

SM and BSM predictions of M_W using different renormalization schemes — FlexibleSUSY implementation and the example of the MRSSM

Dominik Stöckinger

TU Dresden

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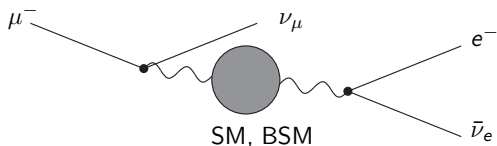
recent work with: [Peter Athron, Markus Bach, Douglas Jacob, Wojciech Kotlarski, Alexander Voigt]

Outline

- 1 SM calculation (intro/review)
- 2 BSM calculations (Problem, solution)
- 3 MRSSM prediction of M_W
- 4 Conclusions

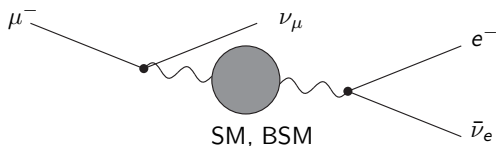
How to calculate M_W ?

$$\mu\text{-lifetime } \tau_\mu = f(M_W; \alpha, M_Z, \dots)$$



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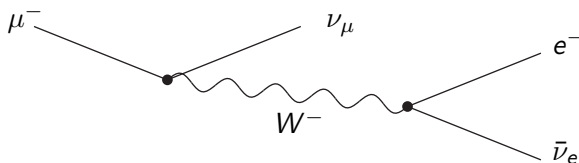


invert \Rightarrow theory prediction for M_W

μ -lifetime $\Leftrightarrow G_\mu$

Convenient: use G_μ : effective matrix element/4-fermion coupling:

$$\tau_\mu^{-1} \equiv \underbrace{G_\mu^2 \frac{m_\mu^5}{192\pi^3} \left(1 + \mathcal{O} \left(\frac{m_e^2}{m_\mu^2}, \frac{m_\mu^2}{M_W^2} \right) + \Delta_{\text{QED}} \right)}_{\text{known, model-independent}}$$



$$G_\mu \stackrel{\text{SM tree}}{=} \frac{\alpha\pi}{\sqrt{2}M_W^2 s_W^2}$$

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Muon-decay and M_W : SM on-shell scheme

(Input) parameters:

$$\alpha = \frac{1}{137. \dots} \quad M_{W,Z} = \text{pole masses} \quad s_W^2 \equiv 1 - \frac{M_W^2}{M_Z^2}$$

Central relation:

$$G_\mu = \frac{\pi\alpha}{\sqrt{2}M_W^2 s_W^2} \left(1 + \Delta r \right)$$

$\Rightarrow \Delta r =$ loop corrections (self-energy, vertex, box, counterterms)

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\Rightarrow model-prediction for M_W in terms of M_Z , α , G_μ , Δr

$$G_\mu = \frac{\pi\alpha}{\sqrt{2}M_W^2 s_W^2} \left(1 + \Delta r \right)$$

Leading loop effects:

- $\Delta\alpha$: shift of finestructure constant (via δZ_e) $\rightarrow \log \frac{m_{\text{light}}}{M_W}$
- $\Delta\rho$: shift of ρ parameter [$\sim T$ -parameter] (via δs_W^2) $\rightarrow \frac{m_t^2}{M_W^2}$

$$\Delta r = \underbrace{\Delta\alpha}_{6\%} - \frac{c_W^2}{s_W^2} \underbrace{\Delta\rho}_{-3\%} + \underbrace{\Delta r_{\text{remainder}}}_{+1\%}$$

full 2-loop + leading 3-loop available

[Sirlin. Freitas, Hollik, Walter, Weiglein, Awramik, Czakon. . .]

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Possible resummation of leading 1-loop terms [Consoli, Hollik, Jegerlehner '89]

$$1 + \Delta r = \frac{1}{(1 - \Delta\alpha)(1 + \frac{c_W^2}{s_W^2} \Delta\rho)} (1 + \Delta r_{\text{remainder}})$$

(generates correct 2-loop terms of $\mathcal{O}(\Delta\alpha^2, \Delta\alpha\Delta\rho, \Delta\rho^2)$)
 (somewhat awkward to combine with full 2-loop computation)

$$G_\mu = \frac{\pi\alpha}{\sqrt{2}M_W^2 s_W^2} \left(1 + \Delta r \right)$$

\leftrightarrow

$$G_\mu = \frac{\pi\alpha(M_Z)}{\sqrt{2}M_W^2 \left(1 - \frac{M_W^2}{\rho M_Z^2} \right)} \left(1 + \Delta r_{\text{remainder}} \right)$$

achieves same resummation, with

$$\alpha(M_Z) \equiv \frac{\alpha}{1 - \Delta\alpha}$$

via $\Pi^\gamma(M_Z^2) - \Pi^\gamma(0)$

$$\rho \equiv \frac{1}{1 - \Delta\rho}$$

via $\Sigma_{W,Z}(0)$

\Rightarrow absorbs leading effects in new parameters! $\longrightarrow \log \frac{m_{\text{light}}}{M_W}, \frac{m_t^2}{M_W^2}$

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Muon-decay and M_W : SM $\overline{\text{MS}}$ -scheme

- $\overline{\text{MS}}$ -scheme with $\mu = M_Z$
- simpler prescription for counterterms, but obscured relation to observables
- here: possibility to similarly absorb leading effects! [Degrassi, Fanchiotti, Sirlin '89]

Muon-decay and M_W : SM $\overline{\text{MS}}$ -scheme

(Input) parameters:

$$\hat{\alpha}, \hat{s}^2 = \overline{\text{MS}} \text{ via gauge couplings} \quad M_{W,Z} = \text{pole masses}$$

Central relation:

$$G_\mu = \frac{\pi \hat{\alpha}(M_Z)}{\sqrt{2} M_W^2 \hat{s}^2} \left(1 + \Delta \hat{r}_W \right)$$

Further relations to observables:

$$\hat{\alpha}(M_Z) \equiv \frac{\alpha}{1 - \Delta \hat{\alpha}} \quad \hat{\rho} \equiv \frac{M_W^2}{M_Z^2 \hat{c}^2} = \frac{c_W^2}{\hat{c}^2} \quad \hat{s}^2 = \left(1 - \frac{M_W^2}{\hat{\rho} M_Z^2} \right)$$

via $\Pi^\gamma(0)_{fin}$ via $\delta M_{W,Z fin}$

$$G_\mu = \frac{\pi \hat{\alpha}(M_Z)}{\sqrt{2} M_W^2 \hat{s}^2} \left(1 + \Delta \hat{r}_W \right) \quad \hat{s}^2 = \left(1 - \frac{M_W^2}{\hat{\rho} M_Z^2} \right)$$

Same structure as resummed form of on-shell scheme!

Contains the same leading loop effects:

- $\Delta \hat{\alpha}$: $\longrightarrow \log \frac{m_{\text{light}}}{M_W}$
- $\hat{\rho}$: $\longrightarrow \frac{m_t^2}{M_W^2}$

$\overline{\text{MS}}$ -scheme calculation promises higher accuracy but involves more intermediate quantities. Status=on-shell:

full 2-loop + leading 3-loop available [... Degraasi, Gambino, Giardino...]

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Definitions and remarks

$$\alpha, s_W^2$$

$$\hat{\alpha}(M_Z), \hat{\rho}$$

$$\hat{\alpha}(M_Z), \hat{s}^2$$

$$\Delta\alpha \equiv \Pi_{\text{light}}^\gamma(0) - \Pi_{\text{light}}^\gamma(M_Z^2)$$

$$\Delta\rho \equiv \frac{\Sigma_Z(0)}{M_Z^2} - \frac{\Sigma_W(0)}{M_W^2}$$

$$\Delta\hat{\alpha} \equiv \Pi^\gamma(0)|_{\text{fin}, \mu=M_Z}$$

$$\frac{1}{\hat{\rho}} = \left(1 + \frac{\delta M_W^2}{M_W^2} - c_{\text{bare}}^2 \frac{\delta M_Z^2}{M_Z^2} \right)_{\text{fin}}$$

Remarks:

- relate low-E (G_μ, α) to high-E ($M_{W,Z}$) [QED running] $\longrightarrow \log \frac{m_{\text{light}}}{M_W}$
- breaking of custodial symmetry $\longrightarrow \frac{m_t^2}{M_W^2}$

Current status

as in [Degrassi, Gambino, Giardino '14/Athron, Bach, Jacob, Kotlarski, Voigt '22]

$$M_W = \begin{cases} 80.357(3) & \overline{\text{MS}}\text{-scheme} \\ 80.351(3) & \overline{\text{MS}}\text{-scheme (update)} \\ 80.363(4) & \text{on-shell scheme} \\ 80.355(3) & \text{on-shell scheme (update)} \end{cases}$$

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 - Non-decoupling problem (in \overline{MS} scheme)
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Overview BSM contributions

typically dominant contributions via $\Delta\rho \sim T$ -parameter. E.g.

- Higgs triplet (tree-level ρ)
- 2HDM
- stop/sbottom, other SUSY particles
- 4th generation (t' , b')

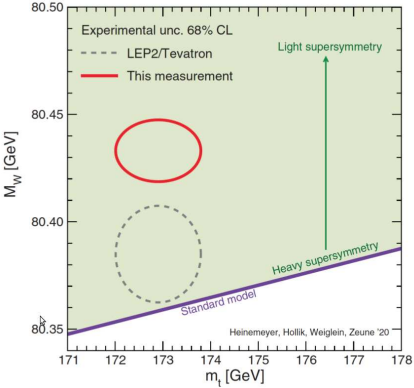
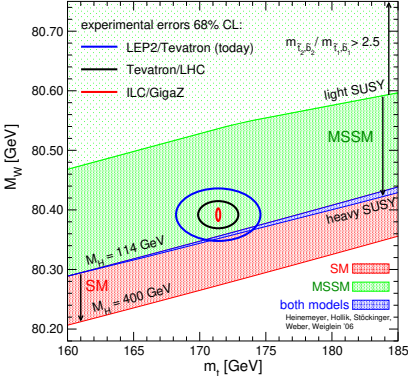
In general, heavy BSM decouples

$$M_W^{\text{BSM}} \rightarrow M_W^{\text{SM}} \quad (\text{if } M_{\text{BSM}} \rightarrow \infty)$$

in particular, no change of QED running below M_W

Example: on-shell calculations in MSSM and 2HDM

MSSM: full 1-loop Δr , 2-loop $\Delta\rho$ [Haestier, Heinemeyer, Hollik, DS, Weber, Weiglein '06... Zeune]



2HDM: full 1-loop Δr and 2-loop $\Delta\rho$ [Lopez-Val, Sola... Hossenberger, Hollik]

How to implement M_W in generic spectrum generators?

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$\overline{\text{MS}}$ -scheme looks natural \rightsquigarrow applied in

- “classical” calculations [Bagger,Matchev,Pierce,Zhang'96]
- Softsusy, Spheno, FlexibleSUSY...

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The non-decoupling problem

In hindsight obvious, but initially obscured! E.g. by light spectra and “compensating bug” in some programs

$$G_\mu = \frac{\pi \hat{\alpha}(M_Z)}{\sqrt{2} M_W^2 \hat{s}^2} \left(1 + \Delta \hat{r}_W \right) \quad \hat{s}^2 = \left(1 - \frac{M_W^2}{\hat{\rho} M_Z^2} \right)$$

$$M_{\text{BSM}} \gg M_W \gg m_{\text{light}} !!$$

The non-decoupling problem

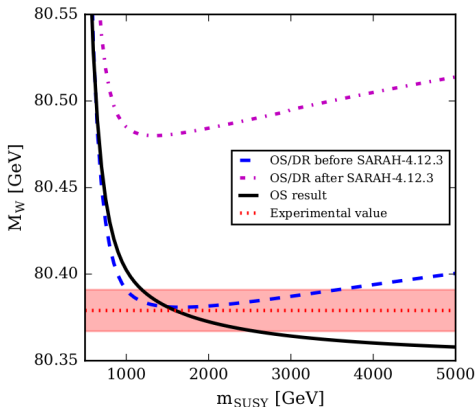
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$$M_{\text{BSM}} \gg M_W \gg m_{\text{light}} !!$$

- $\Delta \hat{\alpha}$: $\rightarrow \log \frac{m_{\text{light}}}{M_W}$ (physical), $\log \frac{M_{\text{BSM}}}{M_W}$ (unphysical, must drop out)
- $\hat{\rho}$: $\rightarrow \frac{m_t^2}{M_W^2}$, BSM sources of custodial sym. break. (physical)

Example of the non-decoupling problem [Diessner, Weiglein '19]



Origin: 1-Loop implementation of $\overline{\text{MS}}$ formula leads to

$$\frac{\dots}{1 - \Delta\hat{\alpha}^{1L}} (1 + \Delta\hat{r}_W^{1L}) \sim 1 + \underbrace{\Delta\hat{\alpha}^{1L} + \Delta\hat{r}_W^{1L} + \dots}_{\text{non-dec. cancels}} + \underbrace{(\Delta\hat{\alpha}^{1L})^2 + \dots}_{\text{incomplete 2L, non-decoup.}}$$

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FlexibleSUSY:

- spectrum generator generator
- e.g. MSSM, NMSSM, MRSSM, 2HDM, other non-SUSY models. . .
- precise computation of BSM spectrum and observables (M_H , a_μ , FlexibleDecay. . .)

M_W :

- SM-contributions fully taken into account from [Degrassi, Gambino, Giardino '14]
- BSM-contributions: strictly expand \overline{MS} formula to 1L-BSM order:

$$M_W^2 = M_W^{2\text{SM}}(1 + \Delta_W)$$

with

$$\Delta_W = \frac{\hat{s}^2}{\hat{c}^2 - \hat{s}^2} \left[\frac{\hat{c}^2}{\hat{s}^2} (\Delta\hat{\rho}_{\text{tree}} + \Delta\hat{\rho}) - \Delta\hat{r}_W - \Delta\hat{\alpha} \right]_{\text{BSM}}^{1\text{L}}$$

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Technical summary of MRSSM

Motivated by new symmetry, $\theta \rightarrow e^{i\alpha}\theta$ (more predictive!)

- \tilde{q}_L : $R=+1$, \tilde{q}_R : $R=-1$, no LR-mixing! Majorana masses forbidden!

Dirac gauginos, new superfields \hat{O} , \hat{T} , \hat{S}

- new Higgs triplet!

Dirac Higgsinos, new superfields \hat{R}_u , \hat{R}_d

- New superpotential terms

$$W_{\text{MRSSM}} = \dots + \mu_u \hat{H}_u \hat{R}_u + \Lambda_u \hat{H}_u \hat{T} \hat{R}_u + \lambda_u \hat{H}_u \hat{S} \hat{R}_u + y_u \hat{Q} \hat{H}_u \hat{U}$$

like top-Yukawa coupling!

Main effects in MRSSM

Higgs triplet VEV and Λ, λ -couplings \Rightarrow impact on M_W (and M_h)

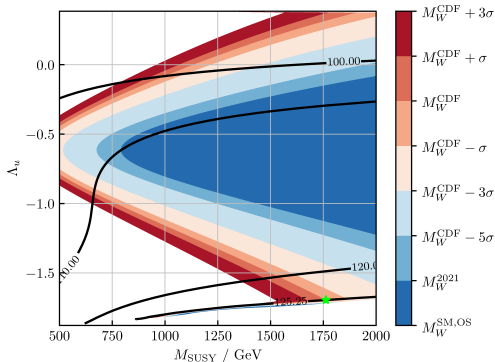
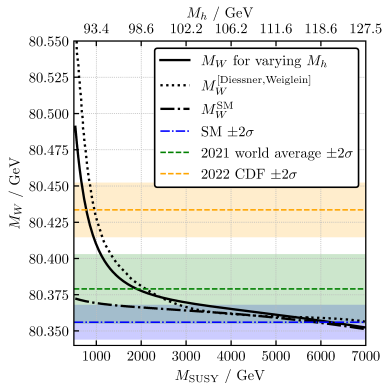
Tree-level

$$\hat{\rho}_{\text{tree}} = 1 + \frac{4v_T^2}{v_d^2 + v_u^2}$$

Main one-loop effect from Λ, λ 's

$$T \approx \frac{1}{16\pi^2 s_W^2 M_W^2} \frac{v^4(4\lambda^4 + 5\Lambda^4)}{10m_D^2}$$

MRSSM results [Athron, Bach, Jacob, Kotlarski, DS, Voigt '22]



$$\tan \beta = 37.6, M_{\text{SUSY}} = 1727 \text{ GeV}$$

$$\Lambda_d = 2.41, \Lambda_u = 1.70, \lambda_d = 2.43, \lambda_u = 0.267$$

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Conclusions

- SM theory prediction very advanced

- ▶ on-shell and $\overline{\text{MS}}$ at 2-loop + beyond
- ▶ uncertainties $\sim \pm 3 \dots 10$ MeV

- BSM predictions tricky (extra scale!)

- ▶ on-shell 2-loop $\Delta\rho$ in some models
- ▶ SM-like $\overline{\text{MS}}$ -scheme non-decoupling

- FlexibleSUSY & applications

- ▶ FlexibleSUSY now implements full SM+strict BSM-1L
- ▶ Decoupling works, accurate prediction
- ▶ MRSSM is promising!

