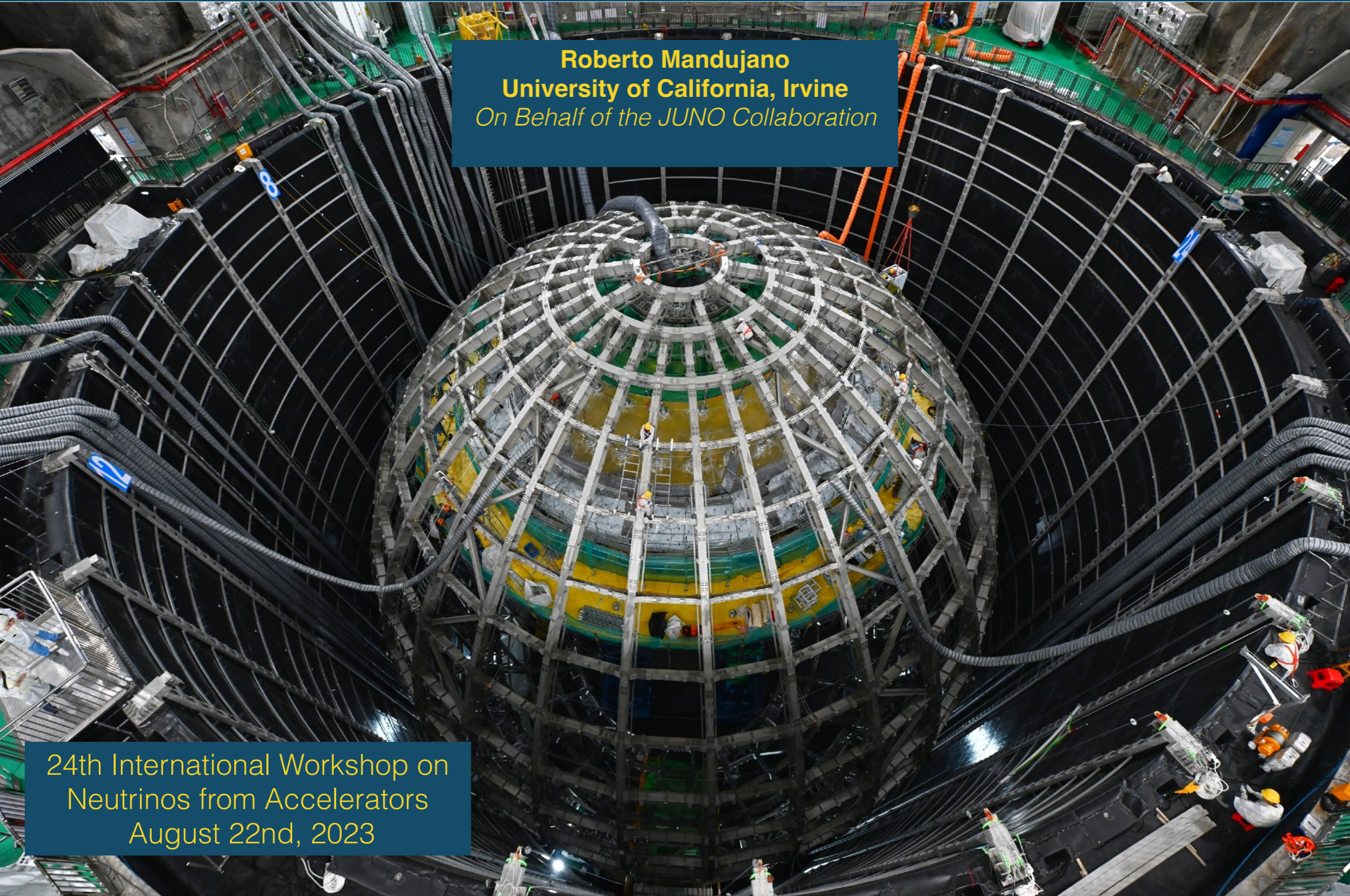


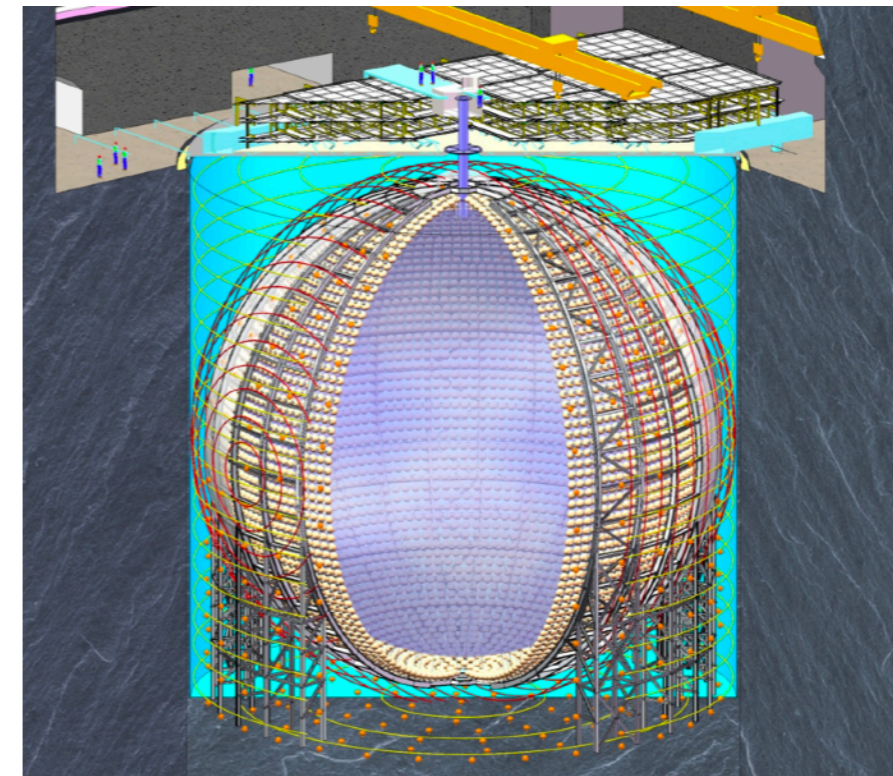
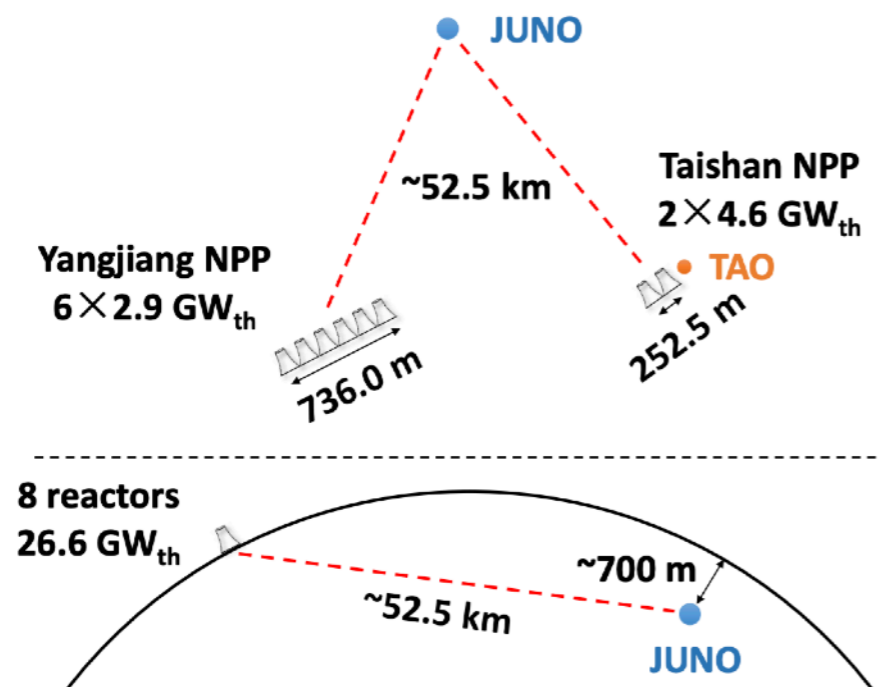
# Reactor Antineutrino Oscillations with JUNO

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*On Behalf of the JUNO Collaboration*

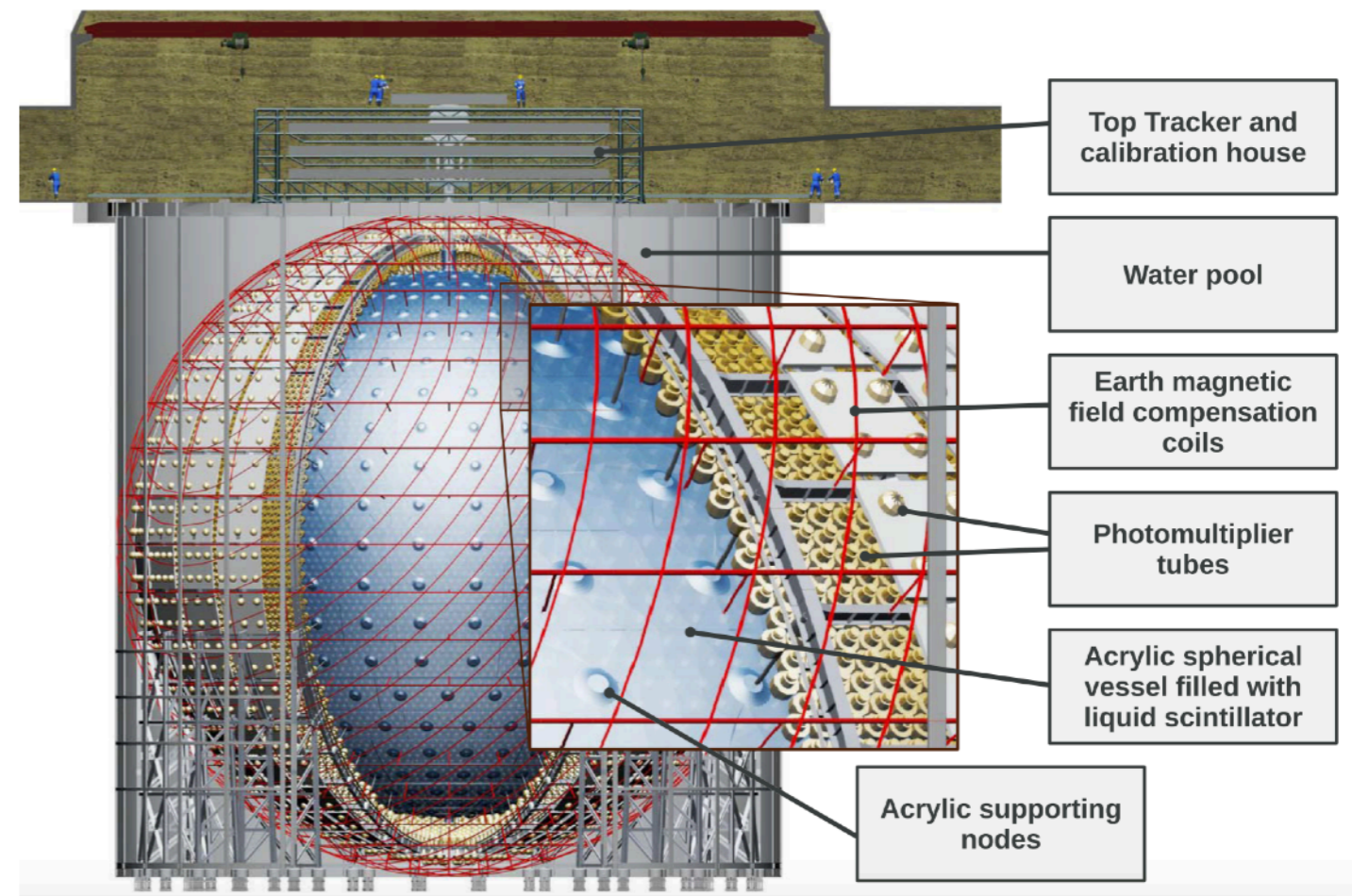
24th International Workshop on  
Neutrinos from Accelerators  
August 22nd, 2023



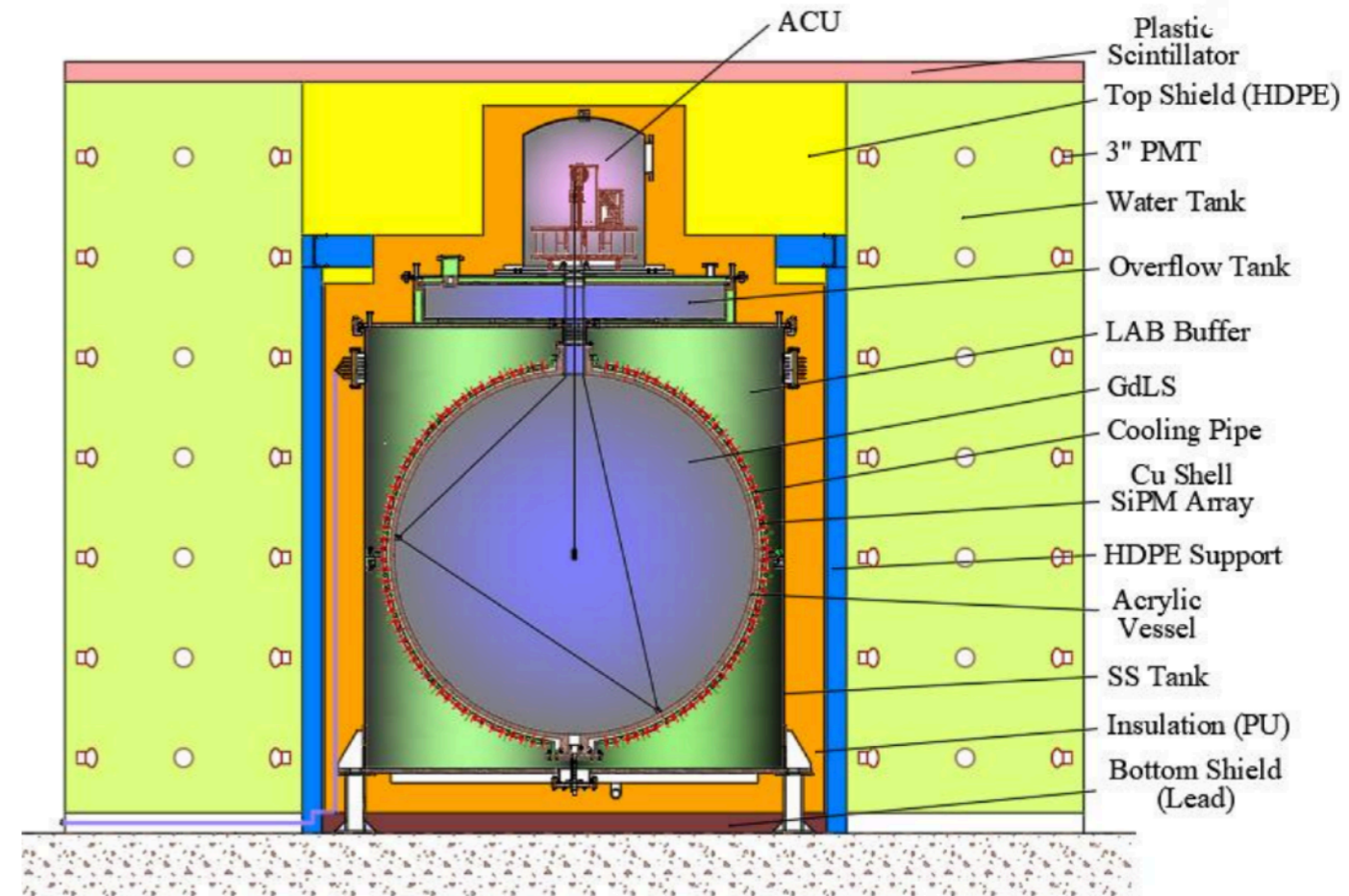
- The Jiangmen Underground Neutrino Observatory (JUNO) is a large liquid scintillator neutrino detector under construction in China
- At 20 kton, it will be the largest detector of its kind ever built
- Multi-purpose experiment with rich physics portfolio
- Main physics goals are neutrino mass ordering (NMO) determination and precision measurement of neutrino oscillation parameters



- 35.4 m diameter acrylic sphere filled with 20 kton of liquid scintillator (LS)
  - ◆ LS designed for high light yield and low attenuation
- 17,612 20" PMTs (LPMT) and 25,600 3" PMTs (SPMT)
  - ◆ ~78% photo-coverage
  - ◆ ~30% detection efficiency (LPMT)
- Instrumented outer water tank and top scintillator panels serve as muon background veto, reconstruct muon tracks, and provide shielding from natural radiation
- Unprecedented 3% energy resolution at 1 MeV

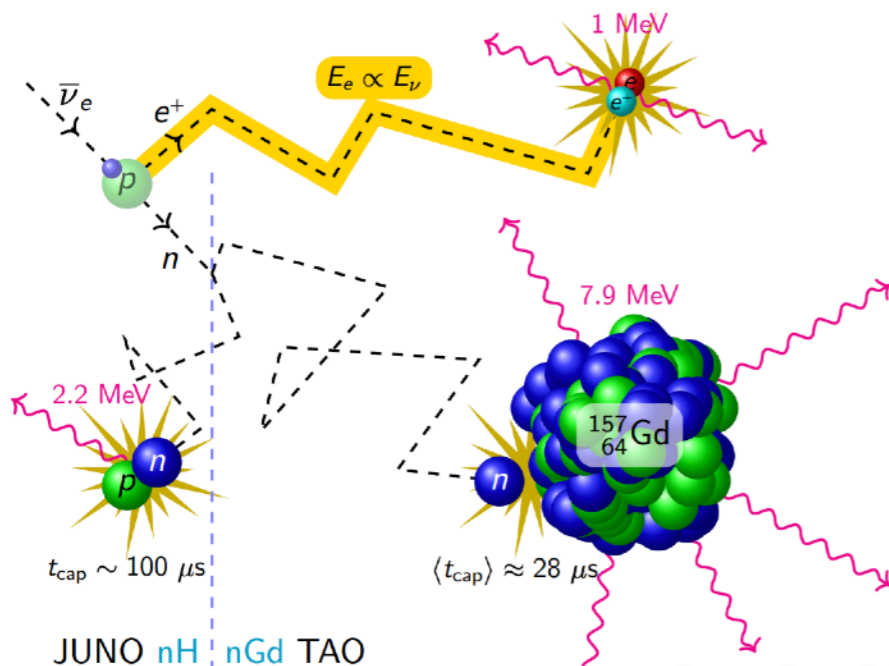
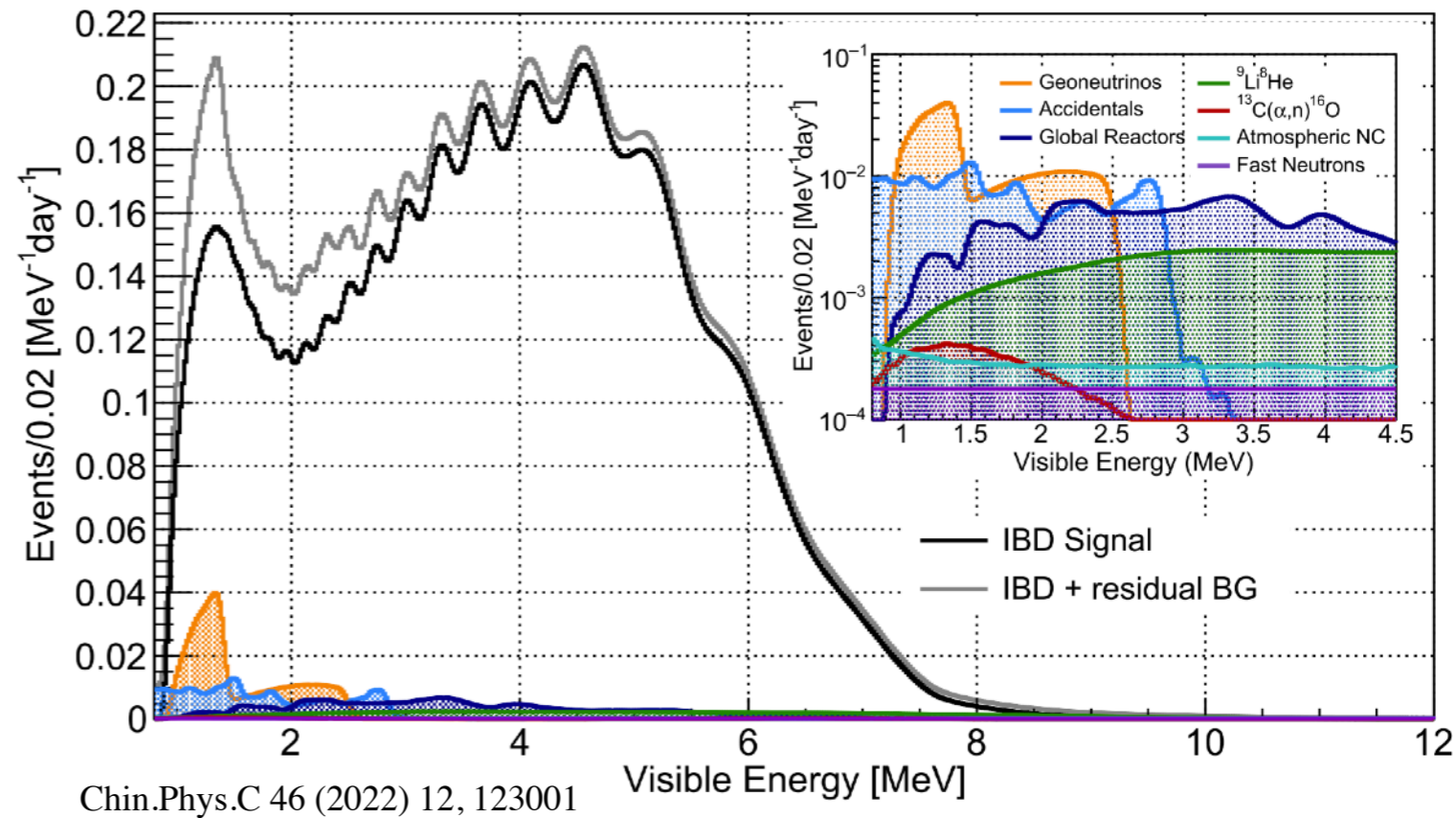


- JUNO-TAO (Taishan Antineutrino Detector) will be a satellite detector
  - ◆ 44 m from 4.6 GW<sub>th</sub> reactor
  - ◆ ~1 ton GdLS fiducial volume
  - ◆ Instrumented with SiPM providing <~2% at 1 MeV energy resolution
  - ◆ Operates at -50 C
  
- **Measure reactor antineutrino energy spectrum** with excellent resolution
  - ◆ Reference spectrum for JUNO, removing possible model dependence from NMO measurement
  - ◆ Search for sterile neutrinos



arXiv: 2005.08745

- Reactor  $\bar{\nu}_e$  detected through Inverse Beta Decay reaction
  - $\bar{\nu}_e + p \rightarrow e^+ + n$
  - Positron (prompt) signal followed by neutron capture (delayed)
  - Temporal and spatial coincidence of prompt and delayed signals is a powerful handle to extract reactor neutrino signal



Event Type	Correlated						Uncorrelated	
	Reactor IBD	Fast-n	<sup>9</sup> Li/ <sup>8</sup> He	Alpha-n	Geo-ν	Global Reactor	Accidentals	Atm.-ν
Rate [per day]	47	0.1	0.8	0.05	1.2	1	0.8	0.16

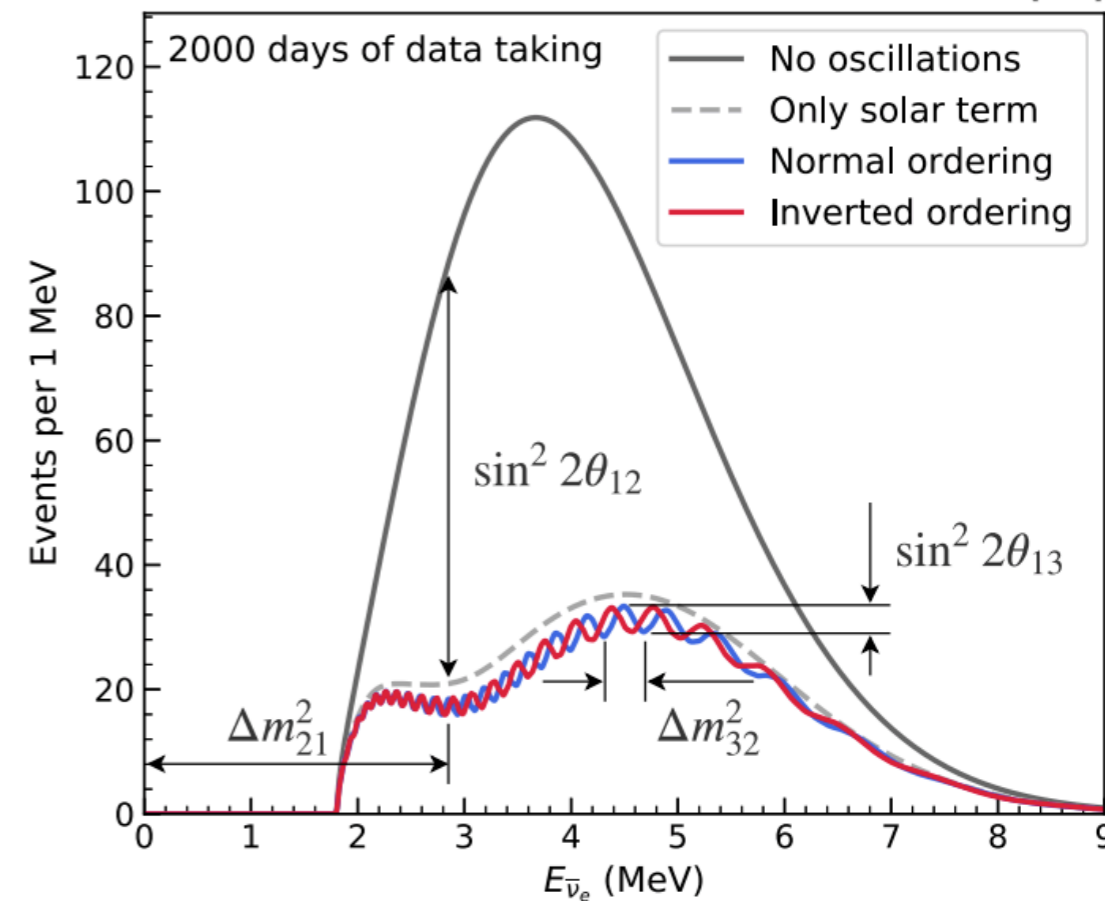
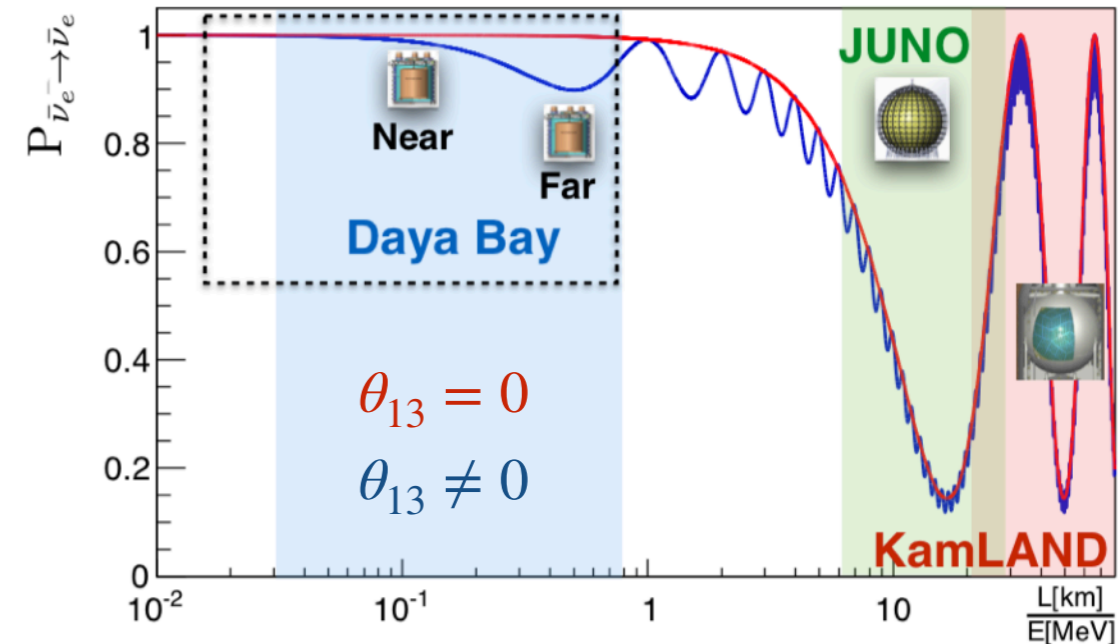
Medium baseline

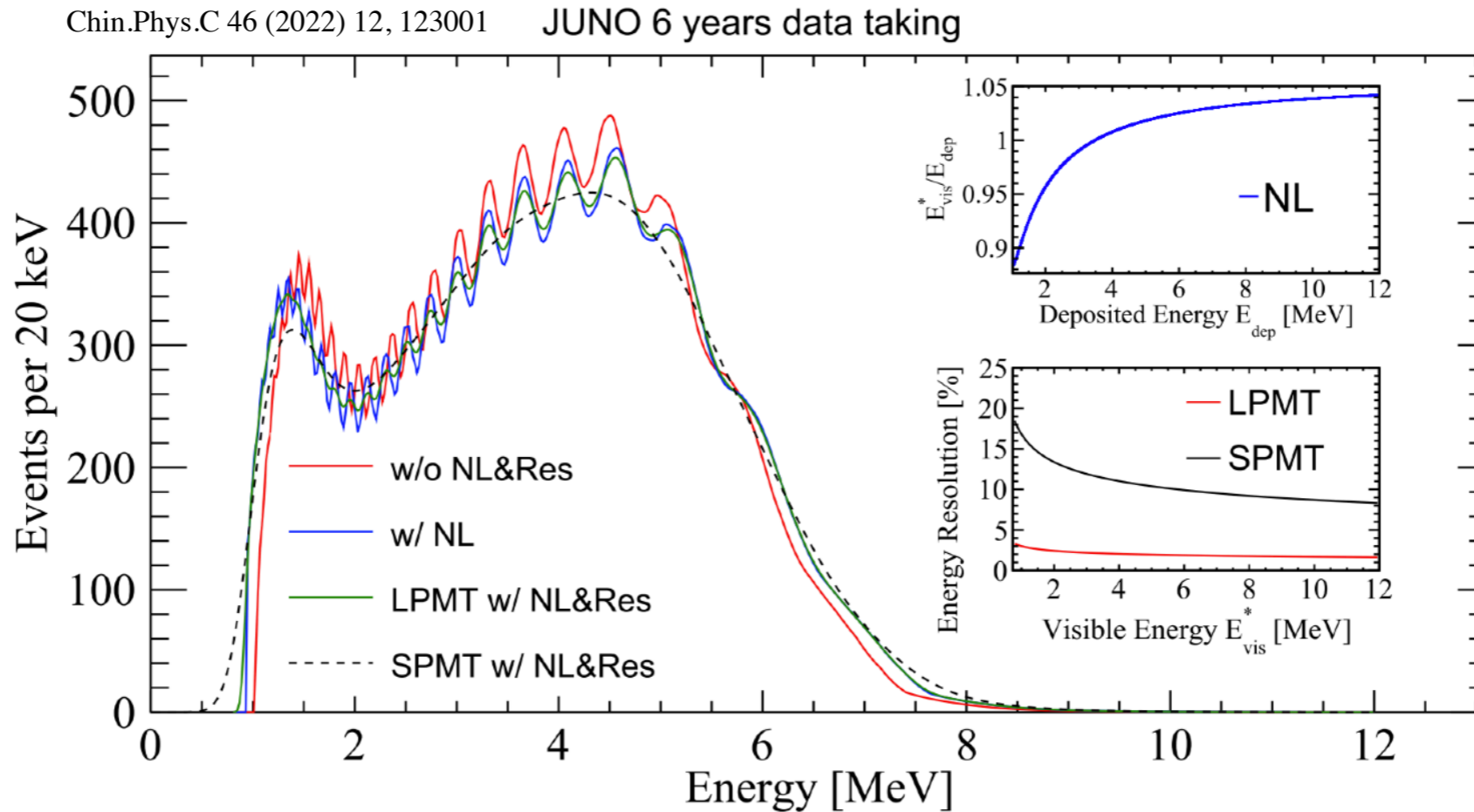
Short baseline

$$P_{\bar{\nu}_e \rightarrow \bar{\nu}_e}(L, E) = 1 - \sin^2 2\theta_{12} \cos^4 \theta_{13} \sin^2 \frac{\Delta m_{21}^2 L}{4E} - \sin^2 2\theta_{13} \left( \cos^2 \theta_{12} \sin^2 \frac{\Delta m_{31}^2 L}{4E} + \sin^2 \theta_{12} \sin^2 \frac{\Delta m_{32}^2 L}{4E} \right)$$

- Observation of  $\theta_{12}$ ,  $\theta_{13}$ ,  $\Delta m_{21}^2$ , and  $\Delta m_{32}^2$  driven oscillations
- Determination of NMO through interference effects in fine structure of oscillated spectrum (allowed by large  $\theta_{13}$ )
  - ◆ JUNO-TAO energy spectrum provides precise reference of un-oscillated JUNO energy spectrum
  - ◆ Independent of  $\delta_{cp}$ , octant of  $\theta_{23}$ , no reliance on matter effects
  - ◆ Complementary to accelerator and atmospheric measurement (different baseline, energy, flavor and technology)

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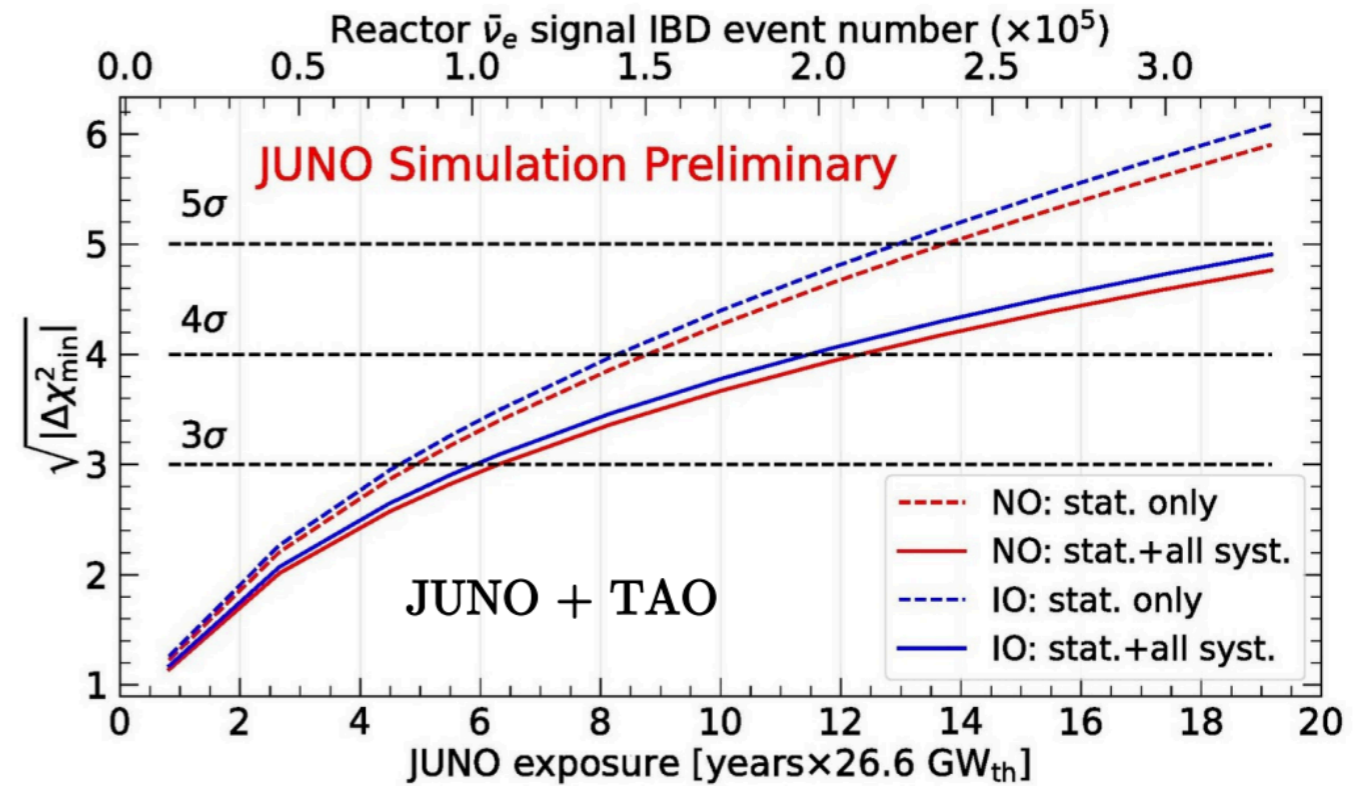


- Recent sensitivity study with full treatment of systematics using updated assumptions and measurements:
  - Updated number of reactors
  - Realistic simulation and veto efficiencies
  - Final detector overburden and location information
  - Spectral shape constraints from JUNO TAO included

- Fit energy spectrum under normal and inverted ordering hypotheses

$$\Delta\chi_{min}^2 = | \Delta\chi_{min}^2(NO) - \Delta\chi_{min}^2(IO) |$$

- ~3σ NMO sensitivity in 6 years
  - Independent of  $\delta_{cp}$ , octant of  $\theta_{23}$
  - Complementary to accelerator measurement (different baseline and technology)
  - Paper under preparation



	$\Delta\chi_{min}^2$	stat. + 1 syst.
Statistics	11.3	
Stat.+Flux error	-0.6	
Stat.+Backgrounds	-1.4	
Stat.+Nonlinearity	-0.4	
Stat.+Others	< -0.05	
Total	9.0	

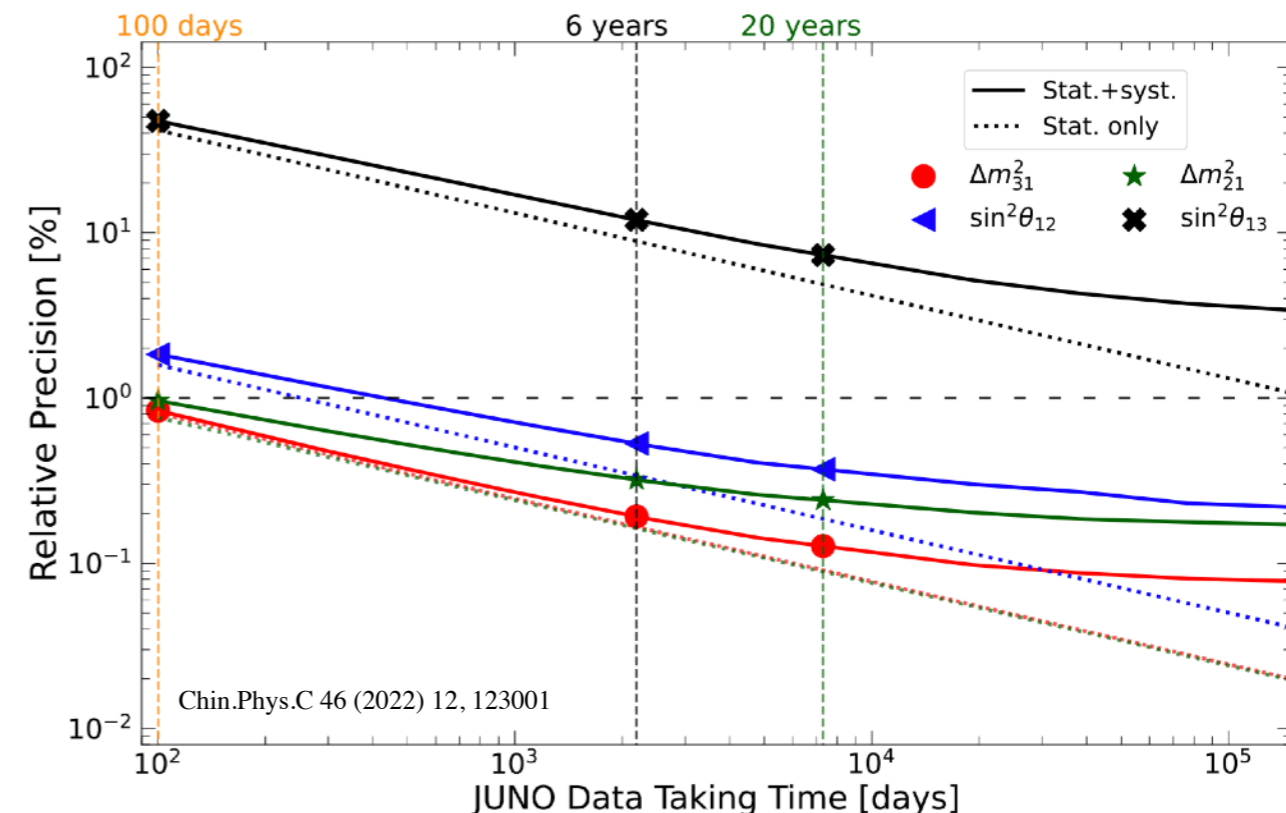
JUNO Simulation Preliminary



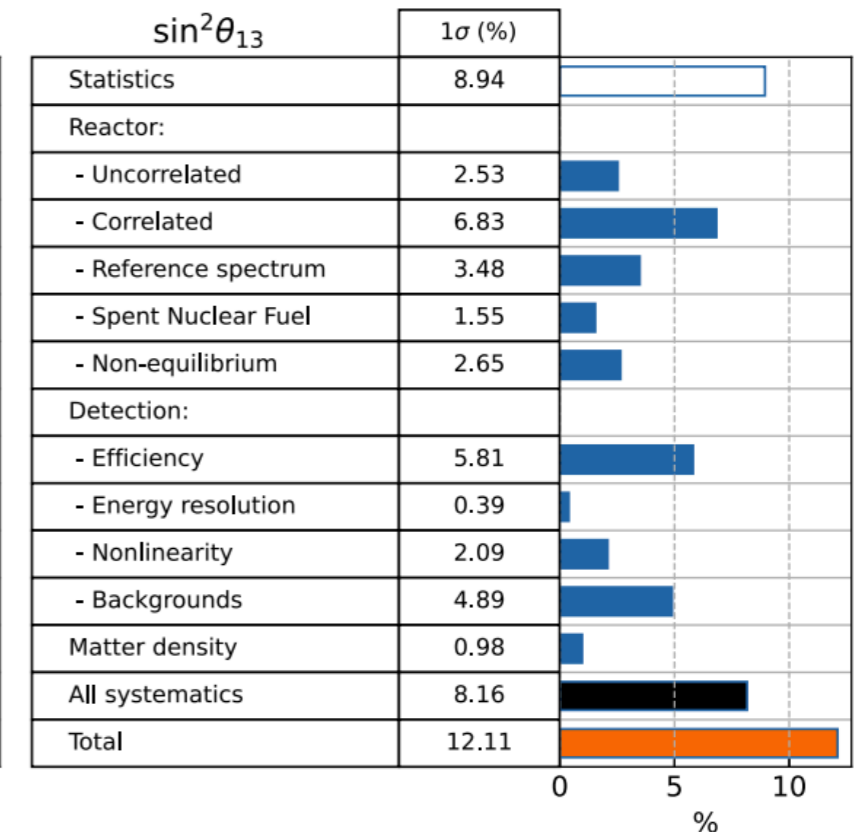
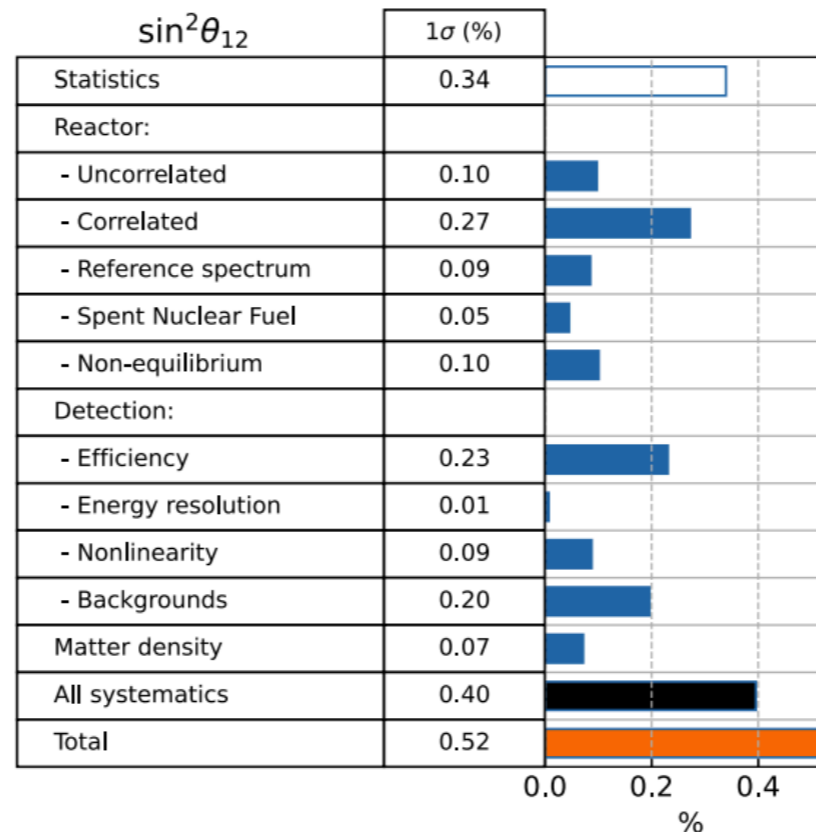
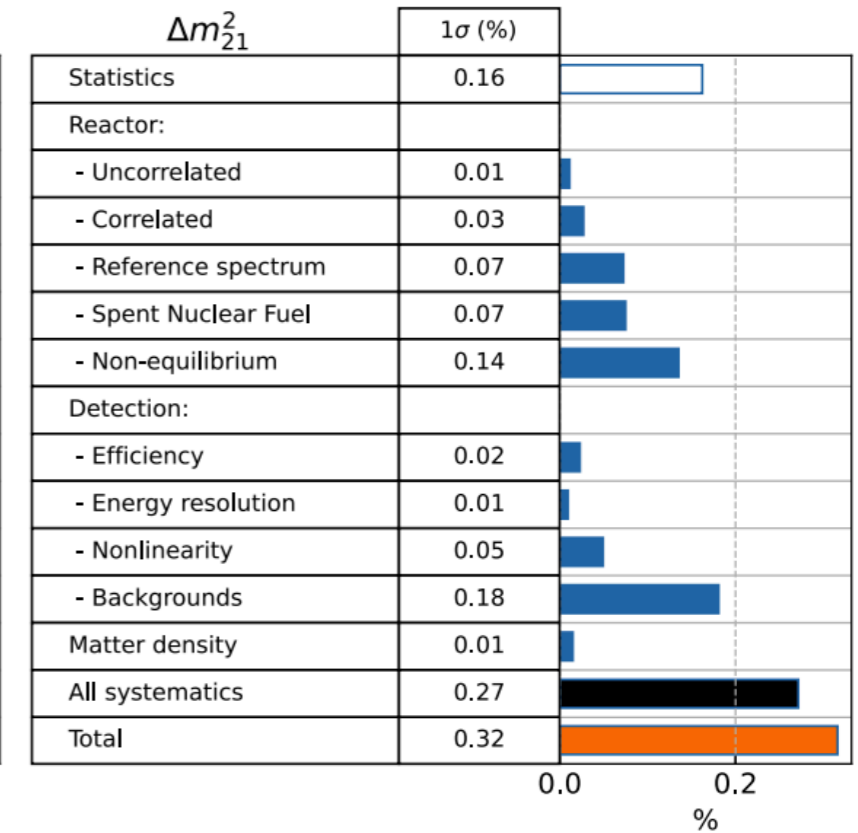
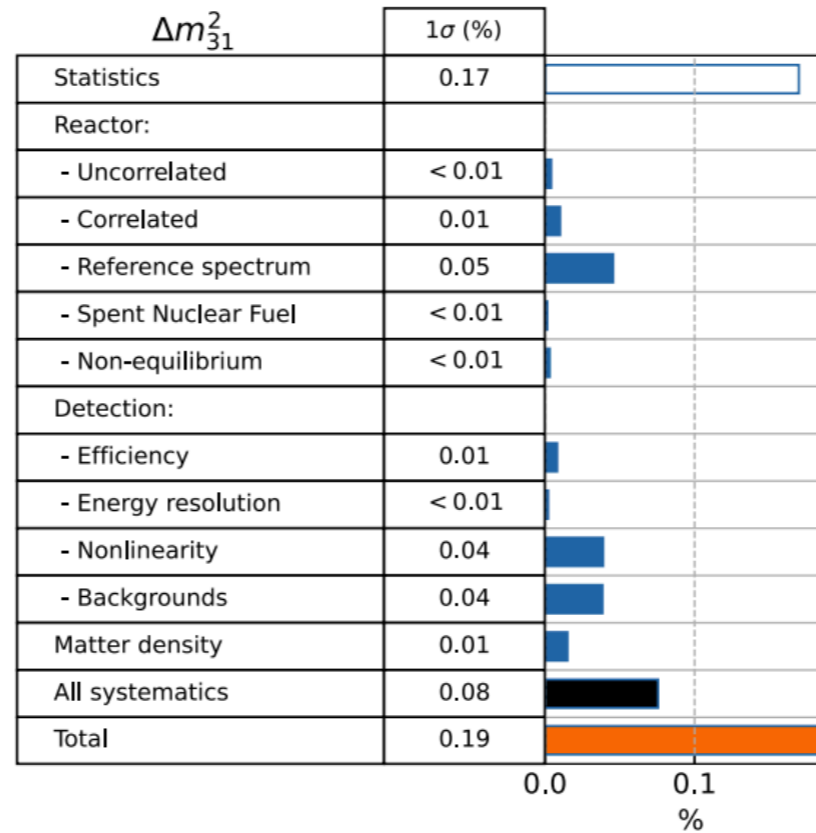
- Measurement of  $\sin^2\theta_{12}$ ,  $\Delta m_{21}^2$  and  $\Delta m_{31}^2$  to **sub-percent precision with O(100 days) data**
  - ◆ 6 years for **order of magnitude improvement over existing constraints**
  - ◆ Precision tests of neutrino oscillations and  $U_{PMNS}$  unitarity (1%)

JUNO Relative Uncertainty vs. Leading Experiments

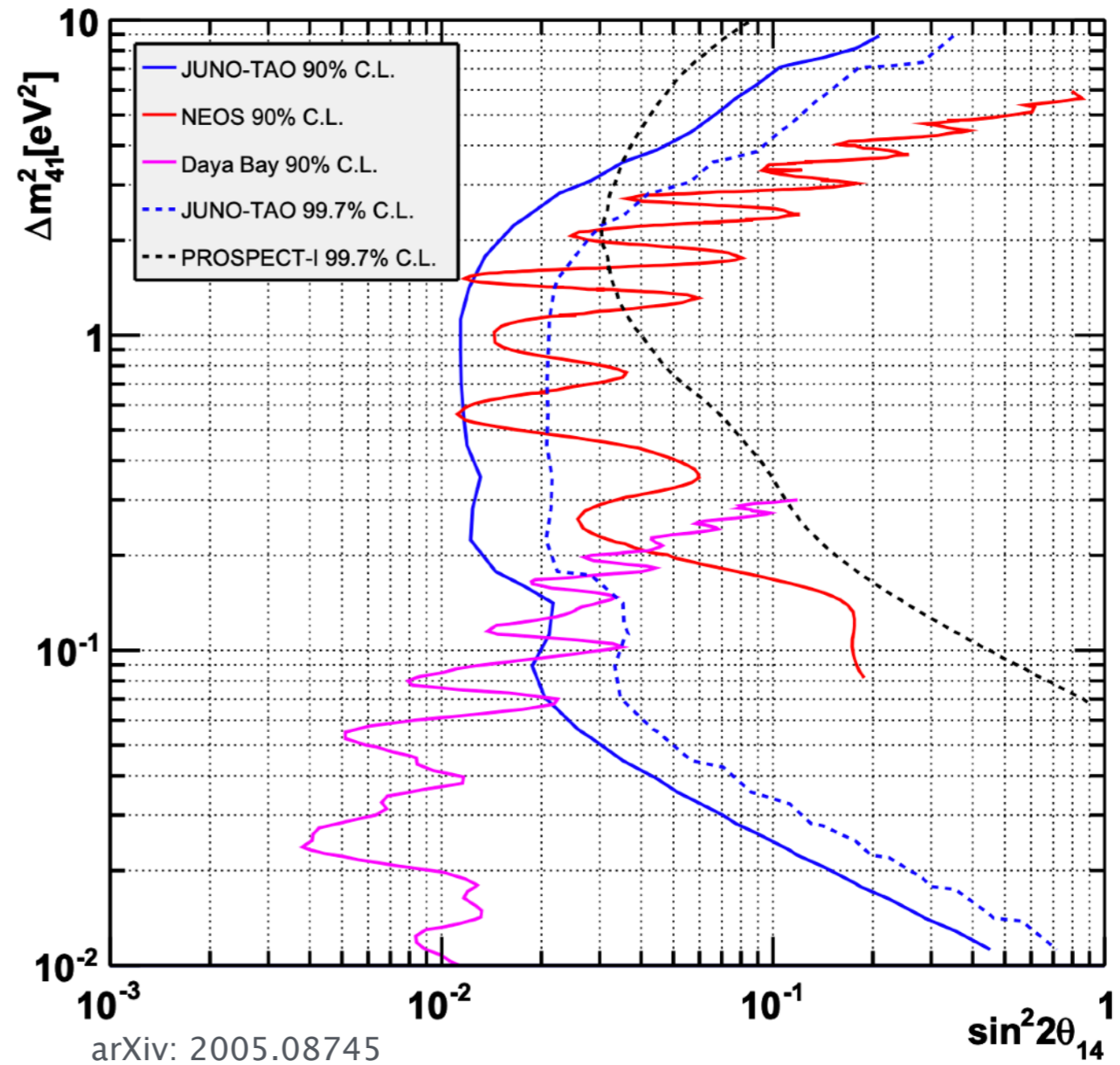
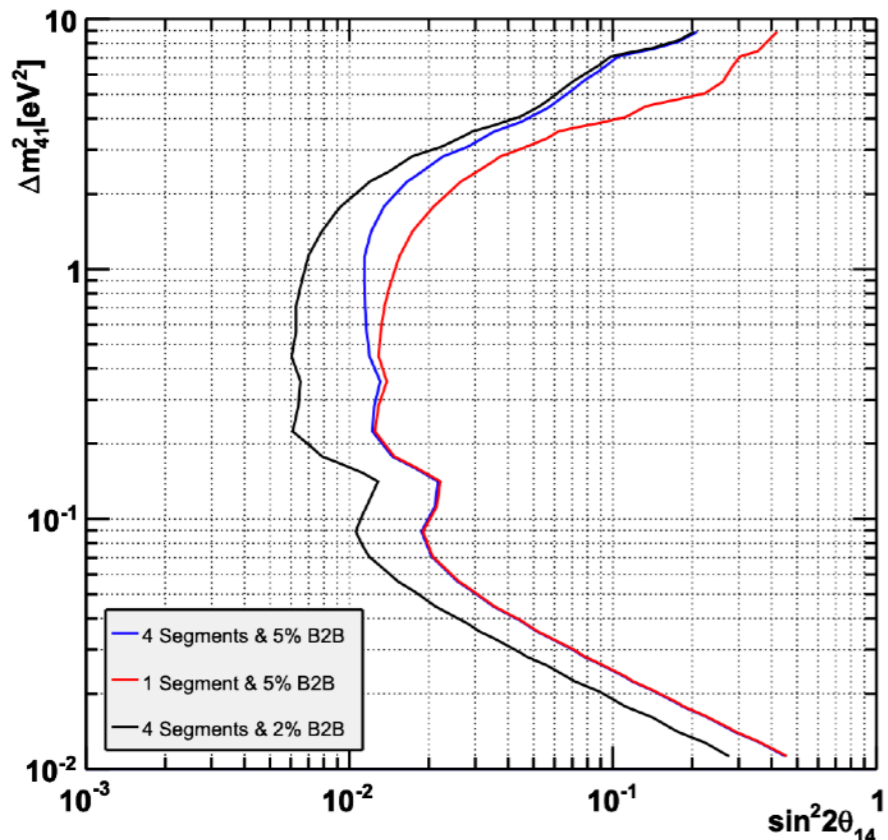
	$\Delta m_{21}^2$	$\Delta m_{31}^2$	$\sin^2\theta_{12}$	$\sin^2\theta_{13}$
<b>JUNO 100 days</b>	0.8%	1%	1.9%	47.9%
<b>JUNO 20 years</b>	0.1%	0.2%	0.3%	7.3%
<b>KamLAND</b>	2.4%	-	-	-
<b>T2K</b>	-	2.6%	-	-
<b>SNO+SK</b>	-	-	4.5%	-
<b>Daya Bay</b>	-	-	-	3.4%



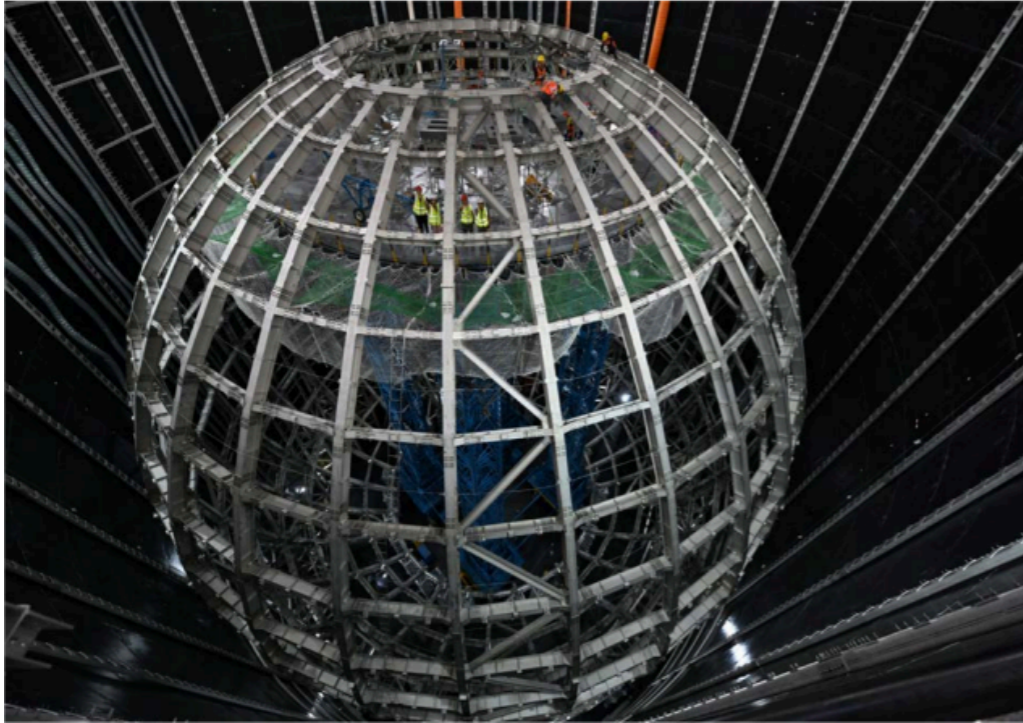
- Updated treatment of systematics
  - Values for 6 year exposure shown
- Rate systematics mitigated by spectral shape constraint on normalization



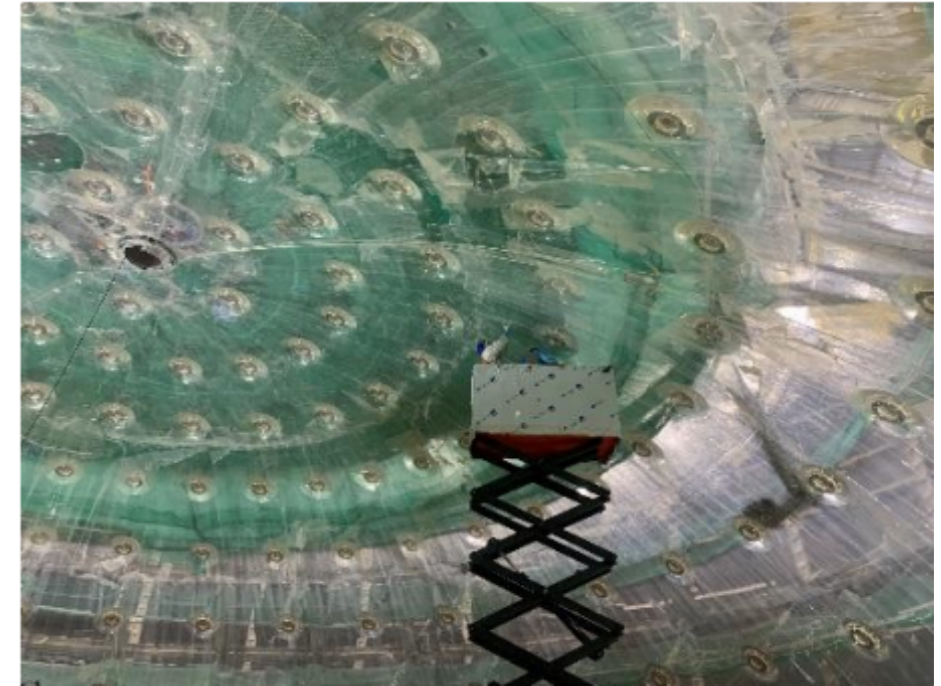
- With its short baseline, TAO has great potential in sterile oscillation searches
- Sensitivity improved by virtual segmentation of detector



Steel Support Structure finished



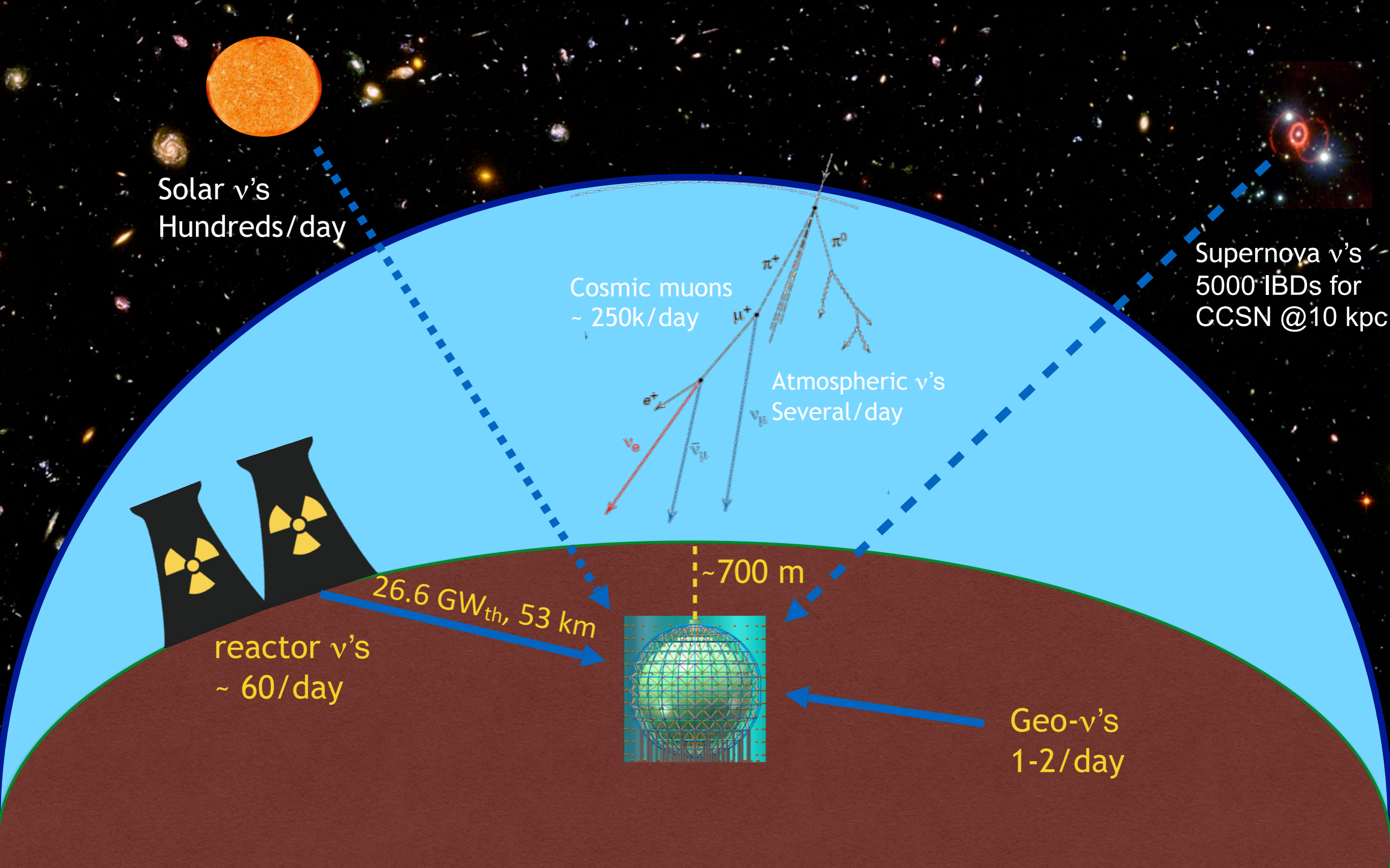
Top hemisphere of acrylic vessel finished



PMT and electronics installation ongoing!



- JUNO is a precision, multi-purpose experiment with a rich portfolio in neutrino, astroparticle, and new physics searches
- Using its reactor  $\bar{\nu}_e$  dataset:
  - NMO measurement to  $3\sigma$  in about 6 years
  - Sub-percent precision for  $\sin^2\theta_{12}$ ,  $\Delta m_{21}^2$  and  $\Delta m_{31}^2$  with as little as O(100) days of data
- Much progress has been done in construction and commissioning, with **data-taking starting next year!**
- Stay tuned for our exciting results!

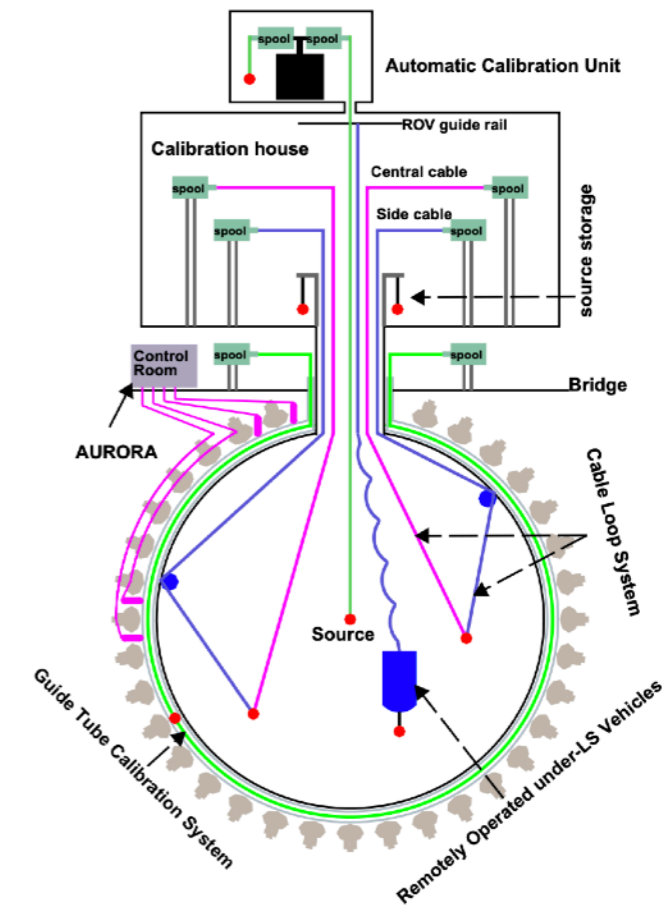




# BACKUP



- Comprehensive calibration strategy
  - ◆ Gamma/neutron sources, cosmogenic  $^{12}\text{B}$  and UV laser
  - ◆ Multi-positional source deployment
- SPMTs serve as linear reference for LPMT non-linearity
  - ◆ Operate in photon-counting mode for  $\sim 1\text{-}10$  MeV
- Dual Calorimetry Calibration compares LPMT charge to SPMT charge under same source
  - ◆ Channel-wise LPMT charge vs. total SPMT charge
  - ◆ UV laser energies span region of interest
  - ◆ Gamma sources match time profile of neutrino (positron) signal
  - ◆ Absolute energy scale uncertainty  $< 1\%$



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