Bayesian probability updates using sampling/importance resampling: Applications in nuclear theory



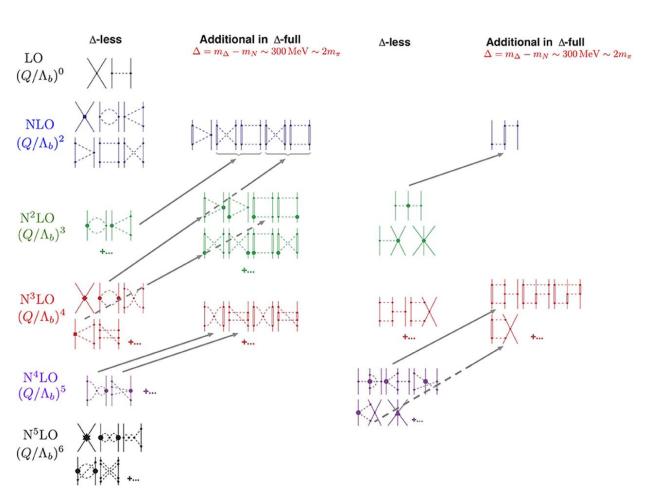
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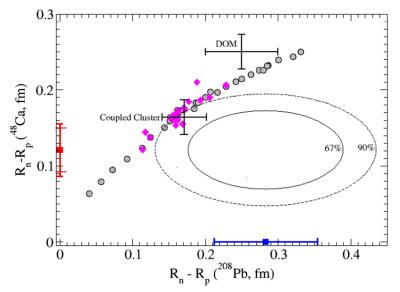


ISNET-9, St. Louis, May 23

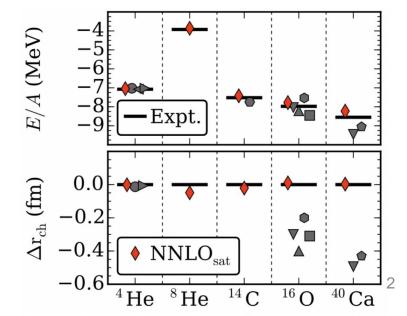
Introduction



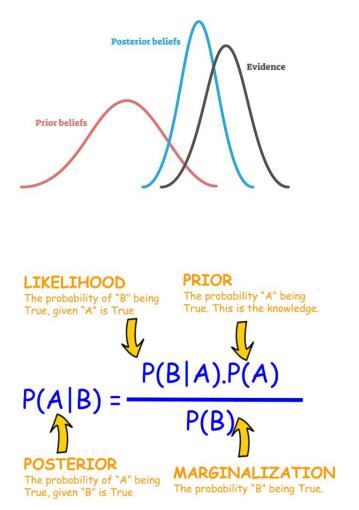
Nuclear interaction based on chiral effective field theory (EFT), parametrized in terms of low energy constants (LECs)



Uncertainty of the nuclear Hamiltonian (nuclear interaction)



BAYESIAN ANALYSIS



Bayesian inference is an appealing approach for dealing with theoretical uncertainties and has been applied in different nuclear physics studies

	Posterior predictive distributions of neutron-deuteron cross sections							
	Sean B. S. Miller, Andreas Ekström, and Christian Forssén							
How Well Do We Know the Neutron-Matter Equation of State at the Densities Inside Neutron Stars? A Bayesian Approach with Correlated Uncertainties								
C. Drischler, R. J. Furn Phys, Rev. Lett. 125 , 2	R. J. Furnstahl, N. Klco, D. R. Phillips, and S. Wesolowski Phys. Rev. C 92 , 024005 – Published 18 August 2015							
Bayesian estimation of the row-energy constants up to rourth order in the nucleon-nucleon sector of chiral effective field theory								
Isak Svensson, Andreas Ekström, and Christian Forssén Phys. Rev. C 107 , 014001 – Published 20 January 2023								
	quantifying model uncertainties in nuclear dynamics							
	D R Phillips ^{9,1} , R J Furnstahl ² , U Heinz ² , T Maiti ³ , W Nazarewicz ⁴ , F M Nunes ⁴ , M Plumlee ^{5,6} , M T Pratola ⁷ , S Pratt ⁴ , F G Viens ³ + Show full author list Published 20 May 2021 • © 2021 IOP Publishing Ltd							
	Journal of Physics G: Nuclear and Particle Physics, Volume 48, Number 7							

Bayesian inference is an excellent framework to incorporate different sources of uncertainty and propagate errors to the model predictions.

• Posterior probability density function (PDF) in Bayes' theorem :

 $\operatorname{pr}(\theta|\mathcal{D}) \propto \mathcal{L}(\mathcal{D}|\theta) \operatorname{pr}(\theta)$

Likelihood function Prior (usually not analytical)

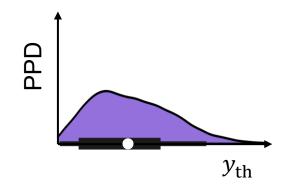
Prior: a priori hypothesis of parameterization θ (e.g. LECs under uniform distribution in a certain range)

Likelihood:

different sources of uncertainty (EFT truncation error, the many-body method error, experimental error...) go in here

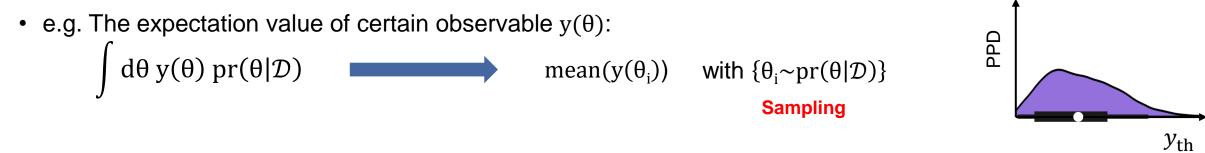
• Posterior predictive distribution (PDD):

 $PDD = \{y_{th}(\theta): \theta \sim pr(\theta|\mathcal{D})\}$



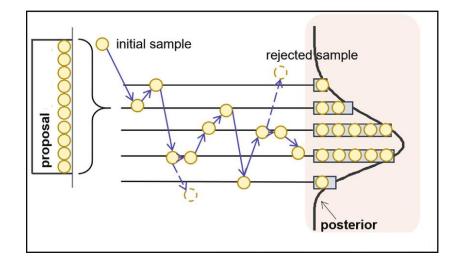
Bayesian Probability and Sampling/Importance Resampling

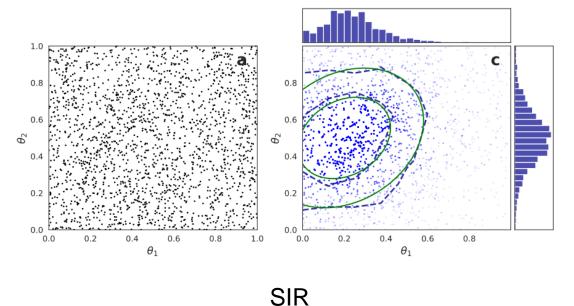
Predicting new observables



• Sampling method:

Markov chain Monte Carlo (MCMC), Sampling/Importance Resampling(SIR)...





MCMC

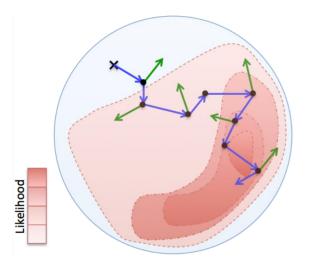
MCMC sampling typically requires many likelihood evaluations, which is often a costly operation in nuclear theory

There are certain situations where MCMC sampling is not ideal or even becomes infeasible:

1) When the posterior is conditioned on some calibration data for which our model evaluations are very costly. Then we might only afford a limited number of full likelihood evaluations.

2) Bayesian posterior updates in which calibration data is added in several different stages. Or in model checking where we want to explore the sensitivity to prior assignments. This typically requires that the MCMC sampling must be carried out repeatedly from scratch.

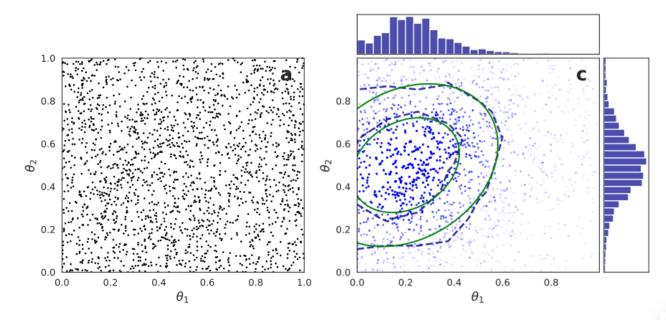
3) Even after we get the pdf using MCMC, the prediction of target observables for which our model evaluations could be very costly and the handling of a large number of MCMC samples becomes infeasible.

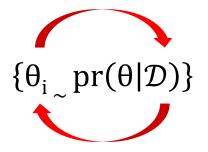


MCMC stochastic processes of "walkers"

The basic idea of SIR is to utilize the inherent duality between samples and the density (probability distribution) from which they were generated

This duality offers an opportunity to indirectly recreate a density (that might be hard to compute) from samples that are easy to obtain.





Bayesian Statistics without Tears: A Sampling-Resampling Perspective Author(s): A. F. M. Smith and A. E. Gelfand Source: *The American Statistician*, May, 1992, Vol. 46, No. 2 (May, 1992), pp. 84–88

weighted bootstrap

Assuming we are interested in the target density $h(\theta) = f(\theta) / \int f(\theta) d\theta$, the procedure of resampling via weighted bootstrap can be summarized as follows:

1) Generate the set $\{\theta_i\}_{i=1}^n$ of samples from a sampling density $g(\theta)$.

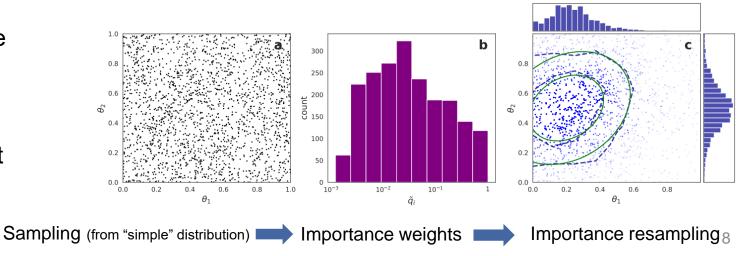
2) Calculate $\omega_i = f(\theta_i) / g(\theta_i)$ for the n samples and define importance weights as: $q_i = \omega_i / \sum_{j=1}^n \omega_j$.

3) Draw *N* new samples $\{\boldsymbol{\theta}_i^*\}_{i=1}^N$ from the discrete distribution $\{\boldsymbol{\theta}_i\}_{i=1}^n$ with probability mass \boldsymbol{q}_i on $\boldsymbol{\theta}_i$.

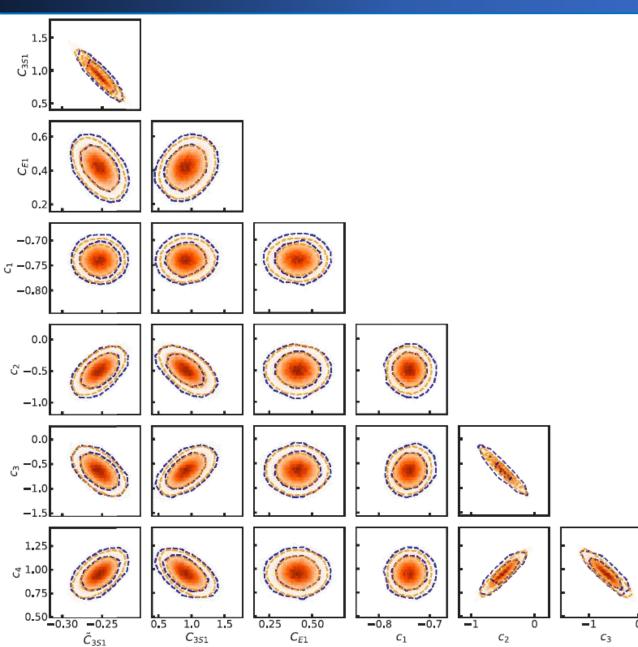
4) The set of samples $\{\theta_i^*\}_{i=1}^N$ will then be approximately distributed according to the target density $h(\theta)$.

Intuitively, the distribution of θ^* should be good approximation of $h(\theta)$ when *n* is large enough. Here we justify this claim *via* the cumulative distribution function of θ^* (for the one-dimensional case)

$$egin{aligned} & \operatorname{pr}\left(heta^*\leq \mathrm{a}
ight) &=\sum_{i=1}^n q_i\cdot H\left(a- heta_i
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ight)\,d heta, \end{aligned}$$



Results and Discussion

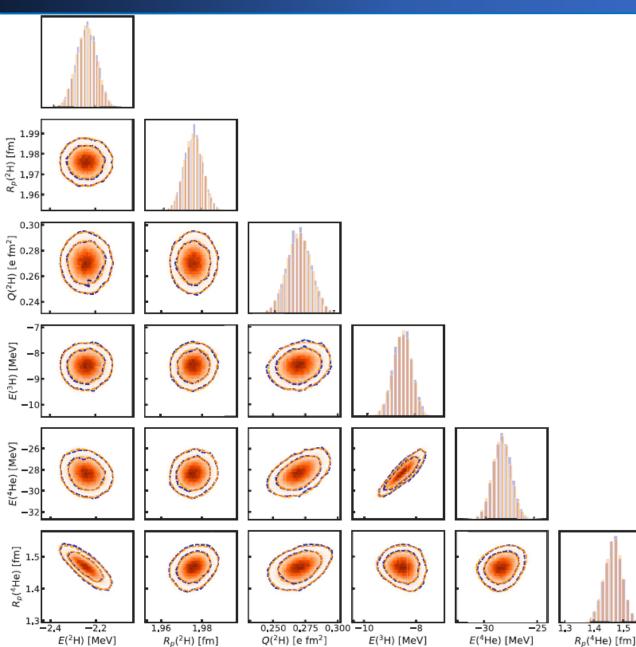


Calibration observables									
Observable	z	$\varepsilon_{\mathrm{exp}}$	$\varepsilon_{\mathrm{model}}$	$\varepsilon_{\mathrm{method}}$	$\varepsilon_{ m em}$				
$E(^{2}\mathrm{H})$	-2.2298	0	0.05	0.0005	0.001%				
$R_p(^2\mathrm{H})$	1.976	0	0.005	0.0002	0.0005%				
$Q(^{2}\mathrm{H})$	0.27	0.01	0.003	0.0005	0.001%				

seven active model parameters: $c_{1,2,3,4}$, \tilde{C}_{3S1} , C_{3S1} , C_{E1} SIR: resample from 2 × 10⁴ samples (uniform distribution)

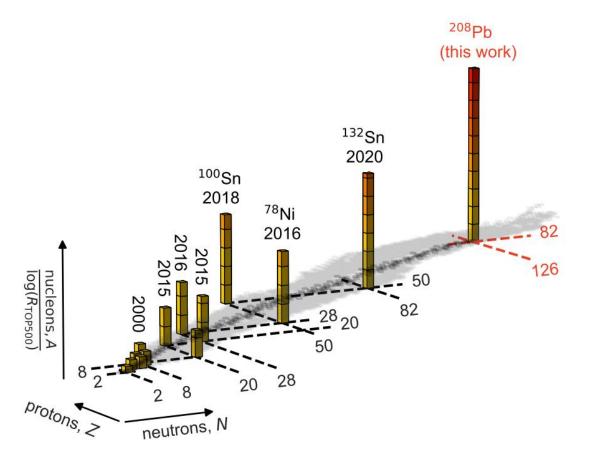
The joint **posterior** of LECs sampled with S/IR (blue) compared with MCMC sampling (orange). The likelihood observables and assigned errors are given in the above Table. The contour lines indicate 68% and 90% credible regions.

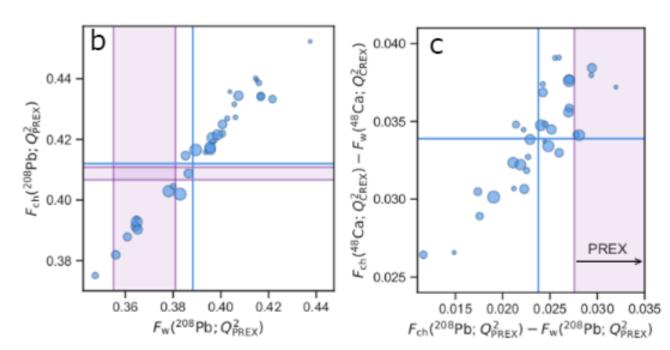
Results and Discussion



The **PPD** obtained from samples of the LECs posterior distribution as shown in previous slide. The bivariate histograms and the corresponding contour lines denote the joint distribution of observables generate by S/IR (blue) and MCMC sampling (orange). The marginal distributions of the observables are shown in the diagonal panels.

Application – neutron skin of ²⁰⁸Pb



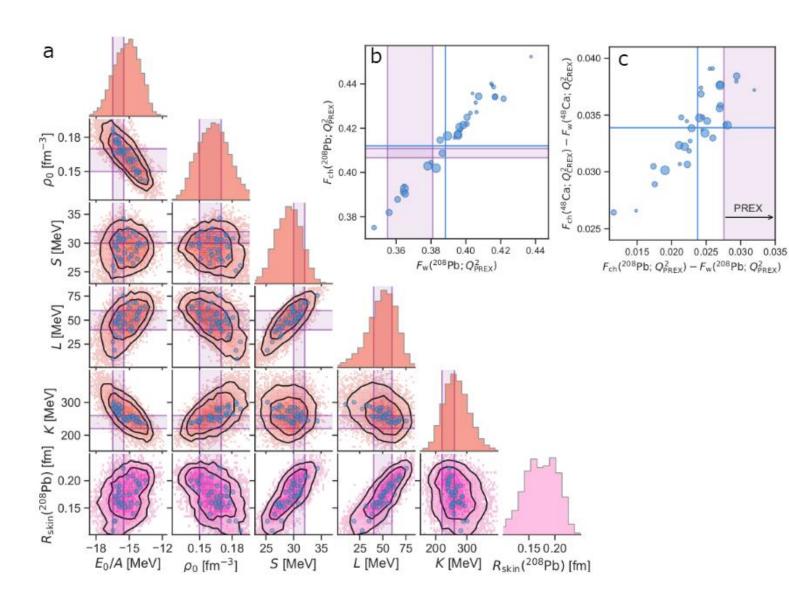


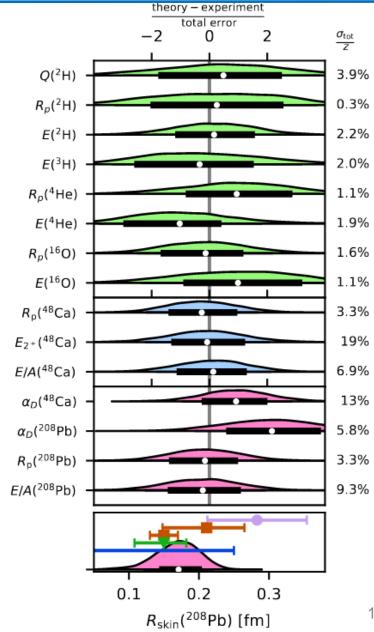
Ab initio calculation of 208 Pb with 3NF up to E_{3max} = 28 by Takayuki etc.

History matching: 34 non-implausible interactions

B. S. Hu*, et al. "Ab initio predictions link the neutron skin of 208Pb to nuclear forces," Nature Phys. 18, 1196-1200 (2022)

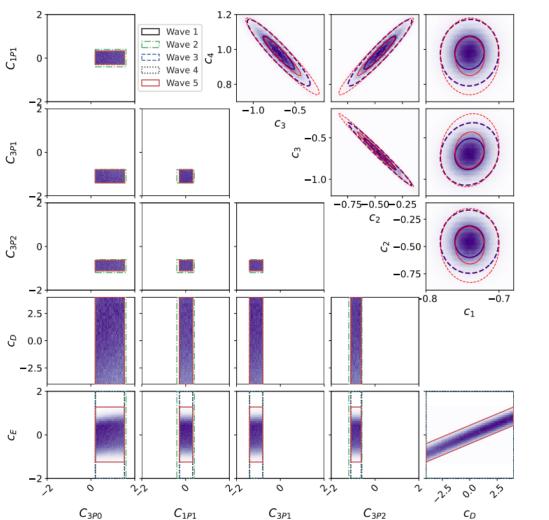
Application – neutron skin of ²⁰⁸Pb





B. S. Hu*, et al. "Ab initio predictions link the neutron skin of 208Pb to nuclear forces," Nature Phys. 18, 1196-1200 (2022)

Application – nuclear matter



After history matching, we acquired 10⁶ non-implausible interaction samples (out of 1 billion)

ig	Observable	z	$\varepsilon_{\mathrm{exp}}$	$\varepsilon_{\mathrm{model}}$	$\varepsilon_{\mathrm{method}}$	$\varepsilon_{ m em}$			
	$E(^{2}\mathrm{H})$	-2.2298	0.0	0.05	0.0005	0.001%			
	$r_p(^2\mathrm{H})$	1.976	0.0	0.005	0.0002	0.0005%			
	$Q(^{2}\mathrm{H})$	0.27	0.01	0.003	0.0005	0.001%			
	$E(^{3}\mathrm{H})$	-8.4818	0.0	0.17	0.0005	0.01%			
	$E(^{4}\mathrm{He})$	-28.2956	0.0	0.55	0.0005	0.01%			
	$r_p(^4\text{He})$	1.455	0.0	0.016	0.0002	0.003%			
	Predicted observables								
	$E(^{6}\mathrm{Li})$	-31.9940	0.0	0.55	0.2000	0.01%			
	$E(^{16}O)$	-127.62	0.0	1.00	0.75	0.5%			
	$r_p(^{16}O)$	2.58	0	0.03	0.01	0.5%			
Bayesian inference with									
00-15-10-10-10-10-10-10-10-10-10-10-10-10-10-				_	uclear prope	matter rties			

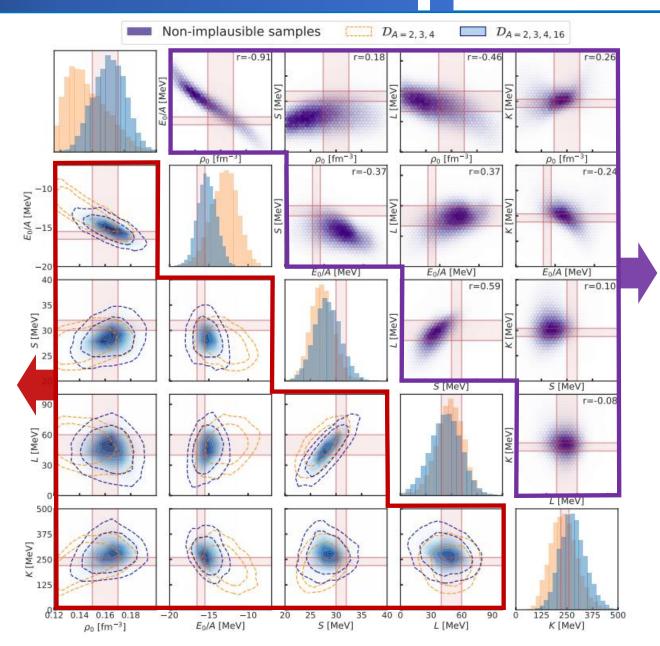
History matching as a good precursor to importance resampling

SIR

Application – nuclear matter

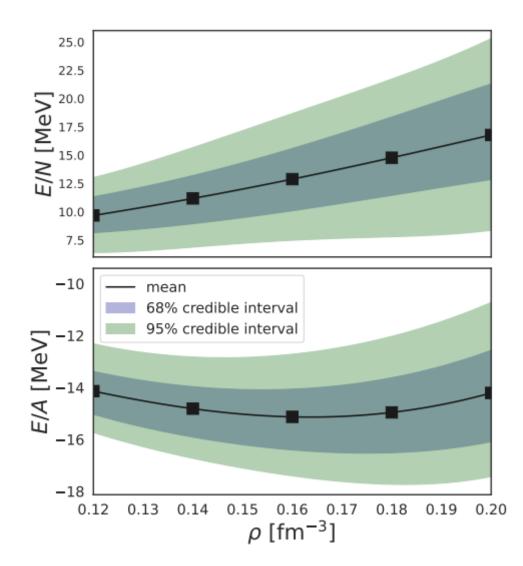
Two PPDs with different PDFs: $D_{A=2,3,4}$, $D_{A=2,3,4,16}$

Note that the same interaction samples are used for different importance resampling stages.

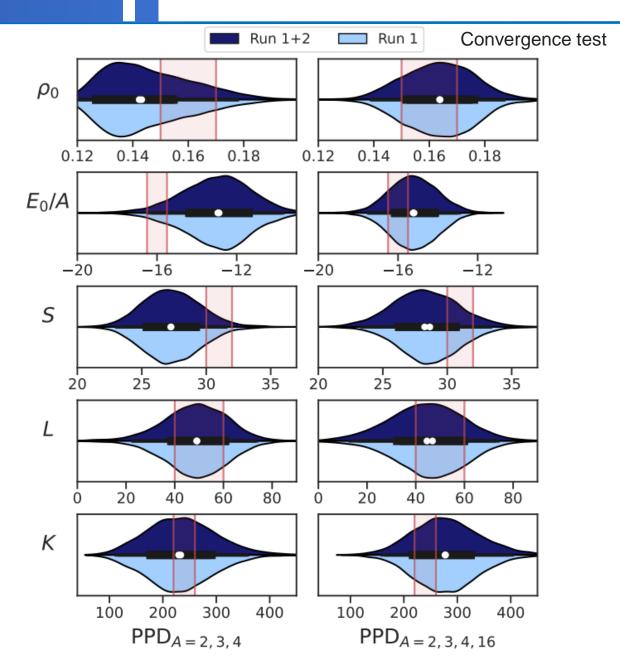


~10⁶ Non-implausible interaction samples form history matching

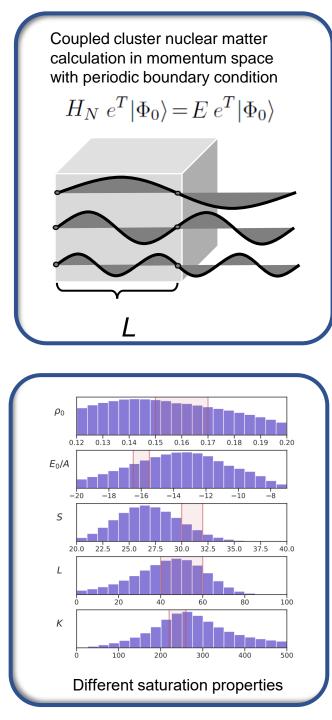
Application – nuclear matter



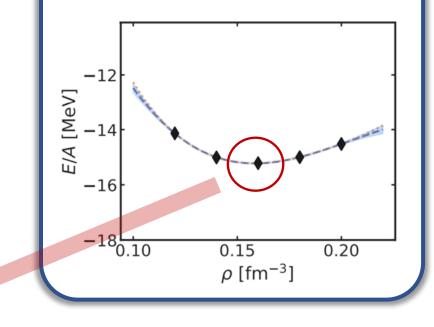
The PPD for the EOS around saturation density



15

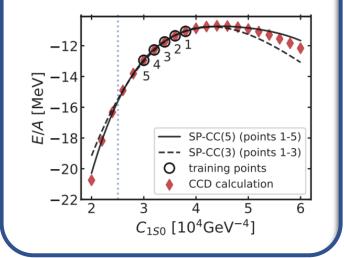


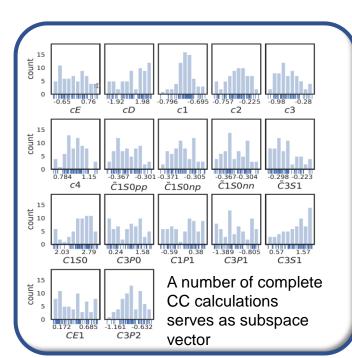
Emulating *ab initio* computations of infinite nucleonic matter



Emulator enables 10^6 times acceleration in this case eg: for SNM (ccd ~200 CPU-hour) vs (emulator ~2ms)

Nuclear matter emulator based on Subspace projected coupled cluster



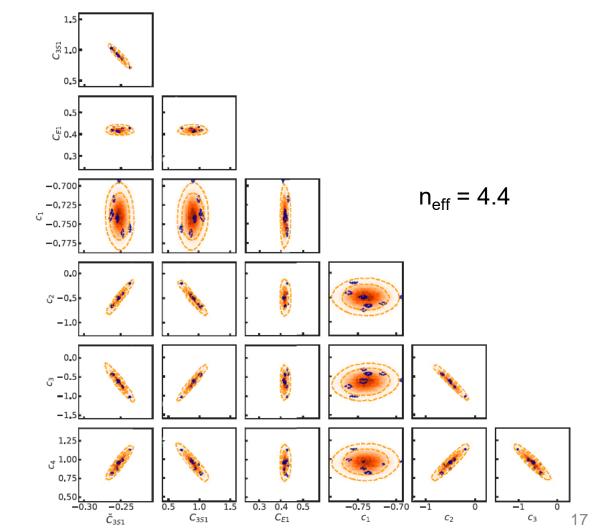


Limitations

- 1. More complex the likelihood function
- 2. Target observables are characterized by very small error assignments

1.5 C^{3SI} S 0.4 n_{eff} = 1589.9 -0.70 J -0.7 -0.80 0.0 S -0.5 -1.0 _ت –0.5. -1.0 -1.5 1.25 J 1.00 0.75 0.50 0.30 0.5 1.0 1.5 0.25 0.50 -0.25 -0.8 -0.7 -1 -1 C_{E1} C351 \tilde{C}_{351} c_1 C2 C3

Metrics: effective number of samples n_{eff} , as the sum of rescaled importance weights, $n_{eff} = \sum_{i=1}^{n} (qi/max(q))$.



- Use the Bayesian method to address the nuclear Hamiltonian (LECs) uncertainty and propagate that to predicted observables.
- In our nuclear physics applications, the Bayesian probability updates are done with sampling/importance resampling to bypass the computational difficulty.
- Limitations of sampling/importance resampling are discussed.
- Open questions: better metrics for the SIR
 other suitable sampling techniques