CMB Bounds from Primordial Black Hole Accretion

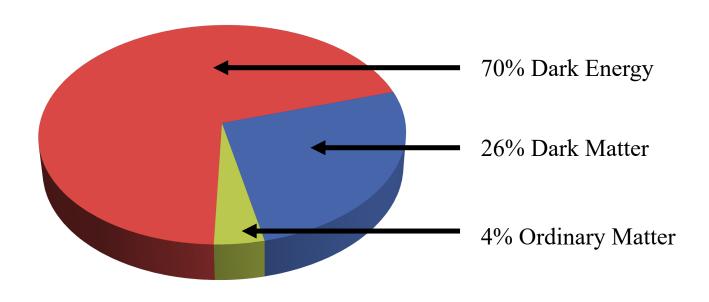
Greg Suczewski
Stony Brook University
May 9, 2023



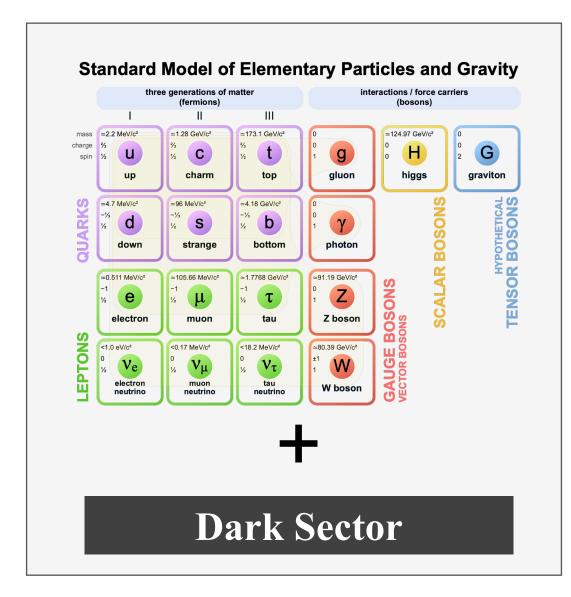
<u>OUTLINE</u>

- 1. Motivation and Background
- 2. PBH Accretion and Energy Deposition
- 3. Ionization and Thermal History
- 4. CMB Bounds and Preliminary Results

What is Dark Matter?



- Extensions of the Standard Model are most explored possibility
- NO DETECTION of DM particle candidates from experiments (yet)



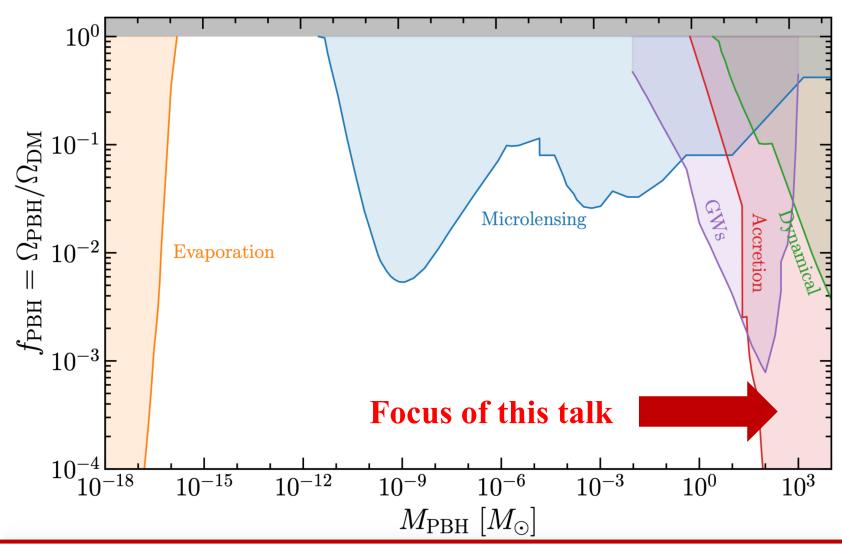
Motivation for PBHs

1. Primordial Black Holes (PBHs) are a macroscopic alternative to particle dark matter

2. PBHs might act as seeds for the formation of super-massive black holes

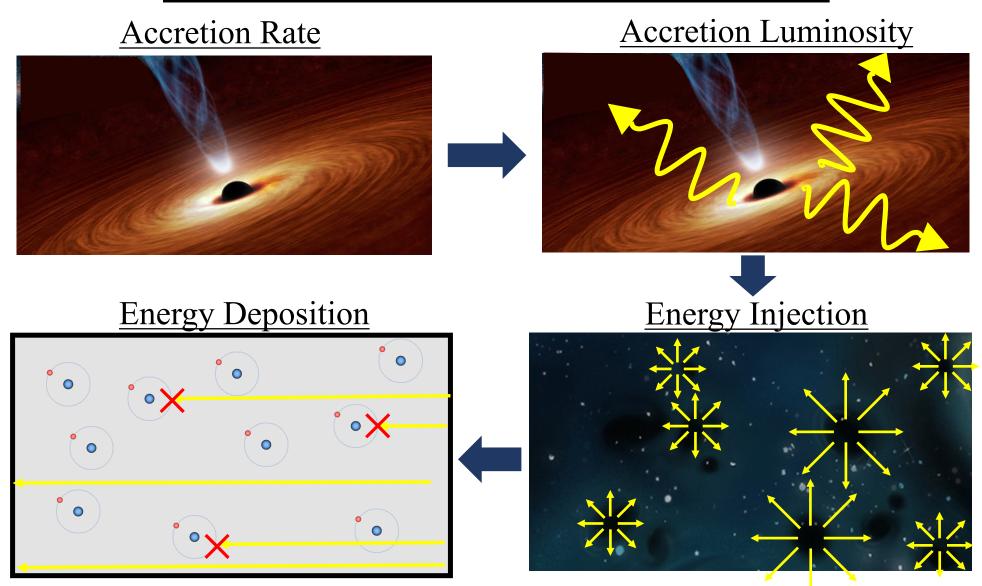
3. PBHs might explain recent gravitational wave detections of black hole mergers

PBH Abundance Constraints



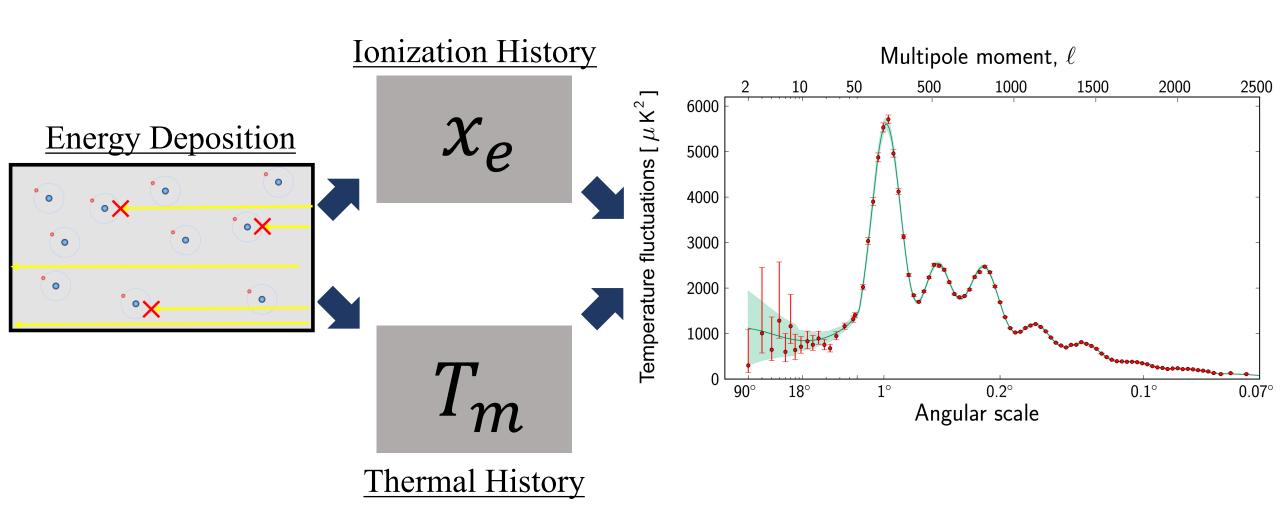
GOAL: TEST THE ROBUSTNESS OF THE ACCRETION BOUND AND APPLY NEW ACCRETION MODEL

Overview of PBH Accretion



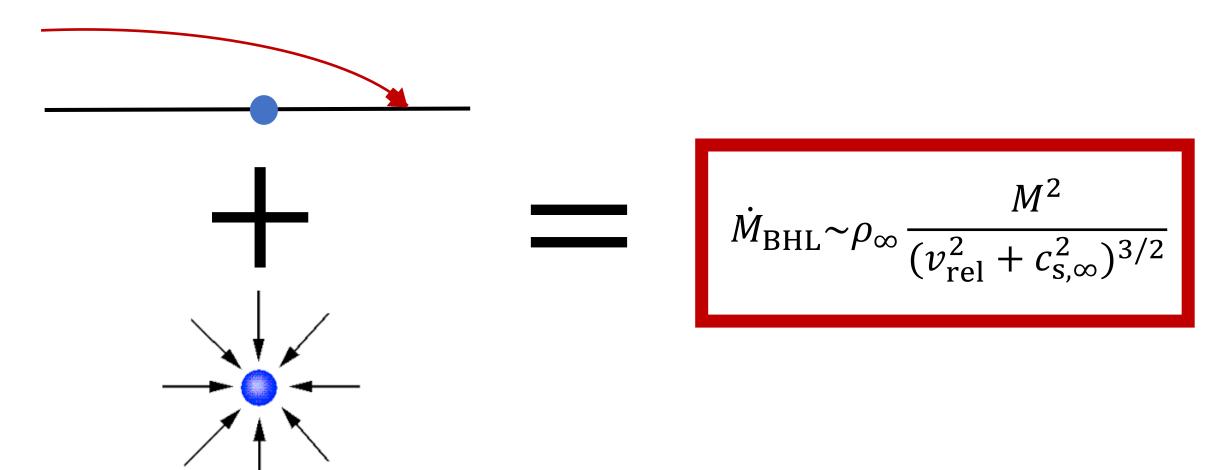
3 DEPOSITION CHANNELS: HEATING, IONIZATION, EXCITATION

Deriving the CMB Bound



Accretion Models (BHL)

- Bondi-Hoyle-Lyttleton (BHL) is the "standard" PBH accretion model
- Typically, BHL has been used to compute PBH abundance bounds



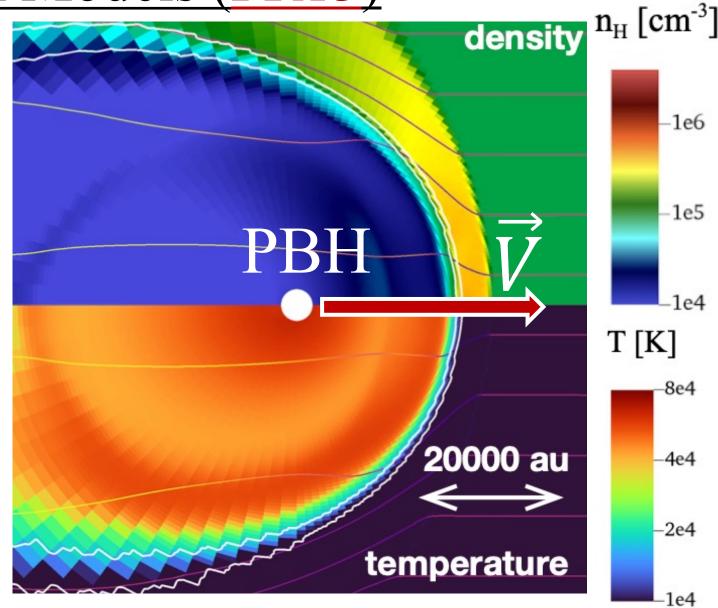
Accretion Models (PR13)

• PBHs form ionization fronts (Sugimura and Ricotti)

Park-Ricotti Model (PR13)

PR13 = Bondi + Ionization Front

$$\dot{M}_{\text{PR13}} \sim \rho_{\text{in}} \frac{M^2}{\left(v_{\text{in}}^2 + \left(c_{\text{s}}^{\text{in}}\right)^2\right)^{3/2}}$$

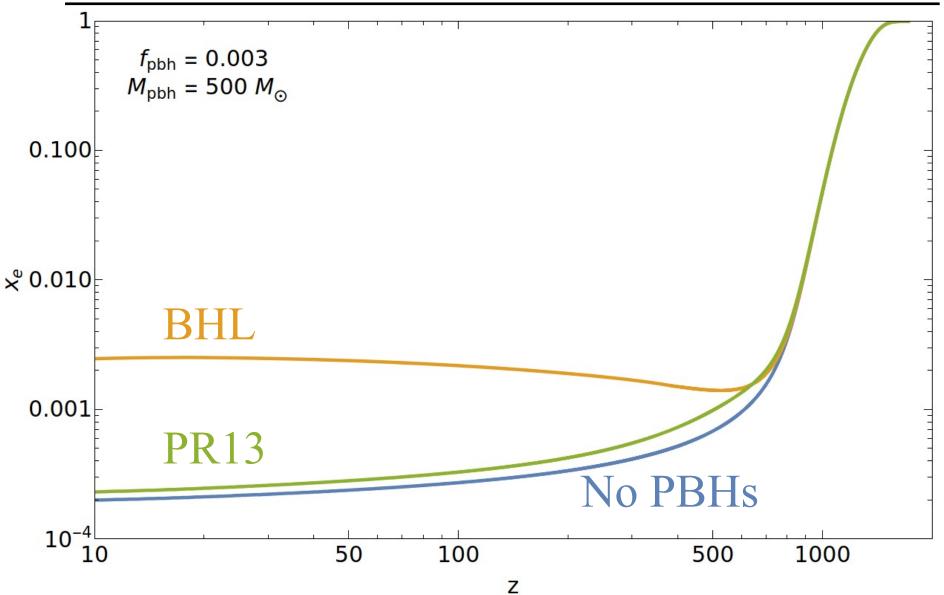


Important Assumptions

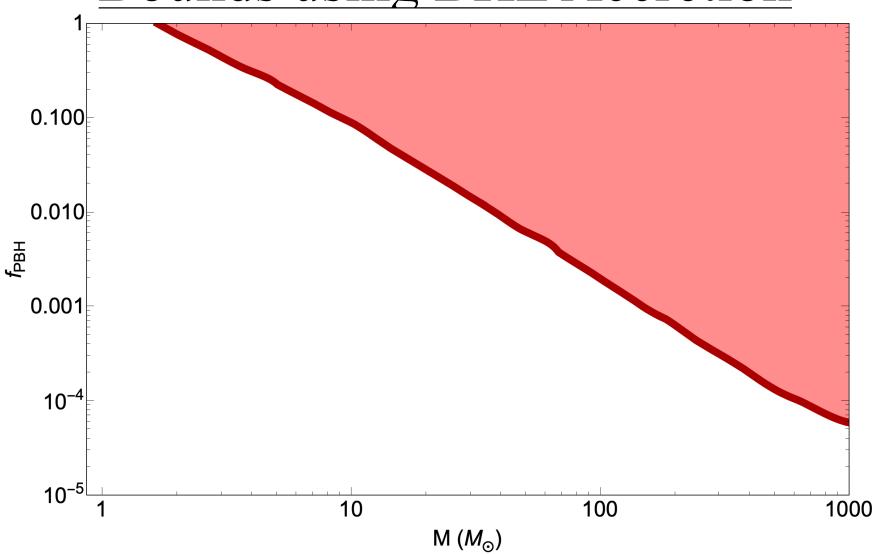
- Accretion Model
- Energy Deposition Functions
- Radiative Efficiency
 - Geometry of Accretion Flow
 - Advection
 - Outflows
 - Jets
 - Turbulence
- PBH Velocity Treatment (Linear Theory)
- Averaging Technique
- Speed of Sound in the Ionized Region (PR13)
- Photon Spectra

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BHL vs PR13 Free-Electron Fraction



Bounds using BHL Accretion

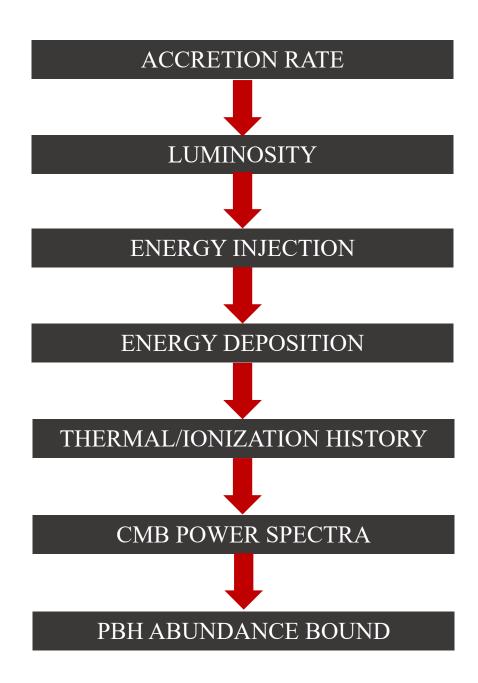


Computing the Bound

PROGRAMS ExoCLASS Dark Ages HyRec MontePython MAIN TASK Ionization and Energy **CMB** Power **MCMC** Deposition Thermal Spectra **Functions** History

Summary

- Computing PBH abundance bounds rely on many assumptions
- PR13 is suppressed compared to BHL
- Expect a **relaxation** in the bound (results coming soon!)



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