

Searching for Compact Binary Coalescences with Deep Learning

Damon Beveridge

The University of Western Australia

01/30/2023



ARC Centre of Excellence for Gravitational Wave Discovery

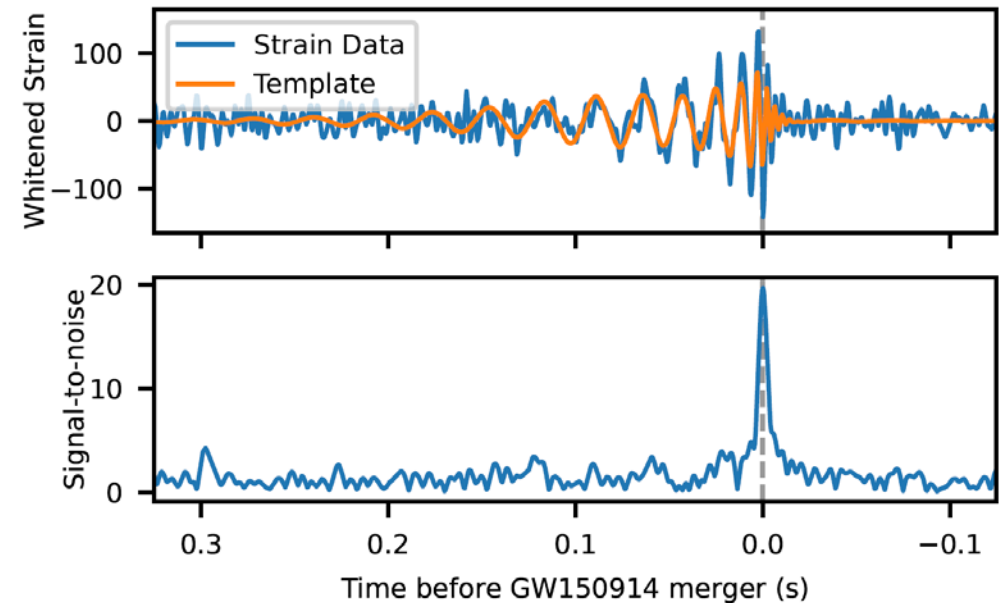


Outline

- Background
- Motivations
- Our Approach
- Binary Black Hole Detection
- Binary Neutron Star Detection
- Challenges
- Future Work
- Conclusion

Background

- Compact binary coalescences (CBCs) involve binary systems with black holes and/or neutron stars that spiral into a merger and emit gravitational waves
- Gravitational waves (GWs) are detected using the LIGO and Virgo detectors
- Current search techniques involve matched filtering to identify signals
 - Considered the optimal method for detecting modelled signals in stationary noise
 - Use a bank of template waveforms to search for correlated signals

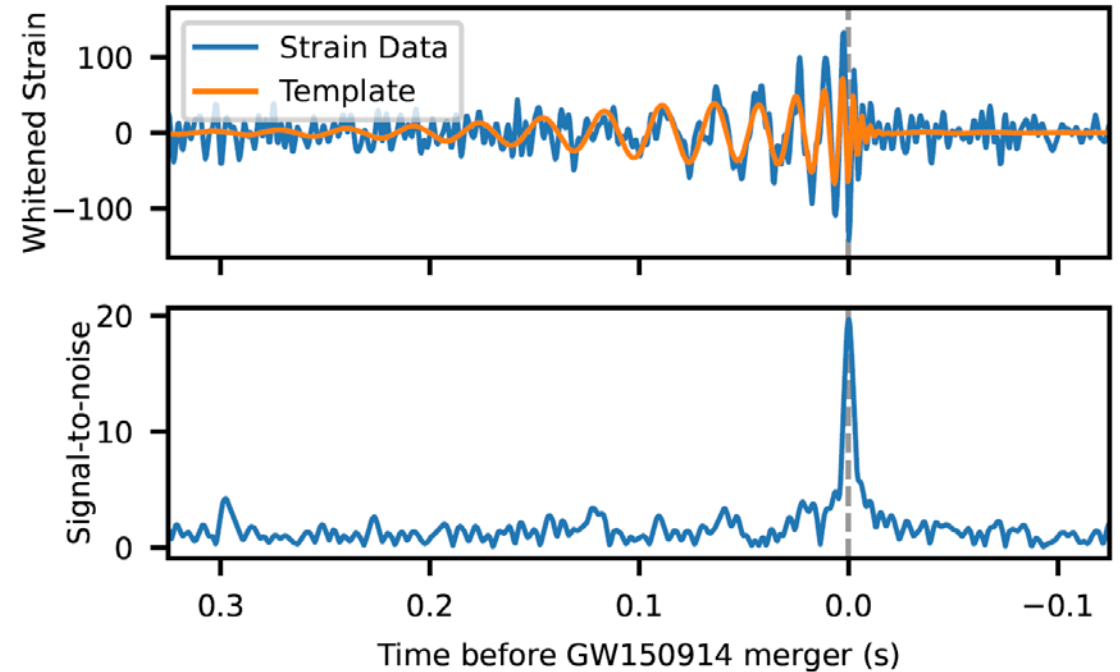


Motivation

- For large parameter spaces, search pipelines require a significant number of templates which adds latency to post-processing
 - Significant computational cost
 - Efforts being made to improve latency, but there is a limit
- Detecting events rapidly after merger, or pre-merger, can lead to multi-messenger astronomy to study these mergers
- Detector data is non-stationary, so the standard matched filtering is not optimal

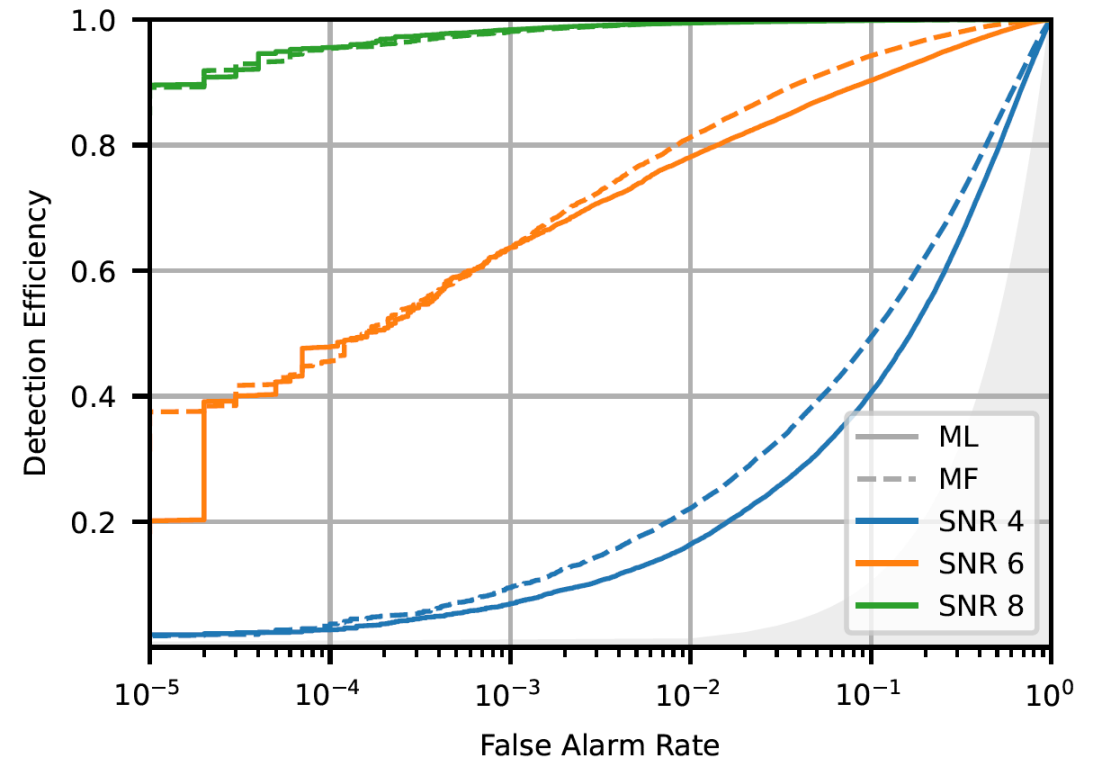
Our Approach

- Signal length increases for lower mass binaries
 - Difficult to design models searching for signals with varying durations
- We use SNR time-series from matched filtering as our networks input
- Proof of concept using Gaussian generated noise
- Use a combination of CNN, ResNet and LSTM layers



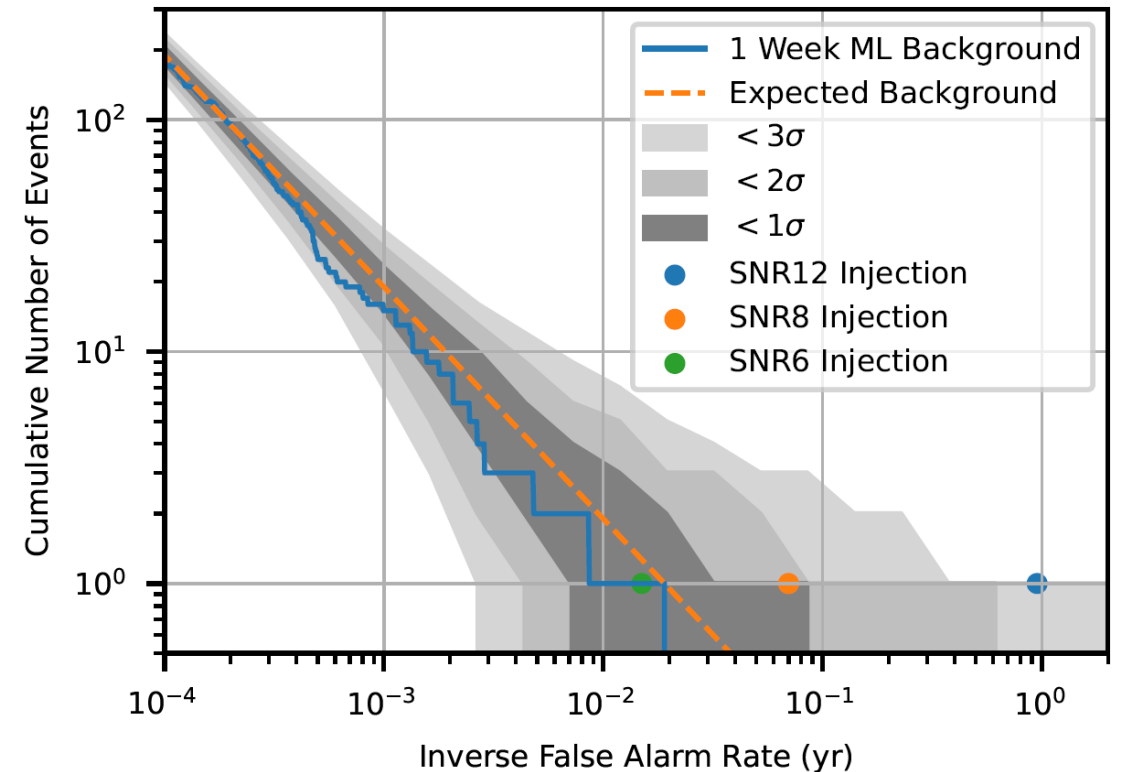
Binary Black Hole (BBH) Detection

- Train on a small set of template banks
- Sample lengths of 1s
- SNR range $\in [5, 50]$
 - Biased to lower SNRs
- Component Masses $\in [10, 80]$
- Dimensionless spin (z-axis) $\in [-1, 1]$



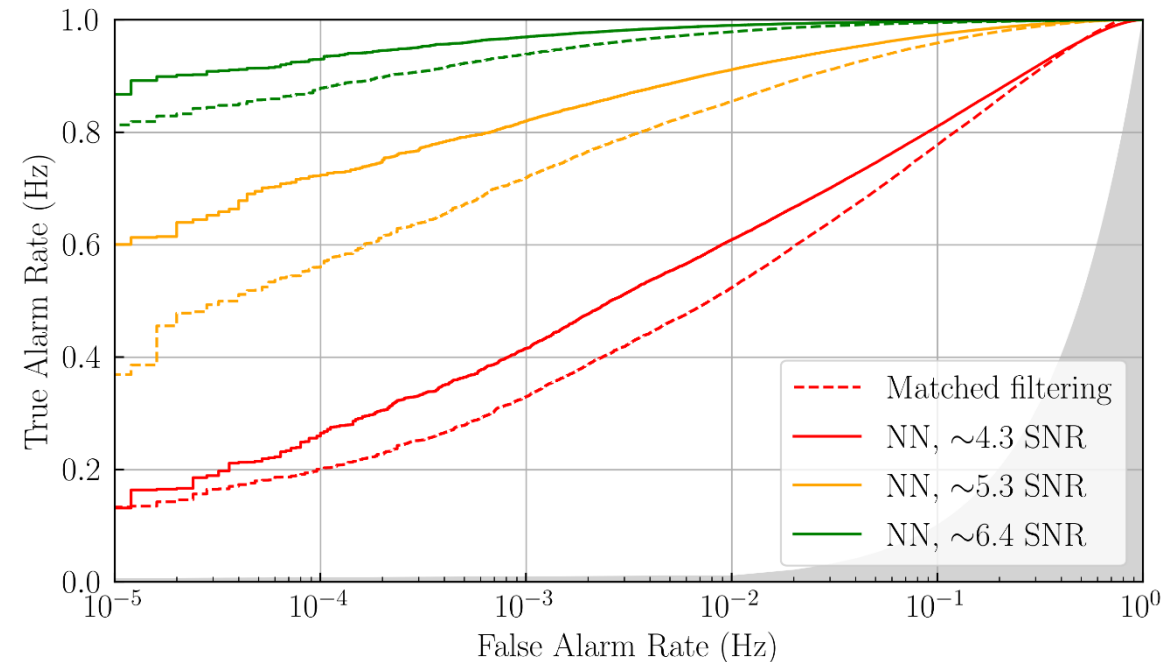
BBH Detection - Testing

- Remove final softmax layer from our model to unbound the ranking statistic
- Verify statistic with inference on 1 year of pure noise



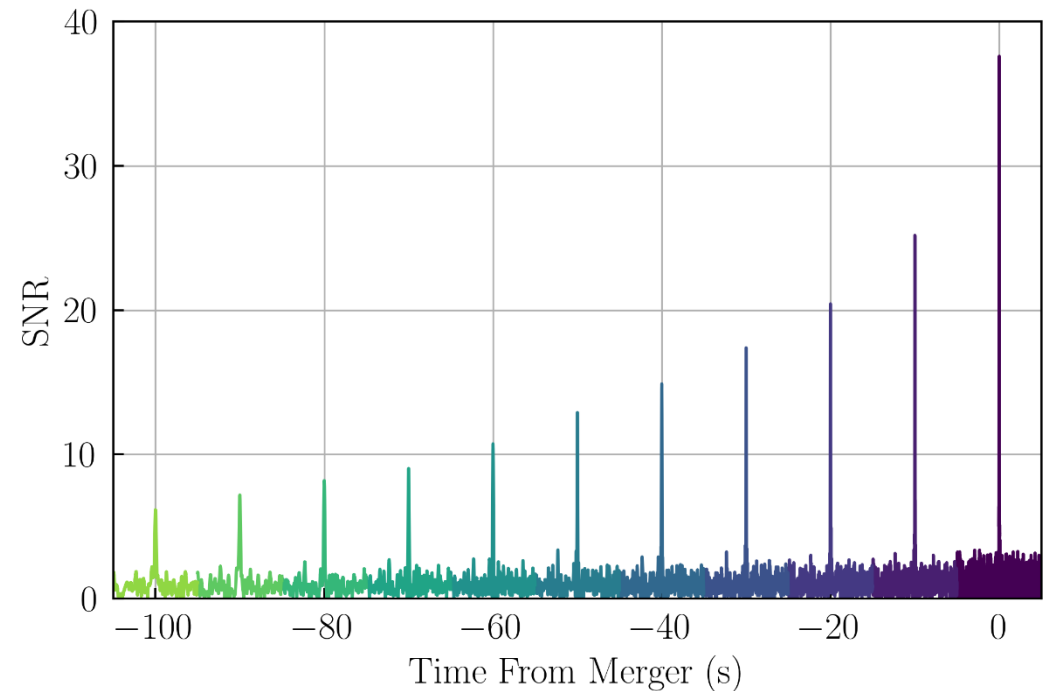
Binary Neutron Star (BNS) Detection

- We have begun similar work for BNS detection using the SNR time series input
- SNR time series response is dependent on signal chirp mass, so we include the template chirp mass as an input
- Tune in to Alistair's talk on Wednesday for more



Pre-merger BNS Detection

- Feasible to conduct pre-merger detection as total signal strength is combined into a smaller window in the SNR time-series
- Goal is to confidently detect events prior to merger so they can be observed with EM telescopes



Challenges

- SNR time series data scales the disk usage linearly with the number of templates we train/test with
 - Cannot all be stored in memory for training
- Inference over many years of data to validate our model requires improved infrastructure
- Limitations when using real data in sample generation

Future Work

- Expand detection to multiple detectors
- Improve data loading/generation infrastructure
- Pre-merger detection
- Implementation offline/online during detector operating runs
 - Integration into the SPIIR pipeline

Summary

- Machine learning can be used to accelerate the detection of GWs from compact binary coalescences
- We have shown that using the SNR time series from matched filtering is a valid approach
- Can match the results of matched filtering search with stationary Gaussian background (where matched filtering is considered optimal)
- We plan to expand this work with real noise, multiple detectors, pre-merger detection, and eventually into online use