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Higgsed dark photons without isocurvatures

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Introduction

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 - $\lim_{k \rightarrow 0} \mathcal{P}_A \sim k^2$
- Light scalars populated via GPP have flat spectrum
- Higgsed $U(1)$'s with “light” Higgs cannot be fully populated via GPP

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- Stabilizes A and prevents kinetic mixing

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$$\frac{\rho_\sigma}{\rho_\gamma} \Big|_{\text{decay}} = \left(\frac{\sigma_0^2}{6 M_{\text{pl}}^2} \right)^{\frac{4}{3}} \left(\frac{M_\sigma}{\Gamma_\sigma} \right)^{\frac{2}{3}} \equiv Q$$

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- Curvaton also sets the curvature perturbations

$$\mathcal{P}_\zeta = \frac{4}{9} \frac{H_*^2}{4\pi^2 \sigma_0^2}, \mathcal{P}_T = \frac{2}{\pi^2} \frac{H_*^2}{M_{\text{pl}}^2}$$
$$r = \frac{\mathcal{P}_T}{\mathcal{P}_\zeta} = 18 \frac{\sigma_0^2}{M_{\text{pl}}^2}$$

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- Since we assume curvaton domination, relevant parameters:

$$\frac{\rho_\sigma}{s} \Big|_{T_R}, \mathcal{B} = \frac{\Gamma_{\text{DM}}}{\Gamma_{\text{SM}} + \Gamma_{\text{DM}}}, M_A, M_\sigma$$

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$$\frac{\rho_{\text{DM}}}{s} = \frac{2M_A}{M_\sigma} \mathcal{B} \frac{\rho_\sigma}{s} \Big|_{T_R} = \frac{2M_A}{M_\sigma} \frac{\mathcal{B}}{1 - \mathcal{B}} \frac{\rho_{\text{SM}}}{s} \Big|_{T_R} \approx \frac{2M_A}{M_\sigma} \frac{\mathcal{B}}{1 - \mathcal{B}} \frac{3}{4} T_R$$

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$$f_\sigma = \frac{\rho_\sigma}{\rho_\gamma + \frac{1}{3}\rho_\gamma + \rho_\sigma} = \frac{3\rho_\sigma}{4\rho_\gamma + 3\rho_\sigma} = \frac{3Q}{4 + 3Q}$$

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- If the curvaton dominates the energy density at decay, ($Q \equiv \frac{\rho_\sigma}{\rho_\gamma} \gg 1$),

$$\zeta = f_\sigma\zeta_\sigma + \mathcal{O}\left(\frac{1}{Q^2}\right)$$

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- Since DM is primarily produced via curvaton decay: $\zeta_{\text{DM}} \approx \zeta_\sigma$
$$\mathcal{S}_{\text{DM}} = 3(\zeta_{\text{DM}} - \zeta) \approx 3(\zeta_\sigma - f_\sigma \zeta_\sigma) = 3(1 - f_\sigma)\zeta_\sigma = \frac{3(1 - f_\sigma)}{f_\sigma} \zeta$$

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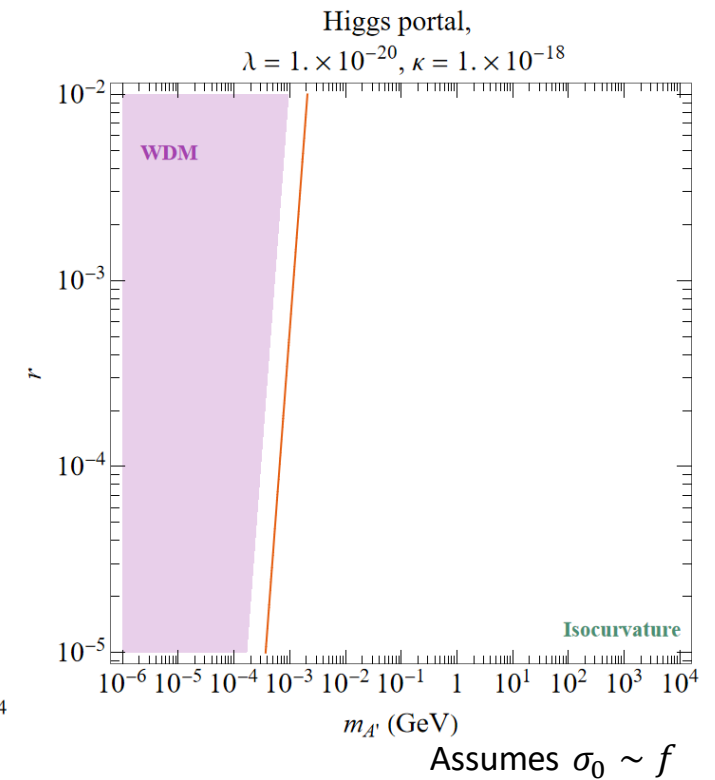
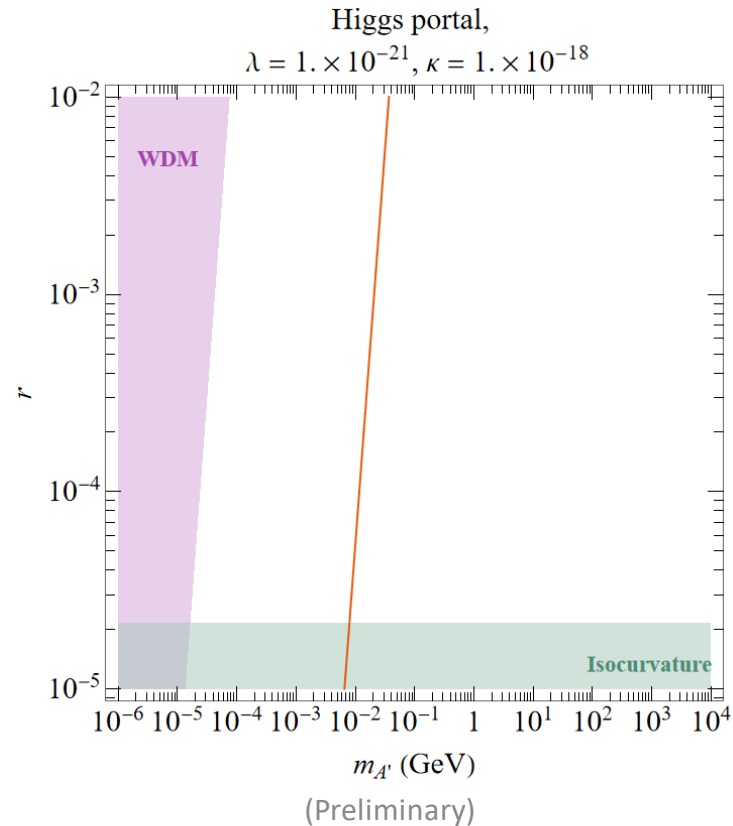
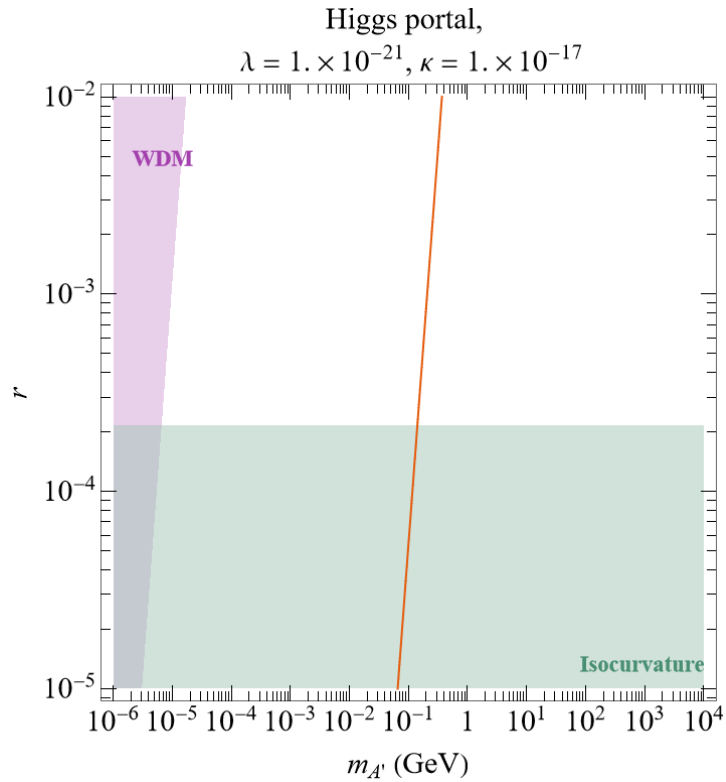
- For isocurvature fully correlated with adiabatic perturbations

$$\frac{\mathcal{S}_{\text{DM}}}{\zeta} = \frac{4}{Q} \approx 0.03$$

Example

- Consider a Higgs portal:

$$\mathcal{L}_{\text{portal}} = -\kappa|\Phi|^2|H|^2$$



Summary

- Higgsed $U(1)$'s with “light” Higgs cannot be populated via GPP without large isocurvatures
- By identifying the Higgs mode as a curvaton, dark sector can still be populated