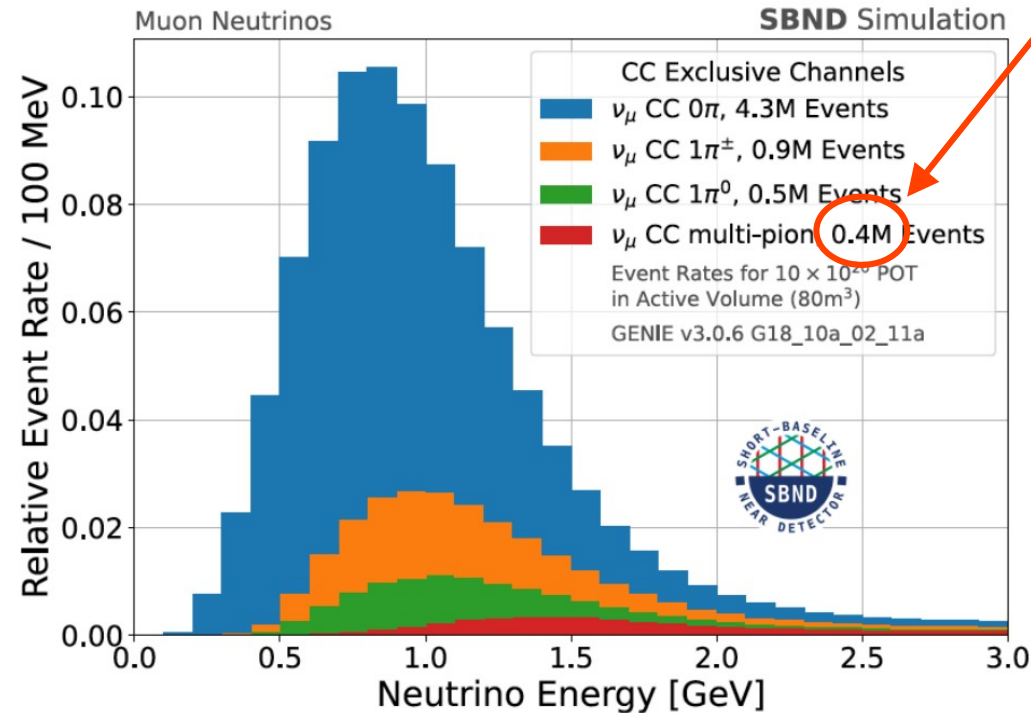


# SBND Event Rates

Sufficient statistics for differential measurements of all topologies

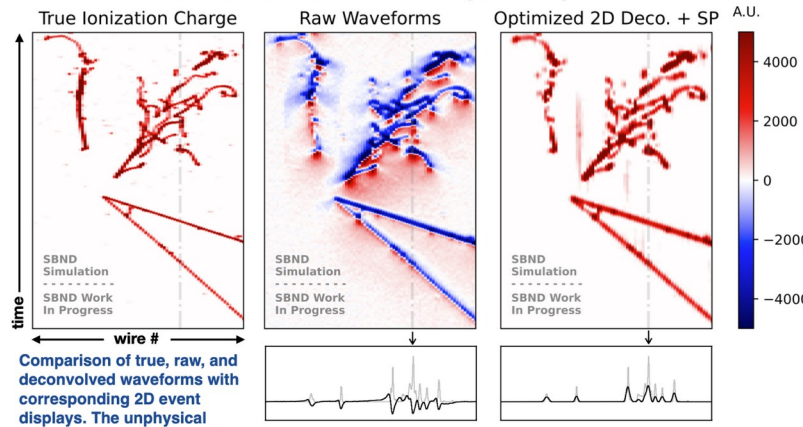
Even electron neutrino multi-pion events (SIS/DIS region)

Compare to MicroBooNE  
~100k CC0 $\pi$ 1p events!



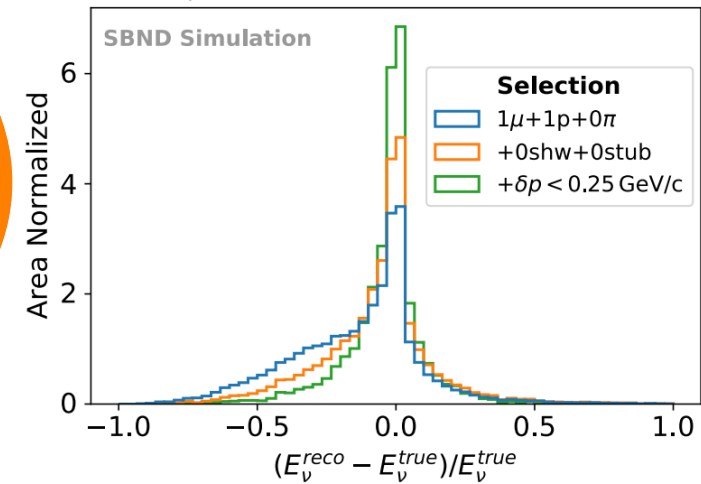
# Expected Performance

Simulated Neutrino Event (U Plane)

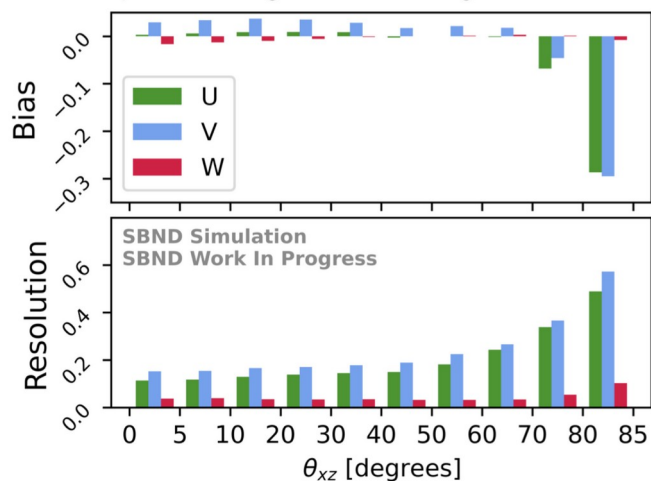


CC0 $\pi$  Event Selection  
Mun Jung  
Poster

$E_\nu$  Reconstruction Fractional Bias

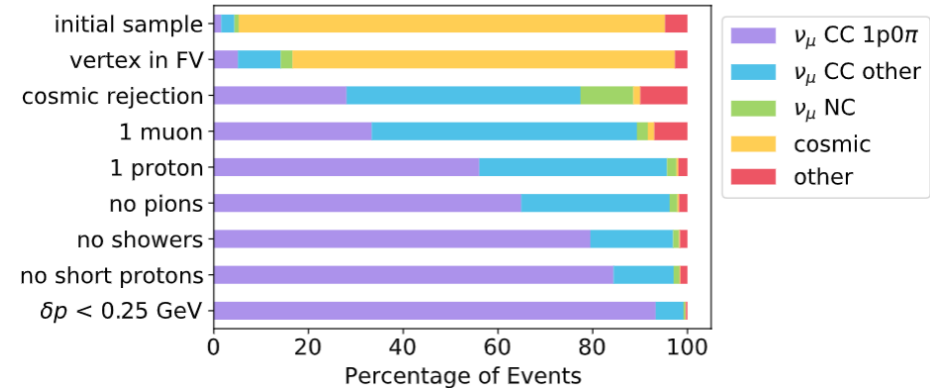


Optimized Signal Processing Performance



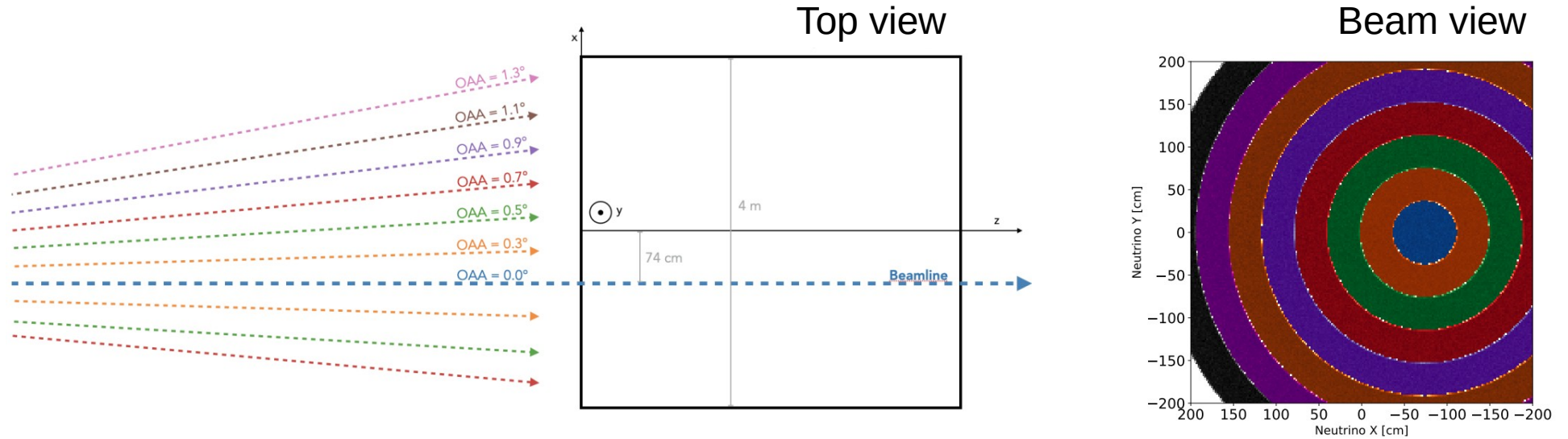
Signal Processing  
Avinay Bhat  
WG6: Friday Morning

SBND Simulation



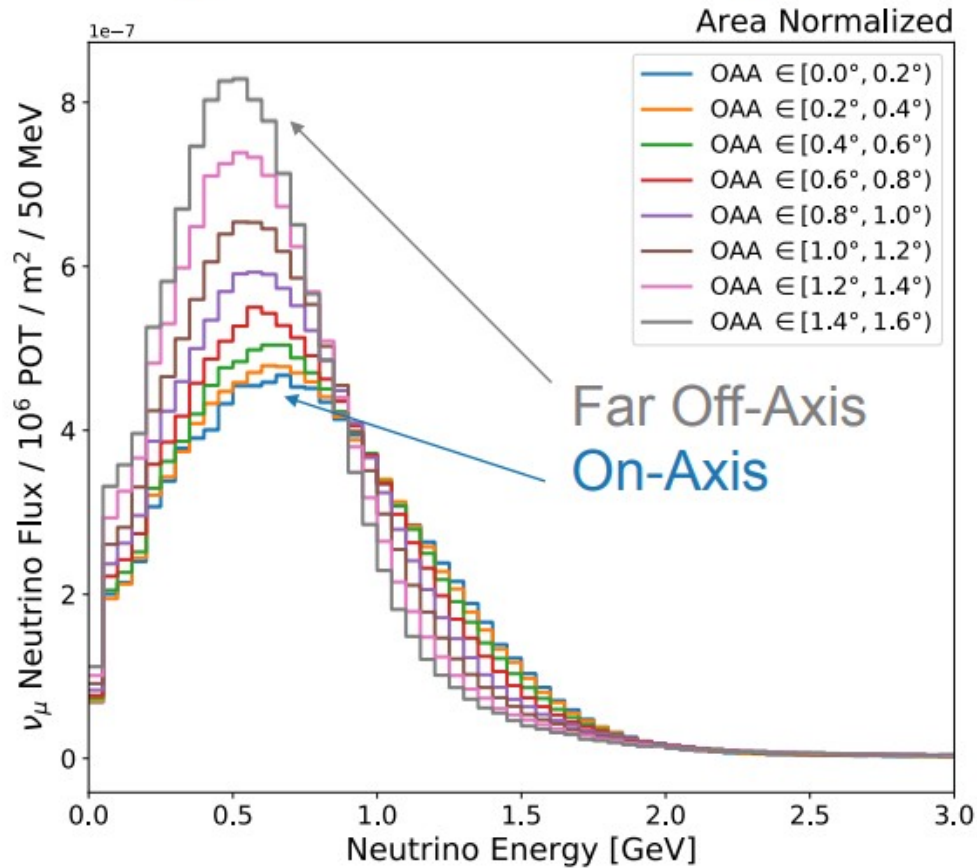


# SBND-PRISM

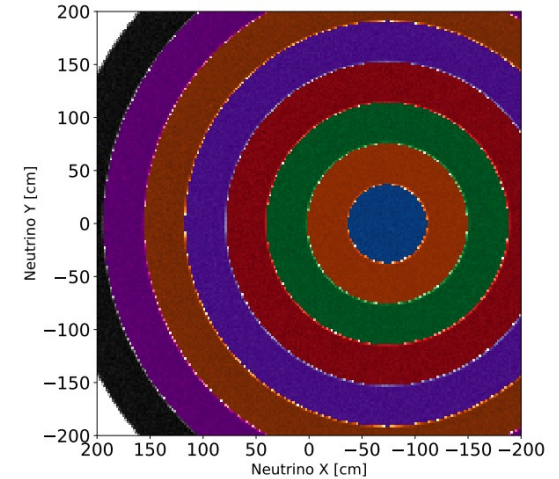


# SBND-PRISM

$\nu_\mu$  flux in each of the OAA regions



Beam view



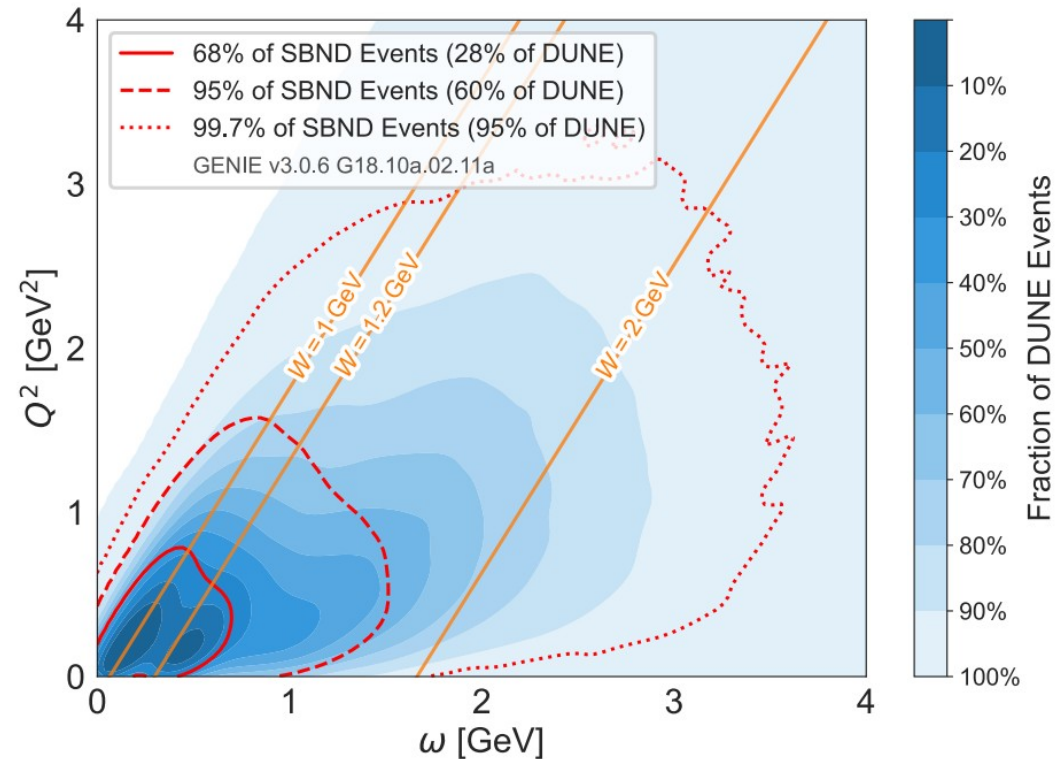
Vary energy dependence by scanning position in detector

Directly determine energy dependence of cross sections



# SBND Kinematic Coverage

- 95% of DUNE phase space covered
- Extremely high statistics



# SBND Status

December 2022 – the big move



# SBND Status

April 2023 – installation



Andrew Furmanski  
University of Minnesota



# SBND Status

May 2023 – a special visit

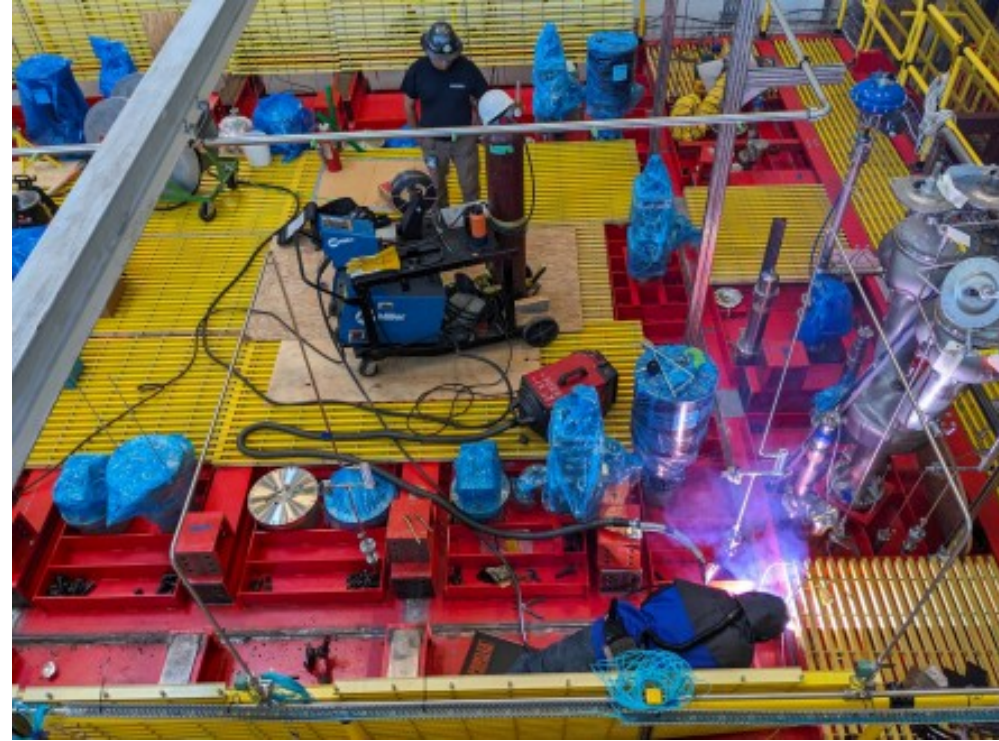


Andrew Furmanski  
University of Minnesota



# SBND Status

June 2023 – first CRT wall installed

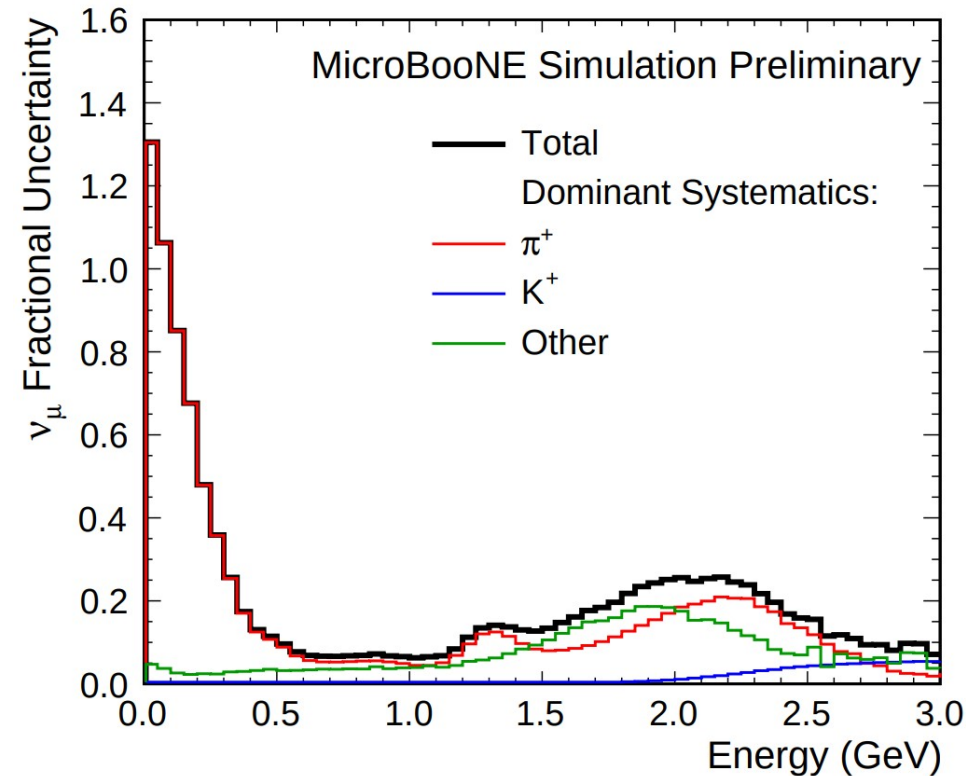


July 2023 – top cap welded in place



# Reducing Uncertainties

- Dominant uncertainty is expected to be the flux
- Campaign underway to reduce these uncertainties:
  - $\nu$ -e scattering (in-situ constraint)
    - O(500) events expected in 3 years: <10% stat. uncertainty
  - New hadron production measurements (ex-situ constraint)





# What's Next?

- Argon fill starts soon!
- Cold commissioning, and CRT installation in late 2023 / early 2024
- Initial Physics Run planned for spring 2024
- After that, **one neutrino every six seconds!**



# Conclusion

- SBND will collect the world's largest dataset of neutrino-argon interactions
  - Within the first month of operations!
- High-granularity TPC with an advanced light collection system
  - Will produce extremely precise measurements
- PRISM effect in a single detector
  - Neutrino energy dependence
- First neutrinos expected next year!



# 감사합니다 !



Andrew Furmanski  
University of Minnesota

