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Hyper-Kamiokande LBL Physics Sensitivity

NuFACT 2023 Seoul

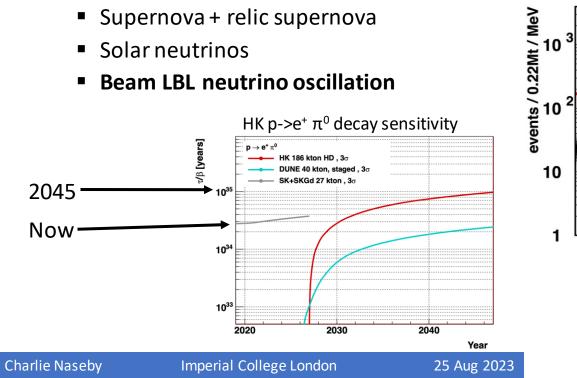
25 August 2023

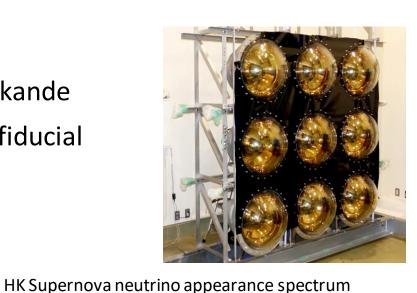
Charlie Naseby

On behalf of the Hyper-Kamiokande collaboration

Hyper-K

- Builds on the success of Super-Kamiokande
- Water Cherenkov detector, 186 kton fiducial
- 20 000 inner detector 50 cm PMTs •
- Rich physics program
 - Proton decay
 - Atmospheric neutrinos
 - Supernova + relic supernova
 - Solar neutrinos
 - Beam LBL neutrino oscillation





10 kpc

+160

10

1

10

20

30

doi:10.3847/1538-4357/abf7c4

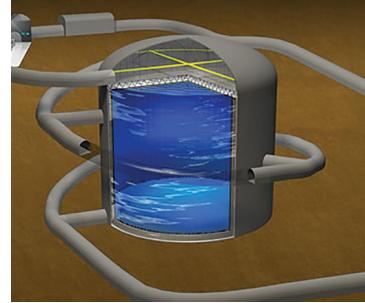
v_+160

40

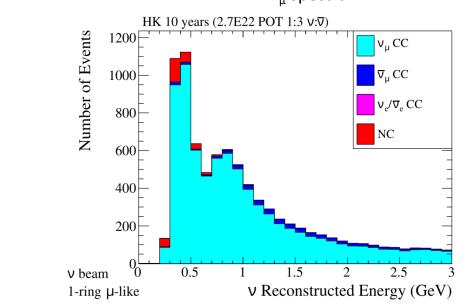
visible energy (MeV)

50

60



HK LBL v_u spectrum

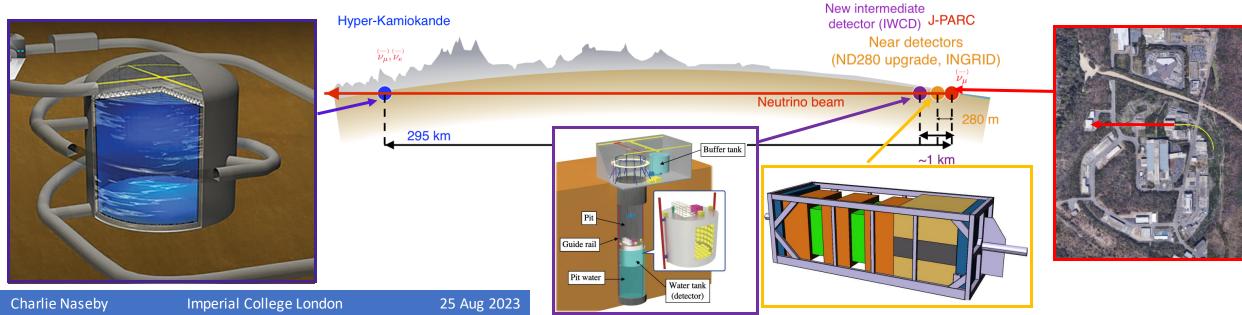


(See talk by N. McCauley for more)



Hyper-K LBL Program

- Builds on success of T2K
- Produce a high-power, selectable v_{μ} or \bar{v}_{μ} dominated neutrino beam at J-PARC
- Measure neutrino flux and cross section with detectors close to beam source, before oscillation
- Observe neutrinos 295km from source in Hyper-K, after they have oscillated
 - Two observable channels: $v_{\mu} \rightarrow v_{\mu}$ and $v_{\mu} \rightarrow v_{e}$
 - Use knowledge of flux and interactions from near detectors to extract oscillation information



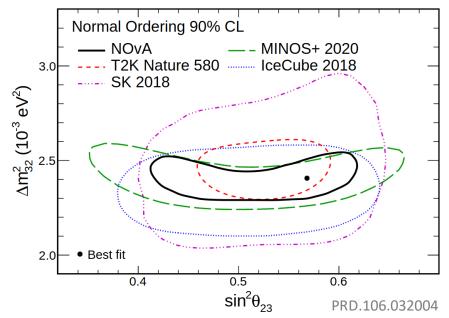
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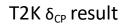
Hyper-K LBL Physics Goals

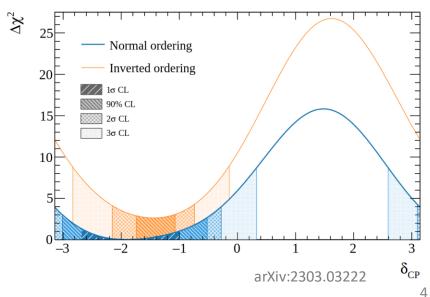
$$U_{PMNS} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \begin{pmatrix} c_{13} & 0 & s_{13}e^{-i\delta_{CP}} \\ 0 & 1 & 0 \\ -s_{13}e^{i\delta_{CP}} & 0 & c_{13} \end{pmatrix} \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

- Does $\sin^2\theta_{23} = 0.5$?
 - If not, which octant is it?
- Precision measurement of Δm^2_{32}
- Independent cross-check on reactor θ_{13} measurements
- Mass ordering with LBL + atmospheric
- Do neutrinos violate CP symmetry?
 - If so, what is the value of δ_{CP}
 - Key information for baryon asymmetry

Global beam & atmospheric results







25 Aug 2023

The Hyper-K LBL Experiment

- To achieve these LBL physics goals, Hyper-K needs:
 - Huge increase in statistics over existing LBL experiments:

186 kton fiducial mass, >8 times that of Super-K

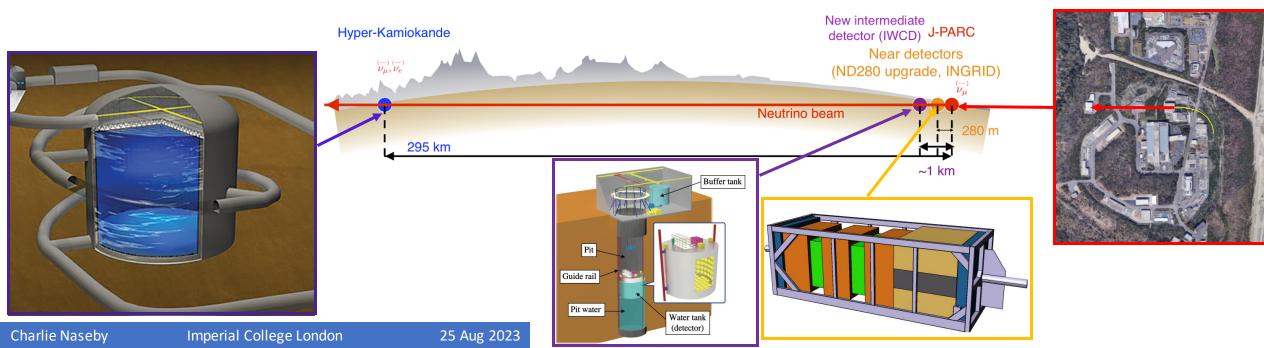
1.3 MW ν_{μ} or $\overline{\nu}_{\mu}$ dominated neutrino beam, more than double T2K beam power

Combined, over 10 years of running, 1:3 v:v mode, expect ≈×40 current T2K far-detector statistics

Improvement in understanding of neutrino flux, cross section and detector effects:

Suite of new and upgraded near detectors

Bottom-up approach to detector systematics



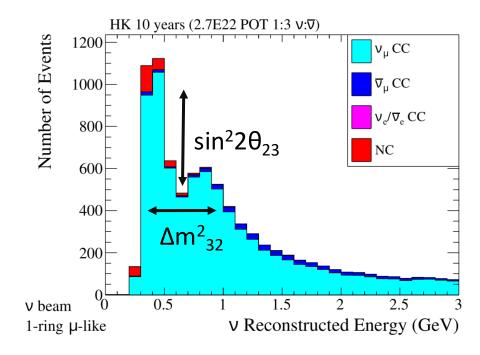
v_{μ} disappearance

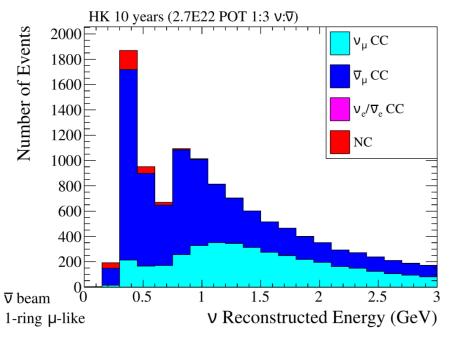
- The majority of events seen at HK will be $v_{\mu} \rightarrow v_{\mu}$
- Position and strength of oscillation dip gives $\Delta m^2_{_{32}}$ and $sin^2 2\theta_{_{23}}$

	v_{μ}	$\overline{\nu}_{\mu}$	v _e	\overline{v}_{e}	$\begin{array}{c} v_{\mu} \\ \rightarrow v_{e} \end{array}$	$\overline{v}_{\mu} \rightarrow \overline{v}_{e}$	NC	Total
v-mode v _µ CCQE-like	8584	480	0.24	0.01	2.32	0.01	283	9349
v-mode v _µ CCQE-like	4399	7688	0.28	0.24	0.33	0.42	286	12375

 $\nu_{\mu}\text{-like}$ events, 10 years, sin² $\theta_{23}\text{=}0.58,\,\Delta m^2{}_{32}\text{=}2.509x10^{\text{-3}}\text{eV}^2,$ normal ordering

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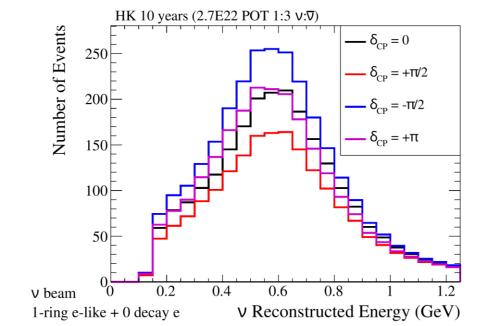


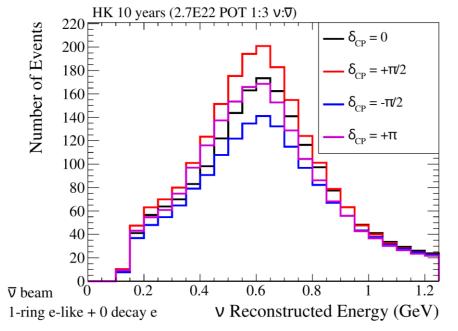
v_e appearance

- Sensitivity to δ_{CP} and octant of θ_{23} comes primarily from $\nu_{\mu} \rightarrow \nu_{e}$ oscillations
- δ_{CP} changes v_{e} and \overline{v}_{e} oscillation probability in opposite directions

	$δ_{CP} = -\pi/2$	$\delta_{CP} = 0$	$\delta_{CP} = +\pi/2$	$δ_{CP} = +π$
v-mode v _e CCQE-like	2740	2285	1846	2301
v-mode v _e CC1π-like	258	223	179	214
⊽-mode v _e CCQE-like	1624	1883	2118	1859

 v_e -like events, 10 years, sin² θ_{23} =0.58, Δm_{32}^2 =2.509x10⁻³eV², normal ordering

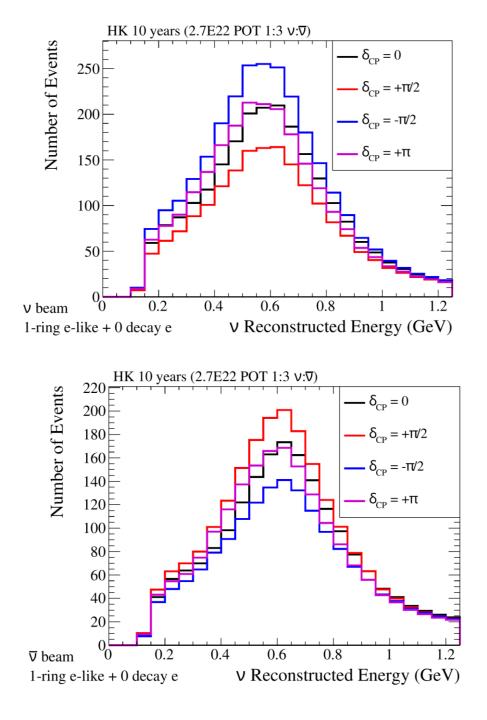






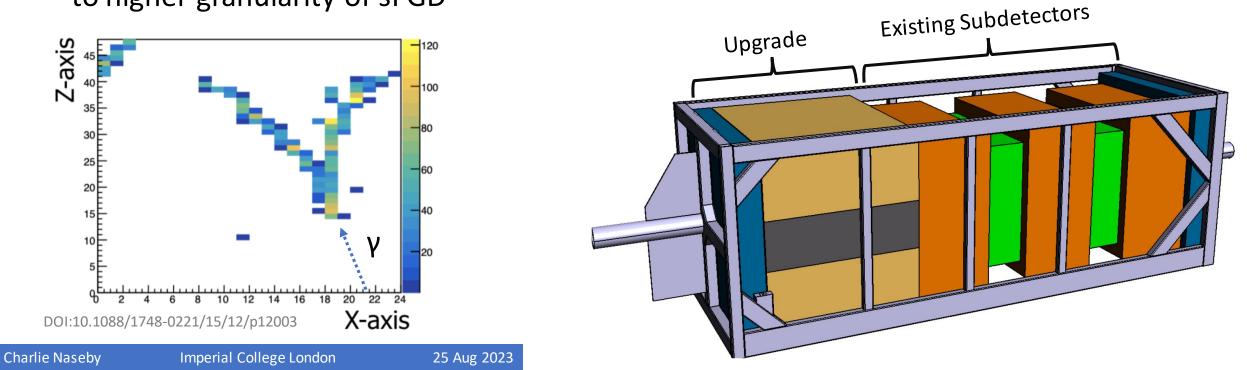
Aside: v_e cross section

- HK's δ_{CP} sensitivity derives from $\nu_e/\overline{\nu}_e$ appearance rate
- v_e/v_e cross-section ratio can also affect this
- T2K currently uses a 4.9% prior uncertainty on this ratio with very little data constraint.
- This would limit δ_{CP} sensitivity, HK intends to measure this ratio as precisely as possible, with current studies focused on a 2.7% precision



ND280

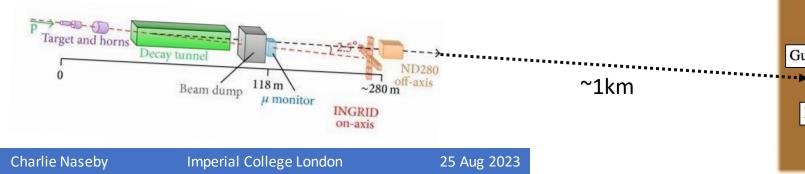
- Magnetised plastic scintillator and TPC near detector
- Constrain neutrino flux, interaction cross section and wrong-sign component
- Currently undergoing upgrade with higher granularity super Fine Grained Detector (sFGD) and high angle TPCs. (See talks by L. Munteanu, D. Nguyen)
- Upgrade provides better e/γ separation and lower proton momentum threshold due to higher granularity of sFGD

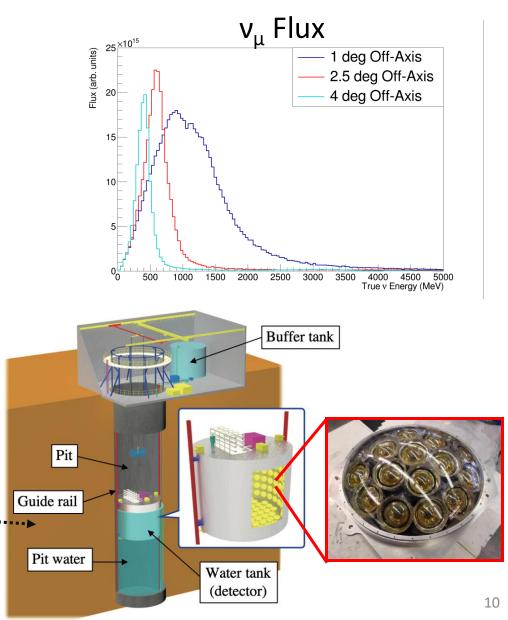


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Intermediate Water Cherenkov Detector

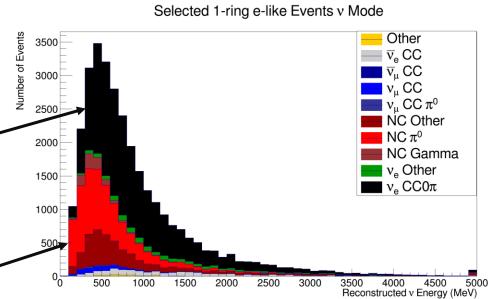
- 600 t water Cherenkov detector ~1km from the beam in a vertical shaft
- Pion/muon decay kinematics results in different neutrino fluxes along the shaft (see M. Wilking's talk)
- Small multi-PMTs provide high-granularity and time resolution in a relatively small detector
- Excellent e/ μ PID, combined with large target mass, IWCD can select the 1% v_e component of the beam

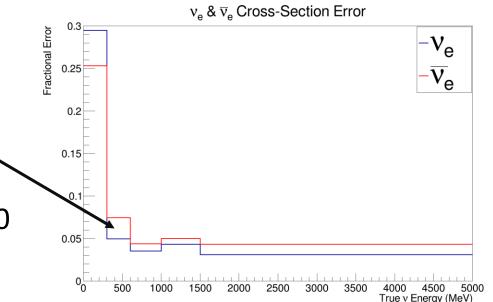




IWCD v_e measurement

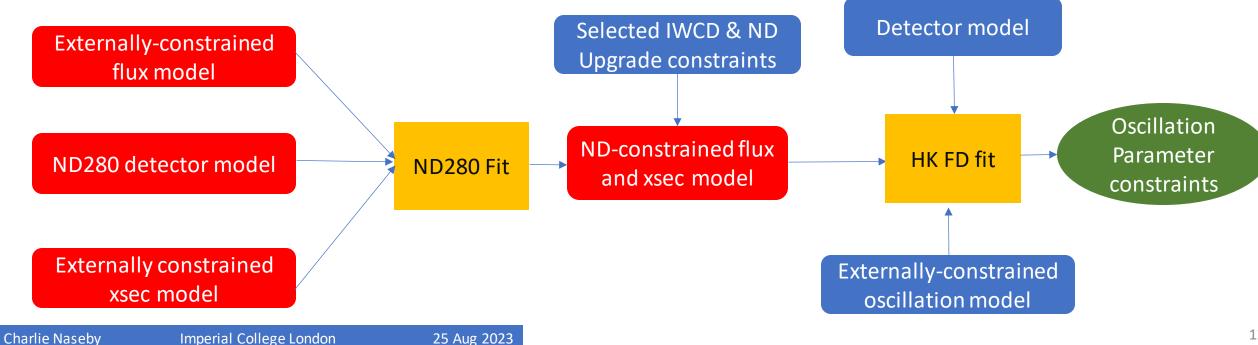
- Current IWCD fit uses six samples: v_e , v_μ and NC π^0 , each in v and \overline{v} mode
- Select over 18000 v_e events in v-mode \sim
- v_{μ} (>1M v mode events) constrains CC cross section & flux, NC π^0 constrains NC background to v_e samples \swarrow
- Integrating over HK v_e appearance spectrum, obtain 2.94% uncertainty on v_e/v_e rate ratio with analysis improvements possible
- Total uncertainty is lower than on each individual parameter due to correlations
- Preliminary studies with ND280 upgrade + IWCD achieve better than 2.7% uncertainty, helped by ND280 charge selection





Analysis Procedure

- To obtain oscillation parameter constraints from these HK spectra, adopt an approach similar to T2K
- Use a near detector to constrain a model of flux and neutrino interaction cross-section
- Add selected constraints from the IWCD fit and ND upgrade studies
- Use this constrained systematic model in an oscillation fit to the far-detector data



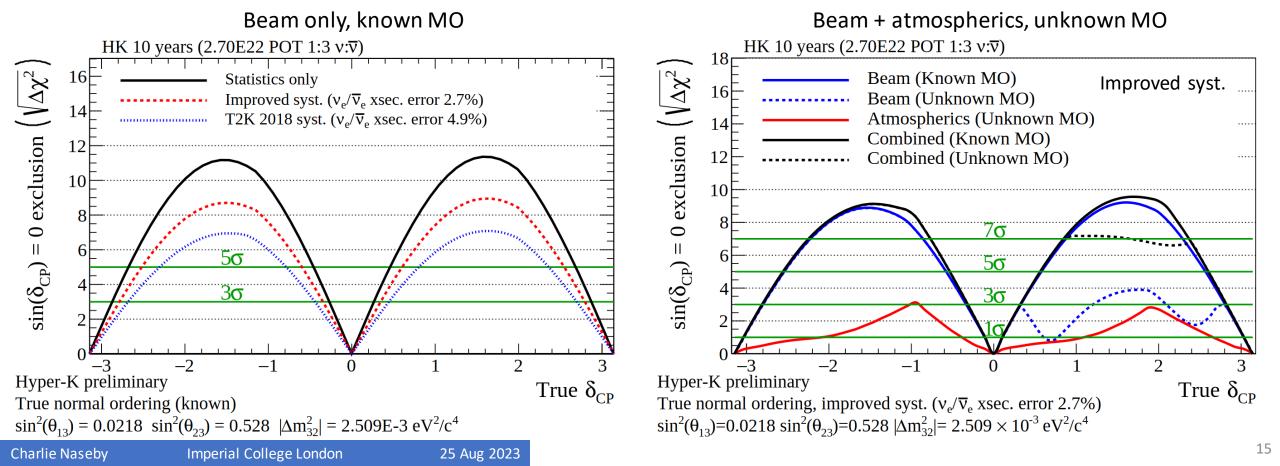
Systematic modelling

- Making use of T2K 2018 model for flux, cross section and detector systematics.
- Three scenarios investigated:
 - Statistics only, no flux, cross section or detector uncertainty
 - T2K 2018 systematics, T2K ND fit with 2018 statistics and 2018 SK detector uncertainties
 - HK Improved systematics, T2K 2018 ND fit systematics scaled by sqrt(N_{T2K}/N_{HK}), preserving correlations
 - A 1% minimum uncertainty is enforced on all parameters
 - Additional constraint from IWCD and ND280 upgrade on specific parameters
 - 2.7% error on v_{e}/\overline{v}_{e} cross-section ratio

Physics Sensitivities

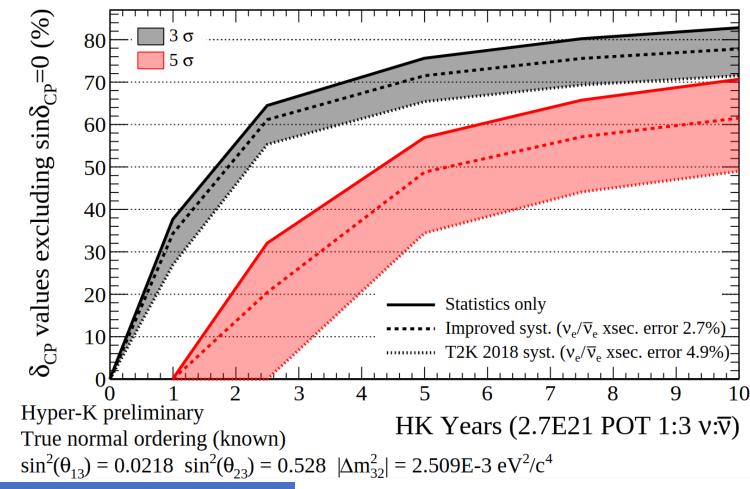
CP violation sensitivity

- Ability to exclude $\delta_{CP} = 0$, $\pm \pi$ as a function of true δ_{CP}
- Systematic uncertainties play a key role in sensitivity
- Beam + atmospherics improves sensitivity over beam alone when MO is unknown



CP violation sensitivity

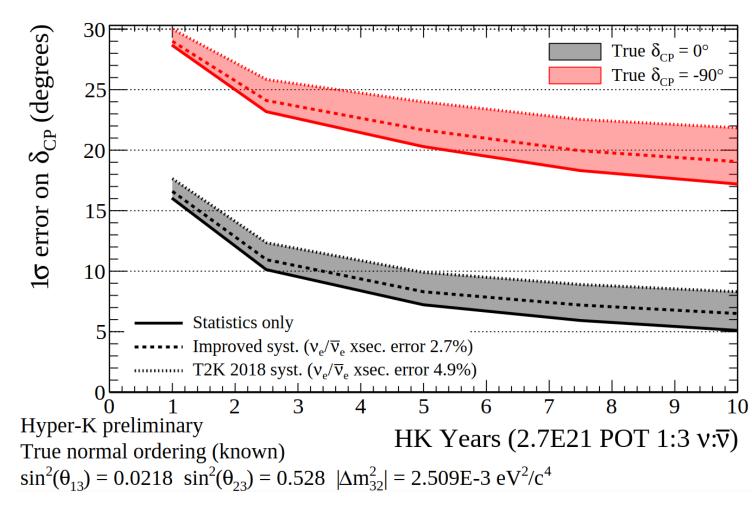
- Fraction of δ_{CP} values excluding CP conservation as a function of exposure
- Exclude CP conservation to 5σ for >60% of true δ_{CP} values, >75% to 3σ



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- Ranges between 6 and 20 degrees
- Lowest precision around maximal CP violation where v_e event rate changes most slowly
- v_e/v
 _e cross-section is leading systematic and most significant close to CP conserving values



θ_{23} Precision & Octant

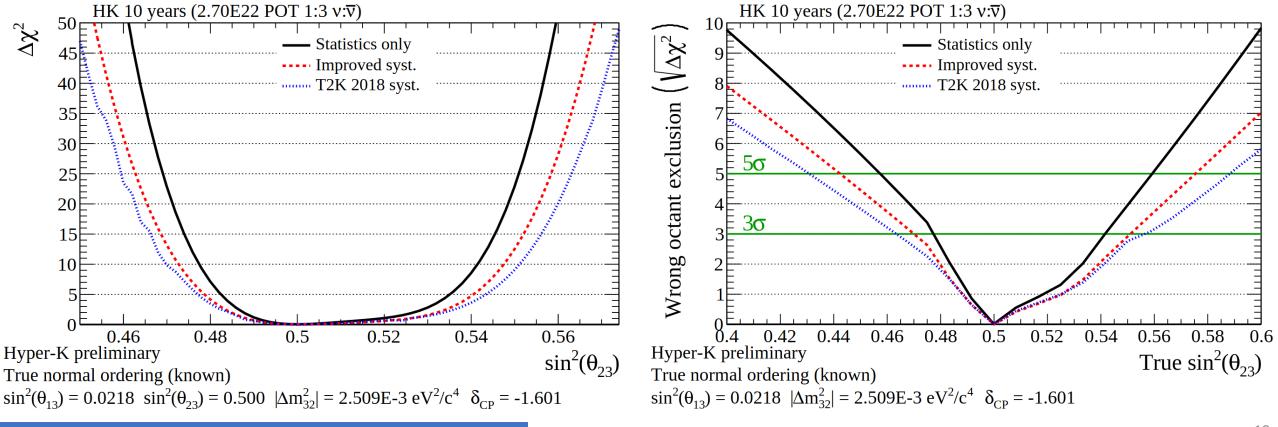
• Measurement precision depends strongly on true θ_{23} value

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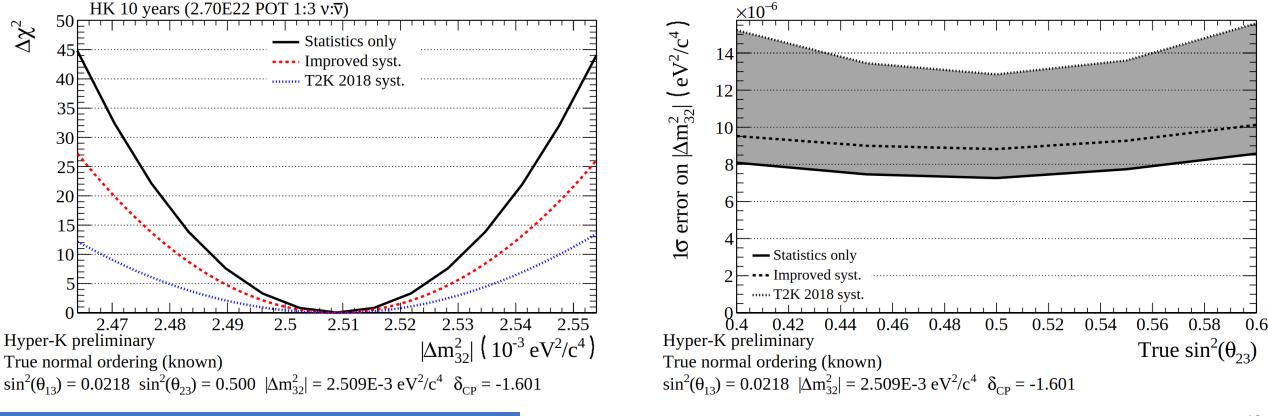
• Beam only, improved systematics can exclude wrong-octant to 3σ for true $sin^2\theta_{23} < 0.47$ or $sin^2\theta_{23} > 0.55$

25 Aug 2023



Δm^2_{32}

- 1 σ precision on Δm^2_{32} has a small dependence on θ_{23} value
- 1 σ precision measurement 9×10⁻⁶ eV² possible, cf. 50×10⁻⁶ eV² from T2K (arXiv:2303.03222)
- Improved systematic model increases precision by 30% over existing systematics



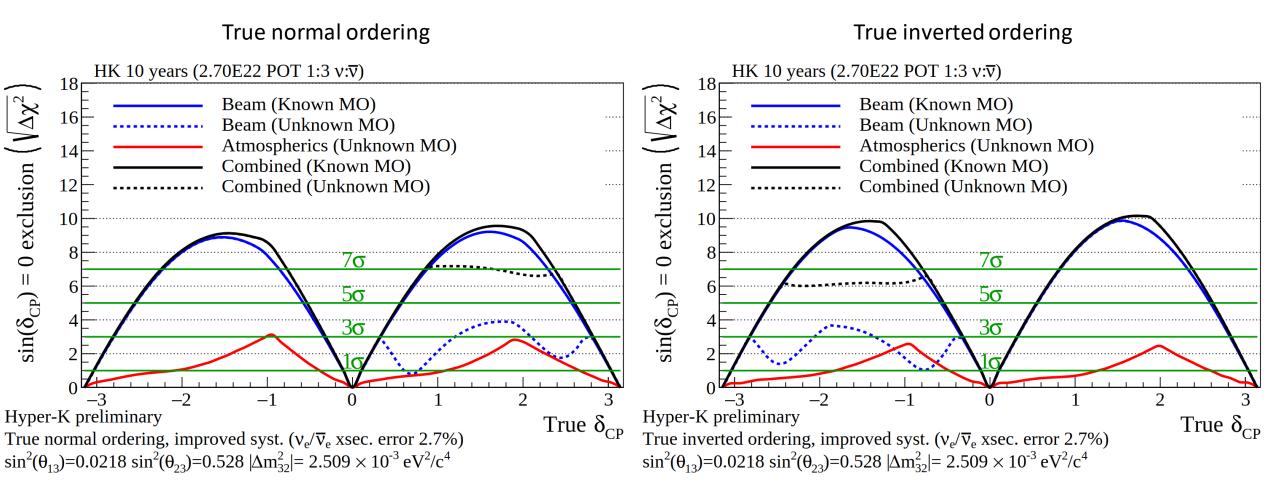
Conclusions

- HK will provide great sensitivity to neutrino oscillation parameters:
 - 5 σ CP violation sensitivity for >60% of true δ_{CP} values
 - 6-20 degree precision on measurement of δ_{CP} depending on true value
 - 3σ rejection of wrong-octant of θ_{23} for $\sin^2\theta_{23} < 0.47$ or $\sin^2\theta_{23} > 0.55$
 - 0.4% precision on Δm^2_{32} measurement
- Ability to constrain systematic uncertainties will be key to precise measurement
- Development of analysis tools for HK is ongoing
- Construction is progressing on track for 2027 data taking

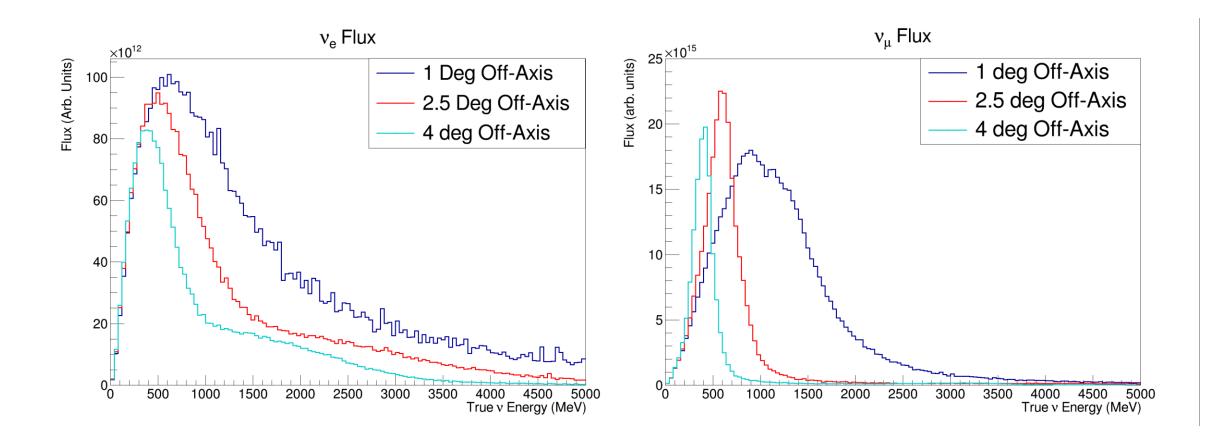


Backups

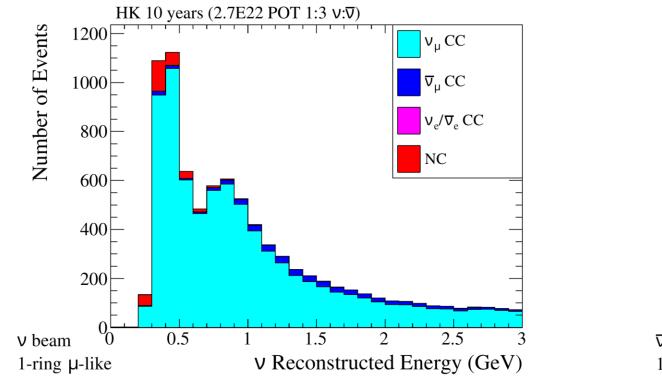
CP sensitivity unknown MO

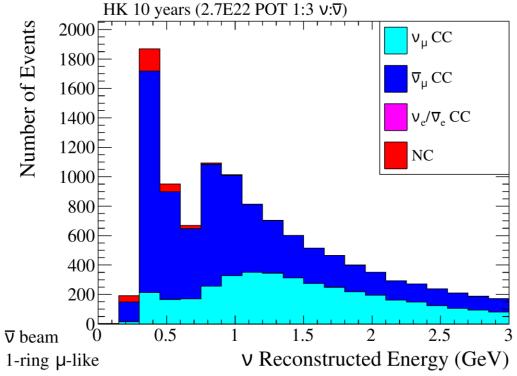


IWCD Fluxes

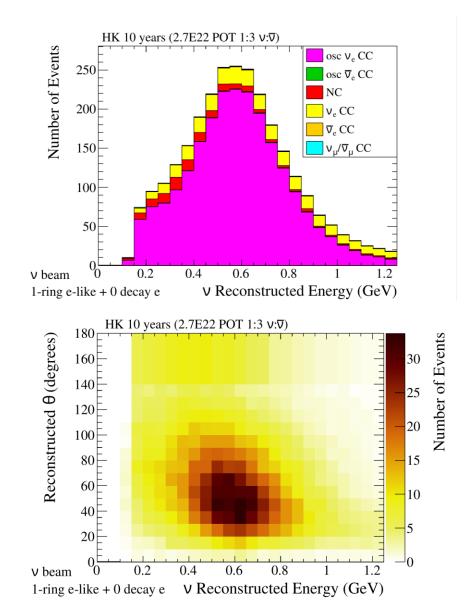


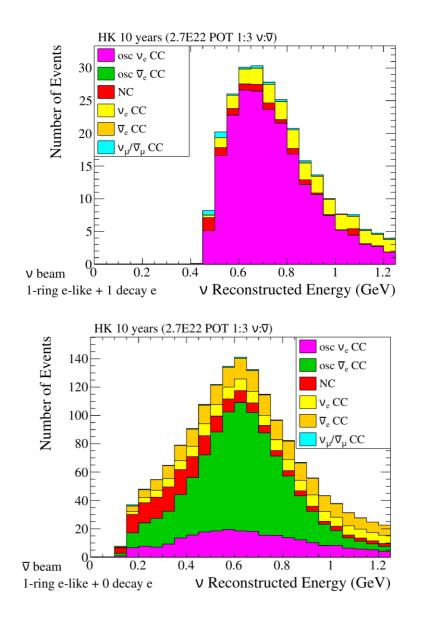
NuMu appearance spectra





NuE appearance spectra





NuMu oscillation probability

$$\begin{split} P\left(\stackrel{(\leftarrow)}{\nu_{\mu}} \to \stackrel{(\leftarrow)}{\nu_{\mu}}\right) &\approx 1 - 4\left((s_{23}^2 c_{13}^2)(1 - s_{23}^2 c_{13}^2)\right)\sin^2\left(\Delta_{32} + \Delta_{21}\frac{s_{12}^2 c_{23}^2}{1 - s_{23}^2 c_{13}^2}\right) \\ &+ \mathcal{O}\left(a^2, \Delta_{21}\right) \\ &\approx 1 - \sin^2 2\theta_{23}\sin^2\left(\Delta_{32}\right) \end{split}$$

NuE Appearance Probability

$$P\left(\stackrel{(\leftrightarrow)}{\nu_{\mu}} \rightarrow \stackrel{(\leftrightarrow)}{\nu_{e}}\right) \approx \sin^{2}\theta_{23}\sin^{2}2\theta_{13}\frac{\sin^{2}\left(\Delta_{31} \mp a\right)}{\left(\Delta_{31} \mp a\right)^{2}}\Delta_{31}^{2}$$

+ $\sin 2\theta_{23}\sin 2\theta_{13}\sin 2\theta_{12}\cos\theta_{12}\frac{\sin\left(\Delta_{31} \mp a\right)}{\Delta_{31} \mp a}\Delta_{31}\frac{\sin a}{a}\Delta_{21}\cos\Delta_{31}\cos\delta_{CP}$
 $\mp \sin 2\theta_{23}\sin 2\theta_{13}\sin 2\theta_{12}\cos\theta_{12}\frac{\sin\left(\Delta_{31} \mp a\right)}{\Delta_{31} \mp a}\Delta_{31}\frac{\sin a}{a}\Delta_{21}\sin\Delta_{31}\sin\delta_{CP}$
+ $\cos^{2}\theta_{13}\cos^{2}\theta_{23}\sin^{2}2\theta_{12}\frac{\sin^{2}a}{a^{2}}\Delta_{21}^{2}$

J-PARC beam

- 30GeV J-PARC proton main ring
- Intend to run at 1.16 s repetition, up to 1.3 MW

0

Three 320 kA focusing horns

