Neutrino physics (Theory)

Yu Seon Jeong (Chung-Ang University)



Outline

- Neutrino oscillation
- Neutrino physics at the LHC
 - Forward experiments
 - Potential for neutrino study

Discovery of neutrino oscillation

- Neutrino oscillation is one of great discoveries in the past three decades.
- Neutrino oscillation is the clear evidence for
 - Non-zero masses of neutrinos
 - Leptonic mixing

Neutrino oscillation in three flavor neutrinos

■ In three flavor neutrino framework: $|\nu_{\alpha}\rangle = \sum_{k} U_{\mu k} |\nu_{k}\rangle$ ($\alpha = e, \mu, \tau$)

$$\begin{pmatrix} \nu_e \\ \nu_{\mu} \\ \nu_{\tau} \end{pmatrix} = \underbrace{\begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu 1} & U_{\mu 2} & U_{\mu 3} \\ U_{\tau 1} & U_{\tau 2} & U_{\tau 3} \end{pmatrix}}_{k} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix} \qquad |\nu_{\alpha}(t)\rangle = \sum_{k} U_{\alpha k} e^{-iE_{k}t} |\nu_{k}\rangle$$

$$U_{\text{PMNS}} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & \cos\theta_{23} & \sin\theta_{23} \\ 0 & -\sin\theta_{23} & \cos\theta_{23} \end{pmatrix} \begin{pmatrix} \cos\theta_{13} & 0 & \sin\theta_{13}e^{i\delta} \\ 0 & 1 & 0 \\ -\sin\theta_{13}e^{-i\delta} & 0 & \cos\theta_{13} \end{pmatrix} \begin{pmatrix} \cos\theta_{12} & \sin\theta_{12} & 0 \\ -\sin\theta_{12} & \cos\theta_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

$$P_{\nu_{\alpha} \to \nu_{\beta}}(L, E) = |\langle \nu_{\beta} | \nu_{\alpha}(t) \rangle|^{2} = \sum_{k,j} U_{\alpha k} U_{\beta k}^{*} U_{\alpha j}^{*} U_{\beta j} \exp\left(-i \frac{\Delta m_{kj}^{2} L}{2E}\right)$$

Neutrino oscillation in three flavor neutrinos

■ Oscillation parameters:

$$\Delta m_{21}^2 = \Delta m_{sol}^2$$
, $\Delta m_{3l}^2 = \Delta m_{atm}^2$

$$\theta_{12}=\theta_{sol}$$
 , $\theta_{13}=\theta_{reactor}$, $\theta_{23}=\theta_{atm}$

 \circ CP phase: δ

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Solar, atmospheric, reactor, accelerator experiments have provided enormous data to determine the parameters.

Parameter	Main contribution from	Other contributions from		
Δm^2_{21}	KamLAND	\mathbf{SOL}		
$ \Delta m^2_{31} $	LBL+ATM+REAC	_		
θ_{12}	SOL	KamLAND		
θ_{23}	LBL+ATM	_		
θ_{13}	REAC	(LBL+ATM) and (SOL+KamLAND)		
δ	$_{ m LBL}$	\mathbf{ATM}		
MO	(LBL+REAC) and ATM	COSMO and $0\nu\beta\beta$		

Current status of oscillation parameters

NuFIT 5.1 (2021)

	Normal Ord	dering (best fit)	Inverted Ordering ($\Delta \chi^2 = 7.0$)		
	bfp $\pm 1\sigma$ 3 σ range		bfp $\pm 1\sigma$	3σ range	
$\sin^2 \theta_{12}$	$0.304^{+0.012}_{-0.012}$	$0.269 \rightarrow 0.343$	$0.304^{+0.013}_{-0.012}$	$0.269 \rightarrow 0.343$	
$\theta_{12}/^{\circ}$	$33.45^{+0.77}_{-0.75}$	$31.27 \rightarrow 35.87$	$33.45^{+0.78}_{-0.75}$	$31.27 \rightarrow 35.87$	
$\sin^2 \theta_{23}$	$0.450^{+0.019}_{-0.016}$	$0.408 \rightarrow 0.603$	$0.570^{+0.016}_{-0.022}$	$0.410 \rightarrow 0.613$	
$ heta_{23}/^{\circ}$	$42.1_{-0.9}^{+1.1}$	$39.7 \rightarrow 50.9$	$49.0^{+0.9}_{-1.3}$	$39.8 \rightarrow 51.6$	
$\sin^2 heta_{13}$	$0.02246^{+0.00062}_{-0.00062}$	0.02060 o 0.02435	$0.02241^{+0.00074}_{-0.00062}$	$0.02055 \rightarrow 0.02457$	
$ heta_{13}/^{\circ}$	$8.62^{+0.12}_{-0.12}$	$8.25 \rightarrow 8.98$	$8.61^{+0.14}_{-0.12}$	$8.24 \rightarrow 9.02$	
$\delta_{\mathrm{CP}}/^{\mathrm{o}}$	230^{+36}_{-25}	$144 \rightarrow 350$	278^{+22}_{-30}	$194 \rightarrow 345$	
$\frac{\Delta m^2_{21}}{10^{-5}~{\rm eV^2}}$	$7.42^{+0.21}_{-0.20}$	6.82 o 8.04	$7.42^{+0.21}_{-0.20}$	6.82 o 8.04	
$\frac{\Delta m_{3\ell}^2}{10^{-3} \text{ eV}^2}$	$+2.510^{+0.027}_{-0.027}$	$+2.430 \to +2.593$	$-2.490^{+0.026}_{-0.028}$	-2.574 ightarrow -2.410	

Current status of oscillation parameters (well measured)

NuFIT 5.1 (2021)

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		Normal Ordering (best fit)		Inverted Ordering ($\Delta \chi^2 = 7.0$)		
		bfp $\pm 1\sigma$	3σ range	bfp $\pm 1\sigma$	3σ range	
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	$\theta_{13}/^{\circ}$	$8.62^{+0.12}_{-0.12}$	$8.25 \rightarrow 8.98$	$8.61^{+0.14}_{-0.12}$	$8.24 \rightarrow 9.02$	$\theta_{13}:9.0\%$
	$\delta_{\mathrm{CP}}/^{\circ}$	230^{+36}_{-25}	$144 \rightarrow 350$	278^{+22}_{-30}	$194 \rightarrow 345$	
	$\frac{\Delta m_{21}^2}{10^{-5} \text{ eV}^2}$	$7.42^{+0.21}_{-0.20}$	6.82 o 8.04	$7.42^{+0.21}_{-0.20}$	$6.82 \rightarrow 8.04$	$\Delta m^2_{21}:16\%$
	$\frac{\Delta m_{3\ell}^2}{10^{-3} \text{ eV}^2}$	$+2.510^{+0.027}_{-0.027}$	$+2.430 \to +2.593$	$-2.490^{+0.026}_{-0.028}$	$-2.574 \rightarrow -2.410$	$ \Delta m^2_{3\ell} :6.7\%[6.5]$

JHEP 09 (2020) 178 [arXiv:2007.14792], NuFIT 5.1 (2021), www.nu-fit.org.

What still need to be determined

NuFIT 5.1 (2021)

	Normal Ord	dering (best fit)	Inverted Ordering ($\Delta \chi^2 = 7.0$)		
	bfp $\pm 1\sigma$ 3 σ range		bfp $\pm 1\sigma$	3σ range	
$\sin^2 heta_{12}$	$0.304^{+0.012}_{-0.012}$	$0.269 \rightarrow 0.343$	$0.304^{+0.013}_{-0.012}$	$0.269 \rightarrow 0.343$	
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 $lacktriang{lacktriang{\mbox{\mbox{\bf Mass ordering:}}} \lacktriang{\mbox{\bf the sign of }} \Delta m_{3l}^2$

lacksquare Octant of θ_{23}

$$\bullet \delta_{CP} = ?$$

Current accelerator experiments for oscillation

Long baseline (LBL) accelerator experiments



■ NOvA

NuMI beam@Fermilab from Ep = 120 GeV.

Peak energy: 1.8 GeV

Baseline: 810km

■T2K (Tokai to Kamioka)

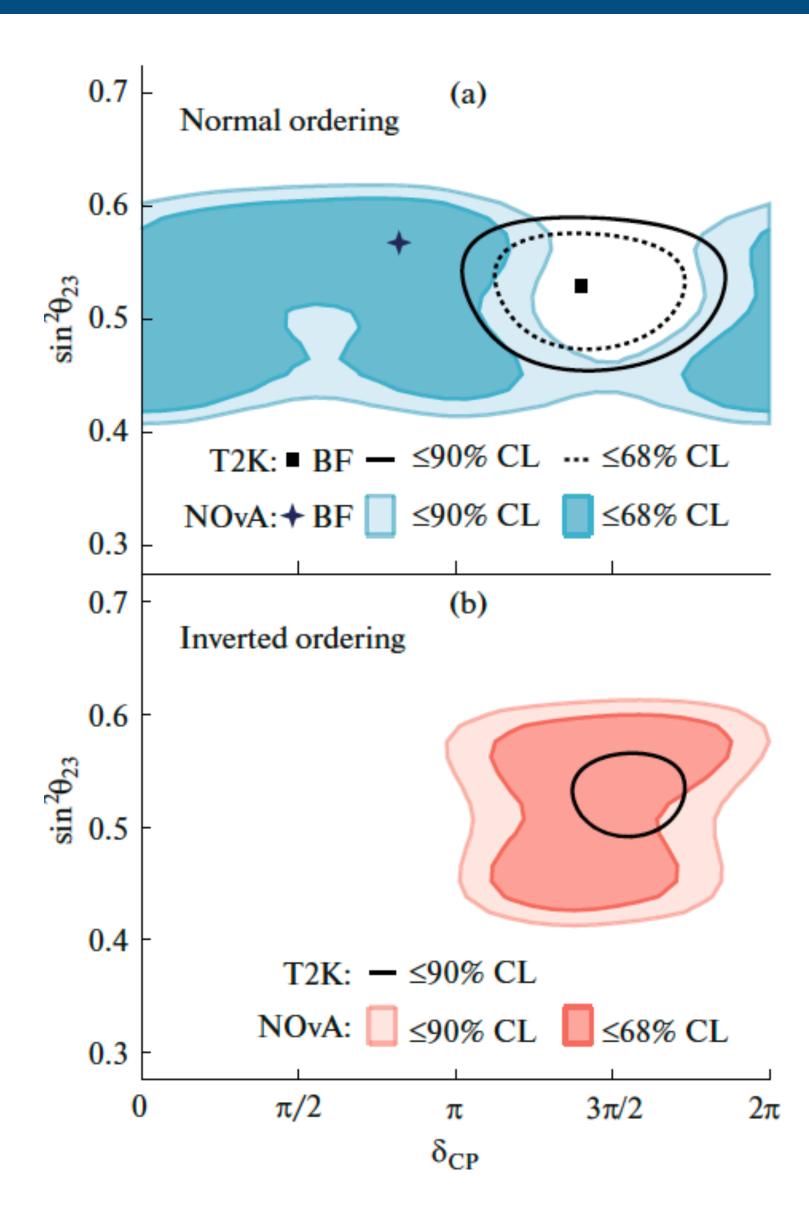
 $^{\circ}$ ν beam is from the 30 GeV proton beam at J-PARC.

Peak energy: 0.6 GeV

Baseline: 295km

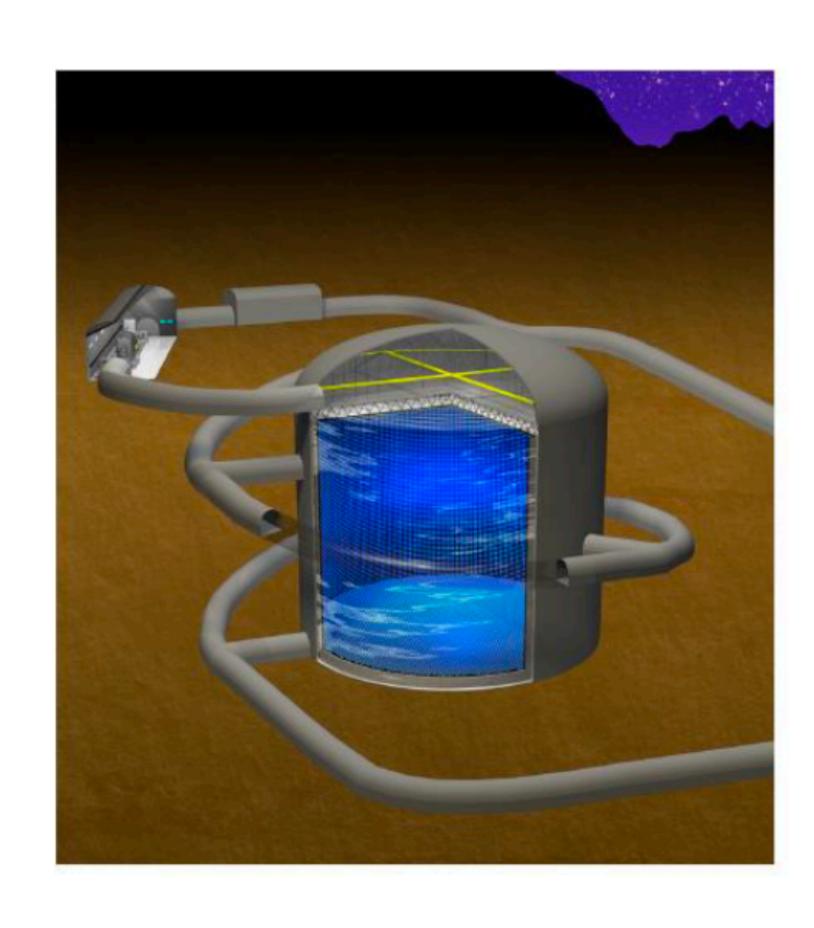


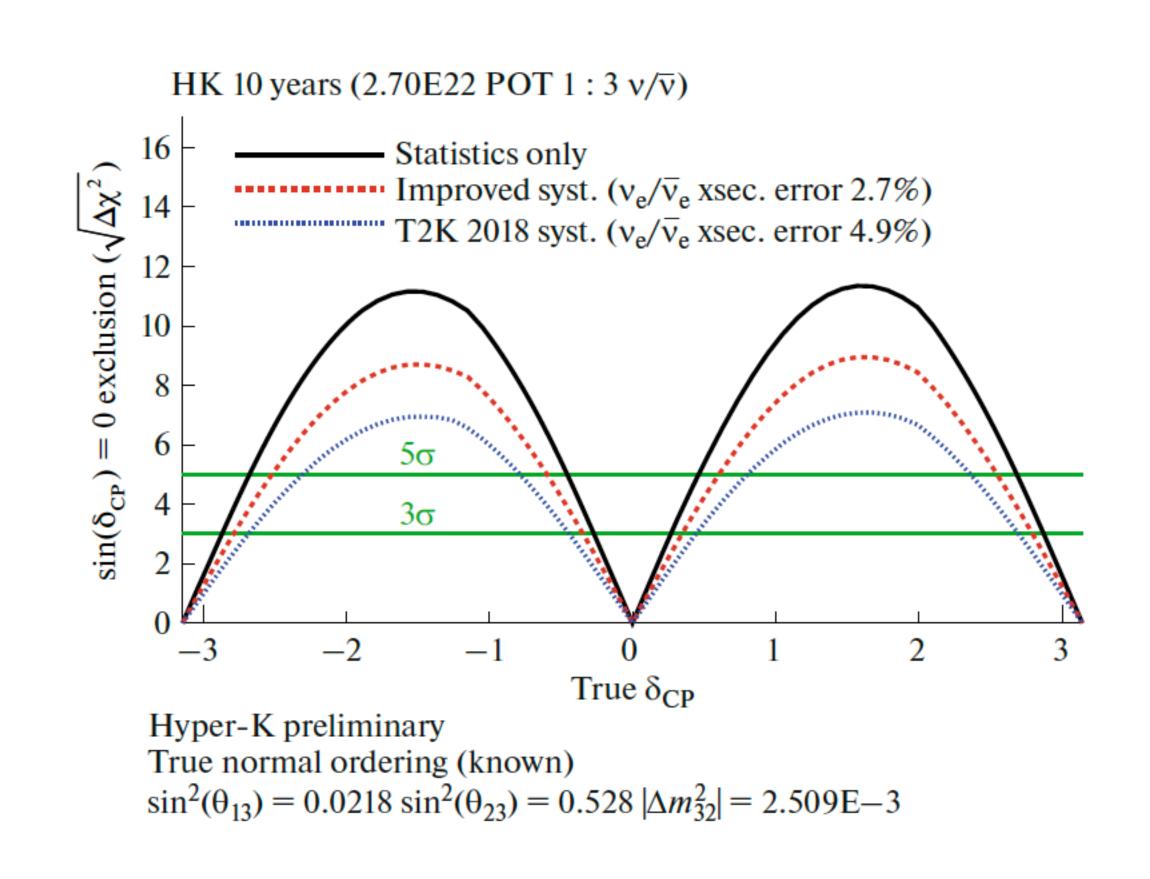
Tension between NOvA vs T2K



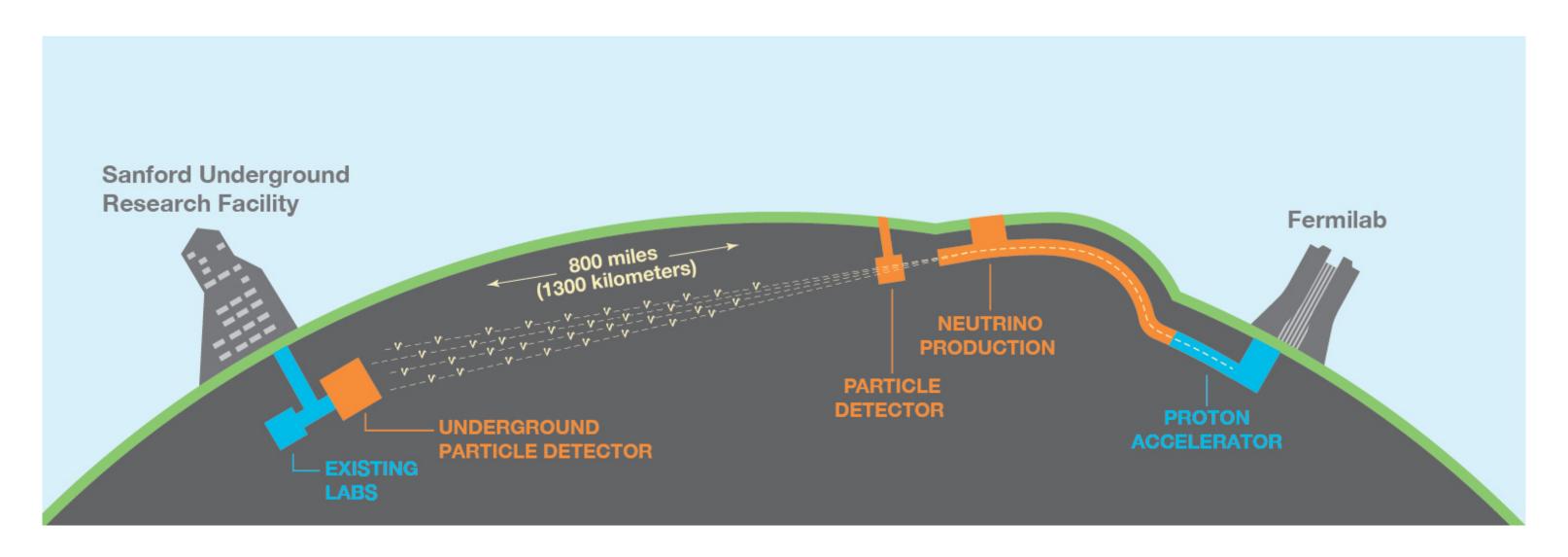
- For mass ordering, both NOvA and T2K prefers normal ordering.
- T2K and NOvA results coincide with the inverted ordering.

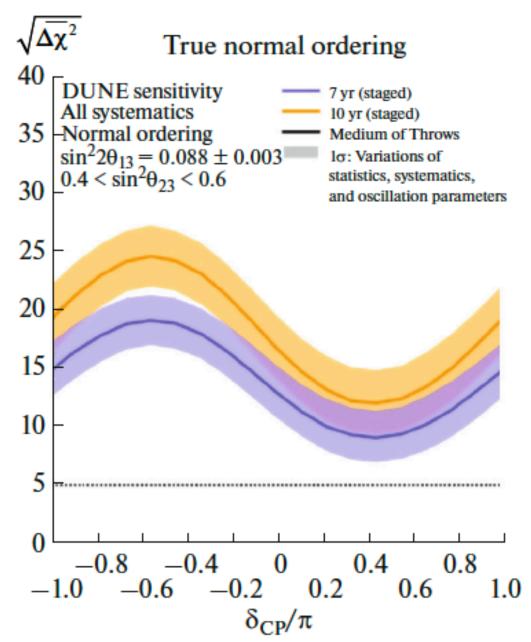
Future accelerator experiments for oscillation — HK

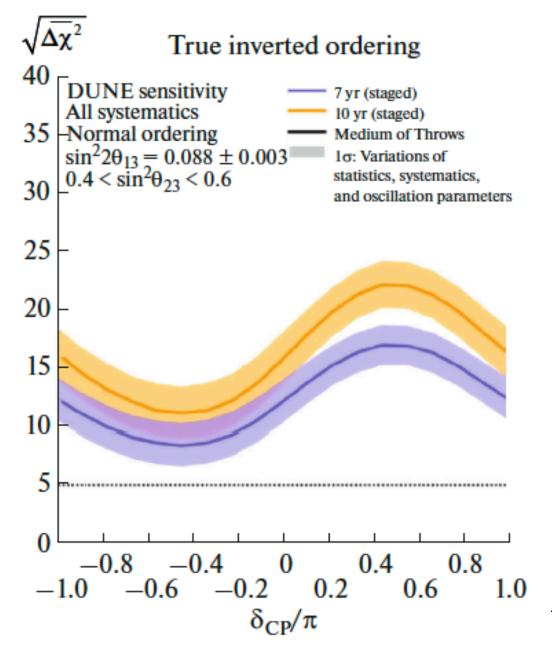




Future accelerator experiments for oscillation — DUNE







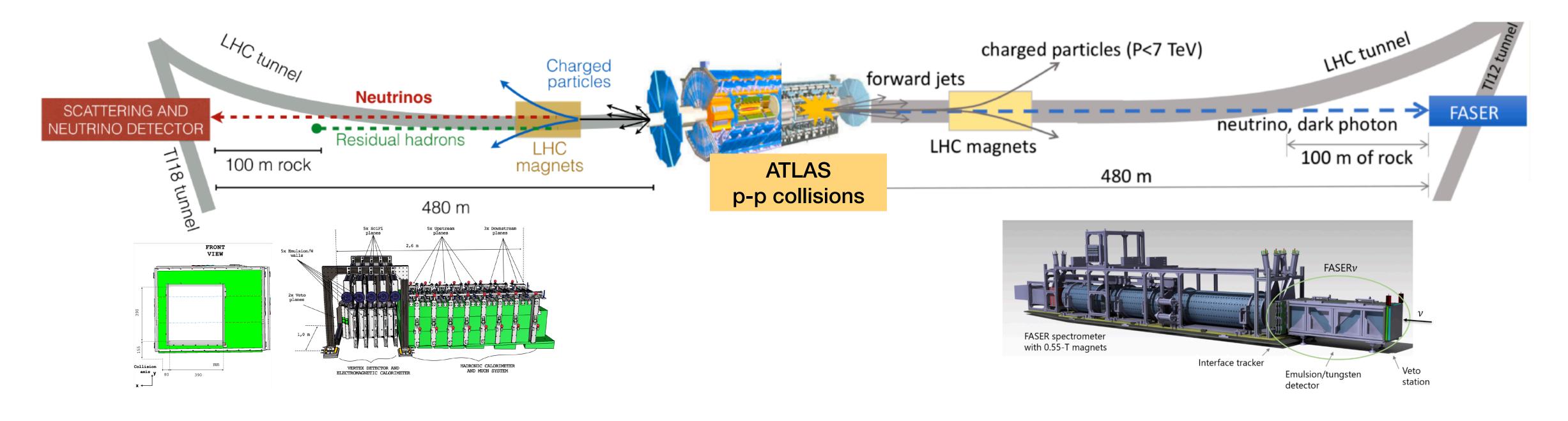
 $\begin{tabular}{ll} \hline & Sensitivity on mass hierarchy \\ & After 7 years running, mass hierarchy can be \\ & determined for all δ_{CP} values. \\ \end{tabular}$

Summary I

- The oscillation data are well described by the three flavor neutrino framework.
- lacktriangle There are still ambiguity in some parameters in mass hierarchy, $heta_{23}$, δ_{CP} .
- Current and future oscillation experiments will provide information for such parameters.

Neutrino physics at the LHC

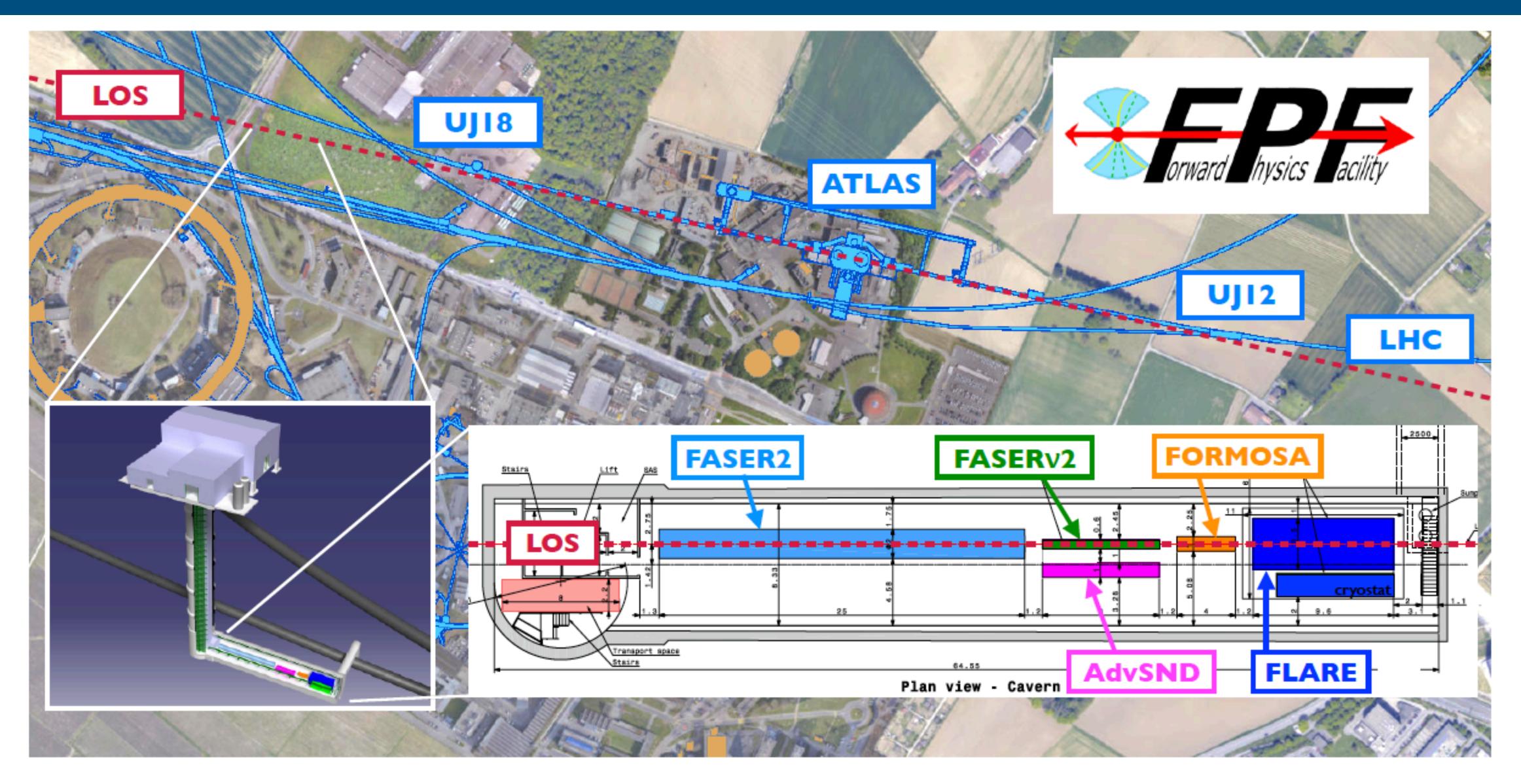
Forward experiments during the Run 3 - SND@LHC



- Rapidity: $7.2 < \eta < 8.4$ (off-axis)
- Target material: Tungsten (w/ emulsion film)
- Target mass: 830 kg
- surface: 39 x 39 cm²

- Rapidity: $\eta \gtrsim 8.5$ (on-axis)
- Target material: Tungsten (w/ emulsion film)
- Target mass: 1.1 tons
- surface: $25 \times 30 \text{ cm}^2$ (1.1 m long)

Forward Physics Facility (FPF, High-Luminosity)



Ref: arXiv: 2203.05090

Forward Physics Facility

Ref: arXiv: 2203.05090

Detector				Number of CC Interactions		
Name	Mass	Coverage	Luminosity	$ u_e + \bar{\nu}_e $	$ u_{\mu} + \bar{\nu}_{\mu} $	$\nu_{ au} + \bar{\nu}_{ au}$
$\overline{\mathrm{FASER} u}$	1 ton	$\eta \gtrsim 8.5$	290 fb ⁻¹	901 / 3.4k	4.7k / 7.1k	15 / 97
SND@LHC	800kg	$7 < \eta < 8.5$	190 10	137 / 395	790 / 1.0k	7.6 / 18.6
${ m FASER} u 2$	20 tons	$\eta \gtrsim 8.5$	3 ab^{-1}	178k / 668k	943k / 1.4M	2.3k / 20k
FLArE	10 tons	$\eta \gtrsim 7.5$	3 ab^{-1}	36k / 113k	203k / 268k	1.5k / 4k
AdvSND	2 tons	$7.2 \lesssim \eta \lesssim 9.2$	3 ab^{-1}	6.5k / 20k	41k / 53k	190 / 754

- The rapidity coverage of the detectors at the FPF is similar to the experiments for the Run 3.
- The increased detector/target mass and the luminosity can yield about two orders of magnitude larger number of events than the first phase experiments.
- The LHC produces large number of neutrinos in the far forward direction for all three different flavors.
 - \blacktriangleright Expected CC events: $10^5~\nu_e$, $10^6~\nu_\mu$, $10^3~\nu_\tau$ in 10 ton scale detector

Neutrino production at the LHC

■ In pp collisions, various hadrons are produced. Some of those hadrons subsequently decay and produce a number of neutrinos.

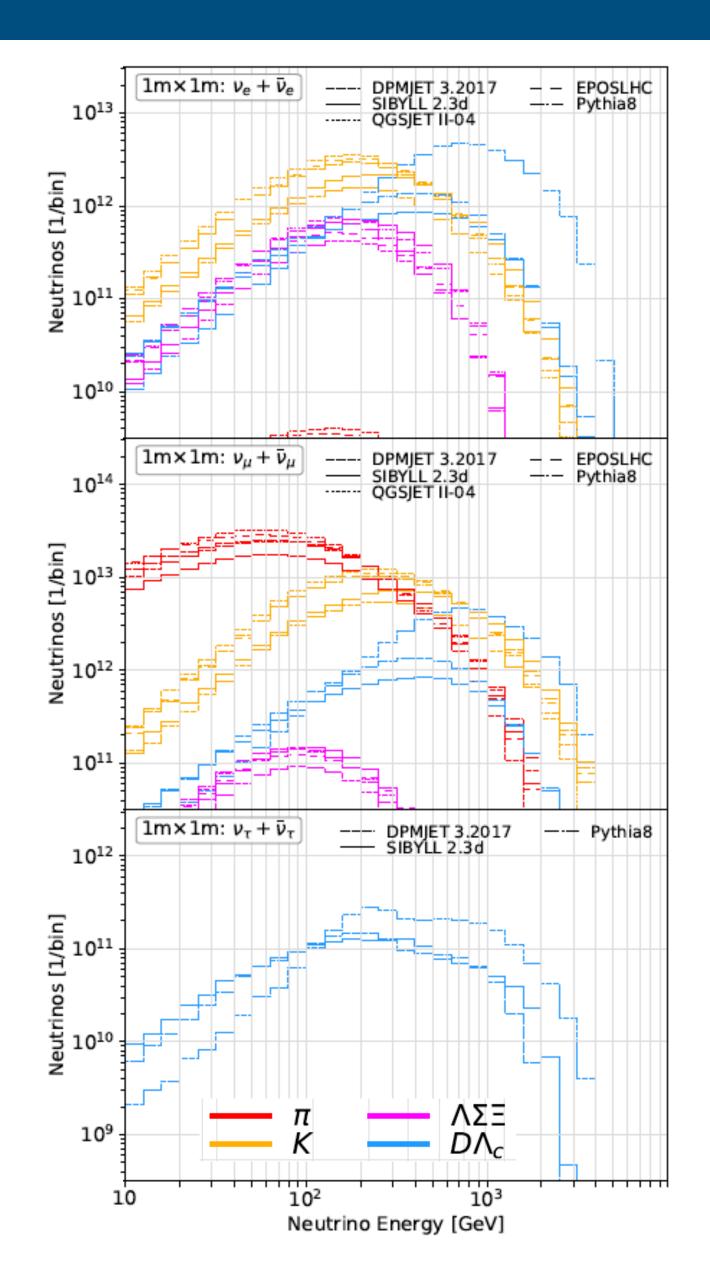
$$p + p \rightarrow h \text{ (hadrons)} + X \text{ (interaction)}$$

$$\nu_a + X' \text{ (}a = e, \mu, \tau) \text{ (decay)}$$

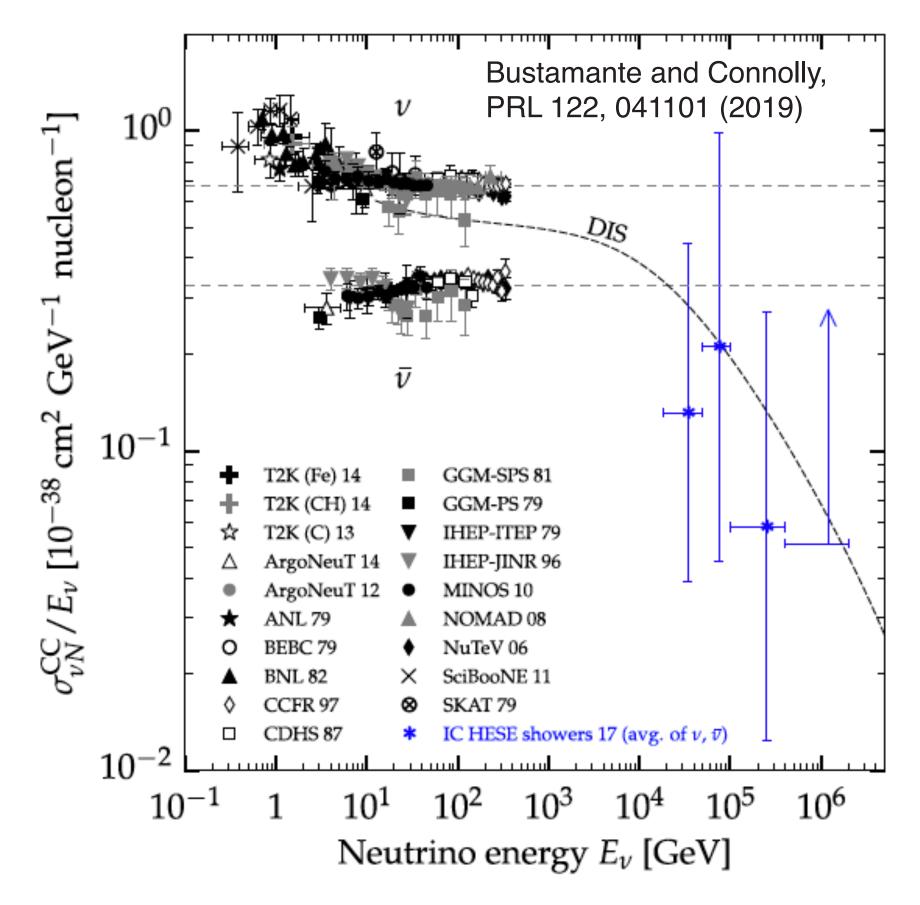
- Neutrinos from the light hadron decays: conventional neutrinos
 - $h = \pi^{\pm}, K^{\pm}, K_{L...}$
- Neutrinos from the heavy flavor hadron decays: prompt neutrinos
 - $h = D^0(\bar{D}^0), D^{\pm}, D_s^{\pm}, B^{\pm}...$

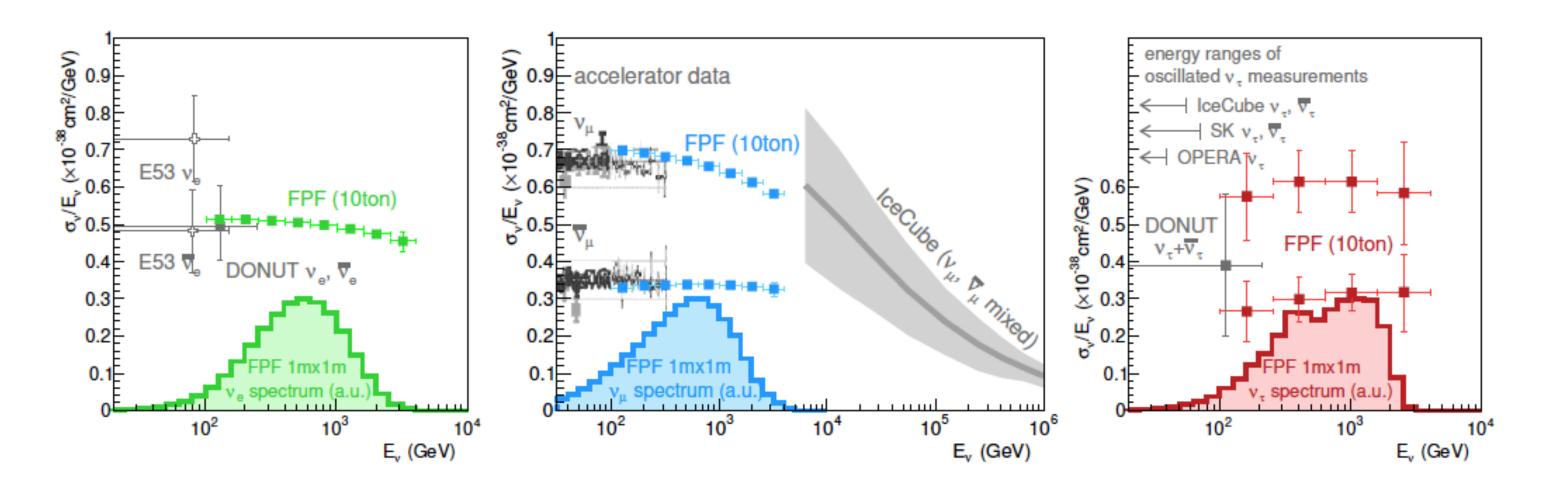
Predicted neutrino fluxes at the FPF

- Fluxes of $\nu_i + \bar{\nu}_i$ incident into the 1m x 1m cross sectional area at the FPF location of 620 m.
- Evaluated using several Monte Carlo generators.
- Neutrinos from light hadrons dominate at relatively lower energies.
- Neutrinos from charm are important at high energy range.
- Tau neutrinos are all from charm mesons.
- The prompt neutrino fluxes have large difference at high energies depending on the generator.



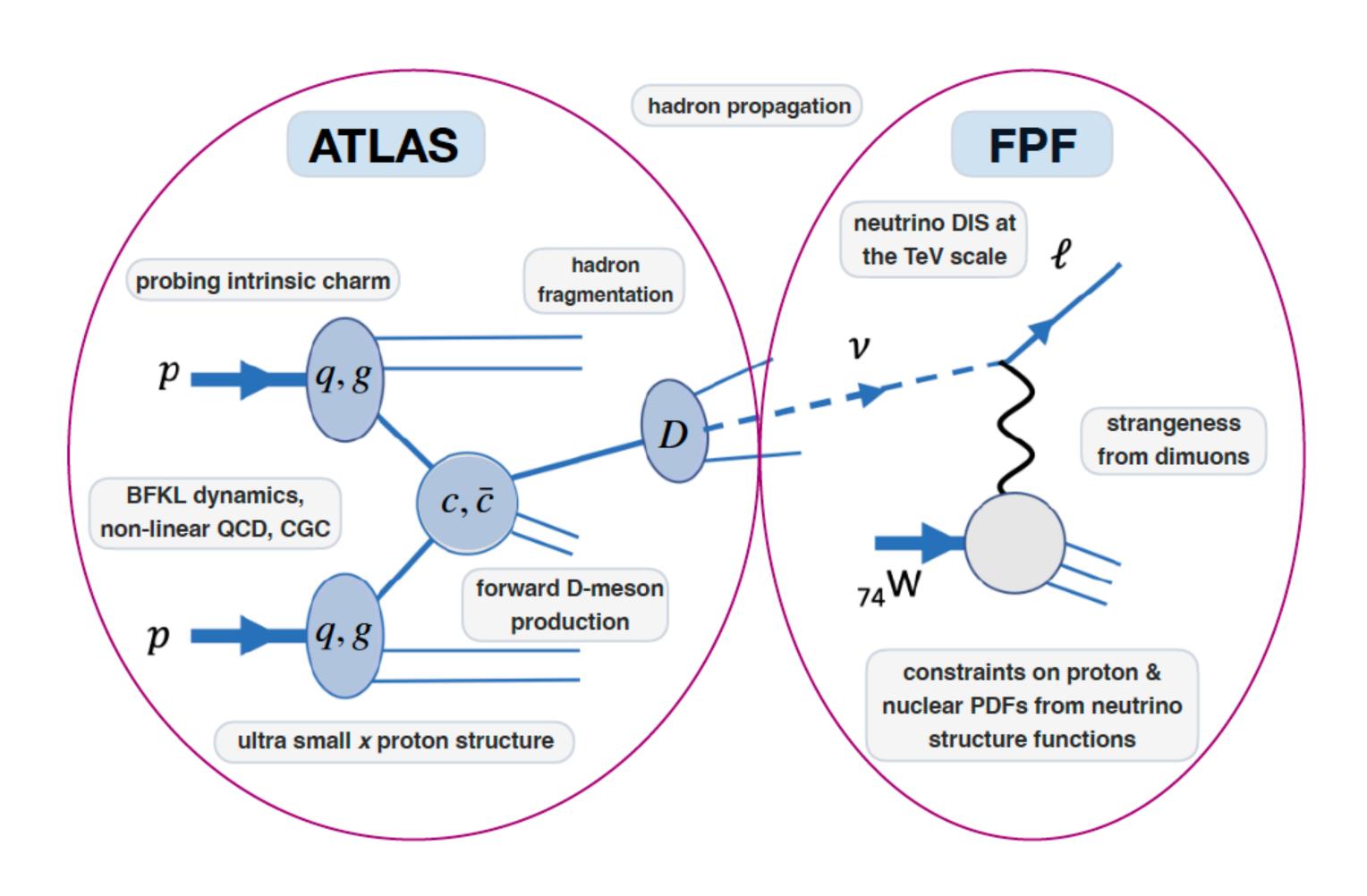
Primary goals for neutrino physics





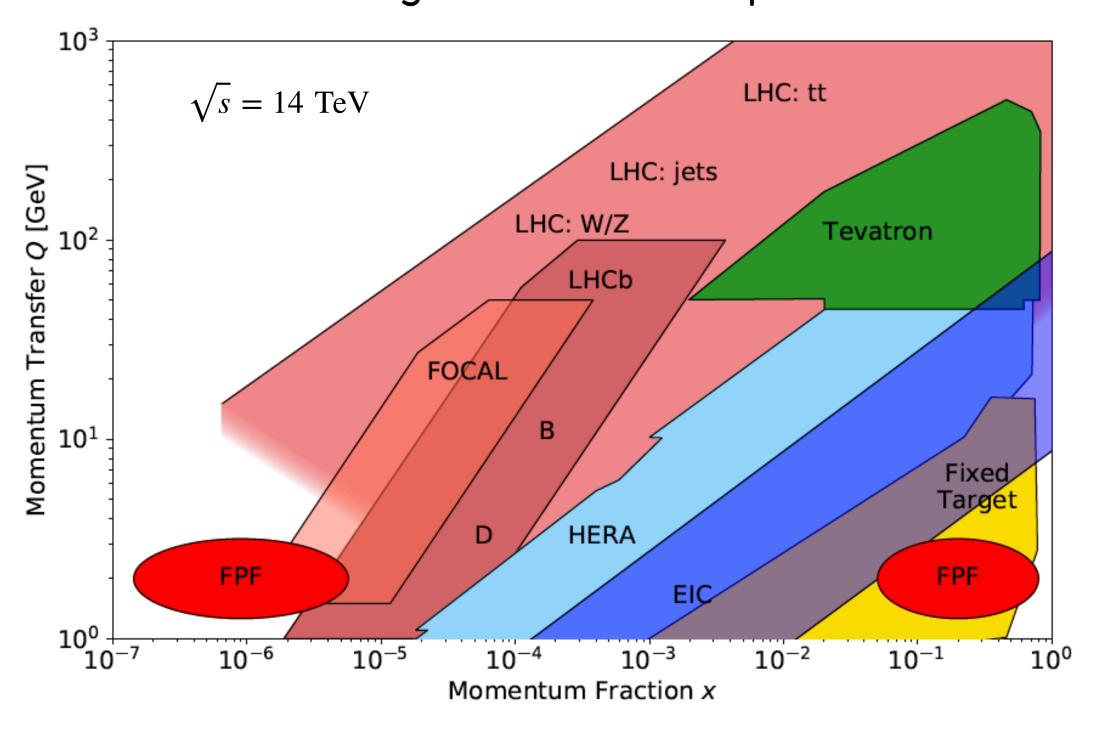
- Measurement of neutrino cross section in the (sub-)TeV energies.
- Measurement of the cross section for all three different flavors.

Prompt neutrino search at the FPF



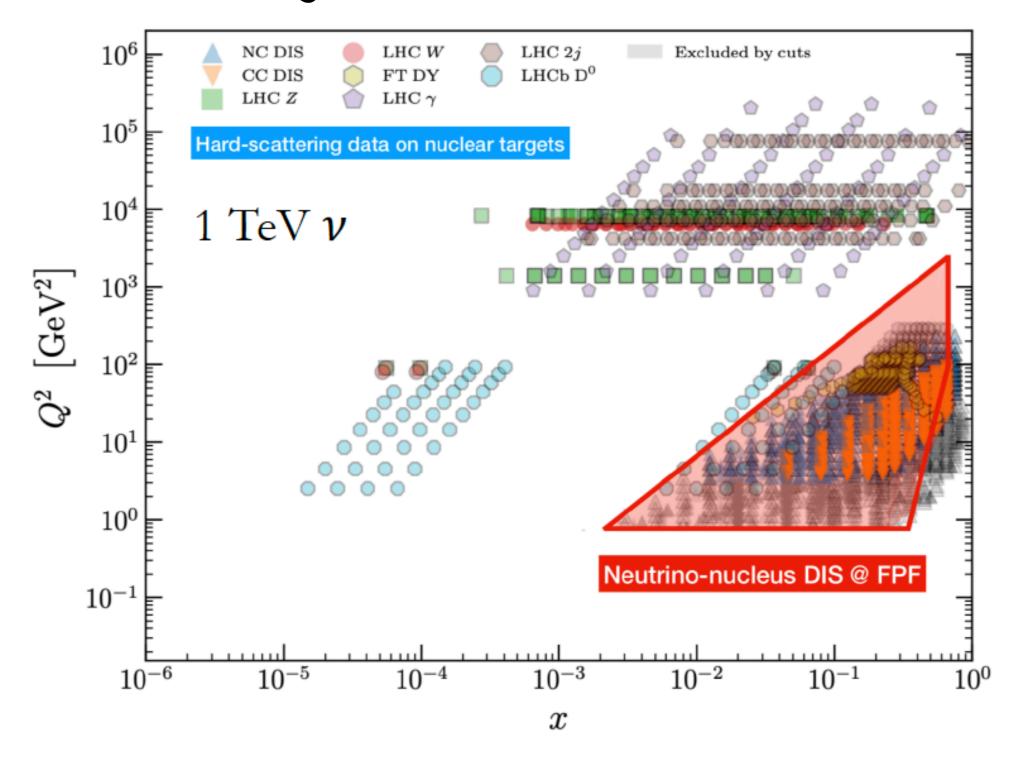
Kinematic region

Kinematic region for D meson production



Small-x involved physics
 e.g.) PDF constraints, gluon saturation,
 HQ production approaches

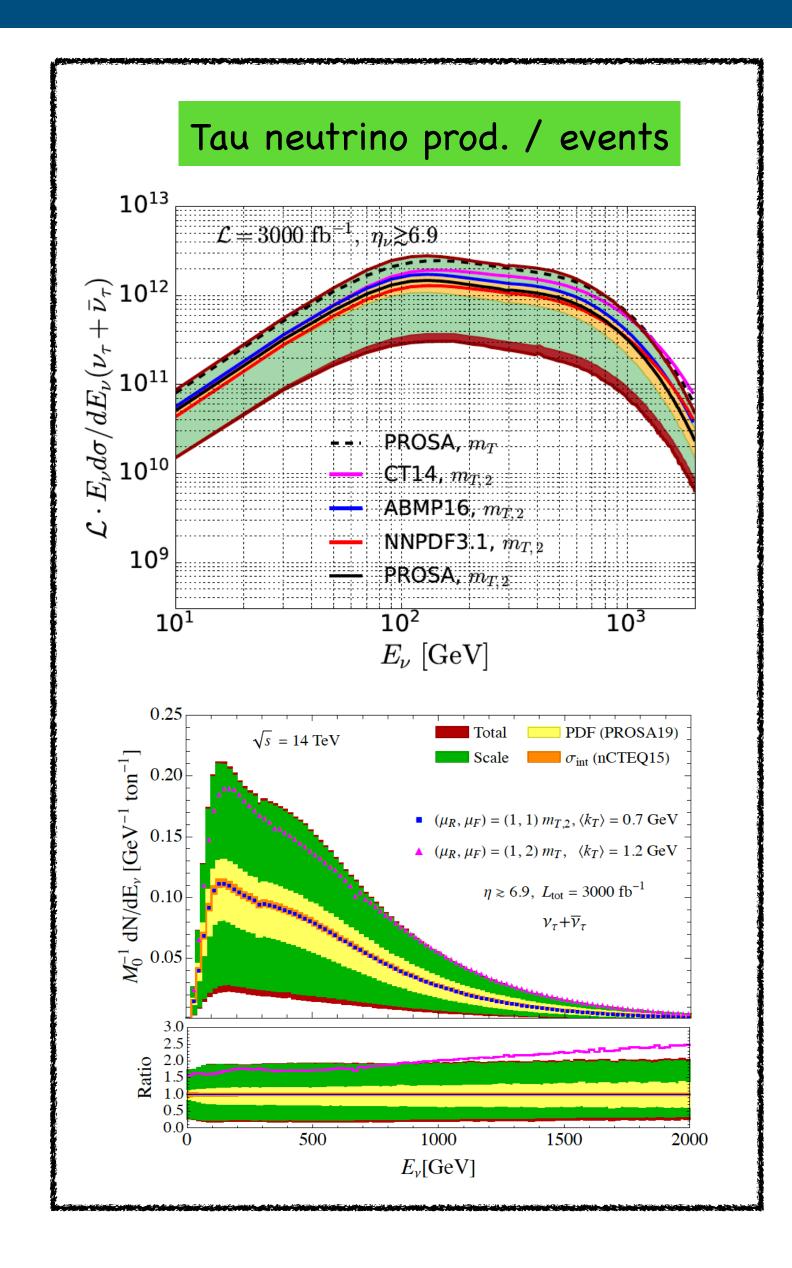
Coverage for νA DIS structure function

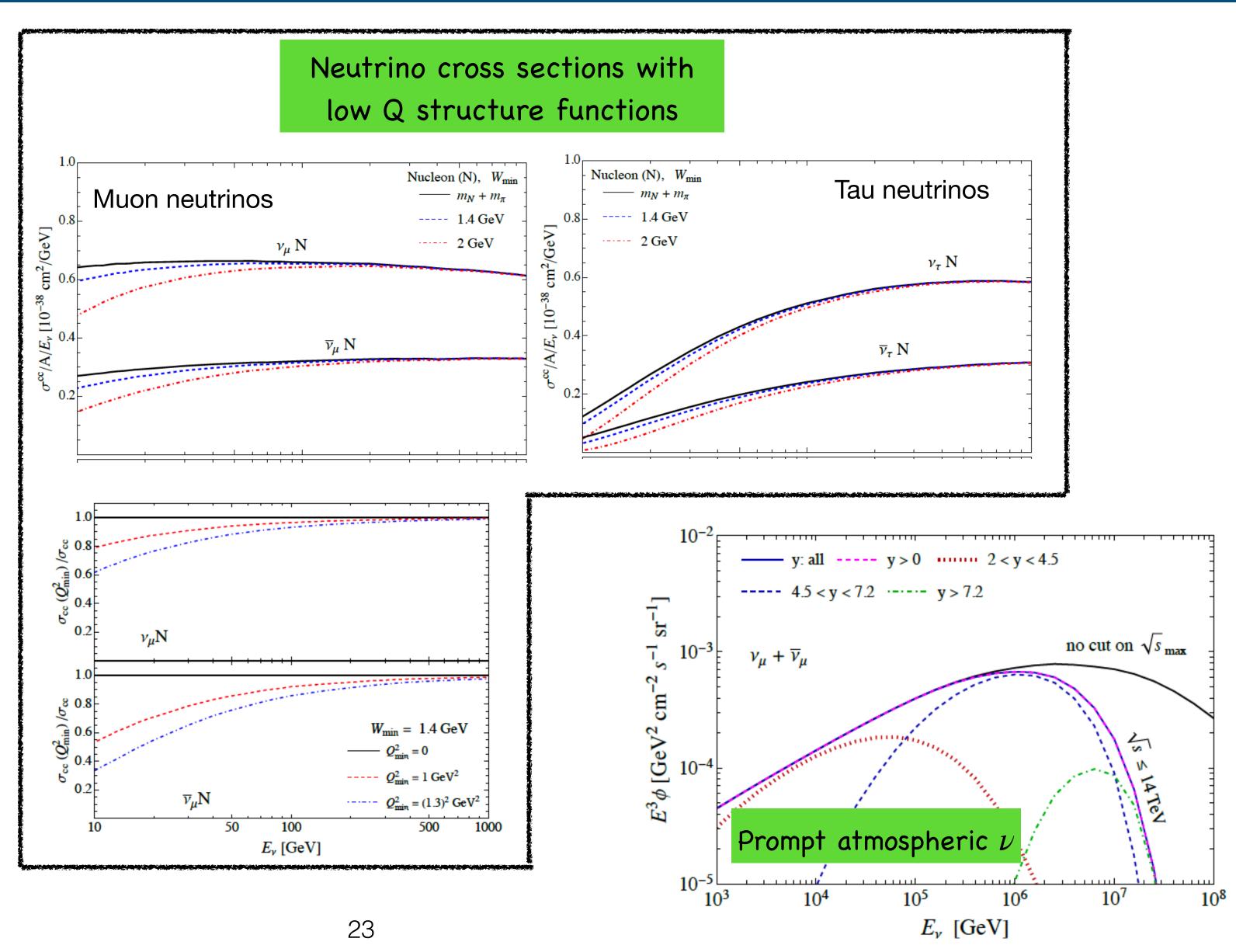


■ To constrain proton and nuclear PDFs.

FPF white paper (arXiv: 2203.05090)

Neutrinos study for the FPF





Summary II

- The LHC produce a large number of neutrinos in the far-forward region.
- The forward experiments at the LHC provide the unique opportunity to study neutrino interactions up to TeV energy range with high precision.
 - DIS is dominant interaction, which can be used to constrain the PDFs (nuclear PDFs, strange PDFs)
- Large fraction of high energy neutrinos and all tau neutrinos are produced from heavy flavor hadron decays (prompt). The measurement of these neutrinos will
 - provide neutrino data capable of probing proton structure and constraining the QCD evaluations for hadron production.
 - help to constrain/improve theoretical prediction on prompt atmospheric neutrinos

Thank you for your attention