

CAST and after-CAST

Konstantin Zioutas

University of Patras / Greece

Begining: October 1999

End: 2024

With Frank Avignone, Georg Raffelt + Juan Collar we defended the

CAST proposal to CERN. 24 years later >> praise CAST >>> !!!

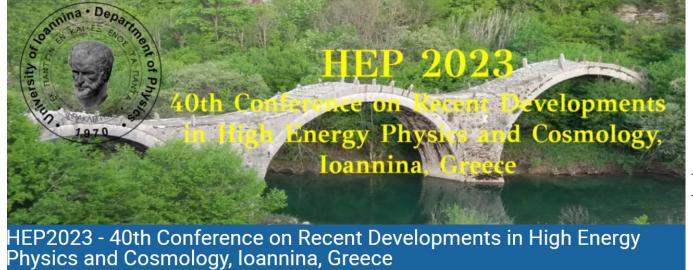
HEP conference ~1997 magnets in the LHC epoch

→ CAST conception ←

Dieter Hoffmann

First -CHF → GSI / Darmstadt
First CHF → FFET / GR

>> Thanking greek representatives



Ioannina, 8th April 2023

CAST conception - approval:

Annual Greek HEP conference ~1997:

LHC large scale magnets!

Via Buenos Aires / Argentina SOLAX collaboration meeting

• 1st suggestion to professor Frank AVIGNONE! -> YES

then proposal to CERN DG - show 1 MCHF - zero

1-2 years later

 nomore zero MCHF

 approved

Alvaro de Rujula



< 1998

AXION SEARCHES

are

- MANDATORY
- O FUN, CREATIVE

O PROCEEDING



https://doi.org/10.109

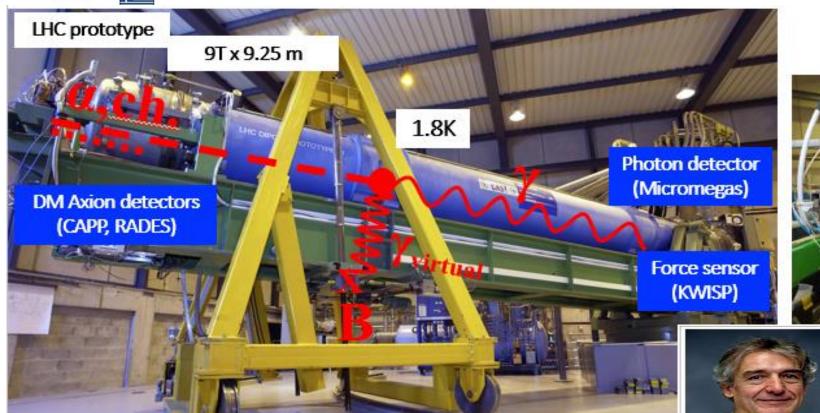
Axion Helioscope

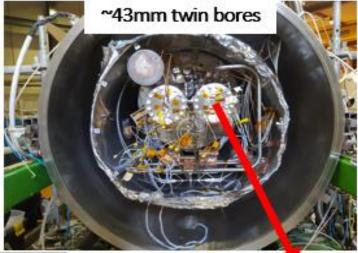
Solar Axions



Axion Haloscope

DM Axions





CAST-CAPP

← Our Father....1983πάτερ ημών...

P. Sikivie

With CAST

From solar axions

- H.E. solar Axions
- Solar Chameleons
- DM axions

unexpected obs's, e.g. planetary relationships

SPINOFFs

Beyond initial proposal

Also:

XRTelescope from DE astrophysics → astropart. Phys.

Phase matched 4 cavities → ∞

~fast scanning ~10 MHz / min (single cavity)

CAST: physics achievements 2NATURE papers

1PRL &



- young people 'start-ups', promotions
- education in astrophysics & cosmology ...

ILIAS >> EU network

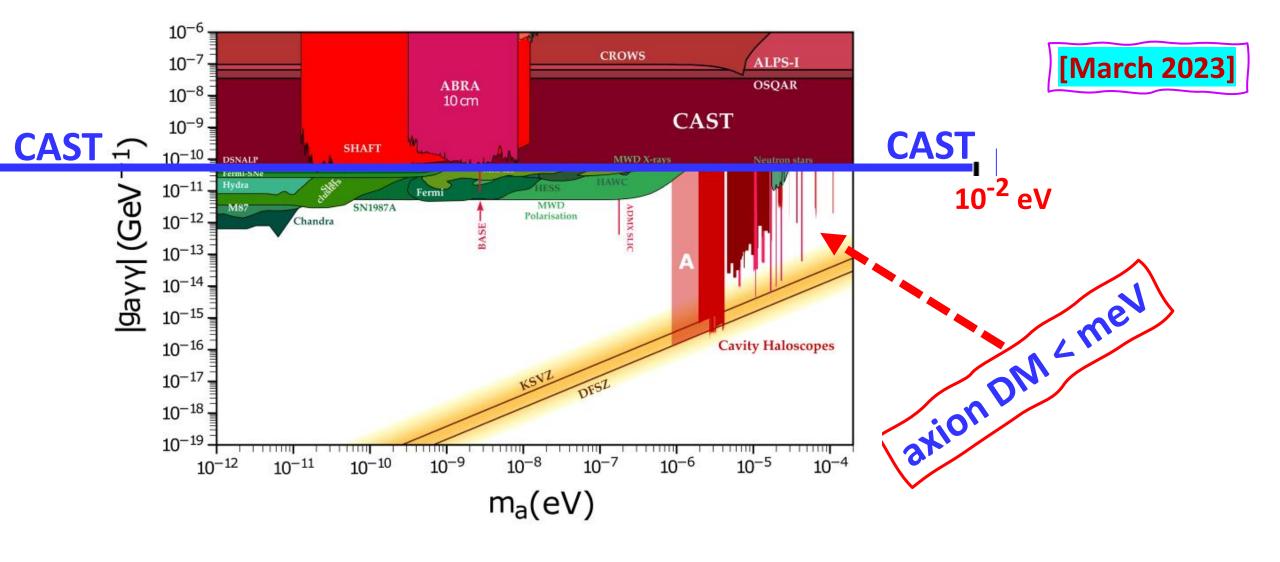


FIG. 1: Limits on the axion electromagnetic coupling $g_{a\gamma\gamma} \equiv g_{\gamma} \frac{\alpha}{\pi} \frac{1}{f_a}$ as a function of axion mass m_a , obtained by various axion dark matter searches. In addition, the light shaded area labeled A indicates the sensitivity, under assumptions spelled out in Section VI, of a search using a reentrant or dielectrically loaded cavity inserted in the "Extended Frequency Range" (EFR) magnet that ADMX plans to operate at Fermilab. The figure was made by modifying Ciaran O'Hare's code

PHYSICAL REVIEW LETTERS

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First Results from the CERN Axion Solar Telescope

K. Zioutas et al. (CAST Collaboration) Phys. Rev. Lett. **94**, 121301 – Published 1 April 2005 2005

Article

References

Citing Articles (276)

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ABSTRACT

Hypothetical axionlike particles with a two-photon interaction would be produced in the sun by the Primakoff process. In a laboratory magnetic field ("axion helioscope"), they would be transformed into x-rays with energies of a few keV. Using a decommissioned Large Hadron Collider test magnet, the CERN Axion Solar Telescope ran for about 6 months during 2003. The first results from the analysis of these data are presented here. No signal above background was observed, implying an upper limit to the axion-photon coupling $g_{a\gamma} < 1.16 \times 10^{-10}~{\rm GeV}^{-1}$ at 95% C.L. for $m_a \lesssim 0.02~{\rm eV}$. This limit, assumption-free, is comparable to the limit from stellar energy-loss arguments and considerably more restrictive than any previous experiment over a broad range of axion masses.

2017



OPEN

New CAST limit on the axion-photon interaction

CAST Collaboration[†]

V. Anastassopoulos¹, S. Aune², K. Barth³, A. Belov⁴, H. Bräuninger⁵, G. Cantatore⁶, J. M. Carmona⁷, J. F. Castel⁷, S. A. Cetin⁸, F. Christensen⁹, J. I. Collar¹⁰, T. Dafni⁷, M. Davenport³, T. A. Decker¹¹, A. Dermenev⁴, K. Desch¹², C. Eleftheriadis¹³, G. Fanourakis¹⁴, E. Ferrer-Ribas², H. Fischer¹⁵, J. A. García⁷†, A. Gardikiotis¹, J. G. Garza⁷, E. N. Gazis¹⁶, T. Geralis¹⁴, I. Giomataris², S. Gninenko⁴, C. J. Hailey¹⁷, M. D. Hasinoff¹⁸, D. H. H. Hoffmann¹⁹, F. J. Iguaz⁷, I. G. Irastorza⁷*, A. Jakobsen⁹, J. Jacoby²⁰, K. Jakovčić²¹, J. Kaminski¹², M. Karuza^{6,22†}, N. Kralj^{22†}, M. Krčmar²¹, S. Kostoglou³, Ch. Krieger¹², B. Lakić²¹, J. M. Laurent³, A. Ljubičić²¹, G. Luzón⁷, M. Maroudas¹, L. Miceli²³, S. Neff¹⁹, I. Ortega^{3,7}, T. Papaevangelou², K. Paraschou¹³, M. J. Pivovaroff¹¹, G. Raffelt²⁴, M. Rosu^{19†}, J. Ruz¹¹, E. Ruiz Chóliz⁷, I. Savvidis¹³, S. Schmidt¹²,

Y. K. Semertzidis^{23†}, S. K. Solanki^{25†}, L. Stewart³, T. Vafeiadis³, J. K. Vogel¹¹, S. C. Yildiz^{8†}, K. Zioutas^{1,3}

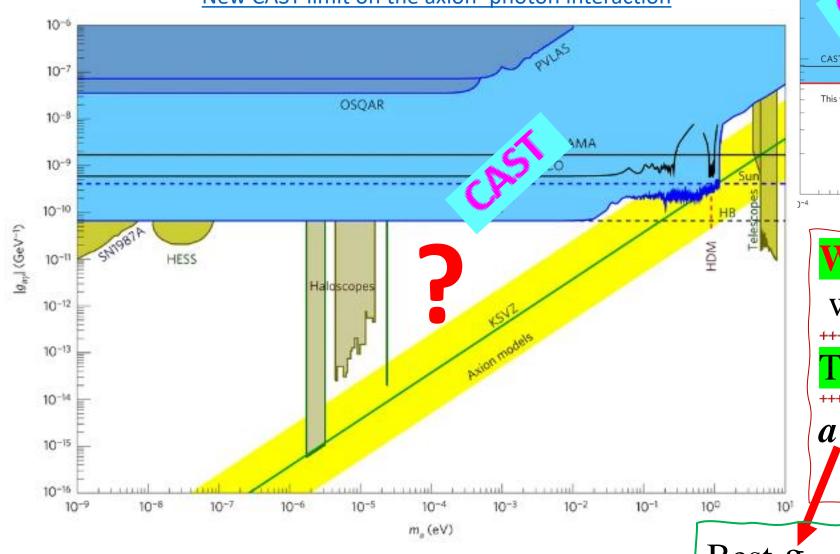
60 authors 25% greeks

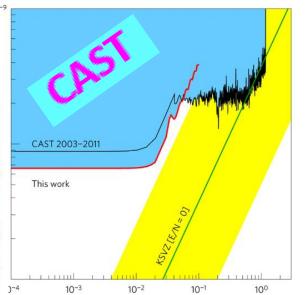
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2017

Constraints on the two-photon coupling $g_{a\gamma\gamma}$ of axions and similar particles depending on their mass m_a .

New CAST limit on the axion—photon interaction





we look(ed) into the Sun?

The arguments:

a priori ≠ *a posteriori* 1999 2022

Best g_{ayy} - limit

This talk

PHYSICAL REVIEW D **85**, 035018 (2012) => Exper. Publication => 2022

Prospects for searching axionlike particle dark matter with dipole, toroidal, and wiggler magnets

Oliver K. Baker, Michael Betz, Fritz Caspers, Joerg Jaeckel, Axel Lindner, Andreas Ringwald, Yannis Semertzidis, Pierre Sikivie, and Konstantin Zioutas Authors Authors Authors Authors Department of Physics, Yale University, New Haven, Connecticut 06520-8120, United States, USA

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University of Patras, Patras, Greece

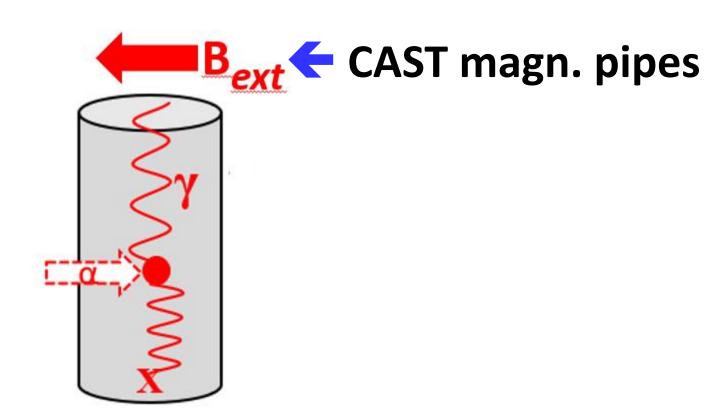
In this work, we consider searches for dark matter made of axions or axionlike particles using resonant radio frequency cavities inserted into dipole magnets from particle accelerators, wiggler magnets developed for accelerator based advanced light sources, and toroidal magnets similar to those used in particle-physics detectors. We investigate the expected sensitivity of such axionlike-particle dark-matter detectors and discuss the engineering aspects of building and tuning them. Brief mention is also made of even stronger field magnets which are becoming available due to improvements in magnetic technology. It is concluded that new experiments utilizing already-existing magnets could greatly enlarge the mass region in searches for axionlike dark matter particles.

(Received 10 November 2011; published 17 February 2012)

DOI: 10.1103/PhysRevD.85.035018 PACS numbers: 14.80.Va, 95.35.+d



→ AXION HALOSCOPE à la Sikivie



nature communications

ි 2022

Article

https://doi.org/10.1038/s41467-022-33913-6

Search for Dark Matter Axions with CAST-CAPP

64 authors 17% greeks

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Check for updates

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M. Vretenar @ 20,30, S. Youn @ 13 & K. Zioutas @

13

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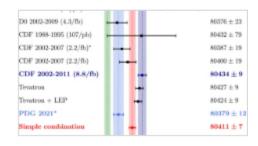
7 Feb 2023

Nature Communications

<u>Hadronic uncertainties versus new physics for the W boson mass and Muon g-2 anomalies</u>

The tension between measured W mass and its Standard Model prediction might arise from uncertainties in the hadronic contribution, and the same is true for the muon g - 2. Here, the authors show that such a common origin for the two anomalies is unlikely, while a model involving leptoquarks might explain them both.

Peter Athron, Andrew Fowlie ... Bin Zhu



Article

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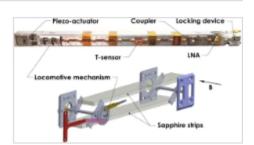
19 Oct 2022

Nature Communications

Search for Dark Matter Axions with CAST-CAPP

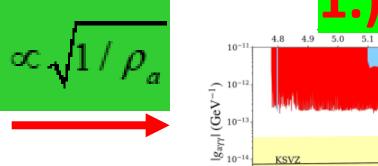
Haloscopes aim at detecting axions by converting them into photons using high-quality resonant cavities, where the cavity resonance should be tuned with the unknown axion mass. Here, the authors improve exclusion limits using four phase-matched resonant cavities and a fast frequency scanning technique.

C. M. Adair, K. Altenmüller ... K. Zioutas



nature communications => October 2022 => https://rdcu.be/cXQ71





2.)

- Phase-matching of all four cavities.
- Fast resonance scanning.
- Unexplored parameter space.

Fig. 5 | CAST-CAPP exclusion limit on the axion-photon coupling as a function of axion mass at 90% confidence level (left), and compared to other axion search results 10,25,10,14-41 within the mass range 1-25 µeV (right). The higher

5.2 5.3

DM streams / clusters -> improve sensitivity

Astroparticle Physics

Volume 20, Issue 1, October 2003, Pages 73-78



Gravitational lensing



C

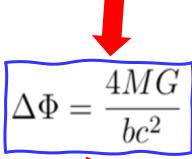
0.001c

Gravitational lensing by the Sun of nonrelativistic penetrating particles

D.H.H Hoffmann ^a ♀ ☒, J Jacoby ^b ☒, K Zioutas ^{a c d} ☒



a "first"





All you need is ..



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https://doi.org/10.1016/S0927-6505(03)00138-5 >

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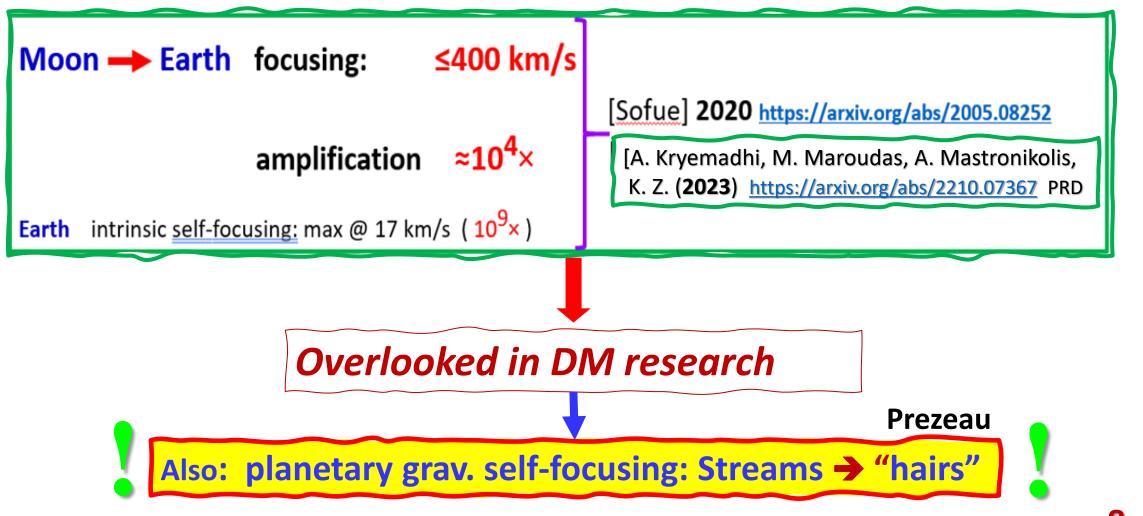
Abstract

Show more V

The flux of weakly interacting particles from celestial sources, moving with a velocity v≈0.2c, can be temporarily amplified at the site of the Earth, due to gravitational lensing effects by the Sun. The effective amplification factor can be as much as $\sim 10^3$ to $\sim 10^4$, for a velocity bin-width of ~0.1%. The theoretically motivated solar Kaluza-Klein axions provide a generic example of particles with a wide velocity spectrum, filling the gap between $v \approx c$ (e.g. neutrinos) and $v \approx 10^{-3} c$ (e.g. dark matter (DM) candidates). If the putative particles come from a direction along the projected path of the Sun in the Sky, within a strip of ~0.1° along the ecliptic, then, time windows of possible enhanced flows can be predicted. This suggestion can be implemented in the (re)-analysis of data from DM-experiments, and, it does not need any major experimental modification. In particular, performing a cross-correlation of data taken over a period more than 1 year, from the same or even also from other experiments, this can result to (un)predictable time windows of interest. Because, if burst-like events re-appear in following years in fixed dates, this will be an unambiguous identification of the cosmic origin of underground events, which were ignored before. Thus, thanks to solar gravitational effects, DM-experiments can be transformed to telescopes of penetrating non-relativistic particles with a field-of-view of ~0.1°, or even more, along the ecliptic.

The missing access to DM-data does not allow us to test this technique. We therefore suggest to the astroparticle physics community to release its data.





Modern Physics Letters A Vol. 29, No. 37 (2014) 1440008 (7 pages) © World Scientific Publishing Company DOI: 10.1142/S0217732314400082





The 11 years solar cycle as the manifestation of the dark Universe

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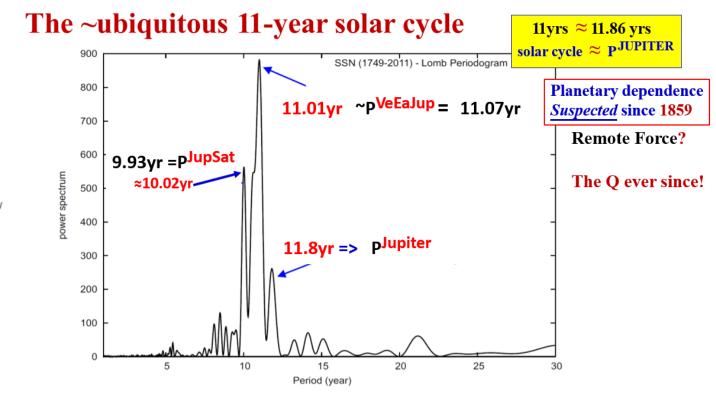
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Received 29 January 2014 Accepted 18 September 2014 Published 26 November 2014

Sun's luminosity in the visible changes at the 10^{-3} level, following the 11 years period. This variation increases with energy, and in X-rays, which should not even be there, the amplitude varies up to $\sim 10^5$ times stronger, making their mysterious origin since the discovery in 1938 even more puzzling, and inspiring. We suggest that the multifaceted mysterious solar cycle is due to some kind of dark matter streams hitting the Sun. Planetary gravitational lensing enhances (occasionally) slow moving flows of dark constituents toward the Sun, giving rise to the periodic behavior. Jupiter provides the driving oscillatory force, though its 11.8 years orbital period appears slightly decreased, just as 11 years, if the lensing impact of other planets is included. Then, the 11 years solar clock may help to decipher (overlooked) signatures from the dark sector in laboratory

experiments or observations in space. Keywords: 11 years solar cycle; dark matter; gravitational lensing.



Discarded... 1967

MERCURY

because inconsistent w' 1/R5 tidal force

http://adsabs.harvard.edu/abs/1967AJ.....72..463B AJ (1967)

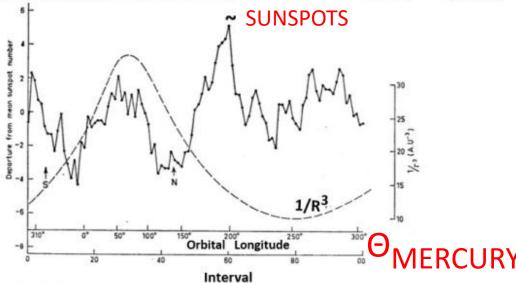
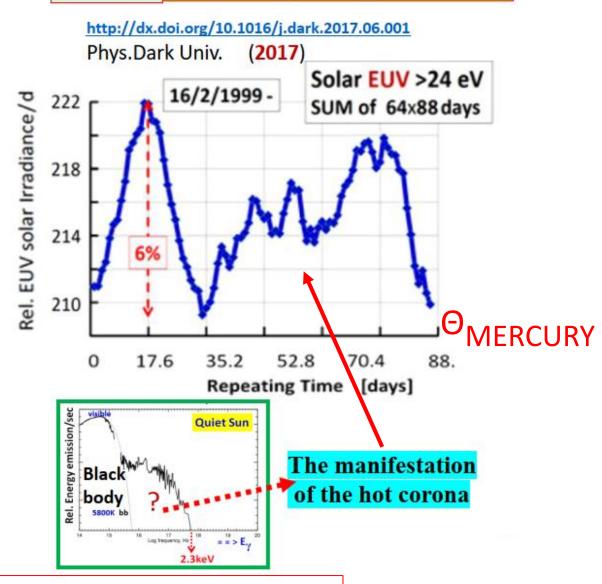


Fig. 4. Departures from mean sunspot numbers as a function of Mercury's position. Equivalent to the mean waveform of the detected signal.

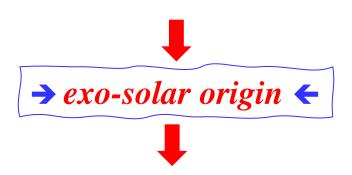
"It is immediately obvious that no simple theory will entirely account for this complex pattern, but one of the maxima occurs near Mercury's closest approach to the sun and the two conspicuous minima occur quite close to the planet's greatest departures from the plane of the earth's orbit (N and S in Fig.)."





WHY solar axions?

- Axion theory
- Sun a permanent big bang~16MK
- From the Sun ~grav. focusing



Best SPINOFF

DM: .001c - Grav. Focusing within solar system => STREAMS D. Hoffmann, J. Jacoby, K.Z, Astroparti. Phys. 20 (2003) 73;

https://www.sciencedirect.com/science/article/pii/S0927650503001385

Sun's dynamical behaviour

=> a multiple mystery: e.g. corona heating & its

T – inversion 1939 – , 11 years cycle, sunspots, flares,

EUV, solar radius variation, ... F10.7 radio line (=solar proxy)

every solar observable

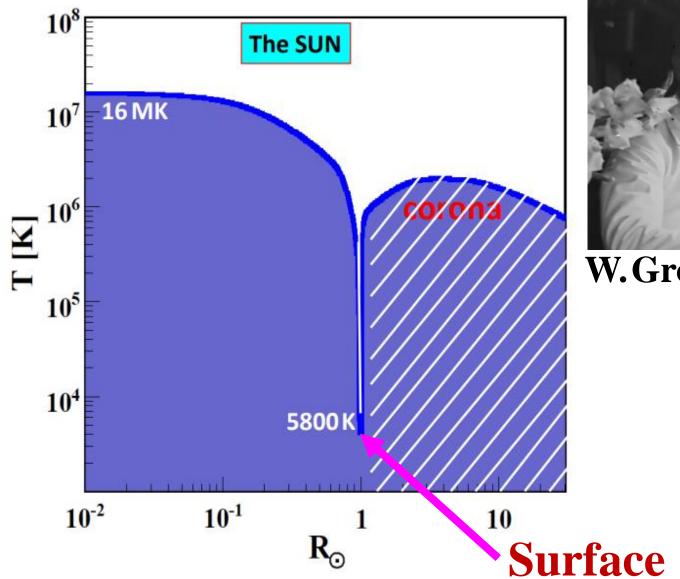


?Best choice?

Frank Wilczek (CENR seminar):

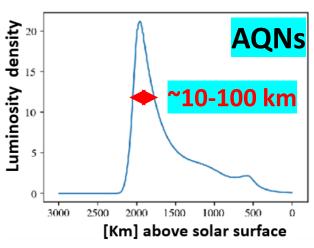
"~focus on anomalies / mysteries"

A mysterious solar observation: $T_{SUN} = f(R)$





W. Grotrian 1939



N. Raza, L. van Waerbeke, **A. Zhitnitsky**, *Solar Corona Heating by the AQN DM*, arXiv:1805.01897 (2018)

The AQN model is being tested by CAST

"Spacecraft Makes <u>Progress</u> on **Solar Heating Mystery"**

October 2022 • Physics 15, 157 Parker Solar Probe config

a long-suspected heat source for the Sun's ... Experts are not sure why the solar and the solar wind are hundreds of times hotter than the surface of the Sun, but they have several theories. Researchers have now confirmed one suspected source of heating.

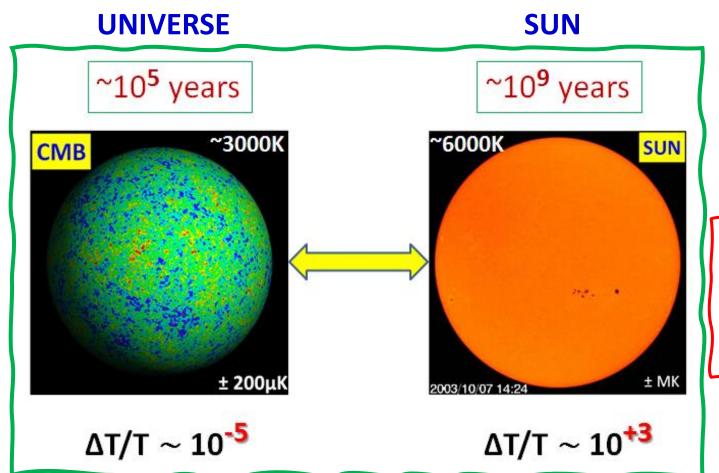
surprisingly hot corona, but there may be others.

https://physics.aps.org/articles/v15/157

It is extremely difficult to simulate the details of coronal heating ... Thus

testing realistic models with observations is a major challenge.

JA. Klimchuk et al., ApJ (2022) accepted 2nd Nov. 2022



SUN:

not thermalized after 4.5 Gyrs

=> WHY_NOT?

look at!

Observation & cosmology invisible streams @ solar system

The dynamical solar behaviour remains a mystery! + inspiring

11 years cycle, Corona anomaly, sunspots, flares, EUV, solar radius variation ... F10.7 radio line (=solar proxy) →

All show planetary dependencies! HOW?

a posteriori CONCLUSION: The Sun is a set of anomalies / mysteries \Rightarrow Look at e.g.: w. an XRT or, w. DM equipment Unexpected.....potential discovery!

Fine-Grained Streams from Cosmological Simulations

The cold nature of dark matter yields particles with nearly zero dispersions (~10⁻¹⁰ c for WIMPs and 10⁻¹⁷ c for Axions) at the last scattering

Mark Vogelsberger and Simon D. M. White used N-body equations of motion:

The DM distribution at a typical point in the halo is described as a superposition of many *fine-grained streams* with discrete velocity distributions, each of which has a very small velocity dispersion.

Mark Vogelsberger and Simon D. M. White, Mon. Not. R. Astron. Soc. 413 (**2011**) 1419

arXiv:2108.11647 [pdf]

The Dark Universe is not invisible

K. Zioutas, V. Anastassopoulos, A. Argiriou, G. Cantatore, S.A. Cetin, A. Gardikiotis, D.H.H. Hoffmann, S. Hofmann, M. Karuza, A. Kryemadhi, M. Maroudas, E.L. Matteson, K. Ozbozduman, T. Papaevangelou, M. Perryman, Y.K. Semertzidis, I. Tsagris,

M. Tsagri, G. Tsiledakis, D. Utz, E.L. Valachovic Phys. Sci. Forum 2021, 2(1), 10

Febr. **2021**

Trillions of DM particles may lurk in Earth's crust

https://doi-org.ezproxy.cern.ch/10.1016/S0262-4079(22)01808-5

Oct.

2022

Dunkle Materie weniger dunkel?

Translated:

Is DM less dark?

https://www.scinexx.de/news/kosmos/dunkle-materie-wer-enigdunkel/

Febr. **2023**

Seminar at CERN 28/3/2023, M. Hostert.

Semi-Visible Dark Photons

https://arxiv.org/abs/2302.05410

Febr. **2023**

The Dark World is not dark!

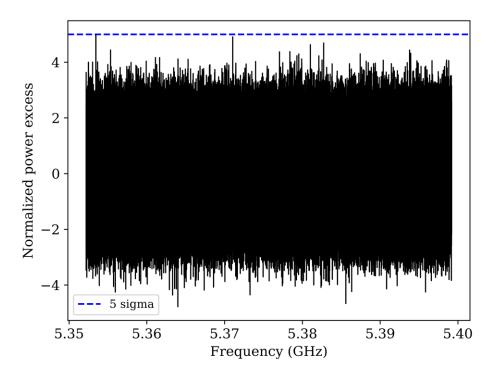


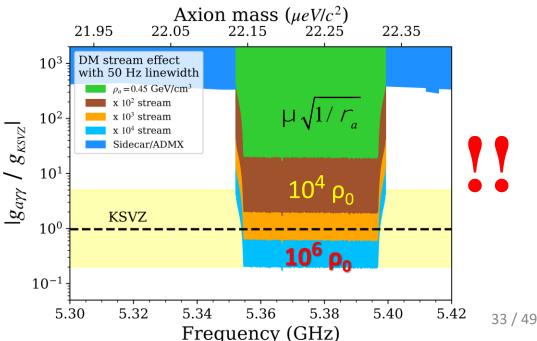
Additional

slides

STREAMING DARK MATTER

- New procedure for transient events such as axion streams or axion-mini clusters.
- Second measurement channel is decisive for such searches.
- Example measurement for 4.5h on 24/11/2020 (19:19 23:53 local time).
- Fast scanning method over 42 MHz (5.3547 GHz 5.3967 GHz).
- \triangleright Combined spectrum (RBW = 50Hz) shows no lines above 5 σ threshold.
- \triangleright "Local" exclusion plot assuming axion streams with modest flux enhancement $10^2 10^4$.
- > Future tests with even smaller RBW including also raw IQ data in time-domain.





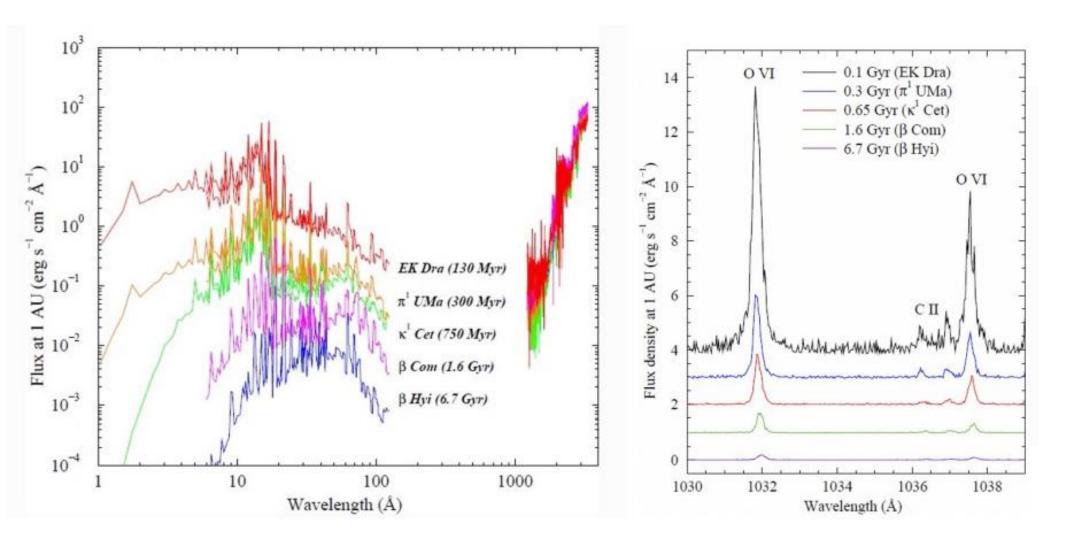
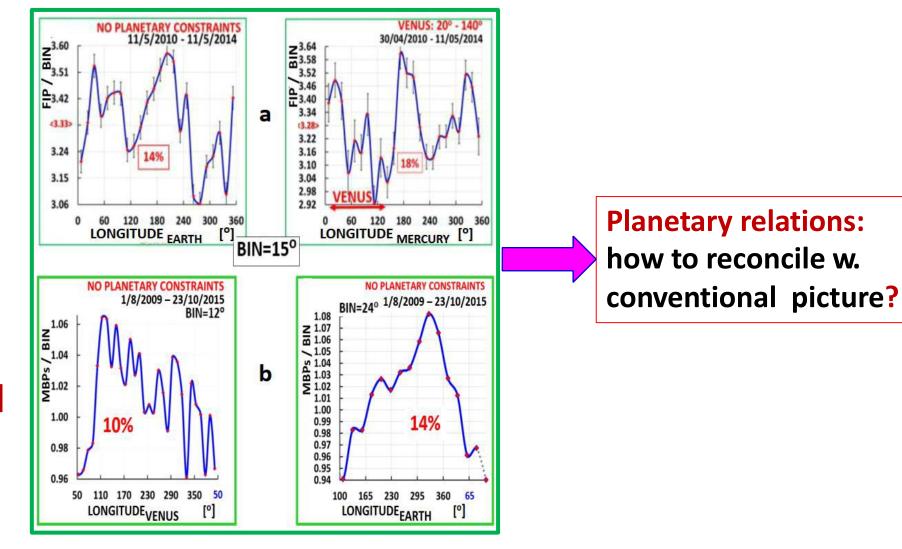


Fig. 9. Levels of emission from Sun-like stars are shown as a function of age on the main sequence. Flux densities at the stellar surface can be derived by multiplying by $(1\text{AU}/R_{\odot} = 215)^2 = 4.6 \cdot 10^4$. Note the logarithmic and linear scales plotted, and the gap between 120 and 1000 Å caused by interstellar absorption. Figures from [22].

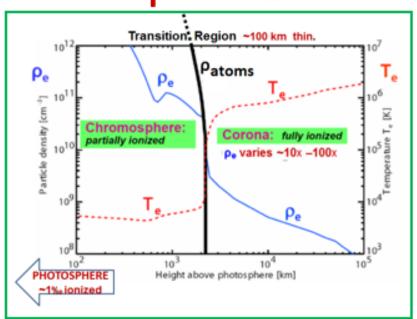
https://arxiv.org/abs/2205.03498

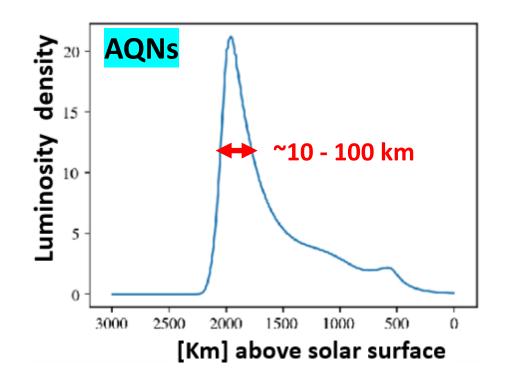
Elemental Composition



- (a) https://www.nature.com/articles/s41467-017-00328-7 NATURE Comm. 2017
- (b) https://arxiv.org/abs/1710.01678 PASJ 2017

Chromosphere ⇔ Corona





AQNs

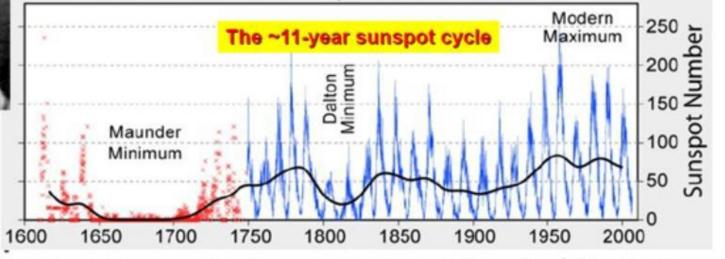
- the <u>only</u> solar atmospheric model reproducing the ~100 km thin Transition Region
- Planetary dependence of the flaring Sun
- more? >>> unexplained obs'?!

The theory of a planetary modulation of solar activity -

Extract of a Letter from Prof. R. Wolf, of Zurich, to Mr. Carrington, dated Jan. 12, 1859.

(Translation.)





the same planets, the conclusion seems to be inevitable, that my conjecture that the variations of spot-frequency depend on the influences of *Venus*, *Earth*, *Jupiter*, and *Saturn*, will not prove to be wholly unfounded. The preponderating planet

Too weak within known physics

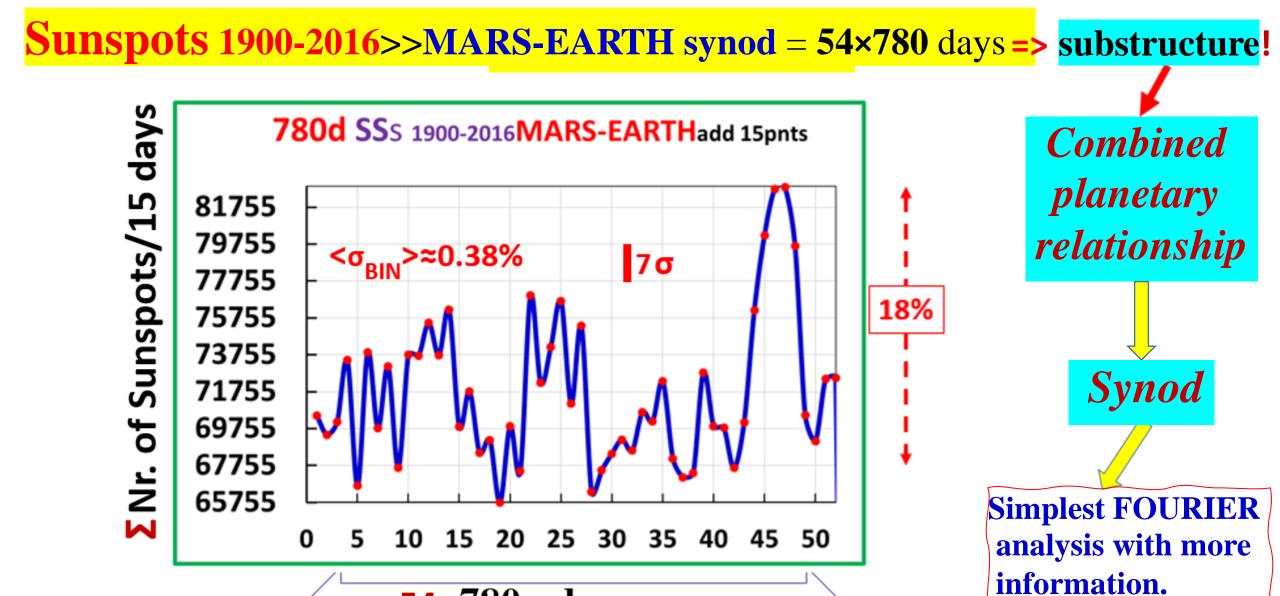
therefore abandoned!

Overlooked!

The variations of sunspot Periodicity depend oninfluences Of Venus, Earth, Jupiter & Saturn

http://people.duke.edu/~ns2002/pdf/Scafetta_AGU-2012.pdf Scafetta, N.: 2012a, J. Atm. Sol. Terr. Phys., 80, 296

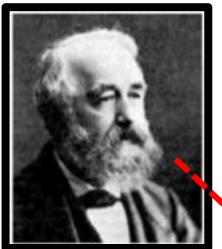
44



days

54×780

R. WOLF

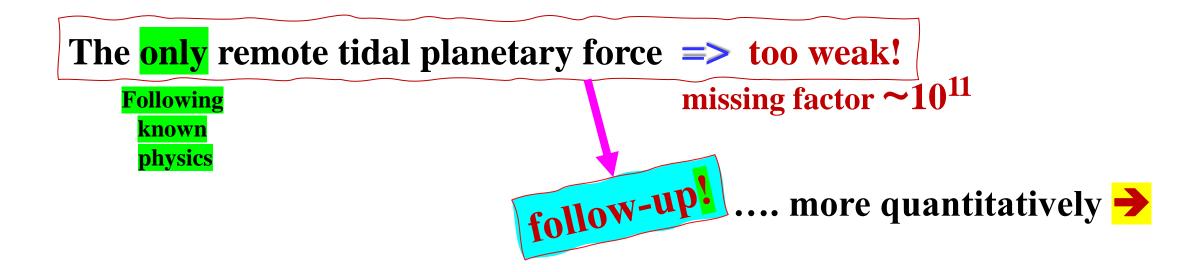


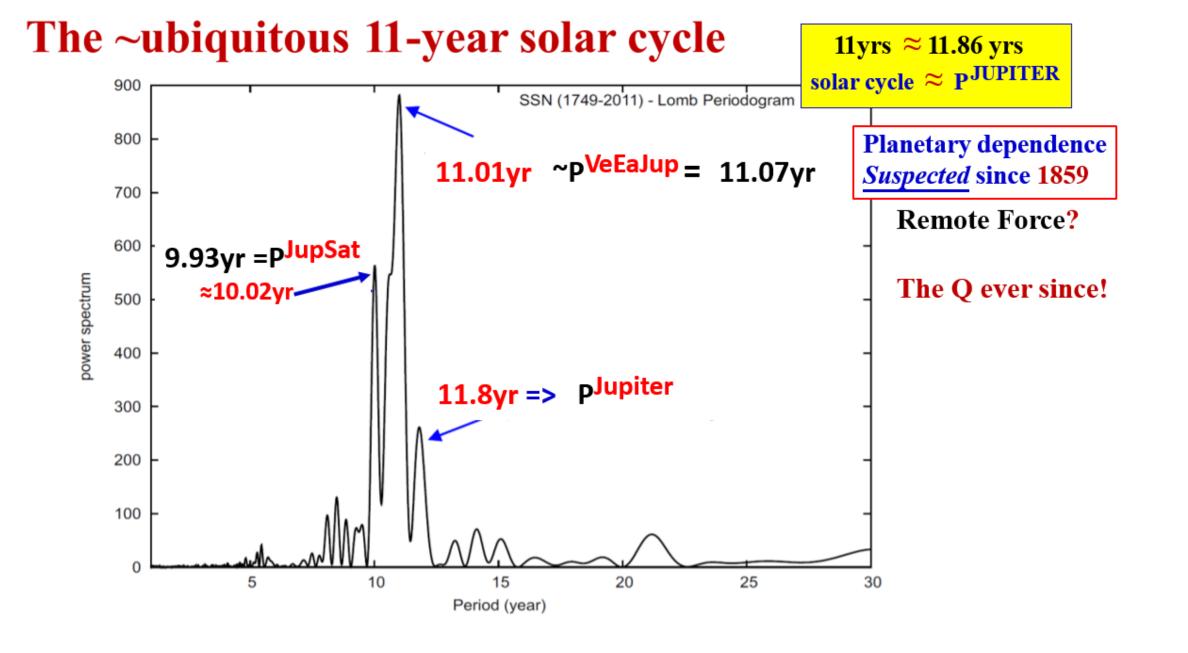
The variations of sunspots depend on Venus, Earth, Jupiter & Saturn

The 11-year solar cycle

PJUPITER = 11.86 years ≈ 11 years

1859- Suspected planetary dependence





Solar Corona 1939- >>> observational mystery.

Sun's upper atmosphere much hotter than its surface => why?

" a major open issue in astrophysics"

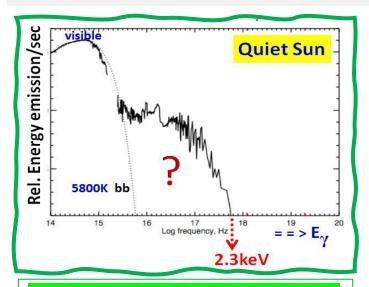
2015

"one of the fundamental outstanding problems in solar physics"

2015

"for 84 years...one of the outstanding unsolved problems in astrophysics" 2015

[http://arxiv.org/abs/1502.07401; http://arxiv.org/abs/1508.05354; DOI: 10.1098/rsta.2014.0269]



The striking **EUV excess** of the quiet Sun is the manifestation of the **solar corona problem**.

H.S. Hudson

Zur Frage der Deutung der Linien im Spektrum der Sonnenkorona,

nachdem schließlich

die Anzeichen dafür sich mehr und mehr verdichten, daß in den äußeren Zonen der Sonnenatmosphäre Bedingungen für die Anregung von Spektrallinien vorliegen, die weit über das hinausgehen, was bei thermischem Gleich gewicht zu erwarten wäre, scheint es nicht mehr völlig ab wegig, die Frage zu diskutieren, ob die Koronalinien als verbotene Linien hochionisierter Atome zu deuten sind.

16. März 1939

http://dx.doi.org/doi:10.1007/BF01488890

Volume 256, Issue 3407 8 October 2022, Page 17

DARK matter can be trapped inside massive objects, and much of it may be closer to the surface of stars and planets than we realised. On Earth, there may be more than 10 trillion dark matter particles in each cubic centimetre of the planet's crust.

Physics

Trillions of dark matter particles may lurk in Earth's crust

DARK matter can be trapped inside massive objects, and much of it may be closer to the surface of stars and planets than we realised. On Earth, there may be more than 10 trillion dark matter particles in each cubic centimetre of the planet's crust.

Dark matter is a hypothetical form of matter that isn't visible because it doesn't seem to interact with light. However, it does interact with regular, or baryonic, matter via gravity, and particles of dark matter may occasionally smash into particles of baryonic matter.

Rebecca Leane at Stanford
University in California and Juri
Smirnov at the University of
Liverpool in the UK calculated how
these collisions would affect the
distribution of dark matter inside
celestial bodies. Our galaxy and
most others are in huge clouds of
dark matter, so a constant stream of
these particles is probably entering
every planet and star in the galaxy.

Leane and Smirnov found that this dark matter doesn't simply sink to the centres of planets and stars as some past research has assumed.

"If you're a dark matter particle, you have gravity pulling you towards the centre of the star or the planet, but as you head down you're bouncing off of all the matter on the way to the core," says Leane. "It turns out that even if you give the dark matter as much time as it likes, some of it still ends up near the surface because of all this bouncing."

They calculated that, in the sun,

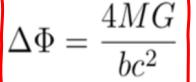
"If there's a bunch of dark matter at the surface of the Earth, that could make it easier to detect" this would result in 100 trillion particles of dark matter or more in each cubic centimetre of the surface.

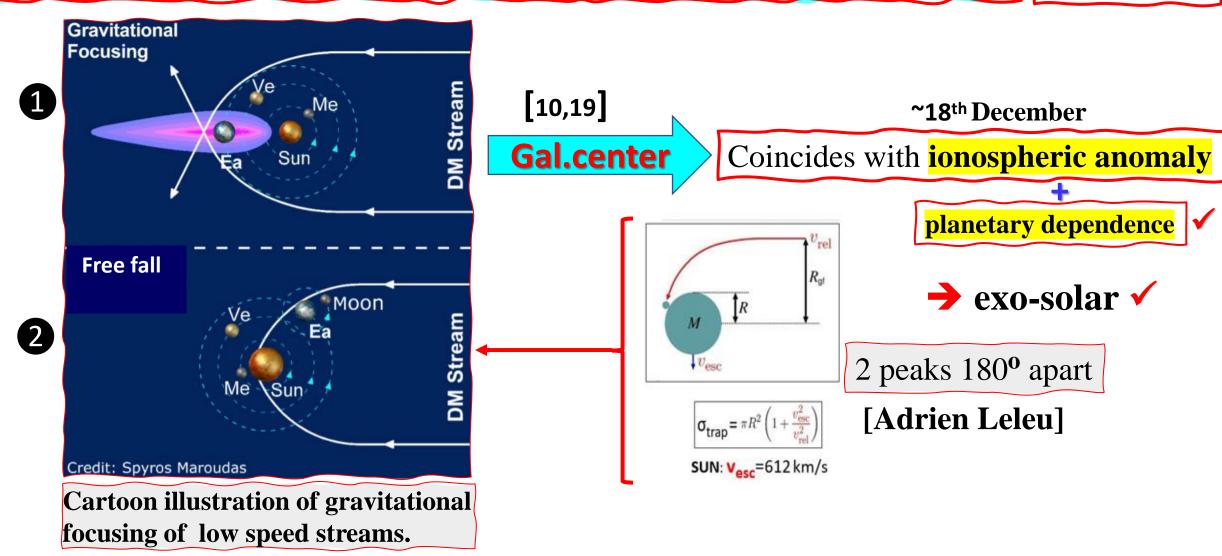
While current detectors aren't built to search for this trapped dark matter – it is expected to move slower than dark matter hurtling in from space, so it would carry less energy and be harder to detect – a high concentration near the surface could help future experiments.

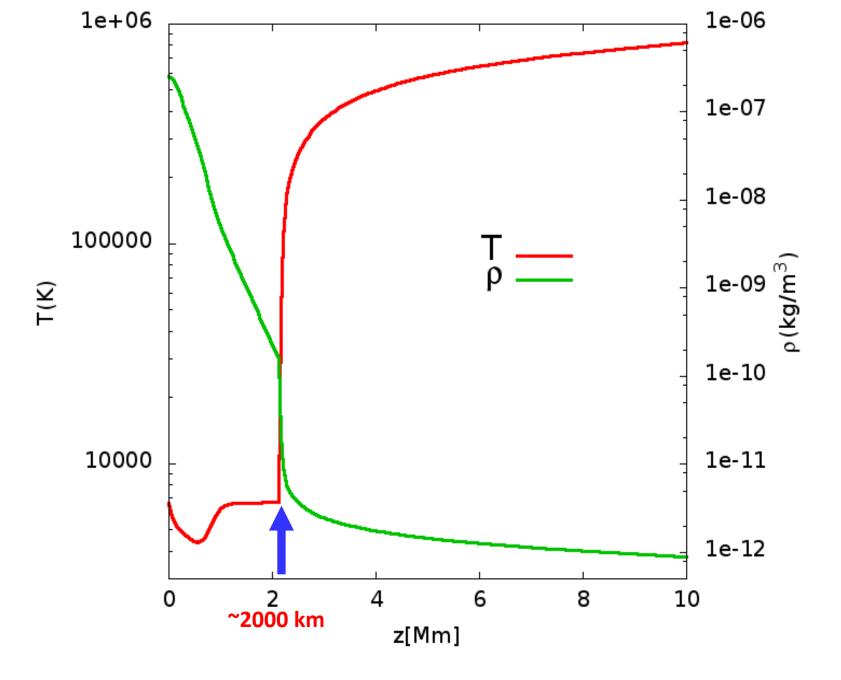
"If there's a bunch of dark matter just sitting at the surface of the Earth, that could make it easier to detect," says Leane.

Leah Crane

Gravitational focusing by the solar system!





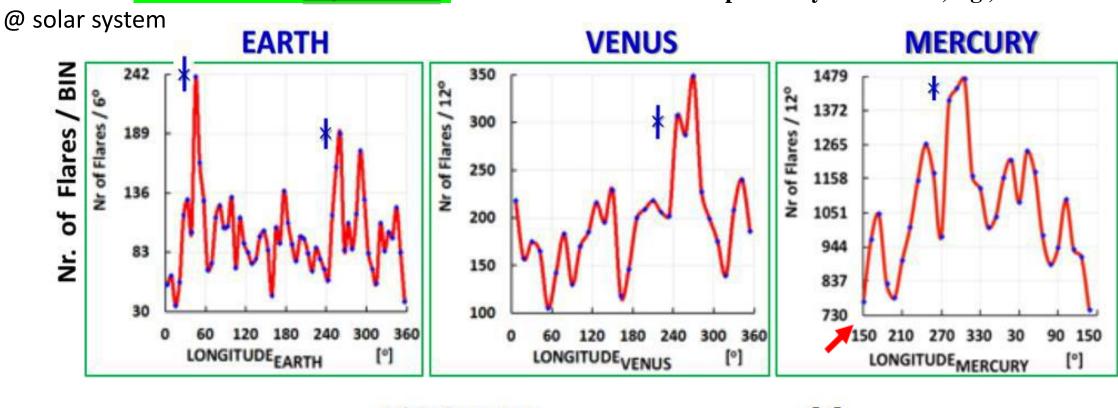


Solar Flares

Peaking planetary relationship,

Biggest and unpredictable "explosions"

excludes remote planetary interaction, e.g., tidal forces



LONGITUDE

[°]

Repeat search of ?? during peaks!

Data from M.J. Aschwanden

DATA TAKING RESULTS

RESULTS:

<u>Data-taking time</u>: 4124 h (172 d)

Frequency range: 660.15 MHz

Axion Masses: 4.77 - 5.43 GHz

<u>Data size</u>: ~ 650 TB !!



QUALITY CHECKS:

Nr.	Parameters	Criteria
1	Frequency stability	$\Delta \nu_0 < 100\mathrm{kHz}$
2	Amplitude variation	$\Delta A_0 < 3\mathrm{dB}$
3	Quality factor	$10^3 < Q_L < 4 \times 10^4$
4	Quality factor shift	$\Delta Q_L < 7 \times 10^3$
5	Temperature variation	$\Delta T_{ m cav} < 3{ m K}$
6	Temperature	$1{ m K} < T_{ m cav} < 273{ m K}$
7	Magnetic field variation	$\Delta ec{B} < 0.1\mathrm{T}$
8	Frequency mismatch	$< 20\mathrm{kHz}$ (before) & $< 80\mathrm{kHz}$ (after)
9	Amplitude mismatch	$< 1\mathrm{dB}$
10	Temperature mismatch	< 3 K



√ Phase-matching of all four cavities.

√ Fast resonance scanning.

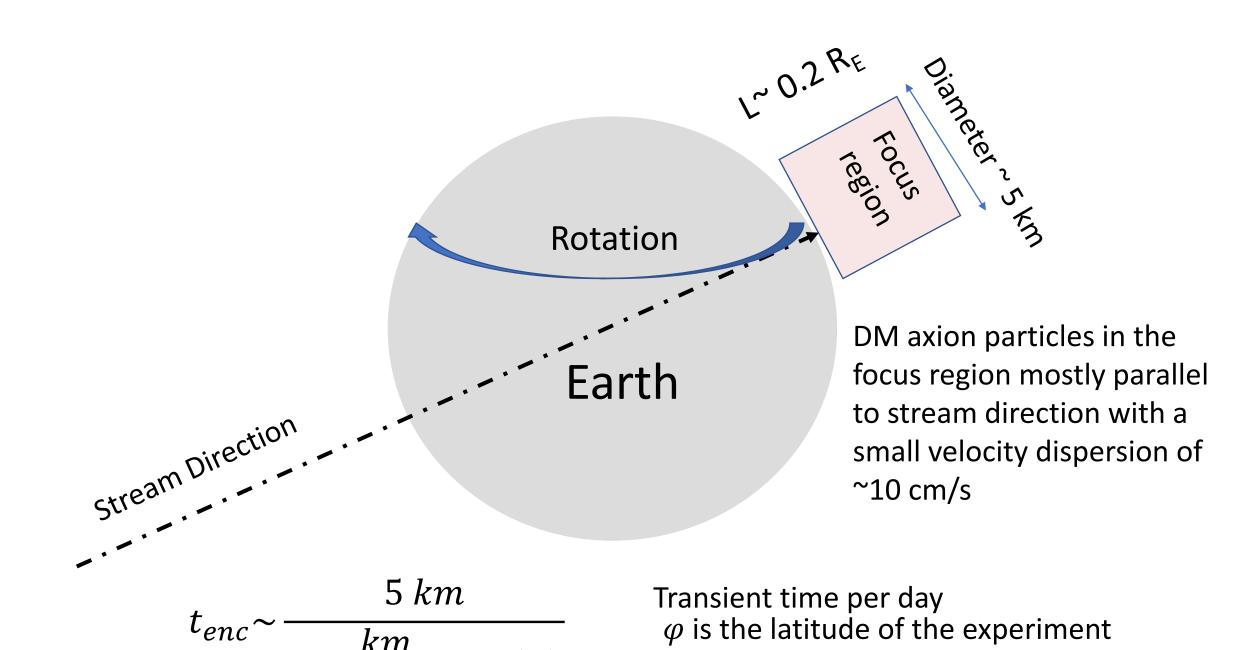
√ Unexplored parameter space.

Total discarded Files: (~4.4%)

SUN:

Plenty of mysteries and anomalies e.g. corona heating & its temperature inversion 1939 - present

Fine grained streams, gravitationally produced high density regions with focus on the surface of the Earth



Most promising cases

Density over nominal of the focus region produced by gravity on a stream	Number of these density regions on the entire surface of the Earth	Probability per that an experi would encound density region	ment ter this
10	20	6e-3	2.2 / <u>yr</u>
100	1.6	5e-4	0.2 / <u>yr</u>
1000	8.0e-02	2.4e-5	0.01 / <u>yr</u>
104	2.0e-03	6e-7	
10 ⁵	4.0e-05	1.2e-8	
10 ⁶	8.0e-07	2.4e-10	



Less than 1 numbers here mean that probabilistically the higher density regions are very unlikely

Lear Professor,

Chinese Journal of Mechanical Engineering (CJME) was launched in 1988. It is a peer-reviewed journal under the

govern of China Association for Science and

Technology (CAST)

and sponsored by Chinese Mechanical Engineering Society (CMES). CJME is indexed and abstracted by SCr5. EI, SCOPUS, DOAJ, etc. As the only official journal of CMES, CJr 4E has been one of the top journals in Mechanical Engineering in China, aiming to become a world-class one. CJME welcomes original research articles, reviews, letters, research highlights, and exitorials.

STREAMING DARK MATTER

Cosmological fine grained streams:

M. Vogelsberger and SDM White, Streams + caustics: the fine-grained structure of Λ CDM haloes, MNRAS 413 (2011) 1419.

- DM could be a superposition of many fine-grained streams.
- Consequence from the collisionless character and the coldness of CDM.
- Each such stream has a very small velocity dispersion.

- > ~10¹⁴ streams in the solar neighborhood.
- \triangleright Half of the local DM density is comprised of the 10⁶ most massive streams.
- \triangleright The most massive individual stream contributing a 1% to the local DM density.

Therefore:

→ Look @ (towards) the Sun!



Steven Weinberg's four tips for aspiring scientists

- 1. You don't have to know everything
- 2. Aim for rough water
- 3. Forgive yourself for wasting time
- 4. History of science as it will your work seem more worthwhile to you. Because, a work in science may not yield immediate results, but to realize that it would be a part of history is a wonderful feeling. CAST!?

Searching for neutrinos from solar flares ... with Super-Kamiokande

No significant solar-flare v signal

above background rate was observed.

https://arxiv.org/abs/2210.12948

2022

Search for Solar Flare Neutrinos with KamLAND.

found no statistical excess of

v's and established upper limits

https://iopscience.iop.org/article/10.3847/1538-4357/ac35d1

2022

Evidence for a New Component of HE Solar γ-Ray Production

The observed multi-GeV γ -ray emission from the solar disk—sourced by hadronic cosmic rays interacting with gas and affected by complex magnetic fields—is **not understood** ... **Most strikingly**, although six γ rays above 100 GeV were observed during the 1.4 yr of solar minimum, none were observed during the next 7.8 yr. These features, along with a 30–50 GeV dip ... were **not anticipated by theory**.

To understand the underlying physics, Fermi-LAT +HAWC obs's of the imminent ... solar Minimum are crucial .

Our work:

>search for planetary dependence!

T. Linden, B Zhou, JF Beacom, AHG Peter, KCY Ng, Q-W Tang, Phys. Rev. Lett. 121 (25th Sept 2018) 131103 https://doi.org/10.1103/PhysRevLett.121.131103

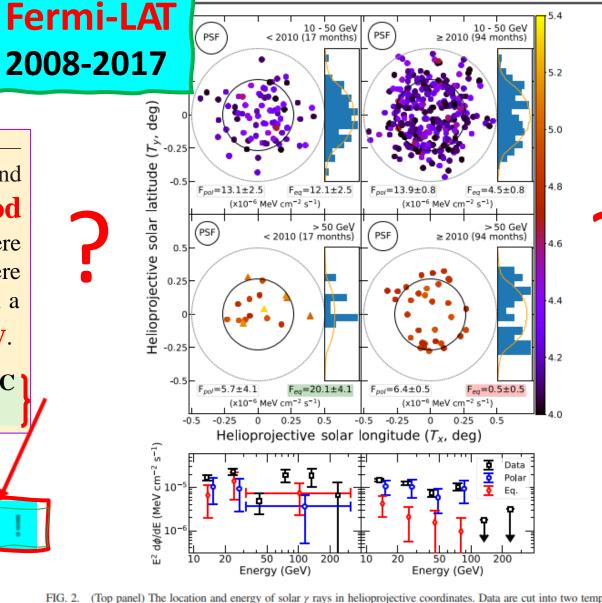
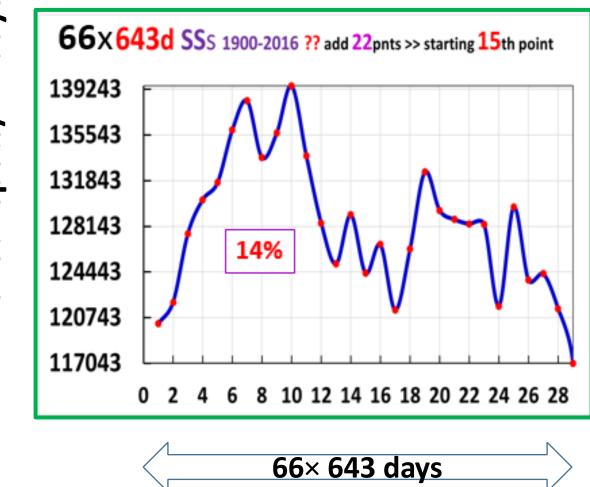


FIG. 2. (Top panel) The location and energy of solar γ rays in helioprojective coordinates. Data are cut into two temporal and two energy bins. The solid disk indicates the solar circle, and the dashed circle indicates the 0.5° ROI. The average 68% containment region of γ rays in each bin is depicted at the top left. The histogram depicts the T_{γ} positions of photons compared to the expectation from isotropic solar emission smeared by the PSF (orange line). Events > 100 GeV are marked with triangles rather than circles. We stress that the exposure after solar minimum significantly exceeds the exposure during solar minimum. Thus, the observed number of counts does not indicate the relative flux. In each bin, we report the flux from the modeled polar and equatorial components, as described in the text. (Bottom panel) The energy spectrum of polar and equatorial emission, divided into regions during (left) and after (right) 54° minimum. The polar emission is approximately constant, while the equatorial emission decreases drastically after solar minimum.

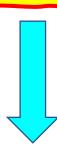
days of Sunspots/22 Z M

NO synod





Gravitational deflection $\Delta \Phi \propto 1/V^2$



- Deflection ($v\sim1\%$ c) $\rightarrow 10^{-6}\times$ Deflection (v=c)
- planetary lensing within inner solar system 🗸



Moon => Earth!

[6,7,10]

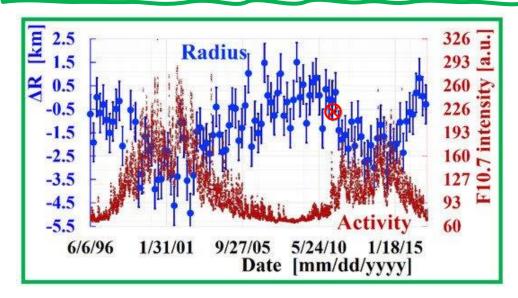


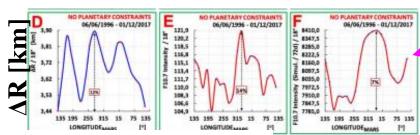
2014

← 2017

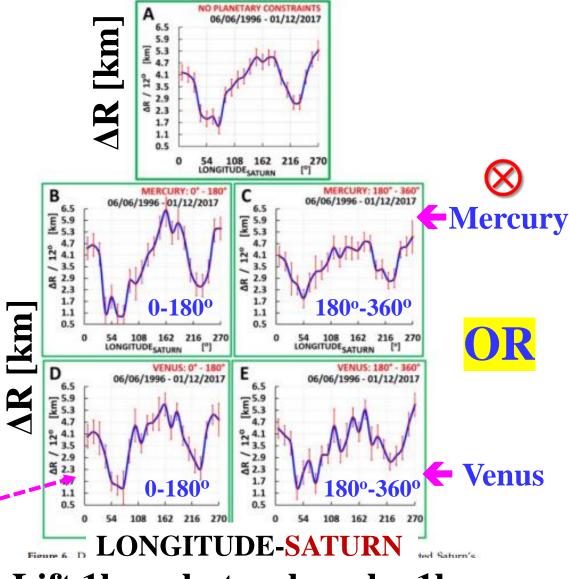
Solar radius variation:

- from helioseismology
 - minimum BIN=72 days
- Planetary relationship





LONGITUDE-MARS



Lift 1km photosphere by 1km: required $\sim 10^{30}$ erg.

AntiQuark Nuggets (AQNs):

dark matter + missing antimatter + (much) more?

https://indico.desy.de/indico/event/20012/session/19/contribution/54/material/slides/0.pdf

N. Raza, L. van Waerbeke, **A. Zhitnitsky**, *Solar Corona Heating by the AQN Dark Matter*, arXiv:1805.01897 (2018),
Phys. Rev. D 98 (2018)103527

Pearls

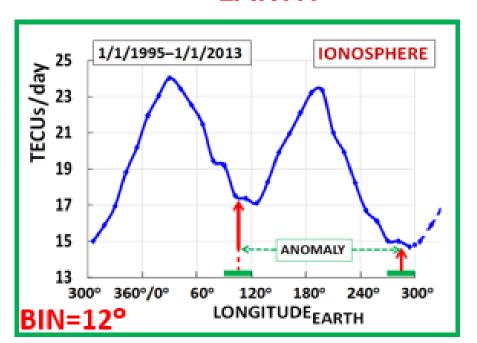
>> see Holger Nielsen' talk, this workshop

Candidates:

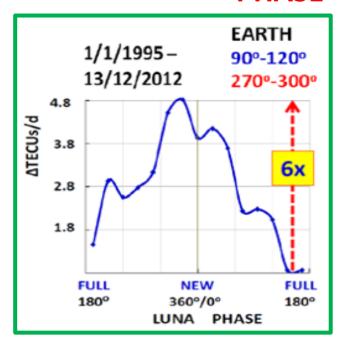
- 1. AQNs
 - Pearls
- 2. Magnetic monopoles
- 3. Dark photons

Or, a combination from + more?

EARTH



EARTH? MOON PHASE



Stream(s) **from G.C. mega-Black Hole?**

Longitude ≈ 266° + TOF (Earth => Sun)

9 18th December

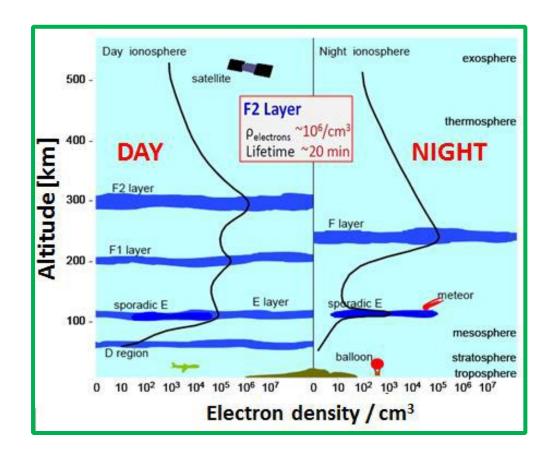
Longitude ≈ 85° + TOF (Moon => Earth)

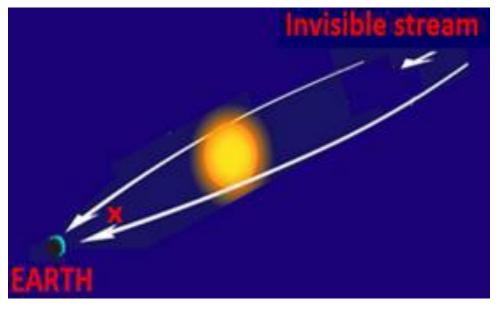
9 17th June, ..?..

EARTH'S IONOSPHERE

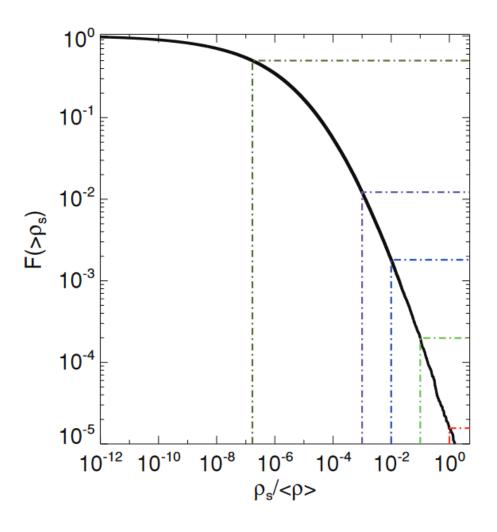
Anomalies lasting for some decades

>>> First obs' 1937/1938





Fine-Grained Stream Abundance



Streams with different densities and their probability in solar neighbourhood.

Density (ρ_s/ρ_0)	Number of streams	Probability (%)
1	1	0.002
0.1	1	0.2
0.01	1	20
10^{-3}	10	100
10^{-4}	500	100
10^{-5}	$2 \cdot 10^{4}$	100
10^{-6}	$4 \cdot 10^5$	100
10^{-7}	$2 \cdot 10^6$	100

Mark Vogelsberger and Simon D. M. White, Mon. Not. R. Astron. Soc. 413 (2011) 1419

2 million streams with density $10^{-7} \rho_0$

STREAMING DARK MATTER

ightharpoonup Conventional halo DM: $U \sim 240 \text{ km/s}$ $aisetarrow 0.45 \text{ GeV/cm}^3$



NO direct detection so far..

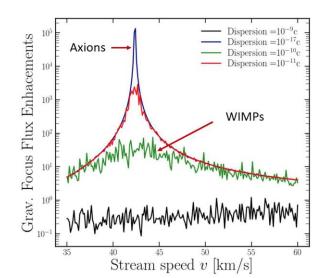
- Streaming DM: For low-speed ($U \sim 0.01c 0.2c$) streams aligned with SUN \rightarrow EARTH: $A \sim 10^{8\pm3}$
- Streaming DM: Earth's self-focusing effects due to intrinsic mass: $A \pm 10^9$
- Streaming DM: Other solar system bodies e.g. the Moon: $A_{axions} \sim 10^5$

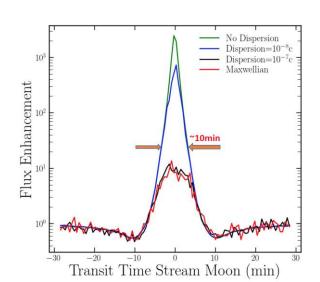


NEW APPROACH:

Wide frequency range.

- Short scanning time per frequency.
- Maximum duty cycle.





Wolf, 1859: solar dynamics is partially driven by planetary tides. a plausible physical mechanism has not been discovered yet... the planetary tidal forces are too small to modulate solar activity. although more complex mechanisms can not be excluded.

N. Scafetta, J. Atm. & Sol.-Terr. Phys. 81–82(**2012**)27

SUMMARY & NOVELTIES

A previously unexplored parameter space has been scanned to extend axion search towards larger rest mass values (19.74 - 22.47 µeV).

- Use of a dipole magnet as a DM axion haloscope.
- No mode crossings over the entire tuning range for the mode of interest.
- Four identical cavities coherently combined through **phase-matching** to increase SNR for the first time in DM axion search.
- State-of-the art tuning mechanism gives a wide mass range (~ 660 MHz) → up to 1 GHz.
- Novel fast-scanning technique to investigate transient events and exclude outliers faster.
- Introduction of a second simultaneous recording channel for **EMI/EMC** parasites.
- New approach and analysis for transient events and search for signal modulations >>> TBDone