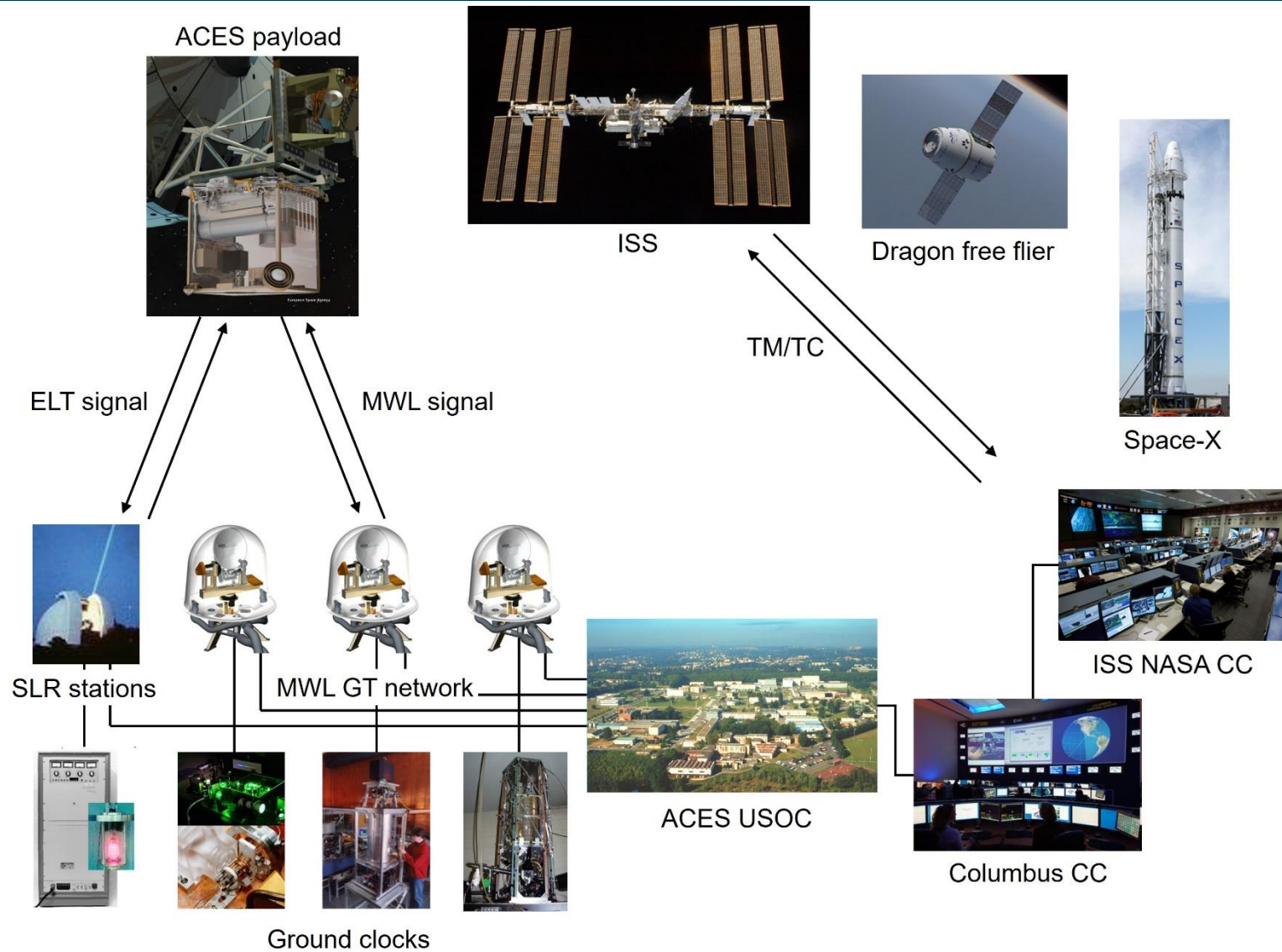




Atomic Clock Ensemble in Space

Luigi Cacciapuoti
9th Symposium on Frequency Standards and Metrology
Kingscliff – Australia
16-20 October 2023

ACES mission concept

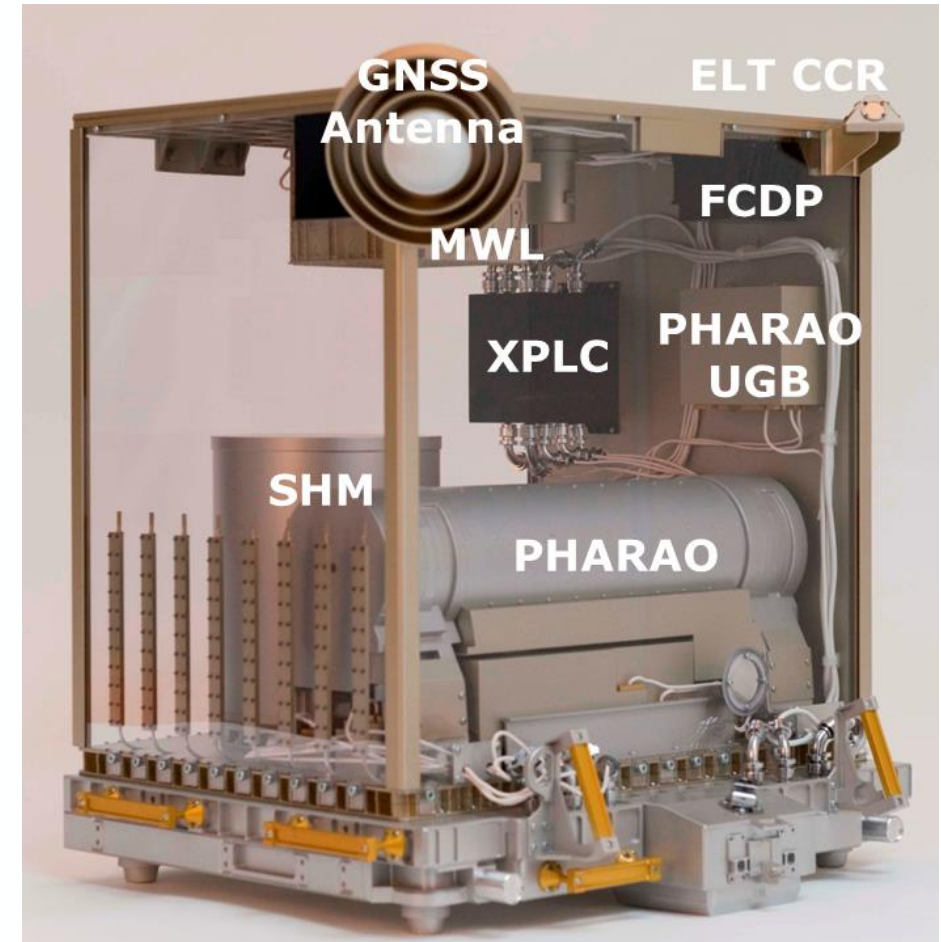


The Columbus module



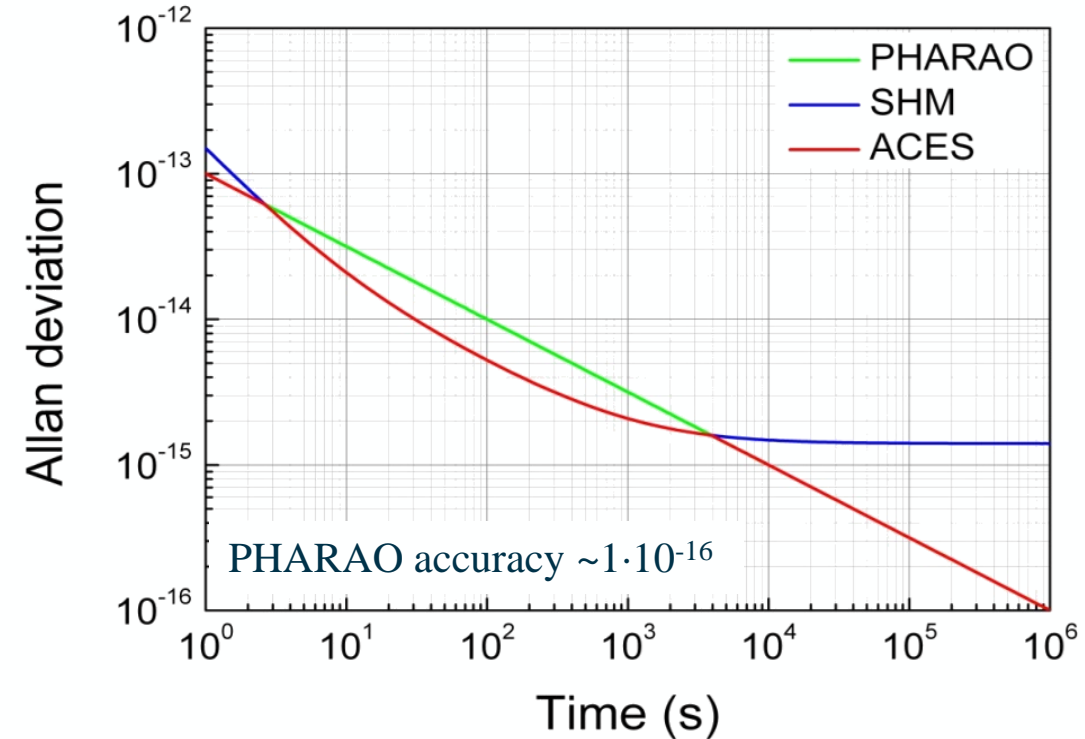
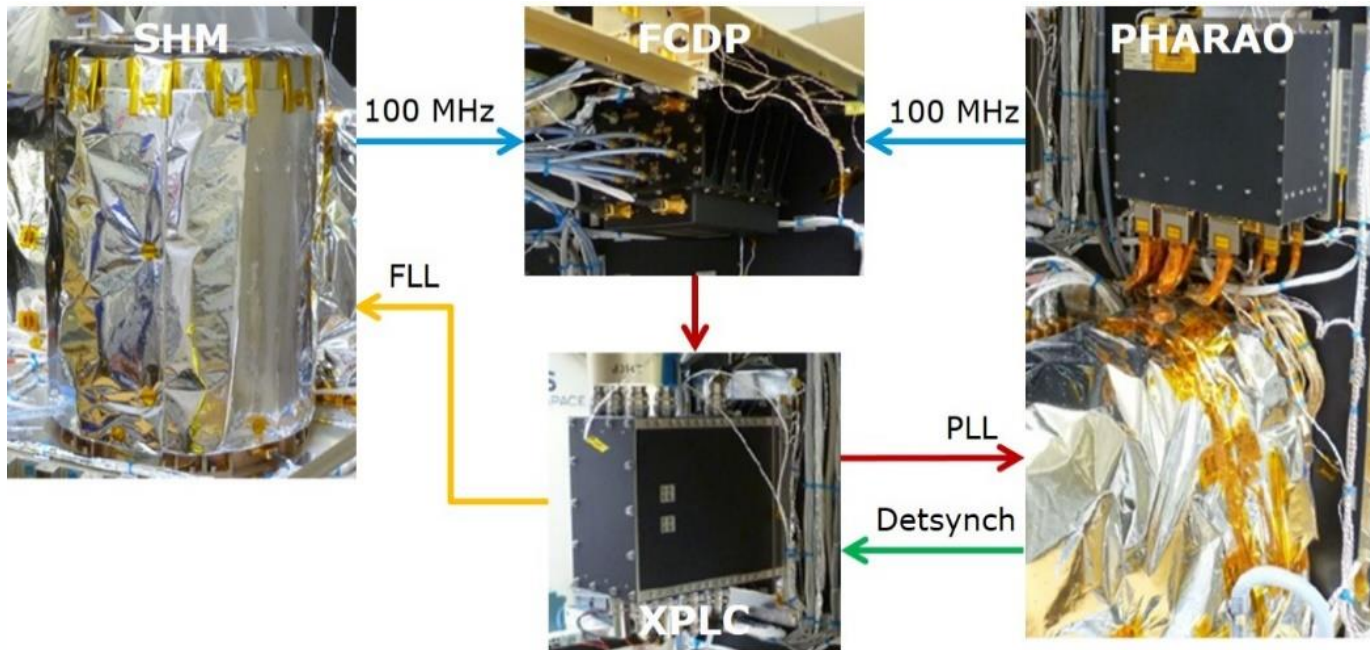
The ACES payload

- **PHARAO** (CNES): Atomic clock based on laser cooled Cs atoms
- **SHM**: Active hydrogen maser
- **FCDP**: Clocks comparison and distribution
- **MWL** : T&F transfer link
- **GNSS** receiver
- **ELT**: Optical link
- **Support subsystems**
 - **XPLC**: External PL computer
 - **PDU**: Power distribution unit,
 - Mechanical, thermal subsystems
 - **CEPA**: Columbus External PL Adapter



Volume: 1172x867x1246 mm³
Mass: 240 kg (w/o CEPA)
Power: 600 W

The ACES clock signal

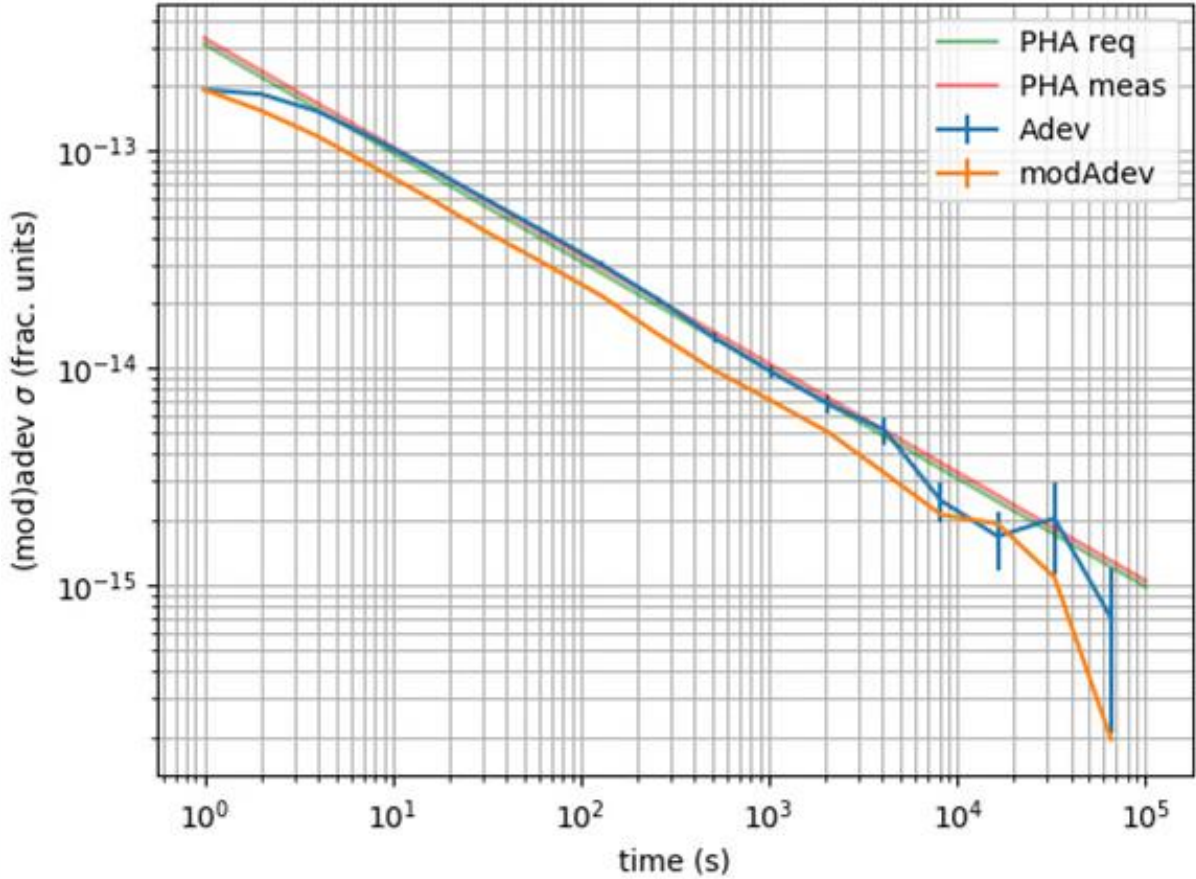


- The ACES clock signal (red curve) combines the good short-to-medium term frequency stability of the active H-maser SHM with the long-term stability and accuracy of the Cs clock PHARAO.
- Two servo-loops control the clocks with two different time constants: the short-term servo-loop (PLL) steers PHARAO local oscillator on SHM (1-2 s time constant); the long-term servo-loop (FLL) corrects SHM against long term drifts (100-500 s time constant).

PHARAO and ACES stability

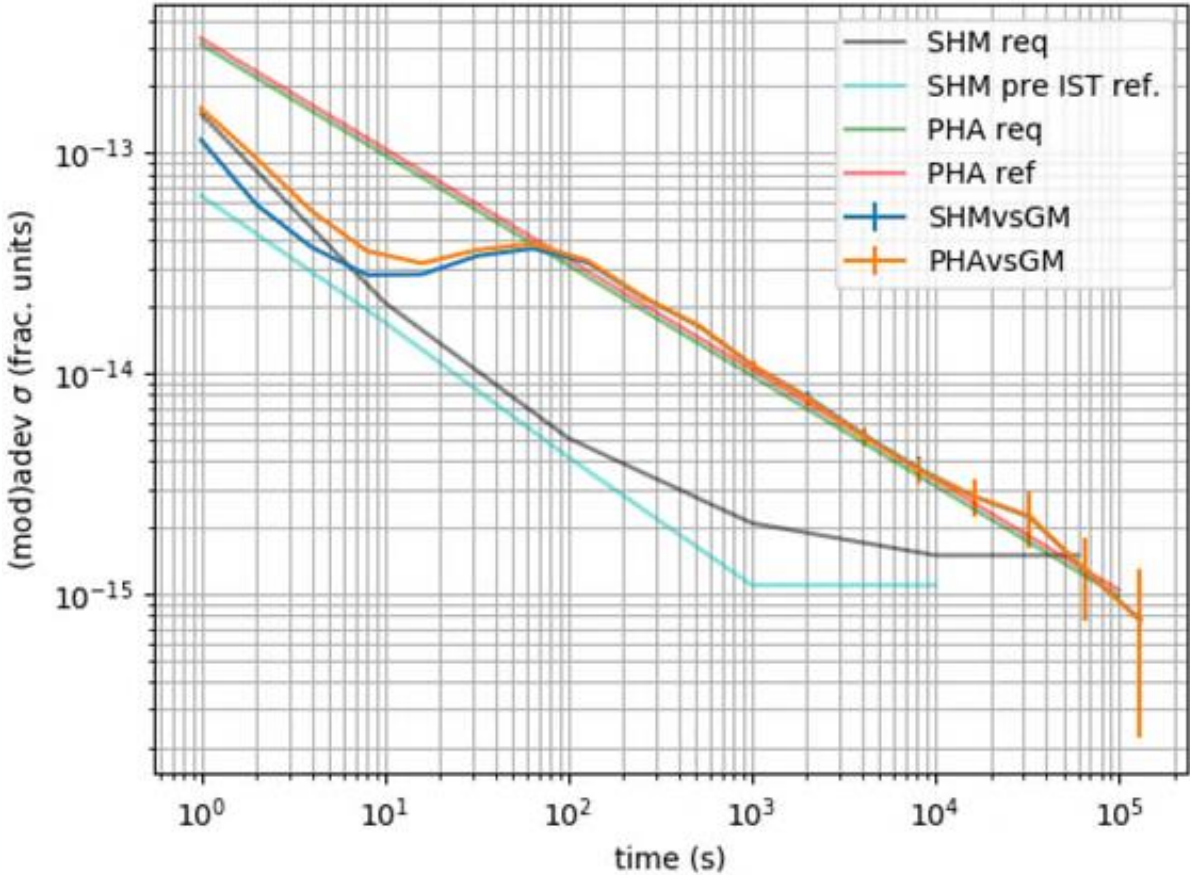


PHARAO_stand_alone
GM vs. PHA



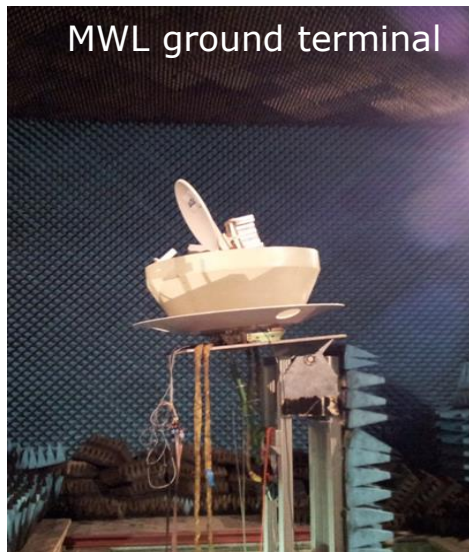
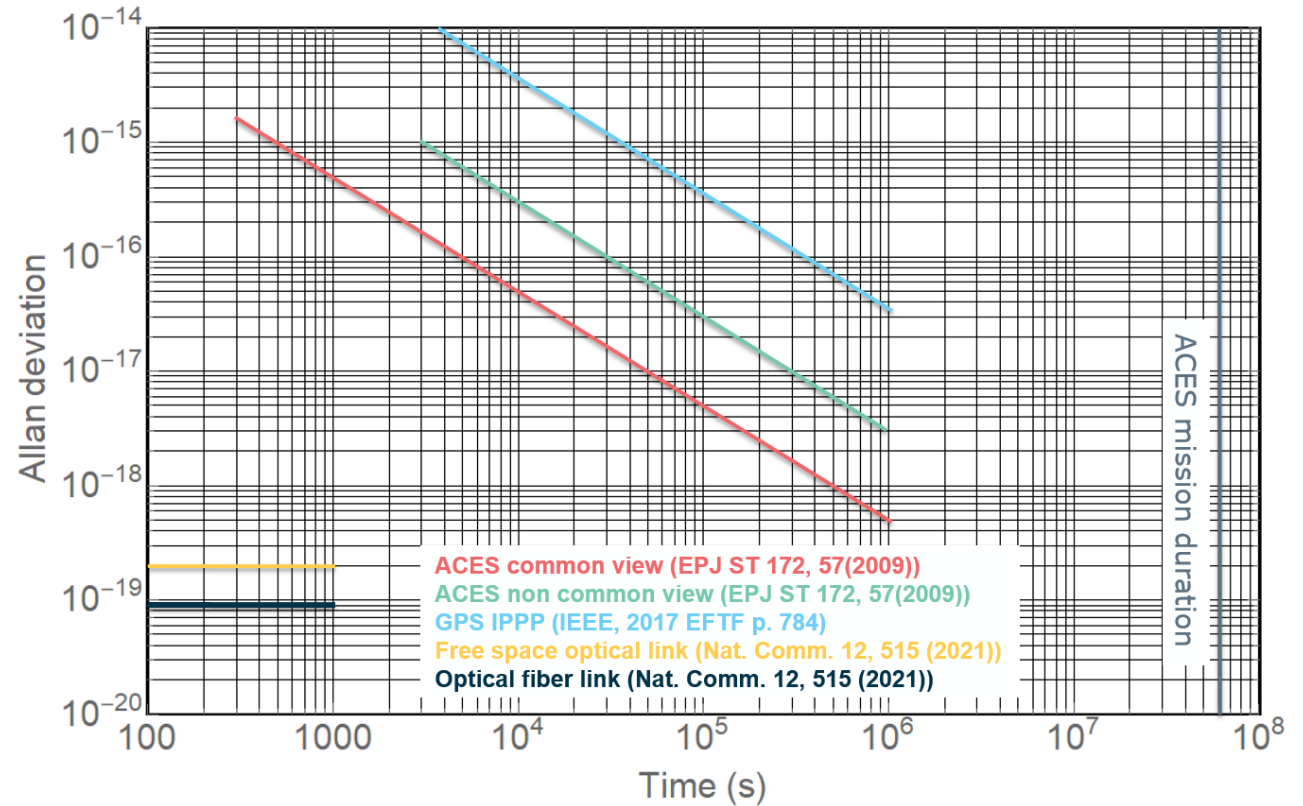
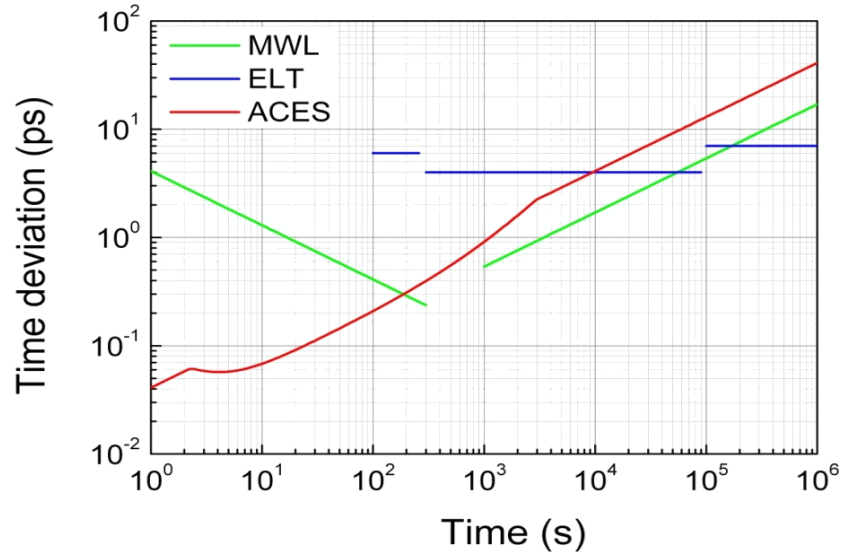
PHARAO stability in autonomous mode

Orb_Osc
SHM vs. GM



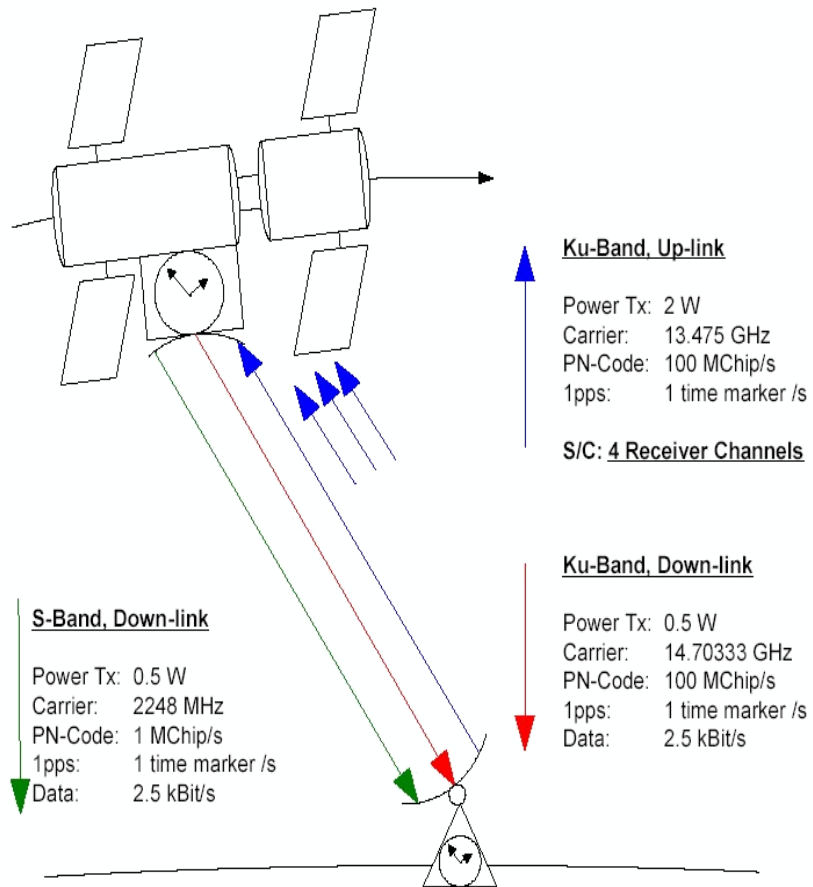
ACES stability with STSL and LTSL closed and in the presence of magnetic and thermal perturbations





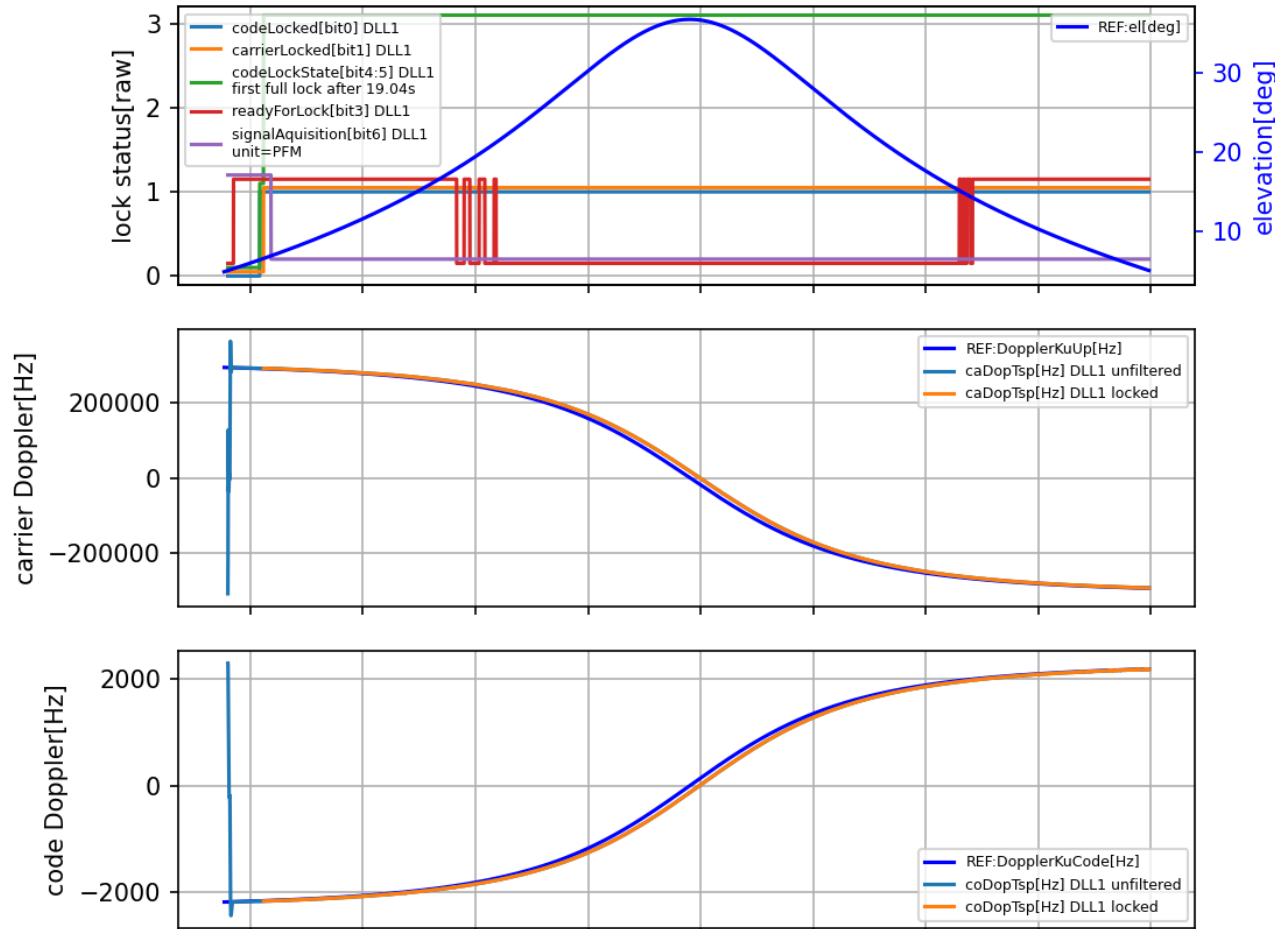
Ground clock comparisons with MWL

- **Common view:** $1.4 \cdot 10^{-15}$ after 300 s (ISS pass duration) and $1 \cdot 10^{-17}$ in less than 1 day.
- **Non common view:** $1 \cdot 10^{-15}$ after 2700 s, $1 \cdot 10^{-17}$ in 3-4 days (calculated assuming a deadtime Δt of 2700 s).



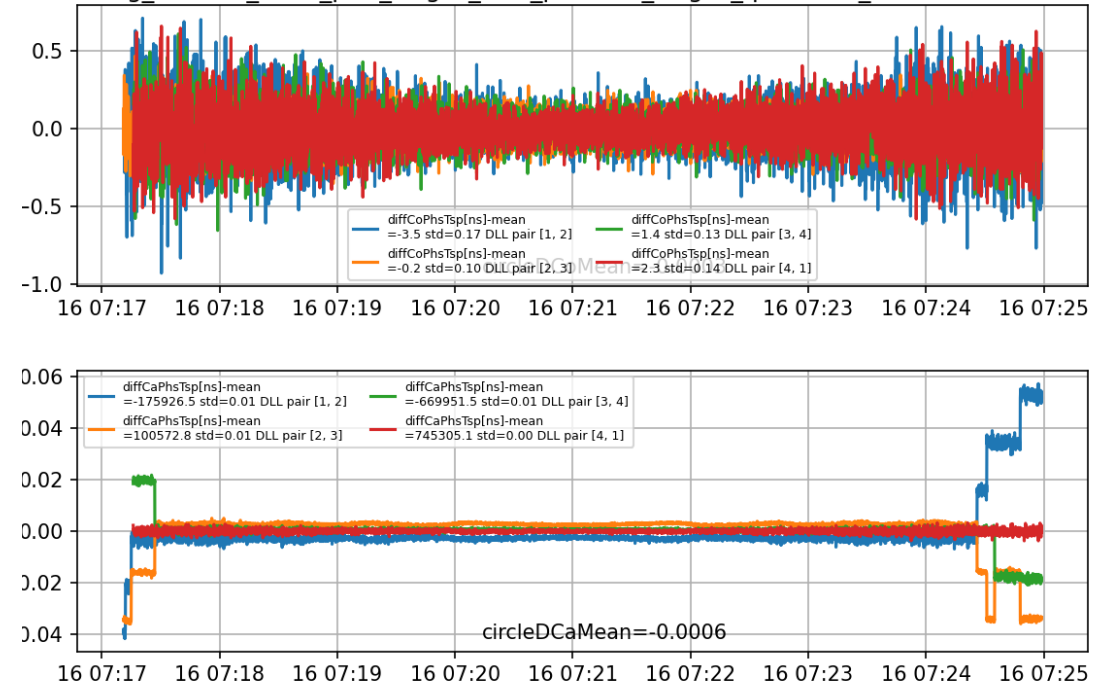
- Two-way link:
 - Removal of the troposphere time delay (8.3-103 ns)
 - Removal of 1st order Doppler effect
 - Removal of instrumental delays and common mode effects
- Additional down-link in the S-band:
 - Determination of the ionosphere TEC
 - Correction of the ionosphere time delay (0.3-40 ns in S-band, 6-810 ps in Ku-band)
- Phase PN code modulation: Removal of 2π phase ambiguity
- High chip rate (100 MChip/s) on the code:
 - Higher resolution
 - Multipath suppression
- Carrier and code phase measurements (80 ms sampling time)
- Data link: 2 kBits/s on the S-band down-link to obtain clock comparison results in real time
- Up to 4 simultaneous space-to-ground clock comparisons

fig_230815_de2e_pfm_rfegse_PFM_pid2932_rfegse_quicklookB_DLL1



4 DLLs simultaneously tracking the same signal

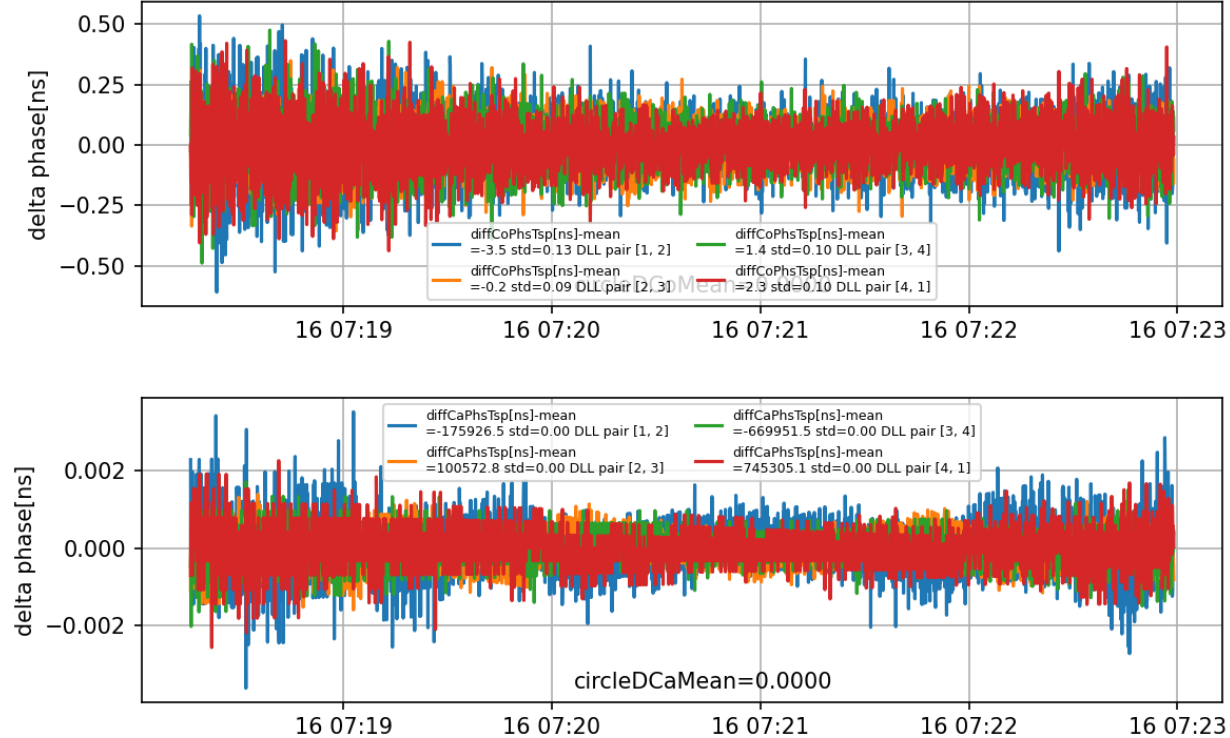
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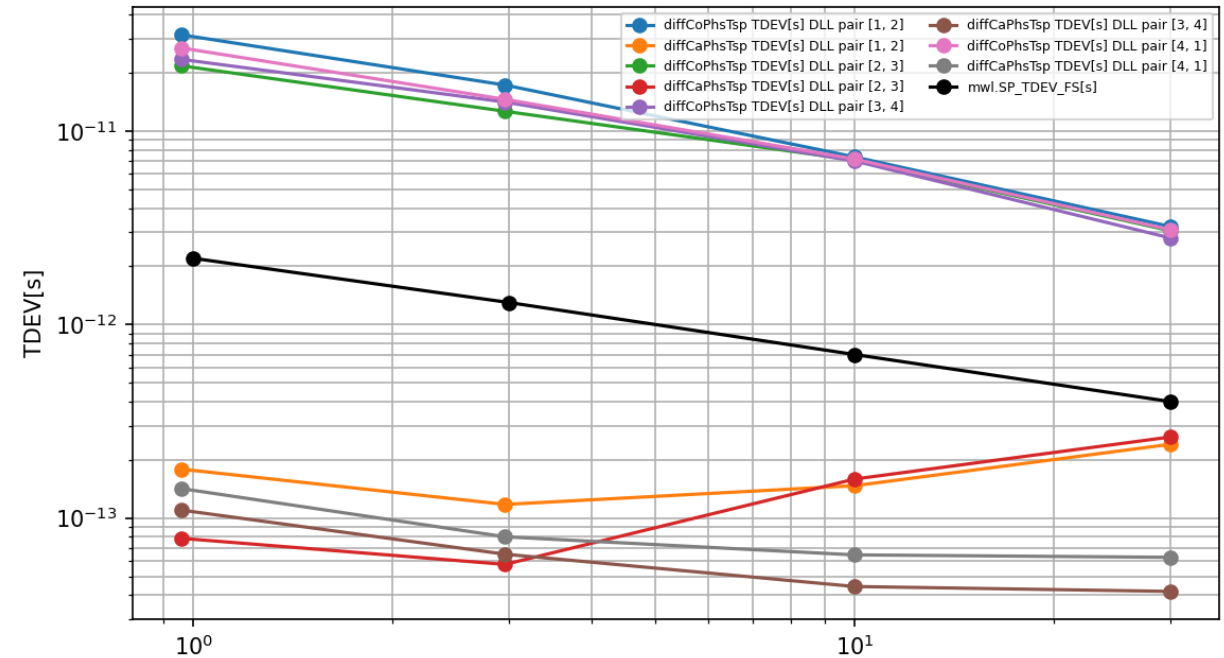
DLLx – DLLy code and carrier phase

MWL PFM performance – Signal dynamics

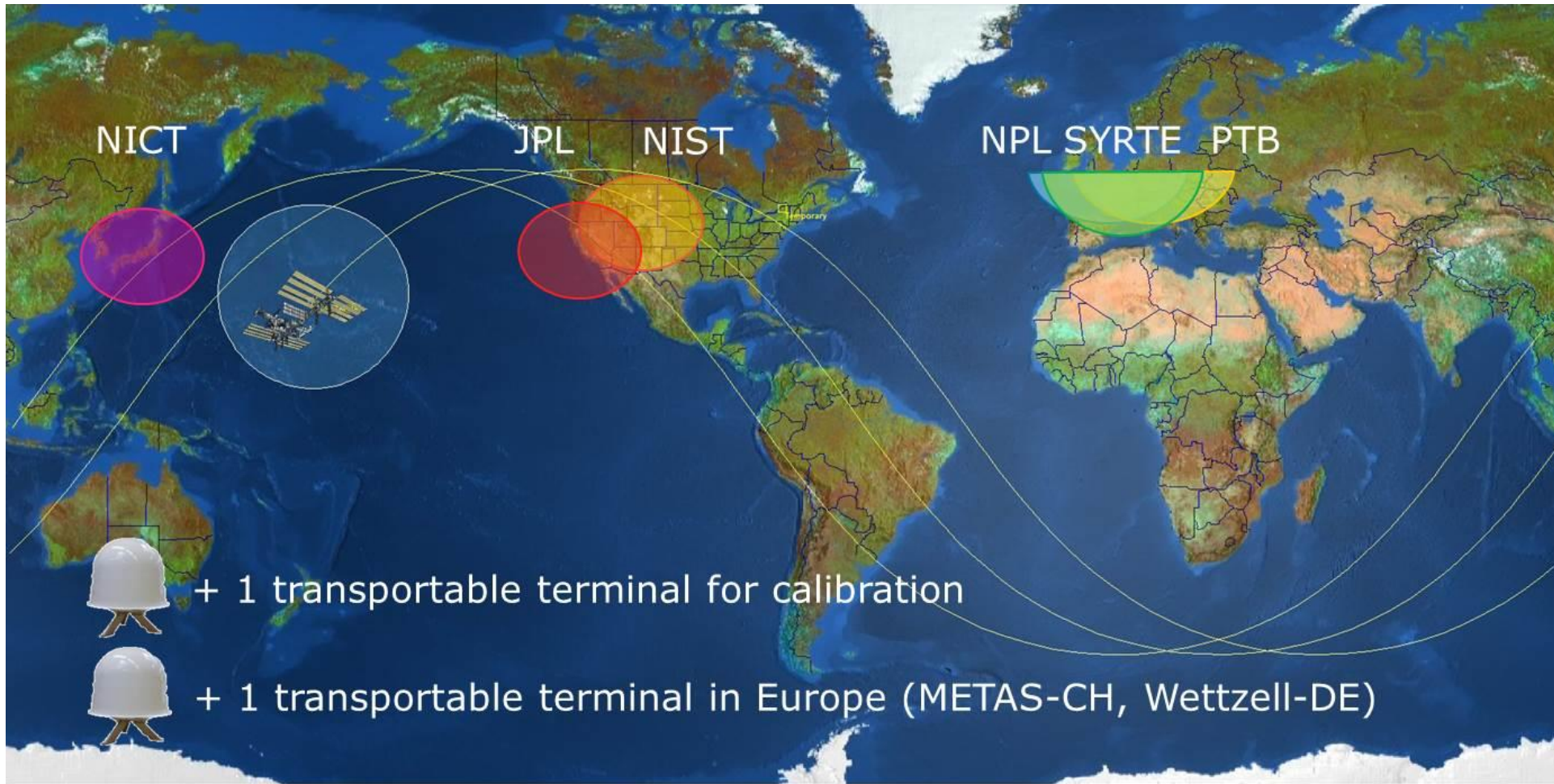
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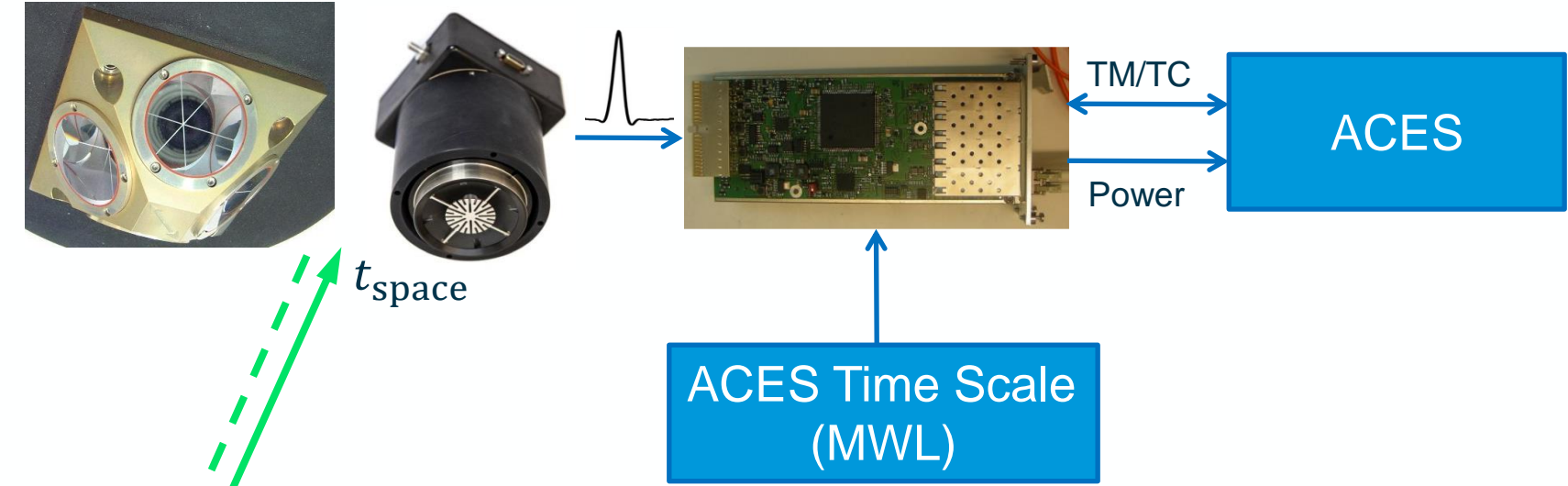
DLLx – DLLy code and carrier phase data after removing data containing carrier cycle slips



Time Deviation [s] of DLLx – DLLy code and carrier phase data

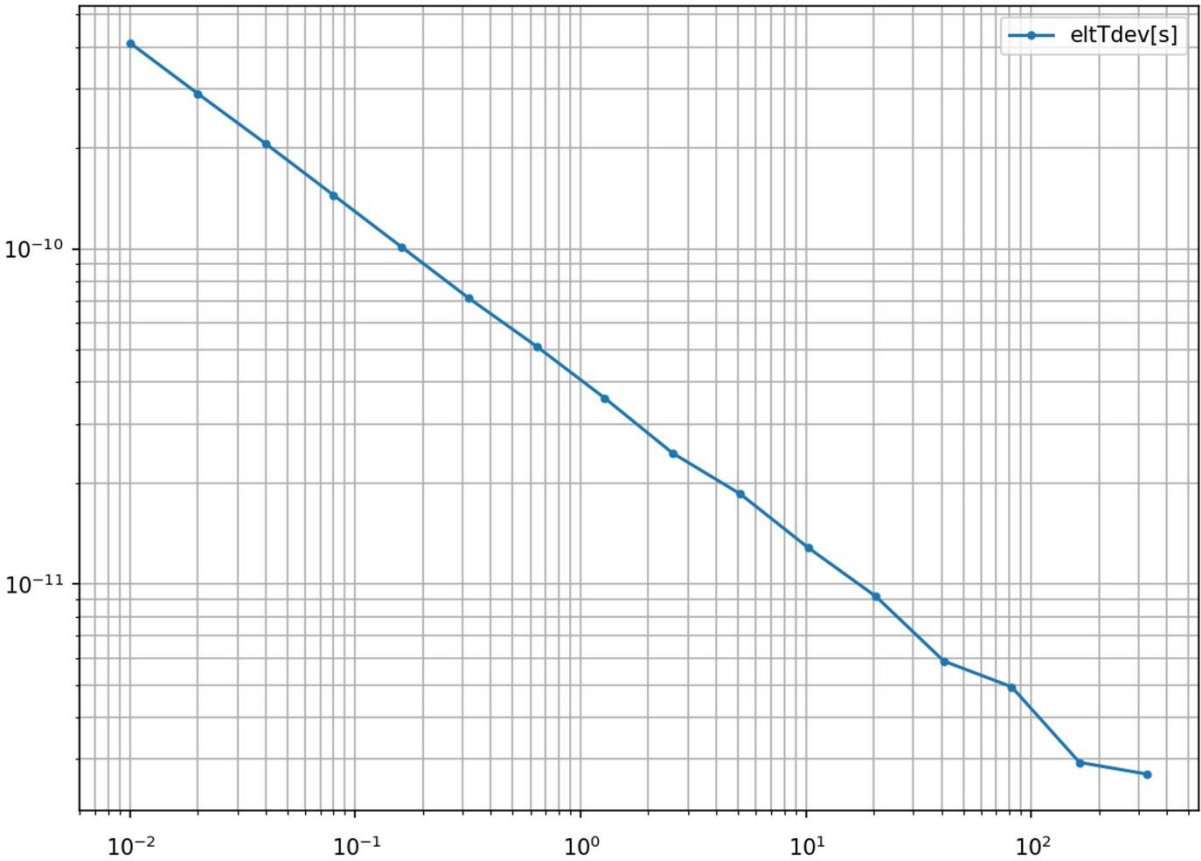


European Laser Timing (ELT)



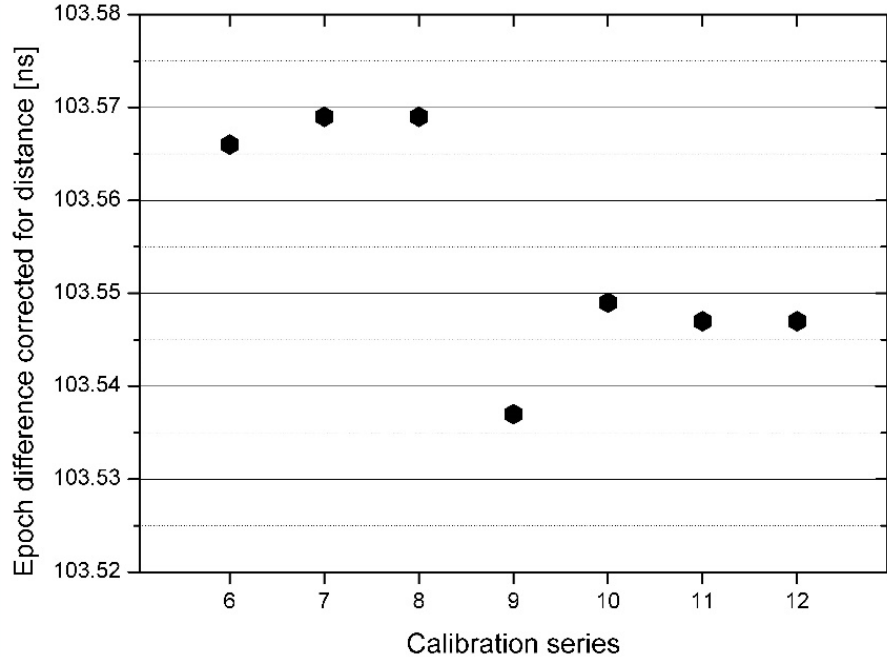
$$\tau_g(t) - \tau_s(t) = \frac{t_{\text{start}} + t_{\text{stop}}}{2} - t_{\text{space}} + \tau_{\text{relativity}} + \tau_{\text{atmosphere}} + \tau_{\text{geometry}}$$

Detector package
500 grams
0.6 Watt



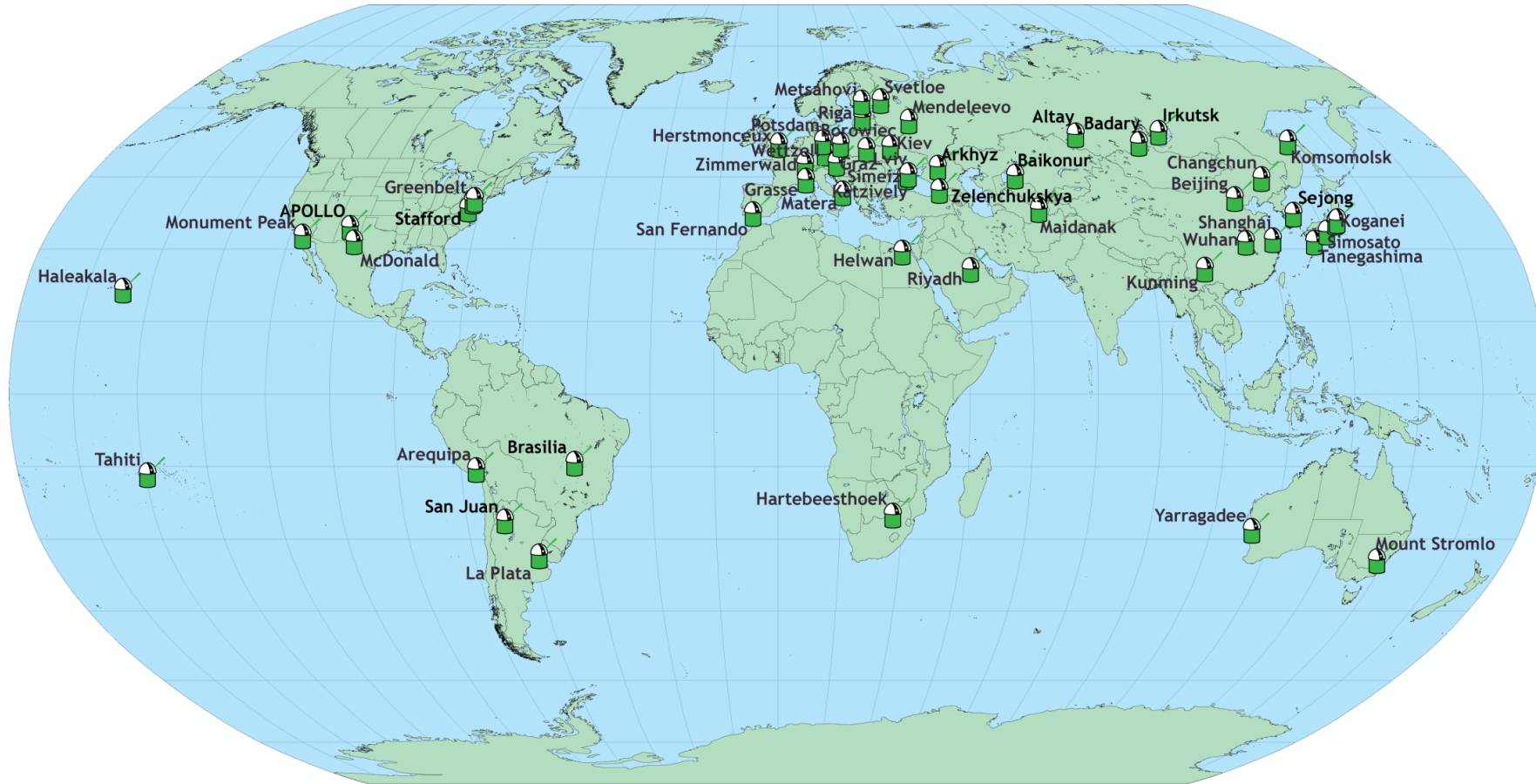
ELT Time Stability [s]: dark counts detection events vs NPET time tagging measurements at 100 Hz gate rate

ELT detector package tested from -55°C to +60°C: $K_T=1$ ps/K



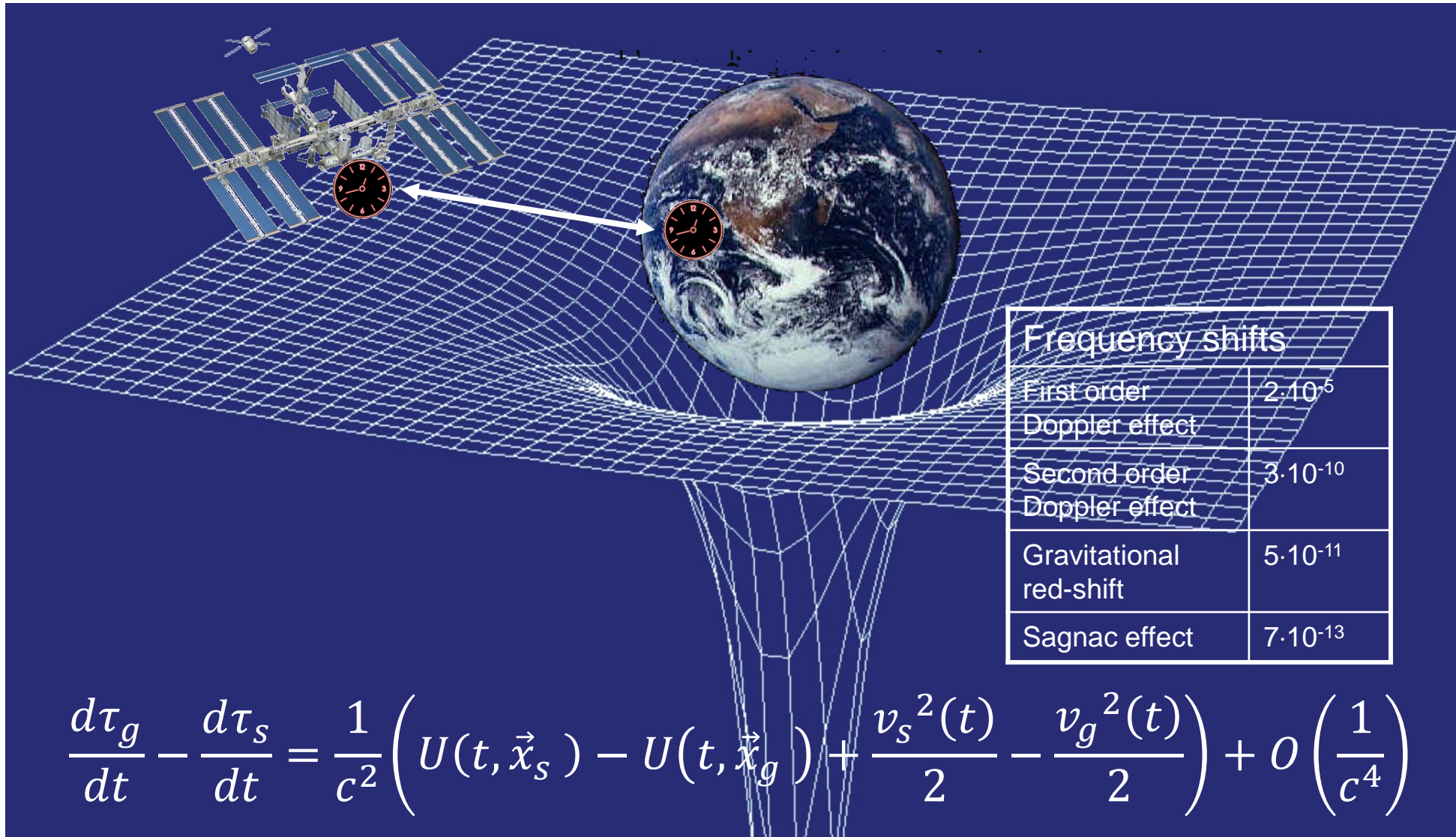
ELT Calibration Campaign at Graz SLR (1 yr)

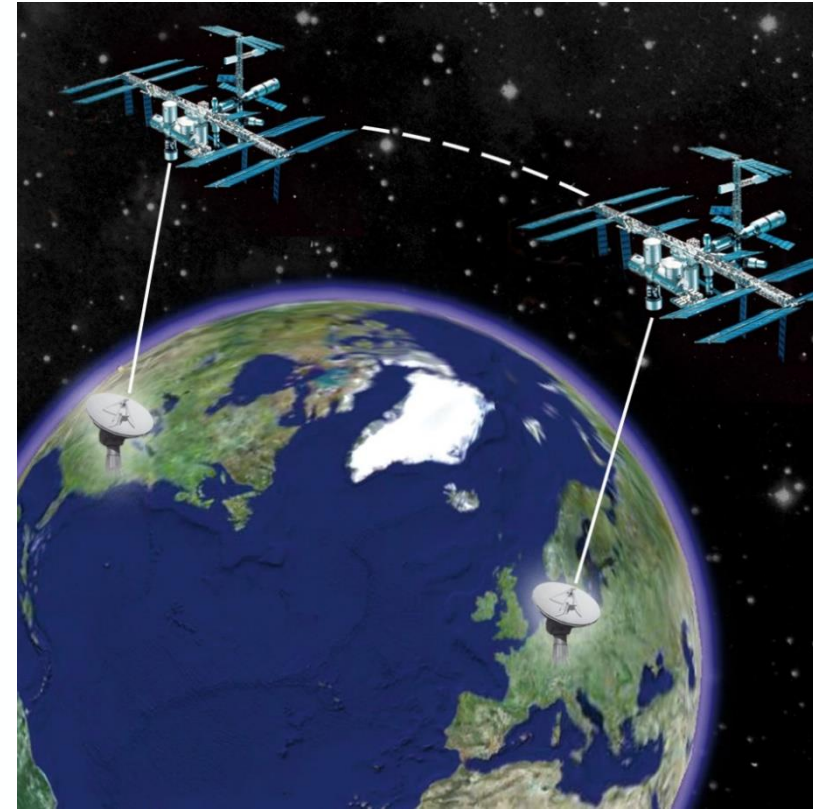
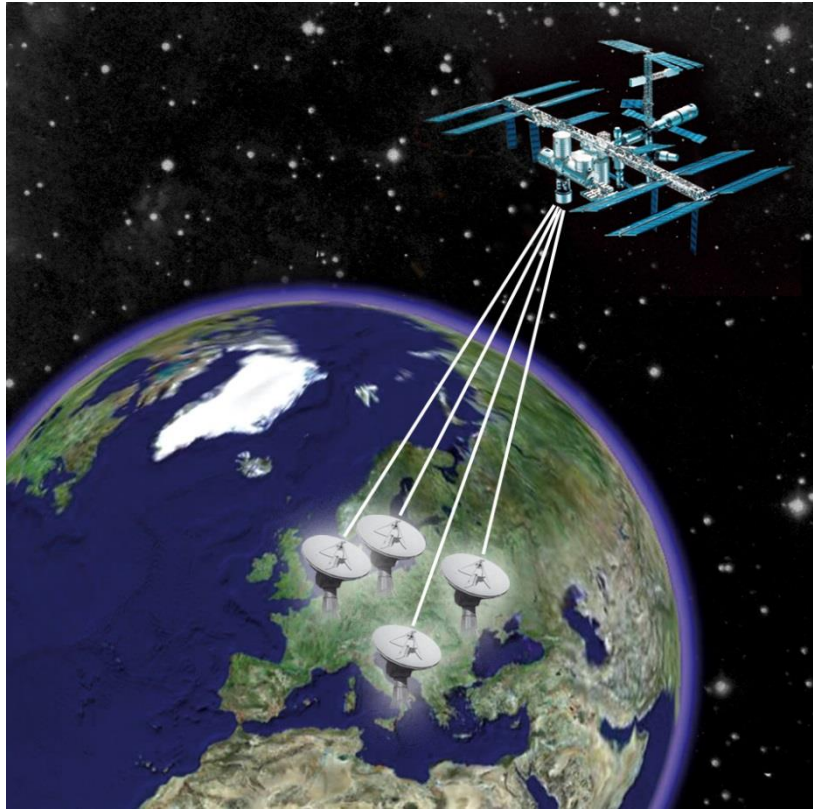
ILRS network of SLR stations



ACES as official ILRS target: **Wettzell** (primary station), **Graz**, **Herstmonceaux**, **Potsdam**, and **Zimmerwald** SLR stations already calibrated; other stations can join provided they comply with ISS safety requirements.

ACES Mission Objectives	ACES performances	Scientific background and recent results
Gravitational red-shift	Absolute measurement of the gravitational red-shift to $< 2 \cdot 10^{-6}$ after 10 days of integration time.	Factor 70 improvement over the GPA experiment and factor 10 over tests involving Galileo 5 and 6 satellites.
Time drifts of fundamental constants	Time variations of α constrained to $\alpha^{-1} \cdot d\alpha / dt < 3 \cdot 10^{-18} \text{ yr}^{-1}$ after 3 years of mission.	Comparisons of clocks based on different atoms and atomic transitions on a worldwide scale to constrain α , m_e/Λ_{QCD} and m_q/Λ_{QCD} .
Dark matter search with atomic clocks	Establish bounds on topological dark matter models based on the comparisons of clocks in the ACES network.	Comparisons via the ACES network imposing limits on the three coupling constants Λ_α , Λ_e , and Λ_q in the model Lagrangian. Measurements over an interval T between encounters of 20 d. Simultaneous observation with several clocks along different baselines providing ways to confirm any observation above the sensitivity threshold and control the measurement systematics. Screening effect on the dark matter field due to the Earth mass reduced to about 0.06 on the space clock PHARAO with respect to ground clocks ($\sim 10^{-7}$).





- ACES as a relay satellite to transfer time and frequency on a worldwide scale.
- Comparisons of ground clocks (microwave and optical) to better than $1 \cdot 10^{-17}$.

Frequency of hyperfine transitions: $\nu_{\text{hfs}}^{(i)} \simeq R_{\infty} c \times \mathcal{A}_{\text{hfs}}^{(i)} \times g^{(i)} \left(\frac{m_e}{m_p} \right) \alpha^2 F_{\text{hfs}}^{(i)}(\alpha)$

Frequency of electronic transitions: $\nu_{\text{elec}}^{(i)} \simeq R_{\infty} c \times \mathcal{A}_{\text{elec}}^{(i)} \times F_{\text{elec}}^{(i)}(\alpha)$

Ratios between atomic frequencies:

$$\frac{\nu_{\text{elec}}^{(ii)}}{\nu_{\text{elec}}^{(i)}} \propto \frac{F_{\text{elec}}^{(ii)}(\alpha)}{F_{\text{elec}}^{(i)}(\alpha)} \quad \frac{\nu_{\text{hfs}}^{(ii)}}{\nu_{\text{elec}}^{(i)}} \propto g^{(ii)} \frac{m_e}{m_p} \alpha^2 \frac{F_{\text{hfs}}^{(ii)}(\alpha)}{F_{\text{elec}}^{(i)}(\alpha)} \quad \frac{\nu_{\text{hfs}}^{(ii)}}{\nu_{\text{hfs}}^{(i)}} \propto \frac{g^{(ii)}}{g^{(i)}} \frac{F_{\text{hfs}}^{(ii)}(\alpha)}{F_{\text{hfs}}^{(i)}(\alpha)}$$

$g^{(i)}$ and m_p depend on the QCD mass scale Λ_{QCD} and $m_q = (m_u + m_d)/2$ thus

$$\delta \ln \left(\frac{\nu^{(i)}}{R_{\infty} c} \right) \simeq K_{\alpha}^{(i)} \times \frac{\delta \alpha}{\alpha} + K_q^{(i)} \times \frac{\delta(m_q/\Lambda_{\text{QCD}})}{(m_q/\Lambda_{\text{QCD}})} + K_e^{(i)} \times \frac{\delta(m_e/\Lambda_{\text{QCD}})}{(m_e/\Lambda_{\text{QCD}})}$$

	$K_{\alpha}^{(i)} \neq 0, K_q^{(i)} \neq 0, K_e^{(i)} \simeq 1$	hyperfine transitions
	$K_{\alpha}^{(i)} \neq 0, K_q^{(i)} \simeq 0, K_e^{(i)} \simeq 0$	electronic transitions
	$K_{\alpha}^{(i)} \simeq 0, K_q^{(i)} \simeq 0, K_e^{(i)} \simeq 1/2$	vibrational transitions

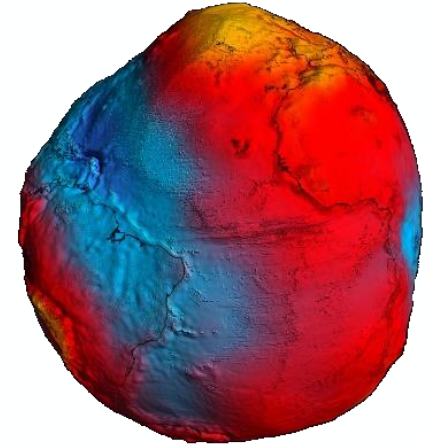
V.V. Flambaum, Laser Spectroscopy, pp. 49-57 (2004)

V.V. Flambaum *et al.*, Phys. Rev. D **69**, 115006 (2004)

- Atomic clock networks can be used to search for dark matter and place boundaries on certain models, e.g. topological dark matter (TDM).
- TDM is represented by a scalar field that couples to fundamental constants thus inducing fluctuations of atomic transition frequencies.
- The ACES network can ensure comparisons of atomic clocks based on different atomic transitions down to $1 \cdot 10^{-17}$, in the optical domain, in the microwave domain, and optical vs microwave.
- Limits on variations for each of the three fundamental constants can be established thus testing different terms in the model Lagrangian and imposing limits on the three energy scales Λ_α , Λ_e , and Λ_q .
- The simultaneous observation with several clocks compared along different baselines will provide ways to confirm any observation above the sensitivity threshold and control the measurement systematics.
- Clock comparisons can be performed continuously on 20 d intervals thanks to the ACES MWL, extending the time T between encounters.
- Screening effect of the dark matter field due to the Earth mass reduced to about 0.06 on the space clock PHARAO with respect to ground clocks ($\sim 10^{-7}$).

Relativistic geodesy

- Relativistic geodesy: mapping of the Earth gravitational potential based on the red-shift measured between two clocks at two different locations.
- ACES intercontinental comparisons of optical clocks at the 10^{-17} level after 4 days, corresponding to a resolution on the local height above the geoid at the 10 cm level.
- The global coverage offered by ACES will complement the results of the CHAMP, GRACE, and GOCE missions.

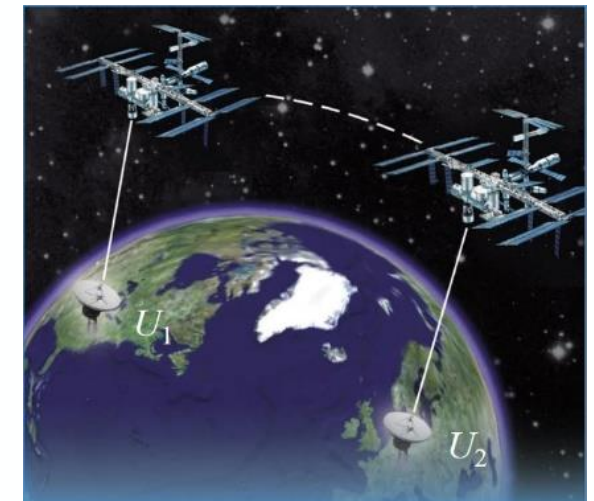


Clocks synchronization

- MWL and ELT clocks synchronization at the 100 ps and 50 ps level, respectively.

Atomic time scales (TAI)

- The PHARAO clock is accurate to $1-2 \cdot 10^{-16}$.
- MWL will provide means to compare atomic clocks on a worldwide scale:
 - PHARAO and primary standards on ground contributing to TAI.
 - Optical clock comparisons to $1 \cdot 10^{-17}$ will help SI second redefinition.



- A first run of the ACES Integrated System Tests (IST) tests has been completed. They included:
 - ACES servo-loops performance
 - ACES clock signal performance evaluation
 - Sensitivity under temperature and magnetic field variations
- IST revealed a major anomaly: SHM getters had reached end-of-life and needed replacement.
- Acceptance status of ACES:
 - PHARAO clock PFM, ELT detector package, ACES GNSS system and FCDP PFM are integrated.
 - SHM PFM getters have been replaced and acceptance tests completed. SHM is now integrated in the ACES payload.
 - MWL PFM is completing qualification tests before final acceptance (by end 2023).
 - ACES payload performance tests planned for summer 2024.
- ACES PFM delivered for launch on SpaceX in end 2024 – beginning 2025.
- MWL GTs deployment will start in mid 2024 (SYRTE, PTB and Wettzell); the other institutes will follow.

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