

# Study of Self-similar Solution of Self-gravitating Non-relativistic Fluids

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# Motivation

- The properties and existence of the dark matter is one of the most fascinating question in cosmology.
- Our main goal of this research is to find **scaling solutions** of the gravitational fields, which can be good candidates to describe the evolution of the Universe or collapse of compact astrophysical objects
- We present a **dark fluid model** described as a non-relativistic and self-gravitating fluid
- We studied these coupled **non-linear differential equation** systems using self-similar time-dependent solutions



# The Model

We consider a set of coupled non-linear differential equations, which describes the **non-relativistic dynamics** of a compressible fluid with zero thermal conductivity and zero viscosity.

$$\partial_t \rho + \nabla(\rho \mathbf{u}) = 0$$

$$\partial_t \mathbf{u} + (\mathbf{u} \nabla) \mathbf{u} = -\frac{1}{\rho} \nabla p + \mathbf{g}$$

$$p = p(\rho)$$

Equation of State (Polytropic and Chaplygin gas)

$$p = -w\rho^n \qquad p = -\frac{A}{\rho^\alpha}$$



# Sedov-Taylor Ansatz

- We are reducing the PDE system into ODE system using Taylor *ansatz*:

$$u(r,t) = t^{-\alpha} f\left(\frac{r}{t^\beta}\right) \quad \rho(r,t) = t^{-\gamma} g\left(\frac{r}{t^\beta}\right)$$
$$\Phi(r,t) = t^{-\delta} h\left(\frac{r}{t^\beta}\right),$$

- $(f, g, h)$  **shape-functions** only depend on  $\zeta = rt^{-\beta}$
- $\alpha, \beta, \gamma, \delta$  similarity exponents
- The  $\beta$  describes **the rate of spread** of the spatial distribution
- Other exponents describe the **rate of decay** of the intensity of the corresponding field



# Result:

