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# Jet quenching in evolving anisotropic matter

Andrey Sadofyev

IGFAE (USC)























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MARIE CURIE



AS, M. Sievert, I. Vitev, PRD, 2021 J. Barata, AS, C. Salgado, PRD, 2022 C. Andres, F. Dominguez, AS, CS, PRD, 2022 J. Barata, AS, X.-N. Wang, arxiv 2022 J. Barata, X. Mayo, AS, CS, 2022 (?)

# Jet tomography

- Jets see the matter in HIC at multiple scales, and essentially X-ray it;
- The existing jet quenching theory is based on multiple simplifying assumptions: large parton energy, static matter, no fluctuations, etc;
- There is a very recent progress on the medium motion and structure effects in jet quenching and this is the focus of this talk;
- The developed formalism can be also applied to include orbital motion of nucleons and some of the in-medium fluctuations to the energy loss in cold nuclear matter;









R. Baier et al, NPB, 1997 B. G. Zakharov, JETP, 1997 R. Baier et al, NPB, 1998 M. Gyulassy et al, NPB, 2000 M. Gyulassy et al, NPB, 2001











# **Color potential**



$$gA_{ext}^{\mu a}(q) = \sum_{i} e^{iq \cdot x_{i}} t_{i}^{a} u_{i}^{\mu} v_{i}(q) (2\pi) \,\delta\left(q^{0} - \vec{u_{i}} \cdot \vec{q}\right)$$

$$(1)$$

$$(2\pi) \,\delta\left(q^{0} - \vec{u_{i}} \cdot \vec{q}\right)$$





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# **Medium averaging**







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# Medium averaging



$$\int d^2 \mathbf{x}_n \, e^{-i(\mathbf{q}_n \pm \overline{\mathbf{q}}_n) \cdot \mathbf{x}_n} = (2\pi)^2 \, \delta^{(2)}(\mathbf{q}_n \pm \overline{\mathbf{q}}_n)$$

$$\uparrow$$

$$\rho \sim T^3$$





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# Medium averaging

















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$$i\mathcal{M}_n(p) = e^{-i\left(\mathbf{u}\cdot\mathbf{p} - \frac{p_\perp^2}{2E}\right)L} \int \frac{d^2\mathbf{p}_0}{(2\pi)^2} \mathcal{G}_n(\mathbf{p}, L; \mathbf{p}_0, 0) J(E - \mathbf{u}\cdot(\mathbf{p} - \mathbf{p}_0), \mathbf{p}_0),$$







$$\mathbf{\Omega} = -\frac{\mathbf{p} - \mathbf{q}}{E} + \frac{\mathbf{q}}{E} \frac{(\mathbf{p} - \mathbf{q})^2 - \mathbf{p}^2}{v(\mathbf{q}^2)} \frac{\partial v}{\partial \mathbf{q}^2}$$

still can be solved analytically







## **Broadening probability**

$$\frac{1}{d_{proj}} \left\langle \operatorname{Tr} \left[ \mathcal{G}^{\dagger}(\boldsymbol{p}_{0}^{\prime}, 0; \boldsymbol{p}_{L}) \, \mathcal{G}(\boldsymbol{p}, L; \boldsymbol{p}_{0}, 0) \right] \right\rangle \equiv (2\pi)^{2} \delta^{(2)}(\boldsymbol{p}_{0} - \boldsymbol{p}_{0}^{\prime}) \, \mathcal{P}(\boldsymbol{p}, L; \boldsymbol{p}_{0}, 0)$$

$$\frac{\partial}{\partial L}\mathcal{P}(\boldsymbol{p},L;\boldsymbol{p}_0,0) = -\int \frac{d^2\boldsymbol{q}}{(2\pi)^2}\,\sigma(\boldsymbol{p},\boldsymbol{q};L)\,\mathcal{P}(\boldsymbol{p}-\boldsymbol{q},L;\boldsymbol{p}_0,0)$$

$$\mathcal{P}(\boldsymbol{p}, 0; \boldsymbol{p}_0, 0) = (2\pi)^2 \delta^{(2)}(\boldsymbol{p} - \boldsymbol{p}_0)$$









$$E\frac{d\mathcal{N}}{d^2\boldsymbol{p}\,dE} = \int \frac{d^2\boldsymbol{p}_0}{(2\pi)^2}\,\mathcal{P}(\boldsymbol{p},L;\boldsymbol{p}_0,0)\left[1-\boldsymbol{u}\cdot(\boldsymbol{p}-\boldsymbol{p}_0)\frac{\partial}{\partial E}\right]E\frac{d\mathcal{N}^{(0)}}{d^2\boldsymbol{p}_0\,dE}$$

- The odd moments are proportional to the transverse flow velocity, the even moments are unmodified;
- The initial and final distributions are not factorized anymore in coordinate space (due to the energy derivative);







#### Jet broadening uniform matter $E \frac{d\mathcal{N}^{(0)}}{d^2\mathbf{p}_0 dE} = f(E)\delta^{(2)}(\mathbf{p}_0)$

Eikonal approximation --  $E \rightarrow \infty$ 

$$\left\langle p_{\perp}^{2k}\boldsymbol{p}\right\rangle = \int \frac{d^2\boldsymbol{p}\,d^2\boldsymbol{r}}{(2\pi)^2}\,p_{\perp}^{2k}\boldsymbol{p}\,e^{-i\boldsymbol{p}\cdot\boldsymbol{r}}e^{-\mathcal{V}(\boldsymbol{r})L} = 0 + \mathcal{O}\left(\frac{\perp}{E}\right)$$

Opacity expansion --  $\chi \equiv C \frac{g^4 \rho}{4\pi \mu^2} L \ll 1$ 

$$\left\langle p_{\perp}^{2k} \boldsymbol{p} \right\rangle \simeq -\frac{\boldsymbol{u}}{2E} \mathcal{C}\rho L \int \frac{d^2 \boldsymbol{p}}{(2\pi)^2} p_{\perp}^{2k+2} \left[ E \frac{f'(E)}{f(E)} v(p_{\perp})^2 + p_{\perp}^2 \frac{\partial v^2}{\partial p_{\perp}^2} \right]$$







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$$\langle \mathbf{p} \rangle \simeq 3 \, \chi \, \mathbf{u} \, \frac{\mu^2}{E} \log \frac{E}{\mu}$$



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# The propagator



$$G(\mathbf{x}_L, L; \mathbf{x}_0, 0) = \int_{\mathbf{x}_0}^{\mathbf{x}_L} \mathcal{D}\mathbf{r} \exp\left(\frac{iE}{2} \int_0^L d\tau \, \dot{\mathbf{r}}^2\right) \mathcal{P} \exp\left(-i \int_0^L d\tau \, t_{\text{proj}}^a v^a(\mathbf{r}(\tau), \tau)\right)$$







$$\left\langle \mathcal{P} \exp\left(-i \int_{0}^{L} d\tau t_{\text{proj}}^{a} v^{a}(\boldsymbol{r}(\tau), \tau)\right) \mathcal{P} \exp\left(i \int_{0}^{L} d\overline{\tau} t_{\text{proj}}^{b} v^{b}(\overline{\boldsymbol{r}}(\overline{\tau}), \overline{\tau})\right) \right\rangle$$
$$= \exp\left\{-\int_{0}^{L} d\tau \left[1 + \frac{\boldsymbol{r}(\tau) + \overline{\boldsymbol{r}}(\tau)}{2} \cdot \hat{\boldsymbol{g}}\right] \mathcal{V}(\boldsymbol{r}(\tau) - \overline{\boldsymbol{r}}(\tau))\right\}$$
$$\stackrel{\uparrow}{\boldsymbol{g}} = \left(\nabla \rho \frac{\delta}{\delta \rho} + \nabla \mu^{2} \frac{\delta}{\delta \mu^{2}}\right)$$

$$\left\langle G(\boldsymbol{x}_{L}, L; \boldsymbol{x}_{0}, 0) G^{\dagger}(\overline{\boldsymbol{x}}_{L}, L; \overline{\boldsymbol{x}}_{0}, 0) \right\rangle$$
  
= 
$$\int_{0}^{\boldsymbol{u}_{L}} \mathcal{D}\boldsymbol{u} \int_{0}^{\boldsymbol{w}_{L}} \mathcal{D}\boldsymbol{w} \exp\left\{ \int_{0}^{L} d\tau \left[ iE \, \dot{\boldsymbol{u}} \cdot \dot{\boldsymbol{w}} - (1 + \boldsymbol{w} \cdot \hat{\boldsymbol{g}}) \, \mathcal{V}(\boldsymbol{u}(\tau)) \right] \right\}$$









$$E\frac{d\mathcal{N}^{(0)}}{d^2\mathbf{p}_0 dE} = \frac{1}{2(2\pi)^3} |J(p_0)|^2 = f(E)\delta^{(2)}(\mathbf{p}_0)$$

inhomogeneous matter

$$\frac{d\mathcal{N}}{d^{2}\boldsymbol{x}dE} \simeq \exp\left\{-\mathcal{V}\left(\boldsymbol{x}\right)L\right\} \left\{ \begin{bmatrix} 1 - \frac{iL^{3}}{6E}\boldsymbol{\nabla}\mathcal{V}\left(\boldsymbol{x}\right) \cdot \hat{\boldsymbol{g}}\,\mathcal{V}\left(\boldsymbol{x}\right) \end{bmatrix} \frac{d\mathcal{N}^{(0)}}{d^{2}\boldsymbol{x}dE} + \frac{iL^{2}}{2E}\,\hat{\boldsymbol{g}}\,\mathcal{V}\left(\boldsymbol{x}\right) \cdot \boldsymbol{\nabla}\frac{d\mathcal{N}^{(0)}}{d^{2}\boldsymbol{x}dE} \right\} \\ \hat{\boldsymbol{g}} \equiv \left(\boldsymbol{\nabla}\rho\frac{\delta}{\delta\rho} + \boldsymbol{\nabla}\mu^{2}\frac{\delta}{\delta\mu^{2}}\right) \\ \rho \sim T^{3}$$

$$\langle \mathbf{p} \, p_{\perp}^2 \rangle \simeq \chi^2 \, \frac{L \nabla T}{2T} \, \frac{\mu^4}{E} \left( \log \frac{E}{\mu} \right)^2$$









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$$\left\langle \mathbf{p} \, p_{\perp}^2 \right\rangle \simeq \chi^2 \, \frac{L \nabla T}{2T} \, \frac{\mu^4}{E} \left( \log \frac{E}{\mu} \right)^2$$





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uniform matter

- Opacity  $\chi \approx 4$
- $u \approx 0.7$  (about  $\pi/4$  to z-axis)
- $\mu = gT$  with  $g \approx 2$  and  $T \approx 500 MeV$

$$\left\langle \frac{p_{\perp}}{E} \right\rangle \simeq 3 \chi \, \frac{u_{\perp}}{1 - u_z} \frac{\mu^2}{E^2} \log \frac{E}{\mu}$$

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What jet energy corresponds to \langle \theta \rangle \approx 1^o?

Jet energies: E < 50 \text{ GeV}
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inhomogeneous matter

- Opacity  $\chi \approx 4$
- $\mu = gT$  with  $g \approx 2$  and  $T \approx 500 MeV$
- $L\nabla T > T$

$$\left\langle \frac{\mathbf{p}}{E} \frac{p_{\perp}^2}{\mu^2} \right\rangle \simeq \chi^2 \frac{L\nabla T}{2T} \frac{\mu^2}{E^2} \left( \log \frac{E}{\mu} \right)^2$$

What jet energy corresponds to 
$$\langle "\theta" \rangle \approx 1^o$$
?  
  
Jet energies:  $E < 100 \text{ GeV}$ 











$$\left\langle \frac{\mathbf{k}}{k_{\perp}^2} \right\rangle \simeq \frac{N_c \chi}{C_F \, x E} \mathbf{u}$$









# Summary

- Jets do feel the transverse flow and anisotropy, and get bended;
- The transverse flow and anisotropy are expected to affect the medium-induced radiation, bending the substructure of jets;
- These effects can be in principle probed in experiment, leading us towards actual jet tomography;
- The initial and final state effects are not factorized anymore;
- One may also expect similar evolution-induced effects for other probes of nuclear matter, e.g. quarkonium;





