Study of Self-similar Solution of Self-gravitating Non-relativistic Fluids

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Motivation

- The properties and existence of the dark matter is one of the most fascinating question in cosmology.
- Our main goal of this research is to find scaling solutions of the gravitational fields, which can be good candidates to describe the evolution of the Universe or collapse of compact astrophysical objects
- We present a dark fluid model described as a non-relativistic and self-gravitating fluid
- We studied these coupled **non-linear differential equation** systems using self-similar time-dependent solutions



The Model

We consider a set of coupled non-linear differential equations, which describes the **non-relativistic dynamics** of a compressible fluid with zero thermal conductivity and zero viscosity.

$$\partial_t \rho + \nabla(\rho u) = 0$$

 $\partial_t u + (u \nabla) u = -\frac{1}{\rho} \nabla p + g$
 $p = p(\rho)$

Equation of State (Polytropic and Chaplygin gas)

$$p = -w\rho^n$$
 $p = -\frac{A}{\rho^n}$

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Sedov-Taylor Ansatz

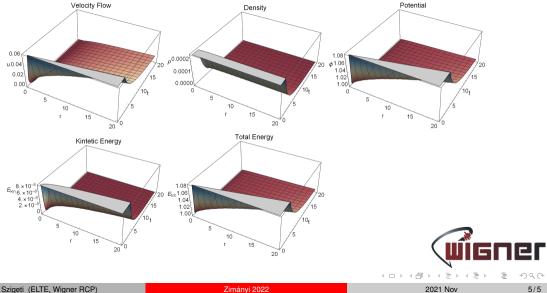
• We are reducing the PDE system into ODE system using Taylor ansatz:

$$\begin{split} u(r,t) &= t^{-\alpha} f\left(\frac{r}{t^{\beta}}\right) \quad \rho(r,t) = t^{-\gamma} g\left(\frac{r}{t^{\beta}}\right) \\ \Phi(r,t) &= t^{-\delta} h\left(\frac{r}{t^{\beta}}\right), \end{split}$$

- (f, g, h) shape-functions only depend on $\zeta = rt^{-\beta}$
- $\alpha, \beta, \gamma, \delta$ similarity exponents
- The β describes the rate of spread of the spatial distribution
- Other exponents describe the rate of decay of the intensity of the corresponding field

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Result:



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