



Deep learning based reconstruction for DUNE

QML for $\nu\beta\beta$

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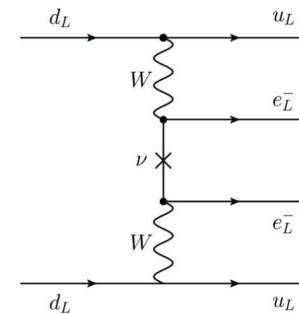


The $0\nu\beta\beta$ decay of ^{136}Xe isotope

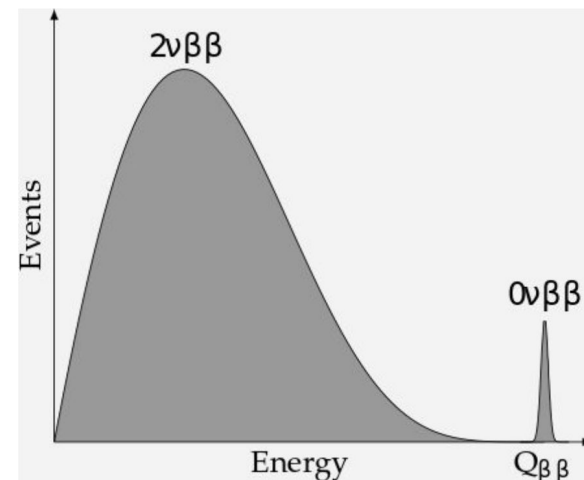
Hypothetical BSM decay process: $(A, Z) \rightarrow (A, Z + 2) + 2e^- + \cancel{2\nu_e}$

Consequences:

- Neutrino and anti-neutrino coincide (Majorana fermions)
- Lepton number violation
- High Q-value, above natural radiation energies: $Q_{\beta\beta} = 2.458 \text{ MeV}$



Feynman diagram



Energy spectrum

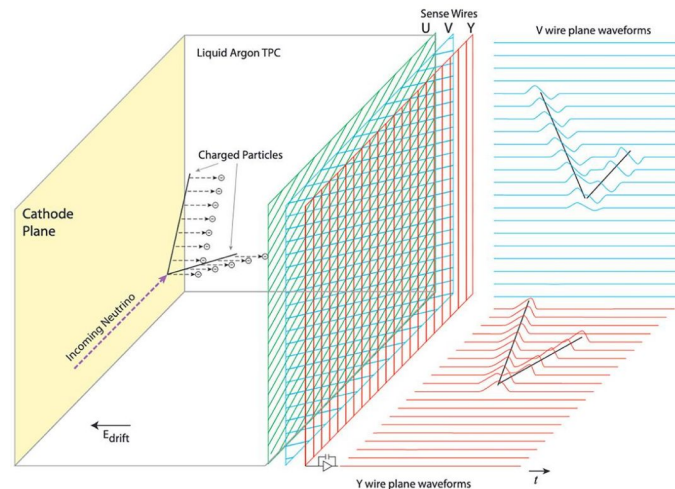
$0\nu\beta\beta$ searches at DUNE experiment

DUNE main goals:

- CP violation in the neutrino sector
- Neutrino mass hierarchy
- Supernova neutrino bursts
- Proton decay

Far Detectors (FD): four 10 ktoms fiducial mass LArTPCs

A 10kt FD module doped at $\sim 2\%$ ^{136}Xe could allow for $0\nu\beta\beta$ search



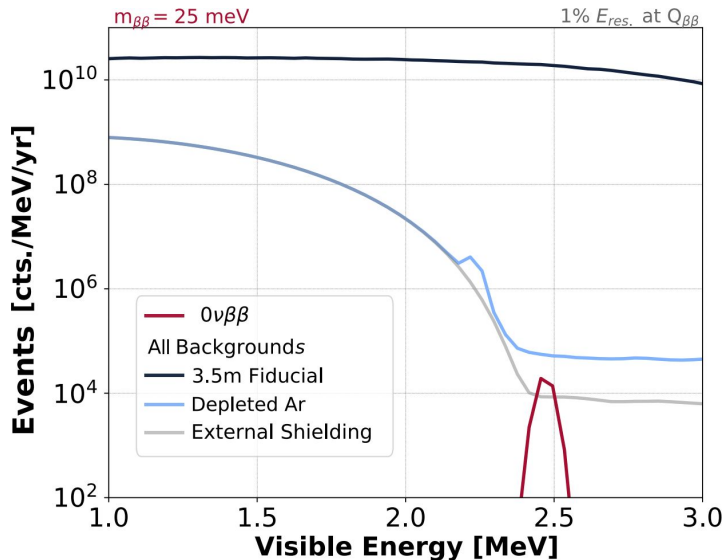
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Background sources - Mitigation strategies

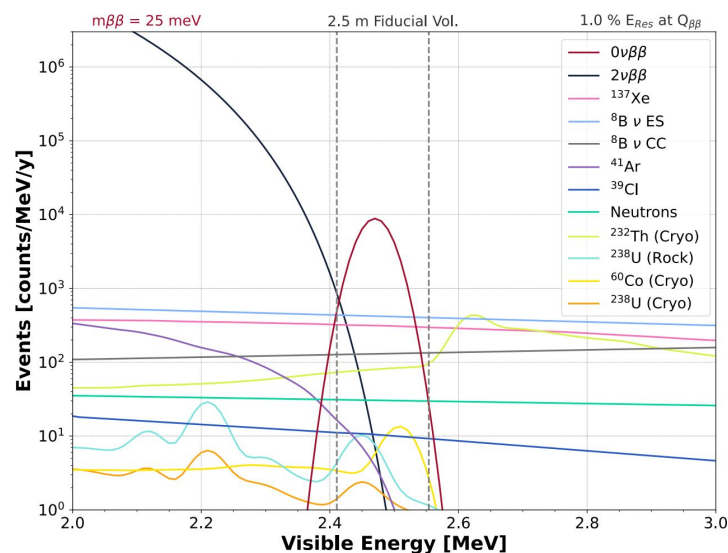
- Environmental radioactivity → VOLUME FIDUCIALIZATION
- Neutrons from (α, n) reactions → PASSIVE SHIELDING
- **Single β emission with $Q_\beta \sim 2.5$ MeV in $^{42}\text{Ar} \rightarrow ^{42}\text{K} \rightarrow ^{42}\text{Ca} \rightarrow$ USE OF DEPLETED ARGON**

- Solar neutrinos (CC and ES interactions) → PHOTON COINCIDENCE TAG
- Cosmogenically-activated radioisotopes (^{41}Ar , ^{137}Xe , ^{39}Cl ...) → COINCIDENCE MUON TIMING VETO
- ^{136}Xe $2\nu\beta\beta$ → IRREDUCIBLE BACKGROUND

Mitigation strategies impact



Signal and Background spectra after mitigation



$0\nu\beta\beta$ vs ^{42}Ar single- β Classification

IN FACT DEPLETED ^{42}Ar CANNOT BE PRODUCED IN HIGH QUANTITIES

NEED OF DIFFERENT STRATEGY \rightarrow MACHINE LEARNING

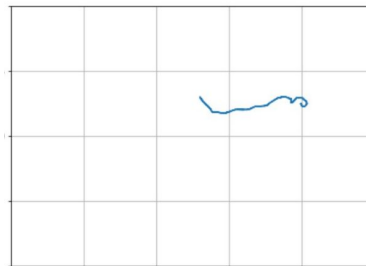
Generate Toy dataset:

- Geant4 simulated high resolution β and $\beta\beta$ tracks at $Q=2.458$ MeV (11k and 10k events respectively)
- DUNE resolution: 3D track voxelization with $5\times 5\times 1$ mm³ bins.
- No detector effects taken into account

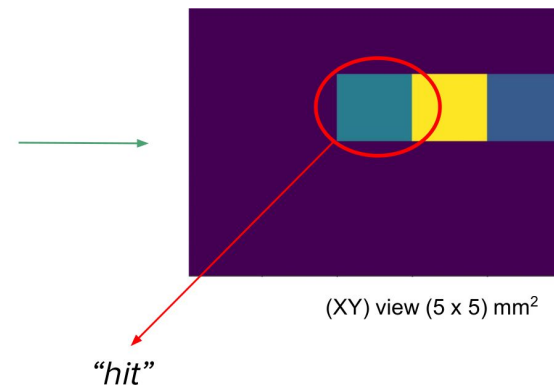
Issues:

- Low energy resolution: ~ 1 MeV
- Short tracks ~ 1.2 cm
- Too few pixels for pattern recognition

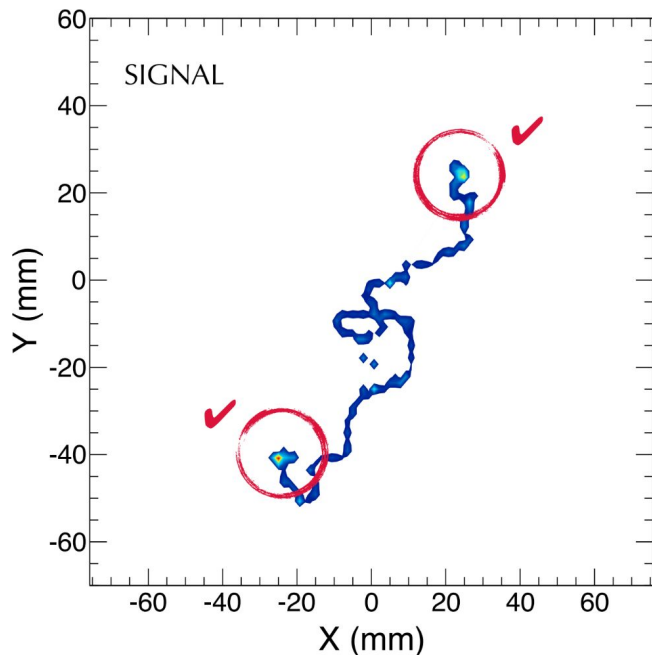
Geant4 track



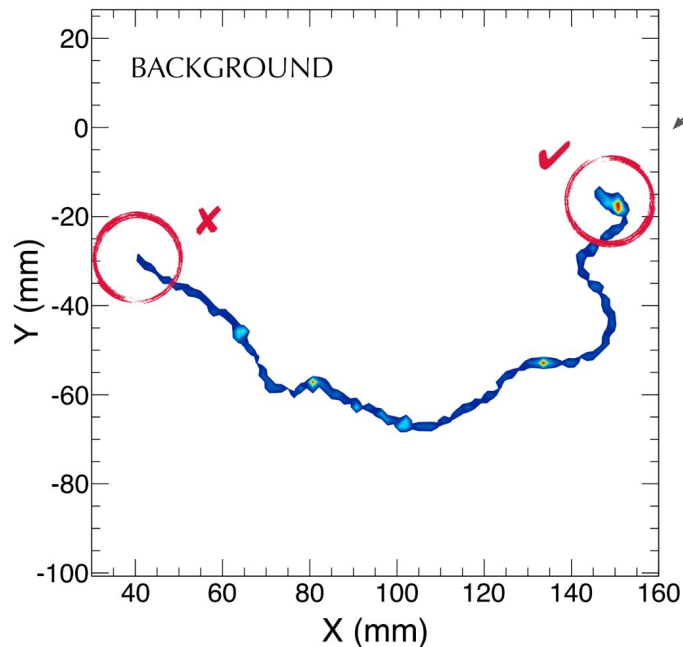
Voxelized projection



Signal vs background

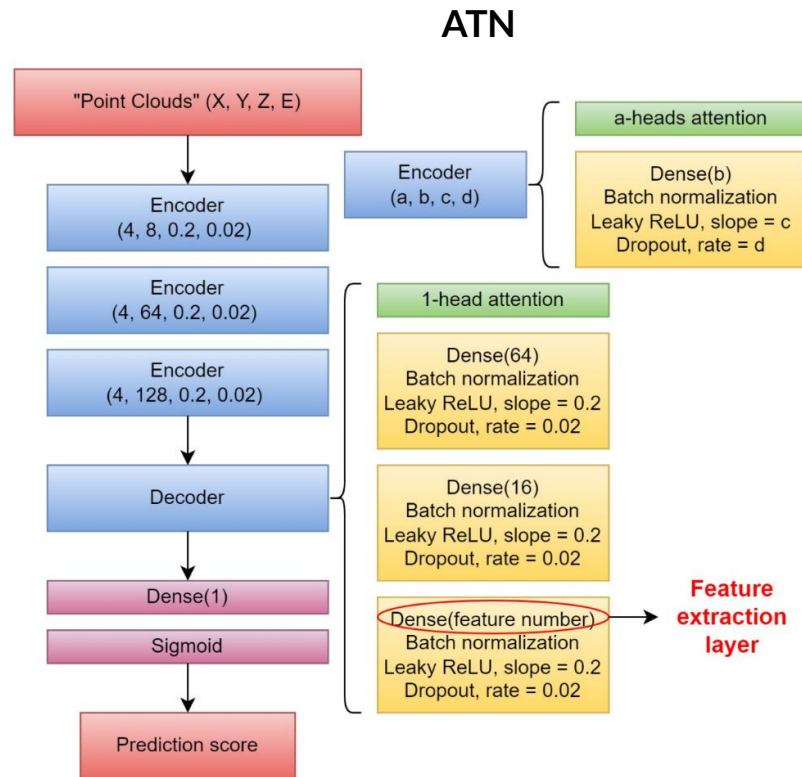
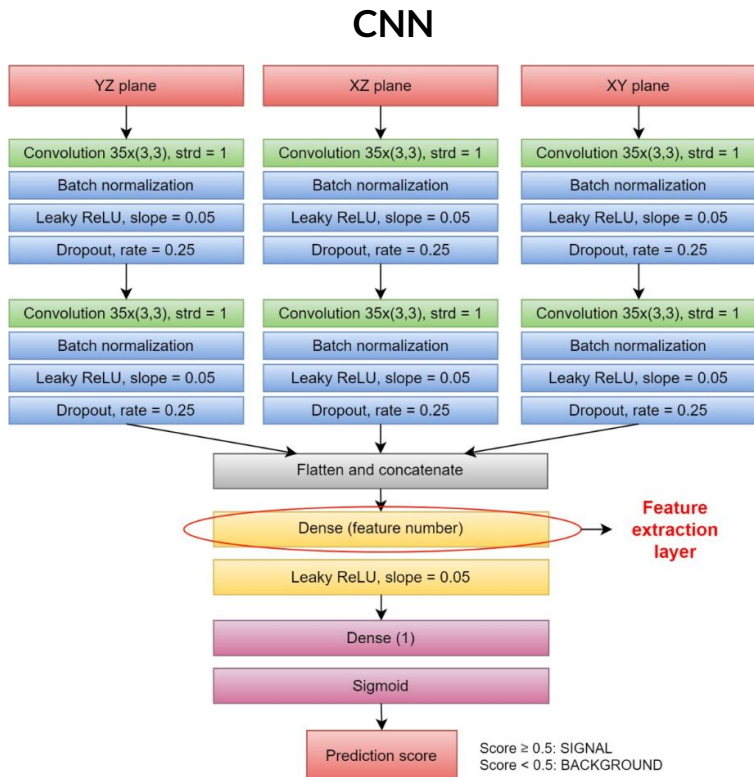


$0\nu\beta\beta$:
Two blobs of energy
deposition at both ends

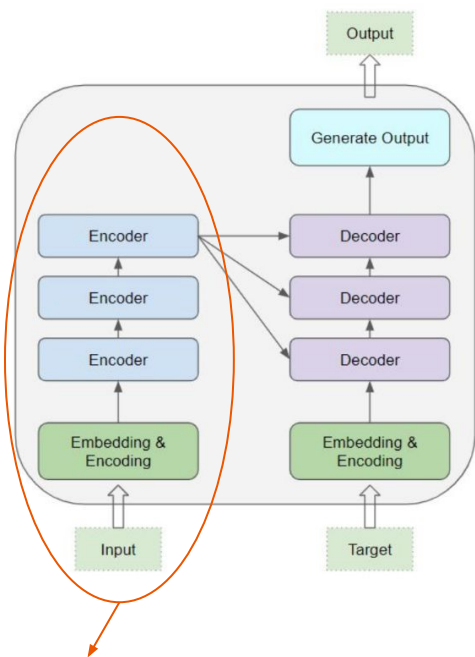


Single β :
Only one blob of energy
deposition at track end

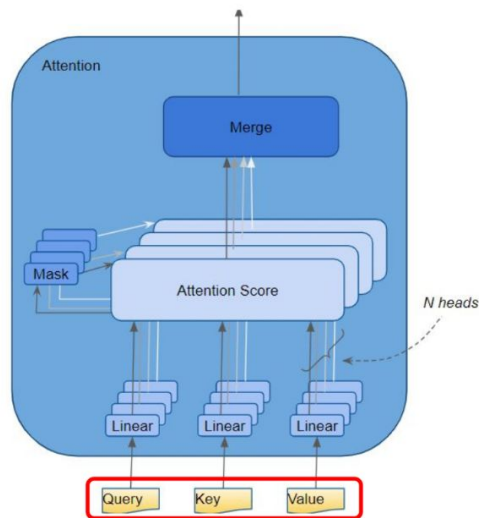
Feature Extractors



Transformer Block



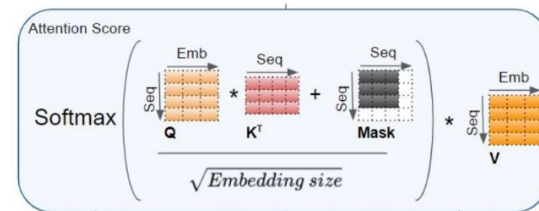
Transformer Encoder



Trainable weights

Attention Mechanism

$$\mathbb{R}^{q \times h} \mapsto \mathbb{R}^{q \times h'} \text{ (self)}$$



Evaluation - Resolution study



Consideration for ProtoDUNE geometry

- DUNE wire pitch is 5 mm on x, y axis and 1 mm on z axis
- DUNE energy sensitivity is ~ 0.1 MeV per voxel
- $0\nu\beta\beta$ events activate very few hits: classification accuracy is poor $\sim 68\%$ with CNN

What if we could build a more efficient detector?

- Increase spatial resolution to 1 mm on all axes
- Increase energy sensitivity: up to 10x factor

Evaluation - Resolution study

Spatial resolution correlates more than energy resolution

Low resolution ($1\text{mm} \leq \text{pitch} \leq 5\text{ mm}$):

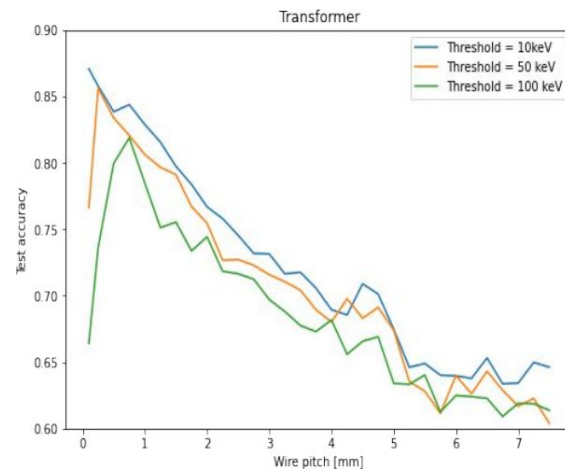
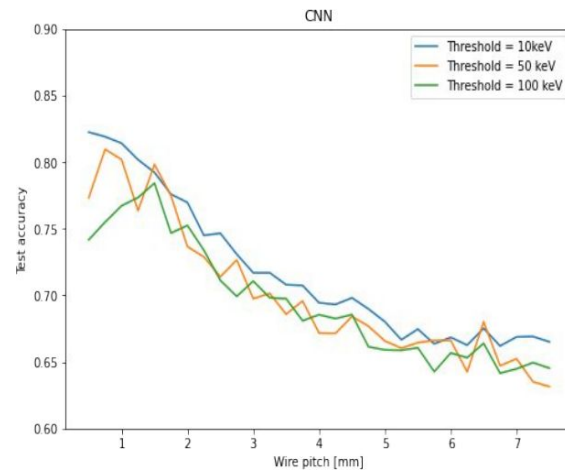
- less voxels
- few active hits

CNN performs better.

High resolution ($\text{pitch} \lesssim 1\text{mm}$):

- more voxels
- more active hits
- accuracy drops for low energy resolution:
energy threshold too high to activate the smaller voxels

ATN is **more efficient** and reaches higher accuracy





Thank you !

Questions?

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