Cosmological SGWB anisotropies

PRD 107 (2023), "New universal property of cosmological GW anisotropies" with E. Dimastrogiovanni, G. Domènech, M. Fasiello and G. Tasinato

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Third EuCAPT Symposium

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Direct observation with gravitational waves

Observation with light



SGWB Anisotropies

GW Production



Primordial source properties imprinted on anisotropies (Inflation, PT, PBH...)



Propagation through large scale density perturbations

$t \approx 0$

See review by LISA CosWG (2022)

GW Propagation

Detection



Characterising SGWB, parameter inference + model constraints

SGWB line-of-sight formalism

CMB-like approach to SGWB anisotropies

$$\underbrace{\Gamma(\eta_0, k, f, \hat{n})}_{\text{``}\Delta T/T" \text{ for GW}} = \Gamma_I + \Phi_I + \int_{\eta_i}^{\eta_0} d\eta \left\{ \Phi'(k, \eta) + \Psi'(k, \eta) \right\} e^{-i\hat{k}\cdot\hat{n}(\eta_0 - \eta)}$$

 $\Gamma_{I} \equiv \Gamma(\eta_{i}, k, f, \hat{n}) \rightarrow \text{initial perturbation}$ $\Phi_{I} \equiv \Phi(k, \eta_{i}) \rightarrow \text{SW}$ $\Phi'(k, \eta) + \Psi'(k, \eta) \rightarrow \text{ISW}$

[Alba & Maldacena 2015, Contaldi 2017; Bartolo et al. 2019a, 2019b]

Adiabatic initial conditions

SGWB produced during an epoch with equation of state w (e.g. kination, eMD..)

$$\Gamma_I = \frac{2\zeta}{5+3w}, \quad \Phi_I = -\frac{3(1+w)}{(5+3w)}$$

No effects in the anisotropies, adiabaticity $\Longrightarrow \zeta$ conserved on superhorizon scales

 10^{1}

 10^{2}

 10^{-10} -

 10^{3}

Isocurvature: curvaton mechanism

Scenario I

$$C_{\ell}^{\Gamma} \propto \left[-\frac{4}{3} \zeta_r \, j_{\ell} [k\eta_0] + \text{ISW} \right]^2$$

4x adiabatic term

Scenario II

independent curvaton fluctuations

- Anisotropies a key property of the SGWB
- Adiabatic anisotropies independent of primordial equation of state
- Distinctive predictions for isocurvature anisotropies based on curvaton mechanism

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