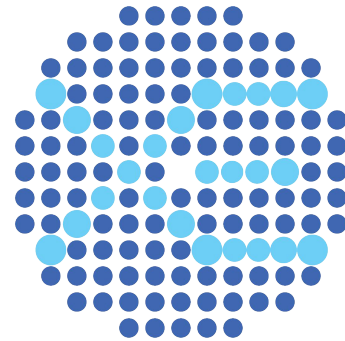


Search for new physics with the XENONnT detector

R. Biondi on behalf of XENON Collaboration



XENON

34th Rencontres de Blois - 16th May 2023

The XENON Collaboration







Last Collaboration Meeting in L'Aquila 1-3 February 2023



- 27 Institutions Worldwide
- More than 180 Scientists

Main goal:

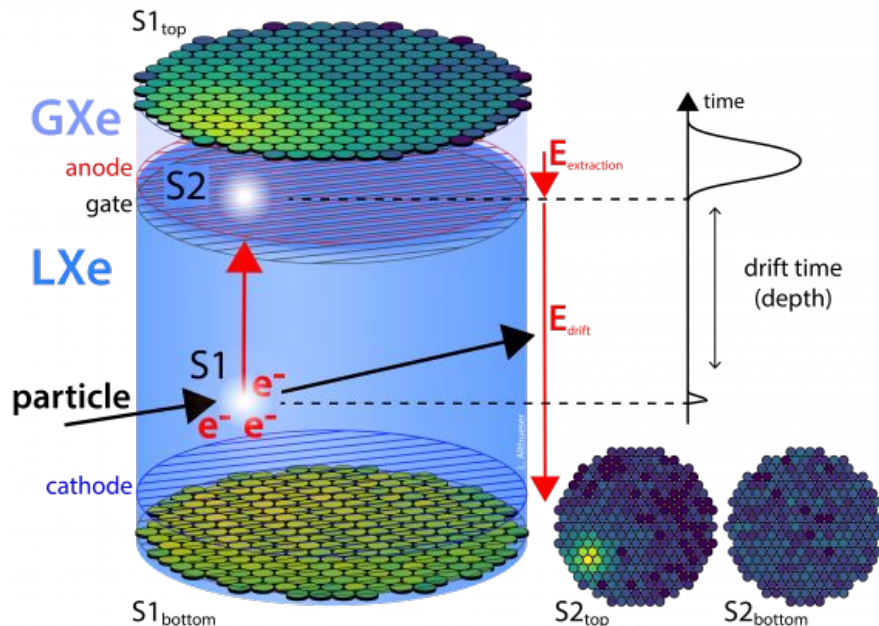
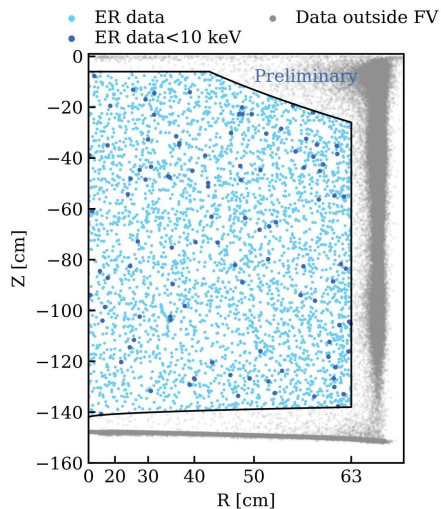
Detection of dark matter particles with a liquid xenon TPC

-  [Xenonexperiment.org](https://xenonexperiment.org)
-  [Xenon_experiment](https://www.instagram.com/xenon_experiment)
-  [XENONexperiment](https://twitter.com/XENONexperiment)
-  [XENONexperiment](https://www.facebook.com/XENONexperiment)

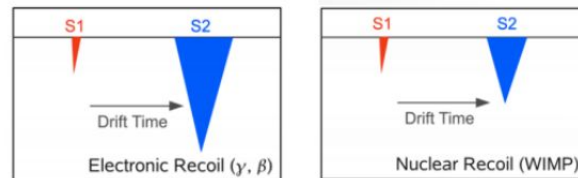
The TPC Detection Principle



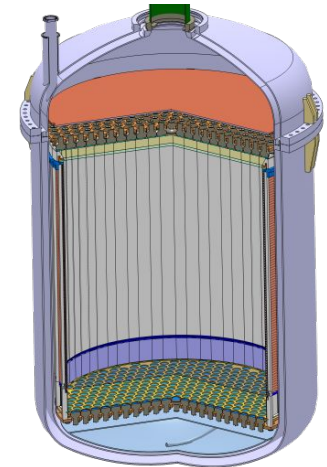
- Dual-phase (liquid+gas)
- Energy reconstruction
- 3D event reconstruction
- Fiducialization
- Event discrimination



$$E = W \left(\frac{S_1}{g_1} + \frac{S_2}{g_2} \right)$$



Evolution of Xe-based DM detectors



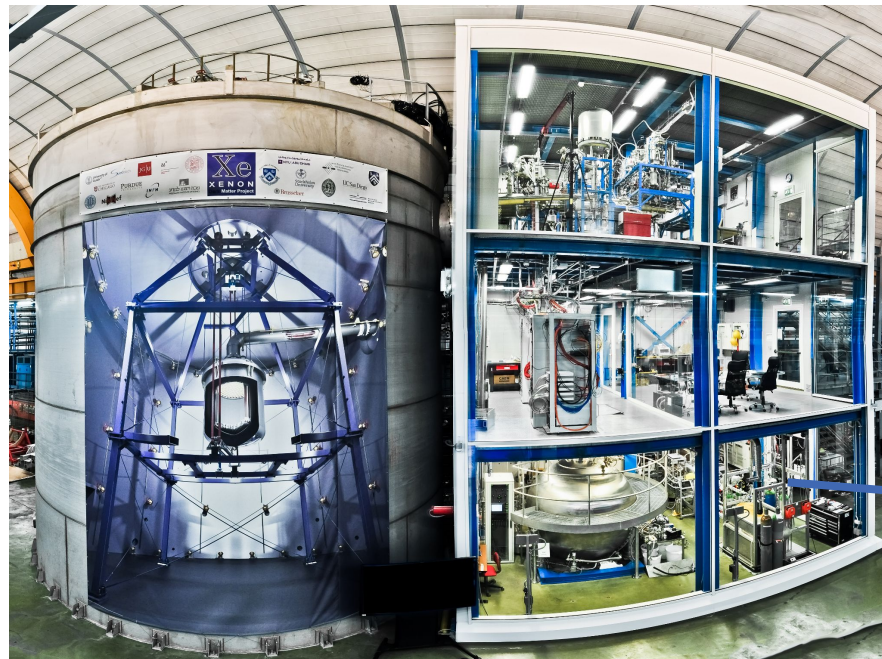
	XENON10	XENON100	XENONIT	XENONnT	DARWIN - XLZD
Period	2005 - 2007	2008 - 2016	2012 - 2018	2019 - 2025	2025 - ...
Dimensions	15 x 20 cm	30 x 30 cm	1 x 1 m	1.5 x 1.3 m	2.6 x 2.6 m
Active mass	14 kg	62 kg	2 tons	5.9 tons	40 tons
Sensitivity	$\sim 10^{-43} \text{ cm}^2$	$\sim 10^{-45} \text{ cm}^2$	$\sim 10^{-47} \text{ cm}^2$	$\sim 10^{-48} \text{ cm}^2$	$\sim 10^{-49} \text{ cm}^2$

The XENON project at LNGS



Laboratori Nazionali del Gran Sasso – INFN (Hall B)

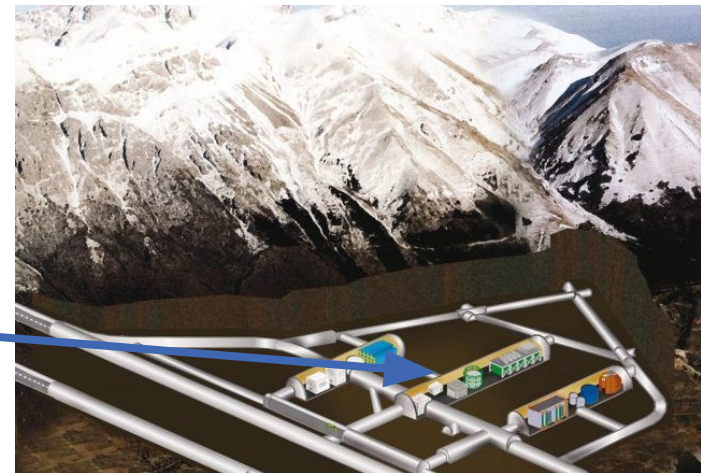
XENONIT



1400 mt of rock: 3.800 m.w.e.

Muon flux $\sim 1 \text{ m}^{-2} \text{ h}^{-1}$

Suppression factor: 10^6



Low-ER excess in XENONIT



From: Phys. Rev. D 102, 072004

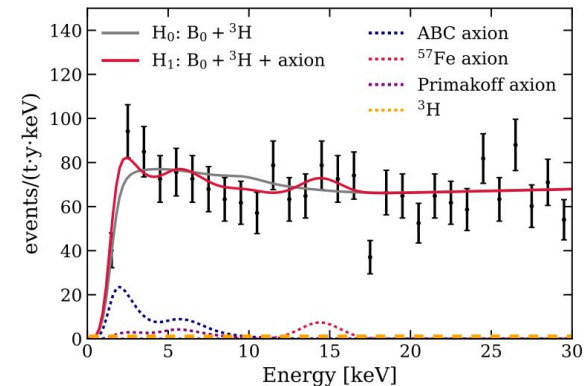
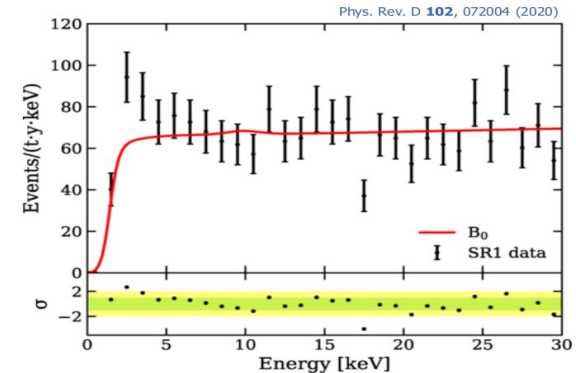
ER search in <30 keV range shown an excess of events over the expected background corresponding to a **3.3 σ fluctuation**

Hypothesis of ^{37}Ar leak is excluded:

- Removed by Kr distillation
- Limits on air leak from other contaminants (Kr, Rn)

Several hypothesis:

- Solar axions (3.4σ over bkg)
- Neutrino magnetic moment (3.2σ over bkg)
- Bosonic DM: ALPs, dark photons (3.0σ over bkg)
- β decay of Tritium
- New Physics?





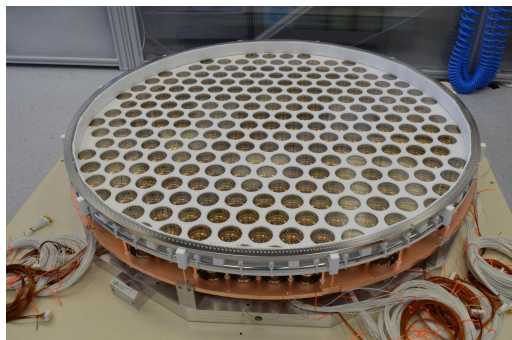
Fast upgrade exploiting the infrastructures from XENONIT



New TPC



- Drift length 1.5 m (XENONIT had 1 m)
- Active **Xe mass 5.9 t** (2 t)
- Double the number of PMTs to 494
- Light detection efficiency: 36%
- Carefully **selected materials** to minimize background
- Field shaping rings with tunable potential



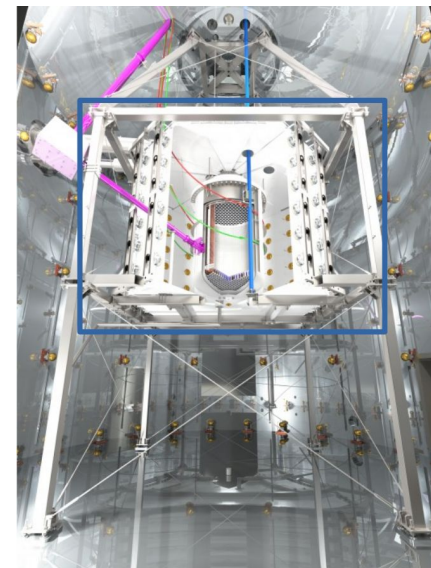
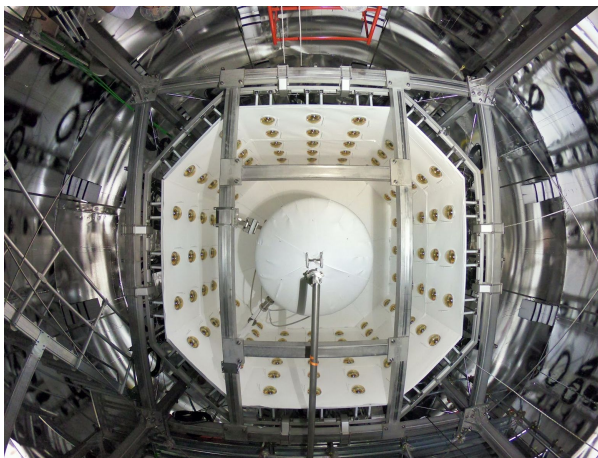
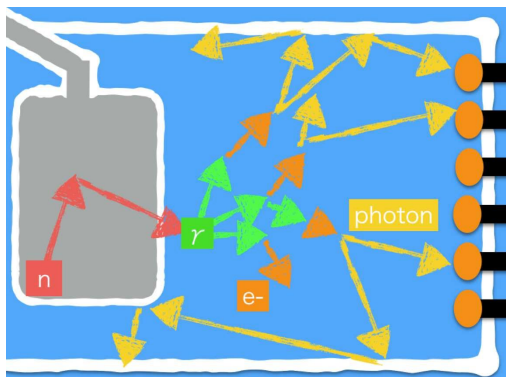
Commissioned in March 2020 during the lockdown due to COVID pandemic

Neutron Veto



Water Cherenkov detector built around the cryostat with 120 PMTs inside an enclosure of reflective panels

To tag **neutrons events** which contribute to background in WIMP search



Running with pure water: **Measured tagging efficiency 68%**

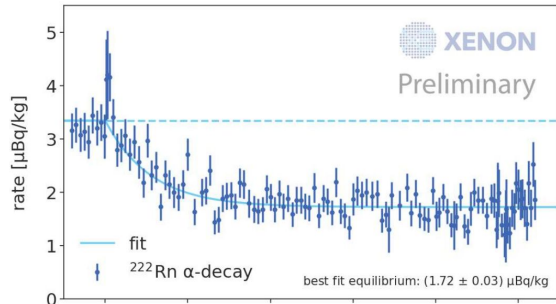
It will be doped with Gd to increase performances (~87% efficiency predicted)

Rn Removal Column



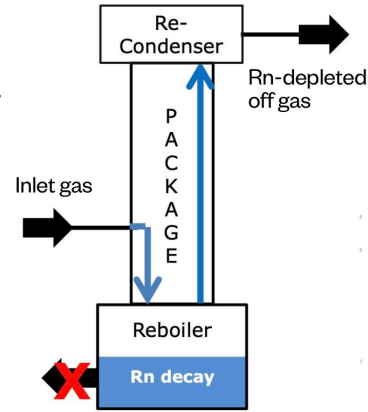
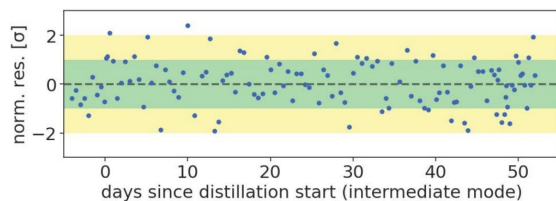
Low-ER search background comes mainly from β -decay of ^{214}Pb originating from ^{222}Rn accumulating into LXe and GXe, mitigated through careful material screening and selection.

Rn atoms are kept in the column and decay



$1.7\mu\text{Bq/kg}$

(GXe-only mode)



Two operation modes: GXe and LXe:

- GXe: Remove Rn emanated from signal cables, cryogenic pipe, ...
- LXe: Remove Rn emanated from TPC body, PMT, ...

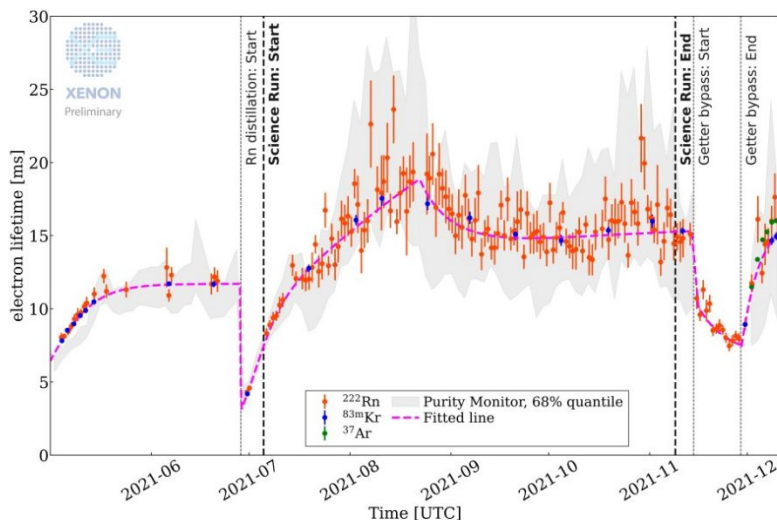
Further x2 reduction achieved for new science run with GXe+LXe: **$<1 \mu\text{Bq/kg}$**

Liquid Xe Purification System



Direct liquid circulation with cryogenic pump To purify faster the full inventory of LXe

- High flow: 2 liters LXe/min
- Replaceable filter units with low Rn emanation
- Online purity measurement by purity monitor



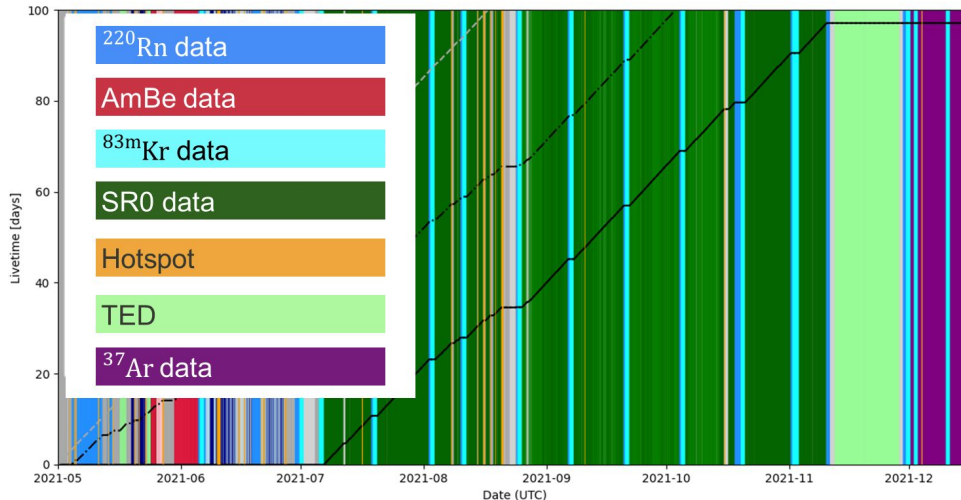
Achieved x10 better purity than XENONIT

Electron lifetime > 10 ms in science run

First Science Run - SR0



97.1 days livetime: July 6th - November 10th 2021



- All but 17 PMTs working
- Triggerless acquisition
- Online processing
- Gain: 2×10^6 stable at 3% level
- Drift field: 23 V/cm due to short circuit between cathode and bottom screen
- Extraction field: 2.9 kV/cm
- Extraction efficiency: ~53%
- Single Electron Gain: ~31 PE
- Internal calibration: ^{83m}Kr , ^{220}Rn , ^{37}Ar
- External calibration: AmBe
- Tritium Enhanced mode (TED)

Main Goal:
Clarify the excess in LOW-ER from XENONIT

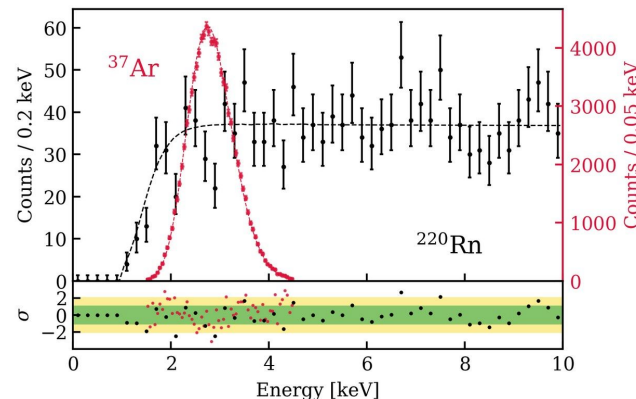
Occasional temporary ramp-downs of the anode, due to localized, high-rate, bursts of electrons (Hotspot).

Calibration

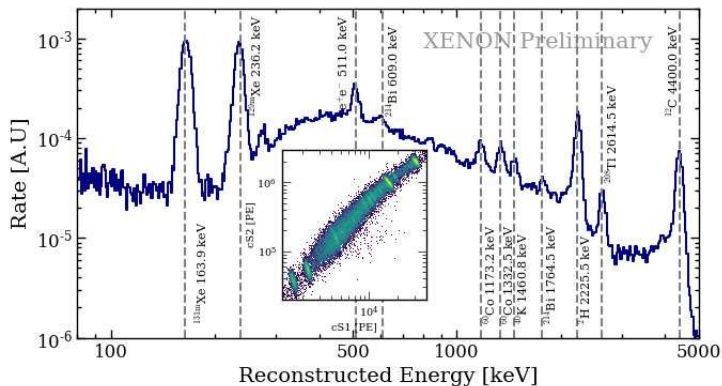


Two **ER calibration** sources at low energies:

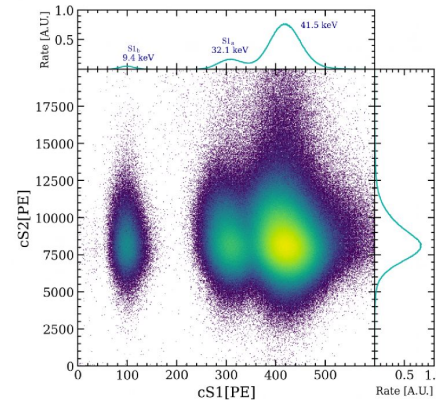
- **^{37}Ar** : mono-energetic peak at 2.82 keV, validates resolution model and energy reconstruction of peaks.
- **^{220}Rn** : its daughter ^{212}Pb provides a flat β -spectrum used to estimate cut acceptances and energy threshold



NR calibration with **AmBe** source:



$^{83\text{m}}\text{Kr}$: bi-weekly calibrations: for corrections and e-lifetime

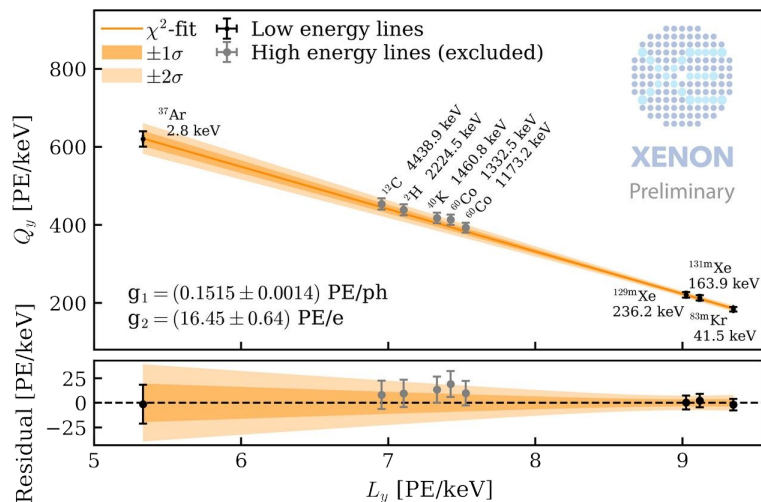
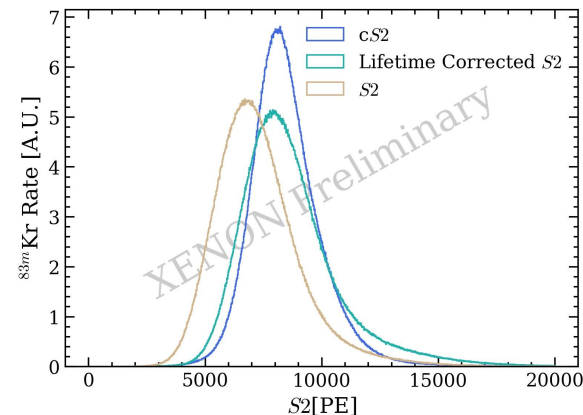


Corrections and Energy Reconstruction



S1 and S2 signals have to be corrected to take into account **position dependent response** of the detector, this is done via the periodical ^{83m}Kr calibration.

Corrected signals cS1 and cS2 are then used in the analysis



Energy reconstruction:

Using: ^{37}Ar , ^{83m}Kr , ^{131m}Xe , ^{129m}Xe calibration

$$E = 13.7 \text{ eV} \left(\frac{cS1}{g_1} + \frac{cS2}{g_2} \right)$$

Low Energy ER Analysis



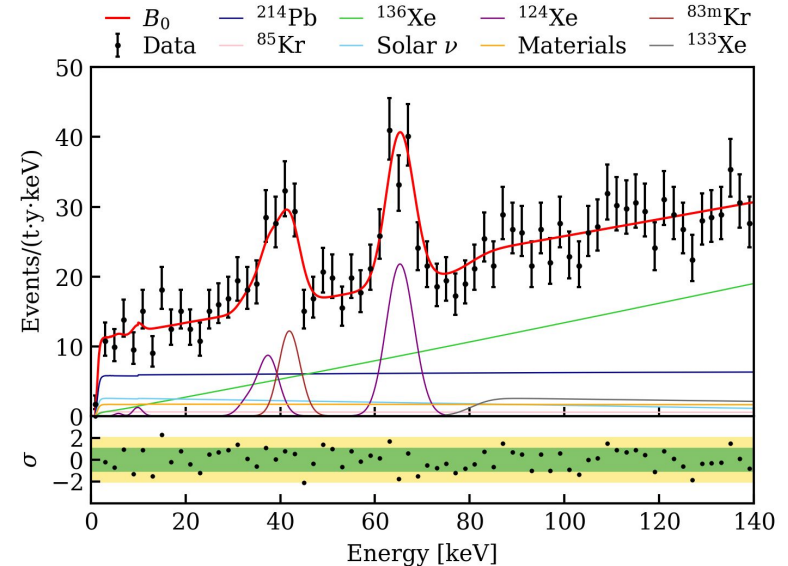
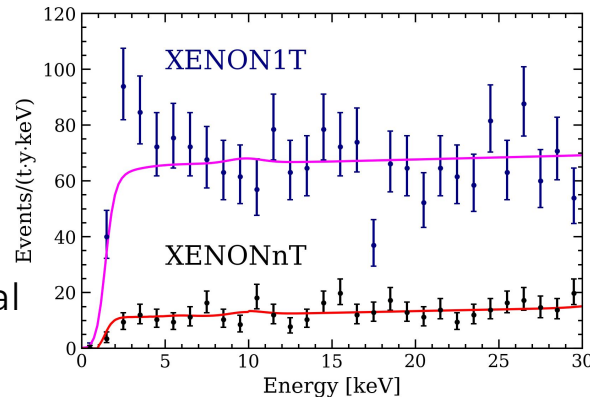
From: PRL 129, 161805 (2022)

Full blind analysis

- Energy range: (1-140) keV
- Fiducial mass: (4.37 ± 0.14) t
- Exposure: **1.16 tons years**

No excess observed

The excess observed in XENON1T could come from either tritium or statistical fluctuation



Background in (1,30) keV:

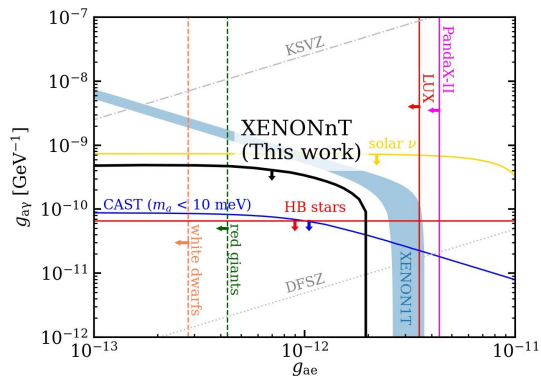
$(15.2 \pm 1.3 \text{ stat}) \text{ events}/(\text{t} \cdot \text{y} \cdot \text{keV})$

Factor ~ 5 reduction w.r.t. XENON1T

Limits on New Physics

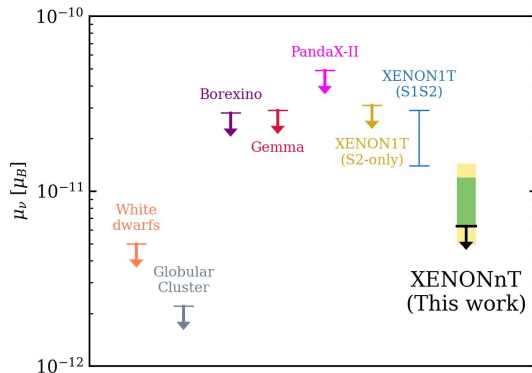


Solar Axions



Limit on 14.4 keV peak for ⁵⁷Fe solar axions is < 20 events/(t*y)

ν Magnetic Moment

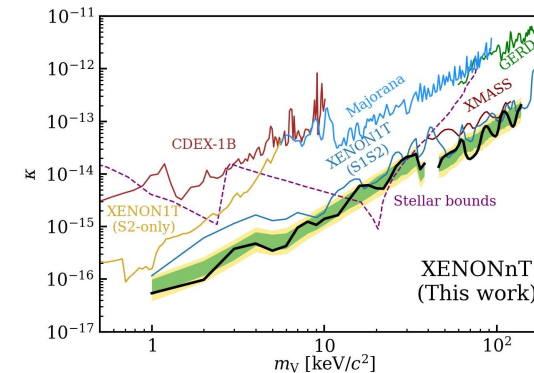
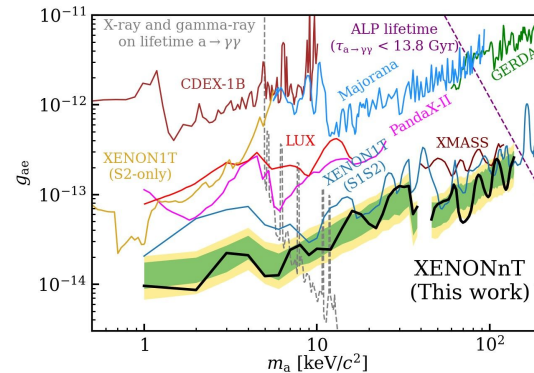


$$\mu_\nu < 6.3 \cdot 10^{-12} \mu_B$$

The most stringent limit from direct detection experiment

Dark Photon

Axion-like particles



WIMPs search



From arXiv: 2303.14729

Background model:

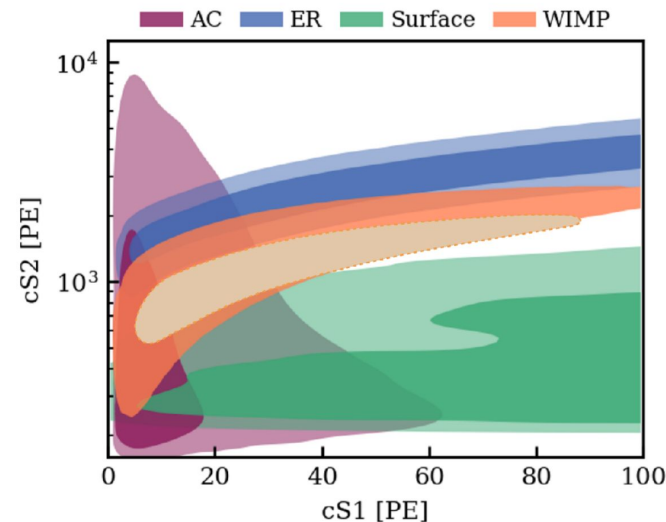
Electronic Recoil (ER): Mainly β decay of ^{214}Pb from ^{222}Rn

Accidental Coincidences (AC): random pairing of small S1 and S2 signals

Surface/Wall: ^{210}Pb plate-out on the PTFE wall of the TPC

Nuclear Recoil backgrounds (same shape as WIMP):

- CEvNS
- Neutrons



Signal-like region containing 50% of a 200 GeV/c² WIMP signal with highest signal-to-noise ratio

WIMPs Results

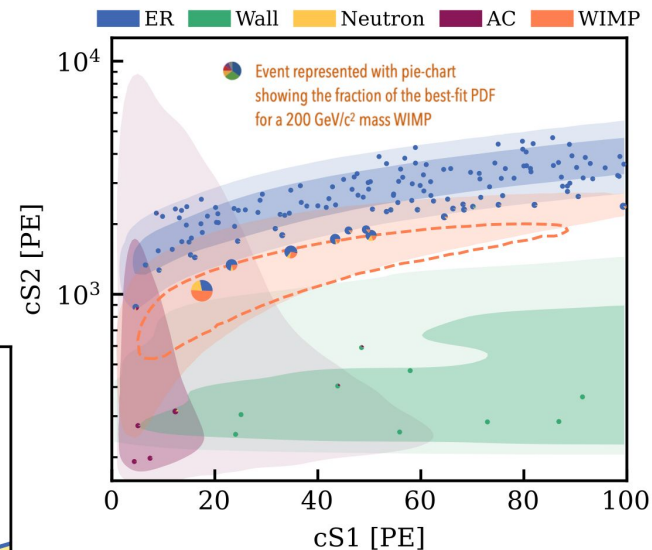
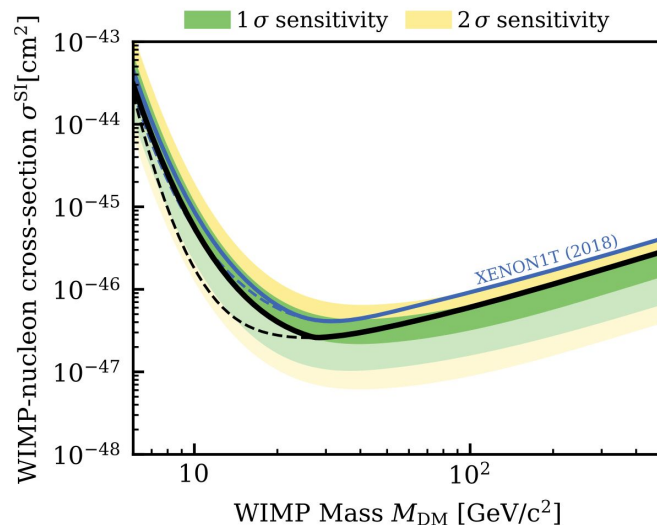


Full blind analysis

- Fiducial mass: (4.18 ± 0.13) t (Same shape of ER, but smaller $R < 61.35$ cm)
- Exposure: **1.1 tons years**

No significant excess

XENONnT 90% C.L.
Power-Constrained
Limit To constrain
large downwards
fluctuations



*Assuming a 200 GeV
WIMP and a best-fit $\sigma = 2.5$
 $\times 10^{-47}$ cm²

Comparison with other LXe Experiments



ER Background:

PandaX-4T PRL 129, 161804 (2022)

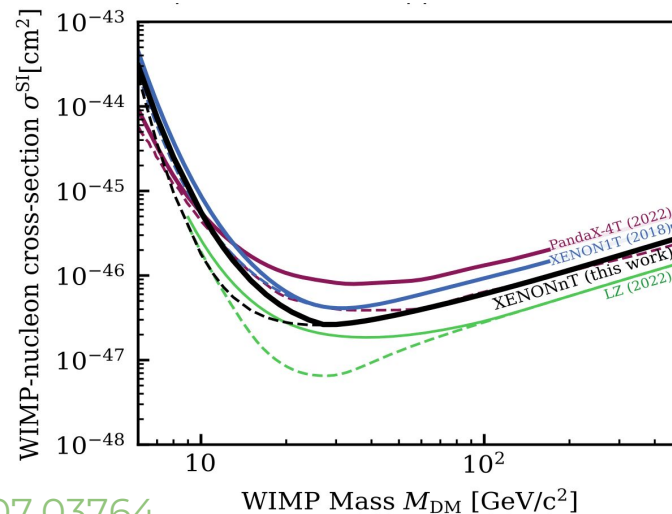
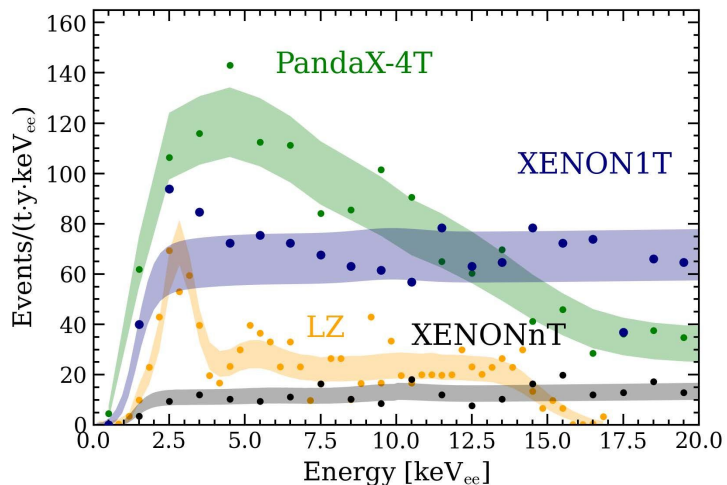
XENON1T PRD 102, 072004 (2020)

LZ arXiv:2207.03764

XENONnT PRL 129, 161805 (2022)

Spin-Independent WIMP-Nucleon cross section:

Same conservative power-constraint applied to results of other LXe experiments from non-blind analyses:



LZ, arXiv:2207.03764

XENON1T, PRL 121, 111302 (2018)

PandaX-4T, PRL 127, 261802 (2021)

Summary and Outlook



We performed a search for new physics exploiting **electronic and nuclear recoil data from XENONnT**

The average ER background rate of $(15.8 \pm 1.3 \text{ stat})$ events/(t·y·keV) in the (1, 30) keV energy region is the **lowest ever achieved in a DM search experiment**

Blind analysis shows **no excess above the background**, excluding BSM interpretations of the XENONIT excess.

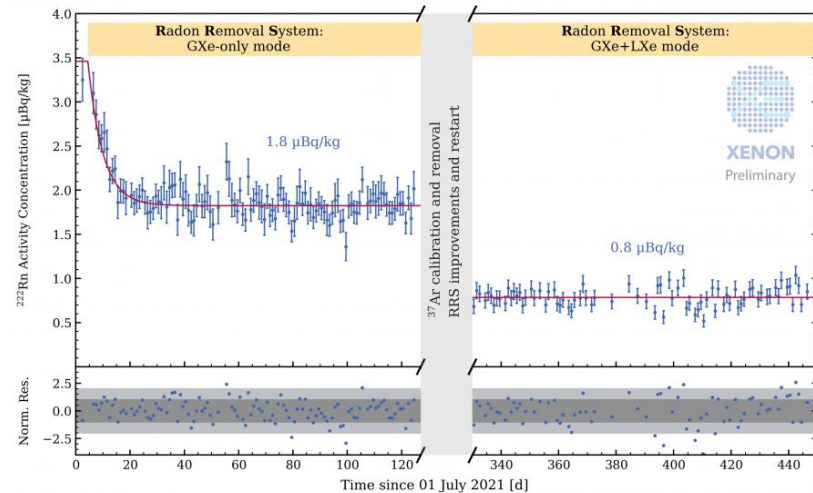
Best fit for WIMP search indicates **no significant excess** we therefore set a new limit for SI WIMP cross section:

$$2.6 \cdot 10^{-47} \text{ cm}^2 \text{ at } 28 \text{ GeV}/c^2 \text{ and } 90\% \text{ CL}$$

Thank You!

What's next:

- Ongoing science run (SR1) with factor 2 lower radon level
- Addition of Gd salt to the Neutron veto





BACKUP

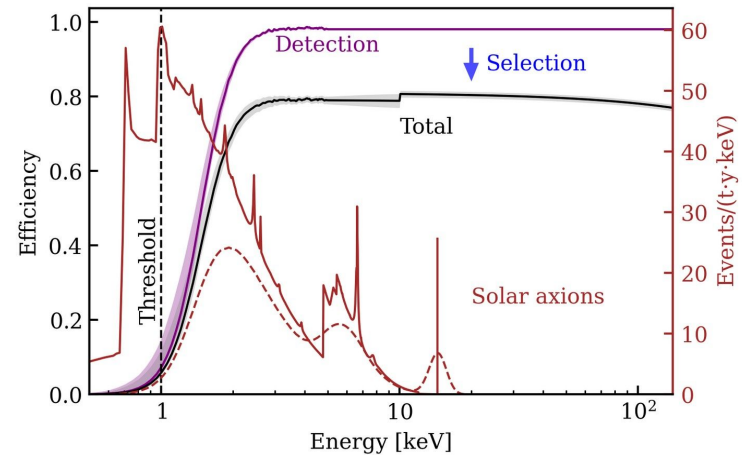
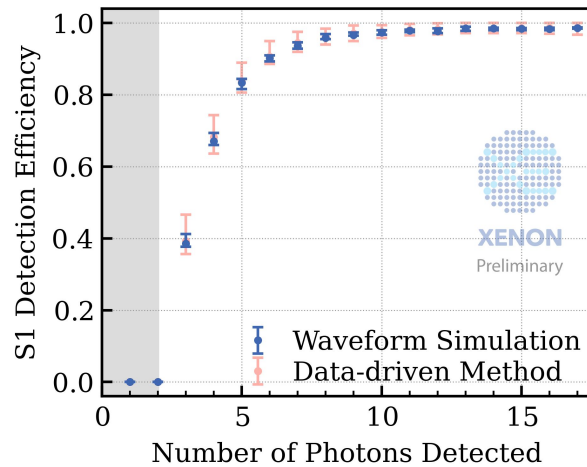
Efficiencies



Detection efficiency validated using **Monte Carlo** and **Data-Driven** methods

Good agreement between the two approaches

Total efficiency takes also **event-selection efficiency** into account

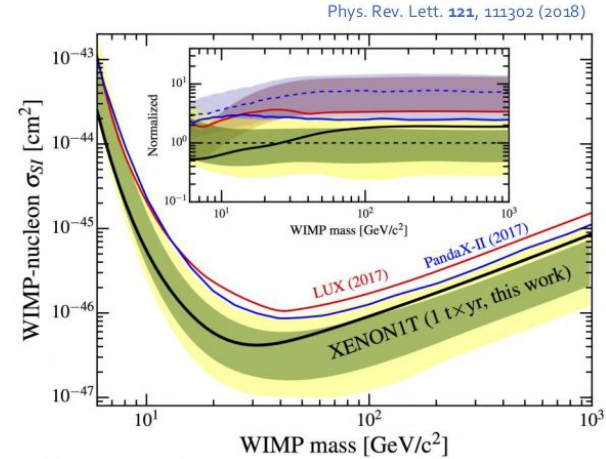
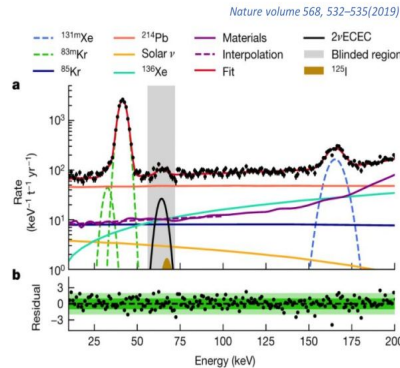
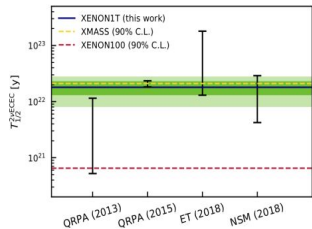
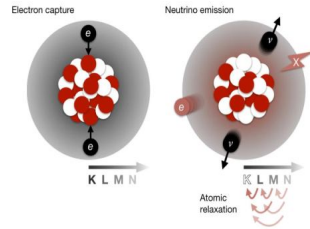


Results from XENONIT



Upper Limit on WIMP-nucleon cross section:

$$4.1 \times 10^{-47} \text{ cm}^2 \text{ @ } 30 \text{ GeV}/c^2 \text{ (90\% CL)}$$



Half-life of **Double Electron Capture in ^{124}Xe**

Longest half-life ever observed:

$$1.8 \times 10^{22} \text{ yr}$$

(4.4 σ Significance)

Tritium Enhanced Dataset (TED)



To Exclude presence of Tritium traces in the detector, XENONnT was operated **bypassing the GXe purification** for 14.3 days at the end of SRO

This would enhance the HT concentration in LXe by a factor 10-100

Data collected in this TED mode were **blinded**

After the unblinding, **no evidence was found for a tritium-like excess**

Tritium is therefore not included in the background model.

